



# Kometen kommer: Progress towards COMET Phase-I

-Phill Litchfield

## Recap of $\mu \rightarrow e$ conversion

### Phase-I and Phase-II

#### 'Frontend' components

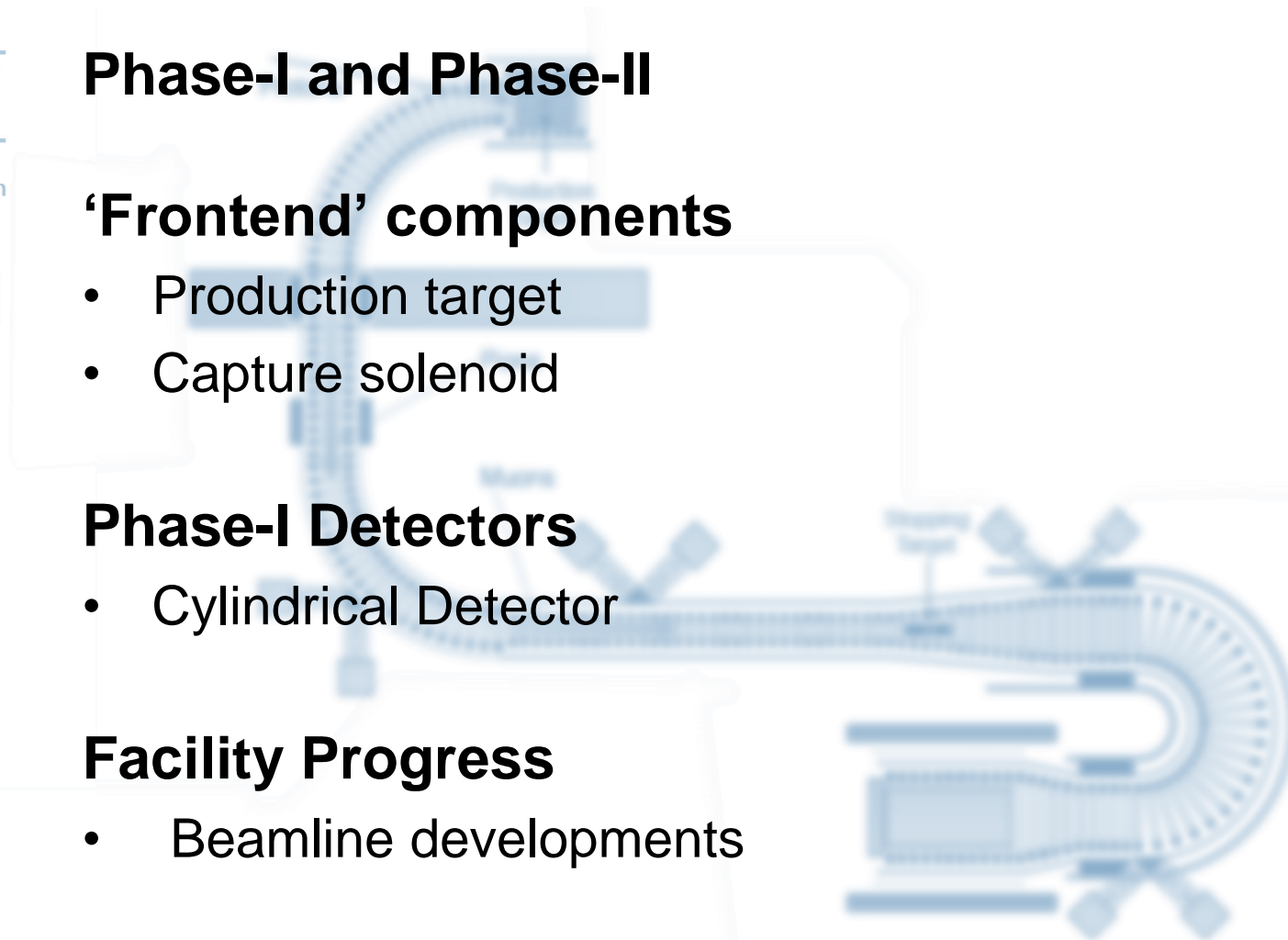
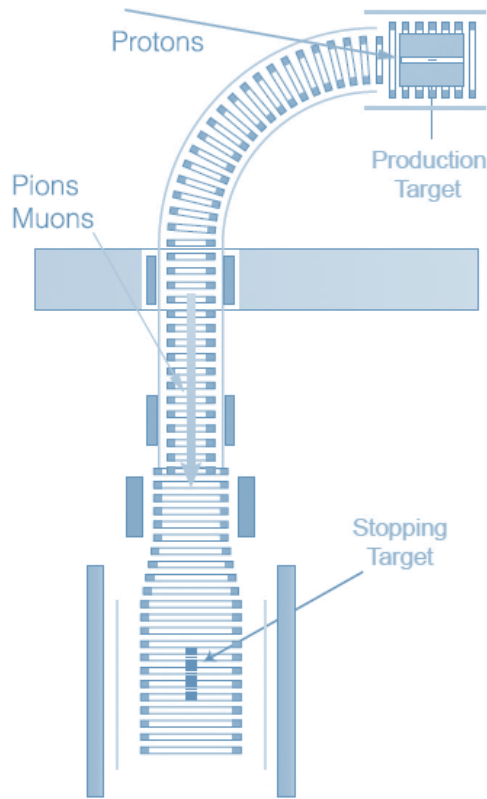
- Production target
- Capture solenoid

#### Phase-I Detectors

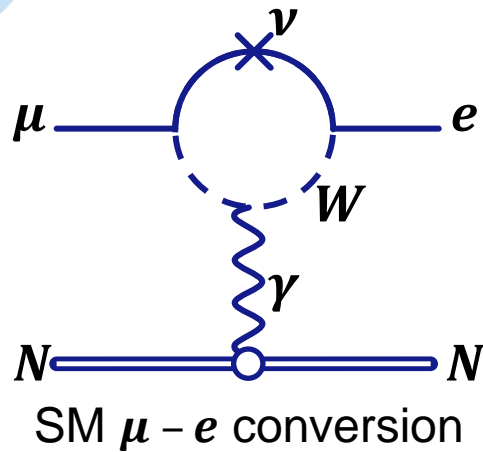
- Cylindrical Detector

#### Facility Progress

- Beamline developments

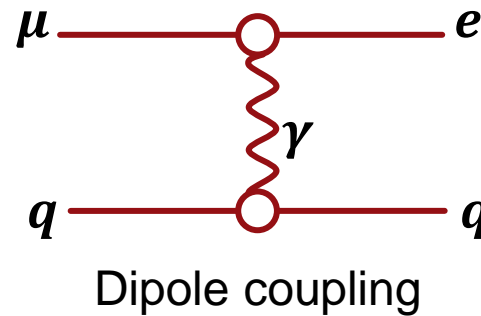
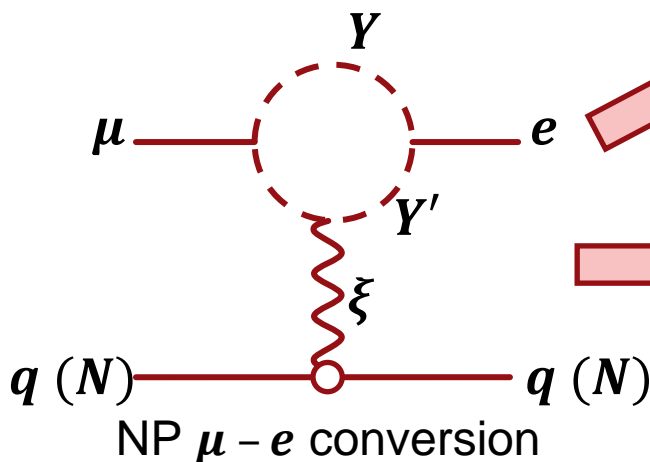


# $\mu$ to $e$ conversion

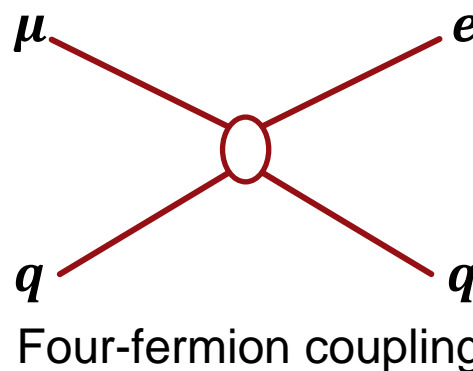


In the **SM**  $\mu N \rightarrow e N$  is heavily suppressed because of the mass disparity between the  $W$  and neutrino.

In **new physics** scenarios this does not usually apply, and other diagrams typically give CLFV much higher than the SM.



$$\mathcal{L}_d \sim \frac{m_\mu}{\Lambda^2} \bar{\mu} \sigma_{\mu\nu} e \cdot F^{\mu\nu}$$



$$\mathcal{L}_4 \sim \frac{1}{\Lambda^2} \bar{\mu} \gamma_\mu e \cdot \bar{q} \gamma_\mu q$$

# A giant leap...

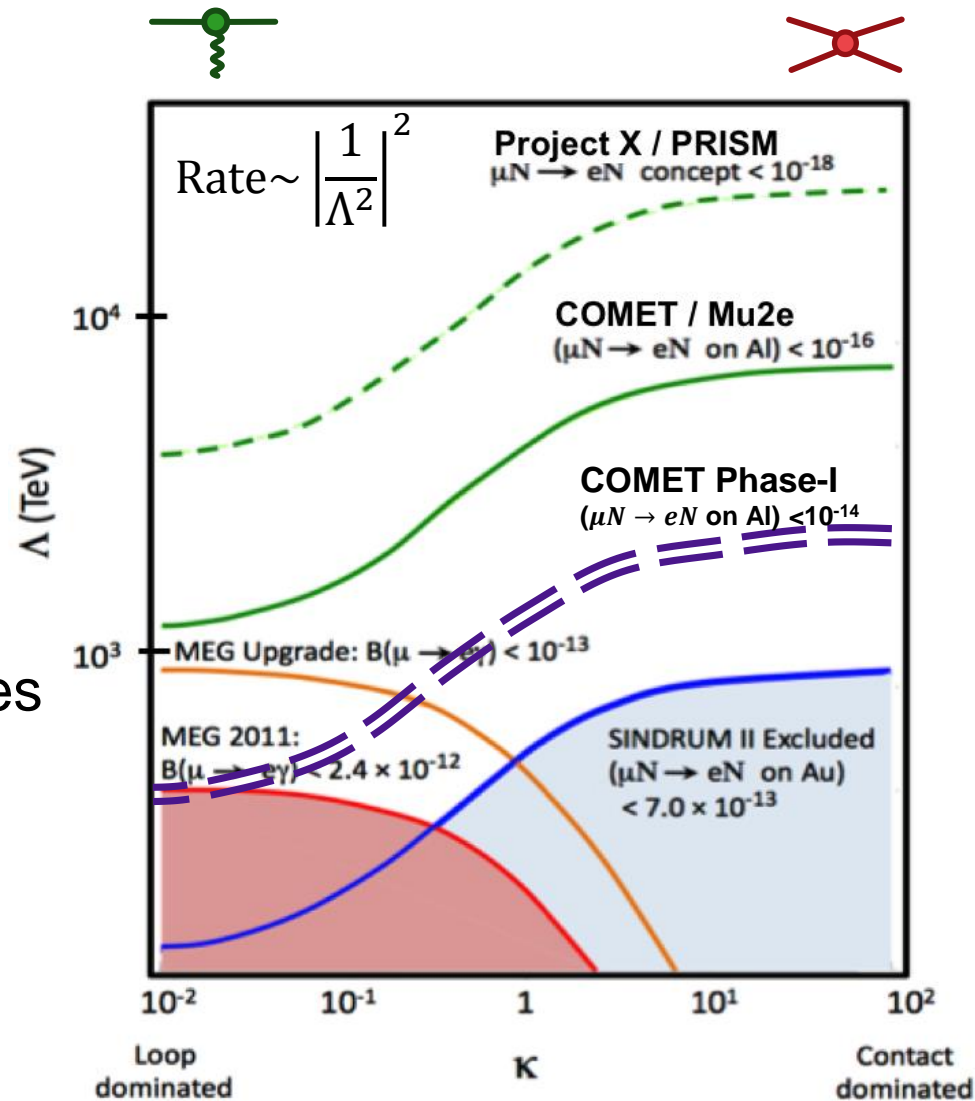
For the full COMET experiment sensitivity improvement over SINDRUM-II is **4 orders of magnitude**.

MC of background processes [especially 'tails'] may not be good enough for optimal design

- Intermediate-scale experiment can measure background sources and inform design.
- Can still do competitive physics with a smaller apparatus

Include in COMET programme:

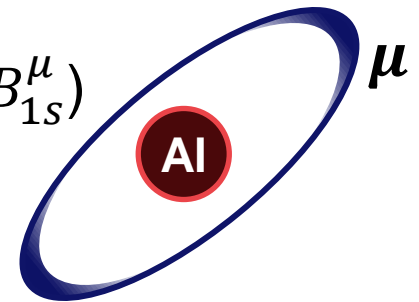
## COMET Phase-1



# $\mu N \rightarrow eN$ and other muon decays

Muons allowed stop in suitable target.

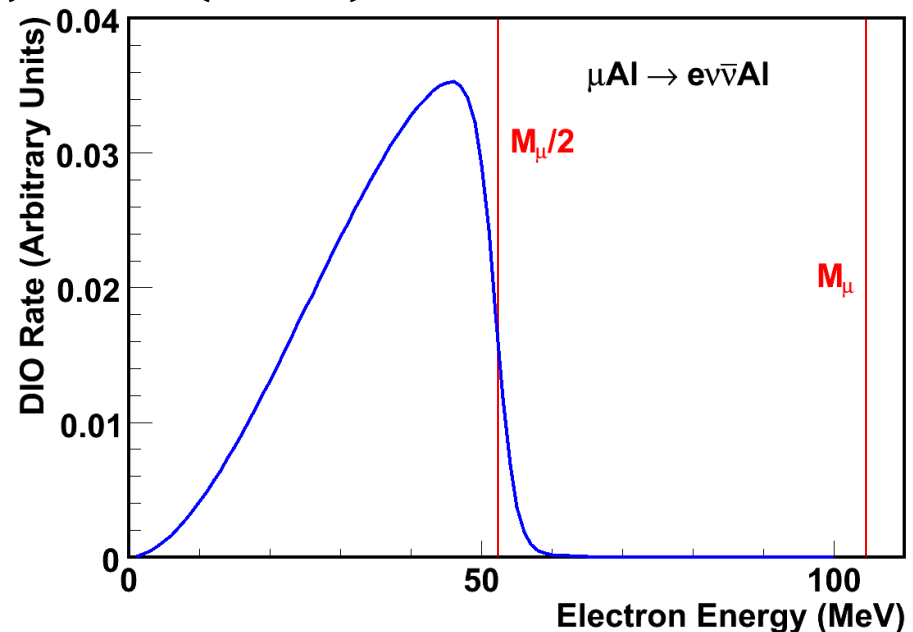
- Initially **Aluminium**, but other materials under study.
- Conversion from 1s orbital:  $\mu N \rightarrow eN$  gives a **mono-energetic electron** at 105MeV ( $\approx m_\mu - B_{1s}^\mu$ )



‘Normal’ decays are backgrounds

- Nuclear muon capture:  $\mu N(Z) \rightarrow \nu N(Z - 1)$
- Decay in Orbit [DIO]:**  
 $\mu N \rightarrow e \nu \bar{\nu} N$

For a free muon, cuts off at  $\frac{1}{2}m_\mu$ , but bound state has a small tail up to  $m_\mu$



Three main background processes:

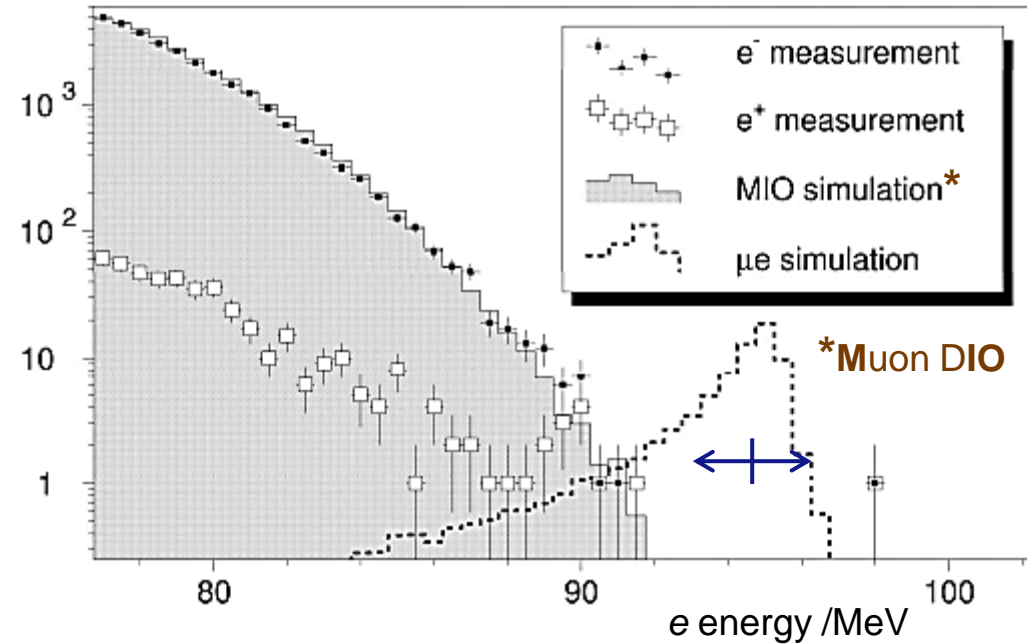
- **Decay in orbit**, as before ▶  
**Energy resolution!**
- **Decay in flight:**  
Electrons from energetic free muons can be boosted to 105MeV.
  - Use momentum selection in muon transport

See friday talk (Kurup)

- **Beam backgrounds:**

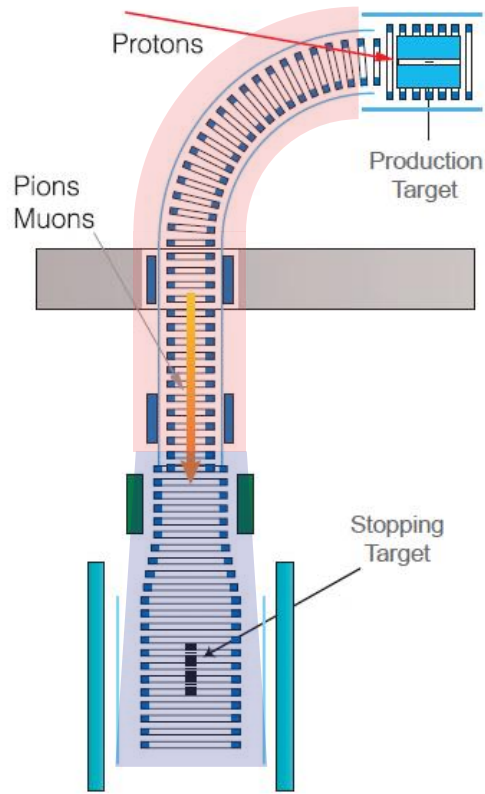
Significant number of prompt  $e^-$  and  $\pi^-$  produced by beam. Can eliminate this with timing *if* we have reliably beam-free time windows.

## Results from SINDRUM-II

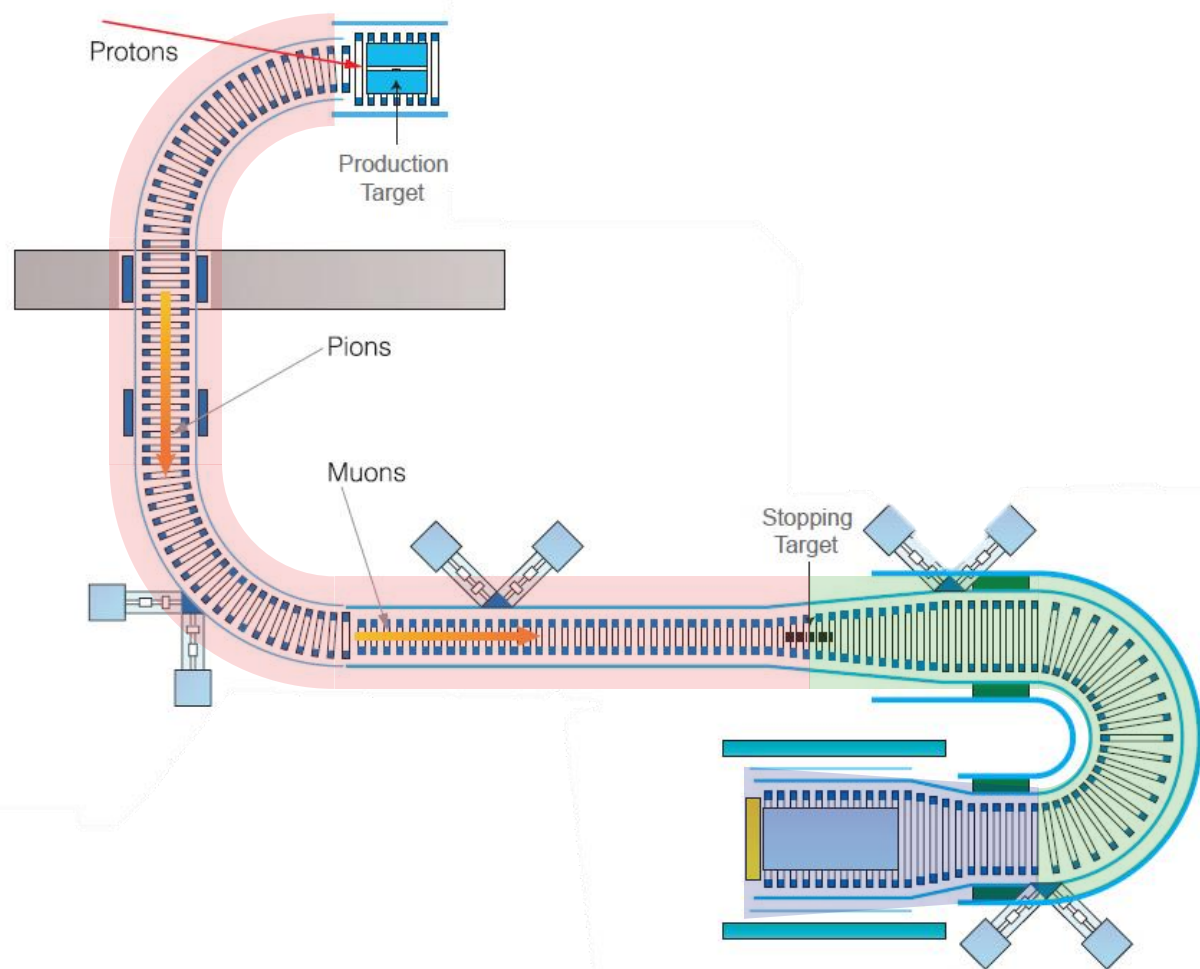


# COMET phases

## Phase I



## Phase II



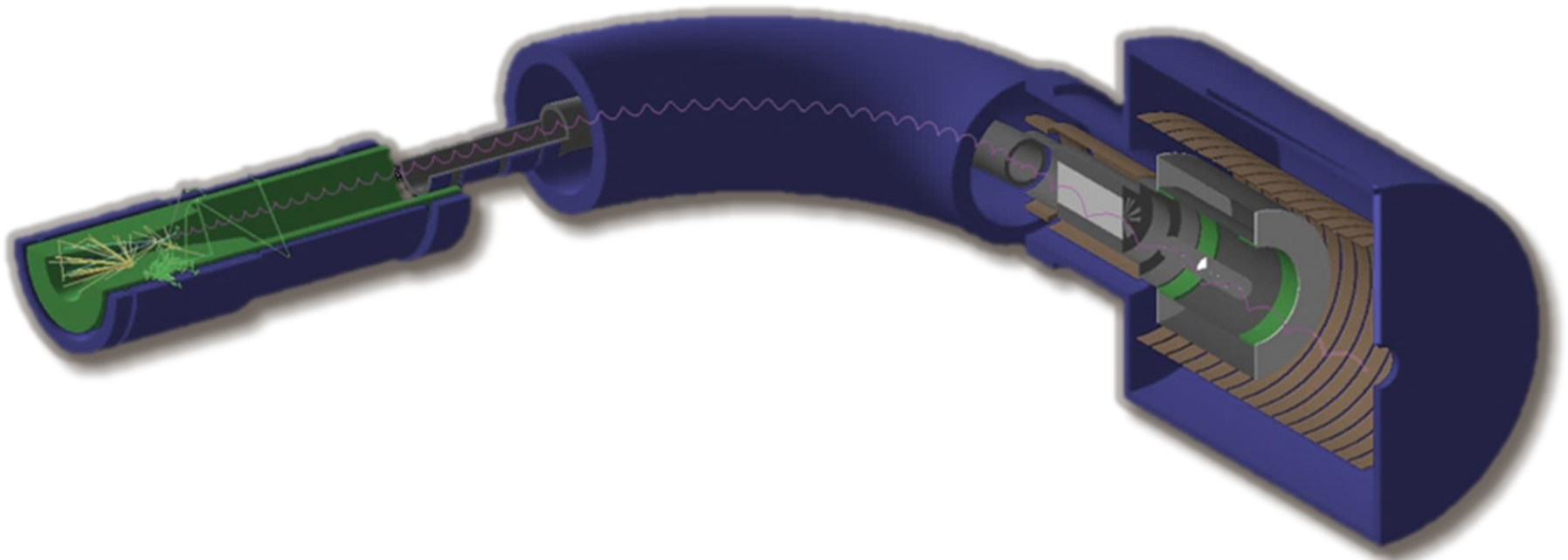
# Phase I

Capture solenoid and first 90° of muon transport will be reused for COMET Phase II.

- See A.K. talk for more on **muon transport**
- **Production target** and **capture solenoid** covered here

Stopping target is 17×0.2mm Al discs (100mm radius), and is surrounded by a **tracking chamber** for physics measurement.

- Not going to cover detector solenoid in any detail



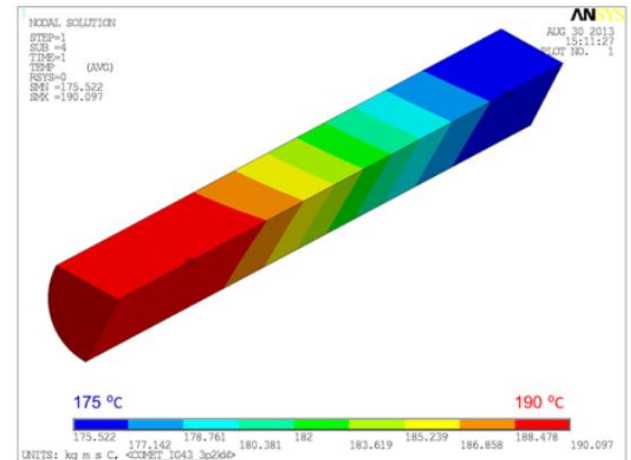
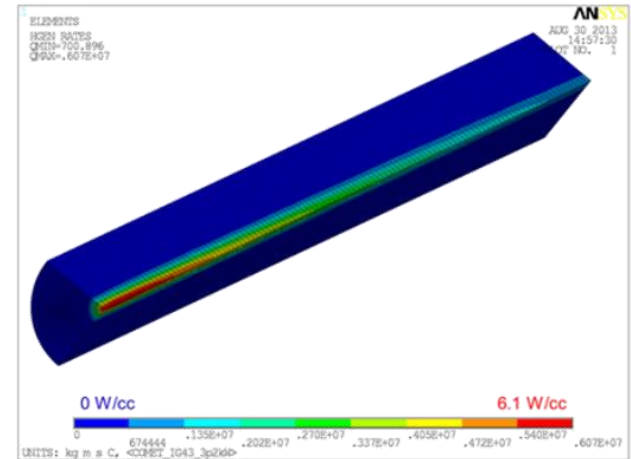


# Production target

In Phase-I we will use a 60cm x 2cm dia. graphite (IG-43) target.

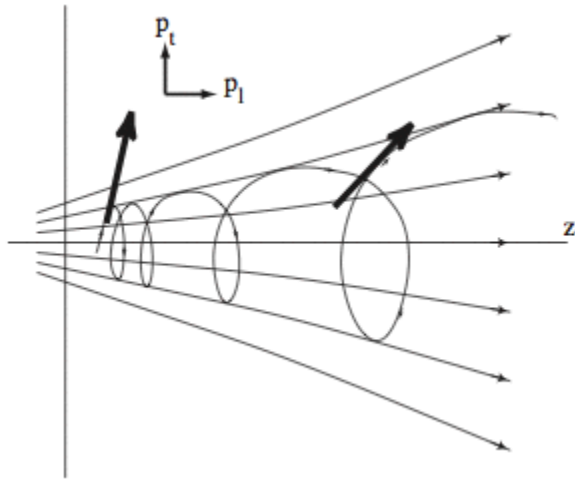
Higher Z is better for pion production, but **graphite** is a 'safer' choice:

- IG-43 is used for T2K target (FX, >200kW beam) so is known to be capable of handling our beam.
- Lower **irradiation** of target and shield makes removal and storage safer in case of replacement in Phase-II
- At Phase-I power, radiative cooling is sufficient for this target.



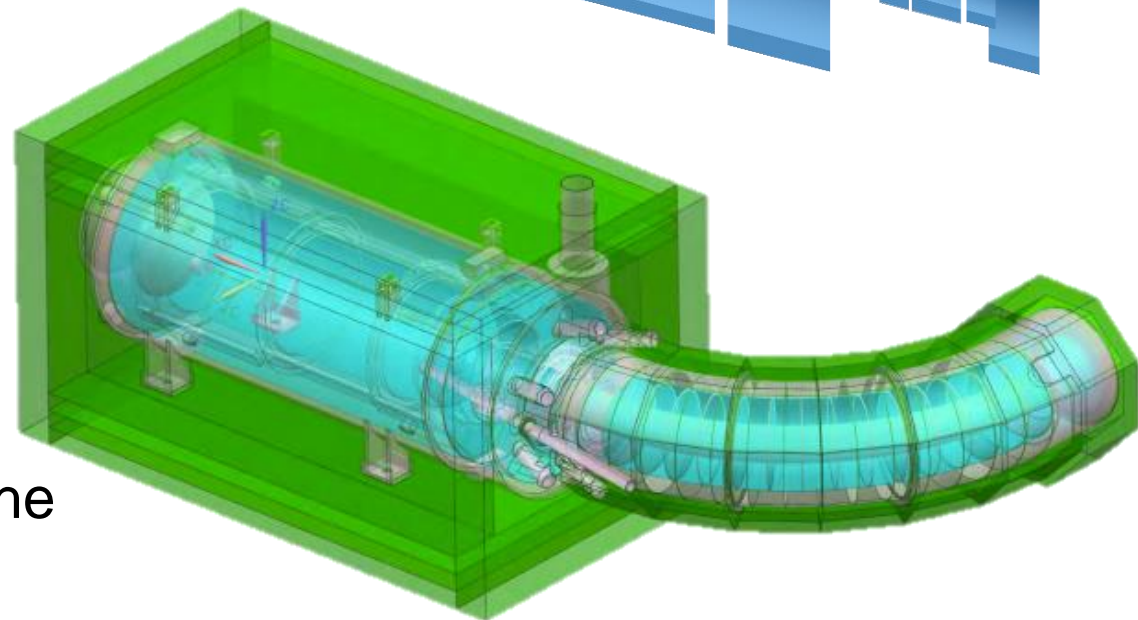
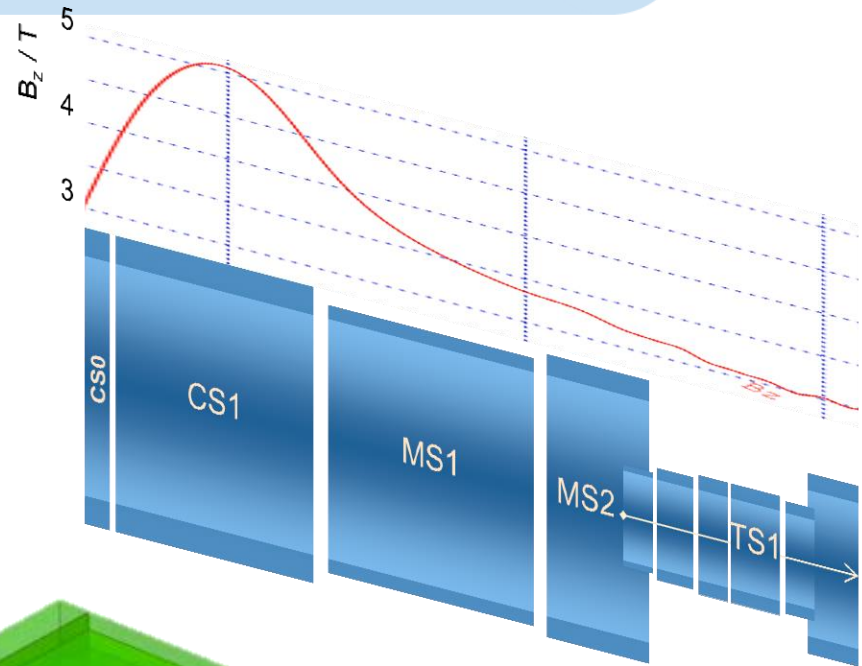
# Capture Solenoid

Comet needs *low energy* pions so collect from **back and sides** of target.

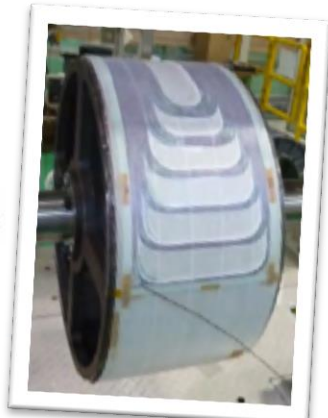
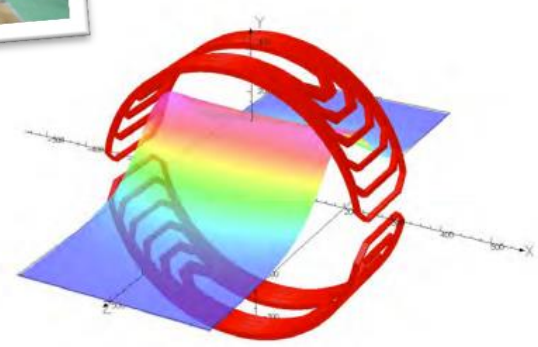
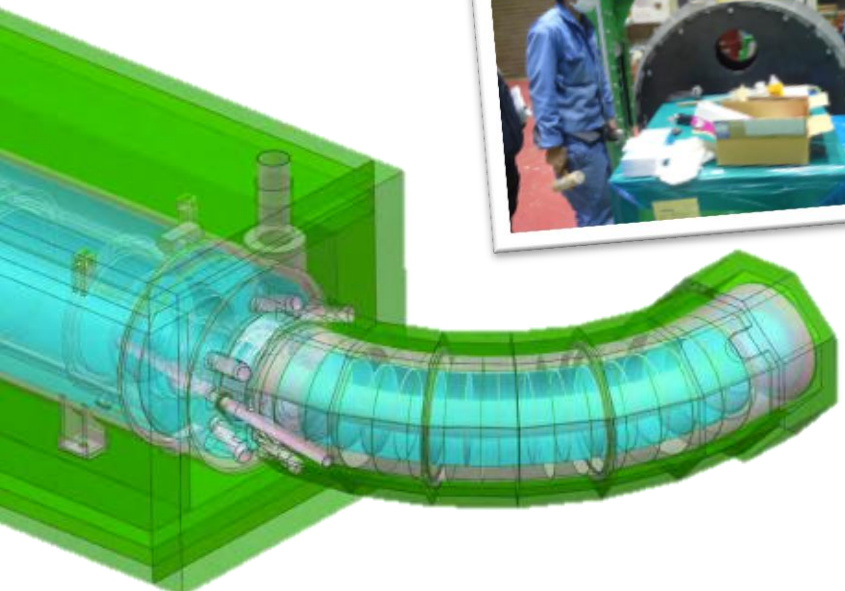
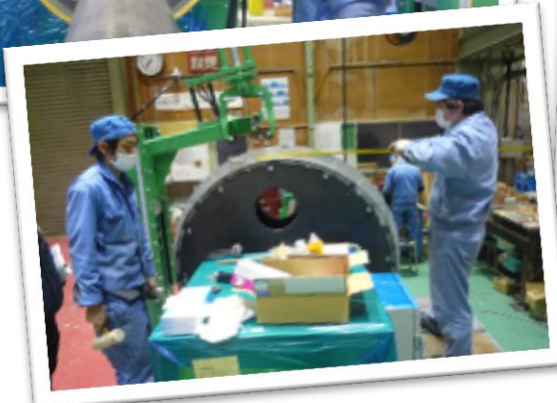
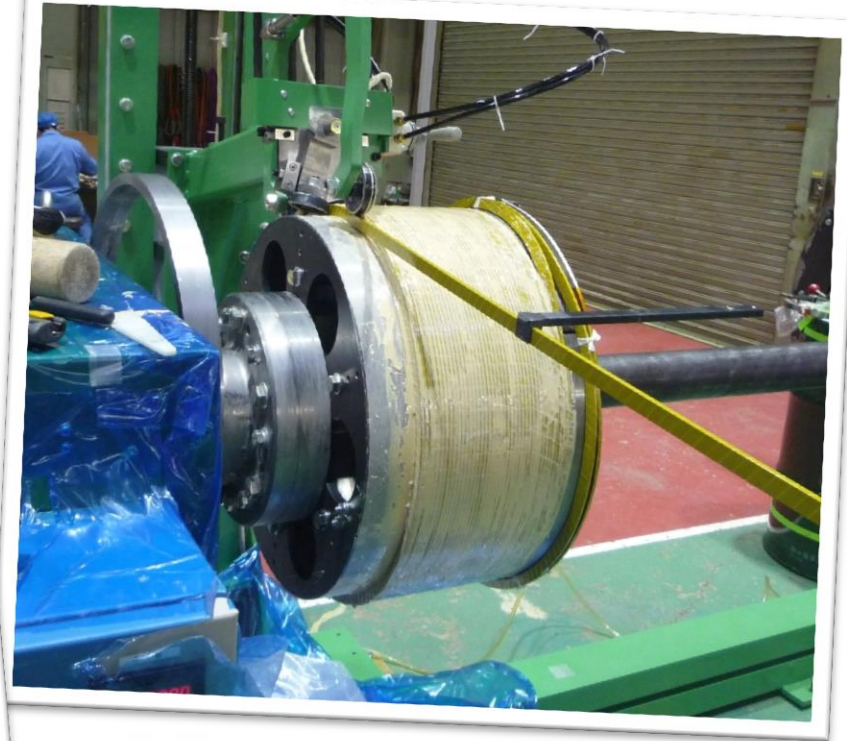


**Gradient field** converts transverse momentum into longitudinal momentum.

- Effectively increases the solid angle aperture into the transport solenoid.



# Coil winding (TS1)



Compensating dipoles

# Cooling and shielding

A 5T solenoid is (unsurprisingly?) superconducting.

- And therefore cryogenically cooled...

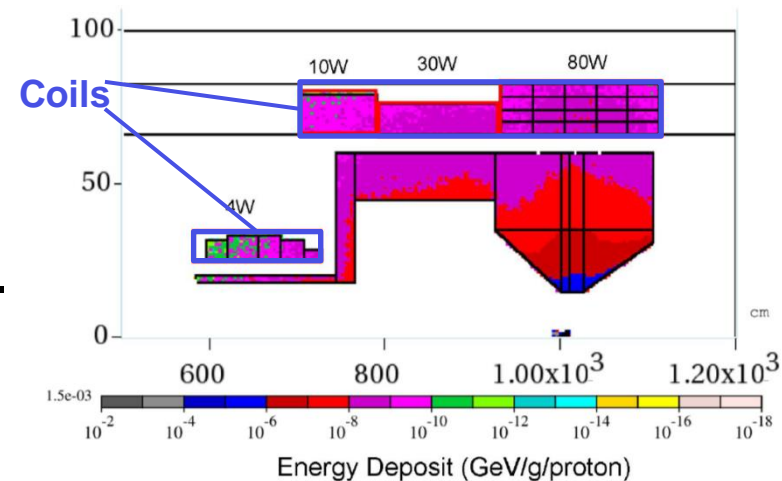
But there is a high power beam hitting a target in the middle!

- Phase I: this heating is estimated up to **30W**
- Phase II: heating can be **120W** [c.f. other sources ~15W]

**Shielding** is needed, for radiation and thermal heating.

- Copper and tungsten shield
- Cooled with water
- Will probably need upgrade for Phase II, gets very (radioactively) hot.

**Non-trivial engineering challenge!**



# Phase-I Detector

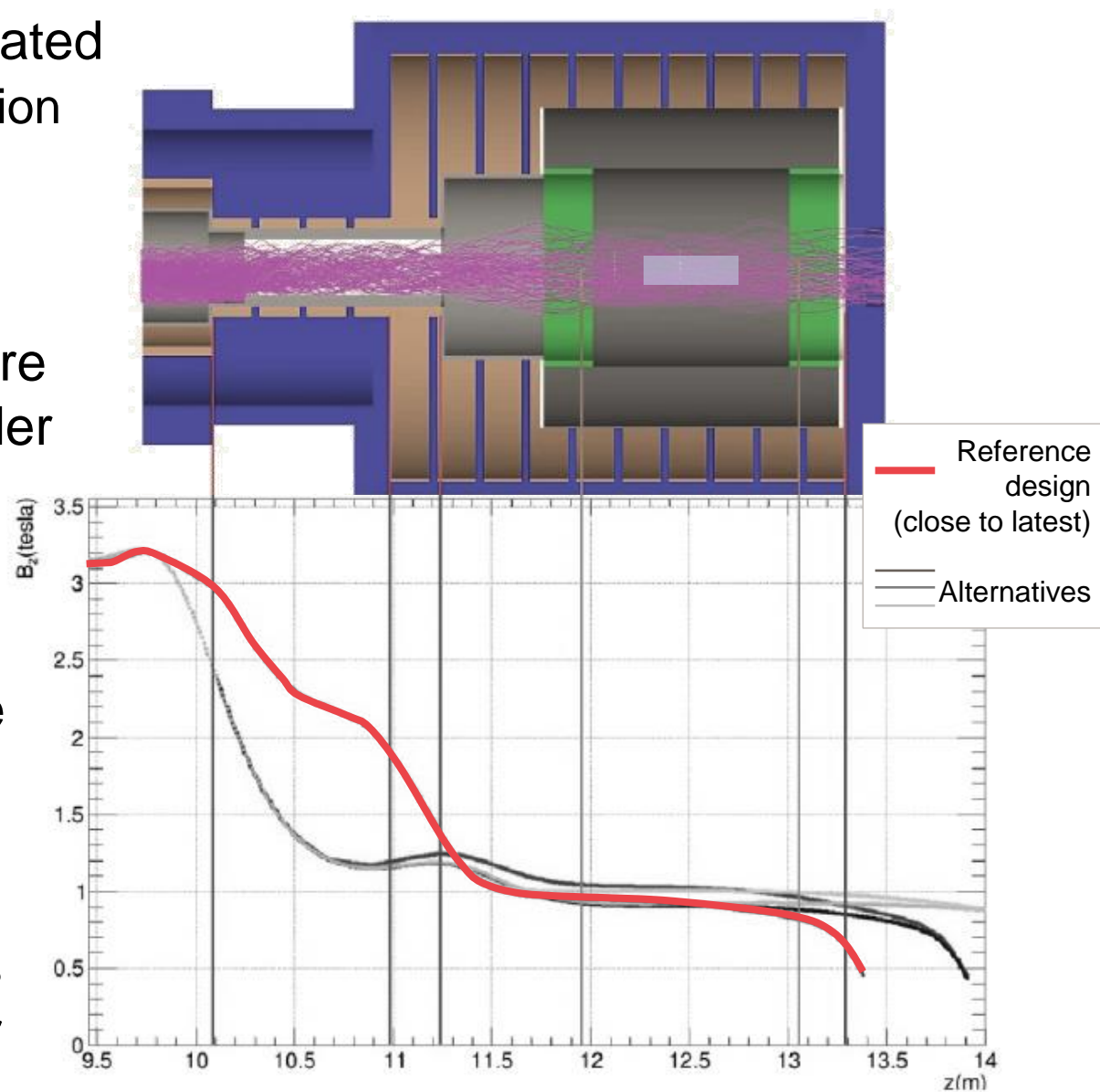
Phase-I will have a dedicated detector for  $\mu \rightarrow e$  conversion measurements.

Because of the charged particle tracks in the centre channel, a co-axial cylinder geometry is used.

↳ **CyDet**

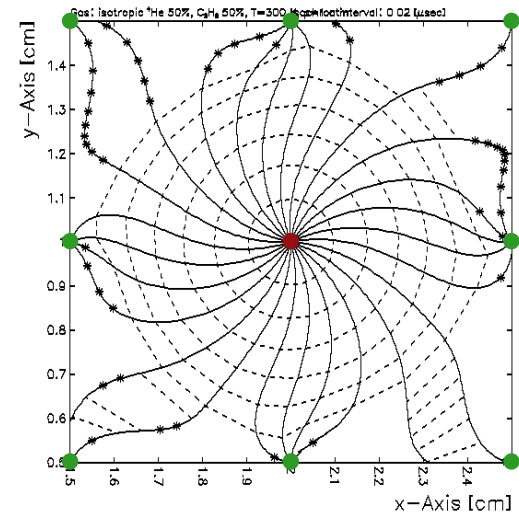
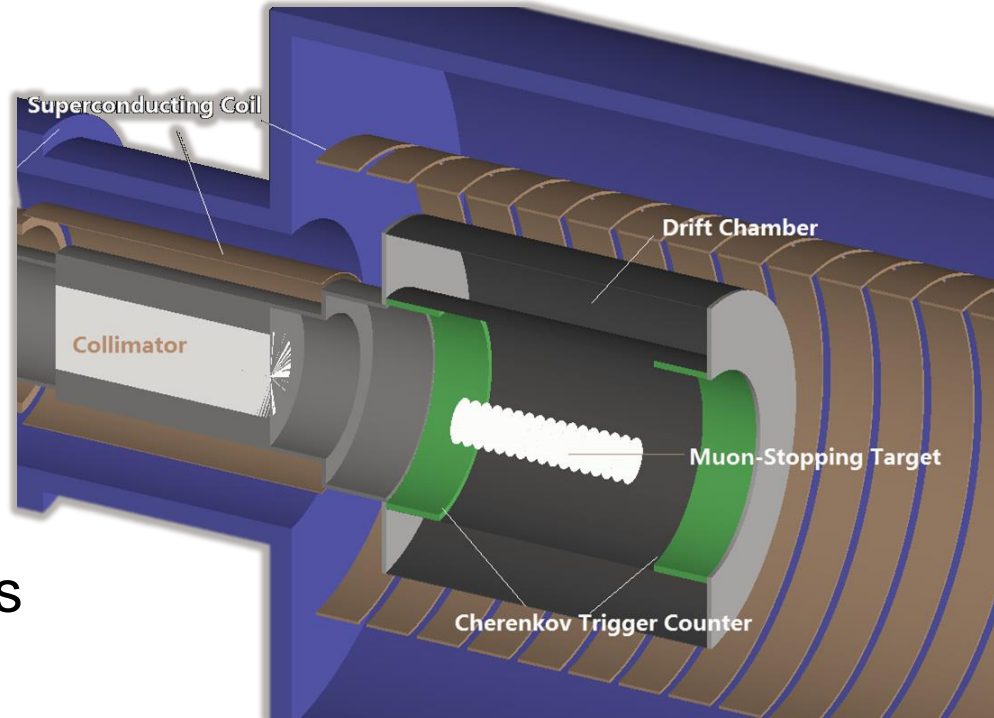
The detector and capture target will sit within a 1T solenoid field.

Low momentum particles do not reach the detector

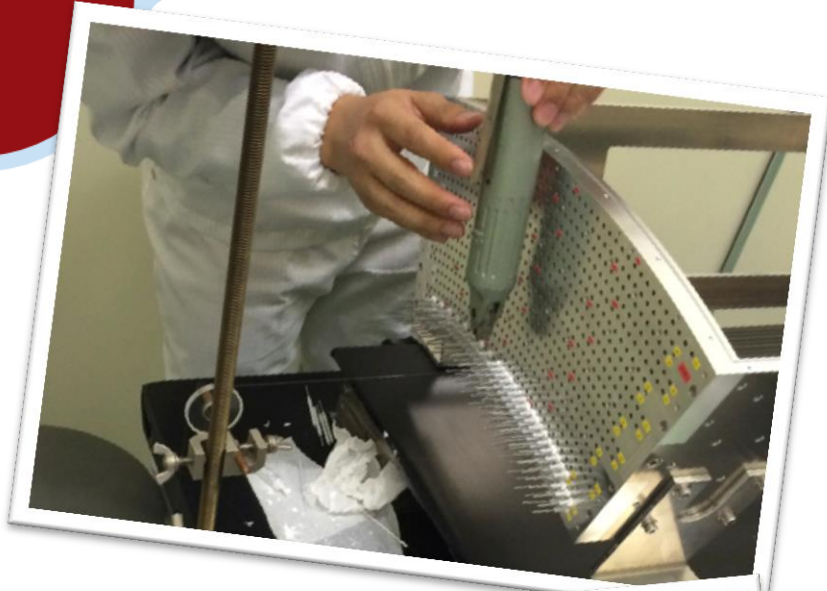


The main part of the detector is a coaxial **drift chamber**

- Helium-based gas mixture to reduce multiple scattering.
  - Resolution  $\sim 200$  keV
- $z$  measurement by stereo layers
- Large inner radius to reduce DIO hit rate
  - Dim: 150cm  $\times$  84cm(outer) // 50cm(inner)
- 19 concentric sense layers
- Triggering from **hodoscopes** at ends

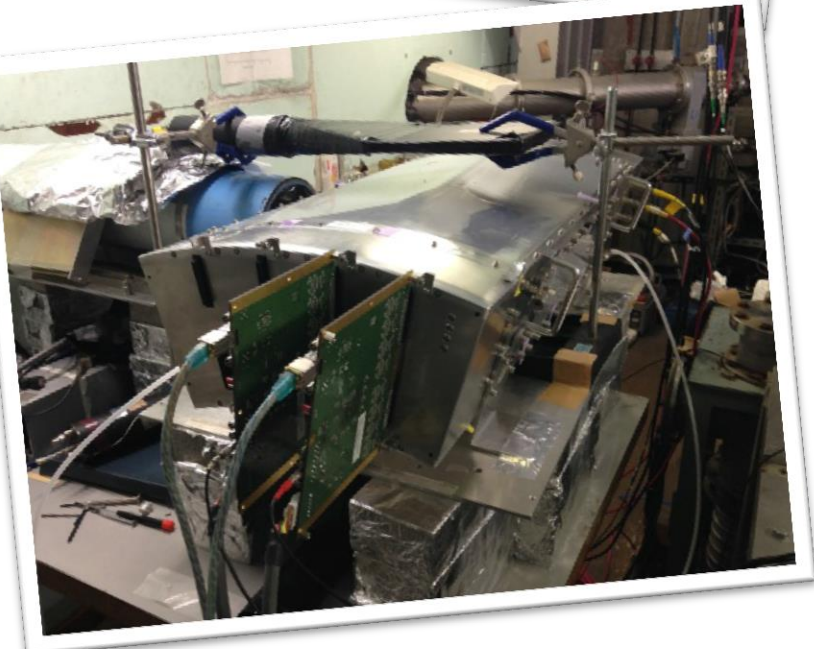
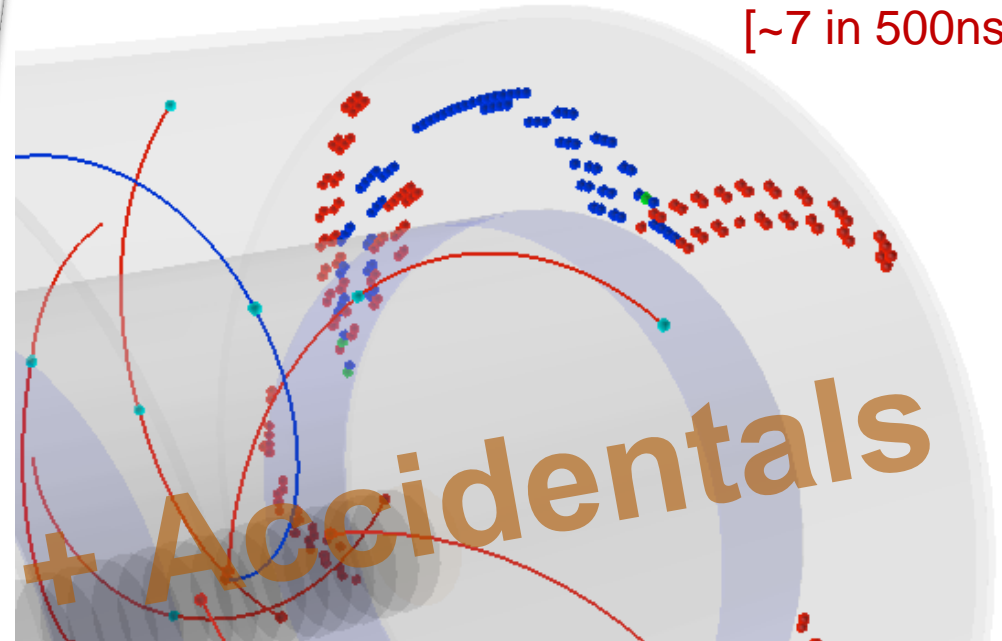


# Drift chamber progress



Electron track [ $\sim 3\text{Hz}$ ]

Proton tracks  
[ $\sim 7$  in  $500\text{ns}$ ]



- ▲ Event display showing event projection
- ◀ Stringing wires and CR test of prototype section

# Facility construction



◀ Magnets laid-out for beam switchyard



June



July

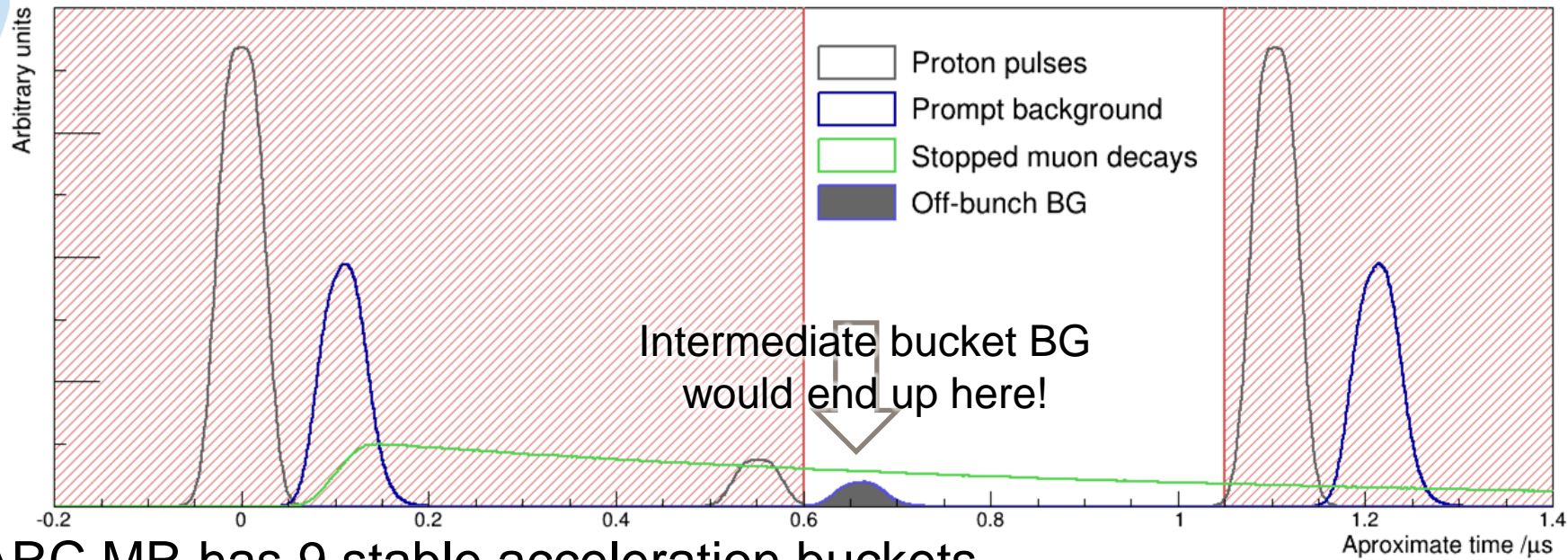


August

◀ COMET hall construction proceeding well



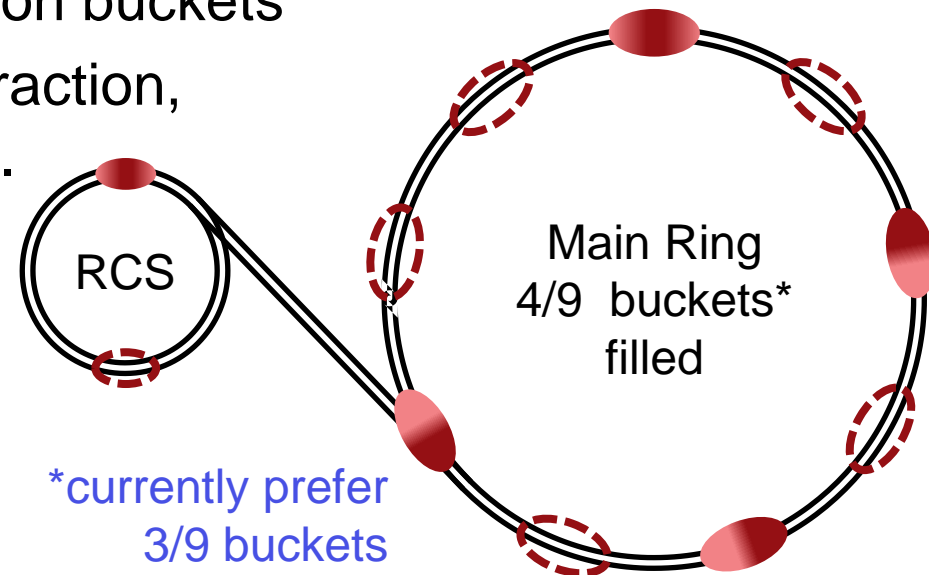
# Beam extinction



J-PARC MR has 9 stable acceleration buckets

- Need to maintain RF during extraction, so that bunch structure remains.
- If RF is not strong enough, protons will 'leak' into empty buckets.

Signal process is rare so even a small leak is a major background

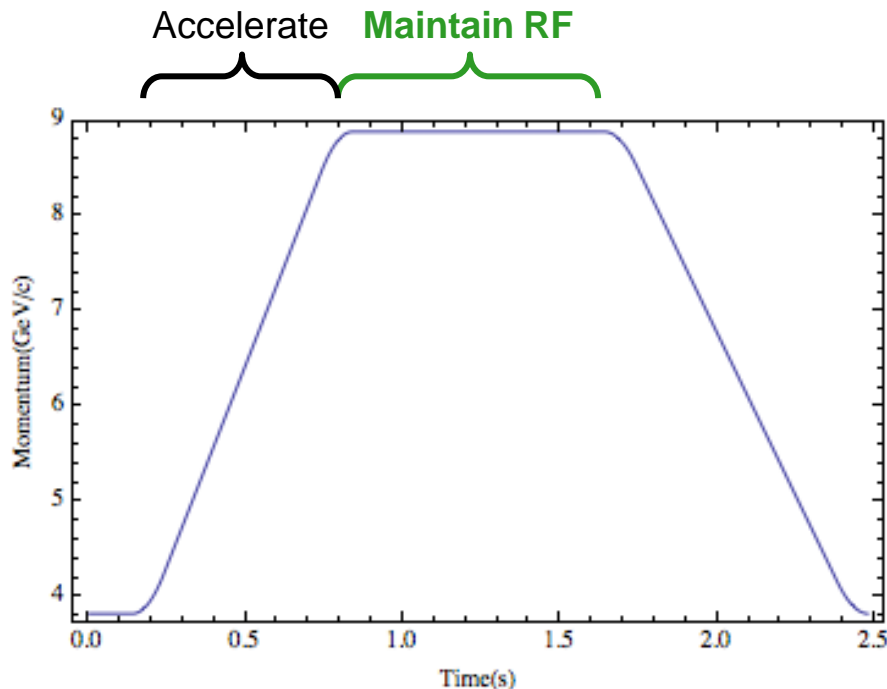
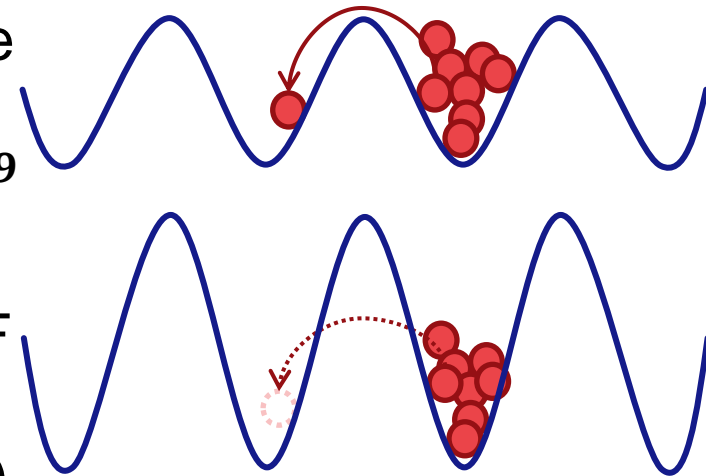


# Extinction measurement

COMET design requires that we can achieve an extinction:

$$E = \frac{N_{Empty}}{N_{Filled}} < 10^{-9}$$

Extinction can be improved by increasing RF voltage, but this heats the cavities.  
(And there is a limit...)



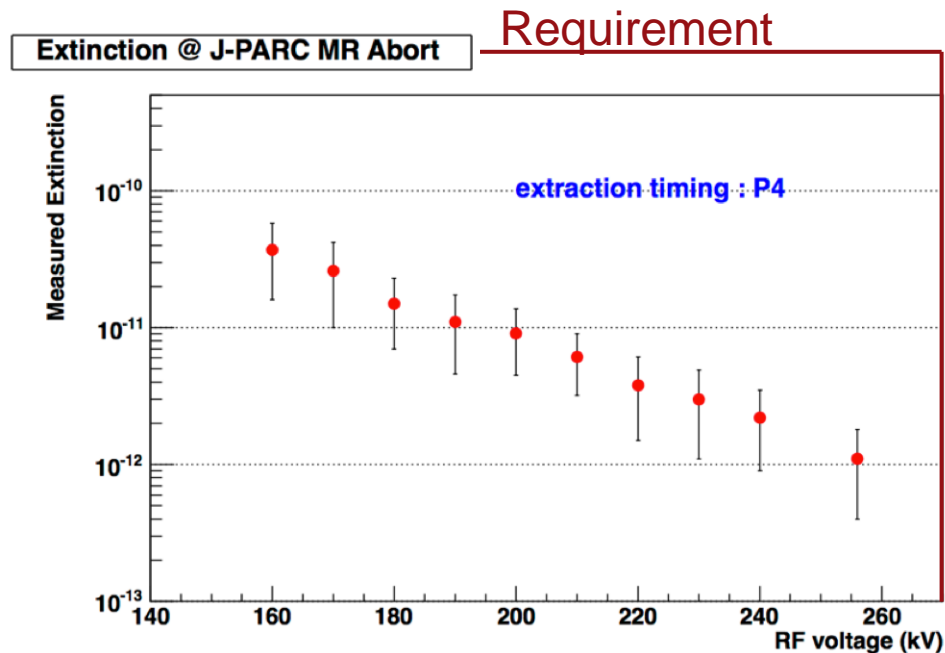
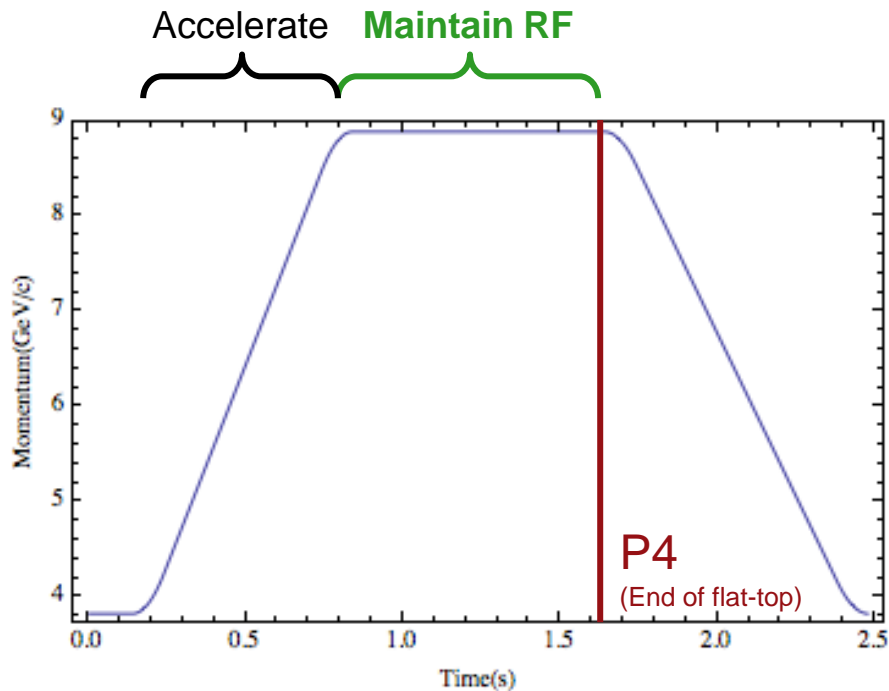
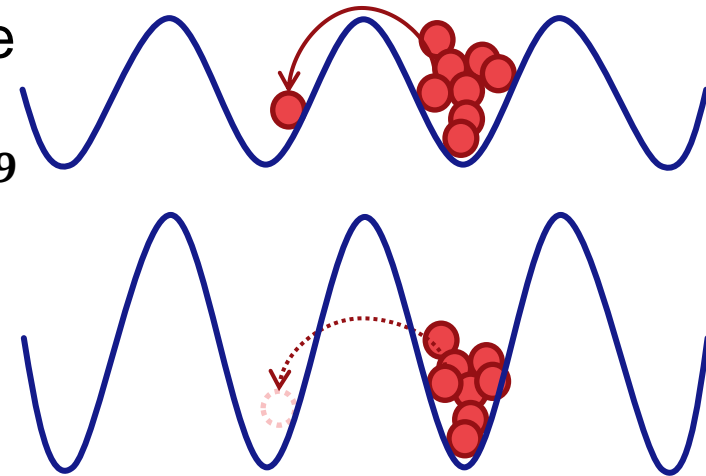
2012 test at 30 GeV demonstrated this is possible for RF > 120kV, But study time at 8 GeV not allocated until this year...

# Extinction measurement

COMET design requires that we can achieve an extinction:

$$E = \frac{N_{Empty}}{N_{Filled}} < 10^{-9}$$

**Result was excellent!** Even low RF voltage provides the required extinction



- **Phase I of COMET is taking shape!**
  - Facility construction progressing quickly.
  - Magnet coils also in construction.
  - CyDet prototype assembled and tested.
- **Beam extinction requirements met comfortably!**
  - Accelerator should be fully capable of (fairly demanding) use cycle.
- **Phase-I projected to be ready in 2016.**
  - Short (few month) run should improve sensitivity by 2 orders of magnitude over previous experiment (S.E.S. of  $3 \times 10^{-15}$ )
  - Will also aid understanding of backgrounds [AS talk] for Phase II [AK talk], so stay tuned for tomorrow's session.

日本語：「ムーミンだにのすいせい」 [The comet of Moomin Valley]

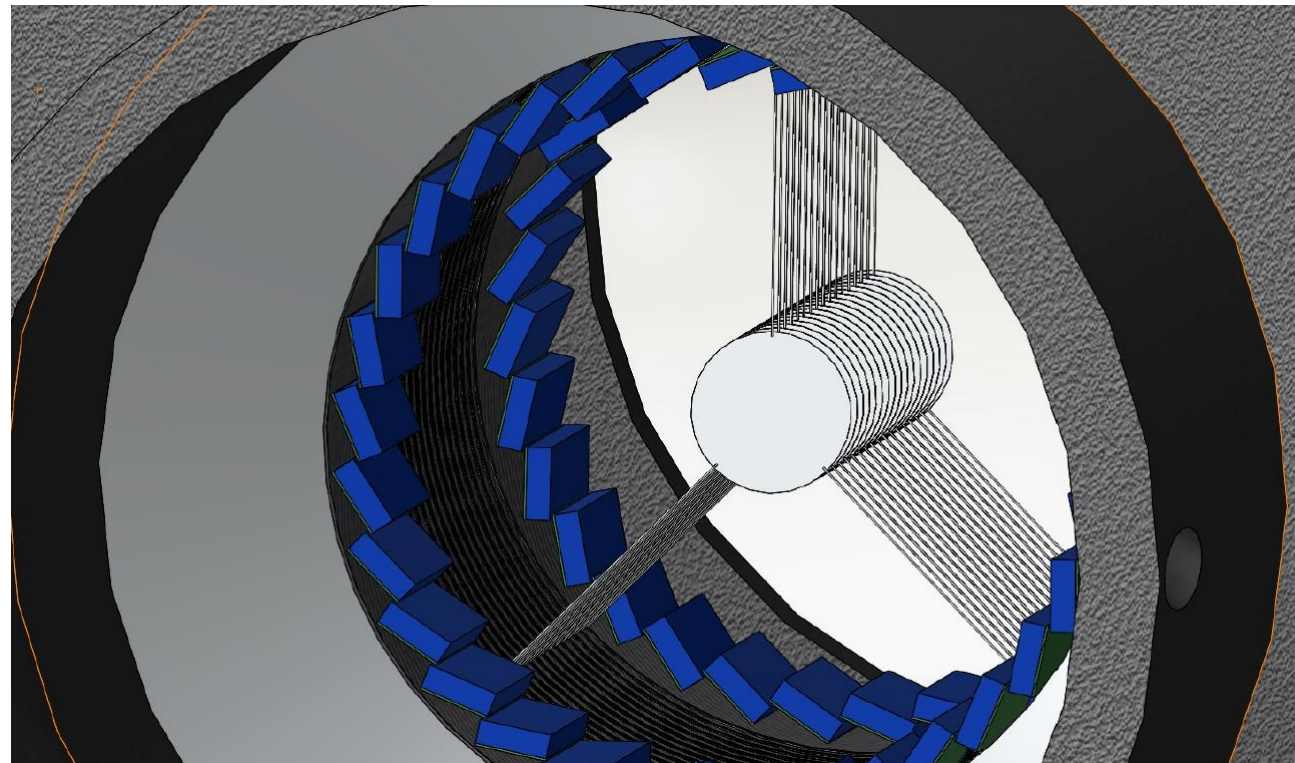


Suomi: “Muumi ja punainen pyrstötähti” [Moomin and the comet chase]

English: Comet in Moominland

# Phase I capture target details

- Target may be changed for Phase 2



Material	Al
Radius	100 mm
Thickness	200 $\mu$ m
N <sup>o</sup> Disks	17
Spacing	50 mm
Total thickness	3.4 mm
Total length	0.8 m

# Extinction vs timing

