



Neutrino Oscillations

in the

OPERA Experiment

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on behalf of OPERA Collaboration

NUFACT2014, XVIth International Workshop on
Neutrino Factories and Future Neutrino Facilities
University of Glasgow, 25th to 30th of August, 2014



The OPERA Collaboration

140 physicists, 28 institutions in 11 countries

Belgium
IIHE-ULB Brussels



Israel
Technion Haifa



Korea
Jinju



Croatia
IRB Zagreb



Italy
Bari
Bologna
LNF Frascati
L'Aquila
LNGS
Naples
Padova
Rome
Salerno



Russia
INR RAS Moscow
LPI RAS Moscow
ITEP Moscow
SINP MSU Moscow
JINR Dubna



France
LAPP Annecy
IPHC Strasbourg



Switzerland
Bern



Germany
Hamburg



Japan
Aichi
Toho
Kobe
Nagoya
Nihon



Turkey
METU Ankara



Oscillation Project with Emulsion-tRacking Apparatus

Aim: establish neutrino oscillations in **direct appearance mode in the $\nu_\mu \rightarrow \nu_\tau$ channel** at atmospheric scale, through **detection of the tau lepton produced in ν_τ Charged Current (CC) interaction on event-by-event basis.** Full coverage of the parameter space indicated by SK, T2K and MINOS.

ν_μ

oscillation

ν_τ

τ^-

τ lepton decays:

$\tau^- \rightarrow \mu^- + \bar{\nu}_\mu + \nu_\tau$	17.7%
$\tau^- \rightarrow e^- + \bar{\nu}_e + \nu_\tau$	17.8%
$\tau^- \rightarrow \text{hadron} + \nu_\tau$	49.5%
$\tau^- \rightarrow \pi^- \pi^- \pi^+ (n\pi^0) + \nu_\tau$	15.0%

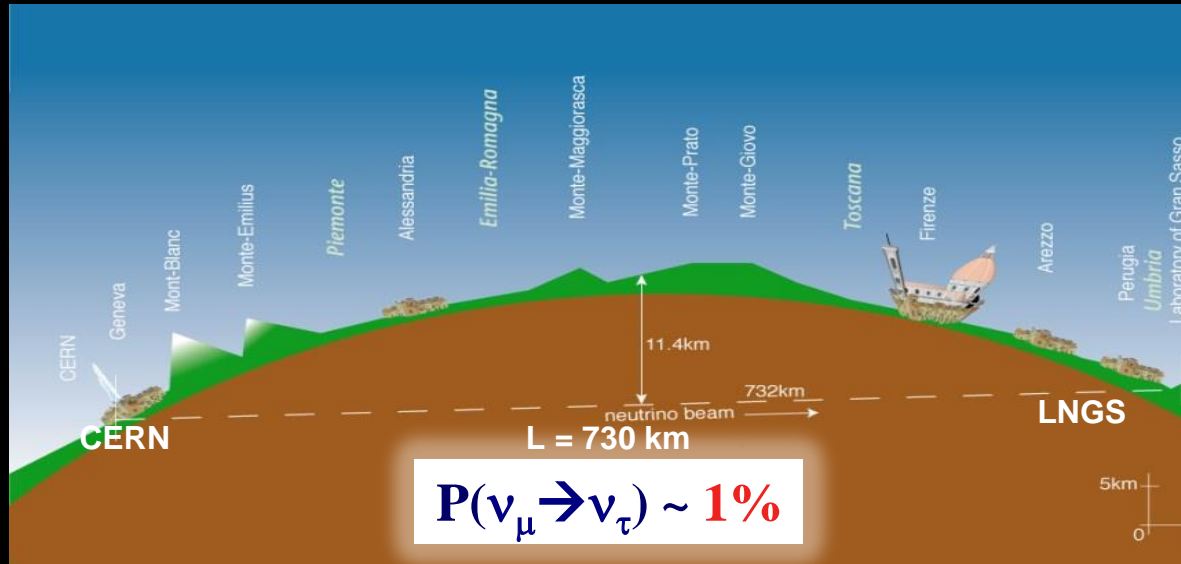
$$P(\nu_\mu \rightarrow \nu_\tau) \sim \sin^2 2\theta_{23} \cos^4 \theta_{13} \sin^2(\Delta m_{23}^2 L/4E)$$

Requirements:

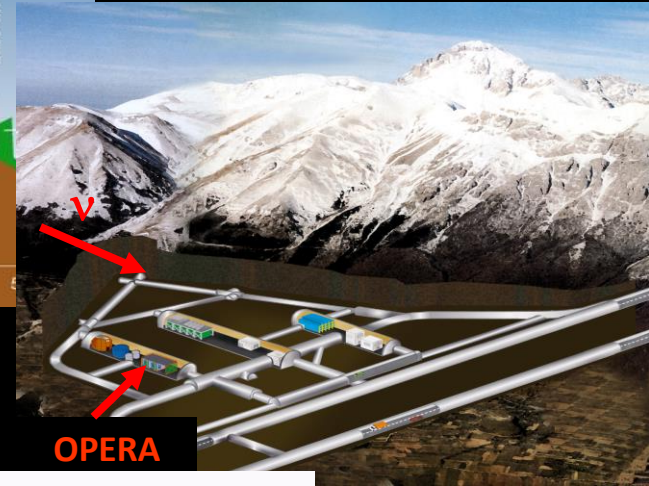
- ✓ High energy and high intensity neutrino beam
Conventional ν_μ beam optimized for ν_τ appearance
- ✓ Long baseline
- ✓ Large target mass and submicron resolution
- ✓ Detector capability to identify short-lived τ lepton

CERN Neutrino to Gran Sasso Beam

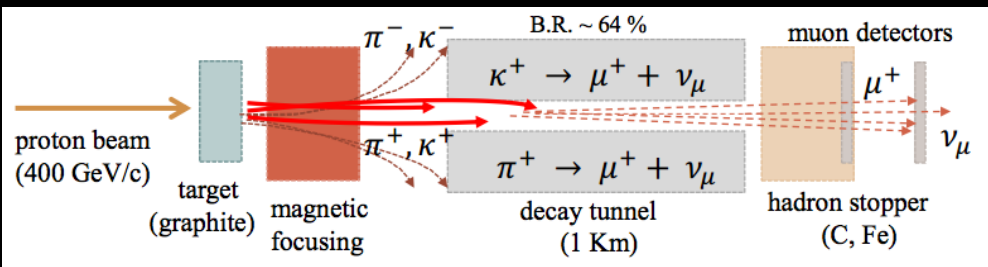
<http://proj-cngs.web.cern.ch/proj-cngs/>



Gran Sasso National Laboratory:
 1400 m rock shield
 (cosmic μ flux reduction by a factor 10^6 w.r.t. surface)

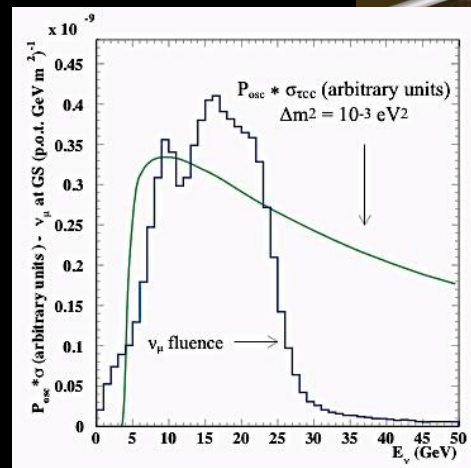


$P(\nu_\mu \rightarrow \nu_\tau) \sim 1\%$



$\langle E_\nu \rangle$	17 GeV
L	730 km
$(\nu_e + \bar{\nu}_e)/\nu_\mu$	0.87 % *
$\bar{\nu}_\mu/\nu_\mu$	2.1 % *
ν_τ prompt	Negligible *

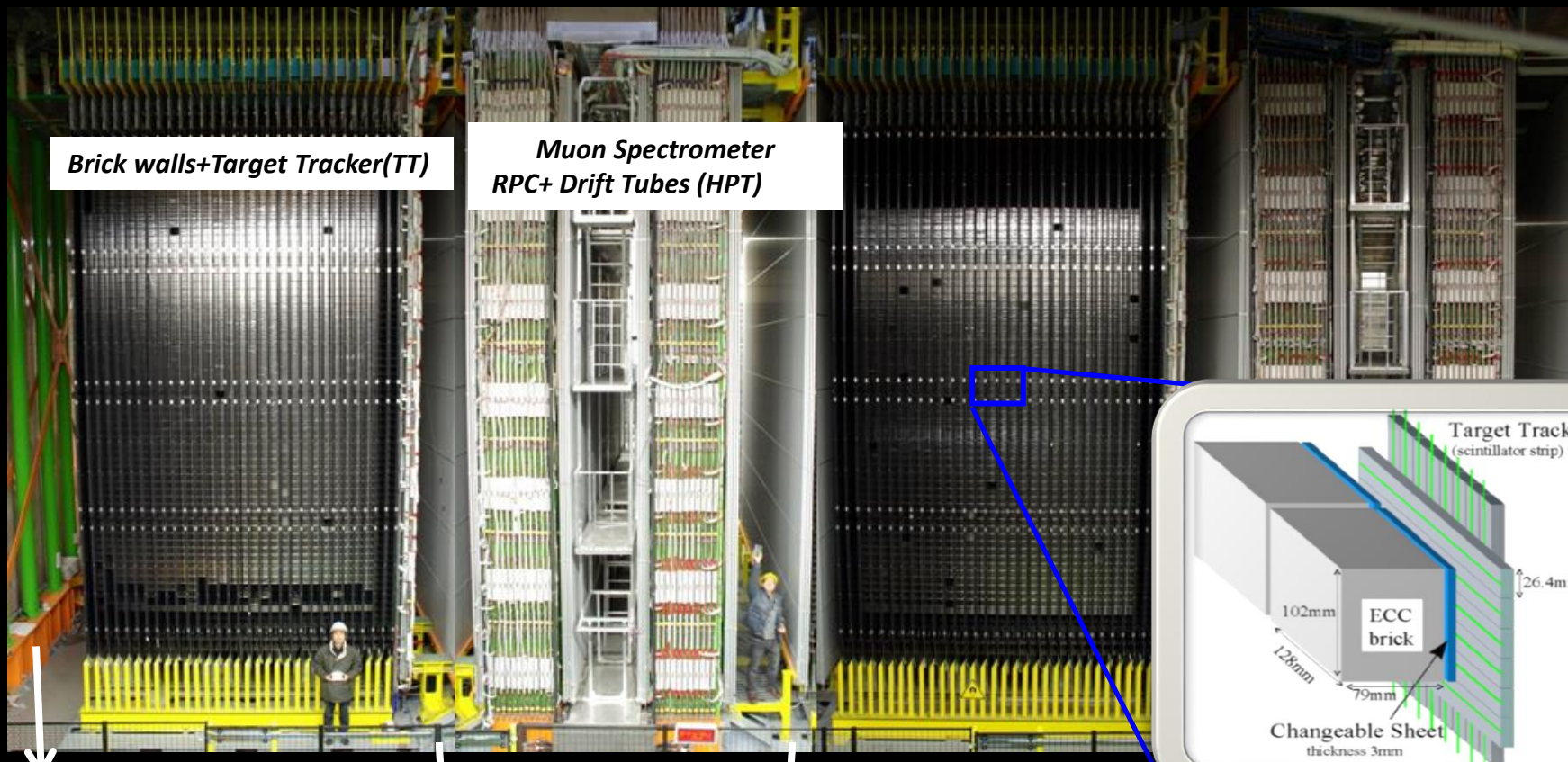
* Interaction rate at LNGS



Optimised to maximise ν_τ CC interactions at LNGS

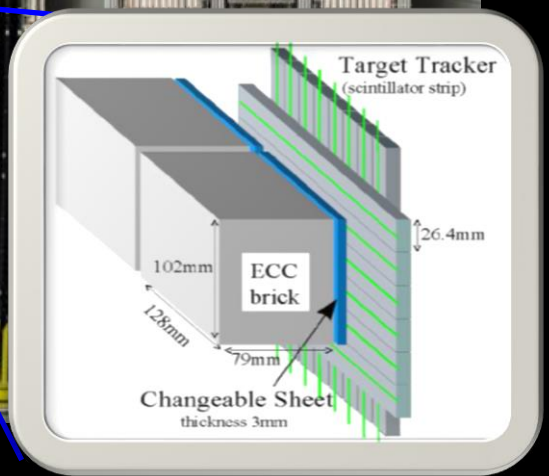
OPERA detector*

* JINST 4 P04018, 2009.



Brick walls+Target Tracker(TT)

Muon Spectrometer
RPC+ Drift Tubes (HPT)



Veto

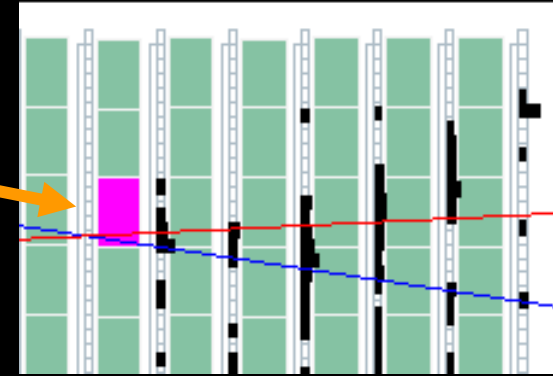
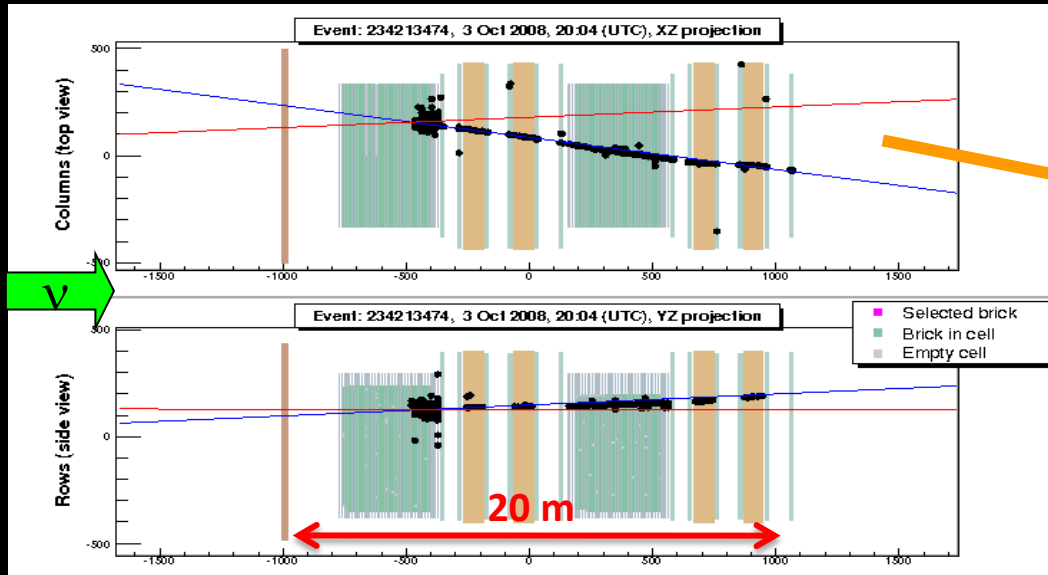
- μ id efficiency $\sim 95\%$ (with TT)
- $\Delta p/p < 20\%$, $p < 50$ GeV/c
- misidentified μ charge prob. : 1.2%

Each brick wall followed by a plane of Scintillator strips in X/Y direction. Target Tracker used for localization of Interaction brick.

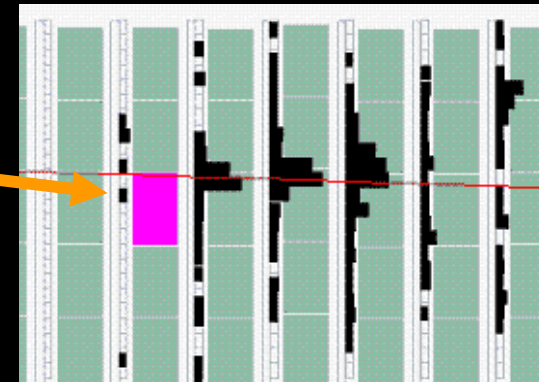
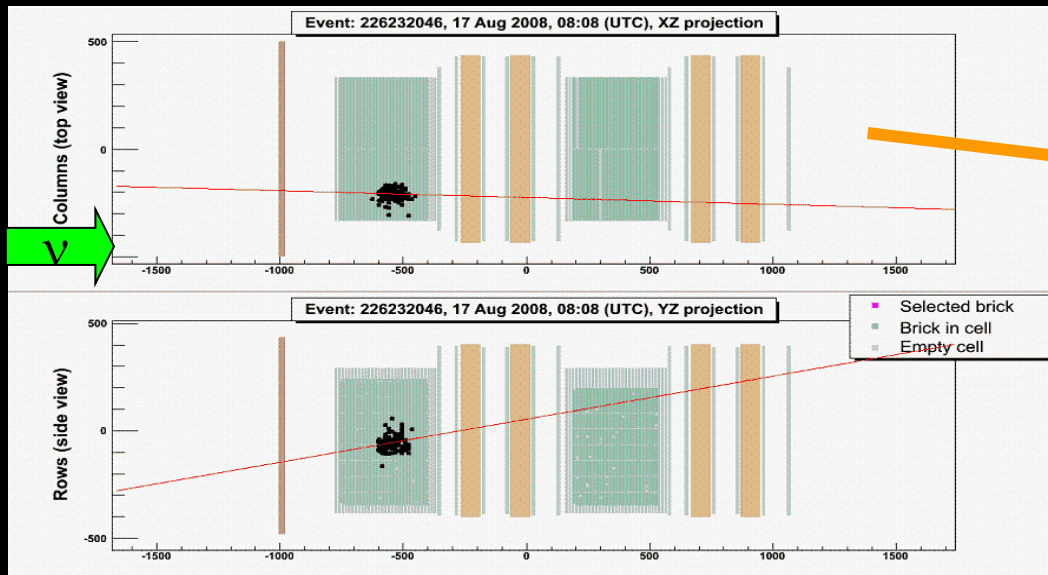
ECC brick= 56 lead plates + 57 films ~ 8 kg ($10X_0$)

- Trigger efficiency $> 99\%$
- Brick finding : $\epsilon \approx 60\div 80\%$

Typical ν_μ CC-like and ν_μ NC-like events

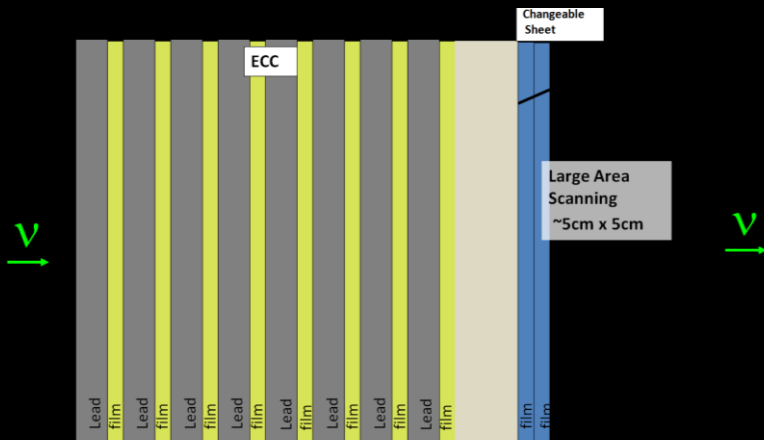


ν_μ CC
 ν_τ CC, $\tau \rightarrow \mu$ (17.7%)

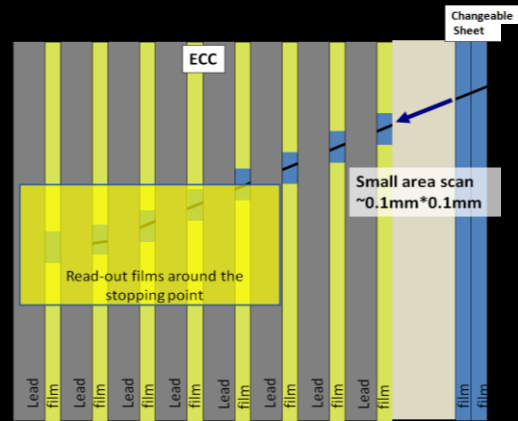


ν NC
 ν_τ CC, $\tau \rightarrow e$ (17.8%)
 ν_τ CC, $\tau \rightarrow h, 3h$ (64.5%)
 ν_e CC

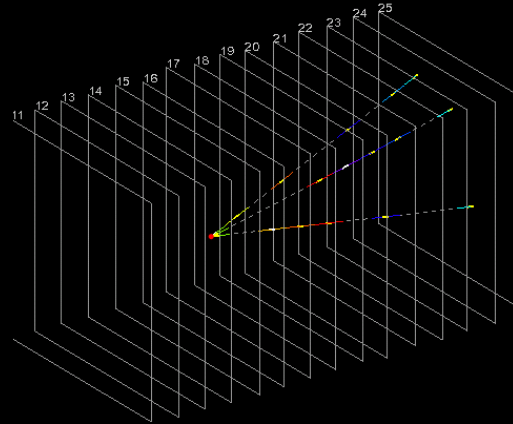
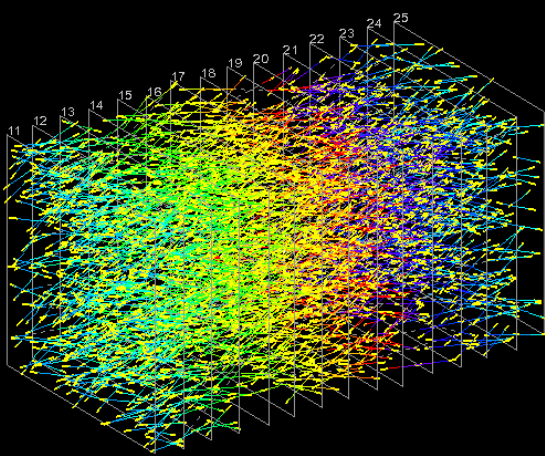
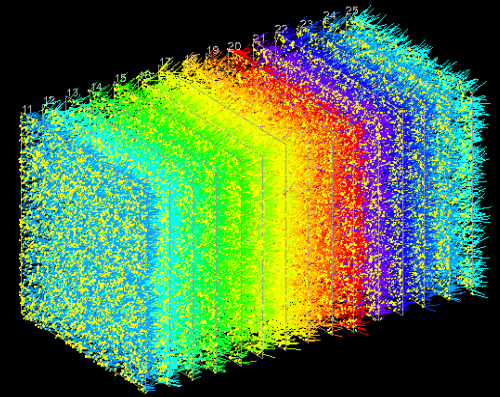
Neutrino interaction location in ECC



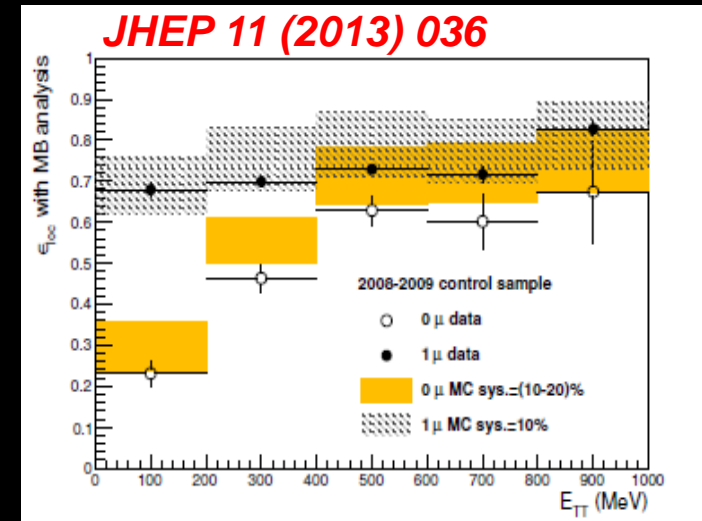
Step 1: CS scanning → confirmation of ν interaction, predictions for brick analysis



Step 2-4: CS-ECC connection, ScanBack procedure, Volume data taking



Step 5-6: Validation of neutrino interaction vertex



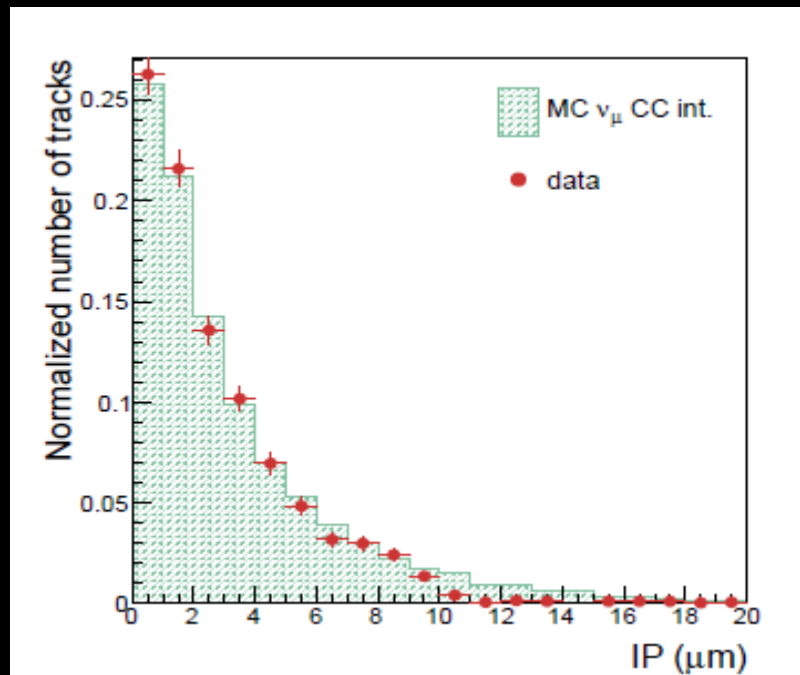
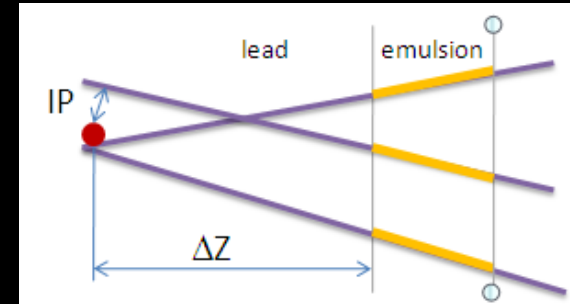
Location efficiency vs event energy

Impact Parameter evaluation

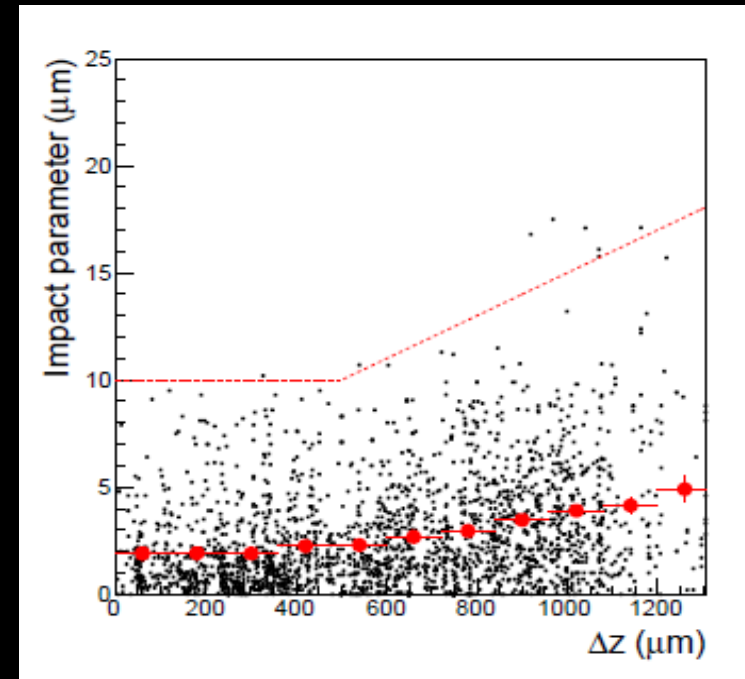
→ crucial point to detect decay topology

Kink search

Parent search



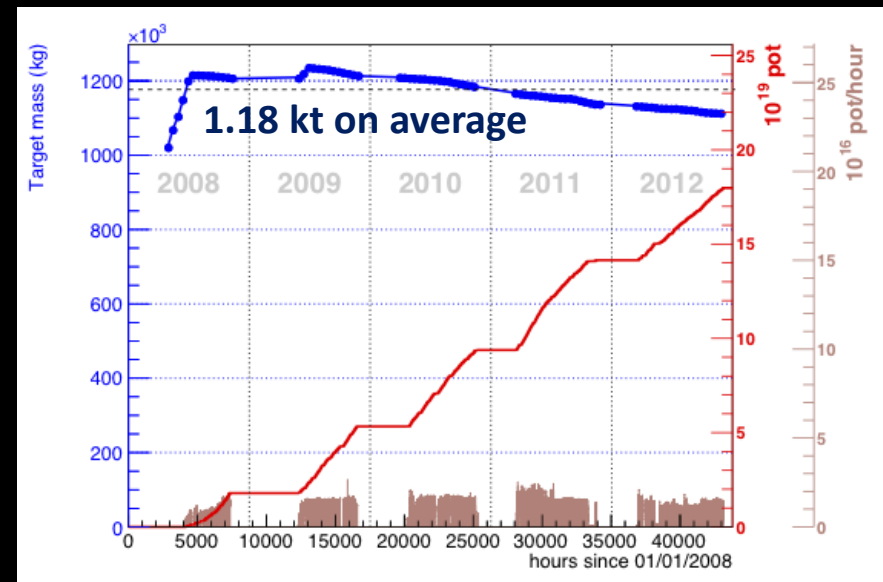
Impact parameter of tracks w.r.t. reconstructed ν interaction vertices



Impact parameter vs. longitudinal distance from reconstructed ν interaction vertex

Status of the Analysis

Run	p.o.t. (10^{19})	SPS efficiency	Beam days	ν interactions
2008	1.7	61%	123	1931
2009	3.53	73%	155	4005
2010	4.09	80%	187	4515
2011	4.75	79%	243	5131
2012	3.86	82%	257	3923
Total	17.97	77%	965	19505



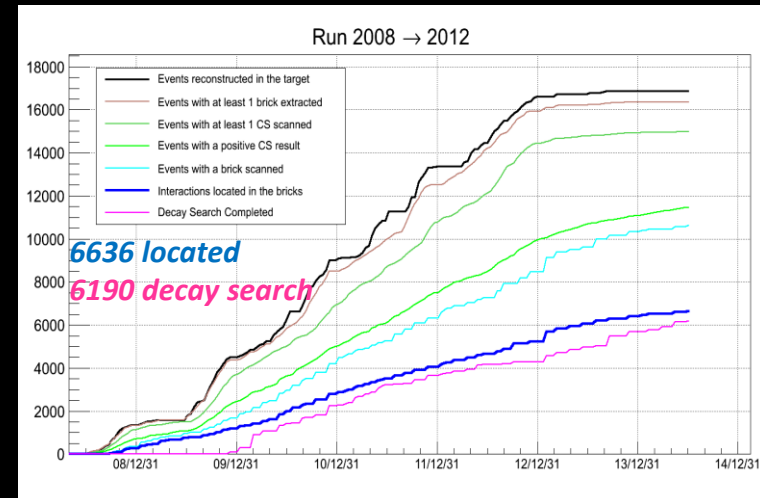
↳ 80% of the design

↳ ~87% predicted in the bricks

Analysis strategy

Bricks ordered by the probability to contain the neutrino interaction

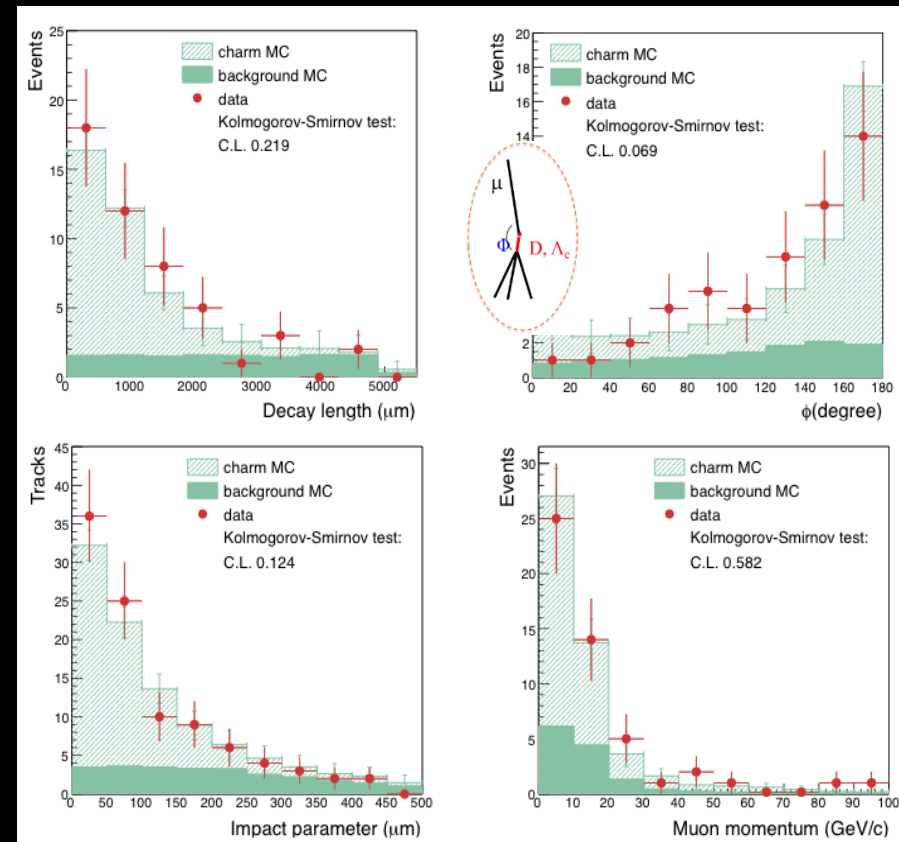
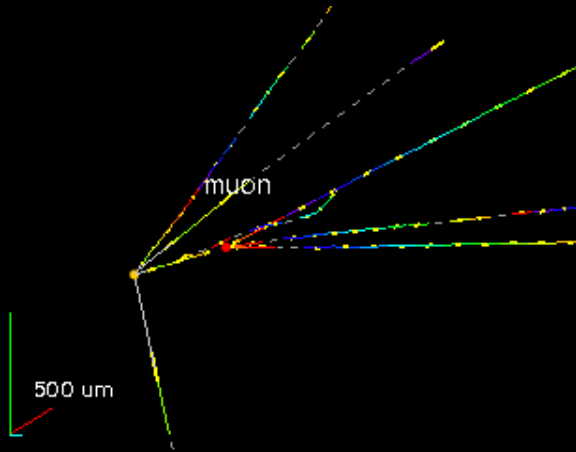
- **Runs 2008-2009:** 1st and 2nd most probable bricks
- **Runs 2010-2012:** 1st brick analysed so far. Extension to less probable bricks in progress.



Search for short-lived particle decays: charmed hadrons as control sample

Eur. Phys. J. C 74 (2014) 2986.

Charmed hadrons produced by ν_μ CC interactions
 → muon at the primary vertex
 Mass and lifetime charmed hadrons \sim tau lepton
 → similar decay topology



Decay topology	Events			Observed
	Expected charm	Expected background	Expected total	
1-prong	21 ± 2	9 ± 3	30 ± 4	19
2-prong	14 ± 1	4 ± 1	18 ± 1	22
3-prong	4 ± 1	1.0 ± 0.3	5 ± 1	5
4-prong	0.9 ± 0.2	-	0.9 ± 0.2	4
Total	40 ± 3	14 ± 3	54 ± 4	50

2008-2010 OPERA data set

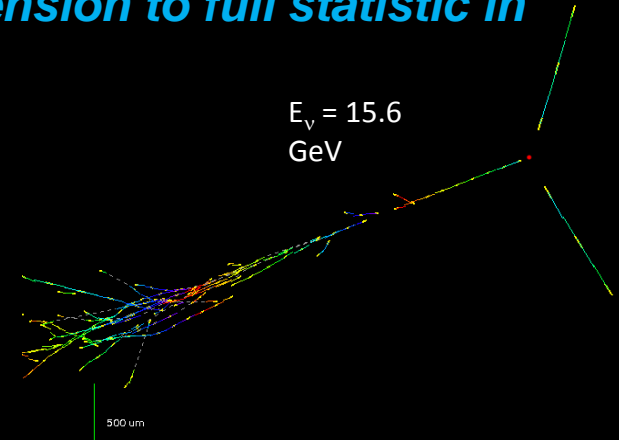
54 ± 4 charm events expected

50 observed in control sample

Good agreement between data and MC.

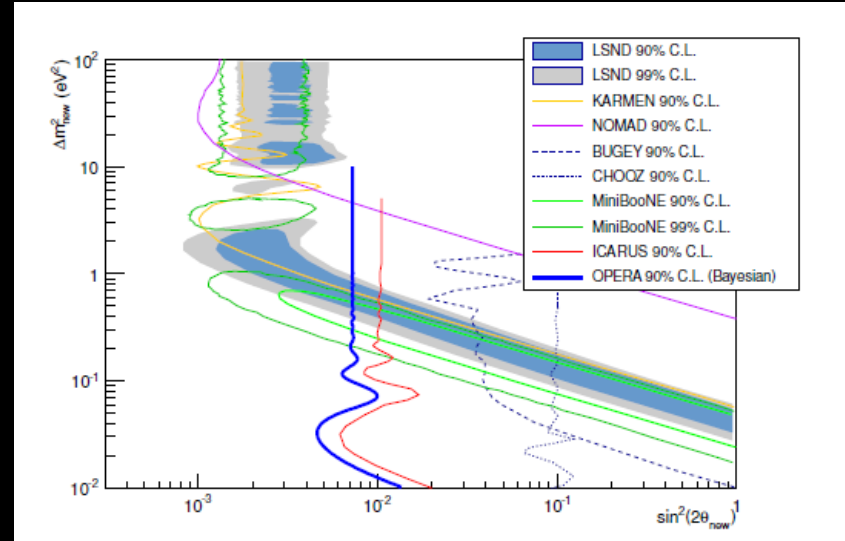
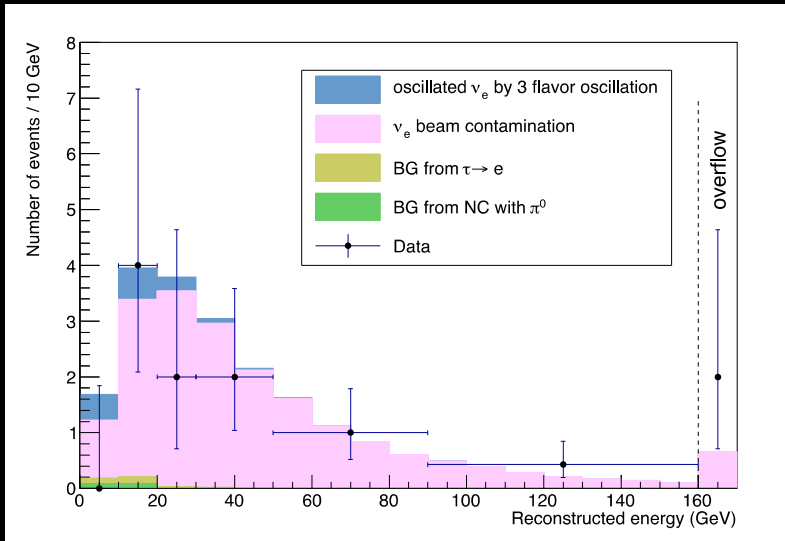
ν_e searched in **505** (data of 2008-2009 run ~ **50% full statistics**) neutrino interaction without the muon in the final state. **Extension to full statistic in progress.**

progress.		20 GeV	30 GeV	No cut
BG common to both analyses	BG (a) from π^0	0.2	0.2	0.2
	BG (b) from $\tau \rightarrow e$	0.2	0.3	0.3
	ν_e beam contamination	4.2	7.7	19.4
Total expected BG in 3-flavour oscillation analysis		4.6	8.2	19.8
BG to non-standard oscillation analysis only	ν_e via 3-flavour oscillation	1.0	1.3	1.4
	Total expected BG in non-standard oscillation analysis	5.6	9.4	21.3
Data		4	6	19



$\sin^2(2\theta_{13}) < 0.44$ @ 90% C.L.

$\sin^2(2\theta_{new}) < 7.2 \times 10^{-3}$ @ 90% C.L.



$\nu_\mu \rightarrow \nu_\tau$ oscillation analysis

Data sample: 4685 located and fully analysed ν interactions

	2008	2009	2010	2011	2012	Total
pot (10^{19})	1.74	3.53	4.09	4.75	3.86	17.97
0 μ events	148	250	209	223	149	979
1 μ events $p_\mu < 15$ GeV/c	534	1019	814	749	590	3706
Total of events	682	1269	1023	972	739	4685

Kinematical selection cuts kept fixed since beginning of the experiment.

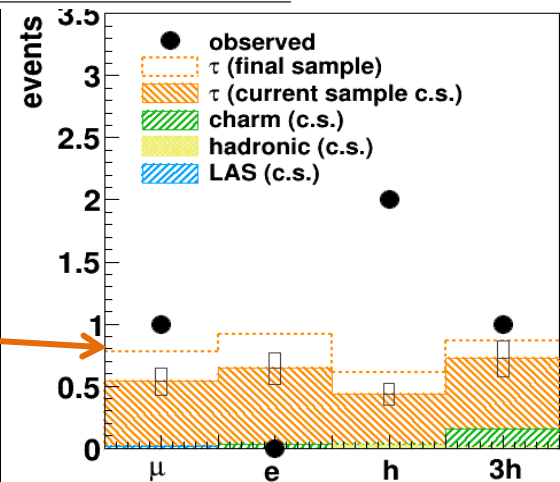
variable	$\tau \rightarrow 1h$	$\tau \rightarrow 3h$	$\tau \rightarrow \mu$	$\tau \rightarrow e$
lepton-tag		No μ or e at the primary vertex		
z_{dec} (μm)	[44, 2600]	< 2600	[44, 2600]	< 2600
p_T^{miss} (GeV/c)	< 1*	< 1*	/	/
ϕ_{lH} (rad)	> $\pi/2^*$	> $\pi/2^*$	/	/
p_T^{2ry} (GeV/c)	> 0.6(0.3)*	/	> 0.25	> 0.1
p^{2ry} (GeV/c)	> 2	> 3	> 1 and < 15	> 1 and < 15
θ_{kink} (mrad)	> 20	< 500	> 20	> 20
m, m_{min} (GeV/c ²)	/	> 0.5 and < 2	/	/

Expected ν_τ events: 2.1 ± 0.4

($\Delta m^2 = 2.32 \times 10^{-3} \text{ eV}^2$, $\theta_{23} = \pi/4$)

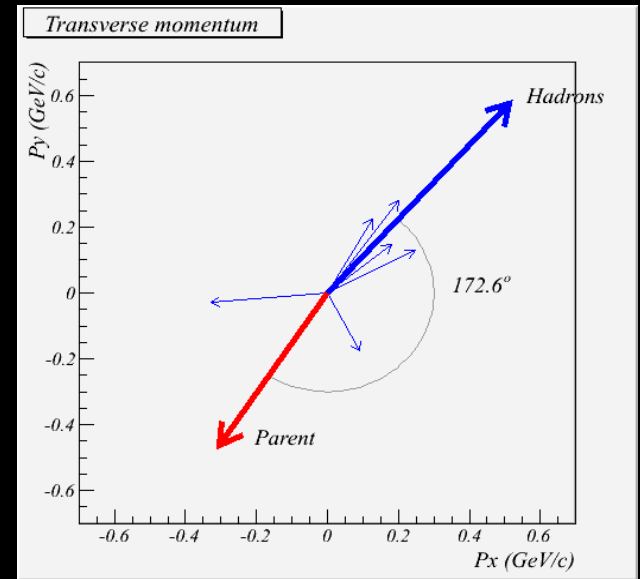
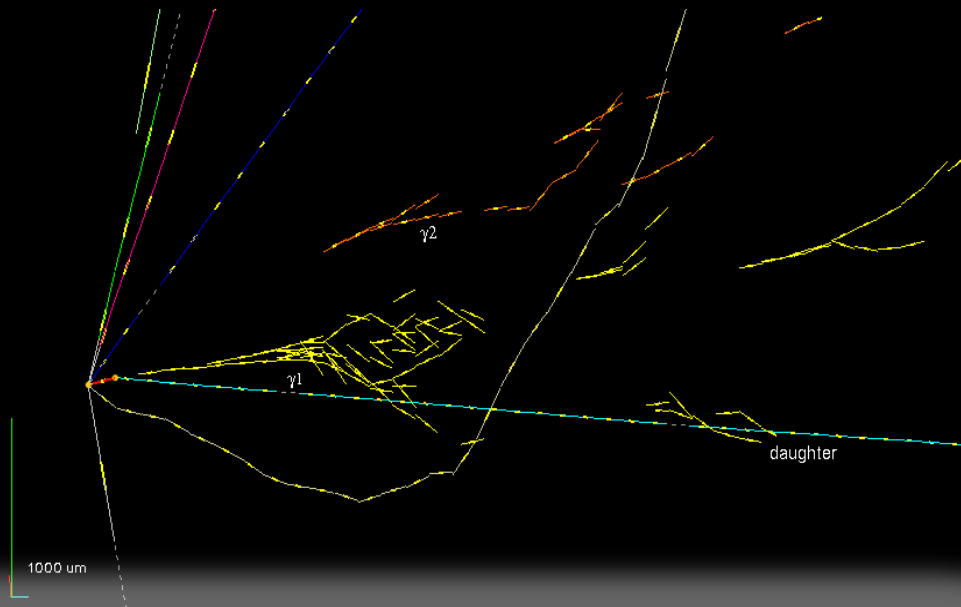
Observed ν_τ events: 4

After completion of 2nd bricks for 2010 ÷ 2012 runs,
expected events : 3.2 events



The first ν_τ candidate event

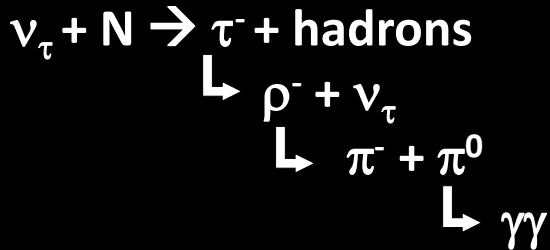
Phys. Lett. B (2010) 138



First direct detection of $\nu_\mu \rightarrow \nu_\tau$ oscillations in appearance mode

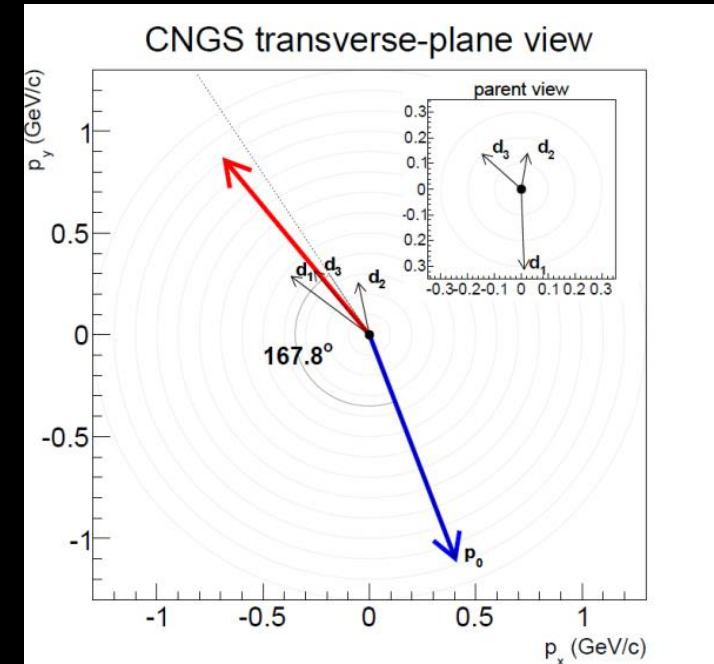
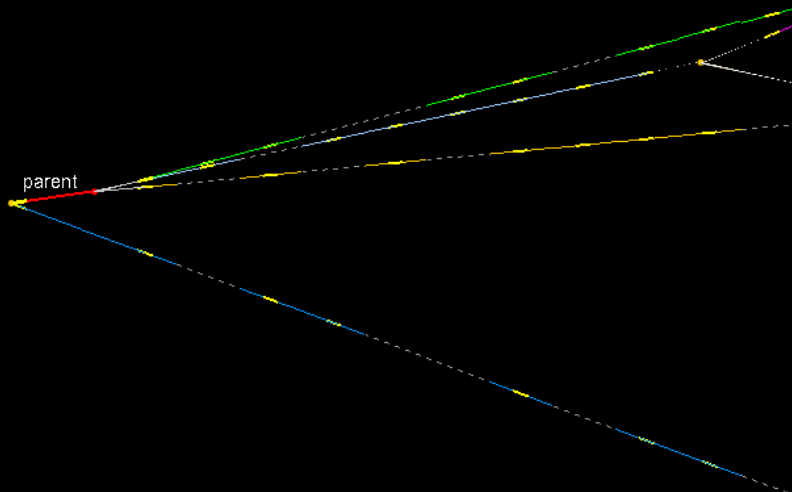
$$\gamma 1 + \gamma 2 \quad 120 \pm 20 \pm 35 \text{ MeV}$$

$$\pi + \gamma 1 + \gamma 2 \quad 640^{+125}_{-80} \text{ } ^{+100}_{-90} \text{ MeV}$$

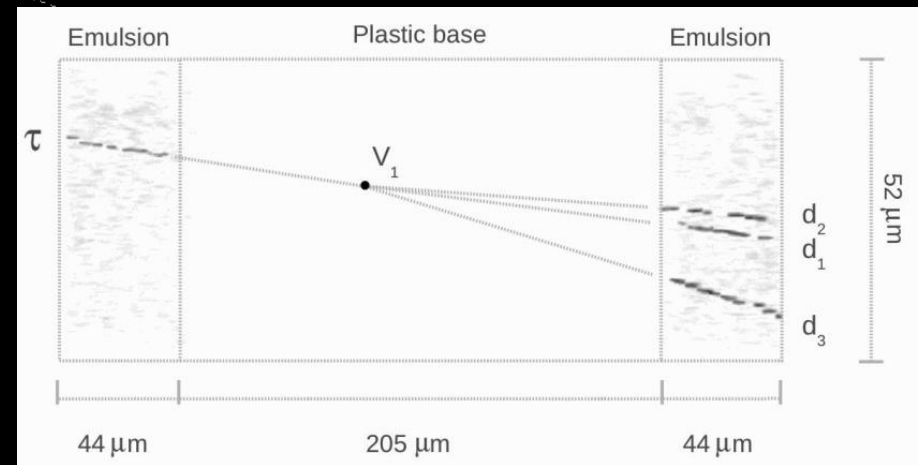
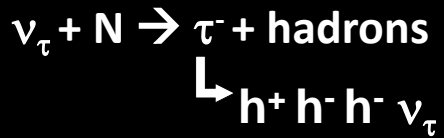


Variable	Value
Kink angle (mrad)	41 ± 2
Decay length (μm)	1335 ± 35
P daughter (GeV/c)	12^{+6}_{-3}
Pt (MeV/c)	470^{+230}_{-120}
Missing Pt at 1ry vertex (MeV/c)	570^{+320}_{-170}
φ (deg)	173 ± 2

The second ν_τ candidate event JHEP 11 (2013) 036

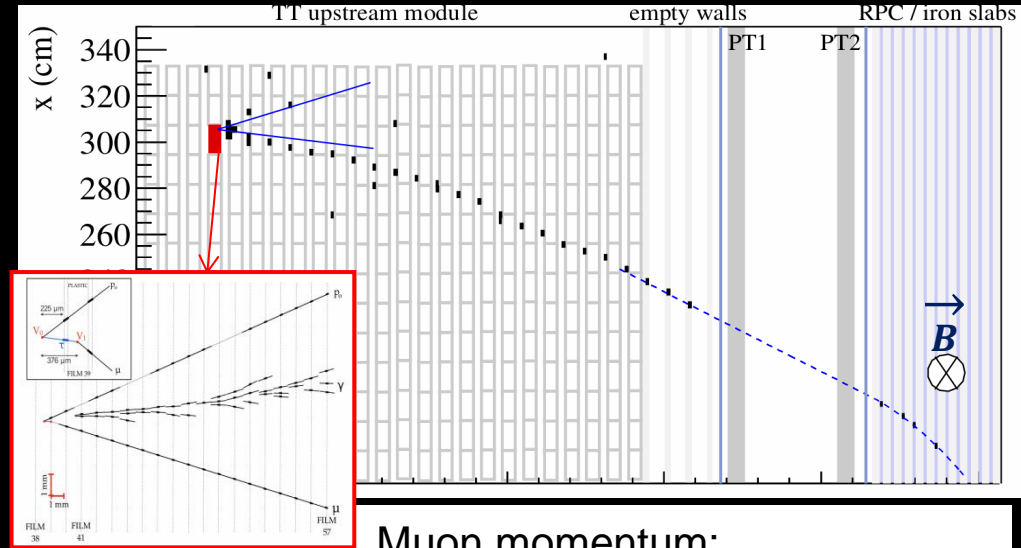
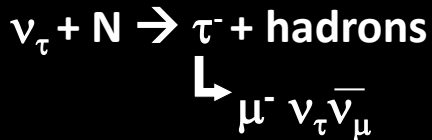


2000
μm



The third ν_τ candidate event

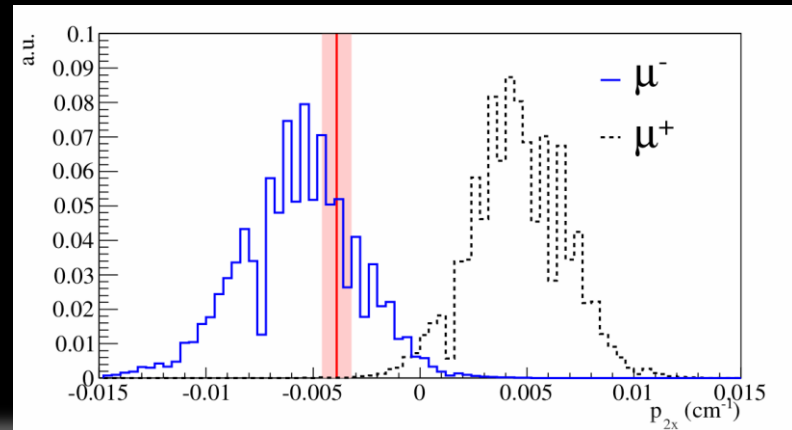
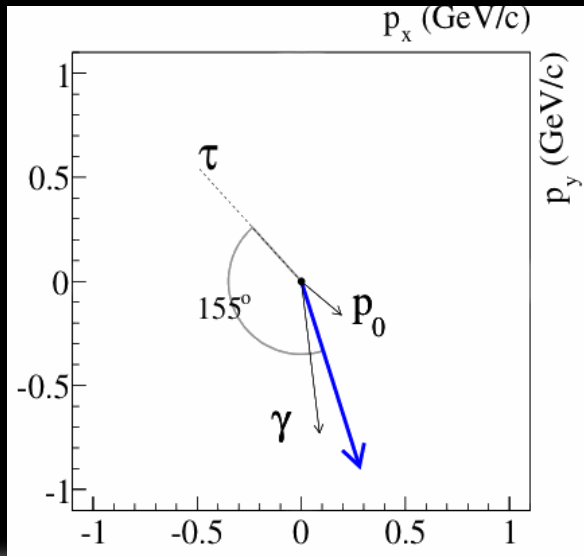
Phys. Rev. D 89 (2014) 051102(R)



Muon momentum:

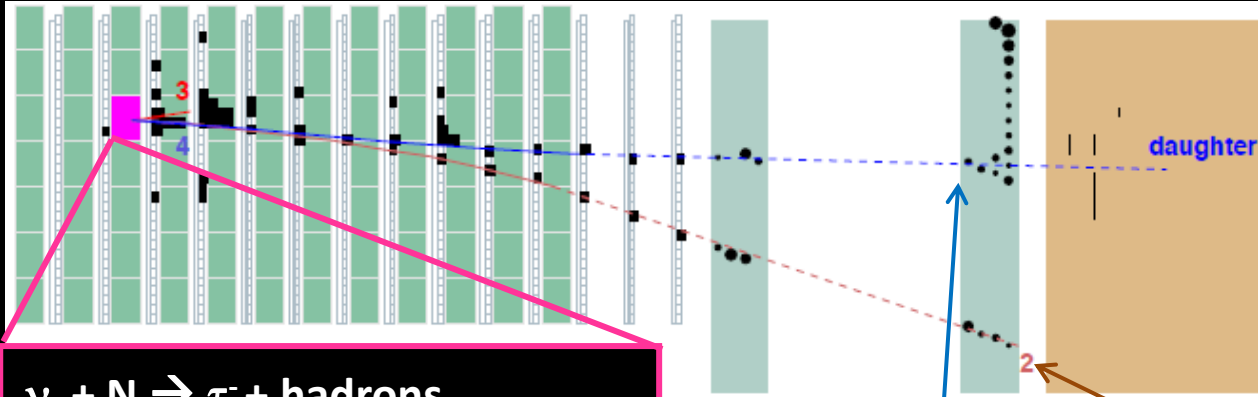
by range (electronic detectors) 2.8 ± 0.2 GeV/c
 by MCS in emulsion 3.1 [2.6, 4.0] GeV/c

Negative muon measured in the muon spectrometer

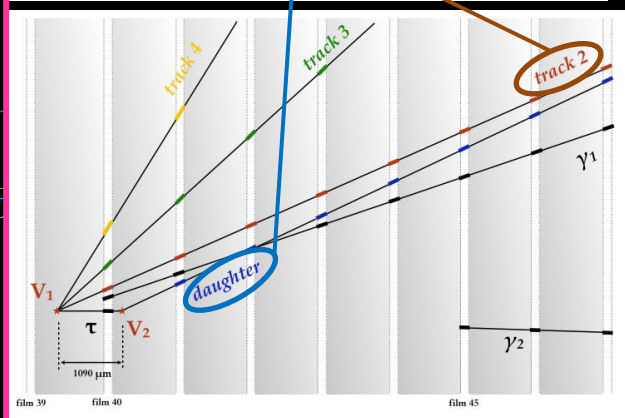
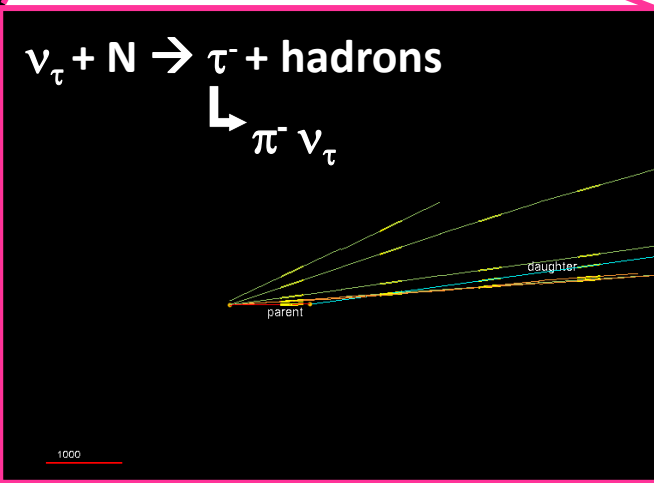


First measurement of the lepton charge in appearance mode

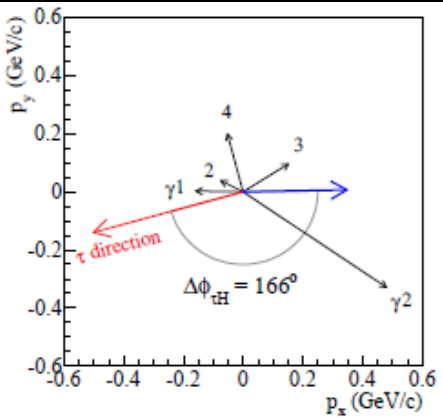
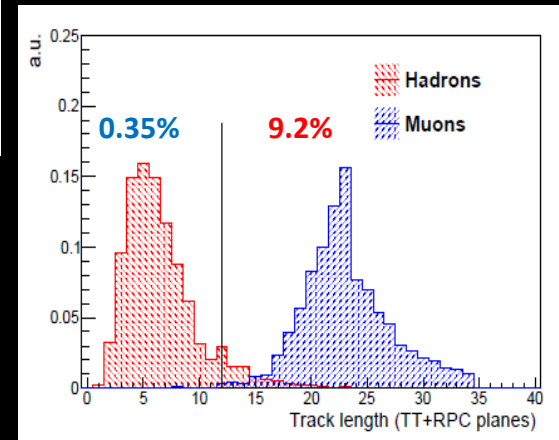
The fourth ν_τ candidate event



Track 3 visible interaction in downstream brick #2
Track 4 crosses 1 wall; identified as a proton based on its ionization



Track 2 from neutrino interaction vertex, $p = 1.9$ GeV crossing 9 wall and stopping in first iron slab of the magnet.

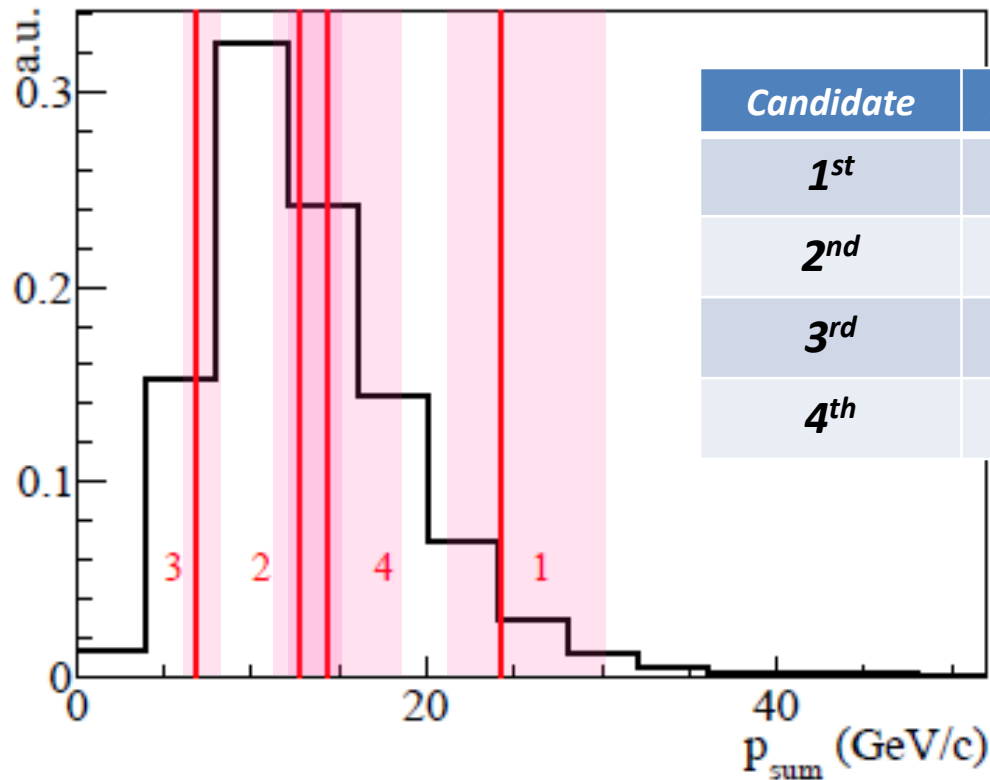


Variable	Selection	Measured value
θ_{kink} (mrad)	> 20	137 ± 4
z_{dec} (μm)	< 2600	406 ± 30
p_{2ry} (GeV/c)	> 2	$6.0^{+2.2}_{-1.2}$
p_T^{2ry} (GeV/c)	> 0.6 (0.3*)	$0.82^{+0.30}_{-0.16}$
p_T^{miss} (GeV/c)	< 1	$0.55^{+0.30}_{-0.20}$
$\Delta\phi_{\tau H}$ (degrees)	> 90	166^{+2}_{-31}

Summary of four events

Visible energy of all ν_τ events:

Scalar sum of momentum and γ energies



Candidate	Decay Mode	$P_{\text{sum}}(\text{GeV}/c)$	year
1 st	$\tau \rightarrow h$	$24.3^{+6.1}_{-3.2}$	2010
2 nd	$\tau \rightarrow 3h$	$12.7^{+2.3}_{-1.7}$	2012
3 rd	$\tau \rightarrow \mu^-$	$6.8^{+0.9}_{-0.6}$	2013
4 th	$\tau \rightarrow h$	$14.4^{+3.9}_{-2.7}$	2014

Significance

The expected signal and background is normalized to **the number of located events**

$$n^{0\mu}(\nu_{\tau}^{CC}) = \frac{\langle \sigma(\nu_{\tau}^{CC}) \rangle}{\langle \sigma(\nu_{\mu}^{CC}) \rangle} \frac{\langle \epsilon^{0\mu}(\nu_{\tau}^{CC}) \rangle}{\langle \epsilon^{0\mu}(\nu_{\tau}^{CC}) \rangle + \alpha \langle \epsilon^{0\mu}(\nu_{\tau}^{NC}) \rangle} n^{0\mu} \quad \alpha = \frac{NC}{CC}$$

Decay channel	Expected signal $\Delta m_{23}^2 = 2.32 \text{ meV}^2$	Total background	Observed
$\tau \rightarrow h$	0.41 ± 0.08	0.033 ± 0.006	2
$\tau \rightarrow 3h$	0.57 ± 0.11	0.155 ± 0.030	1
$\tau \rightarrow \mu$	0.52 ± 0.10	0.018 ± 0.007	1
$\tau \rightarrow e$	0.62 ± 0.12	0.027 ± 0.005	0
Total	2.11 ± 0.42	0.233 ± 0.041	4

Two statistical methods

- Fisher combination of single channel → $p\text{-value} = 1.24 \times 10^{-5}$
- Likelihood ratio → $p\text{-value} = 1.03 \times 10^{-5}$

Observation of tau appearance with 4.2σ

First measurement of Δm_{32}^2 in appearance experiment

$$N_{\nu_\tau} \propto \int \phi(E) \sin^2\left(\frac{\Delta m_{32}^2 L}{4E}\right) \epsilon(E) \sigma(E) dE$$

$$\propto (\Delta m_{32}^2)^2 L^2 \int \phi(E) \epsilon(E) \frac{\sigma(E)}{E^2} dE$$

OPERA Off-peak
 $L/\langle E \rangle \sim 43 \text{ Km/GeV}$
 $(L/\langle E \rangle)_{\text{peak}} \sim 500 \text{ Km/GeV}$

→ Strong dependence on Δm^2 → measure Δm^2 with counting experiment

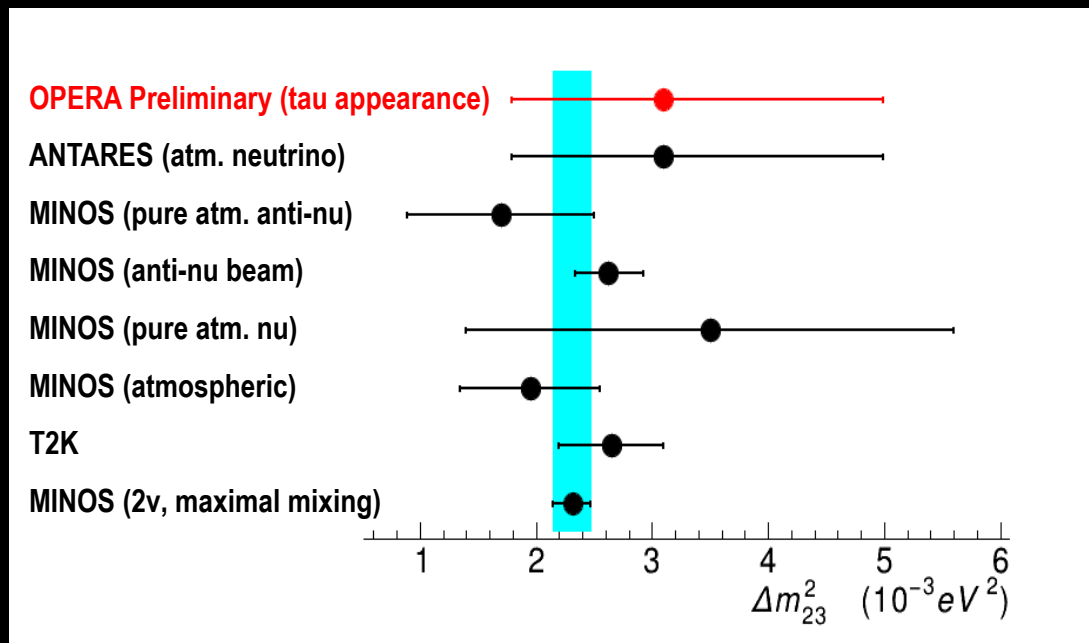
90% CL intervals on Δm_{32}^2 assuming $\sin^2(2\theta_{23}) = 1$

Feldman & Cousin

$[1.8 - 5] \times 10^{-3} \text{ eV}^2$

Bayesian

$[1.9 - 5] \times 10^{-3} \text{ eV}^2$



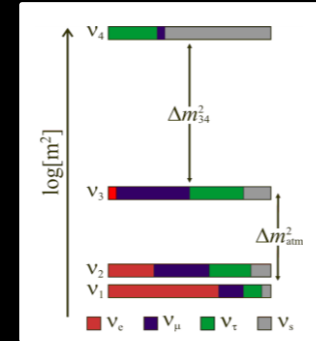
Sterile neutrino searches through $\nu_\mu \rightarrow \nu_\tau$ oscillation

PRELIMINARY

❖ Assume there is an additional sterile neutrino and an additional mass scale

❖ Mixing matrix is extended to:

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \nu_s \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} & U_{e4} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} & U_{\mu4} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} & U_{\tau4} \\ U_{s1} & U_{s2} & U_{s3} & U_{s4} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \nu_4 \end{pmatrix}$$



(3+1) model

$$\Delta m_{41}^2 > |\Delta m_{31}^2|$$

$$\Delta m_{32}^2 = 2.32 \times 10^{-3} \text{ eV}^2$$

For $L = 730 \text{ km}$ and $\langle E_\nu \rangle = 17 \text{ GeV}$
 $P_{\nu_\mu \rightarrow \nu_\tau}$ independent from Δm_{21}^2 since

$$\frac{\Delta m_{21}^2 L}{E} \sim 0$$

$$\sin^2 2\theta_{\mu\tau} = \sin^2 2\theta_{24} \sin^2 \theta_{34}$$

$$P_{\nu_\mu \rightarrow \nu_\tau} = 4|U_{\mu 3}|^2|U_{\tau 3}|^2 \sin^2 \frac{\Delta_{31}}{2} + 4|U_{\mu 4}|^2|U_{\tau 4}|^2 \sin^2 \frac{\Delta_{41}}{2} + 2\Re[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*] \sin \Delta_{31} \sin \Delta_{41} - 4\Im[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*] \sin^2 \frac{\Delta_{31}}{2} \sin \Delta_{41} + 8\Re[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*] \sin^2 \frac{\Delta_{31}}{2} \sin^2 \frac{\Delta_{41}}{2} + 4\Im[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*] \sin \Delta_{31} \sin^2 \frac{\Delta_{41}}{2}$$

$$\Delta_{ij} = \frac{\Delta m_{ij}^2 L}{2E}$$

$$\hat{\delta} = \arg[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*]$$

$$4\Im[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*] = 4\sqrt{|U_{\mu 4}|^2|U_{\tau 4}|^2|U_{\mu 3}|^2|U_{\tau 3}|^2} \sin \hat{\delta} = 2 \sin \theta_{\mu\tau} \sqrt{|U_{\mu 3}|^2|U_{\tau 3}|^2} \sin \hat{\delta}$$

$$8\Re[U_{\mu 4}^* U_{\tau 4} U_{\mu 3} U_{\tau 3}^*] = 8\sqrt{|U_{\mu 4}|^2|U_{\tau 4}|^2|U_{\mu 3}|^2|U_{\tau 3}|^2} \cos \hat{\delta} = 4 \sin \theta_{\mu\tau} \sqrt{|U_{\mu 3}|^2|U_{\tau 3}|^2} \cos \hat{\delta}$$

An upper limit on $\sin^2(2\theta_{\mu\tau})$ at high Δm_{41}^2 values*

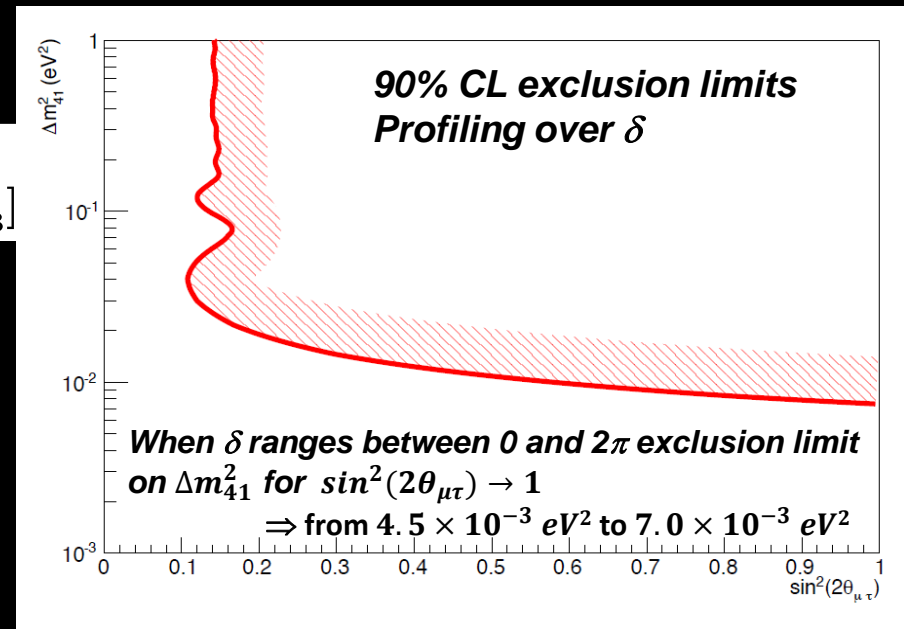
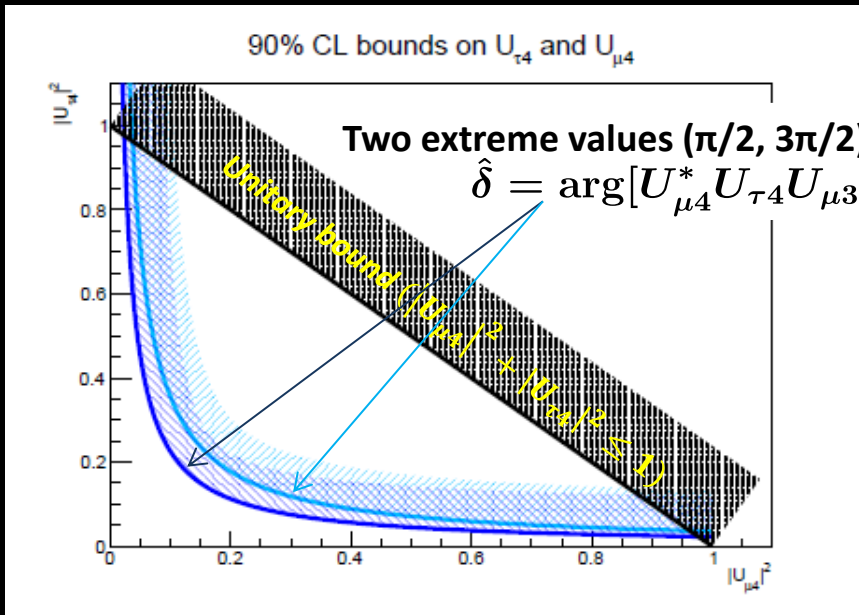
Profile likelihood $\tilde{L}(\delta, \sin^2(2\theta_{\mu\tau})) \rightarrow$ maximizing $L(\delta, \sin^2(2\theta_{\mu\tau}), C^2) = \frac{\mu^n}{n!} e^{-\mu}$ on C^2 between 0 and 1, $C = 2|U_{\mu 3} U_{\tau 3}^*|$, n number of candidate events (4), μ is the expected number of events given by

$$\mu = N_b + \int A \phi(E) P_{\nu_\mu \rightarrow \nu_\tau}(E) \sigma(E) \epsilon(E) dE$$

Expected number of background (0.23), Normalization factor, neutrino flux, oscillation probability, CC ν_τ cross-section, tau detection efficiency.

$$\chi^2 = -2 \ln(\tilde{L}(\delta, \sin^2(2\theta_{\mu\tau})) / L_0) \text{ where } L_0 = \frac{n^n}{n!} e^{-n}$$

PRELIMINARY



Conclusions

❖ OPERA experiment:

successful data taking from 2008 to 2012, $17.97 \cdot 10^{19}$ pot collected.

❖ $\nu_\mu \rightarrow \nu_e$ analysis:

- 19 events (4 events with $E_{rec} < 20 \text{ GeV}$) observed with an estimation of 19.8 ± 2.8 (4.6 with $E_{rec} < 20 \text{ GeV}$) events.
- At large Δm^2 region, 6 events observed with an estimation of 9.4 ± 1.3 events gives an upper limit of $7.2 \cdot 10^{-3}$ at 90% C.L. on $\sin^2(2\theta_{new})$.

❖ $\nu_\mu \rightarrow \nu_\tau$ analysis:

- 4685 located events fully analysed
- 4 ν_τ candidates observed so far with an expectation of 2.1 ± 0.4 events and 0.23 background.
- No oscillation hypothesis excluded at 4.2σ

Observation of $\nu_\mu \rightarrow \nu_\tau$ oscillation in appearance mode

❖ First measurement of Δm_{32}^2 in appearance mode ($\nu_\mu \rightarrow \nu_\tau$ oscillations):

- $\Delta m_{32}^2 = [1.8 - 5] \times 10^{-3} \text{ eV}^2$ (90% C.L.) for $\sin^2(2\theta_{23}) = 1$

❖ Sterile neutrino search in $\nu_\mu \rightarrow \nu_\tau$: PRELIMINARY

- First limits on $|U_{\mu 4}|^2 |U_{\tau 4}|^2$ from direct detection of $\nu_\mu \rightarrow \nu_\tau$ oscillation.



Thank you for your attention

Image taken using OPERA emulsion film with pinhole hand made camera by D. Diferdinando