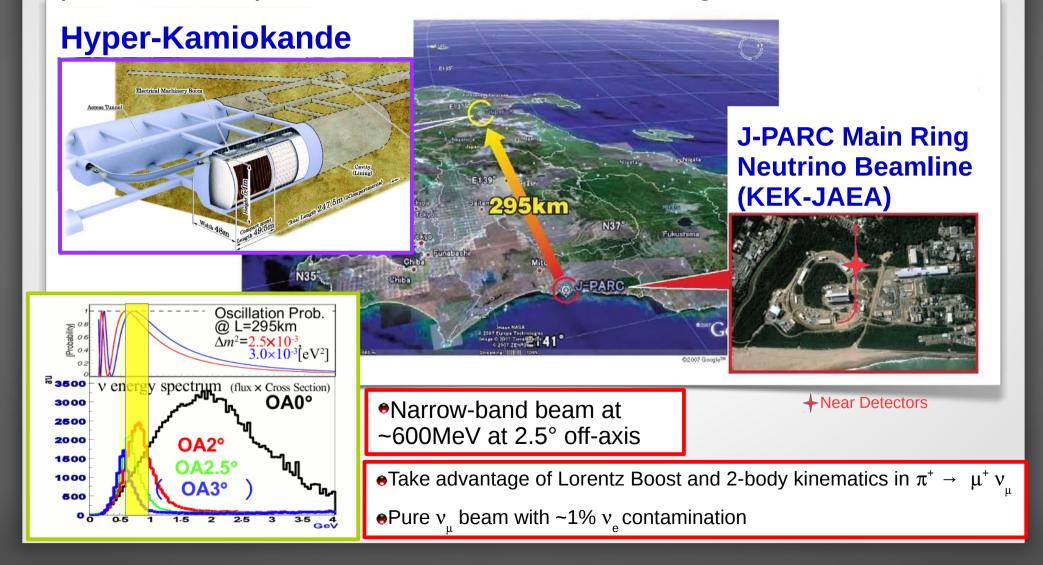




## Tokai to Hyper-Kamiokande

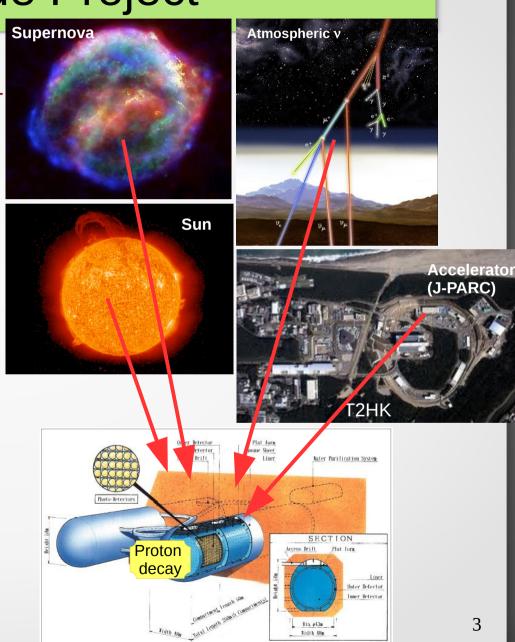
•Use upgraded J-PARC neutrino beam line (same as T2K) with expected beam power ≥750kW, 2.5° off-axis angle.



## The Hyper-Kamiokande Project

Multi-purpose neutrino experiment. Wide-variety of scientific goals:

- •Neutrino oscillations:
  - Neutrino beam from J-PARC
  - Atmospheric neutrinos
  - Solar neutrinos
- Search for proton decay
- Astrophysical neutrinos
   (supernova bursts, supernova relic neutrinos, dark matter, solar flare, ...)
- Neutrino geophysics



### Hyper-K in the World

( http://www.hyperk.org)

- Selected one of the 27 'top projects' in the 'Japanese master plan for large scale research projects' by the Science Council of Japan.
- Open meetings held twice/year since Aug.2012
- International Board of Representatives formed to discuss contributions, cost-sharing and budget request.
- R&D work ongoing internationally.







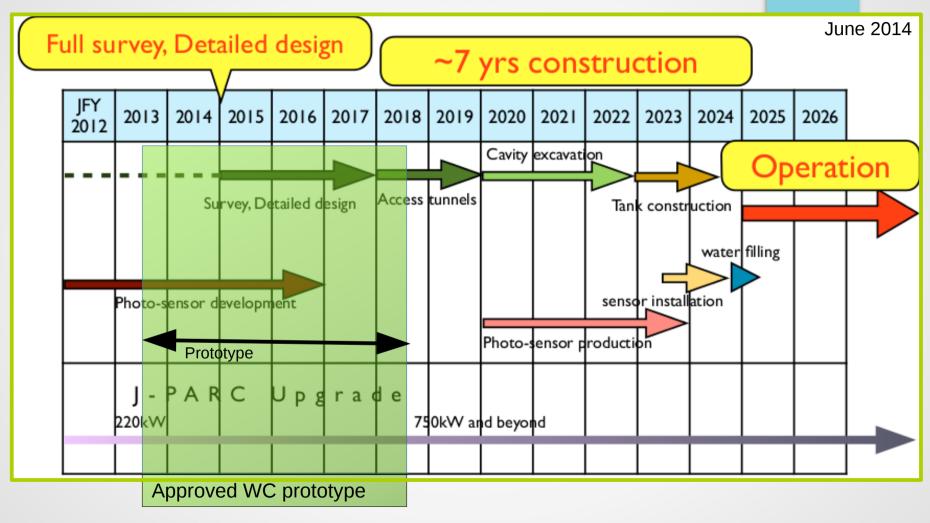






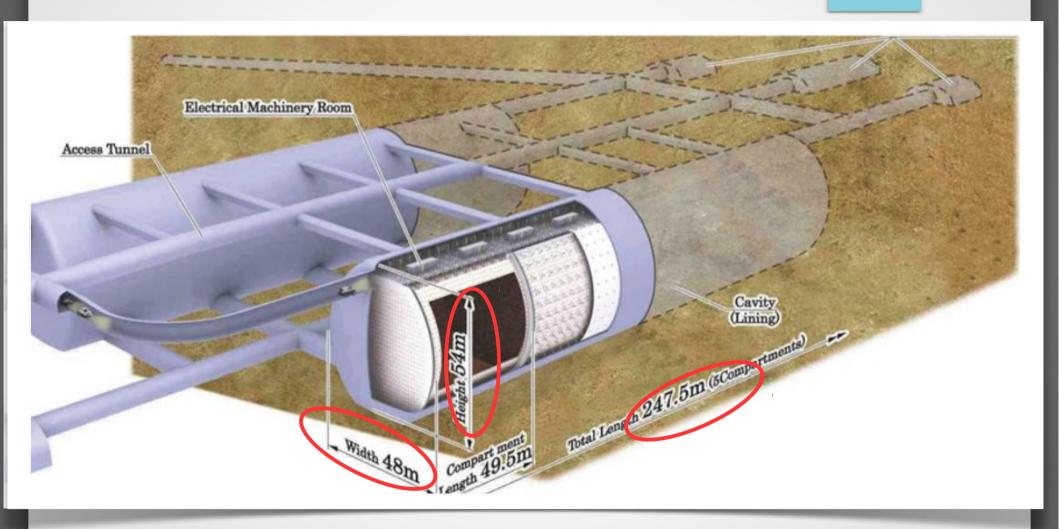


## The Hyper-Kamiokande Timeline



- Approved WC prototype for technical studies (2013-2018) ~\$1.2M
- Major decisions on experiment design expected before 2018.

## The Hyper-Kamiokande Detector



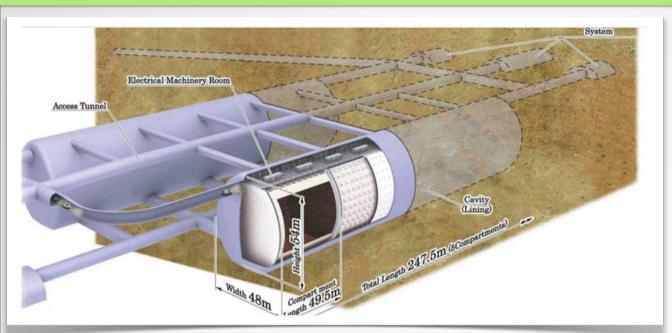
## The Hyper-Kamiokande Detector

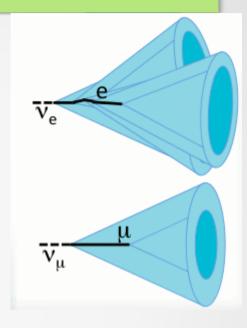


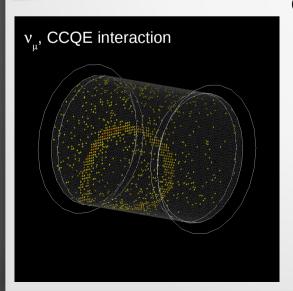
- Water Cherenkov, proven technology & scalability:
  - Excellent PID at sub-GeV region >99%
  - Large mass → statistics always critical for any measurements.

Total Volume	0.99 Megaton
Inner Volume	0.74 Mton
Fiducial Volume	0.56 Mton (0.056 Mton $\times$ 10 compartments)
Outer Volume	0.2 Megaton
Photo-sensors	<ul> <li>•99,000 20"Ф PMTs for Inner Detector (ID)</li> <li>(20% photo-coverage)</li> <li>•25,000 8"Ф PMTs for Outer Detector (OD)</li> </ul>
Tanks	<ul> <li>•2 tanks, with egg-shape cross section ≈</li> <li>48m (w) × 50m (t) × 250 m (l)</li> <li>•5 optically separated compartments per tank</li> </ul>

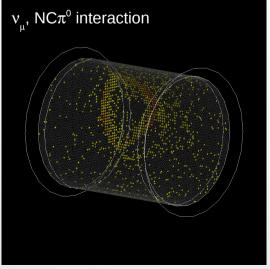
## The Hyper-Kamiokande Detector

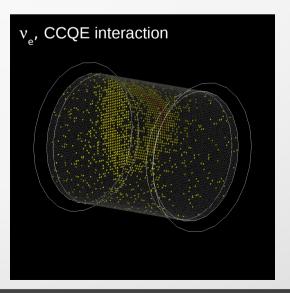






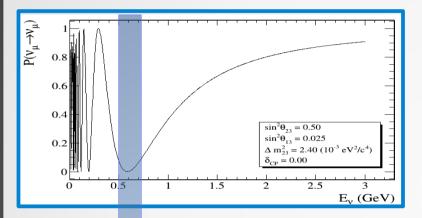
**GEANT4** event displays

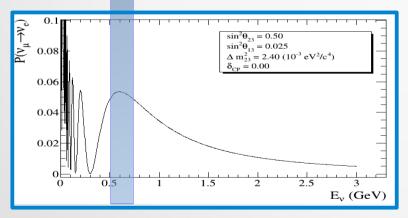




## Oscillation Searches at Hyper-K

HK is optimized for both appearance and disappearance searches





T2K v beam energy peak

ν<sub>μ</sub> Disappearance: determine θ<sub>23</sub> and Δm<sub>32</sub><sup>2</sup>

$$P(ν_μ → ν_μ) ≈ 1 - sin^2 2 θ_{32} sin^2 \left( \frac{\Delta m_{23}^2 L}{4 E_ν} \right)$$

 $v_e$  Appearance: determine  $\theta_{13}$ , constrain  $\delta_{CP}$ 

$$P(\nu_{\mu} \to \nu_{e}) \approx \sin^{2}\theta_{23} \sin^{2}2\theta_{13} \sin^{2}\left(\frac{\Delta m_{31}^{2} L}{4 E_{\nu}}\right)$$

$$-\sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin^{2}\left(\frac{\Delta m_{32}^{2} L}{4 E_{\nu}}\right)$$

$$\sin^{2}\left(\frac{\Delta m_{31}^{2} L}{4 E_{\nu}}\right) \sin^{2}\left(\frac{\Delta m_{21}^{2} L}{4 E_{\nu}}\right) \sin \delta_{CP} + CPC$$

$$\sin^2\left(\frac{\Delta m_{31}^2 L}{4 E_v}\right) \sin^2\left(\frac{\Delta m_{21}^2 L}{4 E_v}\right) \sin\delta_{CP} + CPC$$

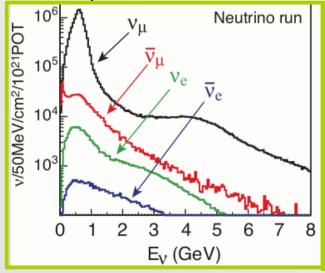
+matter + solar terms

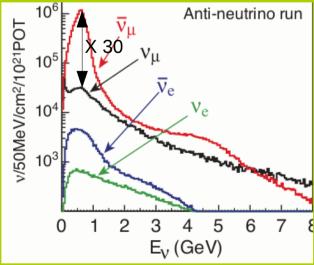
For maximum power fit both data samples jointly

## Neutrino Beam for the Experiment

- Same beam configuration as for T2K
  - > Same off-axis sub-GeV narrow-band beam.
  - Horn current 320kA
- At least 750kW expected at the starting of the experiment.
- •Assumed **7.5MW**  $\times$  **10**<sup>7</sup> **s** (1.56  $\times$  10<sup>22</sup> POT) for the following sensitivity studies
  - ▶ 10 years are needed if 750kW per 10<sup>7</sup>s/year
  - > Higher beam power is under study

Expected unoscillated neutrino flux at Hyper-K





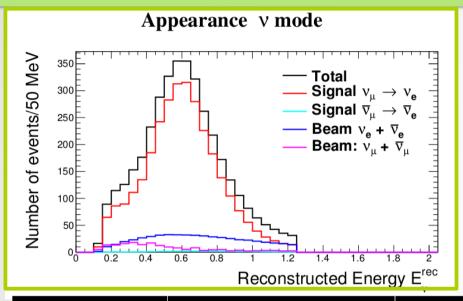
Nominal beam sharing between neutrinos and anti-neutrinos in the following sensitivity plots:

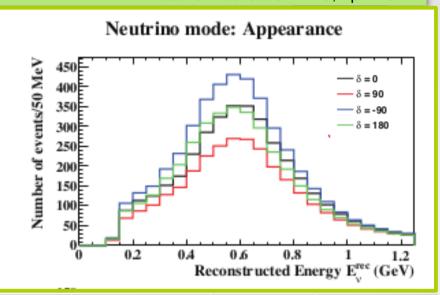
v-mode: v-mode

1y:3y

### **Expected Events**

Letter of Intent to J-PARC, April 2014



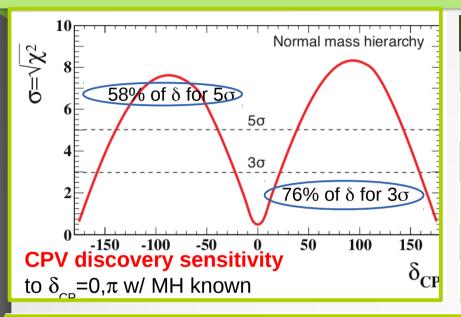


Appearance	Sig	nal		Bac	kgrour	nd		Total
	$V_{\mu} \rightarrow V_{e}$	$\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$	$V_{\mu}$	$\overline{\mathbf{v}}_{\mu}$	$v_{\sf e}$	$\overline{v}_{\mathrm{e}}$	NC	
v mode (	3016	28	11	0	503	20	172	3750
$\overline{v}$ mode	396	2110	4	5	222	265	265	3397
Disappearanc	e v <sub>µ</sub>	$\overline{\overline{v}}_{\mathfrak{\mu}}$	$v_{\rm e}$	$\overline{v}_{e}$		NC	$V_{\mu} \rightarrow V_{e}$	Total
v mode	17225	1088	11	1		999	49	19372
$\overline{v}$ mode	10066	15597	7	7		1281	6	26964

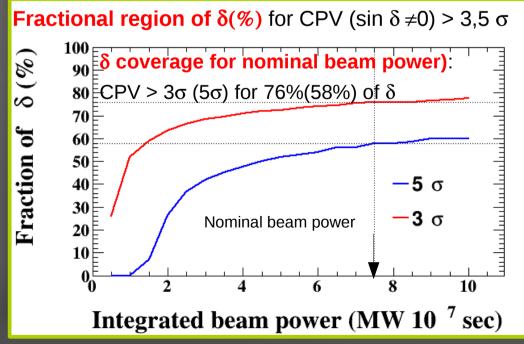
Large expected number of events. NH,  $\sin^2 2\theta_{13} = 0.1$  and  $\delta_{CP} = 0$ 

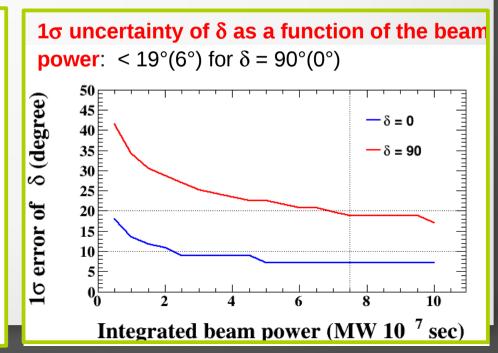
## Hyper-K Sensitivity to $\delta_{_{\text{CP}}}$

Letter of Intent to J-PARC, April 2014



Errors (%) on the expected number of events					
	v mode		$\overline{v}$ mod	le	
	$v_{\rm e}$	${f V}_{\mu}$	$v_{\rm e}$	${f V}_{\mu}$	
Flux & ND	3.0	2.8	5.6	4.2	
ND-independ. xsect	1.2	1.5	2.0	1.4	
Far Detector	0.7	1.0	1.7	1.1	
Total	3.3	3.3	6.2	4.5	

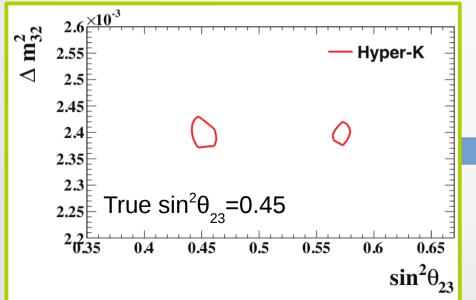


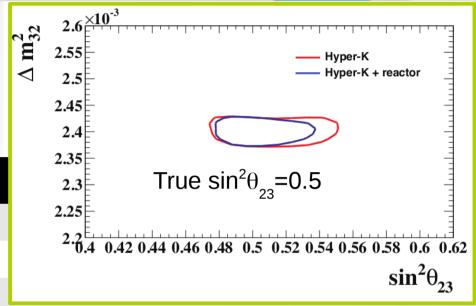


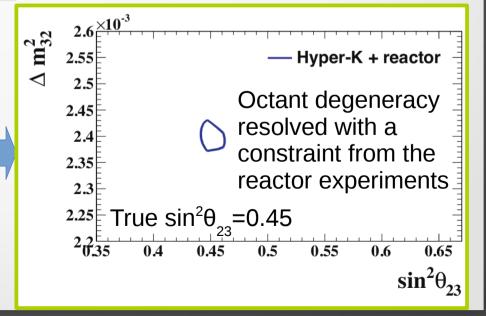
## Sensitivity to $\theta_{23}$

- $\sin^2 2\theta_{23}$  and  $\Delta m^2_{23}$  free parameters as well as  $\sin^2 2\theta_{13}$  and  $\delta_{CP}$  in the fit.
- Octant resolution w/ reactor  $\theta_{13}$ : ~3 $\sigma$  wrong octact rejection for  $\sin^2\theta_{23}$ <0.46 or >0.56

True $\sin^2\theta_{23}$	$1\sigma  err  sin^2\theta_{_{23}}$	$1\sigma \operatorname{err} \Delta m^2_{23} (eV^2)$
0.45	0.006	1.4
0.50	0.015	1.4
0.55	0.009	1.5



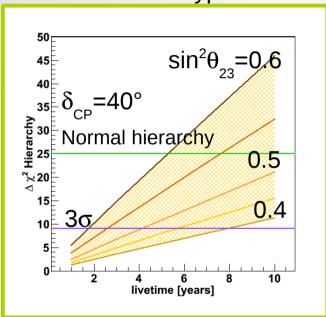


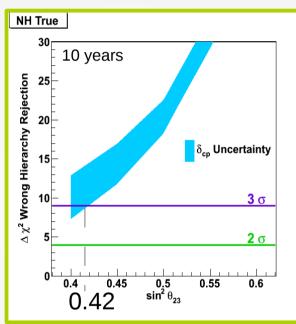


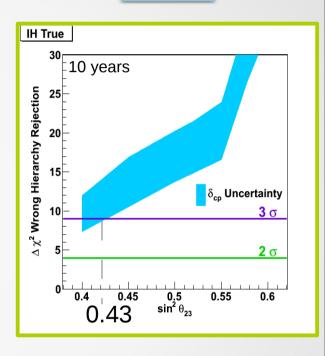
## Hyper-K Sensitivity to MH

arXiv:1109.3262

Significance for MH determination as a function of Hyper-K lifetime



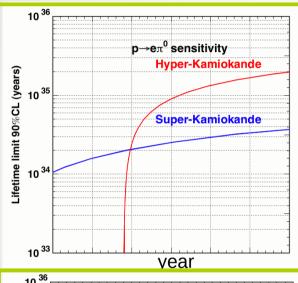


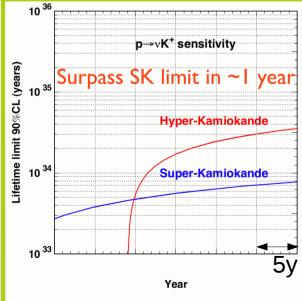


- Use atmospherics for  $3\sigma$  mass hierarchy determination.
- $3\sigma$  mass hierarchy determination for  $\sin^2\theta_{23} > 0.42$  (0.43) for normal (inverted) hierarchy for 10y data taking.
- Also combine with beam data to enhance physics capability.

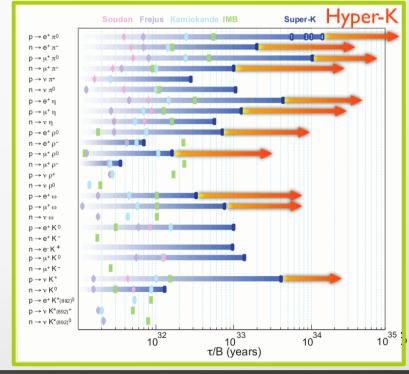
## **Proton Decay Sensitivity**

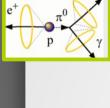
#### Surpass SK limit in ~1 year





- •10 times better sensitivity than Super-K
- Hyper-K surpasses SK limits in ~1y
- Hyper-K is sensitive in every single mode
  - $^{\flat}$  p→e $^{+}$ π $^{0}$ : 1.3 × 10 $^{35}$  y at 90% CL
  - $p \to v K^+$ : 2.5 × 10<sup>34</sup> y at 90% CL
  - Many other modes:
    - $p(n \rightarrow e,\mu) + (\pi,\rho,\omega,\eta); 10^{14}-10^{35}$
    - K<sup>0</sup> modes
    - $v\pi^{0}$ ,  $v\pi^{+}$
    - ....







## 'Other' Physics Topics at Hyper-K

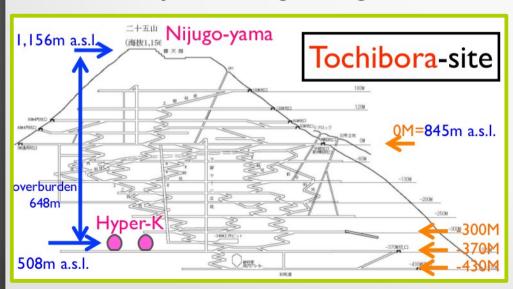
More physics topics can be investigated by Hyper-Kamiokande:

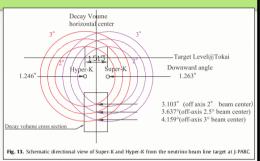
- •Solar Neutrinos: 200 v's / day from Sun  $\rightarrow$  day/night asymmetry of the solar neutrinos flux can be precisely measured at HK.
- Astrophysical neutrinos:
  - 200k ν's from Supernova at Galactic center (10kpc)
    - → time variation & energy can be measured with high statistics. Important data to cross check explosion models
  - Supernova relic neutrinos → possible G<sub>d</sub>-doping of Hyper-K
- •Indirect Searches for Dark Matter: 1) search for excess of neutrinos from the center of the Earth, Sun and galactic centre as compared to atmospheric neutrino background 2) Search for diffuse signal from Milky Way halo.
- •Search for transient astrophysical phenomena: solar flares, GRBs, etc.
- •Neutrino geophysics: neutrino radiography w/ atmospheric neutrinos for surveying the internal structure of the Earth.

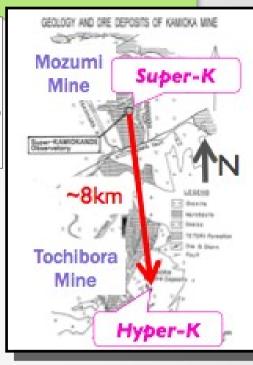
## Site(s) and Cavern(s)

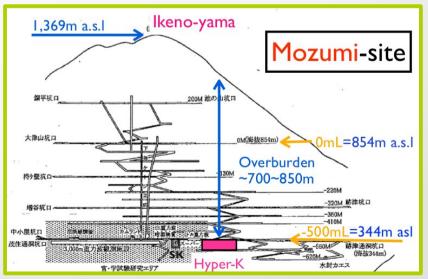
Two sites are being investigated:

- Tochibora mine:
  - ~8km South from Super-K
  - Identical baseline (295km) and off-axis angle (2.5°) to Super-Kamiokande
- Mozumi mine (same as Super-K)
  - Deeper than Tochibora
  - Currently finishing design studies







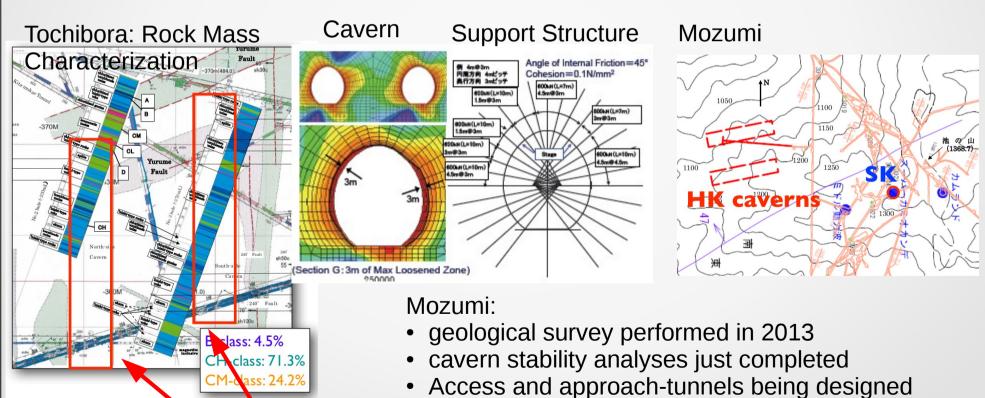


## Site(s) and Cavern(s), cont'd

Rock quality in the two sites is similar.

HK tank location

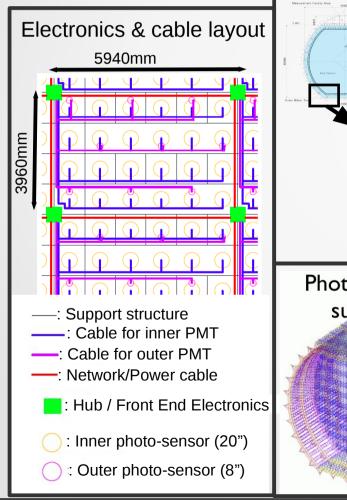
- Design of the cavity, support structure studies based on geological survey
- Confirmed that the HK cavern can be constructed w/ existing techniques
- Construction schedule for the Tochibora mine: ~2y tunnels, ~3y cavern

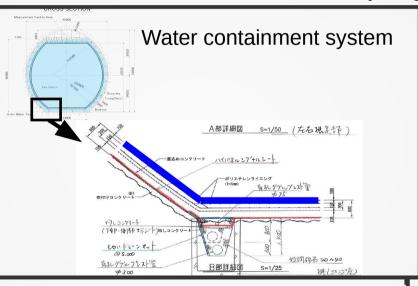


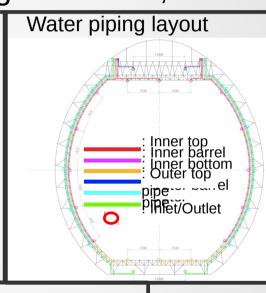
Baseline design expected to finish soon

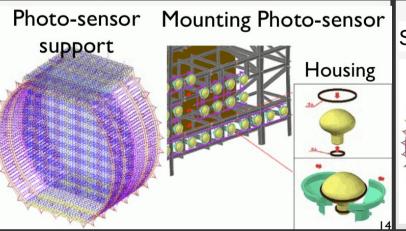
## Tank Design Work

•All major parts of HK tank have been designed: water containment system, photosensors support, layout of water pipes, front-end electronics, cables, calibration holes, plug manholes, etc.









## Photodector Development

#### **PMT**

- Used in Super-K for 18 years
- High reliability



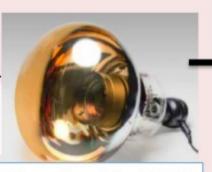
High QE photocathode

Venetian blind PMT (50-cm φ Normal QE)

New photosensor (HPD)

#### **Under viability test**

Test in 200 ton tank



Venetian blind PMT (50-cm φ High QE)



HPD (New ) (20-cm φ Normal QE)

# Under development

 Under evaluation in the air



Box and Line PMT (50-cm φ High QE)

High QE photoca<mark>t</mark>hode

PMT

improvement

Make larger



HPD 50-cm φ High QE)

## Photodector Development

#### **PMT**

- Used in Super-K for 18 years
- High reliability



High QE photocathe

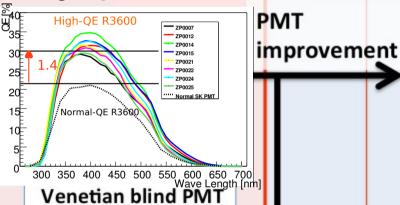
Venetian blind PMT (50-cm φ Normal QE)

New photosensor (HPD)

#### **Under viability test**

Test in 200 ton tank

High-QE SK PMT



(50-cm φ High QE)

#### Under development

Under evaluation in the air



Box and Line PMT (50-cm Φ High QE)



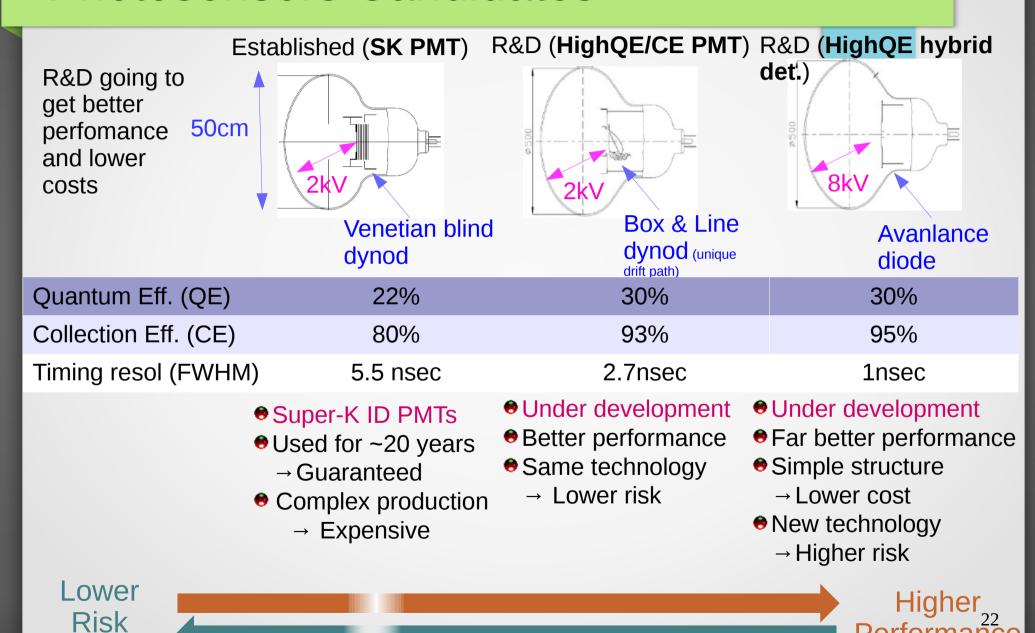
HPD (New) (20-cm φ Normal QE) High QE photocathode

Make larger

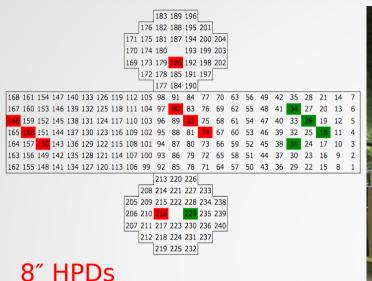


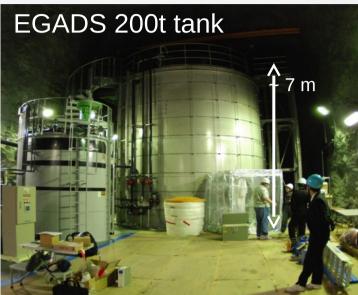
**HPD** 50-cm φ High QE)

### **Photosensors Candidates**



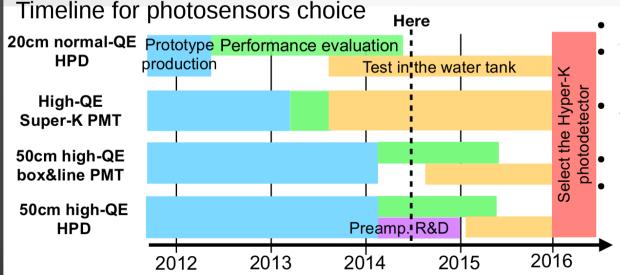
#### Photosensors Tests in Water Tank





- EGADS (for G<sub>d</sub>-doped water tests)
- 240 inward-facing PMTs
- EGADS used to test high-QE PMTs
- 227 PMTs (R3600; currently in SK) for reference for photodetector evaluation
- 8" HPDs, 20" high-QE PMTs

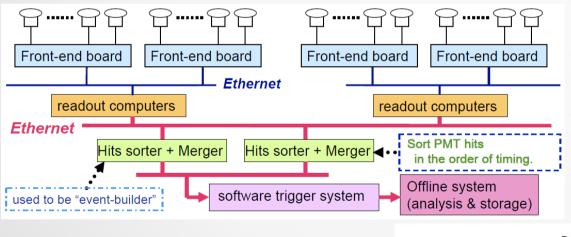
20" high-QE PMTs



- Data taking: Sept 2013, May 2014
- Viability tests performed ongoing process up to 2016.
- Adding (Aug 2014) Box-and-Line PMTs and 2 HPD.
- More tests planned.
  - Photosensor choice will be made in 2016, needed to allow time for making mass production

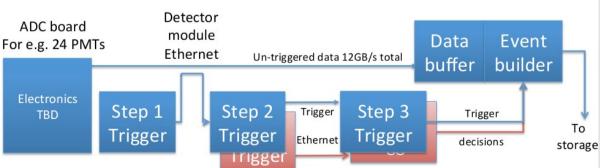
### Electronics/DAQ

- Investigating a few approches for the electronics, eg:
  - > QTC (ADC) + TDC (similar to SK4)
  - > FADC
- •Will evaluate their performance with the WC prototype detector
- Also working on the design of the DAQ



- Digitize all the signal (timing and charge) above ~ 1/4 p.e.
- Define events with software and store the event data.
- Nominal starting point: SK DAQ

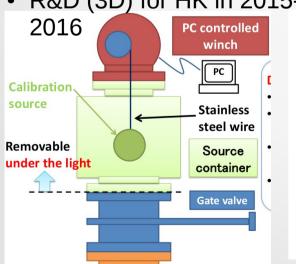
Investigating improved system for triggering (redundant, robust)



### Calibration

- •Review systems used by several experiments (SK, SNO, SNO+, Borexino, KamLAND, Daya Bay) to help in the design of the calibration system for Hyper-K
- Several ongoing R&D activities, some examples:
- Simple semi-automated calibration system (to be deployed in SK)
- Computed controlled.
- Compact and lightshielded.

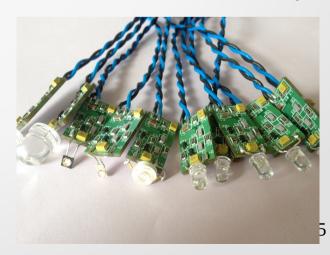
R&D (3D) for HK in 2015-



- Study response & reflection of large photosensors in water (Photosensor Testing Facility at TRIUMF)
- Optical system with laser, monitor and receiver PMTs in place and tested.



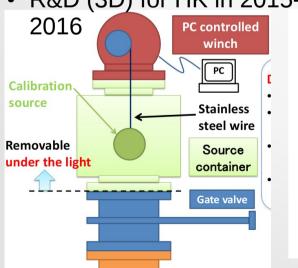
- Use LED as a light source for optical calibration.
- Can build an automated system that can illuminate each PMT with known sources
- Tests of LEDs underway



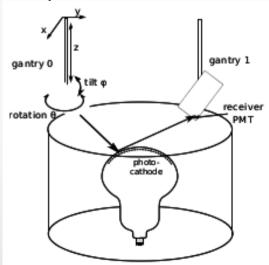
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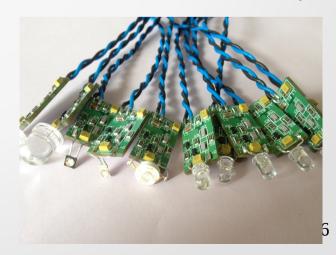
R&D (3D) for HK in 2015-



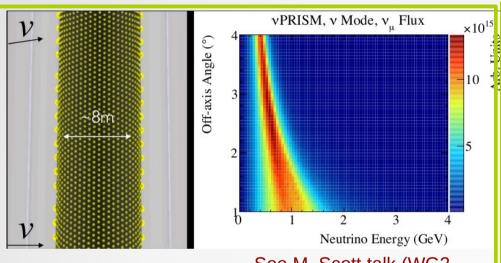
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- Tests of LEDs underway



## New Near Detector Concepts to further enhance the Hyper-K physics potential

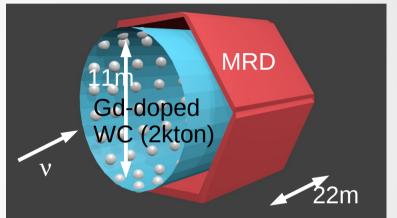


See M. Scott talk (WG2 "ν-PRISM" (~1km) parallel session - Monday)

- tall (~50 m) WC detector spanning wide range of off-axis angles
- effectively isolate response in narrow band of energy by comparing interactions at different off-axis angles

"TITUS" (~2 km)

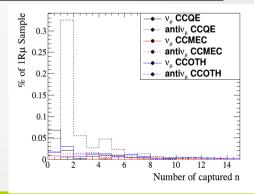
 2 kt Gadolinium-doped WC detector with HPDs and LAPPDs



Muon range detector (MRD)

 $G_d$  for v/v discrimination

See M. Rayner poster (Tuesday)

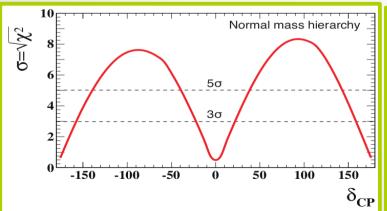


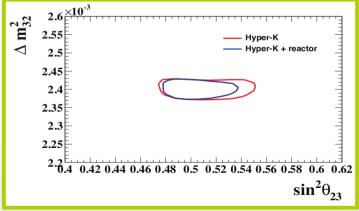
Several new upgrades planned at a 280m distance, either new detectors or upgrading ND280

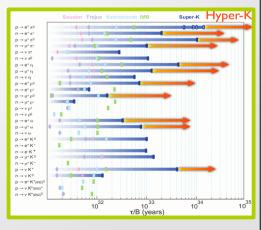
#### Conclusions

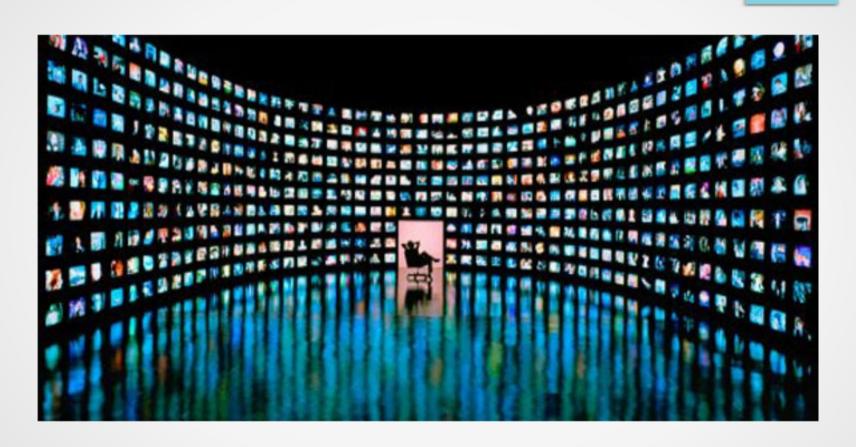


- Next generation multi-purpose experiment
  - Oscillation physics:
    - able to measure  $\delta_{_{\text{CP}}}$  at  $3\sigma$  for 76% of its phase space
    - solve octant degeneracy, mass hierarchy (atmospherics),  $\theta_{32}$ ,  $\Delta m_{32}^2$
  - Astro and other physics:
    - very sensitive to all the proton decay channels, observe supernovas burst and relic supernova neutrinos, indirect dark matter, transient astrophysical phenomena, etc.
- Work ongoing worldwide in all the aspects of the experiment
- Data taking around 2025 with current schedule





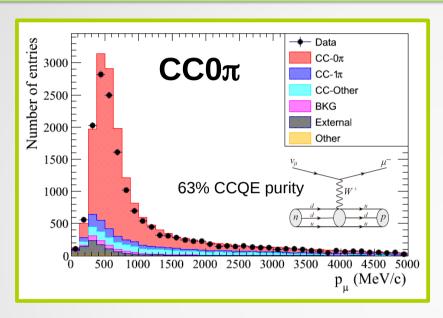


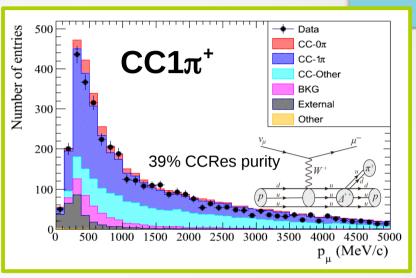


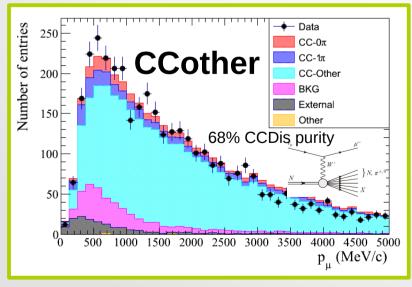
Stay tuned for many more exciting news from Hyper-Kamiokande!

## **Additional Slides**

### Muon Momentum in ND280



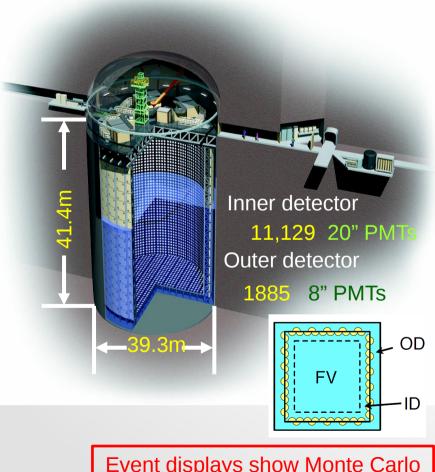




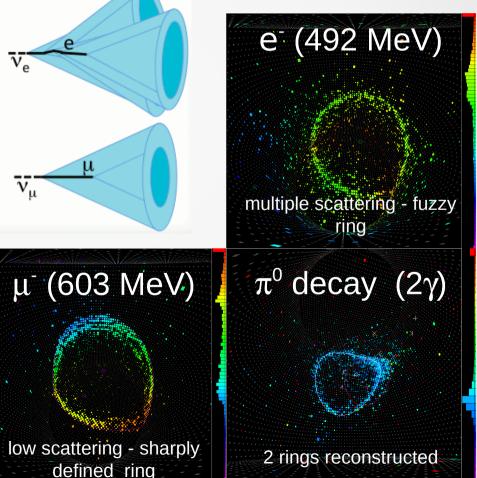
	Purity of each sample						
	CC0π	CC1π	CCother				
CC0π	72.6%	6.4%	5.8%				
CC1π	8.6%	49.4%	7.8%				
CCother	11.4%	31%	73.8%				
Bkg(NC+anti-nu)	2.3%	6.8%	8.7%				
Out of FGD1 Fid Vol	5.1%	6.5%	3.9%				

### T2K Far Detector: Super-Kamiokande

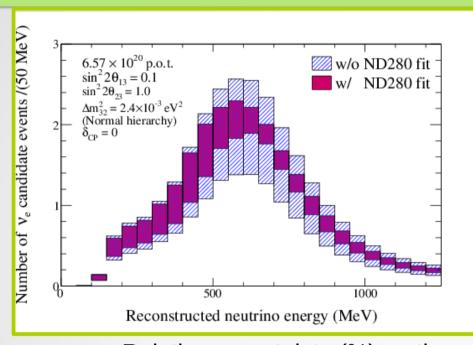
- 50 kton (22.5 kton fiducial) water Cherenkov detector
- Good reconstruction for T2K energy range
- Particle Identification (PID) based on shape of Cherenkov rings

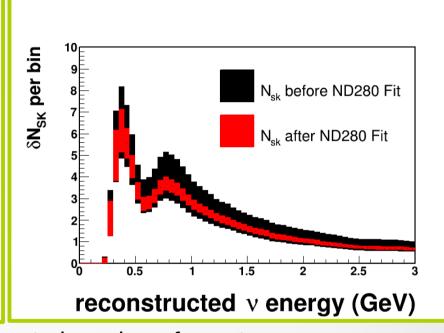






### Results Systematic Errors w/ ND280 Fit





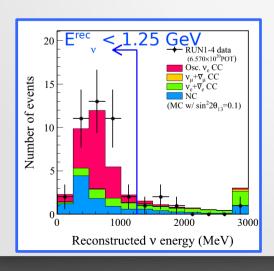
Relative uncertainty (%) on the expected number of events  $(\sin^2 2\theta_{13} = 0.1, \sin^2 \theta_{23} = 0.5, \Delta m_{32}^2 = 2.4 \times 10^{-3} \text{eV}^2, \delta \text{CP} = 0, \text{NH})$ 

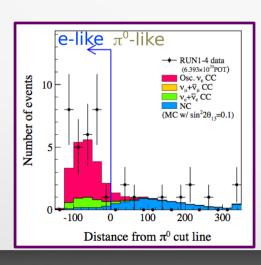
V <sub>e</sub>	Systematic sources	$V_{\mu}$
3.1	Flux & Combined Cross-Sections	2.7
4.7	Independent Cross Sections	5.0
2.4	$\pi$ Hadronic Interactions (FSI)	3.0
2.7	SK Detector Efficiencies	4.0
6.8	TOTAL	7.6

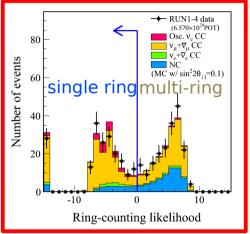
## $\nu_{\mu} \rightarrow \nu_{e}$ Event Selection

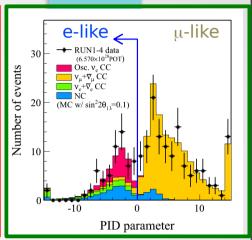
#### **Event selection:**

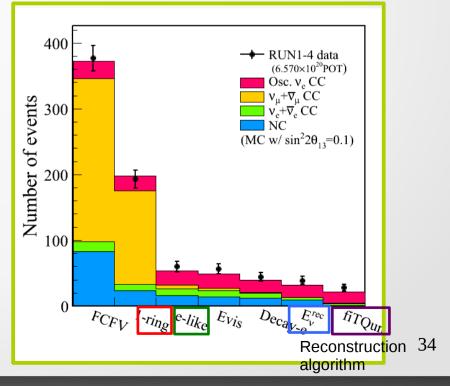
- Fully contained in fid. volume
- Only one reconstructed ring
- Ring is electron-like
- Visible energy > 100MeV
- No Michel Electrons
- Reconstructed energy < 1.25 GeV</li>
- New SK reconstruction (~30% reduction in  $\pi^0$  background)
- 28 events in 6.57× 10<sup>20</sup> POT



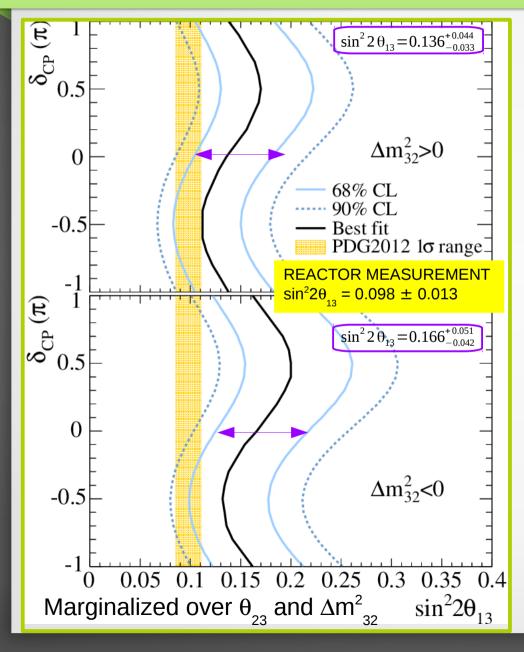




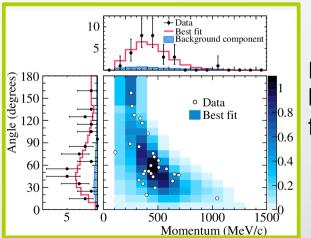




## T2K Observation of $v_e$ Appearance



- $\bullet \nu_{_{e}}$  appearance in a  $\nu_{_{\mu}}$  beam with
- 7.3σ significance
- •Best fit value for  $\sin^2 2\theta_{13}$  larger than the reactor value
- •Due to  $\delta_{\text{CP}}$ -sin<sup>2</sup> $\theta_{13}$  correlation, when applying the reactor constraint, region with sin<sup>2</sup>2 $\theta_{13}$  small as possible is favoured.



Maximum likelihood fit in  $(p_e, \theta_e)$ 

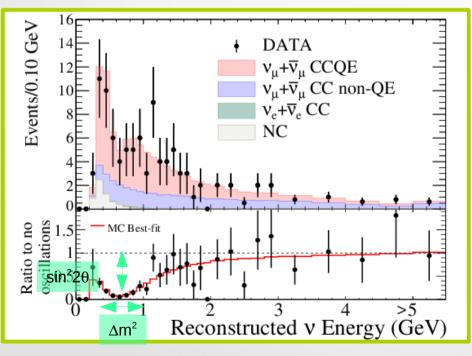
# $V_{\mu} \rightarrow V_{\mu}$ T2K Result

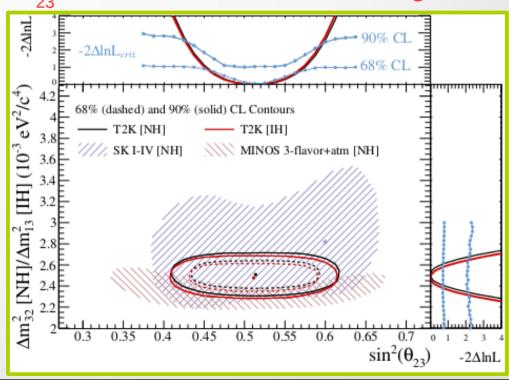
• Best-fit oscillation parameter values:

Oscillation parameter	NH	IH
$\sin^2 \theta_{23}$	$0.514^{+0.055}_{-0.056}$	$0.511 \pm 0.055$
$\Delta m_{_{32}}^{2} (\times 10^{-3} \text{ eV}^{2})$	$2.51 \pm 0.10$	$2.48 \pm 0.10$

• Events: 120 (observed), 446.0 ± 22.5 (no oscillation)

 $\bullet$  Most precise measurement of  $\sin^2\theta_{23}$  and favours Maximal Mixing



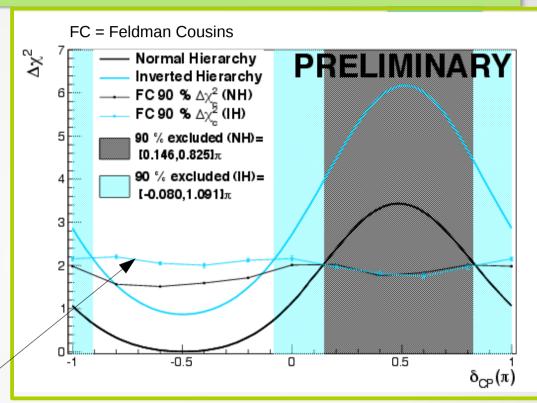


## T2K Joint $v_{\mu} + v_{e}$ Analysis

- •Likelihood ratio fit to both  $\nu_{_{u}}$
- $+ v_e$  event samples
- •Plot includes constraint from reactor experiments as given by the PDG2013 ( $\sin^2 2\theta_{13}$  =
- $0.095 \pm 0.01$ )
- Consistent results with a Bayesian analysis

 $\sin^2\!\theta_{_{23}}$ ,  $\Delta m^2_{_{32}}$  and  $\sin^2\!2\theta_{_{13}}$  are marginalized following the 3D  $\Delta\chi^2$  surface from Run1+2+3+4

T2K has a slight hint for  $\delta_{\rm CP} \sim -\pi/2$ 

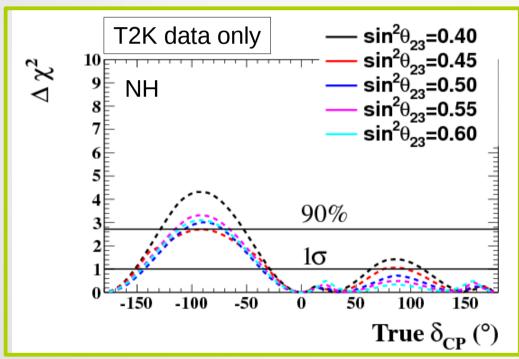


$\delta_{_{\mathrm{CP}}}$ excluded	Preliminary		
regions	Best fit	90%CL (π)	
NH	-0.495	[0.146;0.825]	
IH	-0.495	[-0.080;1.091]	

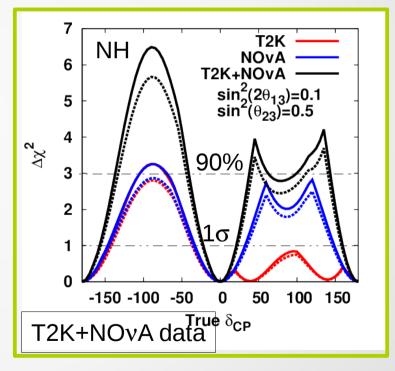
Future: Neutrino & Anti-neutrino running

## T2K Future Sensitivity to $\delta_{_{\text{CP}}}$

Sensitivity studies to resolve  $\sin\delta_{\rm CP}\neq 0$  with  $7.8\times 10^{21}$  POT. Best sensitivity expected for 50%v and 50% anti-v beam running. Projected reactor constraint  $\sin^2 2\theta_{13} = 0.1\pm 0.005$ .



Using joint oscillation analysis with realistic 2012 systematic errors (~10%  $v_{e}$ , ~13%  $v_{u}$ )



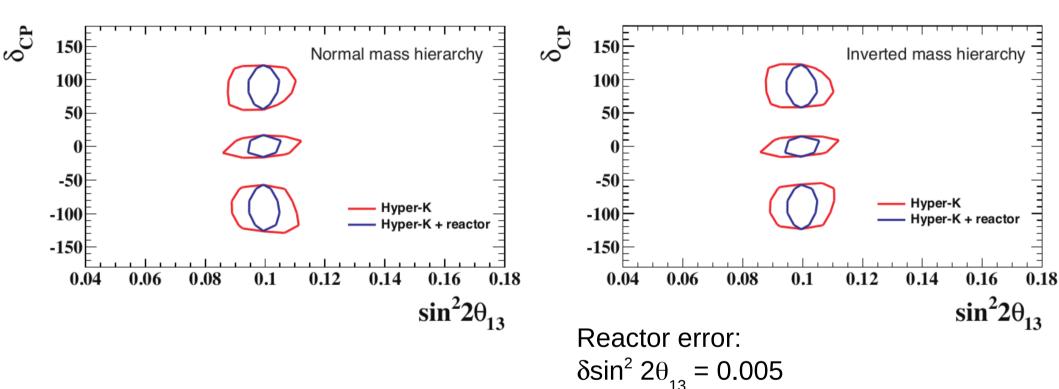
Using GloBES, with 50% v and anti-v running for NOvA too. Solid lines: no syst. error.

## J-PARC v-Beamline



# $\theta_{13}$ - $\delta_{CP}$ allowed region (90%)

 $7.5\text{MW} \times 10^7 \text{sec}$ , v:anti-v=1:3

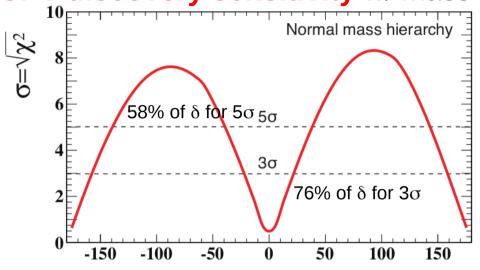


Parameter	sin²2θ <sub>13</sub>	$\delta_{_{ extsf{CP}}}$	$\sin^2 2\theta_{23}$	22		$\sin^2 2\theta_{12}$	$\Delta m_{12}^2$
Nominal (	0.10	90, 0, -90	0.50	_	hierarchy Nominal,	0.8704	7.6x10 <sup>-5</sup>
Treatment I	Fitted	Fixed	Fitted		Inverted Known	Fixed	Fixed 40

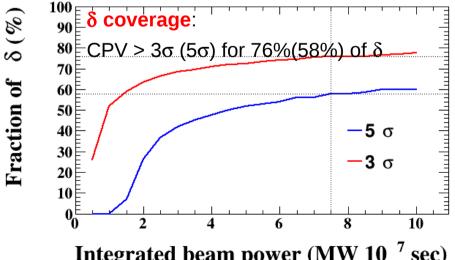
40

# Significance of $\delta_{CD} = 0$

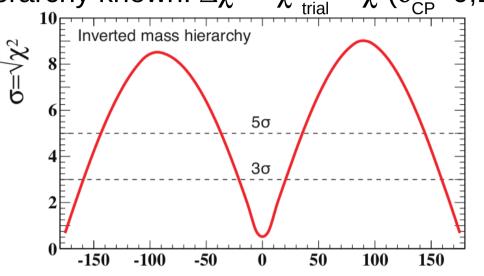
**CPV** discovery sensitivity w/ mass hierarchy known:  $\Delta \chi^2 = \chi^2_{trial} - \chi^2 (\delta_{CP} = 0.180)$ 



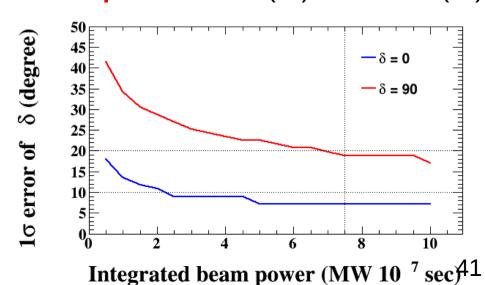
Fractional region of  $\delta(\%)$  for which the CPV ( $\sin \delta \neq 0$ ) significance is >  $3\sigma$ 



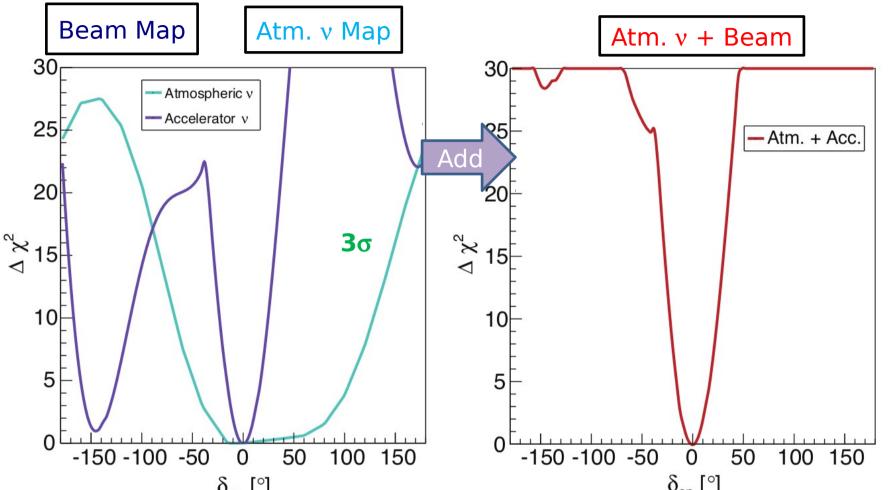
Integrated beam power (MW 10 <sup>7</sup> sec)



 $1\sigma$  uncertainty of  $\delta$  as a function of the **beam power**:  $< 19^{\circ}(6^{\circ})$  for  $\delta = 90^{\circ}(0^{\circ})$ 



# Beam + Atmospheric v: Allowed $\delta_{CP}$

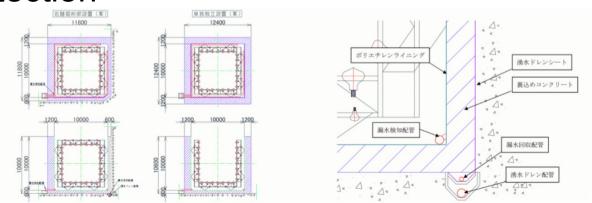


- Without assumption of the prior mass hierarchy knowledge, but NH is true.
- True  $\delta_{\rm CP}=0.0$ ;  $\sin^2\!2\theta_{_{13}}=0.10$ ; Maximal mixing  $\sin^2\!2\theta_{_{23}}=1.0$
- Degenerate solution exists at  $3\sigma$  in the beam only case.
- The physics capability of the project can be enhanced by combining two complementary measurements.

## 1kton WC Prototype

- •Prototype (1kton,  $\sim$ 10×10×10 m³) for R&D test approved in Japan as Grant-in-Aid:  $\sim$  USD 1.2M/5 years (2013-18).
- •It's one of the 27 proposals selected each year from all areas in science.
- •Main feasibility studies:
  - Photosensor and corresponding support structure
  - Liners
  - Leak water collection detection
  - DAQ
  - Electronics
  - Calibration system

>



### **Overall Cost Estimate**

Total	800M USD	
Cavern	300M USD	
Tank & structure	200M USD	
Photo-sensors	200M USD	High QE HPD
Near Detector	30M USD	@Tokai

- •Costs estimated based on the current design and including a new near detector.
- •Proportional sharing of costs between the interested Countries is expected.