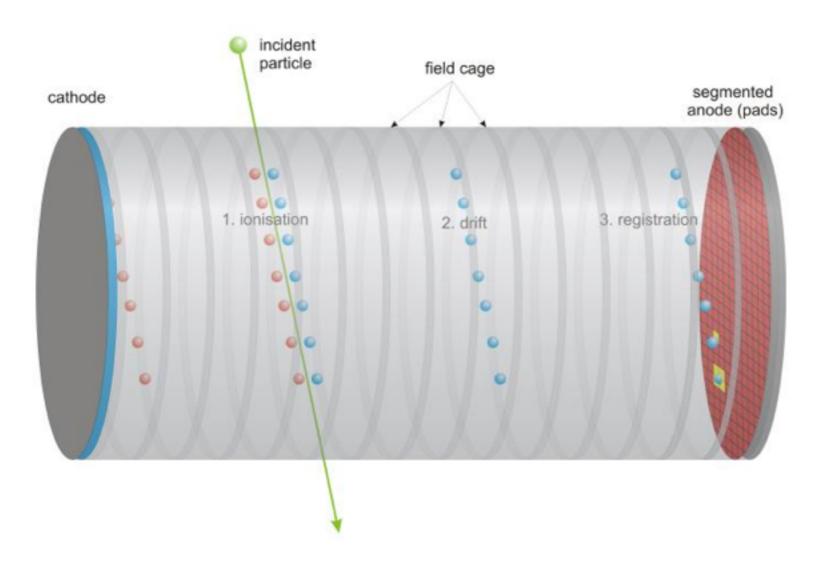
Measuring the neutrino mass

J.J. Gómez Cadenas IFIC (CSIC & UV)

1

St. Andrews, INSS, 2014 Lecture 5

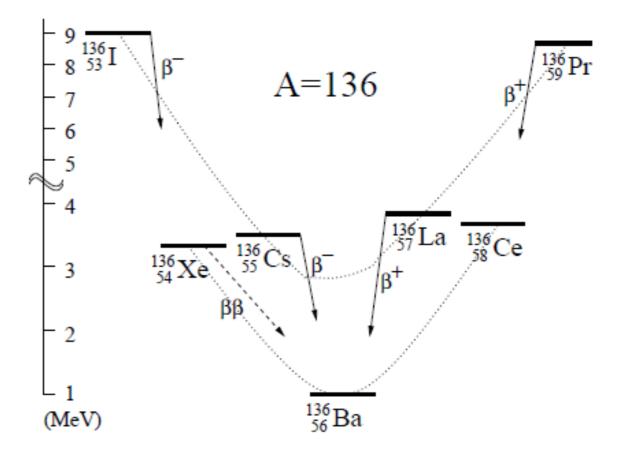
Xenon TPCs



- The Time Projection Chamber (TPC). Invented by Dave Nygren, is one of the most successful detectors in nuclear and particle physics.
- It provides 3D image of tracks and if track contained also energy by calorimetric measurement.

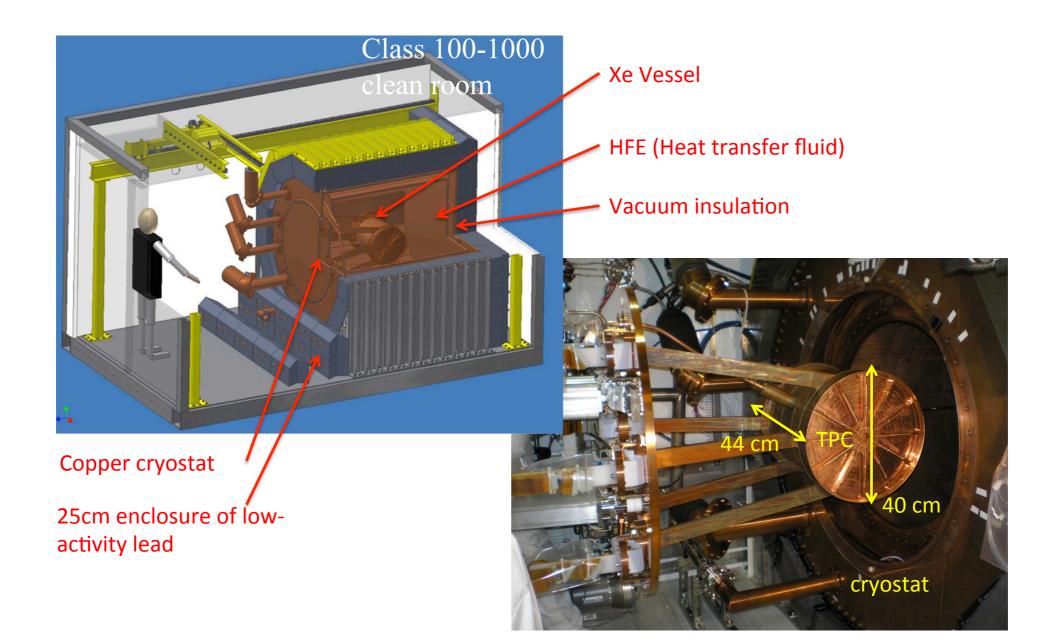
Xenon is a noble gas: one can build at Xenon TPC

- Xenon is a good candidate for bb0n search
 - Q-value larger than energy of gammas from most natural radionuclides
 - Relatively easy to enrich in Xe-136 isotope
 - (no chemistry, centrifuge eff ~ dm=4.7 a.m.u.)
 - No need to grow high-purity crystals, continuous purification is possible (and relatively easy for a noble gas), more easily scalable
 - No long-lived cosmogenically activated isotopes
 - Final state (Ba-136 ion) can, in principle, be tagged, greatly reducing backgrounds

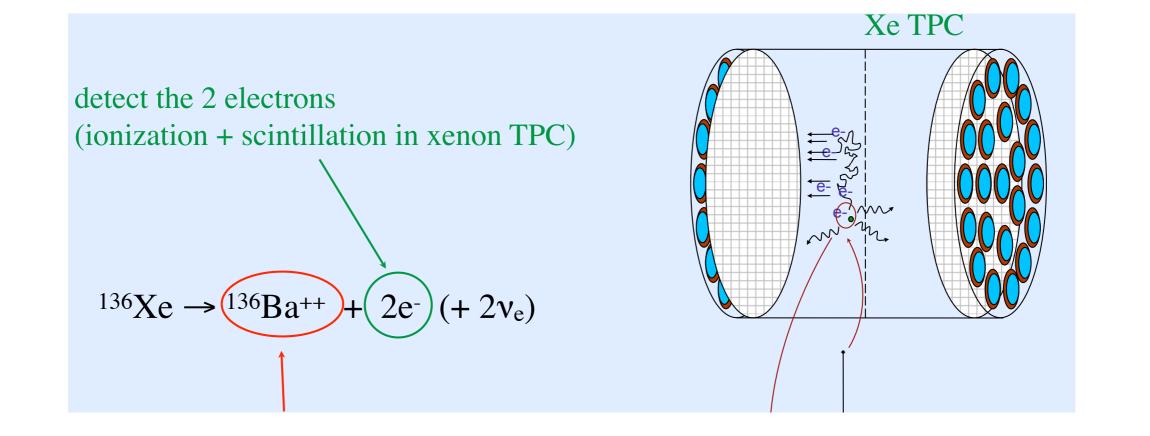




- 200 kg of Xe enriched to 80.6% Xe-136 total procured
 - 175 kg in liquid phase inside a cylindrical Time Projection Chamber
 - ~100 kg current fiducial mass

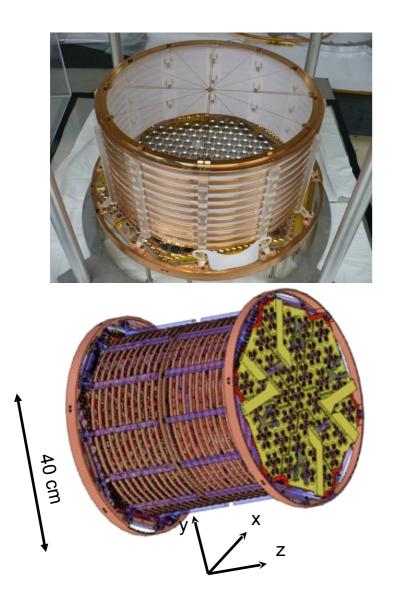


EXO detection strategy



EXO TPC

- 200 kg of Xe enriched to 80.6% Xe-136 total procured
 - 175 kg in liquid phase inside a cylindrical Time Projection Chamber
 - ~100 kg current fiducial mass



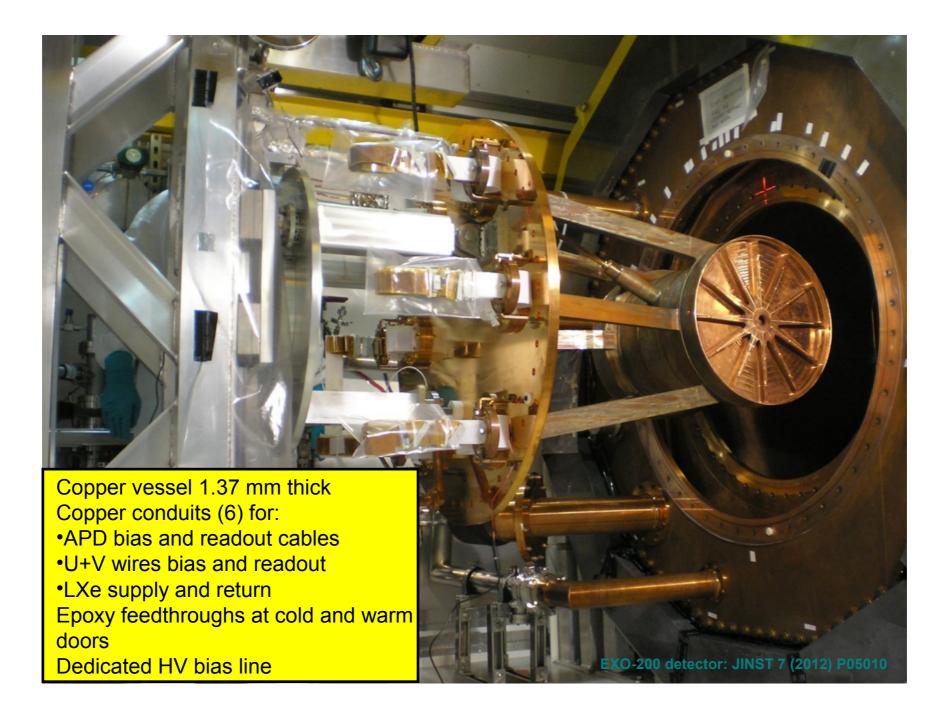
- Common cathode + Two Anodes
 - 376 V/cm drift field
- Each half records both charge and scintillation information with
 - 38 U (charge collection) + 38 V (charge induction) triplet wire channels, crossed at 60 degrees
 - Wire pitch 3 mm (9 mm / channel)
 - Photo-etched Phosphor bronze
 - 234 large area avalanche photo-diodes, in groups of 7 (178 nm Xe light)

APDs

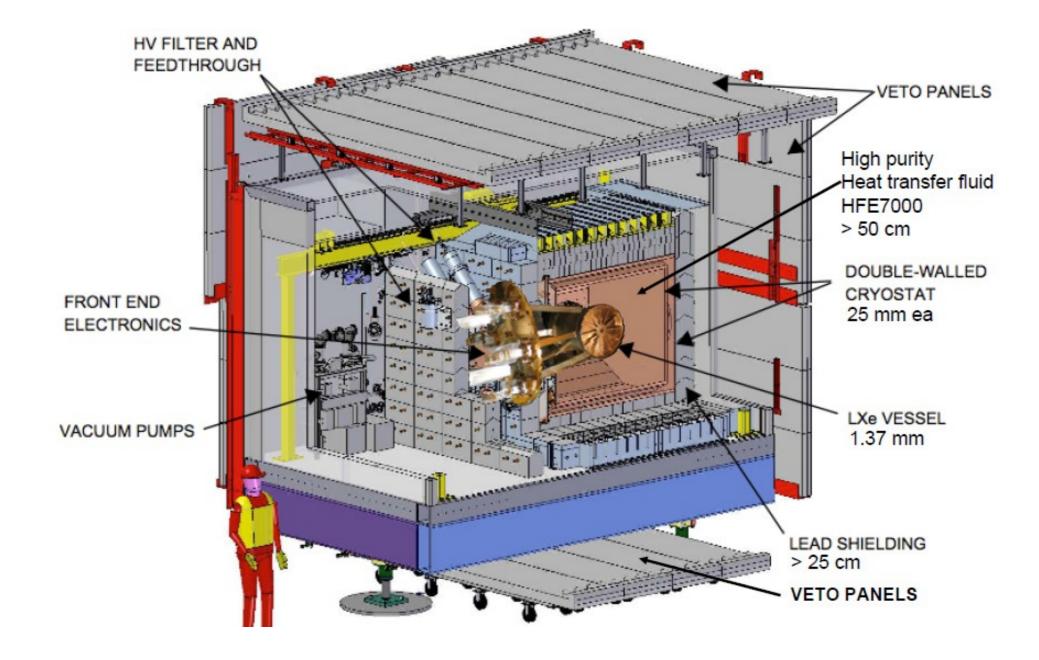


APDs are very clean and light, highly sensitive to VUV Gain is set to ~200 with V~1.5kV Characterization of APDs: NIM A608 68-75 (2009)

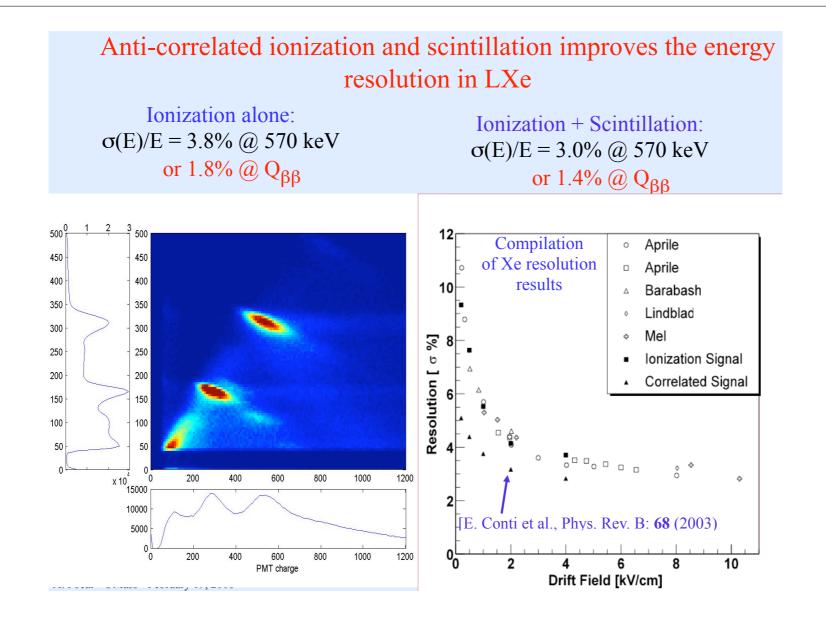
Apparatus at WIPP



Apparatus at WIPP



Energy resolution

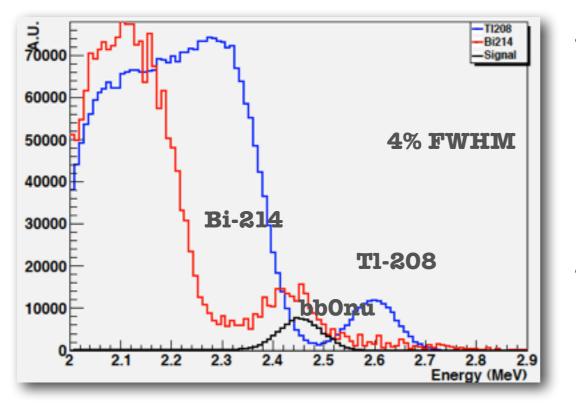


 $\frac{\Delta E}{E}(FWHM) = 2.4 \times \frac{\Delta E}{E}(rms) = 3.6\%$

EXO main backgrounds

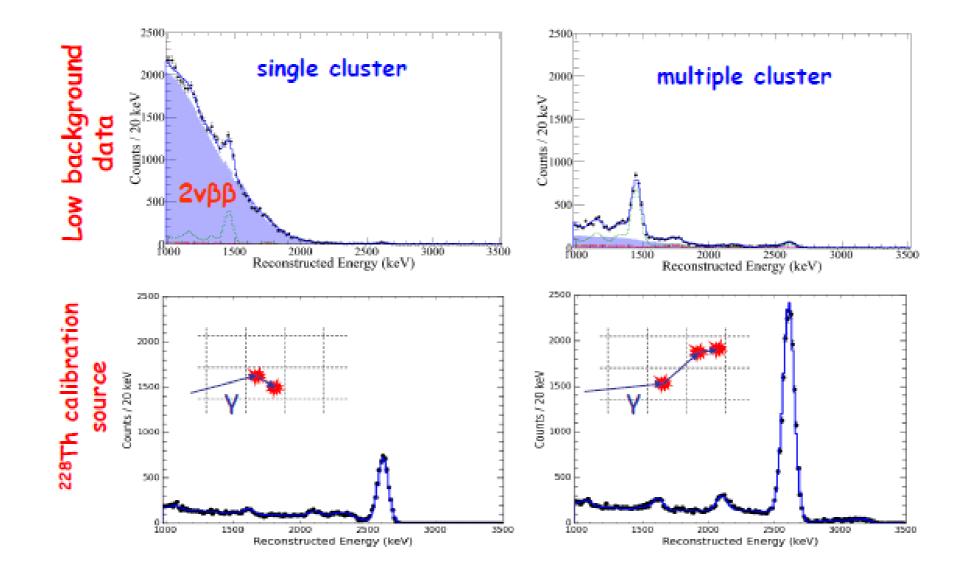
- γ (2449 keV) from ²¹⁴Bi decay (from ²³⁸U and ²²²Rn decay chains)
- γ (2615 keV) from ²⁰⁸Tl decay (from ²³²Th decay chain)
- γ (1.4 MeV) from ⁴⁰K (a concern for the $2\nu\beta\beta$)
- ⁶⁰Co: 1173 + 1333 keV simultaneous γ 's (from ⁶³Cu(α ,n)⁶⁰Co)
- other γ 's in ²³⁸U and ²³²Th chains
- other cosmogenics of Cu (a concern for the $2\nu\beta\beta$)
- in situ cosmogenics in Xe, neutron capture de-excitations, ...

•²²²Rn anywhere (Xe, HFE, air gaps inside lead shield)



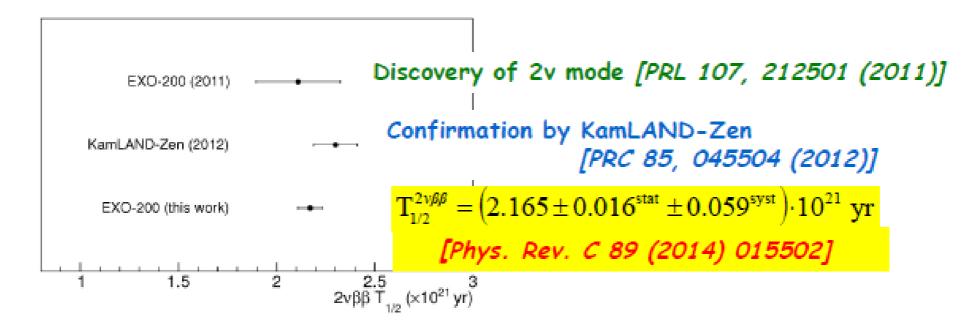
TPC fiducialization (only events in the fiducial volume, away from surfaces), and good 3D location eliminates all alpha background (a concern for Ge, bolometers) leaving only high energy gammas.
 However ~4% FWHM energy resolution does not allow to separate signal peak from leading Bi-214 and TI-208 peaks

Single site vs multiple site: separate signal from backgrounds



Very useful to identify gamma backgrounds!

EXO milestones



- First 0-nu analysis
 - 09/11 04/12 "Run 2a". PRL 109 (2012) 032505, 26.3 kg*yr Xe-136
 - Median expected 90% C.L. U.L.: 0.7e25 yrs. Realized limit: 1.6e25 yrs
- Second Onu-analysis
 - Incorporates several analysis improvements tested in 2nu analysis, more than thrice as much data. Massive effort to assess systematics affecting 0nu search
 - 09/11 09/13 "Run2abc". *Nature 510 (2014) 229*, 100.0 kr*yr Xe-136

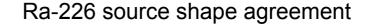
Systematic errors

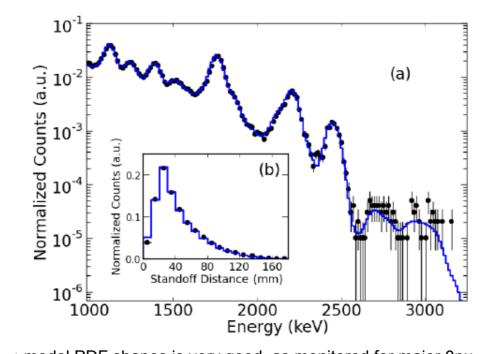
Source	Signal eff. $(\%)$	Error (%)
Event selections		
Summary from [9]	93.1	0.9
Partial reconstruction	90.9	7.8
Fiducial Volume/Rate agreement	-	3.4
Total	84.6	8.6

TABLE I. $0\nu\beta\beta$ signal efficiency and associated systematic errors. 'Partial reconstruction' refers to the requirement that all events be fully reconstructed in X,Y and Z. The summary for event selection from [9] includes all efficiencies and related errors except fiducial volume and partial reconstruction, which have been recalculated in this work for $0\nu\beta\beta$.

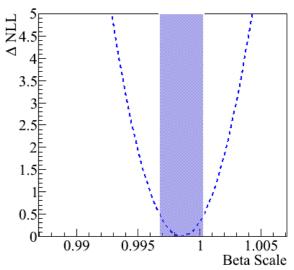
Source	Error (%)
Background shape distortion	9.2
Background model	5.7
Energy resolution variation	1.5
Total	10.9

TABLE II. Systematic errors on the determination of background in the ROI. These arise from incorrect modeling of the background shape (Background shape distortion), incorrect or incomplete background model (Background model), and the residual variation of the energy resolution over time (Energy resolution variation, see e.g. Figure 1).



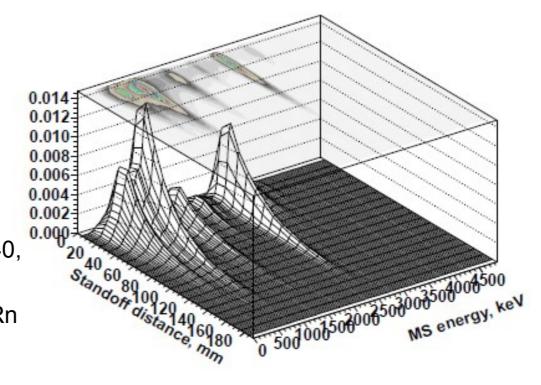


- Energy scales for beta and gammas may not exactly coincide
 - gamma of 1 a.u. of energy deposits energy as several electrons of smaller amount of energy – hence different dE/dX, light/charge ratio, IR/recoil losses
 - even after combining light and charge channels, the effect may not vanish completely due to imperfections of detector response/instrumental effects
 - knowing beta-scale (defined as multiplicative factor B, Eb = B*Eg) is important to accurately define ROI
- Our data is a relatively clean source of beta events - use profile NLL scan to determine the betascale



Fitting the data

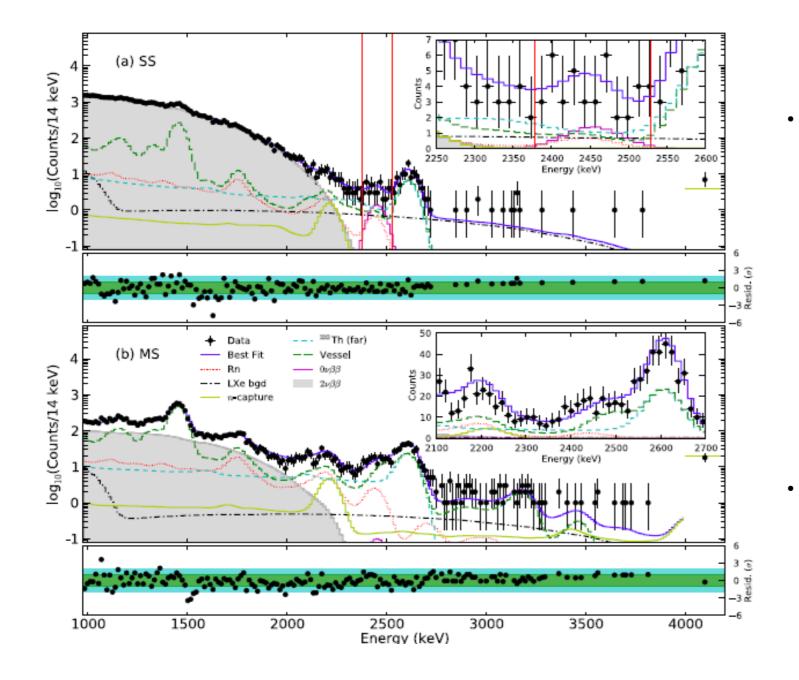
- 3 observables
 - SS and MS energies
 - Standoff distance
- couple dozens parameters
 - Number of events and SS fractions (+ univariate variables for common and ROI specific normalization)
 - Signals: bb2n, bb0n
 - TPC vessel backgrounds: Co-60, K-40, U-238, Th-232, Mn-54, Zn-65
 - LXe backgrounds: Xe-135, Xe-137, Rn
 - Remote backgrounds: Rn in the air gap, Th in cryostat
 - Neutron captures: H in HFE, Cu in cryostat, Xe136 in LXe
- Constraints
 - Normalizations, Rn in LXe, SS fraction (correlated for gammas)



Th232 PDF. True Geant4 energy spectrum was smeared by resolution curve measured from the data

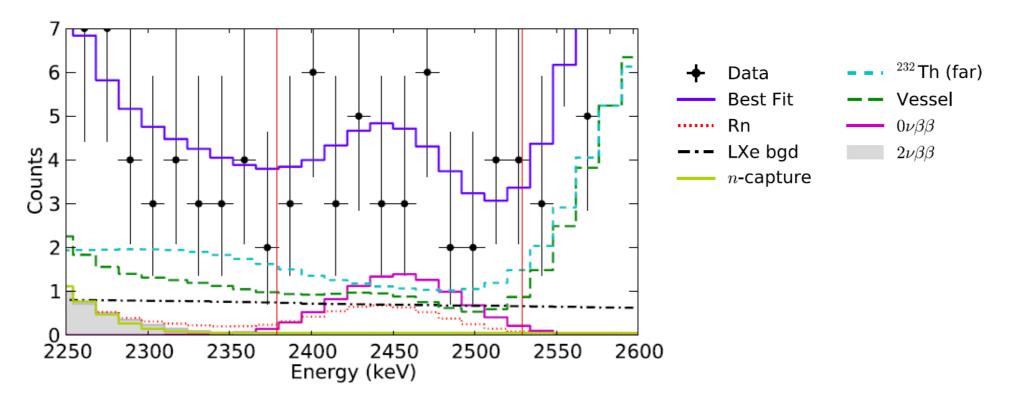
~ ^

Fit results



Notice that EXO spectrum
is in-between that
KamLAND-ZEN (pekas not
resolved) and that of
GERDA (every peak
resolved). Here the
landscape presents broad
peaks corresponding to
~4% FWHM energy
resolution.

Fit results close up



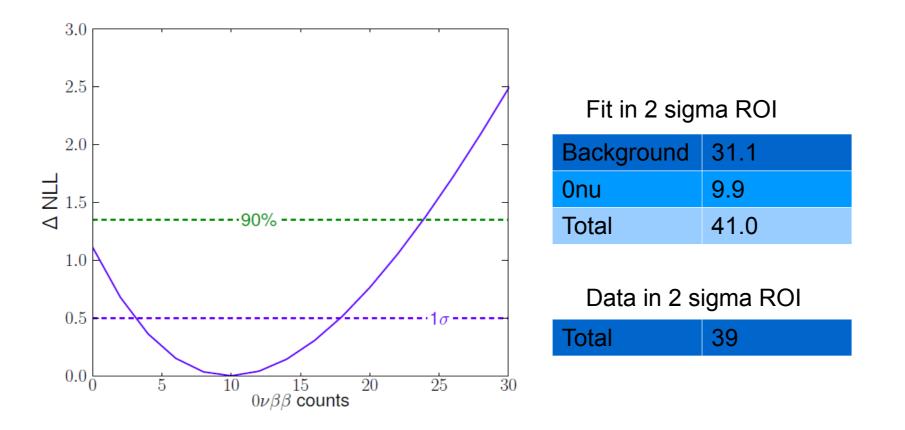
2 sigma ROI breakdown for major backgrounds

	Events	BI, 1e-3 /kg/yr/keV*	
Th-232	16.0		
U-238	8.1		
Xe-137	7.0		
Total	31.1±1.89(stat)±3.3(syst)	1.7±0.2	

• BI ~4 x 10⁻³ ckky

Fit results

•

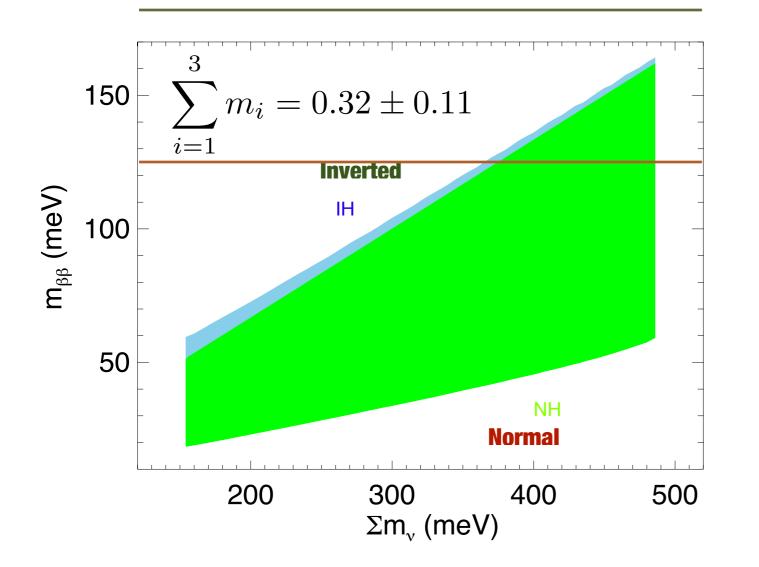


Result is consistent with "background only" hypothesis at ~1.2 sigma

Median 90% C.L. U.L. limit assuming adequate "background-only" model: >1.9e25 yrs Limit from the fit to the actual data: >1.1e25 yrs

Bad luck or hint of a discovery?

Expected: NEXT-100



- · 100 kg yr
- $\Delta E=3.6$ % FWHM at $Q_{\beta\beta}$
- · B=5 x 10⁻³ ckky ⇒
 - · T⁰^v=2 x 10²⁵⇒
 - $\cdot m_{\beta\beta} = 125-352 \text{ meV}$

• Similar result to that of KamLAND-ZEN.





Comming soon to an underground lab near you

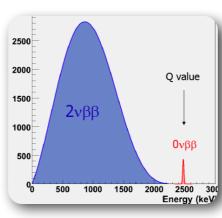
Why NEXT? — Advantages of HPXe technology

 $T_{1/2}^{-1} \propto a \cdot \epsilon \cdot \sqrt{\frac{Mt}{\Delta E \cdot R}}$

Cost



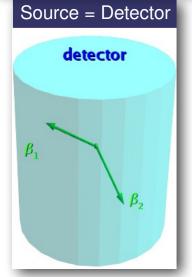
Xenon is the cheapest and easiest to enrich of all ββ isotopes. No long lived radioactive isotopes. There is already 1 ton of enriched xenon in the World.



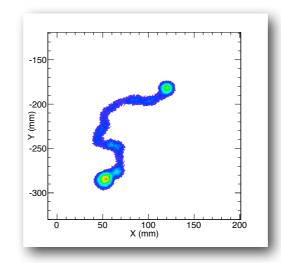


HPXe TPC is the only xenon detector that provides good energy resolution (better 1% FWHM at Qbb)

Scalability



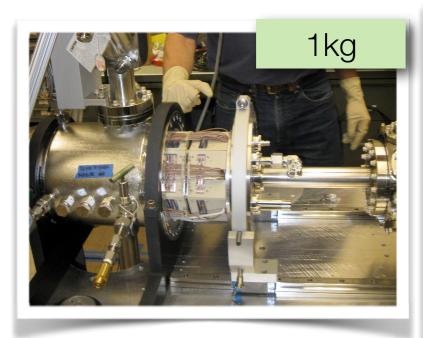
Xenon is a noble gas suitable to build a TPC. No dead areas, S/N improves with L



Background

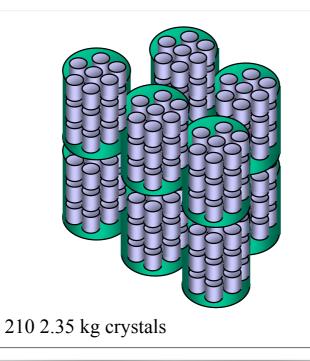
HPXe TPC is the only xenon detector that provides topological signal

Scalability



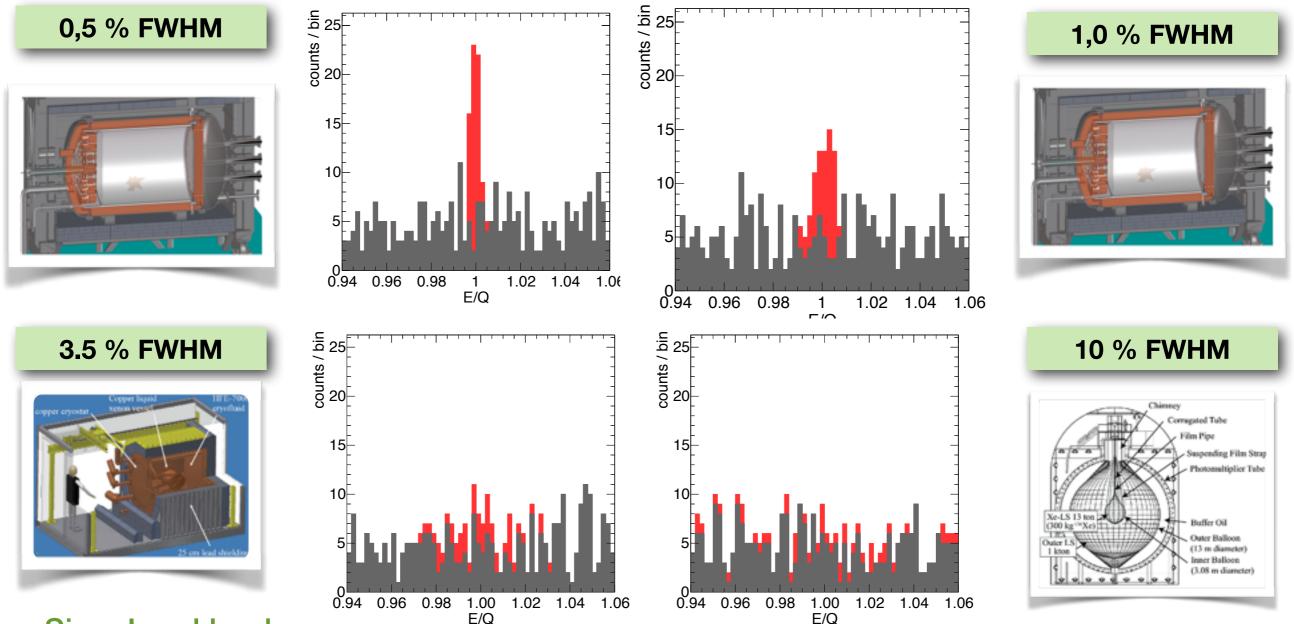






Economy of scale: Double L, signal increases 8 (L³), background increases 4 (L²), S/N improves by a factor 2

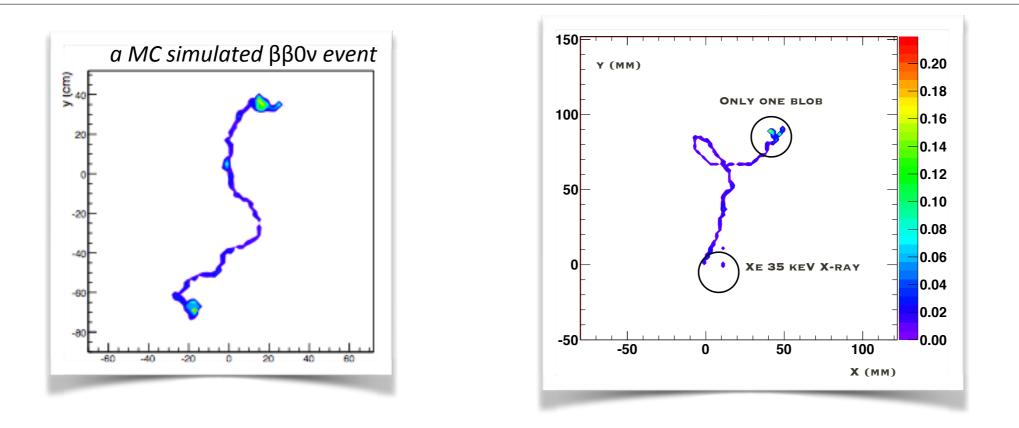
Energy resolution makes a difference



Signal and backg

- •Signal: 50 events, $T^{0v} = 5 \ 10^{25}$ y and an exposure of 1 ton year.
- •Background 1 count/keV/ton/year.

Topological background reduction

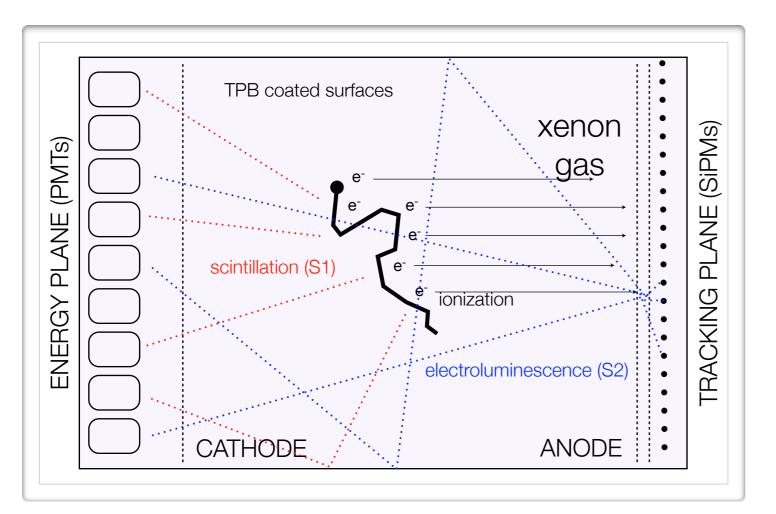


•In xenon gas at 15 bar, a $\beta\beta$ event is a twisted track, 10 cm long, with two energy blobs at the two ends and no additional floating clusters.

 Instead the backgrounds are single electrons, accompanied 85% of the time by X-rays (Xenon de-excitation).

•HPXe TPC offers a signal that looks like a signal: two identified electrons with an energy within 10 keV of Qbb

NEXT: A light TPC



EL mode is essential to get lineal gain, therefore avoiding avalanche fluctuations and fully exploiting the excellent Fano factor in gas It is a High Pressure Xenon
 (HPXe) TPC operating in EL mode.

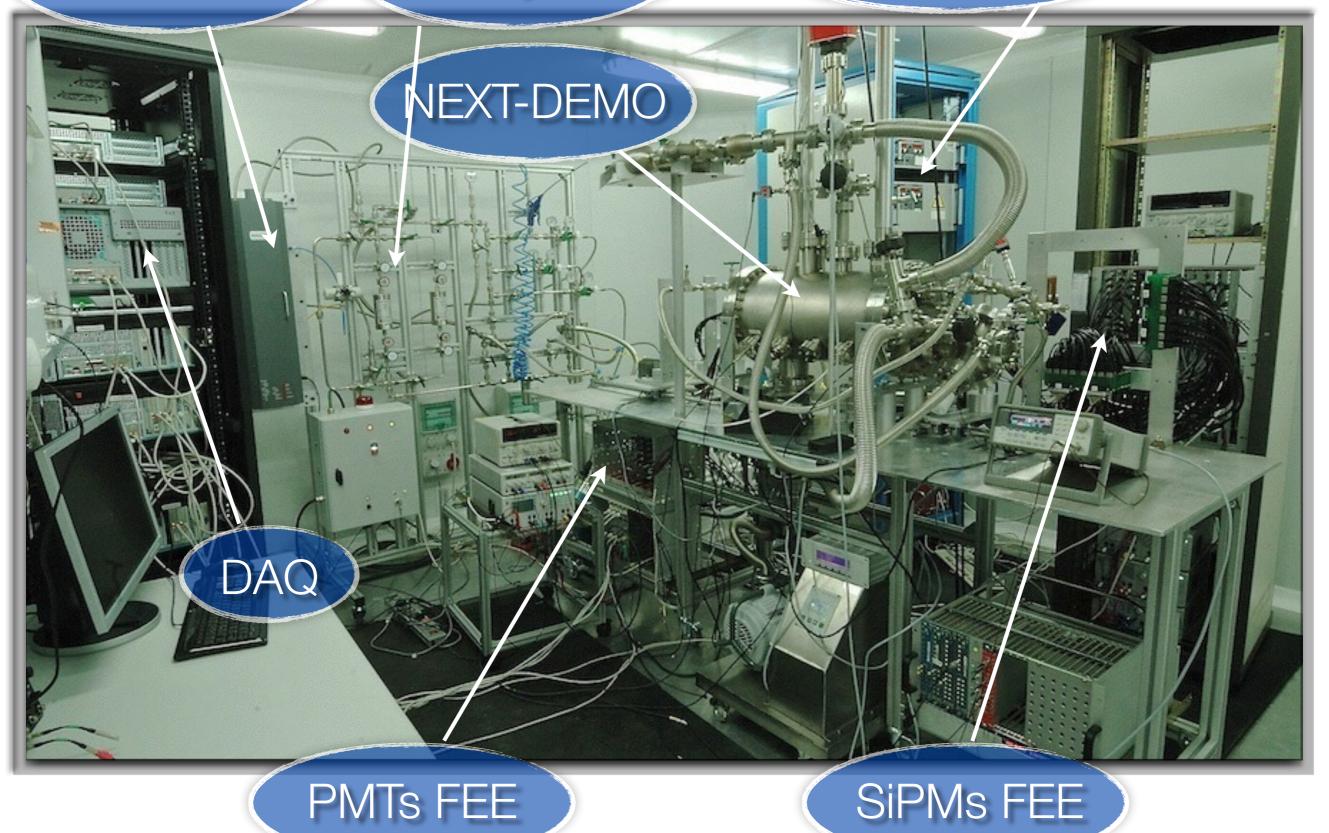
It is filled with 100 kg of Xenon enriched at 90% in Xe-136 (in stock) at a pressure of 15 bar.
The event energy is integrated

•The event energy is integrated by a plane of radiopure PMTs located behind a transparent cathode (energy plane), which also provide t0.

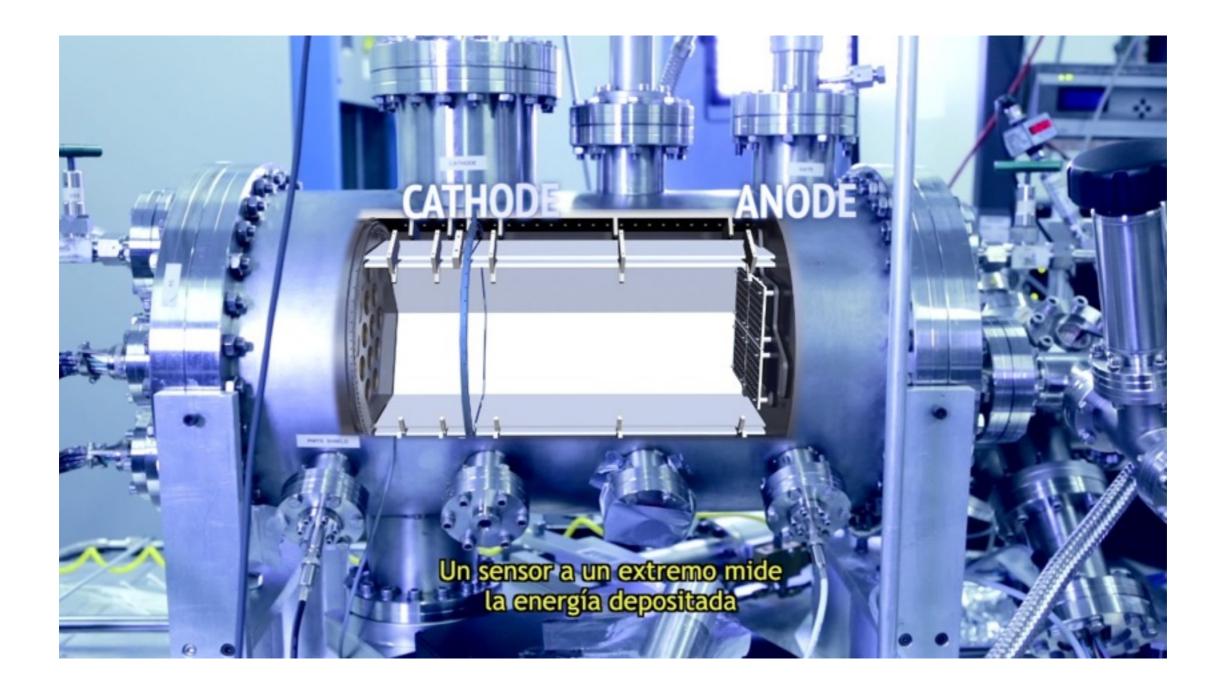
•The event topology is reconstructed by a plane of radiopure silicon pixels (MPPCs) (tracking plane).

Hot Getter Gas System

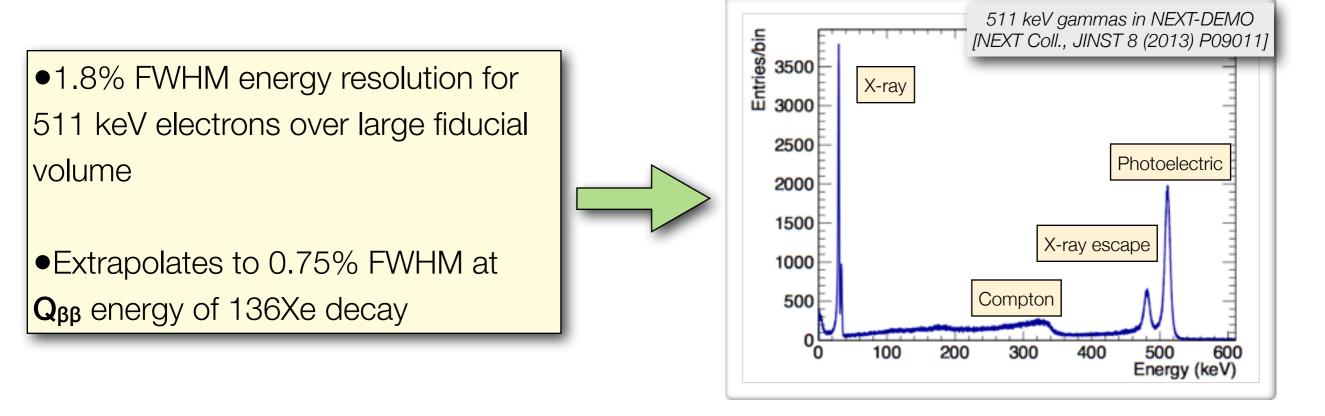
HHV modules



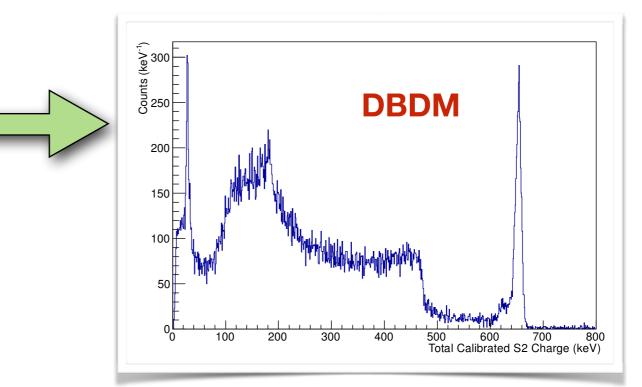
Anatomy of NEXT-DEMO



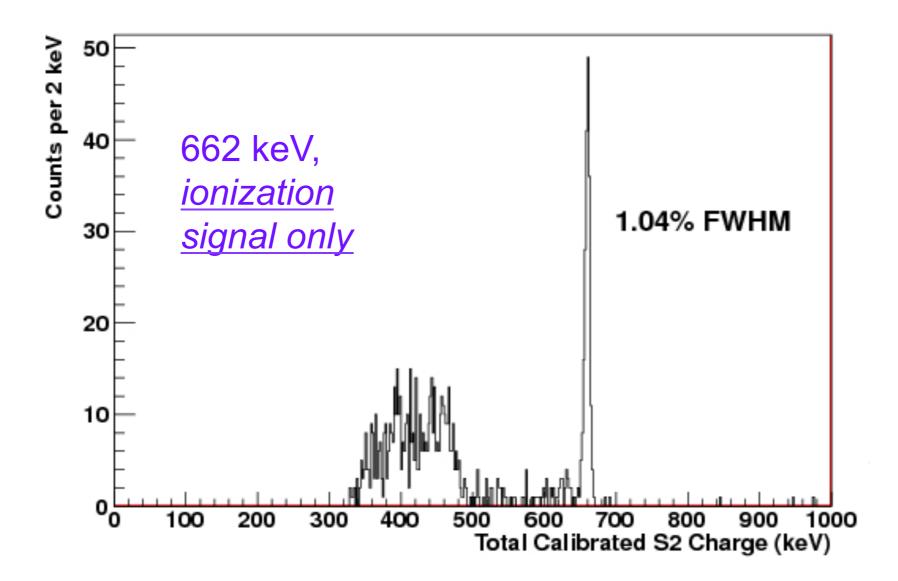
NEXT R&D: detector performance achievements



The DBDM prototype at LBNL
 extrapolates to 0.5 % FWHM at Q_{ββ}
 using 660 Cs-137 electrons

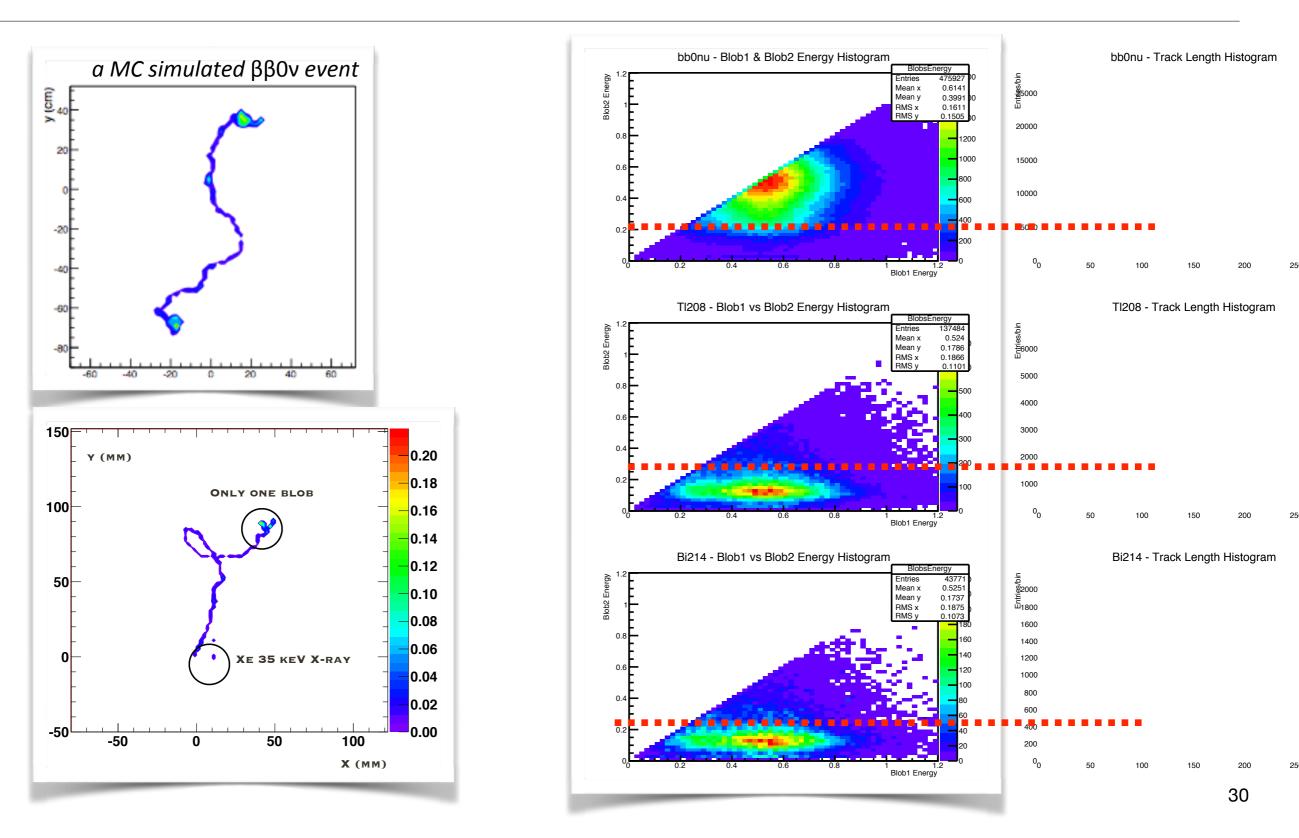


The beauty of resolution

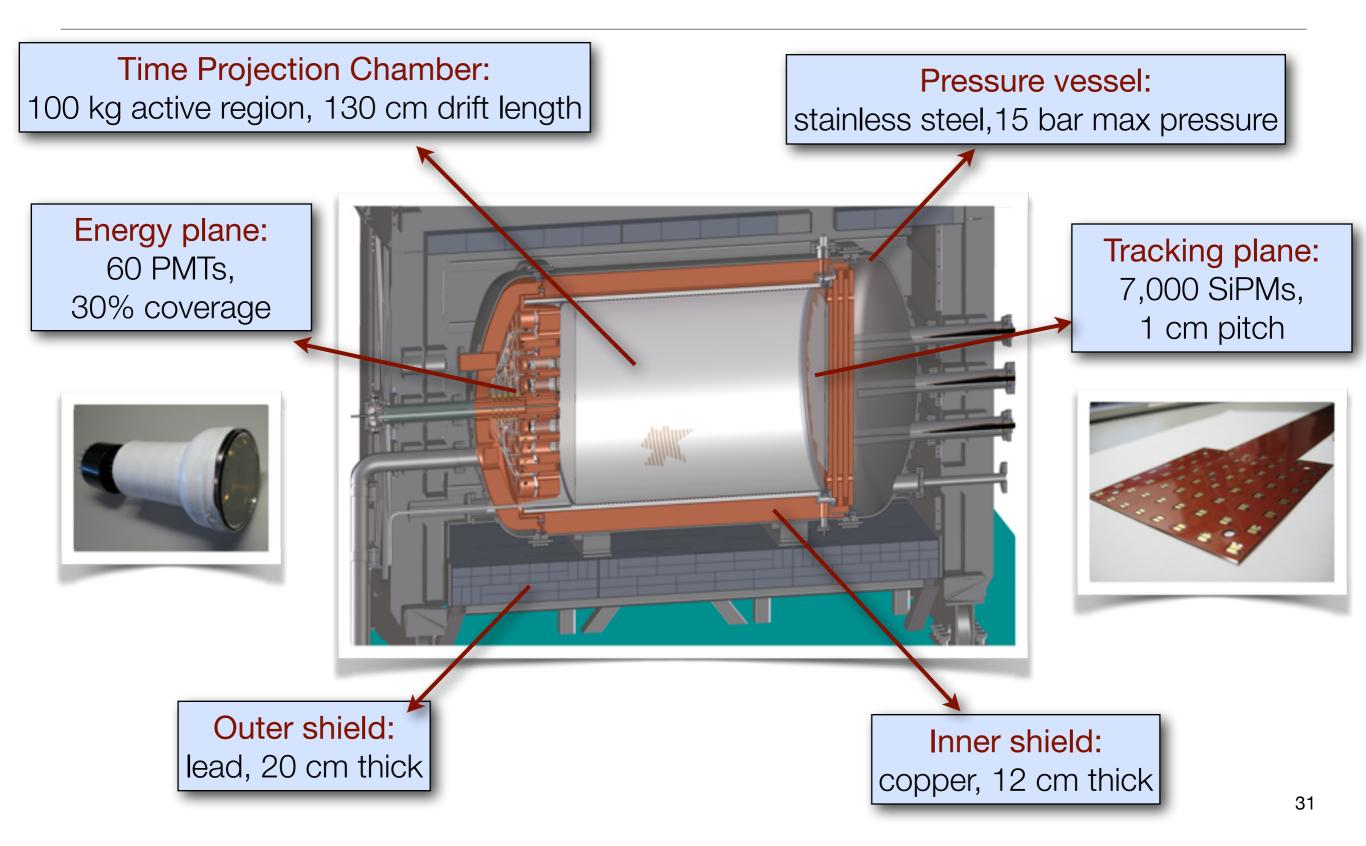


•Extrapolates to 0.5 % at $Q_{\beta\beta}$

Topological background reduction



NEXT 100 kg detector at LSC: main features



NEXT 100 kg radioactive budget

Vessel

- Stainless steel 316Ti; 1121 kg
- Activity TI-208: <0.150 mBq/kg
- Activity Bi-214: <0.460 mBq/kg

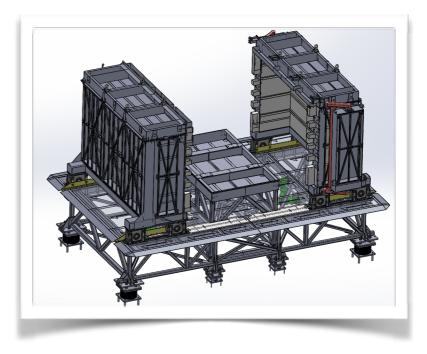
Shielding

- Lead; 13000 kg
- Activity TI-208: <0.031 mBq/kg
- Activity Bi-214: <0.35 mBq/kg

Lead Castle and Pressure Vessel: Activity shielded by ICS.



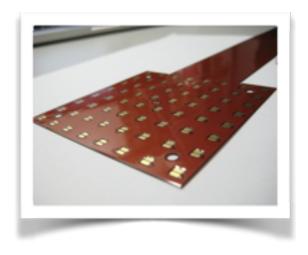
Activity shielded by ICS



NEXT 100 kg radioactive budget

Kapton Dice boards

- Kapton and copper;107 units
- Activity TI-208: -0.040 mBq/unit
- Activity Bi-214: -0.030 mBq/unit



PMTs

- Hamamatsu R11410-10; 60 units
- Activity TI-208: -0.140 mBq/unit
- Activity Bi-214: -0.500 mBq/unit

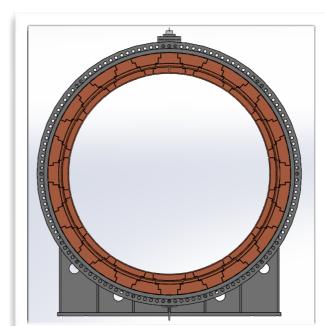


Sensors: Activity level ~3 mBq per plane. Actual measurements. Not shielded.

NEXT 100 kg radioactive budget

ICS, and support plates

- Copper (CuA1); ~9500 kg
- Activity TI-208: <0.001 mBq/kg
- Activity Bi-214: <0.012 mBq/kg



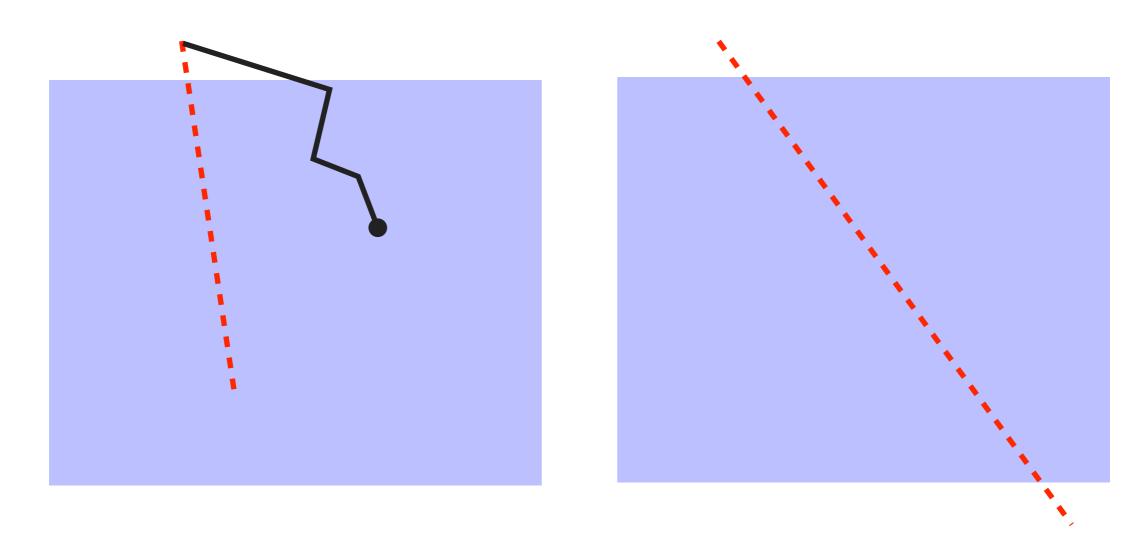
Copper

Electroformed commercial copper. Current measurements show our stock to be very radio pure, but only limits so far.



Residual radioactivity of ICS partially shielded (self-shielding)

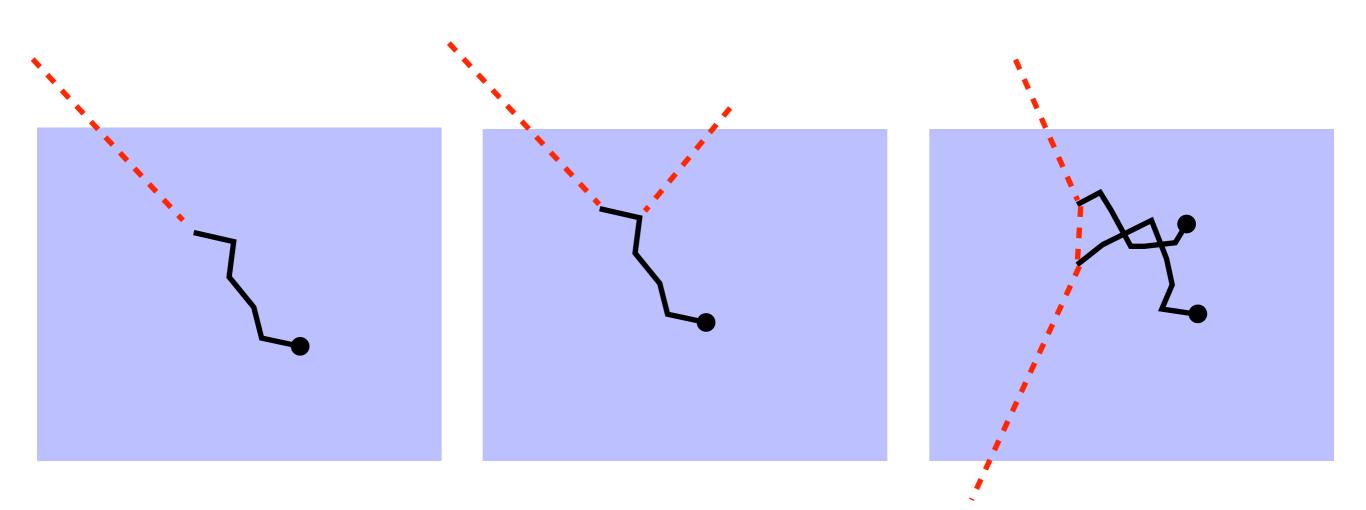
NEXT100 rejection of backgrounds



A transparent target, away from surfaces

 Veto of effectively all charged backgrounds entering the detector (left). High-energy gammas have a long interaction length (>3 m) in HPXe.

NEXT100 rejection of backgrounds



The 2-electron signature

 Interaction of high-energy gammas (from TI-208 and Bi-214) in the HPXe can generate electron tracks with energies around the Q value of Xe-136. However, electron often accompanied of satellite clusters and single blob deposit

NEXT100 rejection of backgrounds

	0νββ	TI-208	Bi-214
Fiducial E>2 MeV	67.86%	0.25%	0.01%
ROI	95.52%	8.99%	64.66%
1 track	74.60%	1.86%	12.54%
2 blobs	73.76%	9.60%	9.89%

The 2-electron analysis

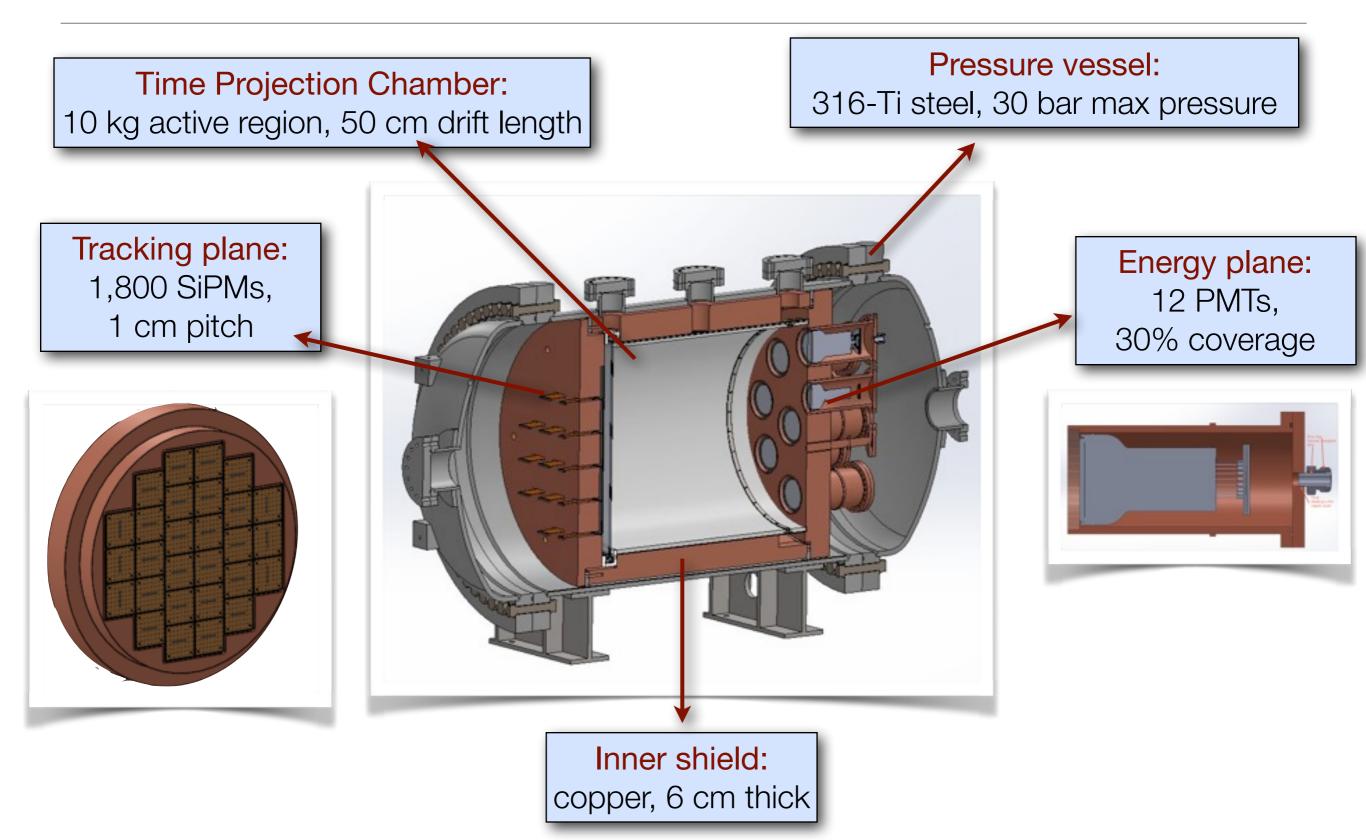
•Effect of the filters (cuts) defining an event with 2 electrons and energy in a ROI of 2σ around $Q_{\beta\beta}$.

- •Efficiency for signal ~35% for suppression factors $4-8 \times 10^{-7}$
- •Topology rejection is the product of 1 track x 2 blobs conditions

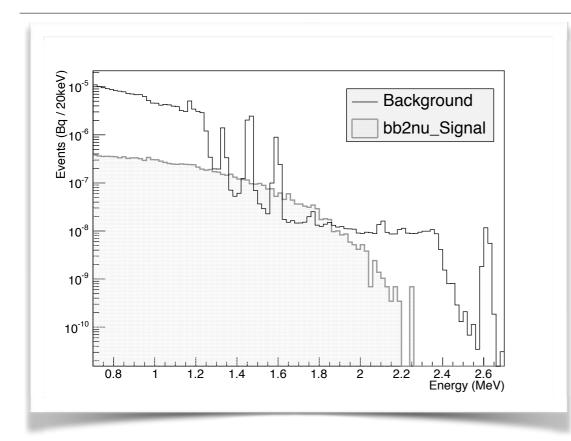
NEXT 100 expected background

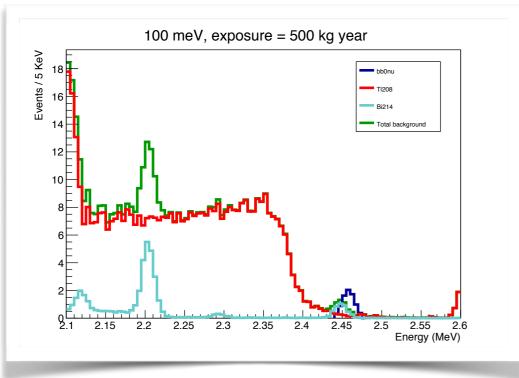
	Activity (Bq)		Rejection Factors		Final rate (ckky)	
	TI-208	Bi-214	TI-208	Bi-214	TI-208	Bi-214
Dice Boards	4,28E-03	3,21E-03	7,90E-07	8,85E-07	3,047E-05	2,560E-05
PMTs	8,40E-03	3,00E-02	3,30E-07	2,68E-07	2,498E-05	7,244E-05
Field Cage	4,38E-03	1,53E-02	5,30E-07	8,02E-07	2,091E-05	1,107E-04
ICS	1,326E-02	1,105E-01	1,100E-07	8,400E-08	1,315E-05	8,365E-05
Vessel	1,66E-01	5,16E-01	1,10E-08	2,80E-09	1,644E-05	1,301E-05
Shielding Lead	6,266E-01	1,084E+00	2,000E-09	1,000E-10	1,129E-05	9,763E-07
SUBTOTAL	8,23E-01	1,76E+00			1,172E-04	3,063E-04
TOTAL BKGND	2,58E+00				4,24E-04	

NEW (NEXT-WHITE) at glance



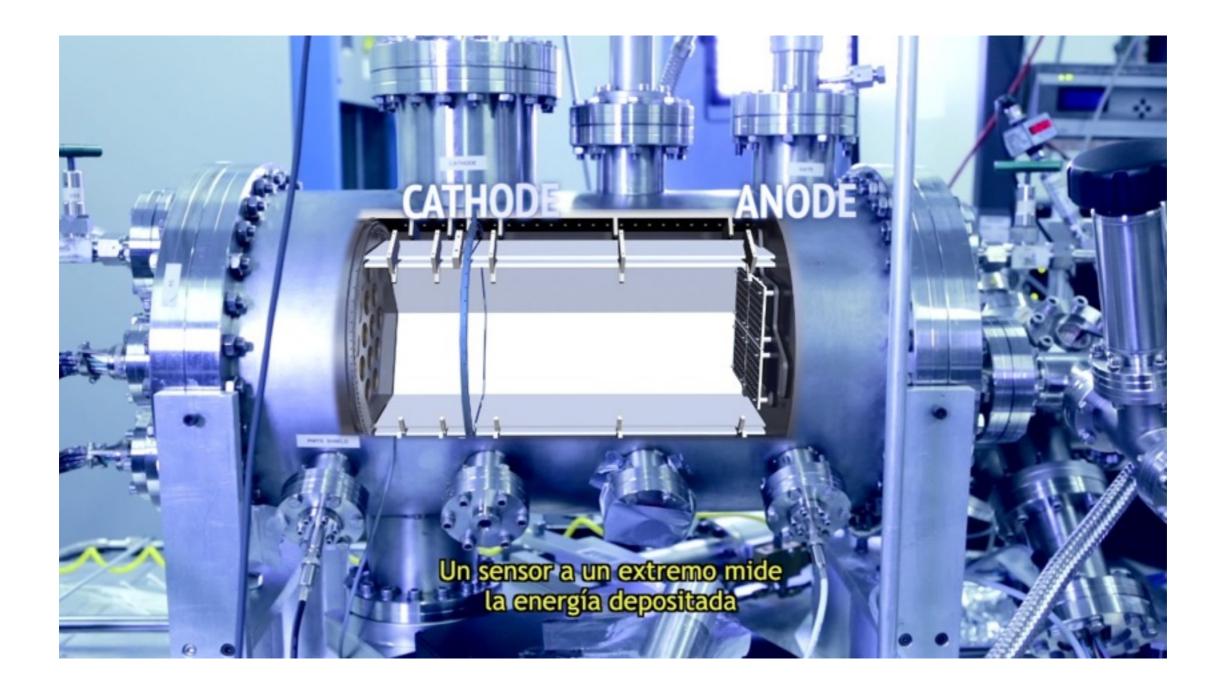
Goals of NEW



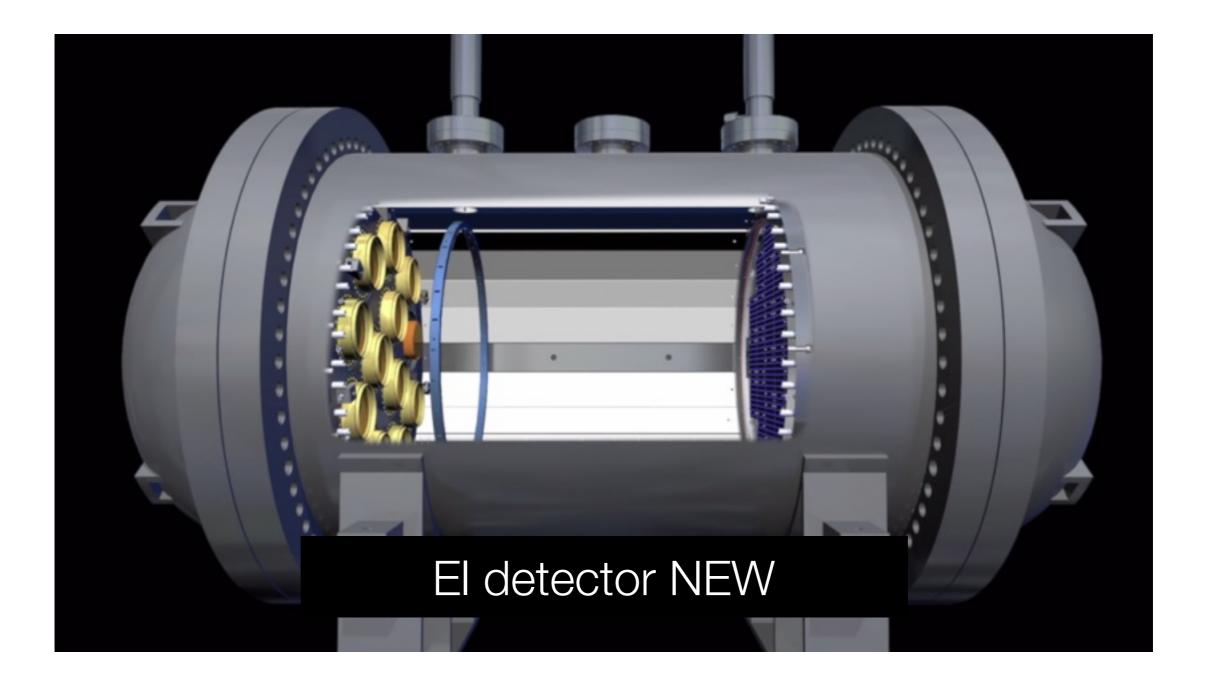


- •Measure the expected backgrounds from the different isotopes, but specially Bi-214 and TI-208.
- •Validate NEXT background model using measurement.
- Identify any unexpected source of background (correct if needed)
- •Observe $\beta\beta2\nu$ signal.
- •Demonstrate energy resolution: our goal is to reach 0.5 % FWHM in the large detector.
- Demonstrate topological signature from data ($\beta\beta$ 2v and TI-208 double escape peak).
- •Certify technology and underground operation with enriched xenon.

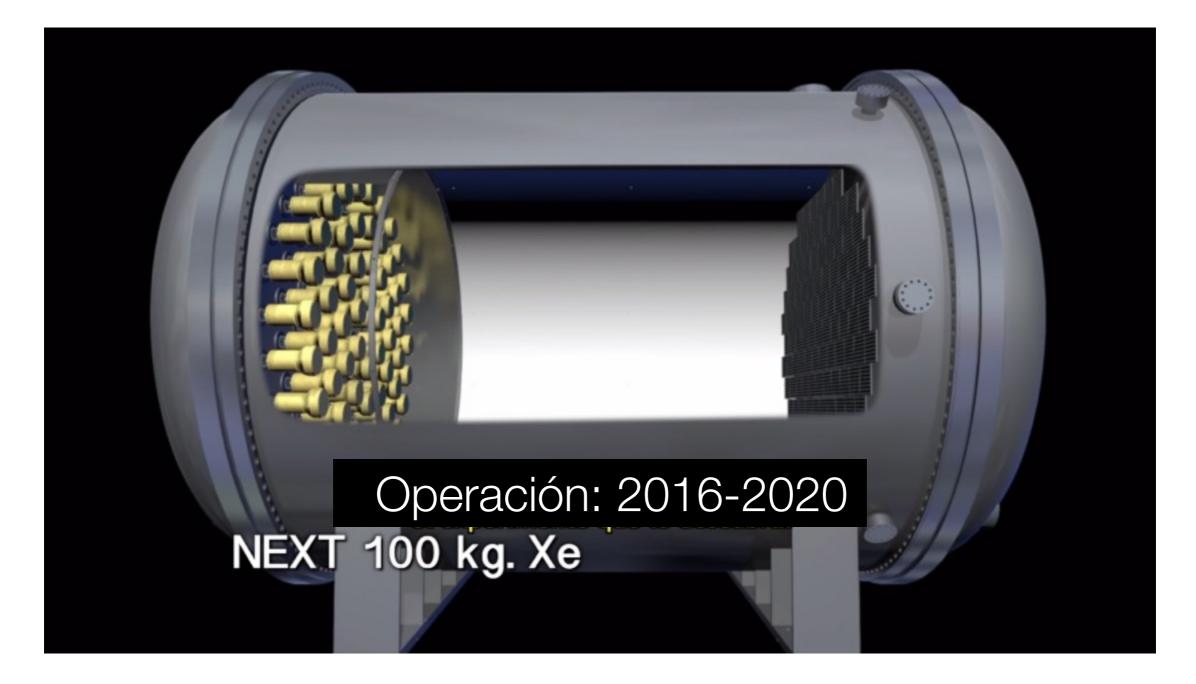
Anatomy of NEXT-DEMO



NEW

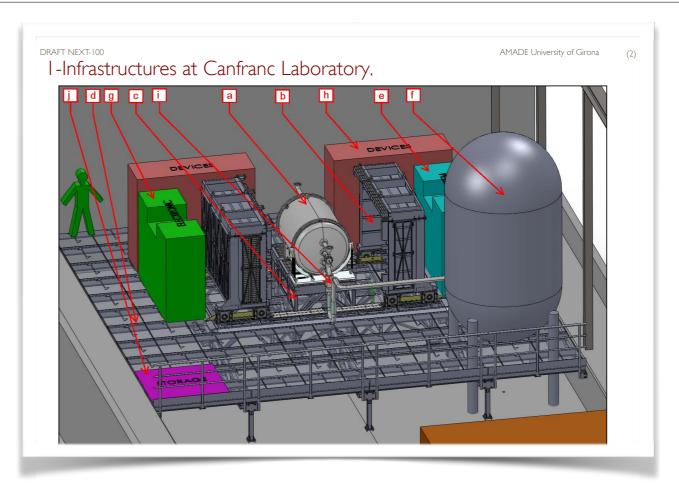


NEXT-100



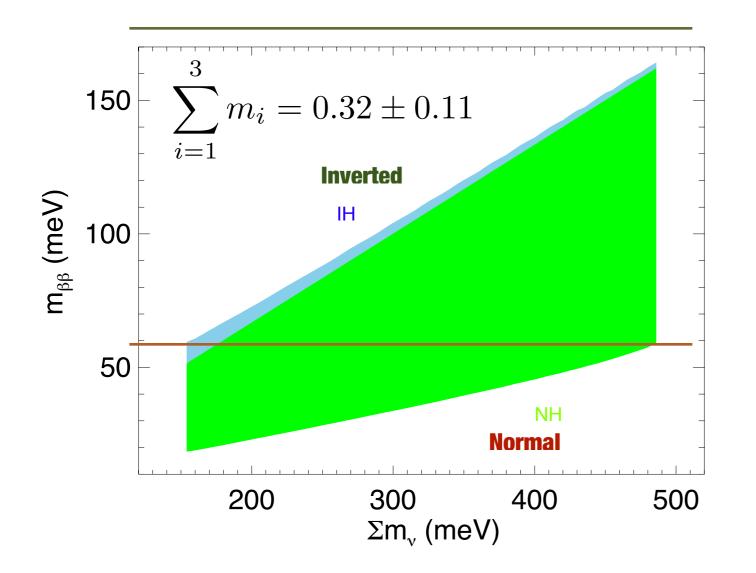
NEXT at LSC





Infrastructures: platform, lead castle, gas system, emergency recovery system, completed. First phase of experiment starts in 2015. In stock, 100 kg of enriched xenon and 100 kg of depleted xenon.

Second phase: the 600 kg run



- · 300 kg yr in 5 years
- $\Delta E=0.7$ % FWHM at $Q_{\beta\beta}$
- B=5 x 10⁻⁴ ckky \Rightarrow
 - · T⁰^v=6 x 10²⁵⇒
 - · $m_{\beta\beta} = 70-196 \text{ meV}$

• Will Cover a significant fraction of cosmological region.