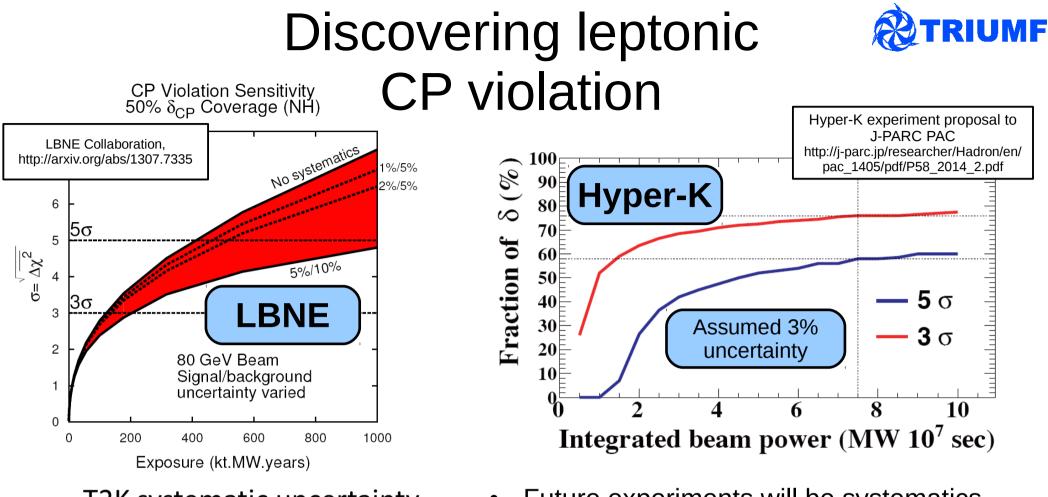


# vPRISM: A new way of probing neutrino interactions

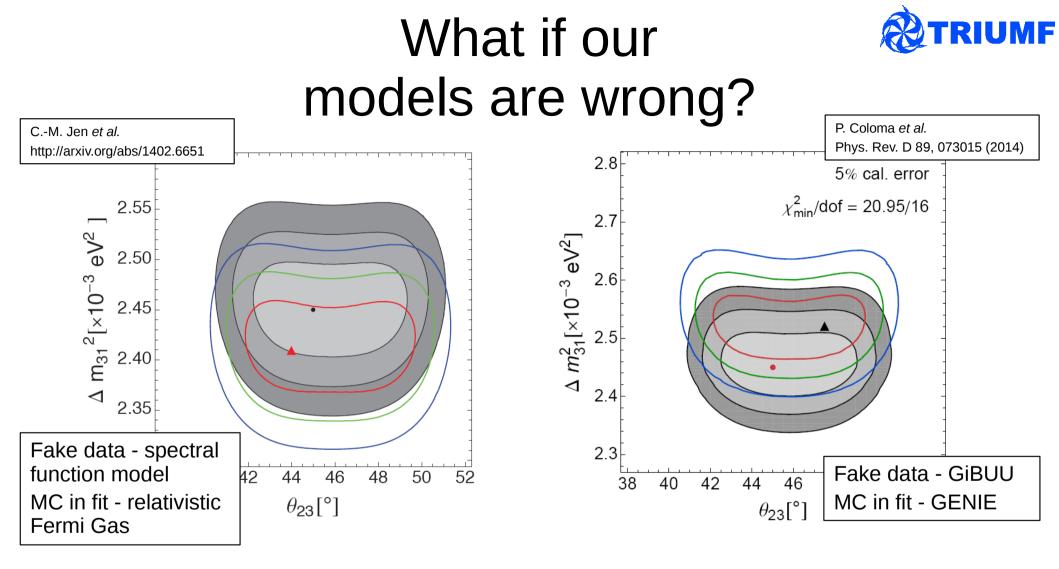
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#### T2K systematic uncertainty

			$v_{\mu}$ sample	$v_{e}$ sample	
$\boldsymbol{\nu}$ flux and cross section	w/o ND measurement		21.8%	26.0%	
	w/ ND measurement		2.7%	3.1%	
$\nu$ cross section due to difference of nuclear target btw. near and far			5.0%	4.7%	
Final or Secondary Hadronic Interaction			3.0%	2.4%	
Super-K detector			4.0%	2.7%	
total	w/o ND measurement		23.5%	26.8%	
	w/ ND measurement		7.7%	6.8%	
F	ractional error on number-of	-eve	nt prediction		

- Future experiments will be systematics limited
- To meet P5 requirement (75%  $\delta$ CP coverage at  $3\sigma$ ):
  - Hyper-K = 3% total uncertainty
  - LBNF = 1% total uncertainty
- Neutrino interaction uncertainties dominate •

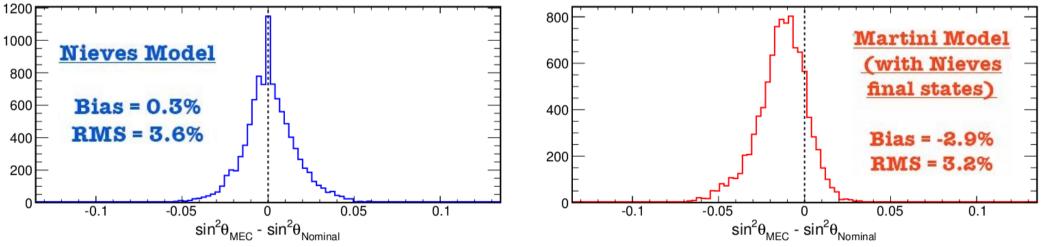


- Simulate a "T2K-like" experiment for muon neutrino disappearance analysis:
  - Shaded regions perform fit using same model as fake data
  - Coloured lines Fit results using a different nuclear model
  - Triangle/Dot Best fit point

### T2K multi-nucleon study



- MC-based analysis using full detector simulation, full systematics etc.
- Three fake datasets
  - Nominal NEUT MC
  - NEUT + meson exchange current (MEC) events from Nieves' model -Phys. Rev. C, 83:045501, Apr 2011
  - NEUT + MEC events from Martini's model -Phys. Rev. C, 81:045502, Apr 2010
  - Perform disappearance fit to extract  $\theta_{23}$  in each case and compare



• Both models give ~3.5% RMS in  $\sin^2 \theta_{23}$ , Martini model introduces ~3% bias

• Effects much smaller than current statistical uncertainty, but maybe large for future analyses

Mark Scott, NuFact 2014, Glasgow



# The problem with neutrinos...

- Measuring neutrino interactions is hard
- Want to know cross section as function of interaction variables:
  - Neutrino energy
  - Momentum transfer (Q<sup>2</sup>)
- Very hard (impossible) to measure these experimentally usually rely on the models we are trying to tune!
- Is there a better way?

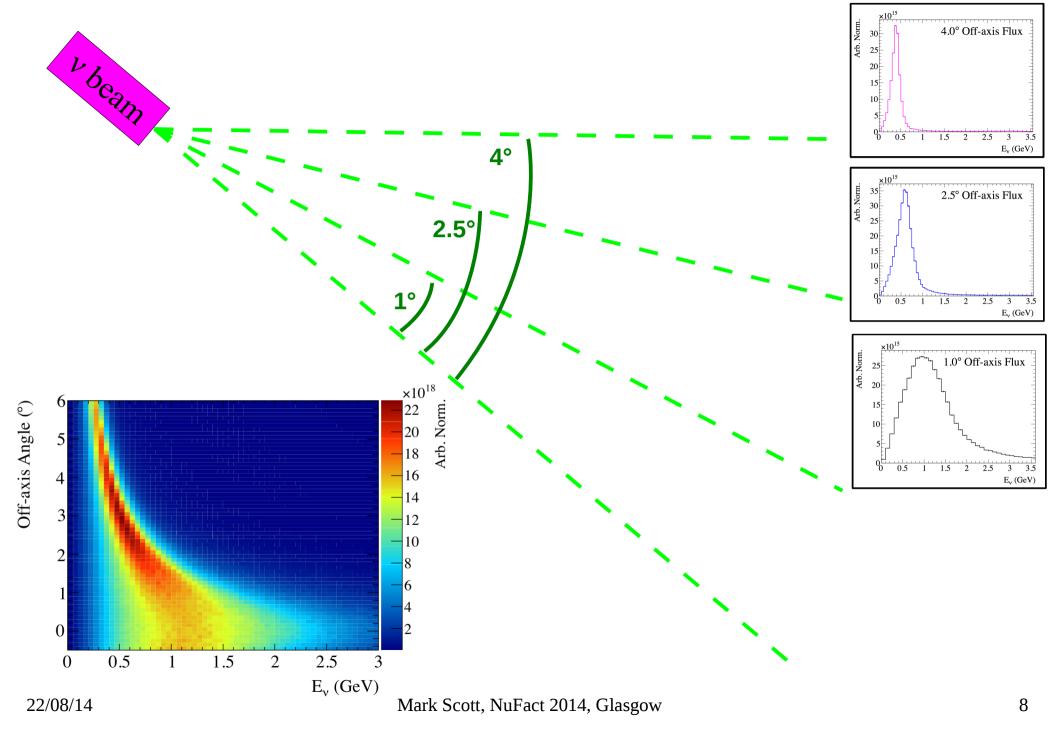


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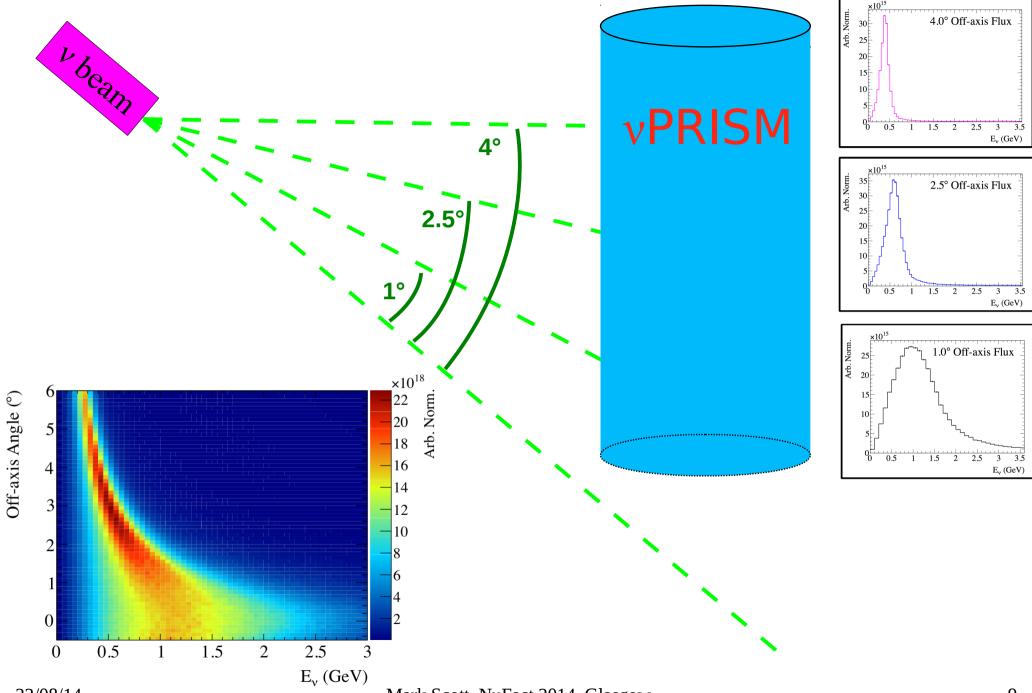
v beam

v beam **4**° **2.5**° • Works with any neutrino beam from pion decay So far studied in the • context of T2K



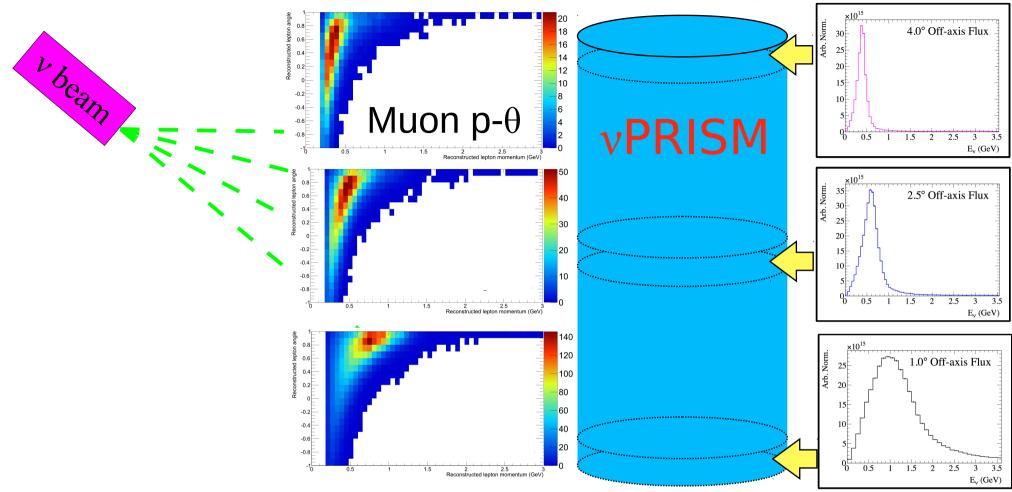


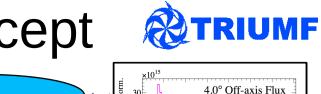


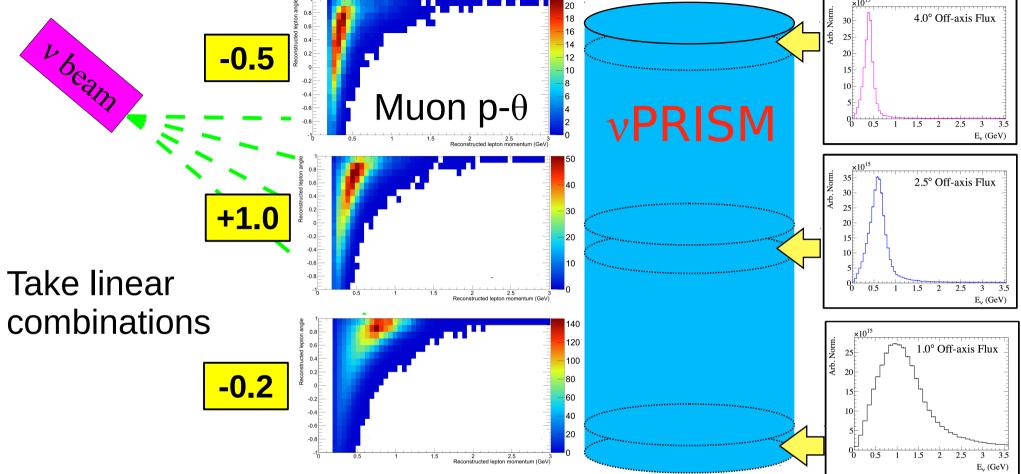


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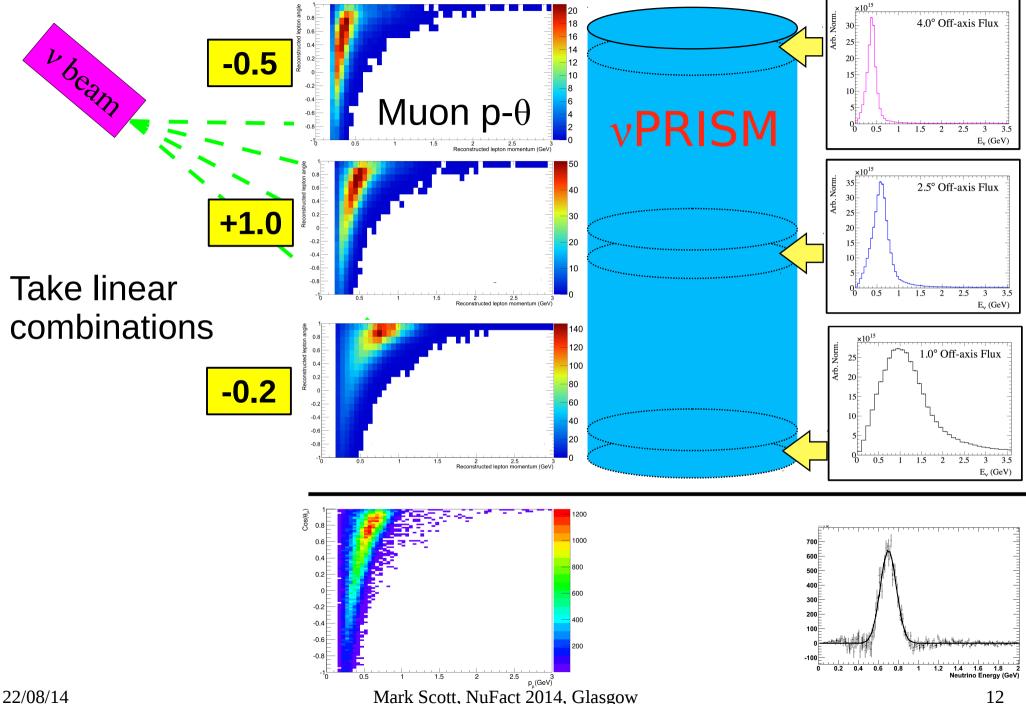
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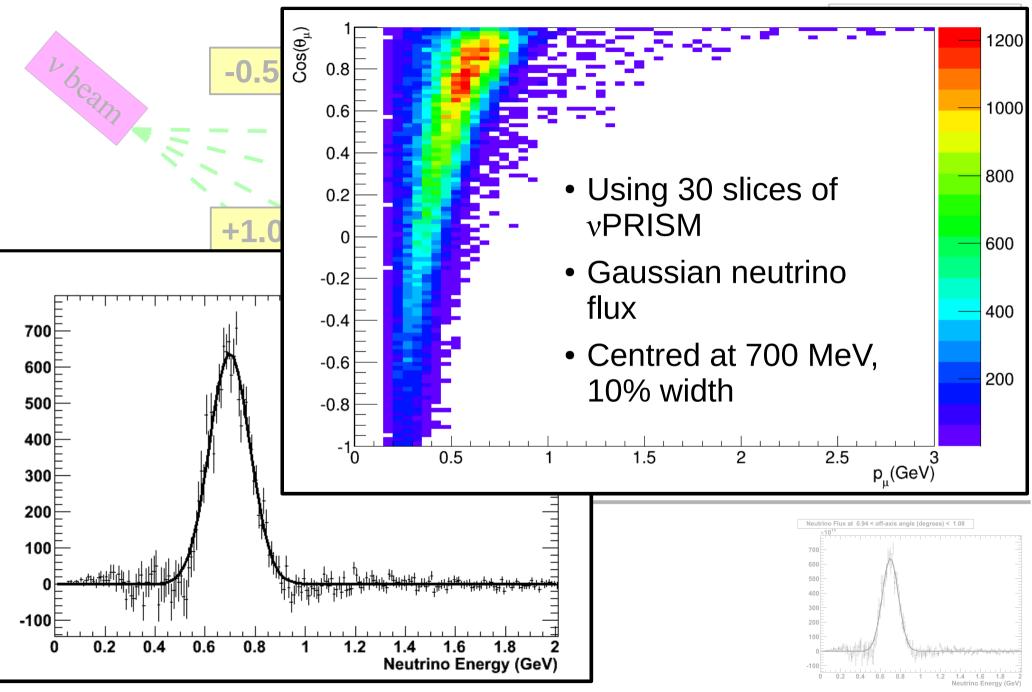




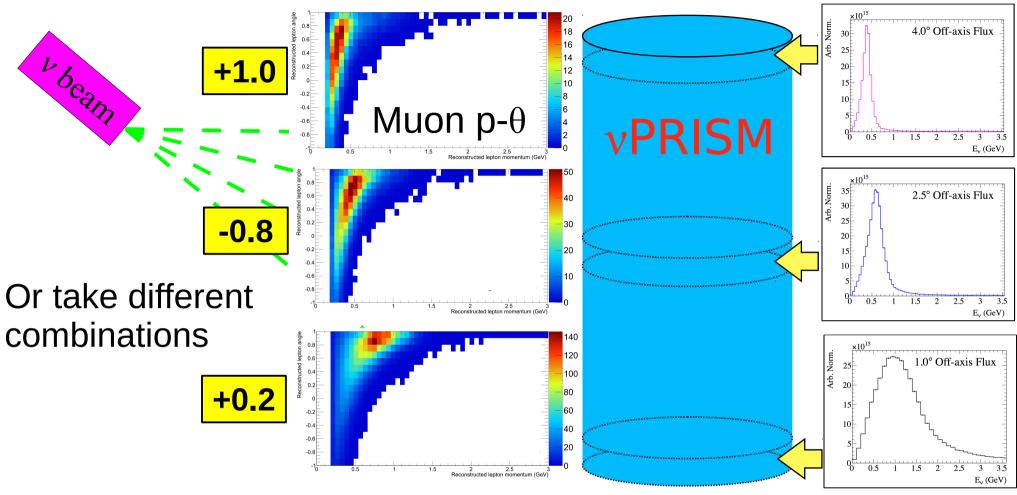


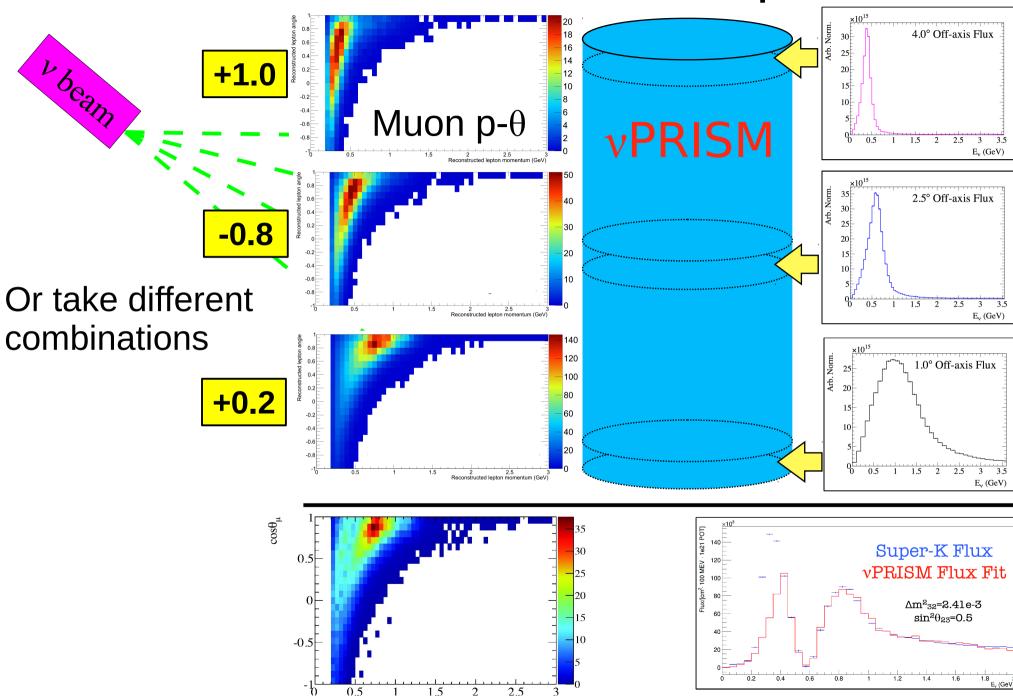




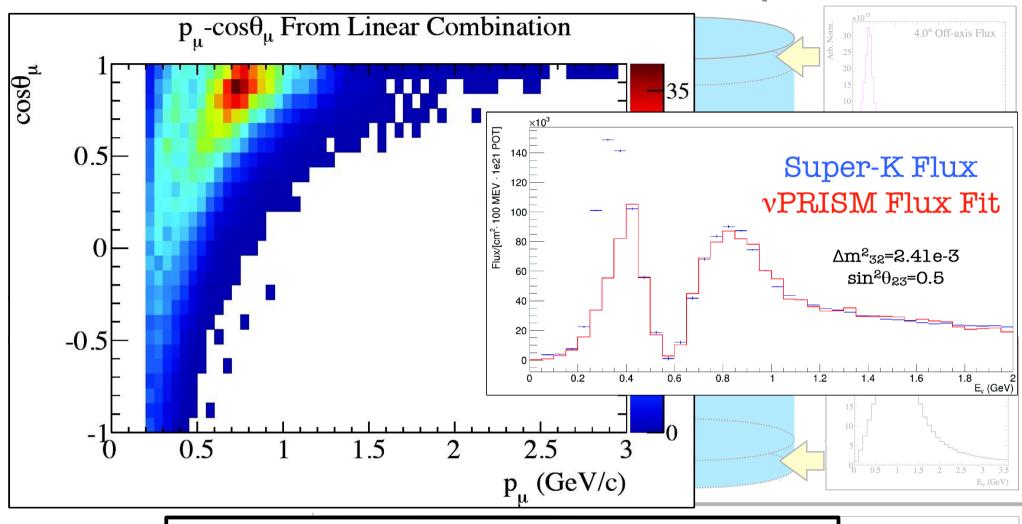


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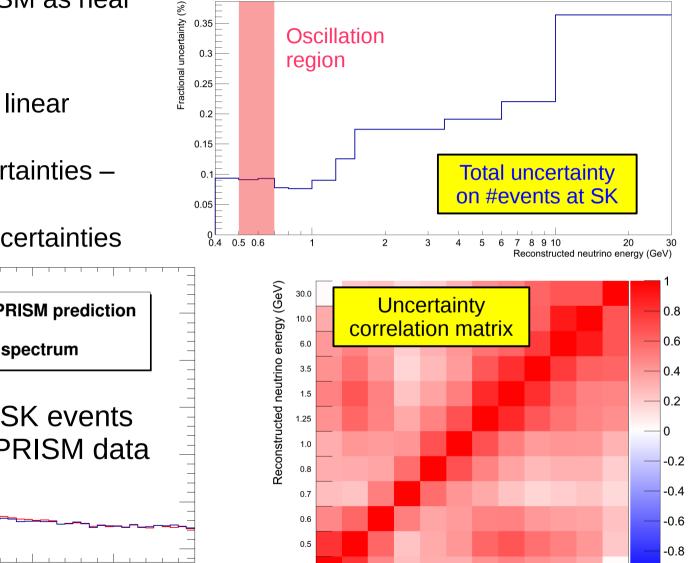
- Recreate oscillated neutrino flux at SK using near detector
- Directly measure muon  $p-\theta$  for given value of oscillation parameters

**RISM Flux Fit** 

 $\Delta m^{2}_{32}=2.41e-3$ 

### vPRISM disappearance analysis

- Full analysis using vPRISM as near detector for T2K
- Take into account:
  - Statistical error from linear combinations
  - Neutrino beam uncertainties direction, flux etc.
  - Interaction model uncertainties

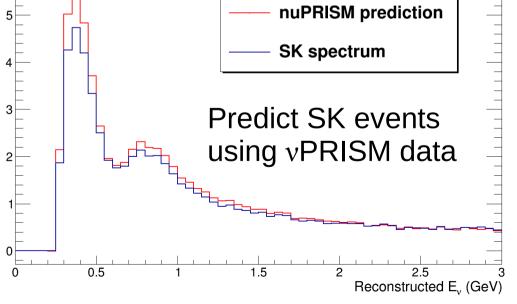


0.4

0.4 0.5 0.6 0.7

0.8

1.0 1.25 1.5 3.5





# Events at SK

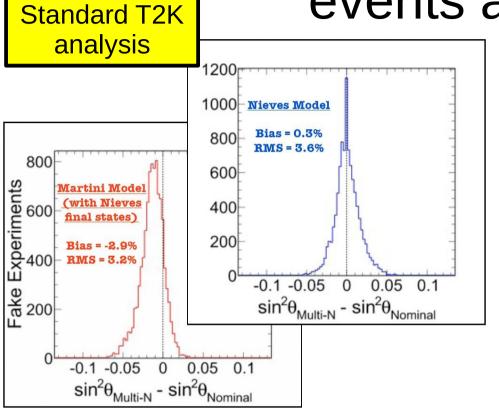
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10.0 30.0

6.0

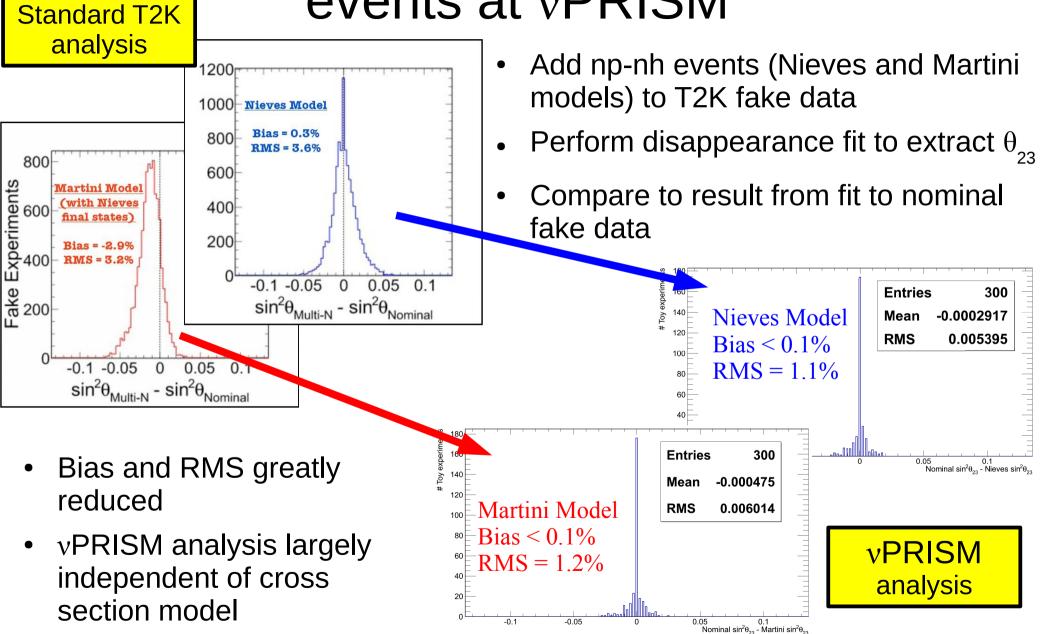
Reconstructed neutrino energy (GeV)

# Effect of multi-nucleon & events at vPRISM



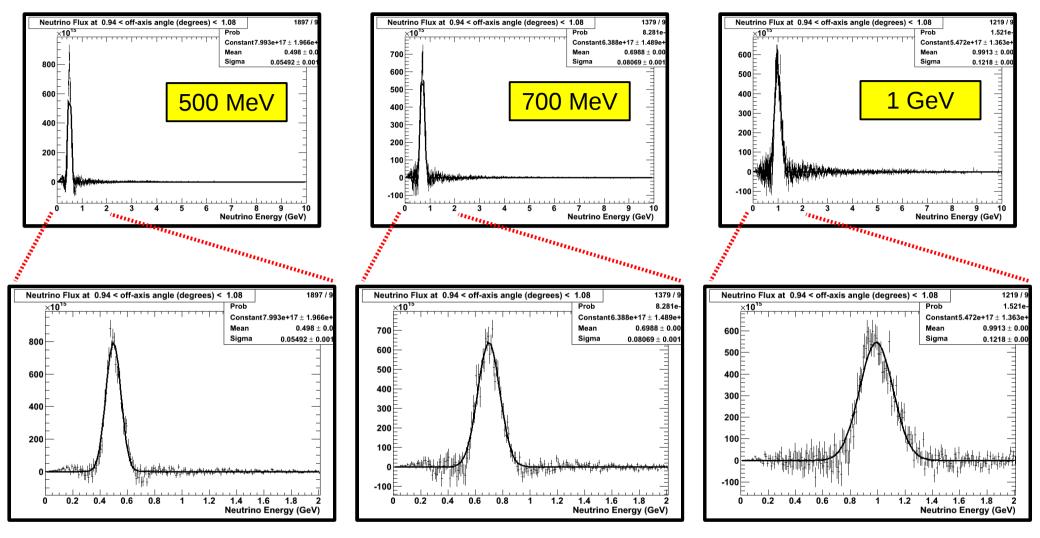
- Add np-nh events (Nieves and Martini models) to T2K fake data
- Perform disappearance fit to extract  $\theta_{23}$
- Compare to result from fit to nominal fake data

# Effect of multi-nucleon events at vPRISM





### A neutrino spectrometer



- Gaussian spectra from ~0.4 GeV to ~1 GeV
  - Depends on off-axis span of vPRISM:  $6^{\circ}$  0.25 GeV,  $0^{\circ}$  1.2 GeV
- High energy tail cancelled in all cases



# The problem with neutrinos...

- Measuring neutrino interactions is hard
- Want to know cross section as function of interaction variables:
  - Neutrino energy
  - Momentum transfer (Q<sup>2</sup>)
- Very hard (impossible) to measure these experimentally usually rely on the models we are trying to tune!
- Is there a better way?



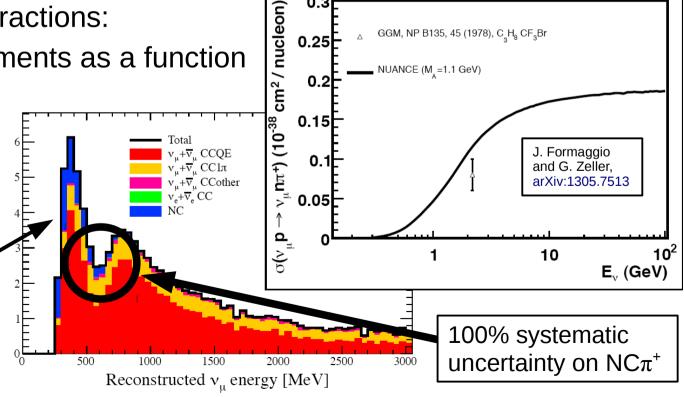
# The problem with neutrinos...

- Measuring neutrino interactions is hard
- Want to know cross section as function of interaction variables:
  - Neutrino energy
  - Momentum transfer (Q<sup>2</sup>)
- Very hard (impossible) to measure these experimentally usually rely on the models we are trying to tune!
- Is there a better way?

### Cross sections with a neutrino spectrometer



- We now know the neutrino energy!
  - Can calculate  $Q^2$  directly lepton information
- For charged current interactions:
  - Measure same cross section across a range of neutrino energies
  - Know correlations between energies
- For neutral current interactions:
  - First ever measurements as a function of  $E_v$
- fiTQun reconstruction algorithm  $\rightarrow$  measure pion kinematics
- Predicted oscillated spectrum at T2K far detector – broken down by interaction

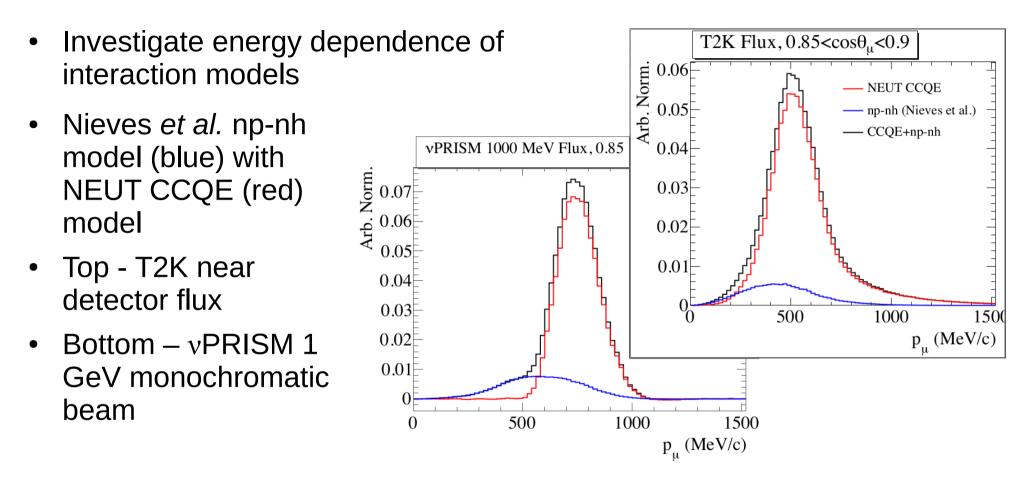


0.3

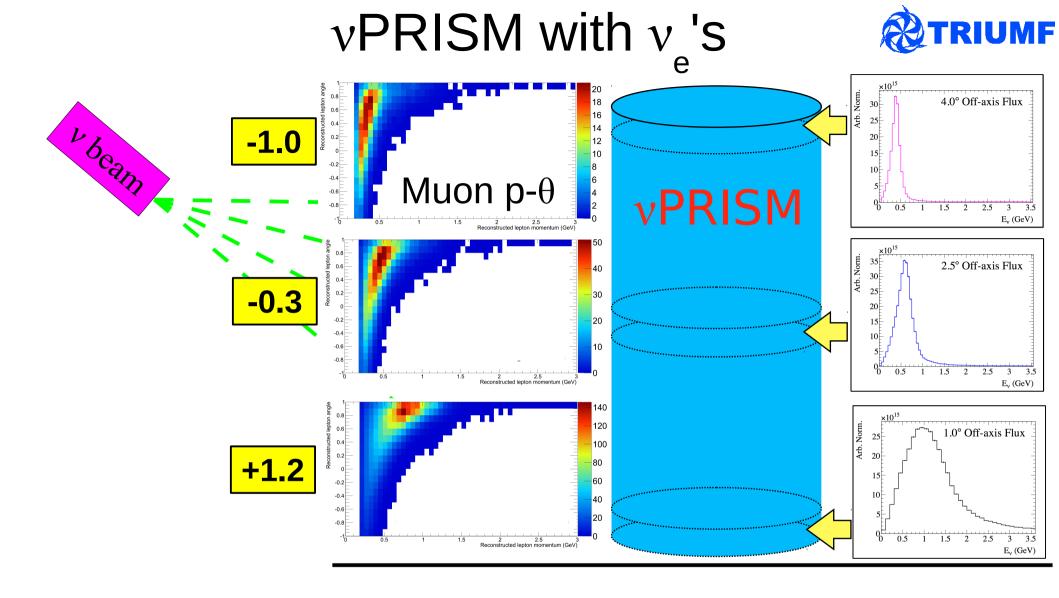
Mark Scott, NuFact 2014, Glasgow

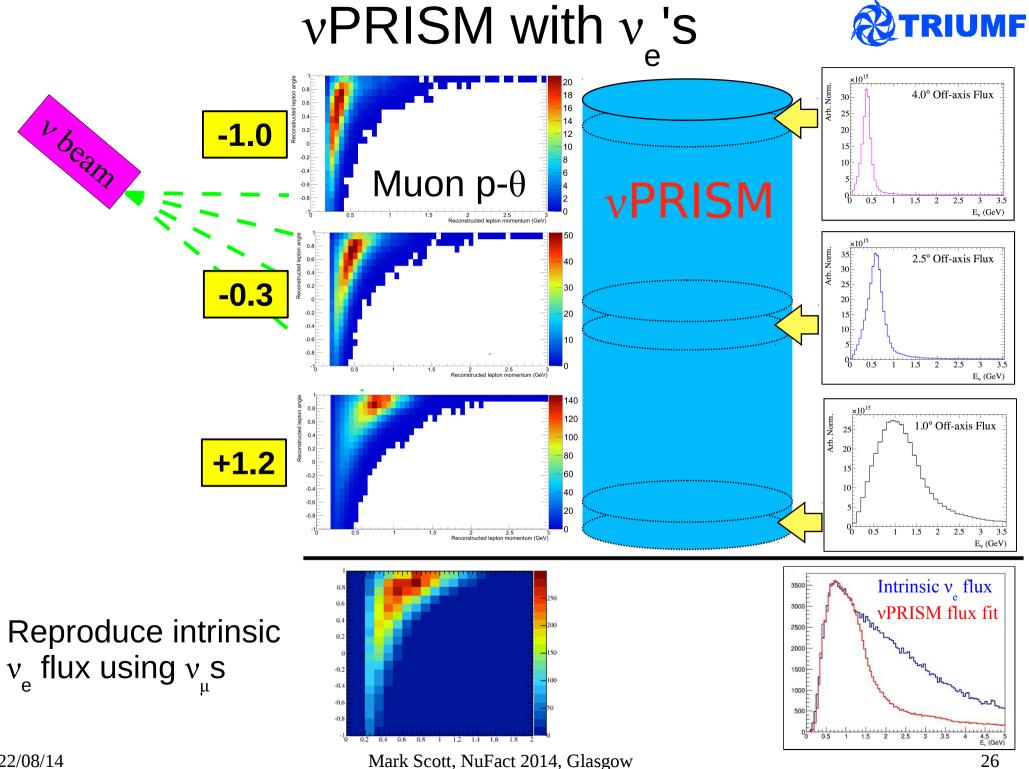


### Separating models

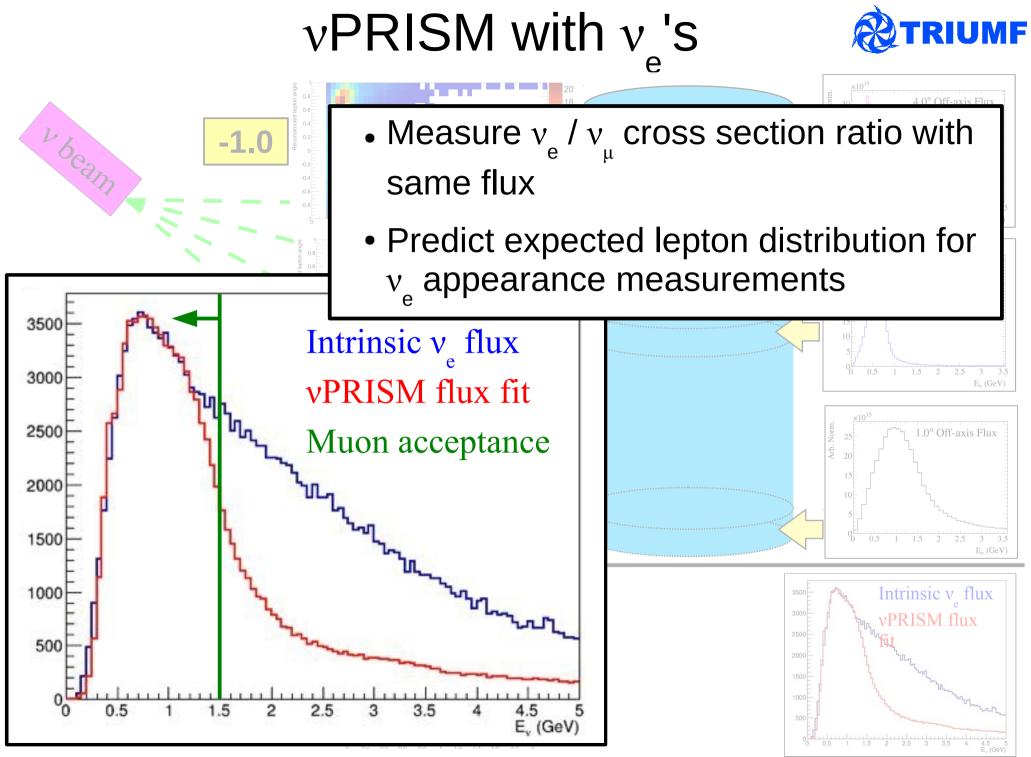


- np-nh model clearly different from CCQE model in vPRISM
- Measure across neutrino energy and lepton kinematic space lots of power to separate models





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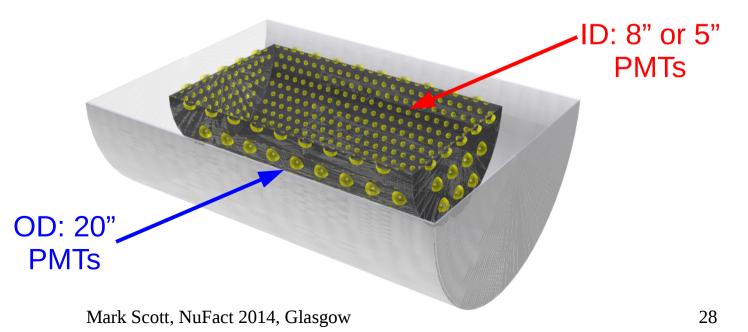


Mark Scott, NuFact 2014, Glasgow

# vPRISM-lite

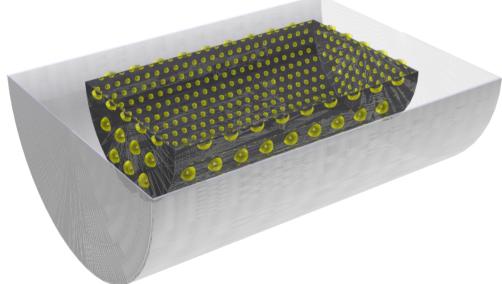


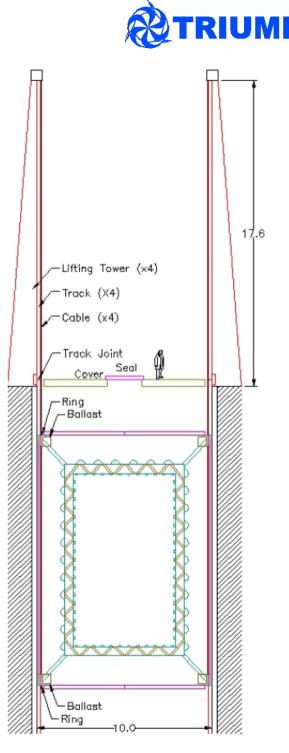
- Water Cherenkov detector spanning  $1^{\circ} 4^{\circ}$  off-axis
  - T2K beam points 3.6° downward, so get 4° off-axis for free
  - 52.5m tall if 1km from neutrino target
  - Instrument movable cylinder:
    - Inner Detector (ID): 6 or 8m diameter, 10m tall
    - Outer Detector (OD): 10m diameter, 14m tall
  - OD surrounded by scintillator panels external veto



### Current status

- Expression of Interest (EoI) document written:
  - Detailed  $v_{\mu}$  disappearance results
  - Discussion of other physics applications
    - (CP violation, anti-v, sterile neutrinos See talk by J. Caravaca in oscillation session)
  - Preliminary detector design
- Currently preparing a full proposal





# Summary

- Precision neutrino physics requires a better understanding of neutrino interactions
- The vPRISM detector concept provides this:
  - Removes model bias in neutrino oscillation measurements
  - Measure NC cross section vs  $\mathsf{E}_{v}$
  - Separate models that are otherwise degenerate
  - Directly compare  $v_e$  and  $v_{\mu}$  cross section
- Hope to have a working example for T2K in 2019
  - Concept useful for any accelerator-based neutrino experiment







### **Backup Slides**

# $v_{\mu}$ Disappearance Analysis

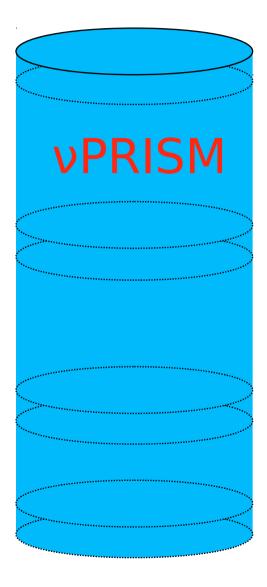


- Event selection
- vPRISM predicted SK spectrum
- Systematic uncertainties
- Statistical uncertainties
- Oscillation fit
- Effect of multi-nucleon events

### vPRISM Design



• Baseline design used in the oscillation studies



- 3m radius inner detector
- 52.5m tall spanning 1-4 degrees off axis
- 1km from neutrino target
- vPRISM-lite:
  - Instrument 14m movable cylinder
  - Take data at different off-axis angles over run
  - Studies assumes 4.5 x 10<sup>20</sup> POT in each off-axis slice of vPRISM



20

18

16

14

12

10

8

6

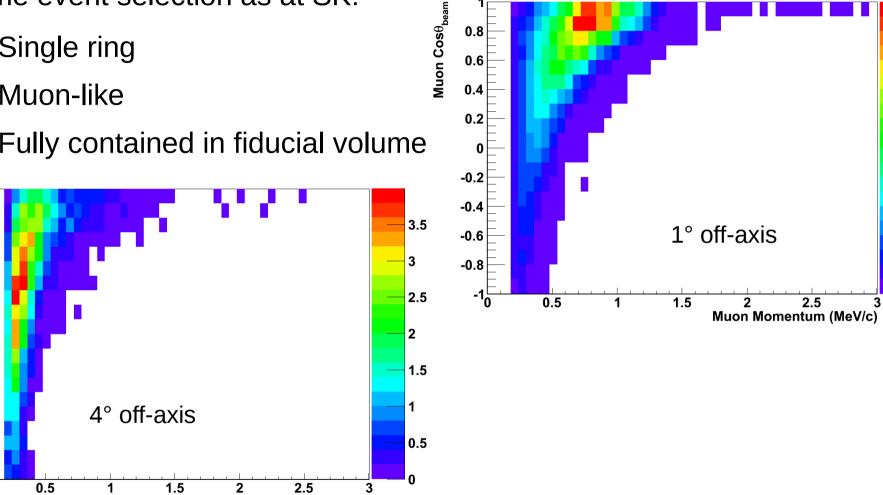
2

0

### **Event Selection**

0.8

- Same event selection as at SK:
  - Single ring
  - Muon-like
  - Fully contained in fiducial volume ۲



Record the off-axis angle of the interaction, using the reconstructed • vertex position

Muon Momentum (MeV/c)

Muon Cosθ<sub>beam</sub>

0.8

0.6

0.4

0.2

0

-0.2

-0.4

-0.6

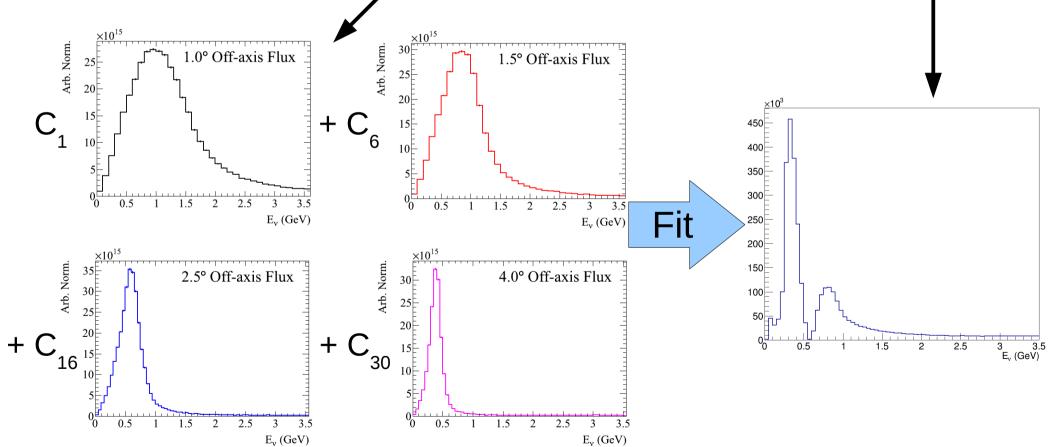
-0.8

-1<mark></mark>₀



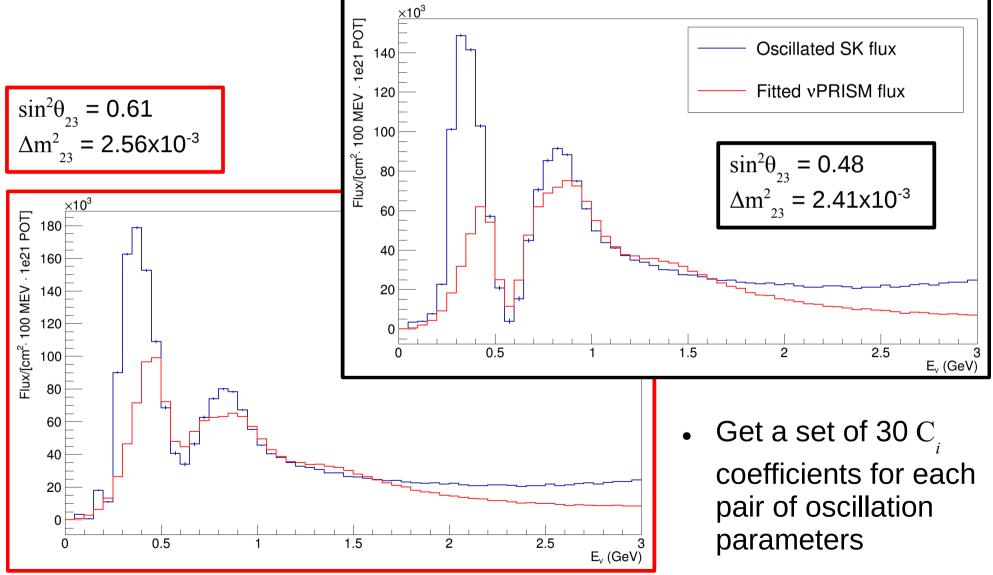
## Building the oscillated flux

- All based on simulated neutrino flux at SK and vPRISM
- Slice vPRISM into 30 slices of 0.1 degree assign each a weight
- MINUIT  $\chi^2$  fit between sum of weighted vPRISM slices and oscillated SK flux



# Building the oscillated flux

Perform fit for all combinations of oscillation parameters used in the oscillation fit

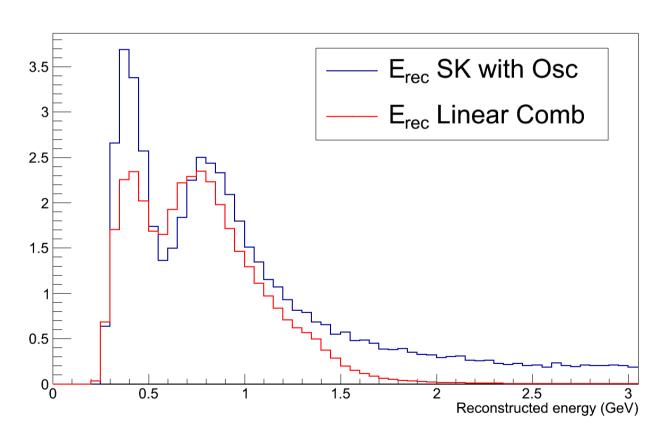


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## SK prediction



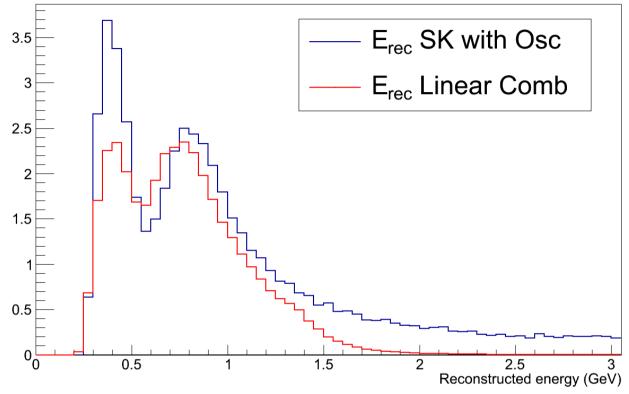
- Apply these weights to the selected events in each off-axis slice of  $\nu\text{PRISM}$
- Now looking at reconstructed neutrino energy events smeared into oscillation dip by nuclear effects and energy resolution



- To vPRISM data:
  - Background subtraction
  - Efficiency correction
  - Addition of selected SK background
- Introduce some model dependence

## Additive correction





- Final step additive correction
- Subtract selected SK spectrum from vPRISM prediction
- Add this difference to the vPRISM prediction
- If our MC exactly reproduces nature, vPRISM prediction will exactly match selected SK spectrum

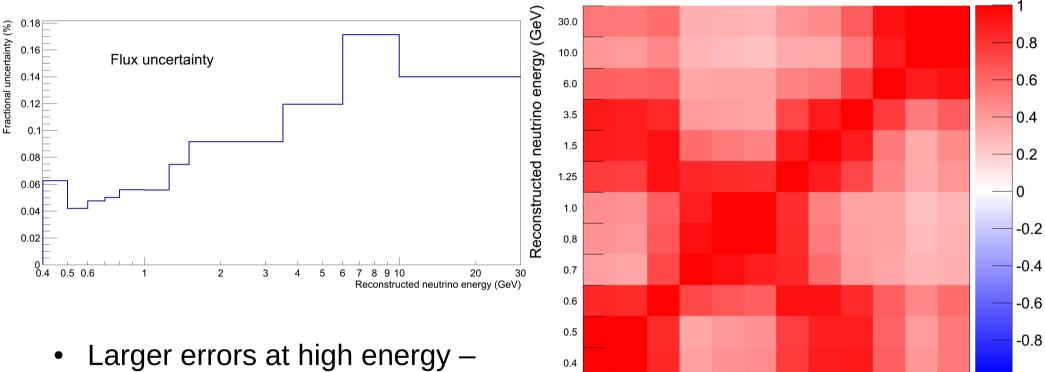


## Systematic uncertainties

- Every correction made to the vPRISM prediction is calculated from our nominal MC all are constant corrections
- To calculate systematic uncertainties:
  - Apply a variation to the vPRISM and SK MC
  - Changes number of selected events at both detectors
  - Apply corrections (from the unvaried, nominal MC)
  - Calculate change in the vPRISM prediction
  - Use this to calculate fractional covariance matrix for vPRISM prediction
- This analysis takes flux and cross section uncertainties into account
  - Conservative detector systematics coming soon!

## Flux uncertainty

- Flux uncertainties calculated in same ways as for T2K, evaluated at 1km
- Fractional error on left, correlation matrix on right



- Larger errors at high energy no vPRISM events
- Error at oscillation dip around 4-5%

0.4

0.5

0.6

0.7

0.8

1.25

1.0

3.5

Reconstructed neutrino energy (GeV)

1.5

6.0

10.0

30.0

UMF

## Flux uncertainty

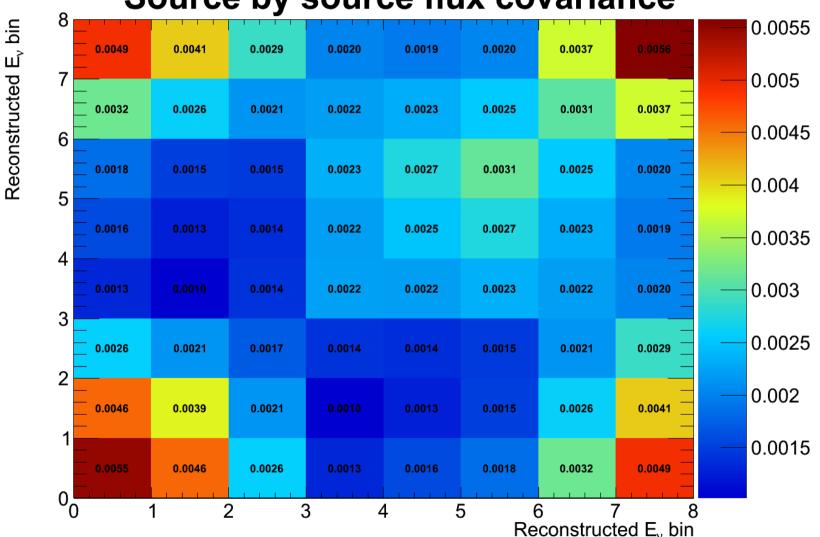


- Flux uncertainties come from 26 sources
  - Proton beam alignment
  - Hadron production
  - Etc.
  - Expect to be independent of one another
- Can calculate a flux covariance matrix in two ways:
  - From each source separately, then combine in quadrature
  - Apply variation from each source at the same time and calculate a covariance for the entire flux uncertainty in one step
- These should give the same answer



#### Separate sources

 Oscillation analysis performed using 12 uneven bins in reconstructed neutrino energy – the 8 shown cover 0 – 3 GeV

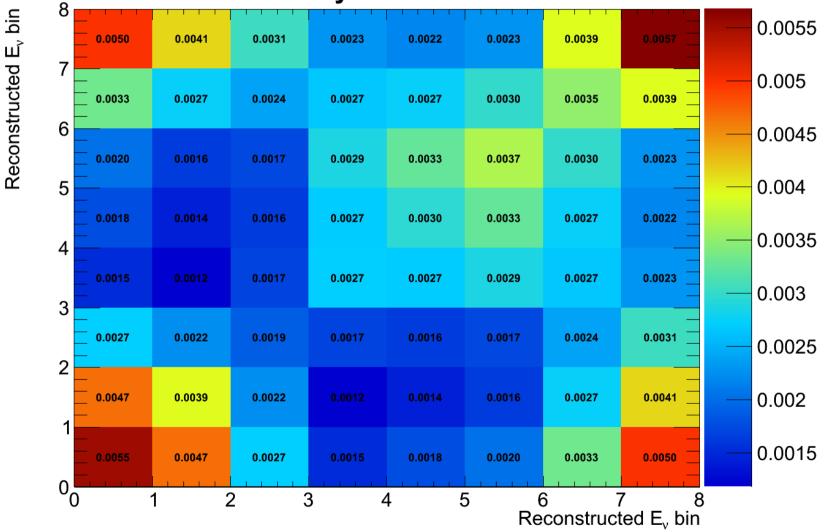


#### Source by source flux covariance



## Simultaneous variation

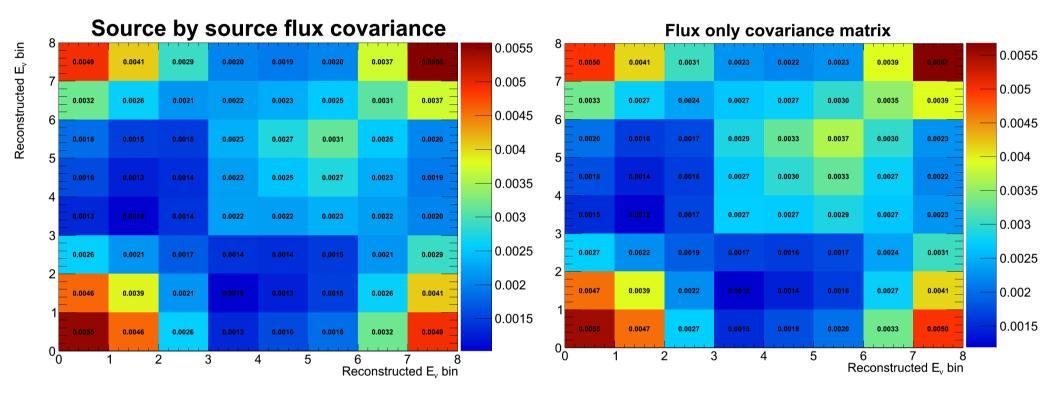
- Larger errors at high and low energy no vPRISM events
- Error at oscillation dip (bin 3) around 5%



#### Flux only covariance matrix

## Comparing flux uncertainty

• Source by source matrix on left, simultaneous matrix on right

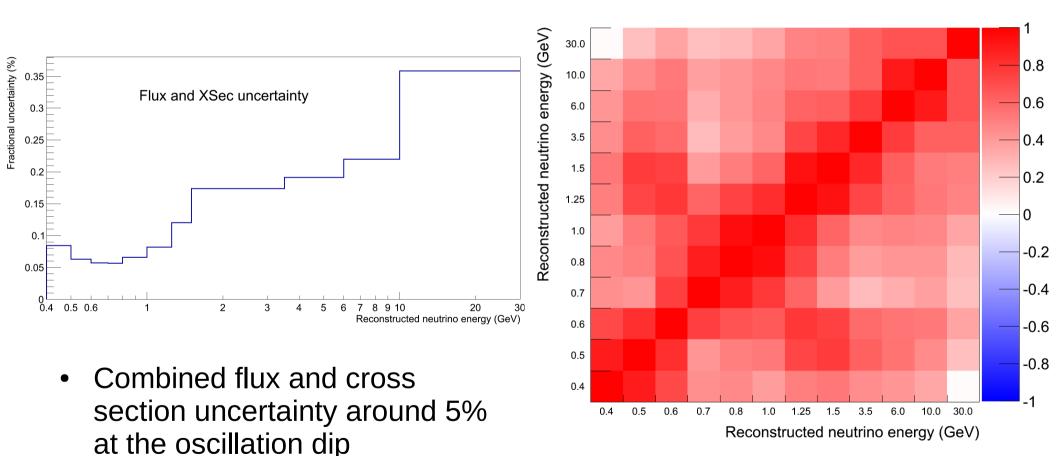


- Very good agreement between the two methods
- Confident flux uncertainties are being applied correctly

**TRIUMF** 

# Flux and XSec uncertainty

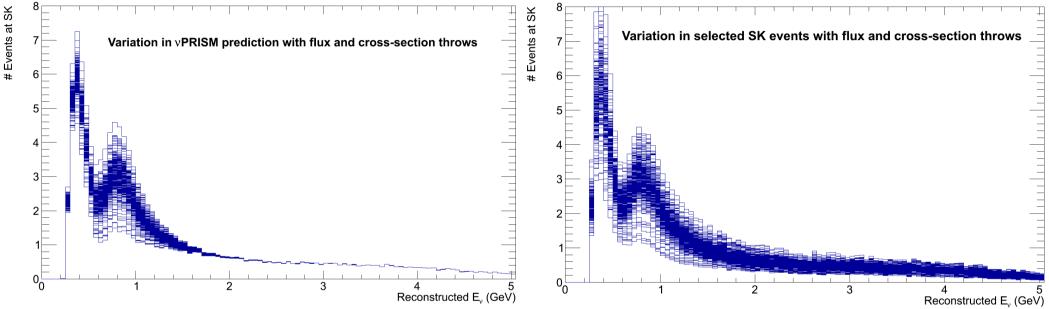
- Xsec uncertainties should largely cancel at vPRISM amount of cancellation depends on how well flux combination matches SK flux
- Need to throw flux and cross section uncertainties together



## Systematic throws



Look at fake data throws of both flux and cross section uncertainties

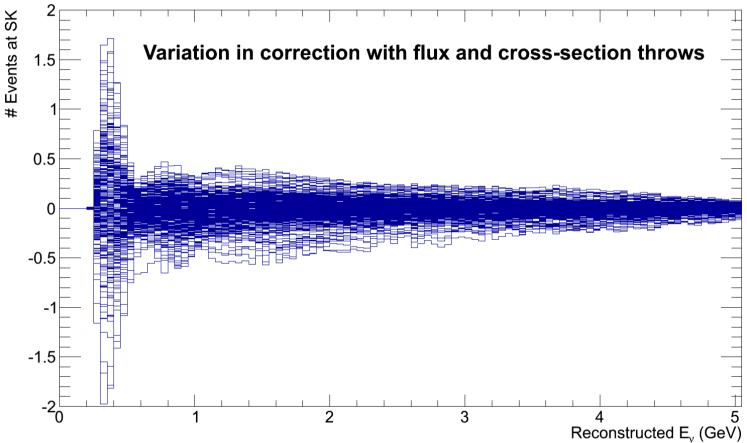


- Plots show all 300 throws of the vPRISM prediction (left) and selected SK events (right)
- vPRISM very few events at low or high energy, little variation
- In oscillation region variations similar at SK and vPRISM
- Spectra are ~Gaussian distributed about the central value

# Systematic throws



- Plot difference between selected SK events and vPRISM prediction for each throw

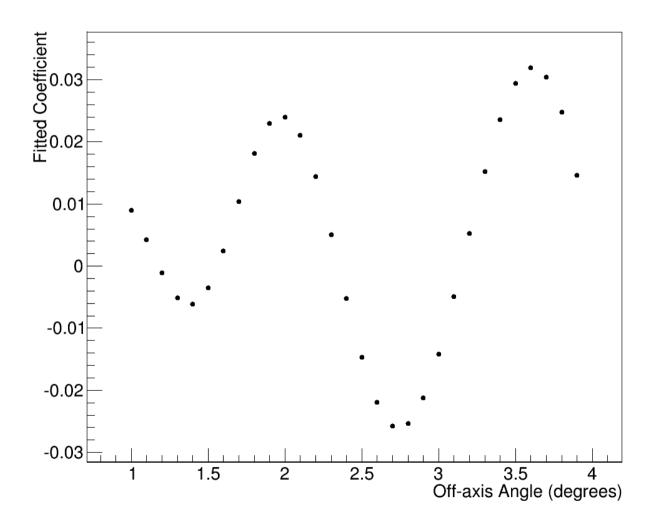


- Most of spectrum shows less than 0.5 event difference between SK and  $\nu \text{PRISM}$  prediction
- Systematic uncertainties are cancelling between the two detectors

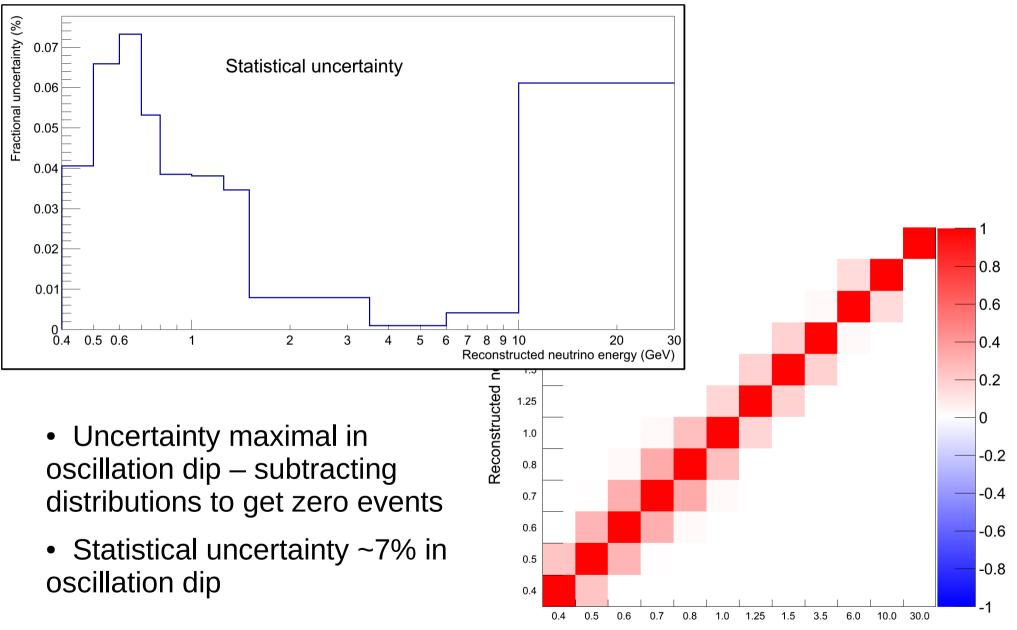


### Statistical uncertainties

- Potential to be large due to linear combination
- Smooth linear combination variations in neighbouring slices cancel out to large extent



### Statistical uncertainties



Reconstructed neutrino energy (GeV)

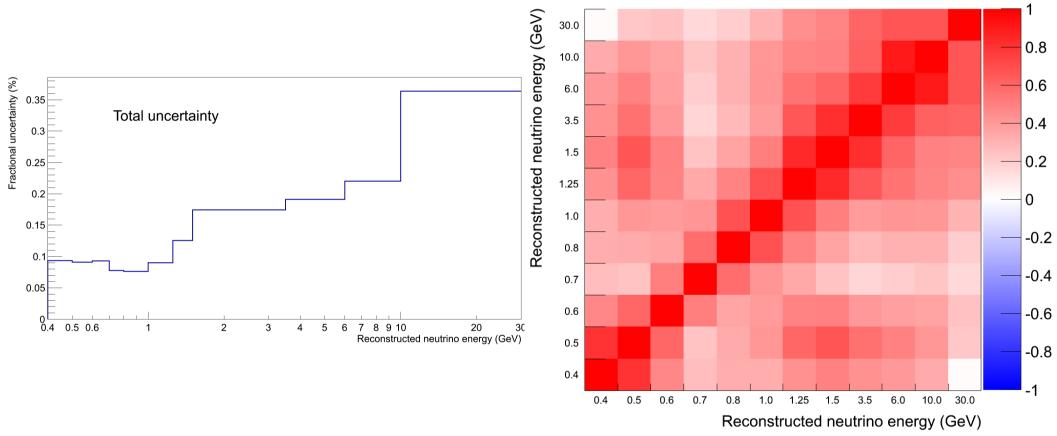
Mark Scott, NuFact 2014, Glasgow

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## Total uncertainty

• Total uncertainty on the predicted event spectrum at SK, including statistical and systematic sources



- Total uncertainty is <10% at oscillation peak
- ~7% statistical, 6% systematic

## Oscillation fit



- Calculate covariance matrix and vPRISM prediction for various points in  $\theta_{_{23}}$  and  $\Delta m^2$  phase space
  - °20.0029 70 0.0028 60 0.0027 50 0.0026 0.0025 40 0.0024 30 0.0023 20 0.0022 0.0021 10 0.002 0 0.65 0.4 0.45 0.5 0.55 0.6 0.35  $sin^2\theta_{23}$

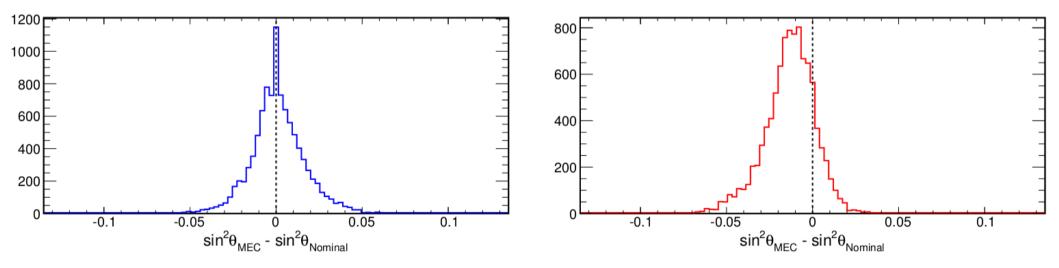
#### -log(L) surface for nominal MC

- Use Simple Fitter to calculate likelihood (L)
- Plot ln(L) for all points in  $\theta_{23}$  and  $\Delta m^2$
- Minimum bin gives best fit oscillation parameters



### Multi-Nucleon effect

- Add meson exchange current (MEC) interactions to the same  $\nu\text{PRISM}$  and SK fake data sets, using Nieves and Martini models
- Re-calculate vPRISM prediction of SK distribution do not change any of the corrections!
- Find the best fit oscillation point for each fake data set compare to best fit point without MEC

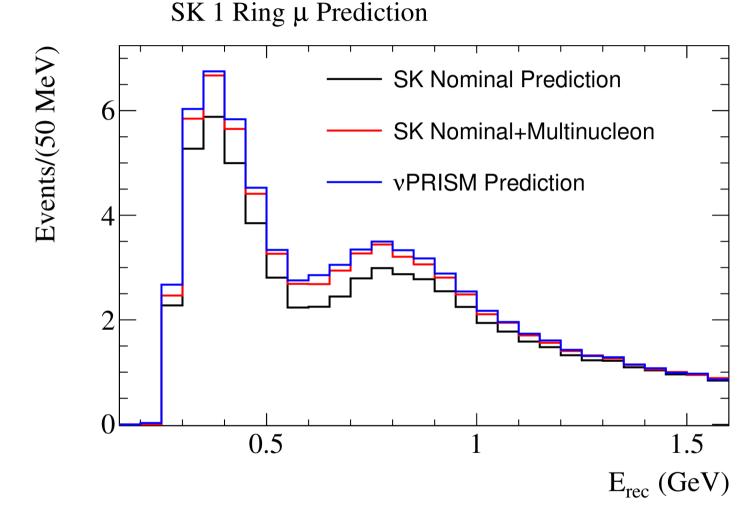


- Plots above show the result of the same analysis performed by T2K
- Using Nieves' MEC prediction on left, Martini mock up on right
- Both show ~3.5% spread, with a bias in the Martini case



## Multi-Nucleon example

• Add multi-nucleon events to the nominal MC to make fake data



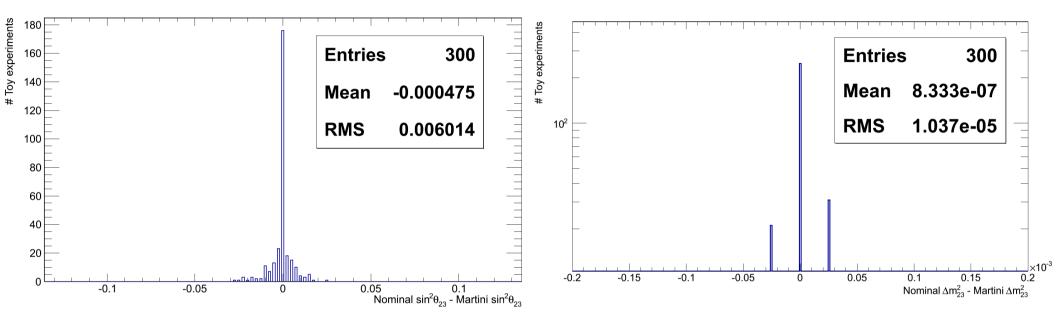
 See vPRISM prediction still reproduces oscillated SK spectrum when multi-nucleon events are present

Mark Scott, NuFact 2014, Glasgow



## Martini MEC result

• Look at effect of adding MEC events to 300 fake data sets

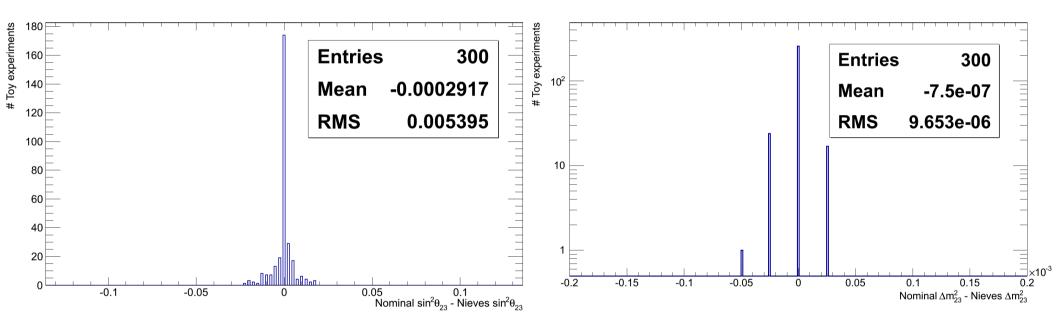


- Much smaller RMS in  $\theta_{_{23}}$  (left) and  $\Delta m^2$  (right) than in T2K analysis
- No bias seen in  $\theta_{23}$  plot
- vPRISM will provide the first data driven constraint on the effect of multi-nucleon events in oscillation measurements



#### Nieves' result

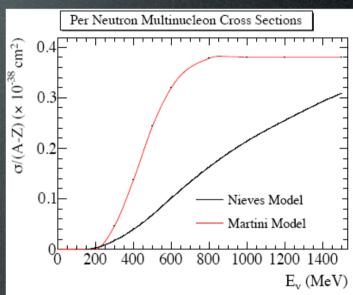
 Look at the difference in best fit oscillation parameters between the nominal MC and the MC with additional Nieves MEC events

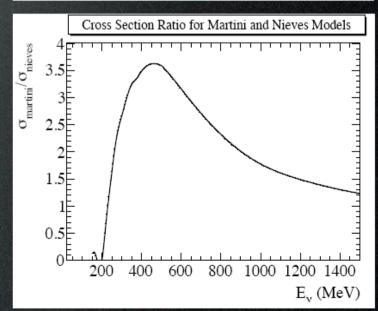


- Much smaller RMS in  $\theta_{_{23}}$  (left) and  $\Delta m^2$  (right) than in T2K analysis
- Large spike at 0 difference in both plots

# How Well are the New Models Understood?

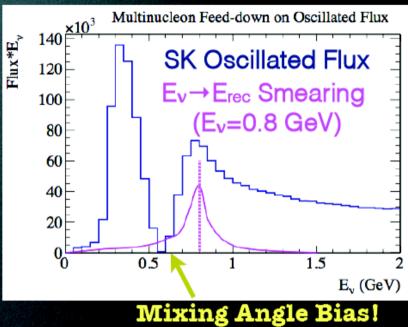
- It is very difficult to answer this question without a direct measurement
- However, the two most commonly used "new" models can be compared
  - J. Nieves, I. Ruiz Simo, and M. J.
    Vicente Vacas, PRC 83:045501 (2011)
  - M. Martini, M. Ericson, G. Chanfray, and J. Marteau, PRC 80:065501 (2009)
- Cross section differs by a factor of 2 to 3 over a large range of neutrino energies
- Which model is correct?
  - Is either model correct?
- Nuclear physics at 1 GeV is difficult

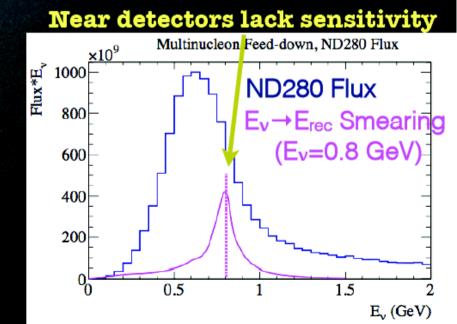




# Isn't This is Why Oscillation Experiments Build Near Detectors?

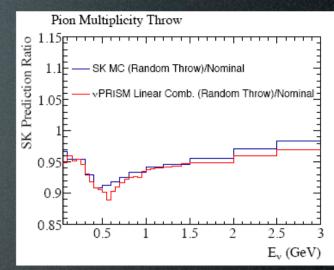
- Shouldn't cross section systematics cancel in a near/far fit?
  - Some errors, like total normalization, will cancel
- However, multi-nucleon effect causes feed-down of events into oscillation dip
  - Cannot disentangle with near detectors
    - Energy spectrum is not oscillated
- More multi-nucleon = smaller dip
  - Multi-nucleon effects are largely degenerate with mixing angle effect!

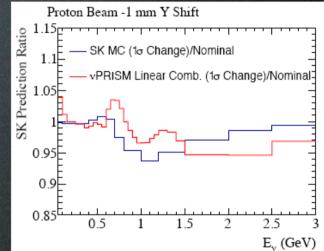


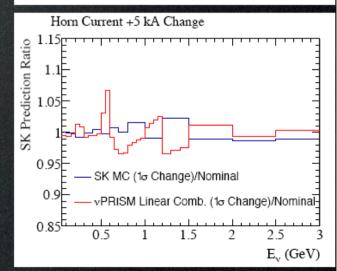


# Beam Errors

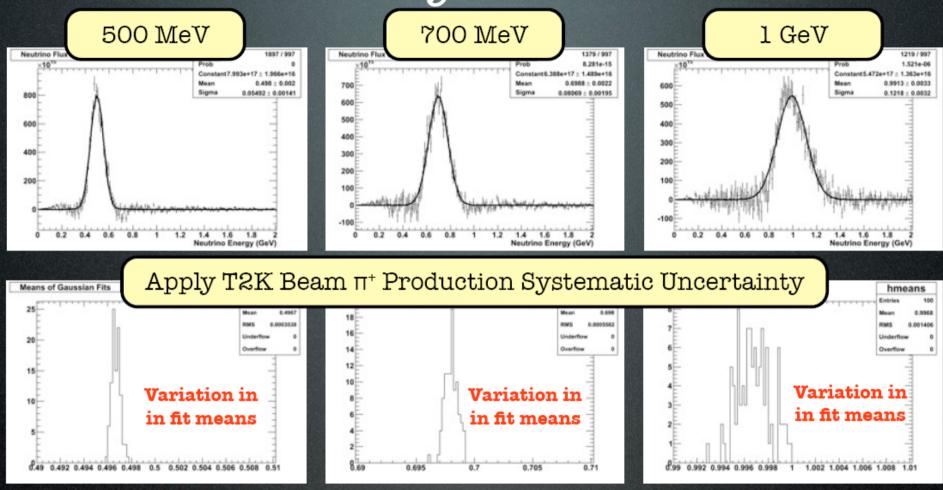
- Haven't we just replaced **unknown cross section errors** with **unknown flux errors**?
  - Yes! But only relative flux errors are important!
  - Cancelation exist between vPRISM and far detector variations
- Normalization uncertainties will cancel in the vPRISM analysis
  - Cancelations persist, even for the vPRISM linear combination
  - Shape errors are most important
- For scale, 10% variation near the dip means
  <sup>~</sup>1% variation in sin<sup>2</sup>2θ<sub>23</sub>
  - Although this region is dominated by feed down
- Full flux variations are reasonable
  - No constraint used (yet) from existing near detector!
    - Uncertainties set by NA61 and T2K beam data





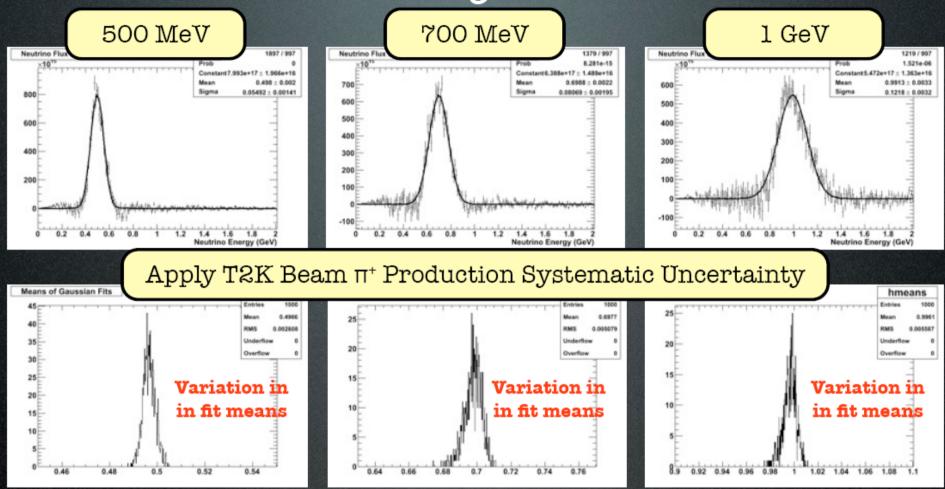


#### **Beam Systematics**



- Apply **T2K**  $\pi^+$  production variations to flux linear combinations
  - This is expected to be the dominant normalization uncertainty for T2HK
- Spread in neutrino energy due to π<sup>+</sup> production uncertainty is O(0.1%)
  - More detailed study needed, but so far looks promising

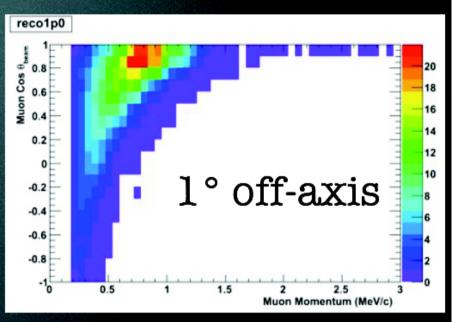
#### **Detector Systematics**



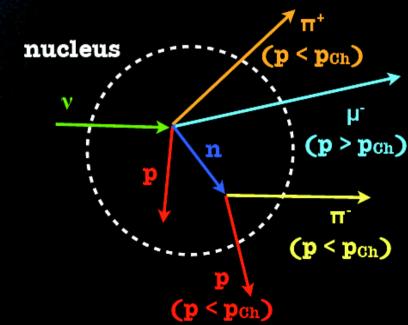
- Efficiency was randomly varied by 5% in each slice
  - The resulting variations in the fit means are still all below 1%
- Continuous variations across the detector can cause problems
  - Need homogeneous detector, and good monitoring & calibration

# Signal Selection/Definition

- Same signal selection as used at Super-K
  - Single, muon-like ring
- Signal events are defined as all true single-ring, muon-like events
  - A muon above Cherenkov threshold
  - All other particles below Cherenkov threshold
- vPRISM can measure single muon response for a given E<sub>v</sub> spectrum
  - Signal includes CCQE, multinucleon, CCπ<sup>+</sup>, etc.
  - No need to make individual measurements of each process and extrapolate to T2K flux

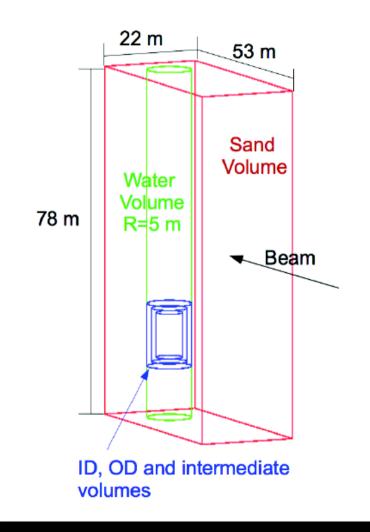


#### **Example Signal Event**



# Event Pileup at 1 km

- Full GEANT4 simulation of water and surrounding sand
  - Using T2K flux and neut cross section model
- 8 beam bunches per spill, separated by 670 ns with a width of 27 ns (FWHM)
- 41% chance of in-bunch OD activity during an ID-contained event
  - Want to avoid vetoing only on OD light (i.e. using scintillator panels)
- 17% of bunches have ID activity from more than 1 interaction
  - 10% of these have no OD activity
  - Need careful reconstruction studies
    - (but multi-ring reconstruction at Super-K works very well)



#### Pileup Rates at 1 km Look Acceptable!

# PMTs

- For the ID, both 8" and 5" PMTs are being considered
  - Perhaps with highquantum-efficiency (HQE) coating
  - Also considering Hyper-Kstyle hybrid photodetectors (HPD)
- Initial Hamamatsu estimate for basic 8" R5912 PMT is much more expensive that assumed for 2km detector
  - US \$4.3M for 3,000 PMTs
- UK/Texas company ETEL/ ADIT has also been consulted
  - Basic 8" PMT is \$1775
  - No HQE or HPD option available

#### Hamamatsu Estimates

Name	Туре	QE%	Quantity	Price/PMT	Total Cost	Delivery Year
5" PMT	R6594-WPassy	25	8000	103,500	828M	
5" PMT HQE		35	5714	123,700	707M	
8" PMT	R5912-WPassy	25	3215	143,000	460M	
8" PMT HQE		35	2296	170,500	391M	
8" HPD HQE	R12112-WPmodule	35	2296	264,000	606M	2014
		35	2296	236,500	543M	2015
		35	2296	209,000	480M	2016
20" PMT HQE	R12860-WPassy	30	508	604,500	307M	2014
		30	508	572,000	291M	2015
		30	508	539,500	274M	2016
20" HPD HQE	R12850-WPmodule	30	508	715,000	363M	2014
		30	508	617,500	314M	2015
		30	508	520,000	264M	2016
20" HPD HQE	R12850-WPmodule	30	140	770,000	108M	2014
		30	140	665,000	93M	2015
		30	140	560,000	78M	2016
20" PMT	R12860-WPassy	30	140	651,000	91M	2014
		30	140	616,000	86M	2015
		30	140	581,000	81M	2016