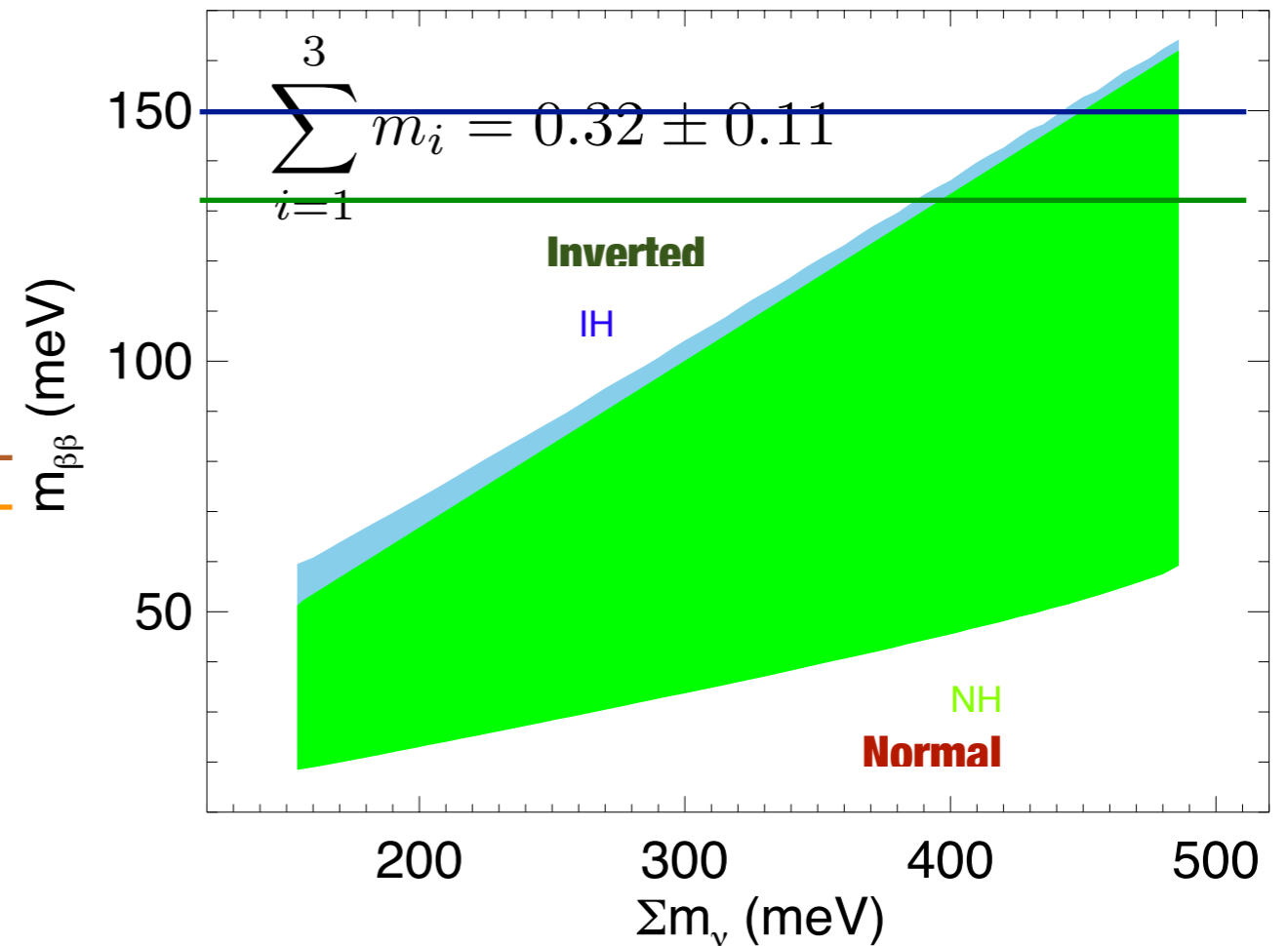
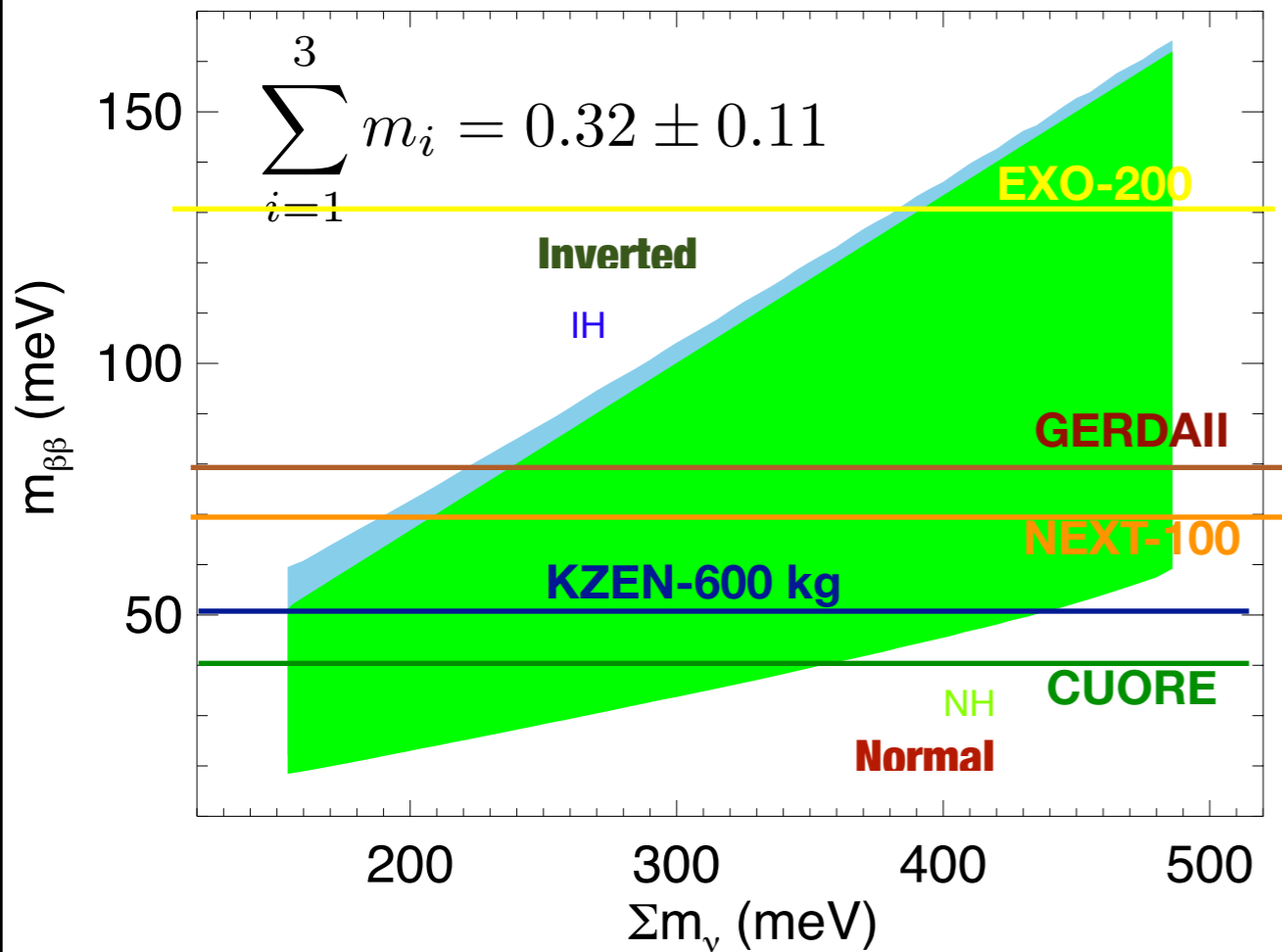


Measuring the neutrino mass

J.J. Gómez Cadenas
IFIC (CSIC & UV)

St. Andrews, INSS, 2014
Lecture 5

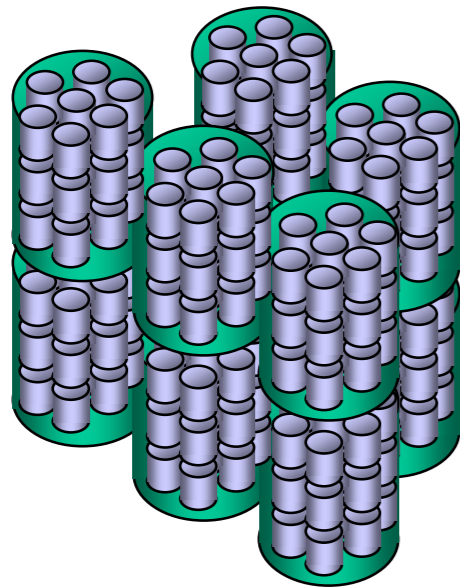
Majorana landscape circa 2020



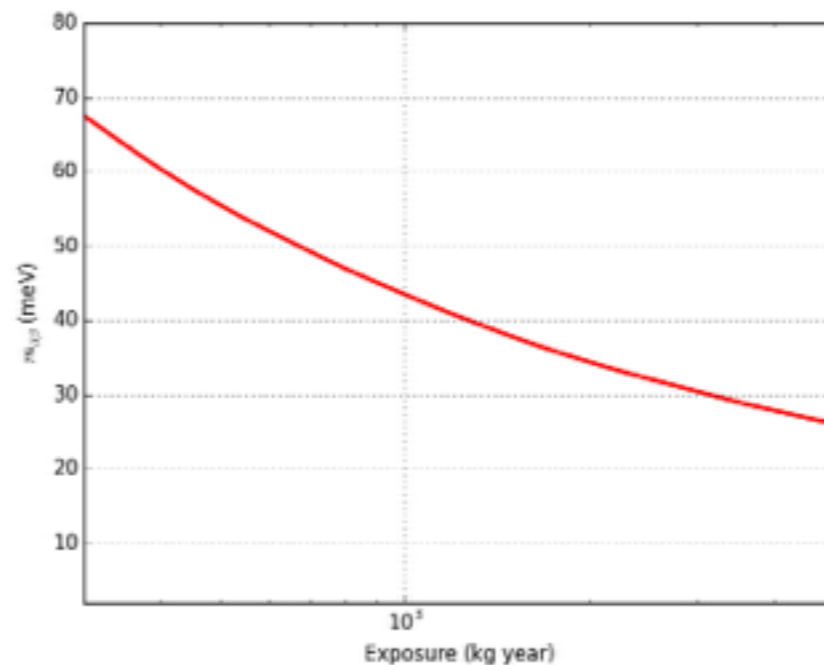
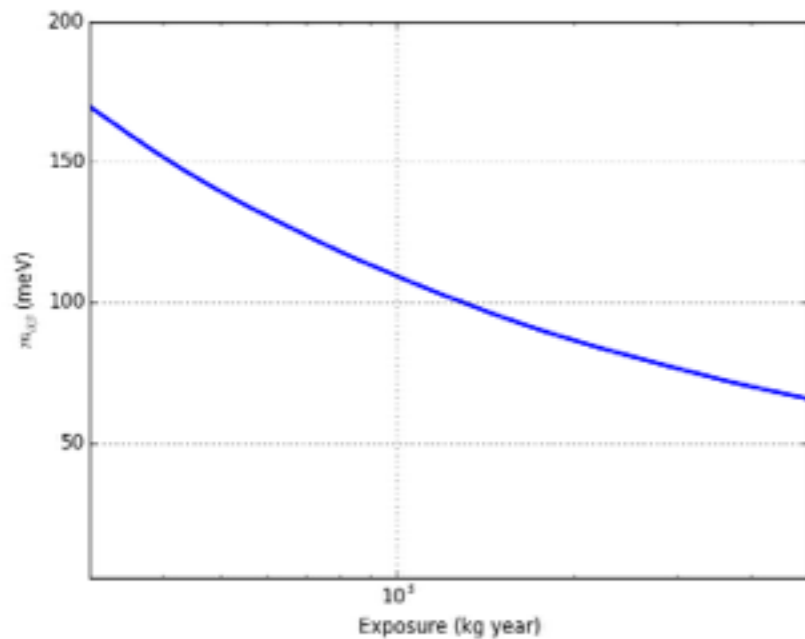
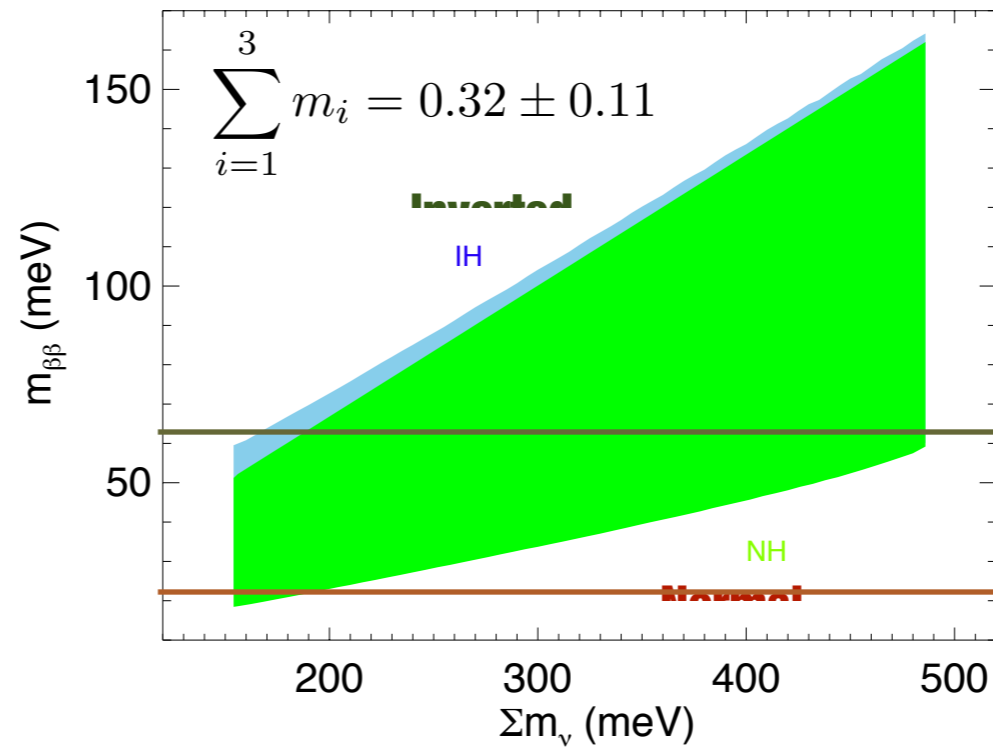
- **Larger NME**
- **Smaller NME**
- **How do we cover the cosmo-region even for small NME?**

SuperGerda

- Mass ~500 kg = 10 x GERDA II
- Total running time = 15 years (!!)
- Effective exposure = 5 ton year



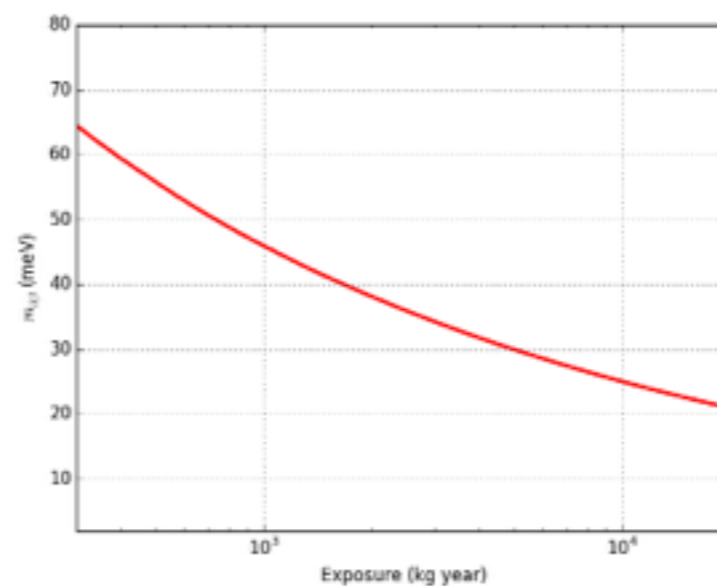
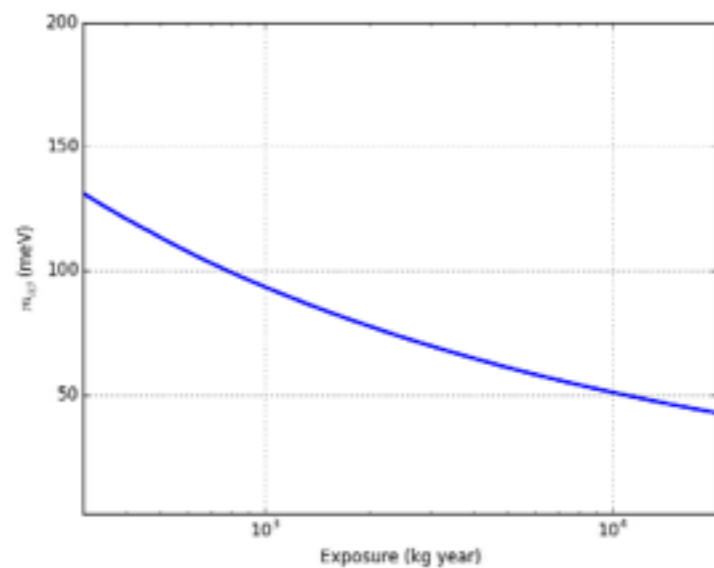
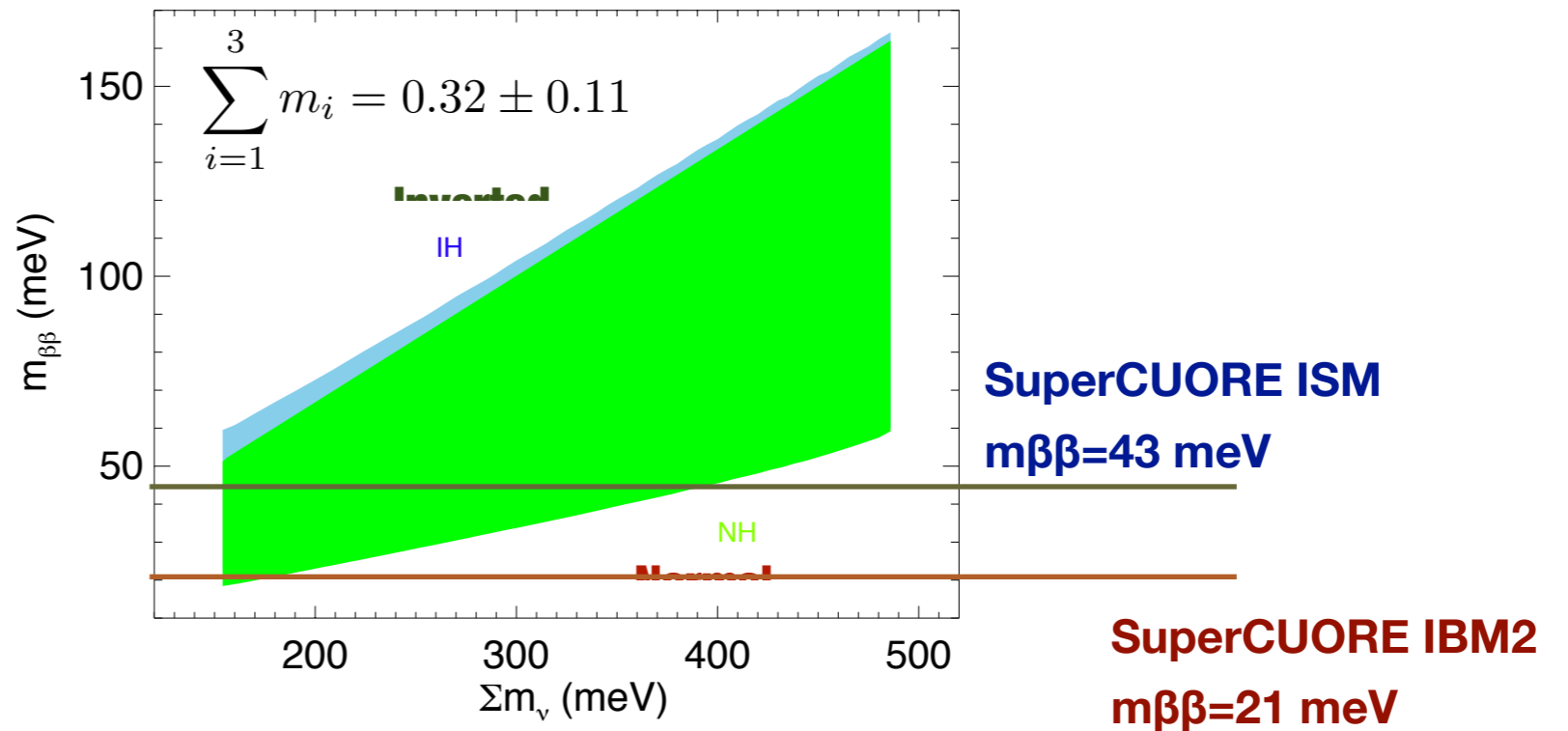
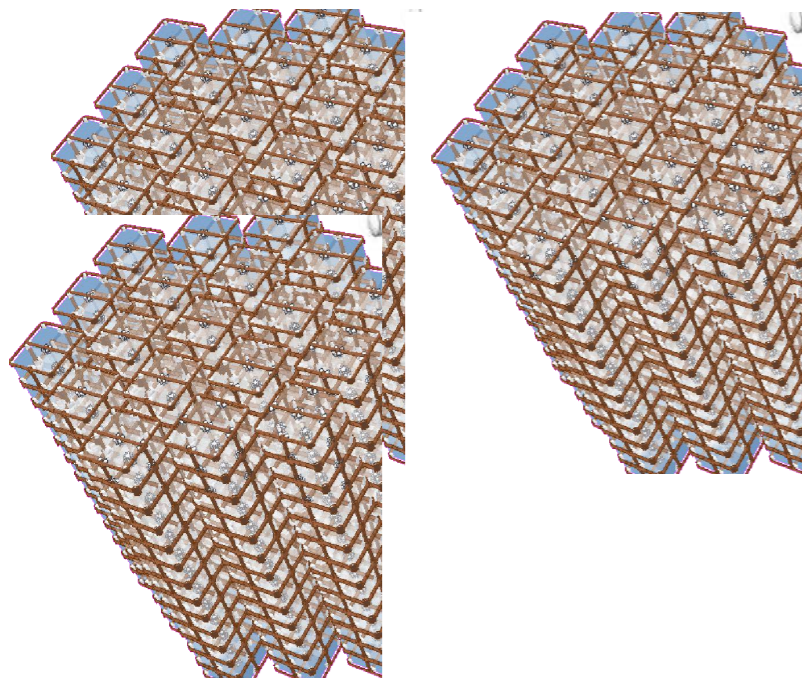
210 2.35 kg crystals



efficiency = 0.62
 $\Delta E = 0.1$ % at 2.6 MeV
 $B = 1. \times 10^{-3}$ ckky

SuperCuore

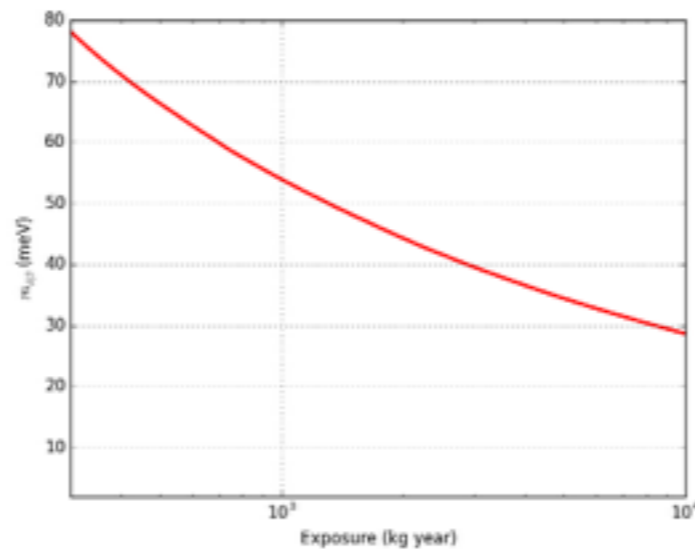
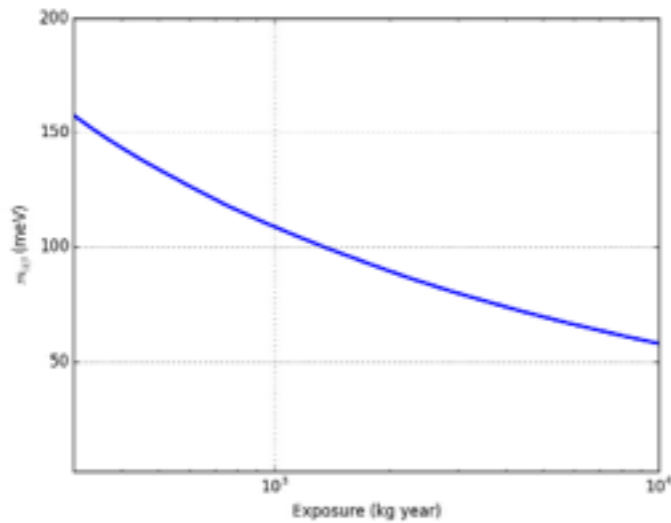
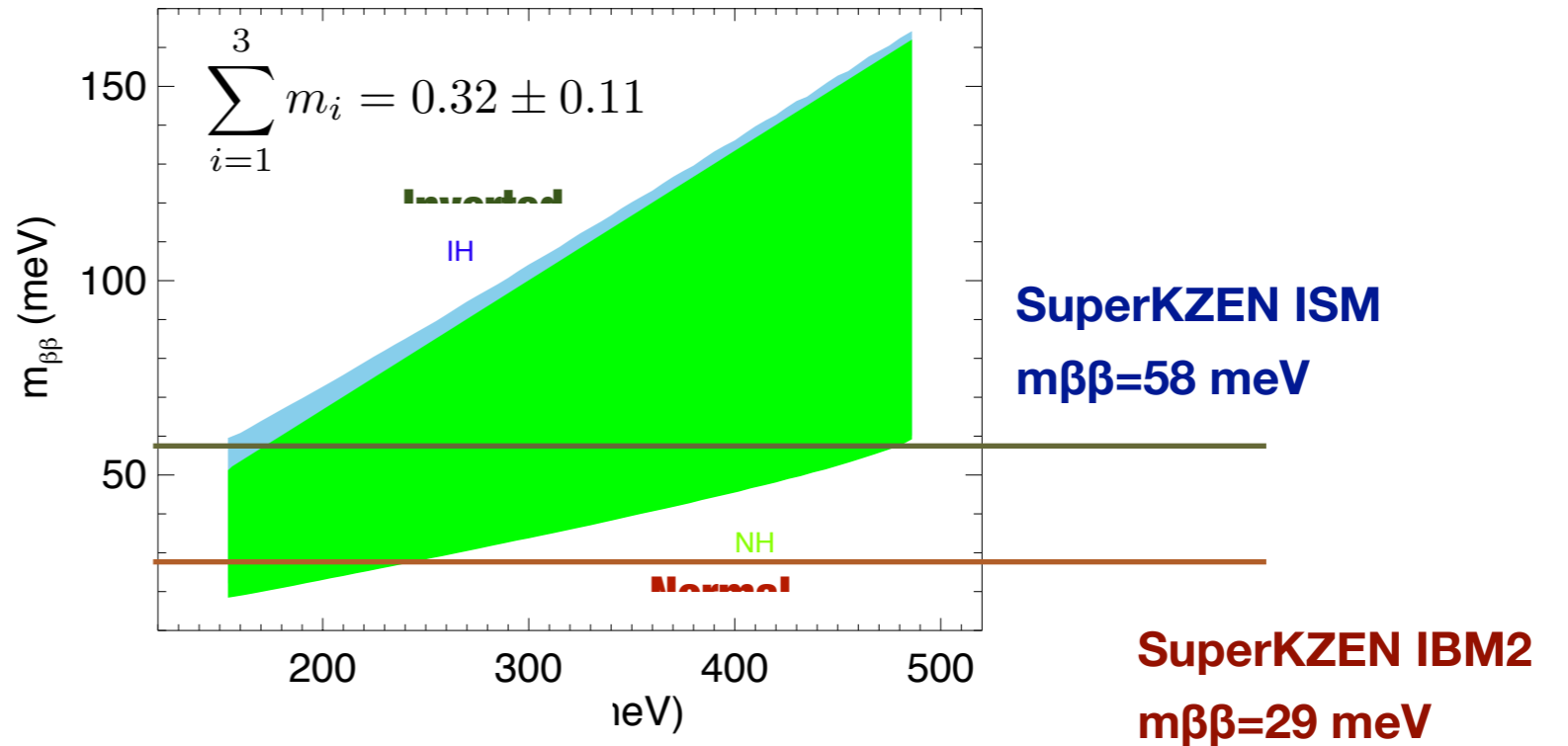
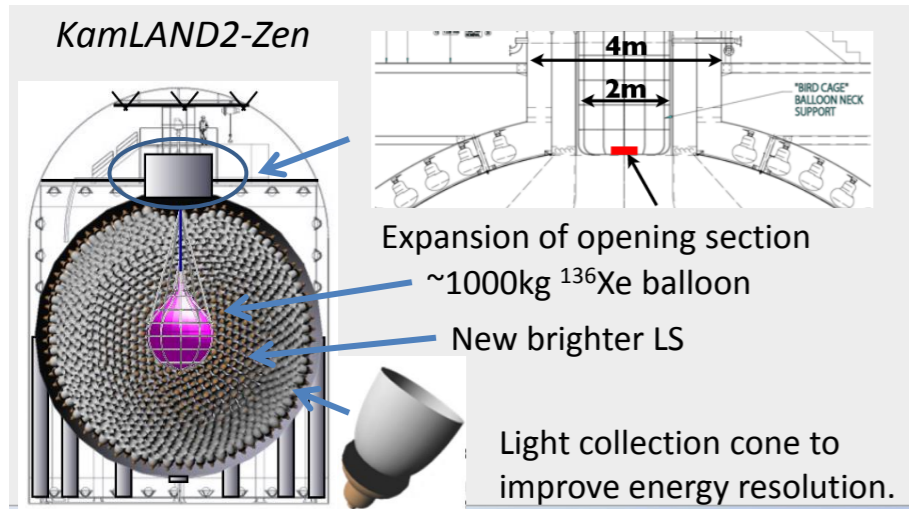
- **Mass (isotope) ~2 ton = 10 x CUORE**
- **3 super-towers + enriched TI**
- **Effective exposure = 20 ton year**



- efficiency = 0.87
- $\Delta E = 0.2 \%$ at 2.6 MeV
- $B = 1. \times 10^{-2} \text{ ckky}$

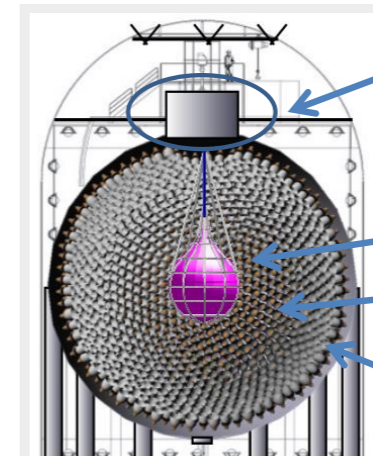
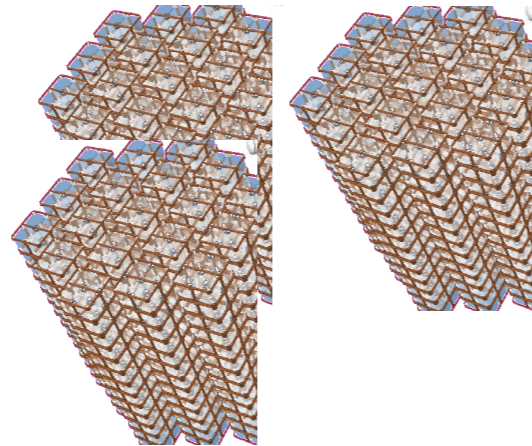
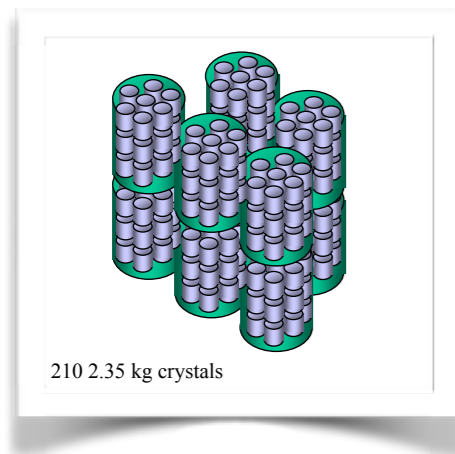
KZen-II

- Mass (isotope) ~1 ton = 10 x KZENI
- Improve resolution to 6% FWHM
- Effective exposure = 10 ton year



- efficiency = 0.55
- $\Delta E = 6\%$ at 2.5 MeV
- $B = 1. \times 10^{-4} \text{ ckky}$

Assessment: Calorimeters



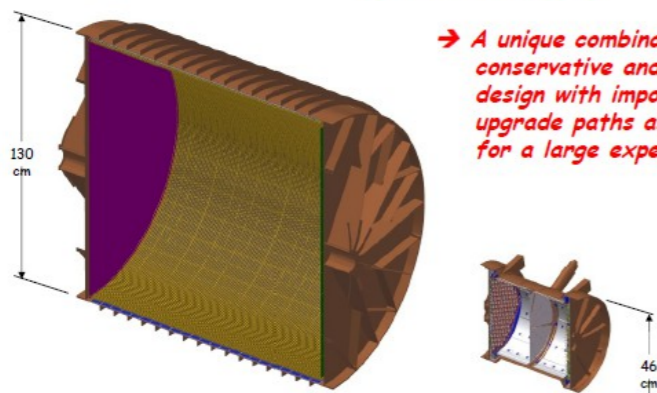
- **GERDA & CUORE: increase the mass by a factor ~10**
 - **Maintain or moderately improve the best expected performance (resolution and B index) of GERDA II and CUORE.**
 - **Challenge: effort & cost to scale up detectors to very large masses.**
 - **Advantage: clear signature if signal is observed**
- **KamLAND-Zen**
 - **Increase the mass up to allowed load in LS (~1 ton)**
 - **Improve BI by a factor 6-7 wrt KamLAND-ZEN I and improve resolution by a factor ~2**
 - **Advantage: Moderate cost (xenon already in hand), feasibility**
 - **Challenge: Counting experiment, no clear signature, normalisation**

NEXO

- Mass ~5 ton = 20 x EXO-200
- Effective mass: ~3 ton Effective exposure = 30 ton year

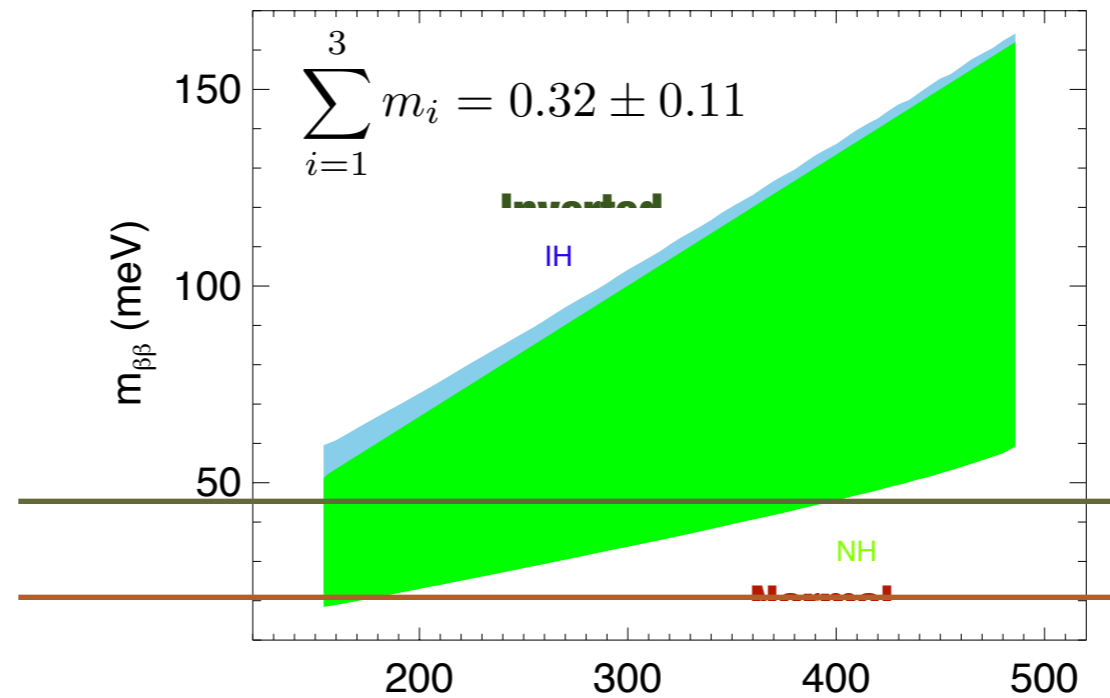
nEXO: a 5 tonnes of enriched Xe

- LXe TPC "as similar to EXO-200 as possible"
- Provide access ports for a possible later upgrade to Ba tagging



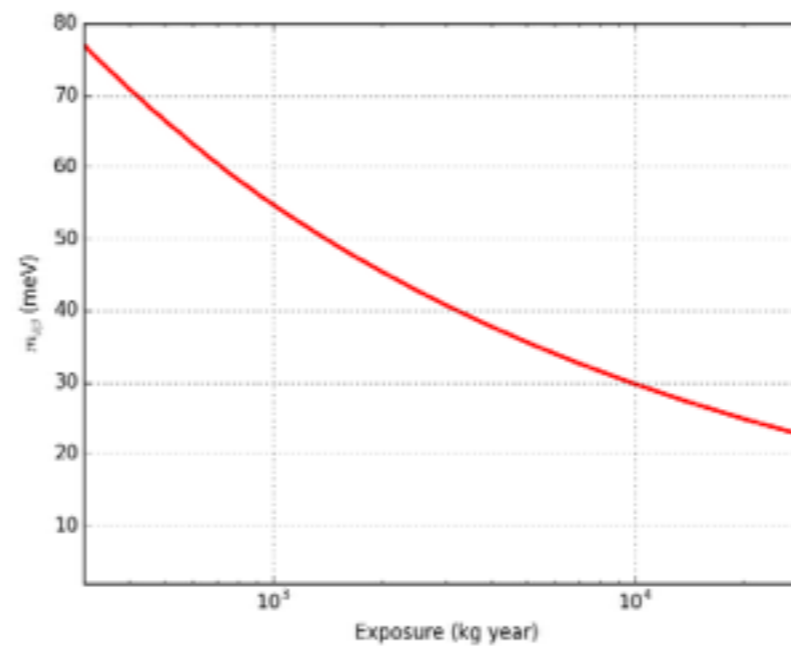
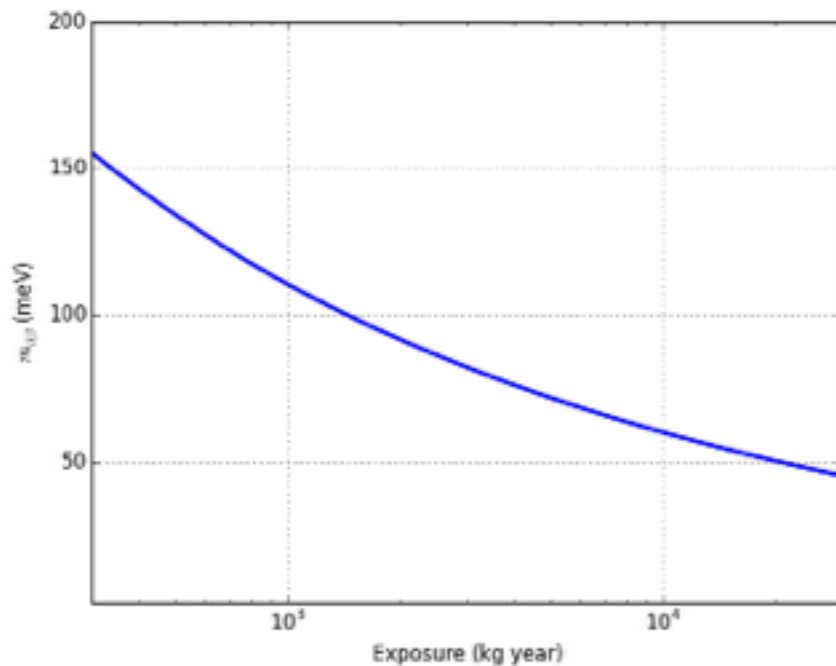
→ A unique combination of conservative and aggressive design with important upgrade paths as desirable for a large experiment

expect better self-shielding, tracking



NEXO ISM
 $m_{\beta\beta}=45$ meV

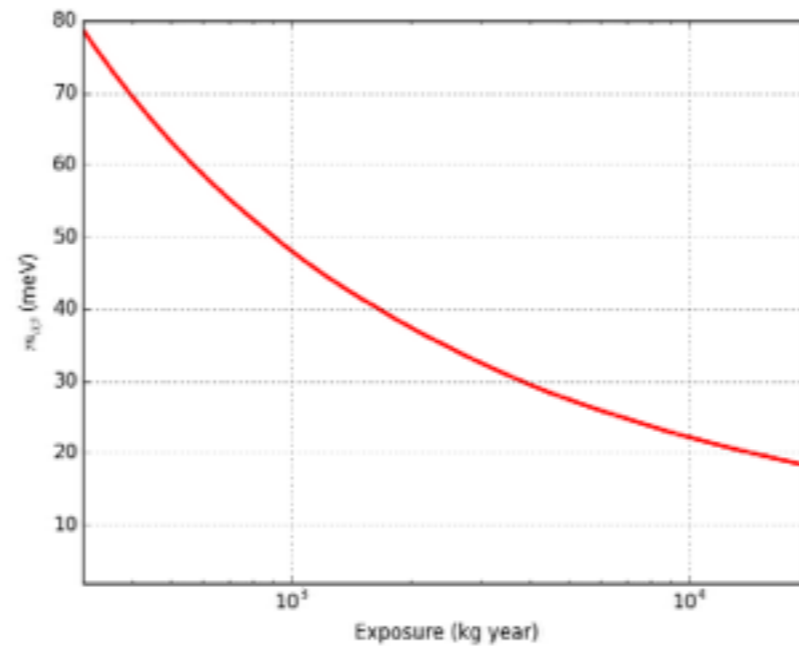
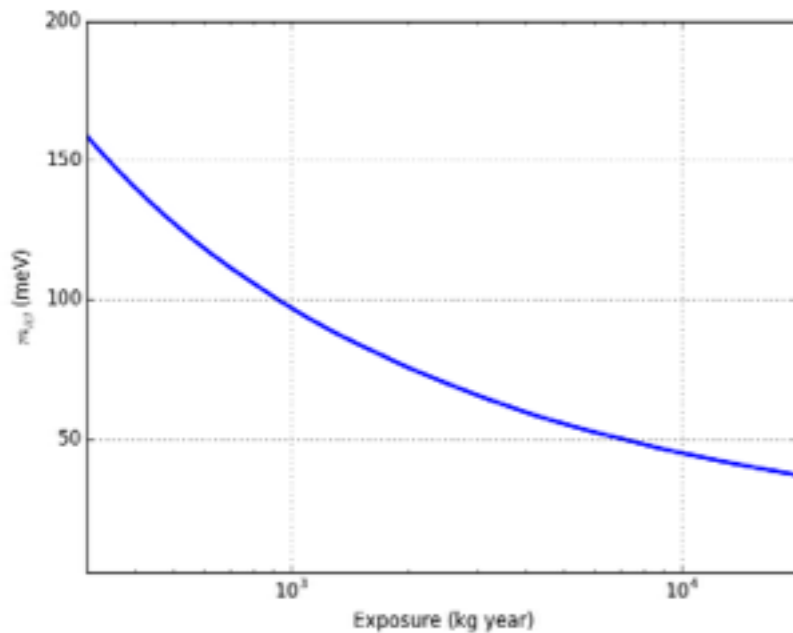
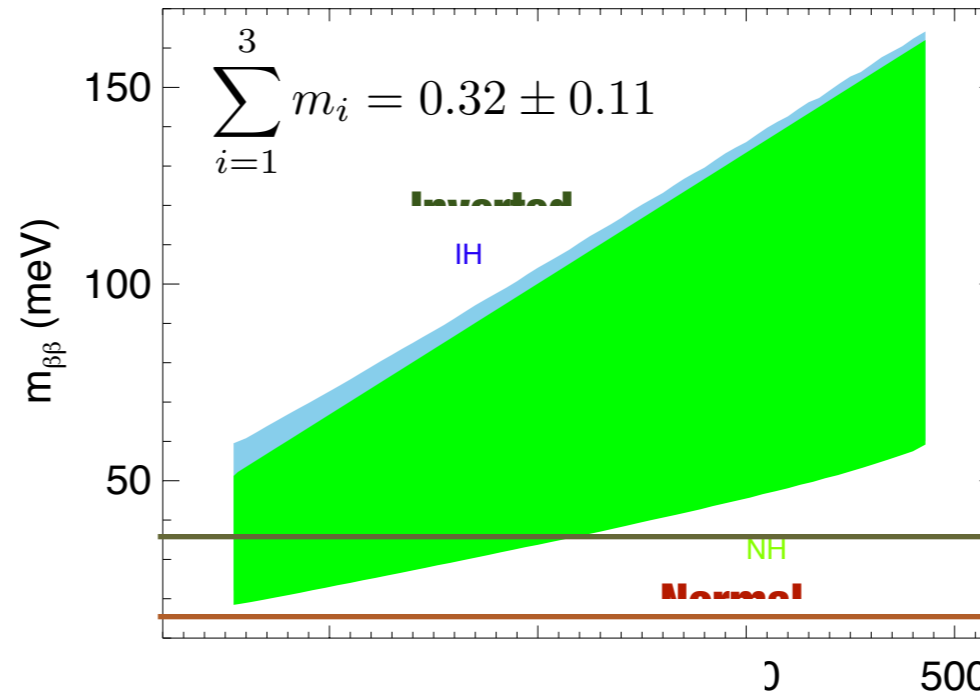
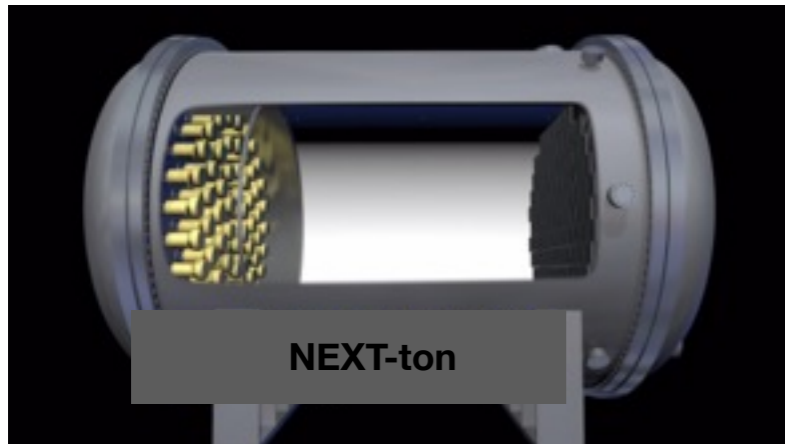
NEXO IBM2
 $m_{\beta\beta}=23$ meV



- efficiency = 0.85
- $\Delta E = 3.6$ % at 2.5 MeV
- $B = 5. \times 10^{-4}$ ckky

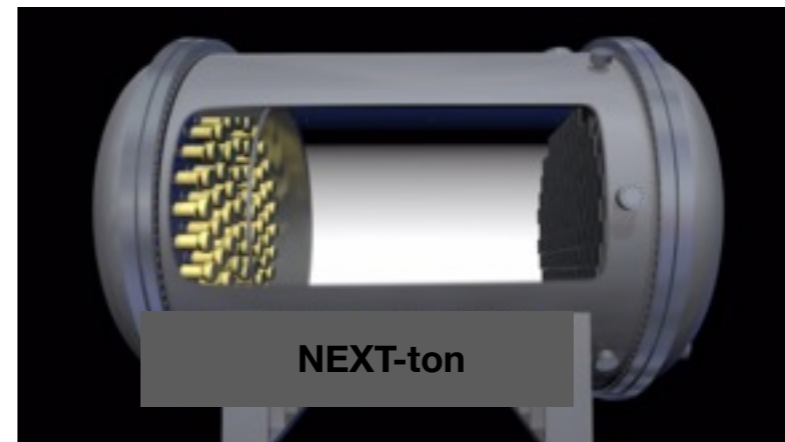
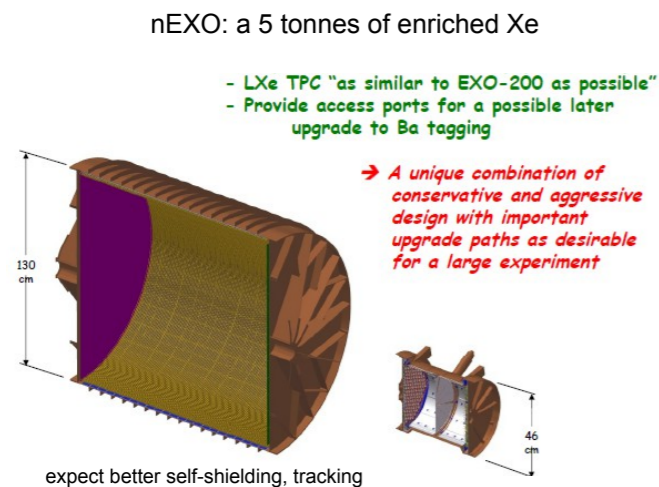
NEXT-ton

- Mass ~2 ton = 20 x NEXT-100
- Increase pressure to 20 bar
- Effective exposure = 20 ton year



- efficiency = 0.3
- $\Delta E = 0.5$ % at 2.5 MeV
- $B = 1. \times 10^{-4}$ ckky

Assessment: TPCs



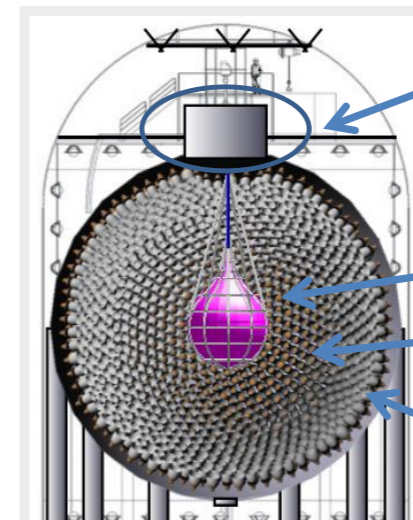
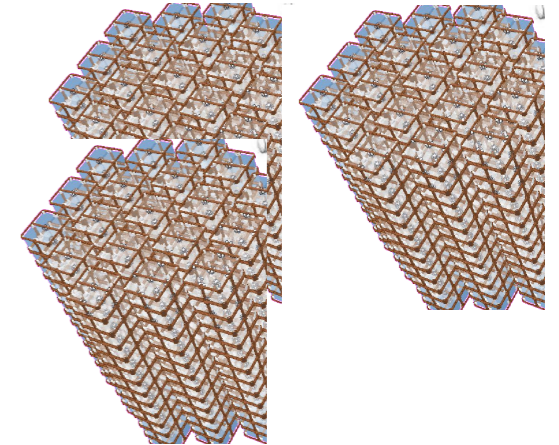
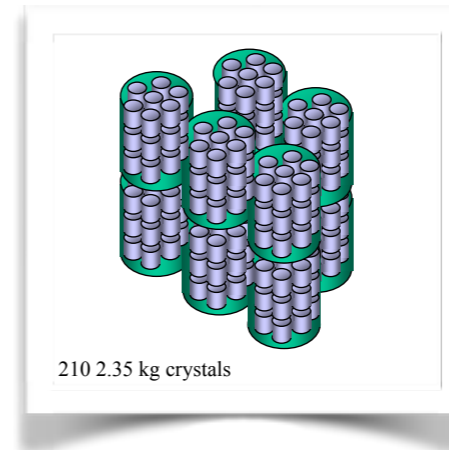
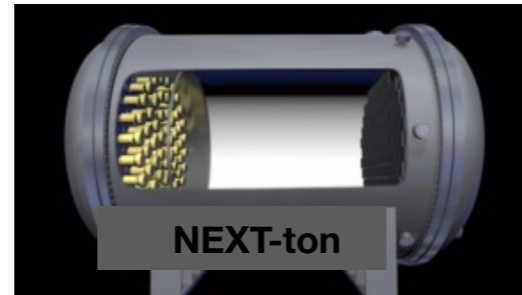
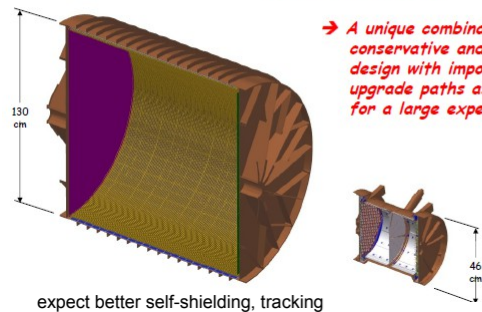
- **NEXO: increase the mass by a factor ~20 (5 ton total)**
 - Maintain or moderately improve energy resolution
 - Improve BI ~10 by self-shielding (give up 2 of 5 ton)
 - **Challenge: light absorption, normalisation. No striking signature**
 - **Advantage: excellent economy of scale**
- **NEXT-TON**
 - Moderately improve energy resolution (wrt projected in NEXT-100)
 - Improve BI by (at least) a factor 5 wrt NEXT (scale, improve topology)
 - **Advantage: clear signature, including topological signal**
 - **Challenge: large TPC.**

Assessment: The next generation

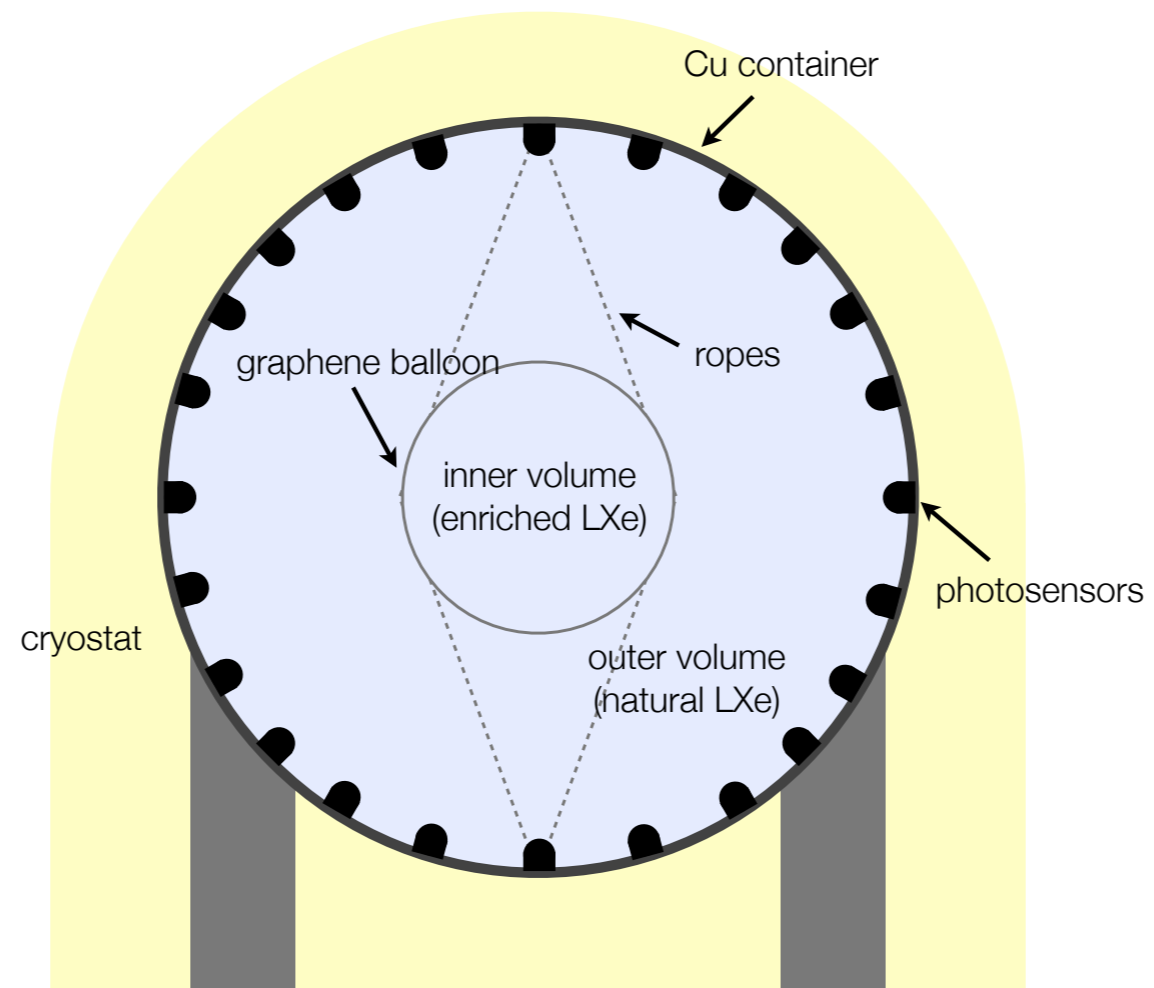
nEXO: a 5 tonnes of enriched Xe

- LXe TPC "as similar to EXO-200 as possible"
- Provide access ports for a possible later upgrade to Ba tagging

→ A unique combination of conservative and aggressive design with important upgrade paths as desirable for a large experiment



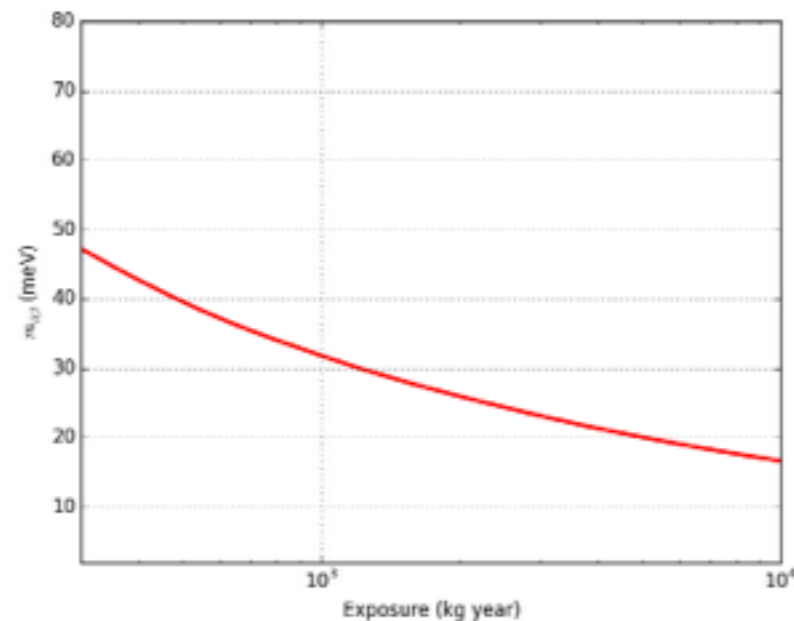
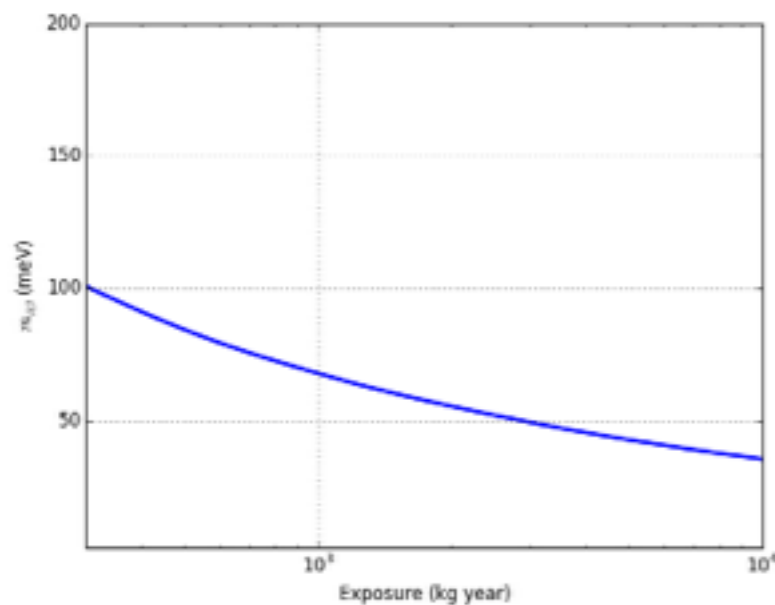
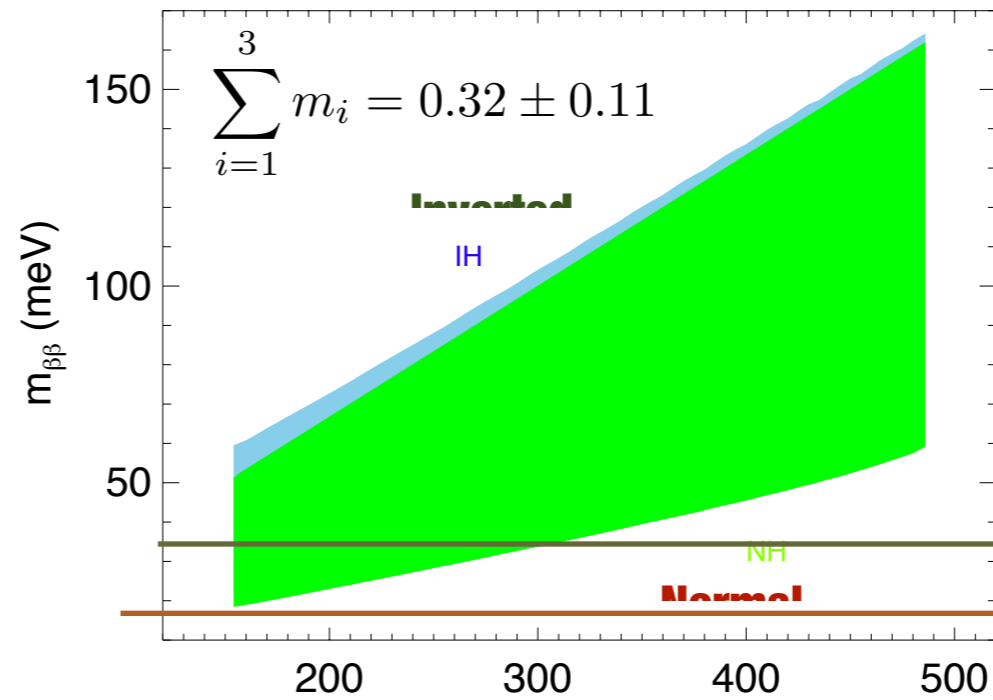
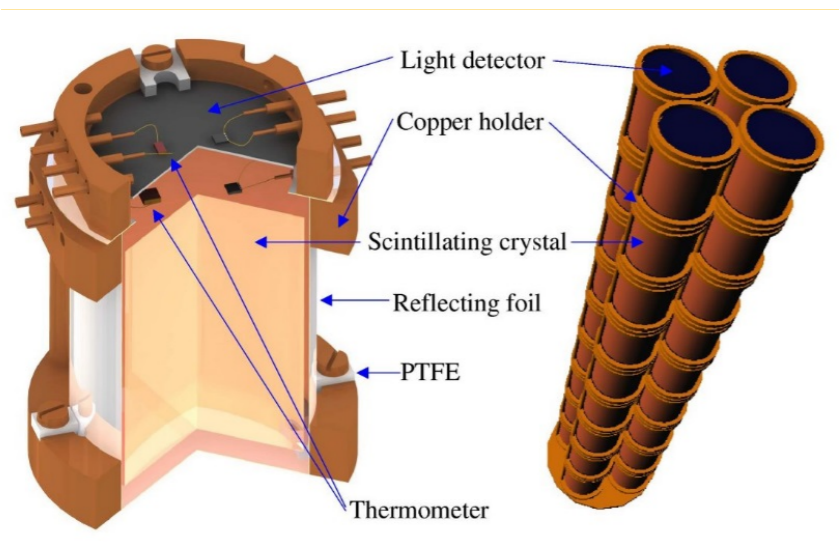
- **Super-Gerda: 26-66 meV**
- **Super-Cuore: 21-43 meV**
- **KamLAND-Zen-II: 29-48 meV**
- **NEXO: 23-45 meV**
- **NEXT-TON: 18-37 meV**
- **Cover the cosmological relevant region (inverse hierarchy) only if NME is high**



Exploring new ideas

GrandLucifer

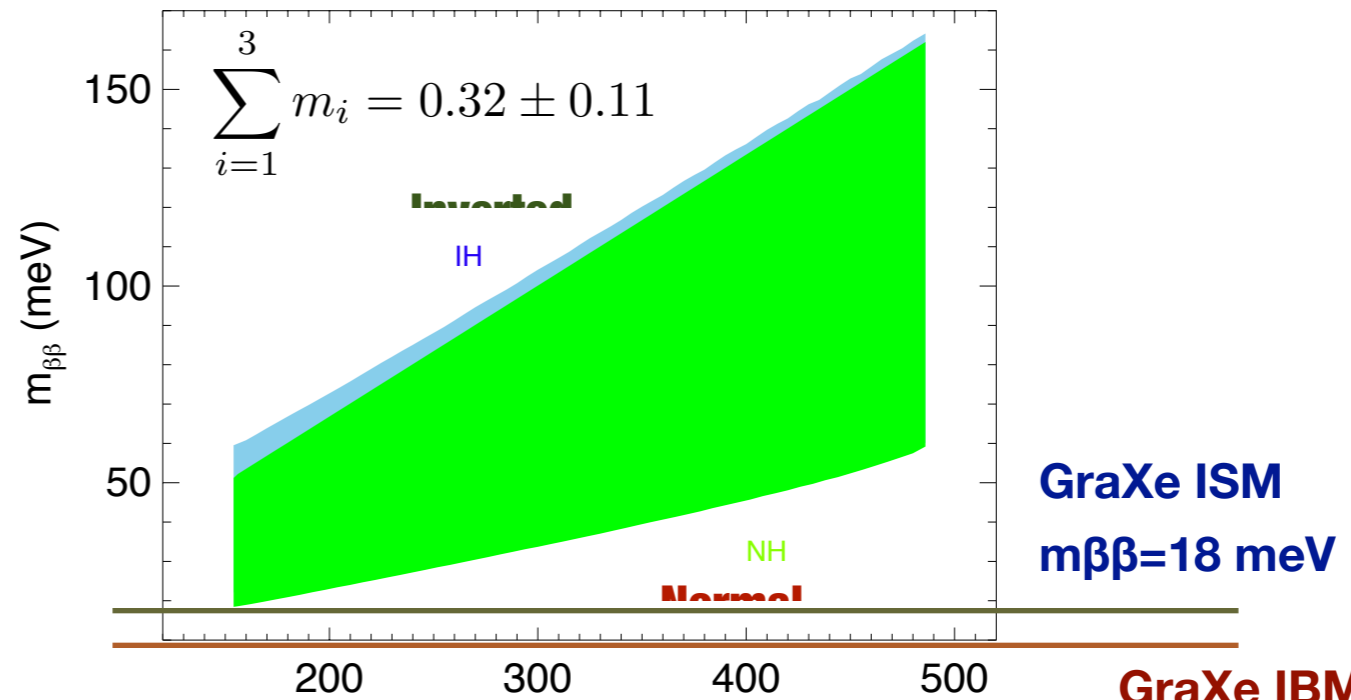
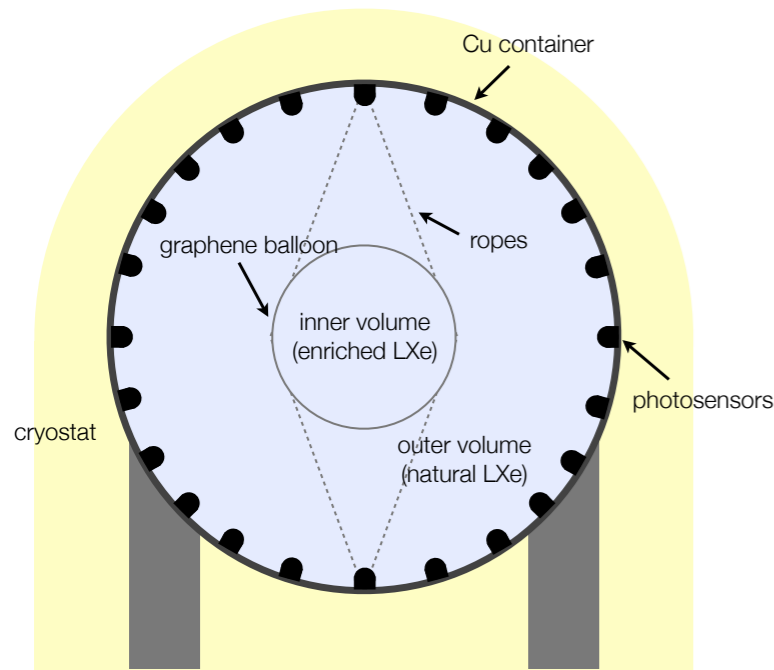
- Mass ~1ton = 50 x Lucifer
- Effective exposure = 10 ton year



- efficiency = 0.65
- $\Delta E = 0.03\%$ at 2.9 MeV
- $B = 1. \times 10^{-3}$ ckky

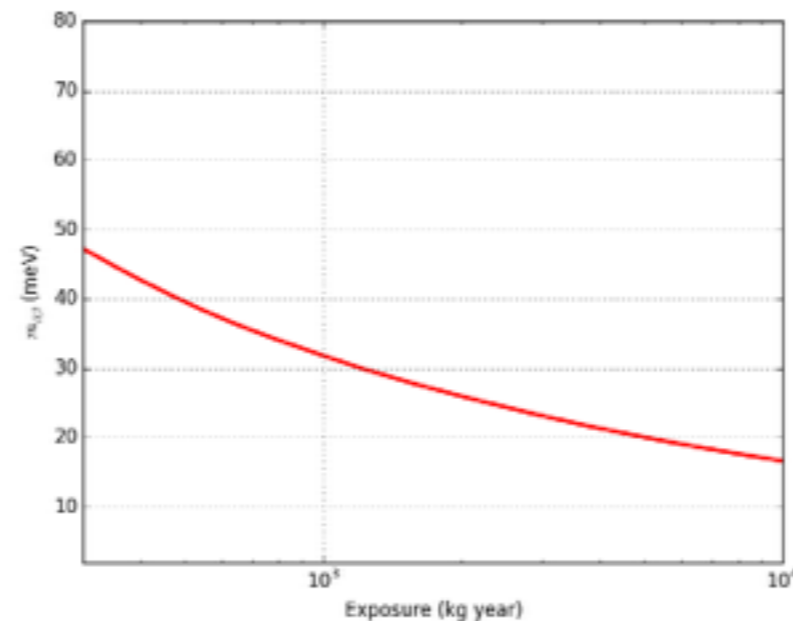
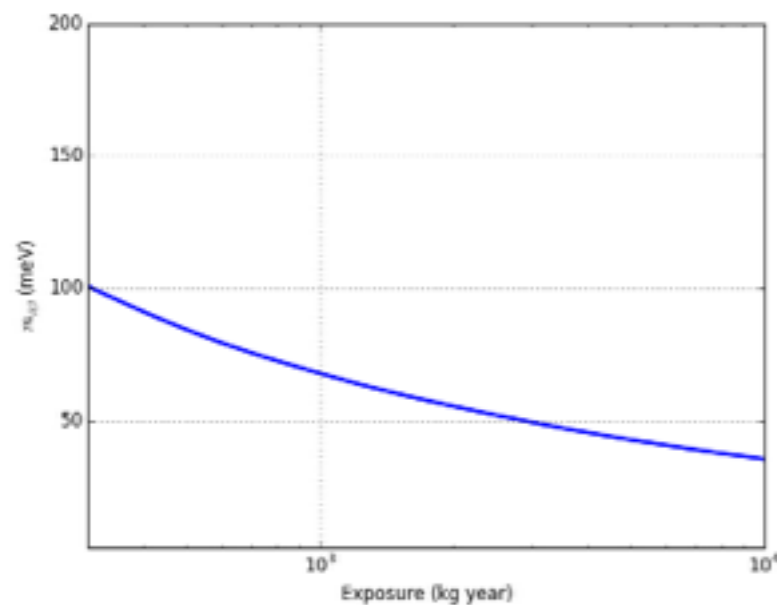
GraXe

- Mass ~3ton enriched in balloon, 20 ton LXe shield
- Effective exposure = 30 ton year



GraXe ISM
 $m_{\beta\beta}=18$ meV

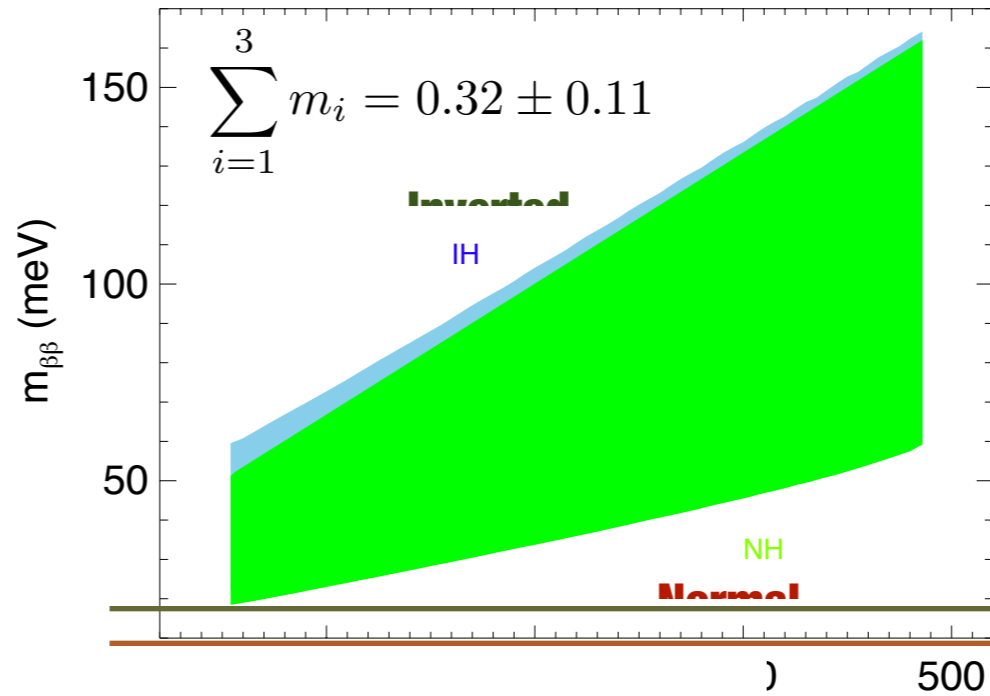
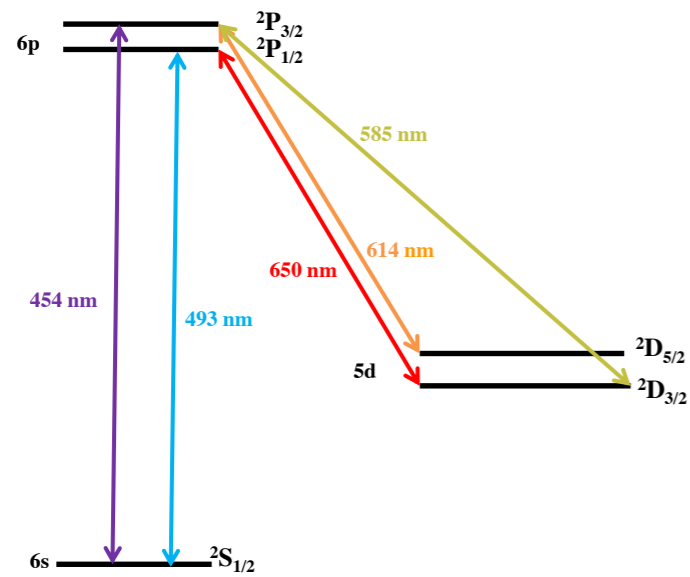
GraXe IBM2
 $m_{\beta\beta}=9$ meV



- efficiency = 0.85
- $\Delta E = 3.6\%$ at 2.5 MeV
- $B = 1. \times 10^{-5}$ ckky

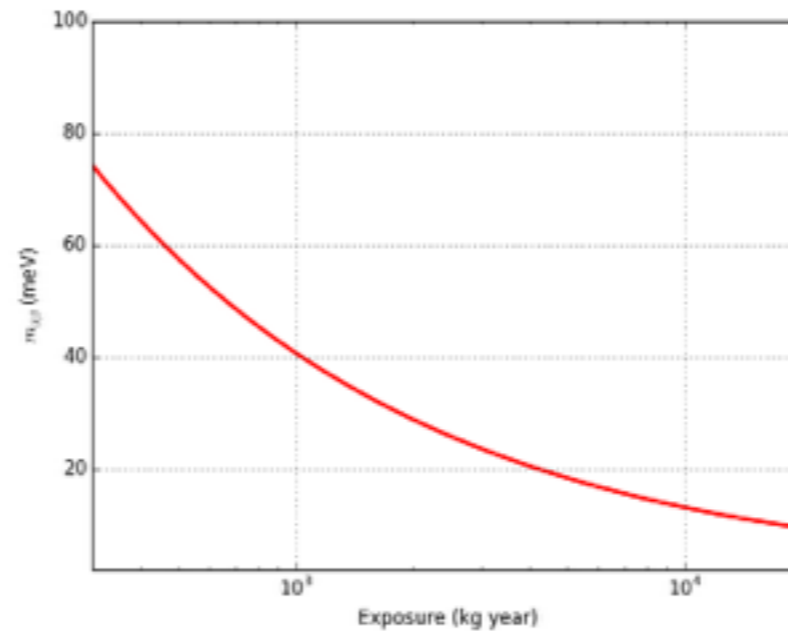
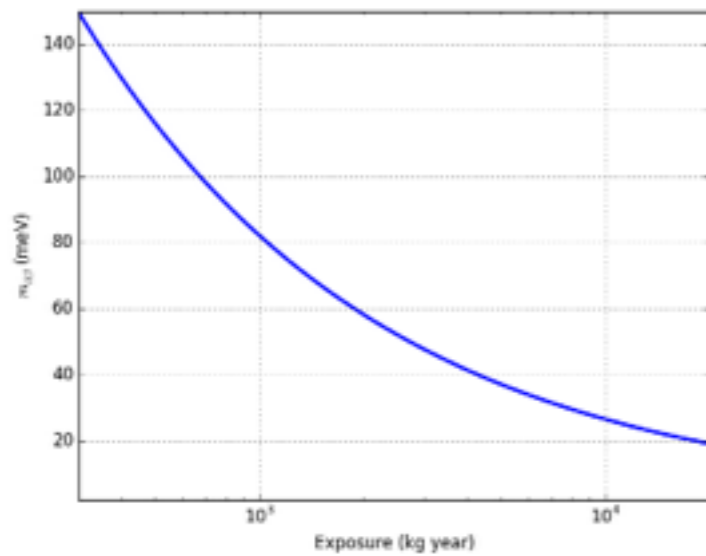
BEXT

- Mass ~2 ton Effective exposure = 20 ton year



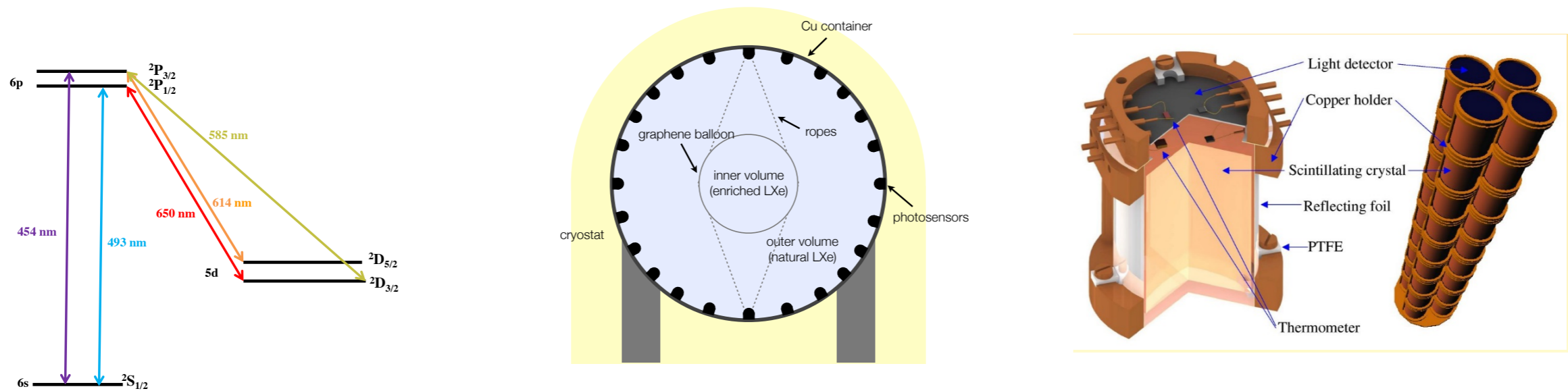
BEXT ISM
 $m_{\beta\beta}=19$ meV

BEXT IBM2
 $m_{\beta\beta}=9$ meV



- efficiency = 0.3
- $\Delta E = 0.5$ % at 2.5 MeV
- $B = 1. \times 10^{-6}$ ckky

Assessment: NEW ideas



- **The three experiments discussed essentially cover the inverse hierarchy even for pessimistic case of NME.**
- **They are examples of ideas that require technological break-throughs (large graphene balloon, barium tagging, large-scale scintillation bolometers) and large target masses.**
- **Discovering that the neutrino is a Majorana particle is possible, but requires a brave heart.**