

# Accelerator Neutrino (2)

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- **1<sup>st</sup> Lecture**
  - What can we learn by Long baseline accelerator experiments
  - Latest status(1)
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  - Near future prospect
- **3<sup>rd</sup> Lecture**
  - Technologies in the long baseline accelerator experiments
  - Future prospect

# What is the next step?

- $\sin^2 2\theta_{13}$  was determined precisely by reactor experiments.
- $\nu_e$  appearance via  $\sin^2 2\theta_{13}$  was also observed by an accelerator experiment.
- This is not the end of the story, but actually start of the new story!

# $\nu_e$ appearance probability Leading term

$$P(\nu_\mu \rightarrow \nu_e) = 4 C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \Phi_{31}$$

Leading term

$$C_{ij} = \cos \theta_{ij}, S_{ij} = \sin \theta_{ij}$$

$$\Phi_{ij} = \Delta m_{ij}^2 \frac{L}{4E_\nu}$$

# $\nu_e$ appearance probability (exact formula in vacuum)

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & 4C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \Phi_{31} && \text{Leading term} && \Theta_{13} \\
 & + 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cos \Phi_{32} \sin \Phi_{31} \sin \Phi_{21} && \text{CPC} \\
 & - 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \Phi_{32} \sin \Phi_{31} \sin \Phi_{21} && \text{CPV} \\
 & + 4S_{12}^2 C_{13}^2 (C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta) \sin^2 \Phi_{21} && \text{Solar}
 \end{aligned}$$

$$C_{ij} = \cos \theta_{ij}, S_{ij} = \sin \theta_{ij}$$

$$\Phi_{ij} = \Delta m_{ij}^2 \frac{L}{4E_\nu}$$

replace  $\delta$  by  $-\delta$  for  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

CP violating term introduced by interference btw.  $\theta_{13}$  and  $\theta_{12}$

# $\nu_e$ appearance at around oscillation maximum

$$P(\nu_\mu \rightarrow \nu_e) = 4C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \Phi_{31}$$

$\theta_{13}$

$$+ 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cos \Phi_{32} \sin \Phi_{31} \sin \Phi_{21}$$

CPC

$$- 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \Phi_{32} \sin \Phi_{31} \sin \Phi_{21}$$

CPV

$$+ 4S_{12}^2 C_{13}^2 (C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta) \sin^2 \Phi_{21}$$

Solar

$$C_{ij} = \cos \theta_{ij}, S_{ij} = \sin \theta_{ij}, \quad \Phi_{ij} = \Delta m_{ij}^2 \frac{L}{4E_\nu}$$

$$P(\nu_\mu \rightarrow \nu_e) \cong 4C_{13}^2 S_{13}^2 S_{23}^2 - 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \Phi_{21}$$



Observation

Q. Calculate how much  $P(\nu_\mu \rightarrow \nu_e)$  changes between  $\delta=0$  and  $\delta=90^\circ$

$$P(\nu_\mu \rightarrow \nu_e) \cong 4C_{13}^2 S_{13}^2 S_{23}^2 - 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \Phi_{21}$$

$$C_{ij} = \cos \theta_{ij}, S_{ij} = \sin \theta_{ij}, \quad \Phi_{ij} = \Delta m_{ij}^2 \frac{L}{4E_\nu}$$

$$\theta_{12} = 34^\circ (\cos \theta_{12} = 0.83, \sin \theta_{12} = 0.56)$$

$$\theta_{23} = 45^\circ$$

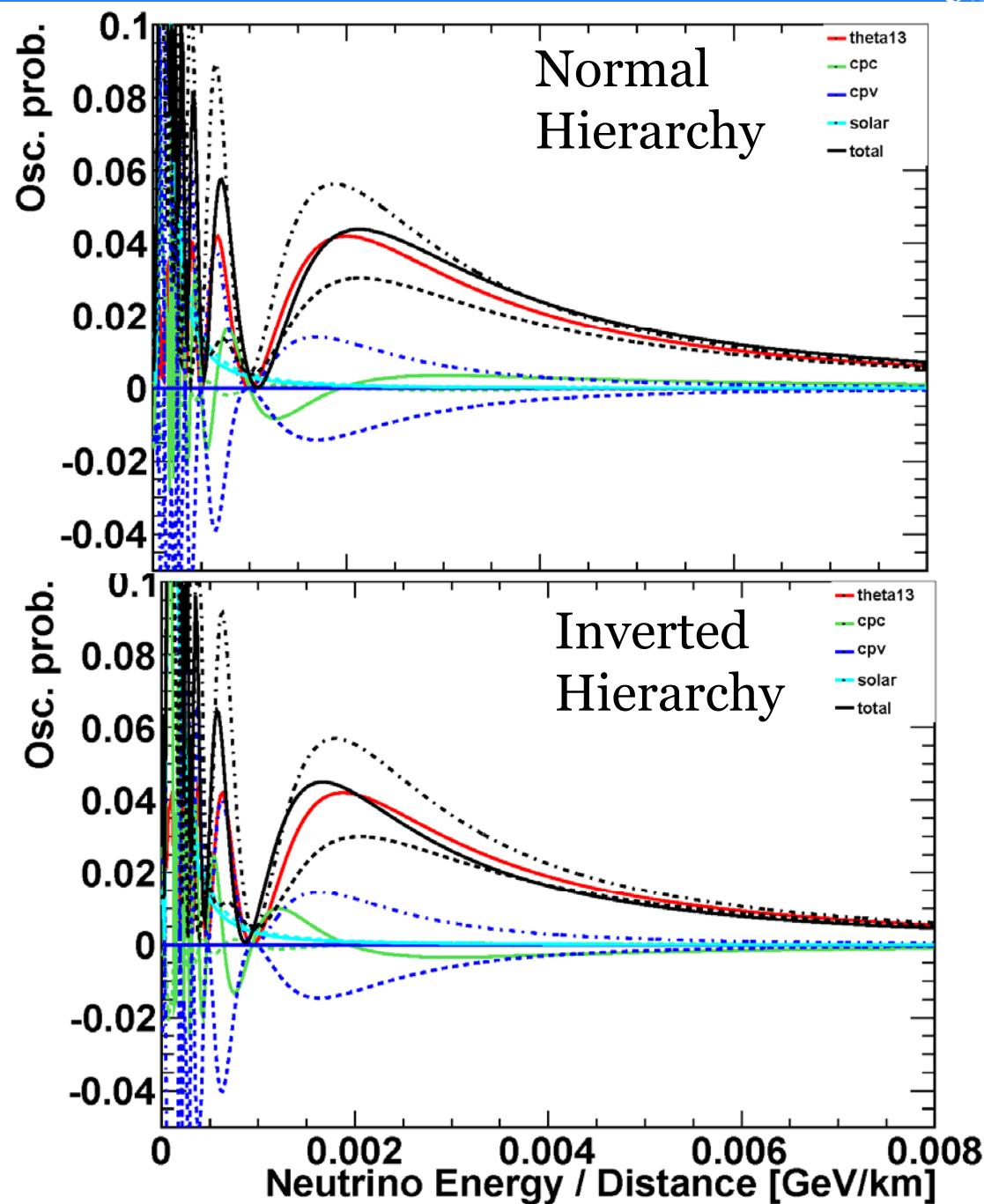
$$\theta_{13} = 9.0^\circ (\cos \theta_{13} = 0.99, \sin \theta_{13} = 0.16)$$

Max. 27% asymmetry (violation) by CP phase

## $\nu_\mu \rightarrow \nu_e$ oscillation probability

- $\Delta m^2_{23} = 2.4\text{e-}3 \text{ eV}^2$
- $\Delta m^2_{12} = 7.59\text{e-}5 \text{ eV}^2$
- $\theta_{12} = 34^\circ$
- $\theta_{23} = 45^\circ$
- $\sin^2 2\theta_{13} = 0.084$
- w/o matter effect

$\theta_{13}$  CPC CPV Solar Total  
 —  $\delta=0$   
 - - -  $\delta=\pi/2$   
 - · -  $\delta=-\pi/2$

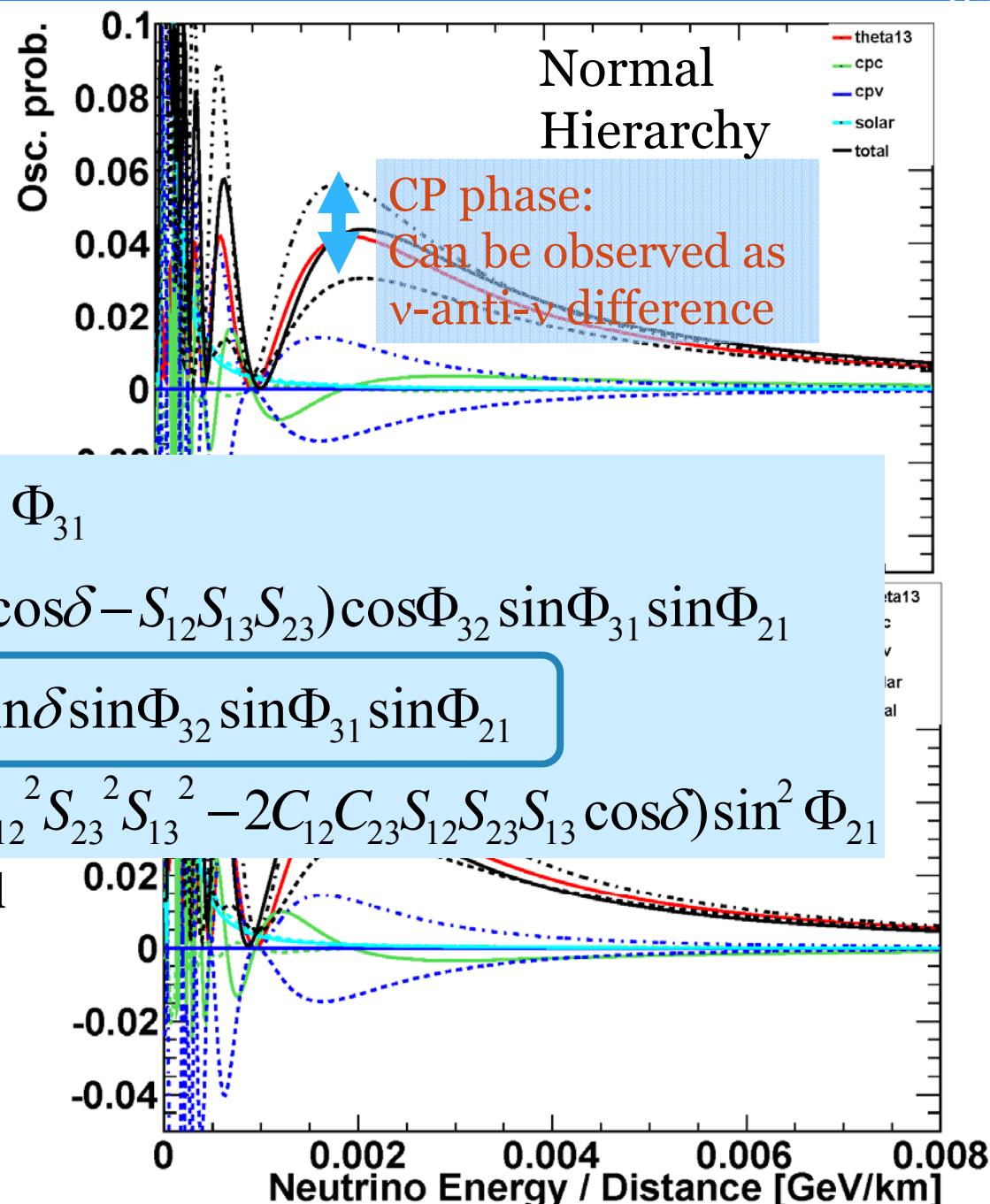


## $\nu_\mu \rightarrow \nu_e$ oscillation probability

- $\Delta m^2_{23} = 2.4 \text{e-3 eV}^2$
- $\Delta m^2_{12} = 7.59 \text{e-5 eV}^2$
- $\theta_{12} = 34^\circ$
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$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & 4C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \Phi_{31} \\
 & + 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cos \Phi_{32} \sin \Phi_{31} \sin \Phi_{21} \\
 & \boxed{-8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \Phi_{32} \sin \Phi_{31} \sin \Phi_{21}} \\
 & + 4S_{12}^2 C_{13}^2 (C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta) \sin^2 \Phi_{21}
 \end{aligned}$$

—  $\delta=0$   
 - - -  $\delta=\pi/2$   
 - · -  $\delta=-\pi/2$



## $\nu_\mu \rightarrow \nu_e$ oscillation probability

- $\Delta m^2_{23} = 2.4 \text{e-3 eV}^2$
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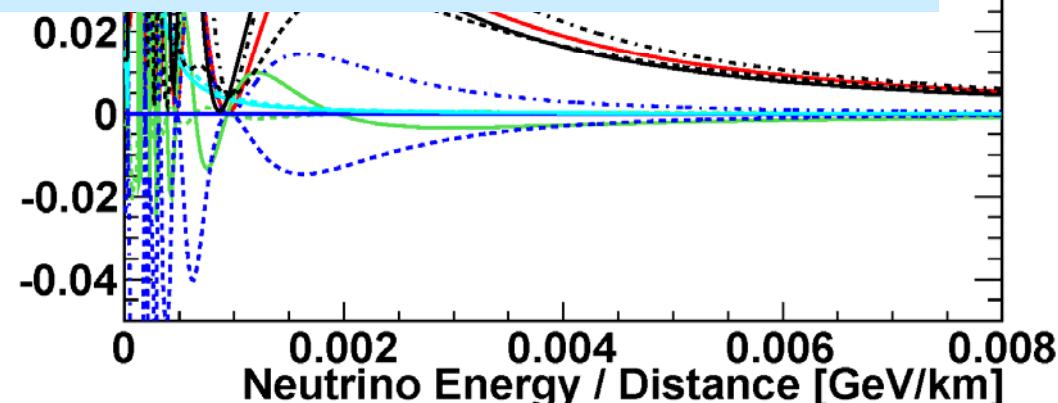
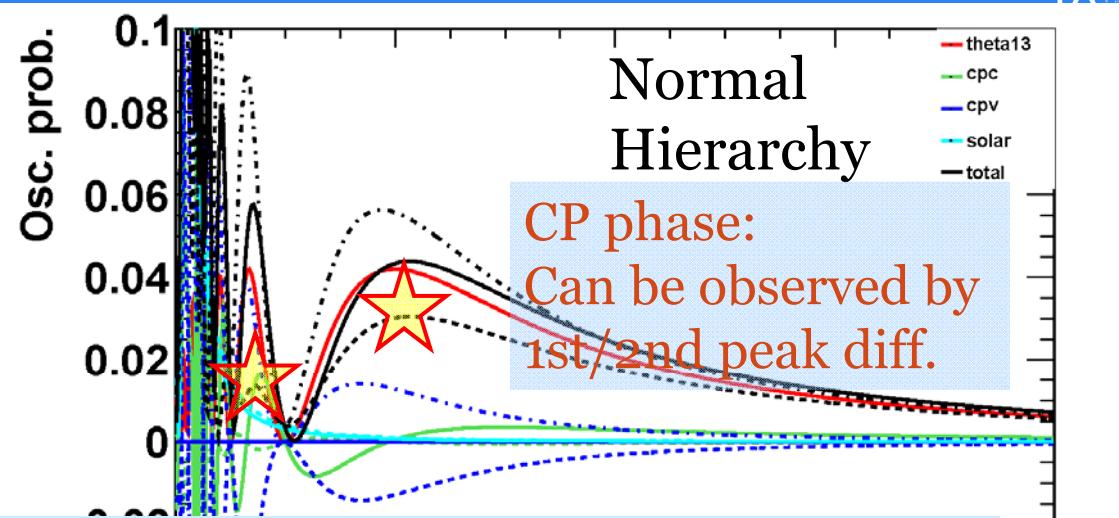
$$P(\nu_\mu \rightarrow \nu_e) = 4C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \Phi_{31}$$

$$+ 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cos \Phi_{32} \sin \Phi_{31} \sin \Phi_{21}$$

$$- 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \Phi_{32} \sin \Phi_{31} \sin \Phi_{21}$$

$$+ 4S_{12}^2 C_{13}^2 (C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta) \sin^2 \Phi_{21}$$

Total  
 —  $\delta=0$   
 - - -  $\delta=\pi/2$   
 - · -  $\delta=-\pi/2$



Q. Calculate how much  $P(\nu_\mu \rightarrow \nu_e)$  changes between  $\delta=0$  and  $\delta=90^\circ$

$$P(\nu_\mu \rightarrow \nu_e) \cong 4C_{13}^2 S_{13}^2 S_{23}^2 - 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \Phi_{21}$$

$$C_{ij} = \cos \theta_{ij}, S_{ij} = \sin \theta_{ij}, \quad \Phi_{ij} = \Delta m_{ij}^2 \frac{L}{4E_\nu}$$

$$\theta_{12} = 34^\circ (\cos \theta_{12} = 0.83, \sin \theta_{12} = 0.56)$$

$$\theta_{23} = 45^\circ$$

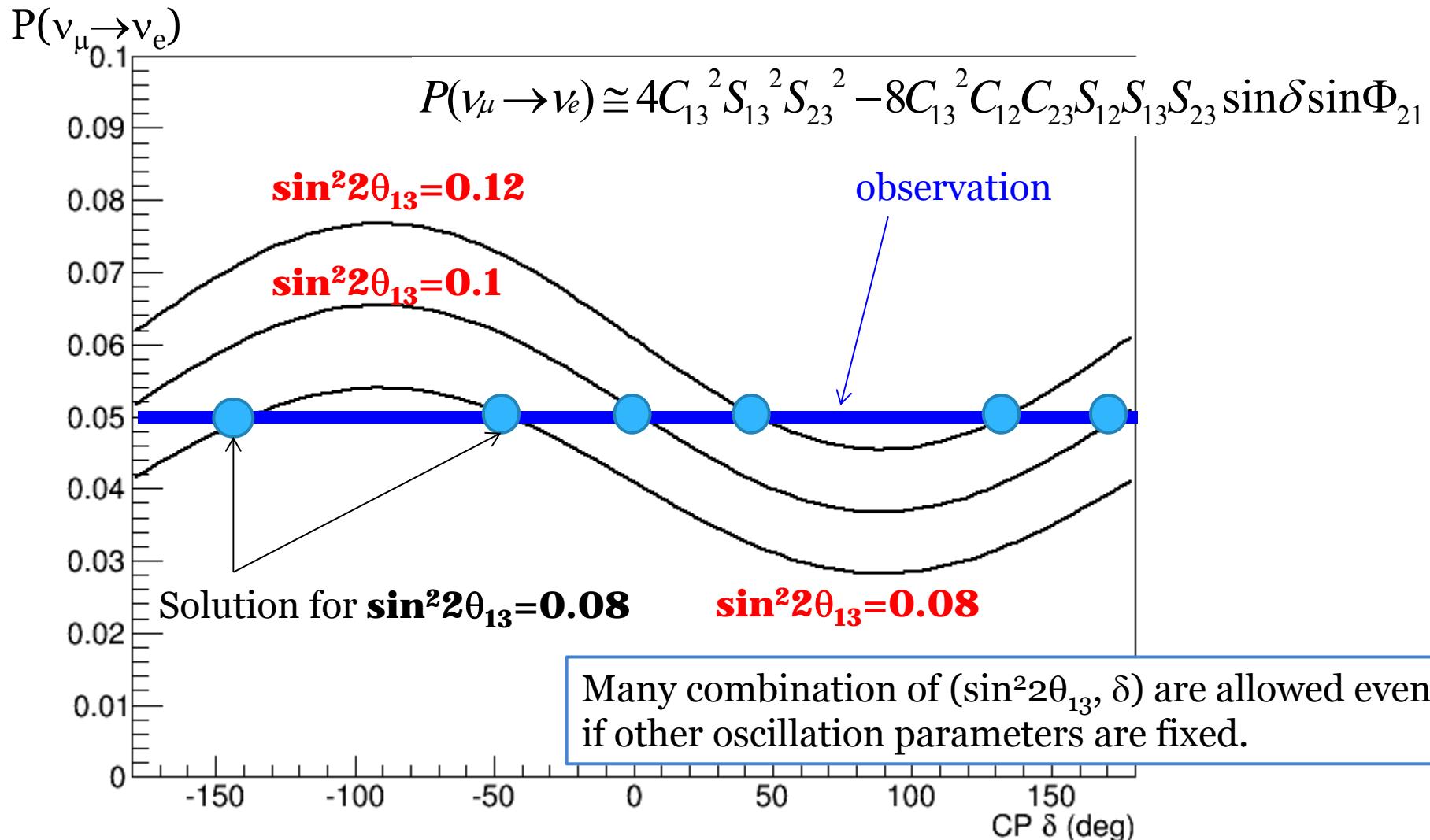
$$\theta_{13} = 9.0^\circ (\cos \theta_{13} = 0.99, \sin \theta_{13} = 0.16)$$

Max. 27% asymmetry (violation) by CP phase

# $\nu_\mu$ to $\nu_e$ oscillation probability

## at oscillation maximum

### $\sin^2 2\theta_{23} = 1$ in vacuum



# T2K Run1-4 allowed region of $\sin^2 2\theta_{13}$ for various $\delta_{CP}$ values

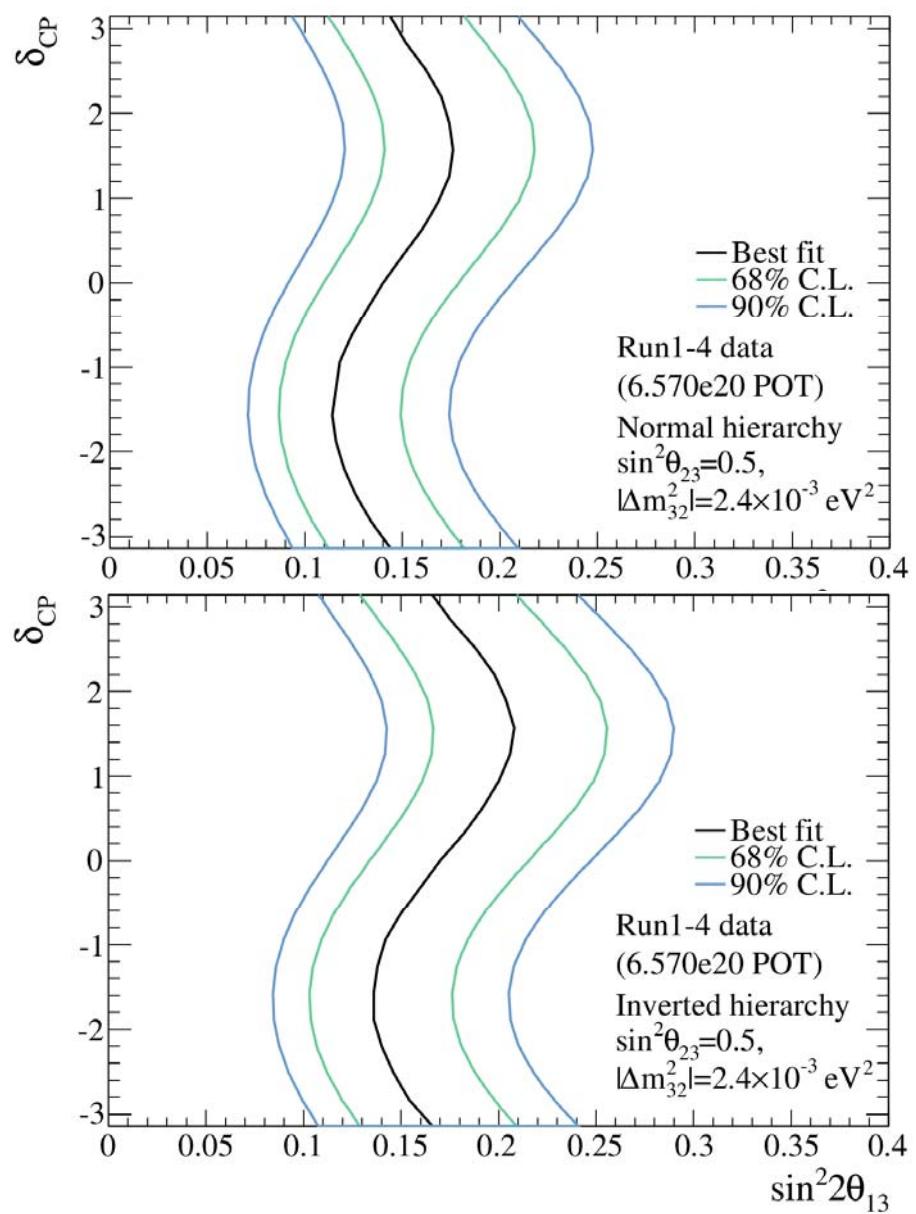
Normal Mass Hierarchy

Inverted Mass Hierarchy

**NOTE:**

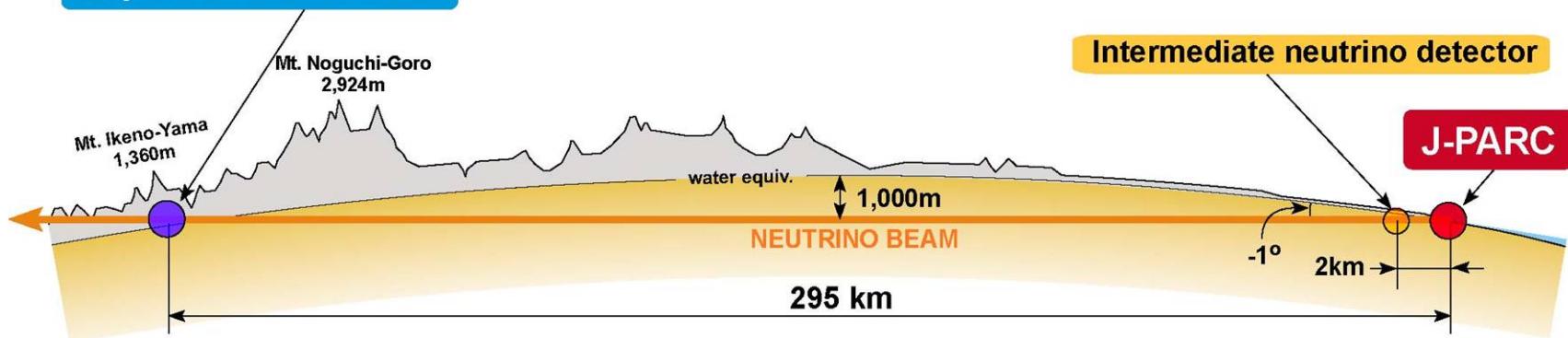
w/o reactor info. on  $\sin^2 2\theta_{13}$   
 $\sin^2 2\theta_{23}$  is fixed 1.

Why allowed region is different for different mass hierarchy?



# Earth is not symmetric about flavor nor CP

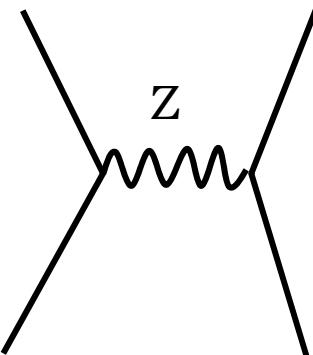
## Super-KAMIOKANDE



Interactions not changing final state shift the phase of propagation. (Work as potential)

$\nu_e$  (or  $\nu_\mu$ ,  $\nu_\tau$ )

$e^-$  or p or n



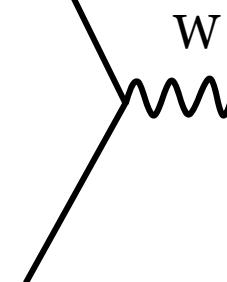
$\nu_e$  (or  $\nu_\mu$ ,  $\nu_\tau$ )

$e^-$  or p or n

NC interaction

$e^-$

$\nu_e$  (or  $\nu_\mu$ ,  $\nu_\tau$ )



$\nu_e$  (or  $\nu_\mu$ ,  $\nu_\tau$ )

$e^-$  or p or n

CC interaction

Time evolution of wave function  $\exp(-iHt)$   
 Hamiltonian in vacuum

$$-U \begin{pmatrix} p_1 & 0 & 0 \\ 0 & p_2 & 0 \\ 0 & 0 & p_3 \end{pmatrix} U^\dagger \simeq -p_1 + \frac{1}{2E} U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} U^\dagger$$

Hamiltonian of the interaction with matter

$$\begin{pmatrix} \sqrt{2}G_F n_e & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad n_e: \text{electron density}$$

(Opposite sign for } v \text{ and } \bar{v} \text{)}

The part which affect  
 the relative phase is,

$$H \approx U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \frac{\Delta m_{21}^2}{2E} & 0 \\ 0 & 0 & \frac{\Delta m_{31}^2}{2E} \end{pmatrix} U + \begin{pmatrix} \frac{a}{2E} & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

This cannot be solved analytically.

$$a \equiv 2\sqrt{2}G_F n_e E = 7.56 \times 10^{-5} \text{ eV}^2 \frac{\rho}{\text{gcm}^{-3}} \frac{E}{\text{GeV}}$$

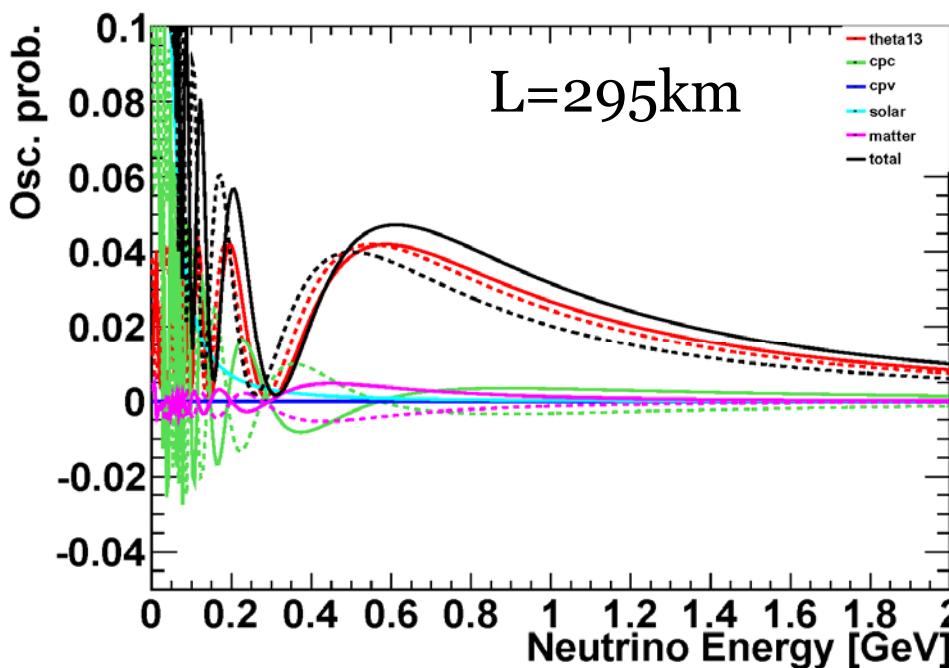
# More complete eq. of $\nu_e$ appearance (1<sup>st</sup> order for matter effect)

$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) = & 4C_{13}^2 S_{13}^2 S_{23}^2 \sin^2 \Phi_{31} \left( 1 + \frac{2a}{\Delta m_{31}^2} (1 - 2S_{13}^2) \right) && \text{Leading including matter effect} \\
 & + 8C_{13}^2 S_{12} S_{13} S_{23} (C_{12} C_{23} \cos \delta - S_{12} S_{13} S_{23}) \cos \Phi_{32} \sin \Phi_{31} \sin \Phi_{21} && \text{CP conserving} \\
 & - 8C_{13}^2 C_{12} C_{23} S_{12} S_{13} S_{23} \sin \delta \sin \Phi_{32} \sin \Phi_{31} \sin \Phi_{21} && \text{CP violating} \\
 & + 4S_{12}^2 C_{13}^2 (C_{12}^2 C_{23}^2 + S_{12}^2 S_{23}^2 S_{13}^2 - 2C_{12} C_{23} S_{12} S_{23} S_{13} \cos \delta) \sin^2 \Phi_{21} && \text{Solar} \\
 & - 8C_{13}^2 S_{13}^2 S_{23}^2 (1 - 2S_{13}^2) \frac{aL}{4E} \cos \Phi_{32} \sin \Phi_{31} && \text{Matter effect (This is small)}
 \end{aligned}$$

replace  $\delta$  by  $-\delta$  and  $a$  by  $-a$  for  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

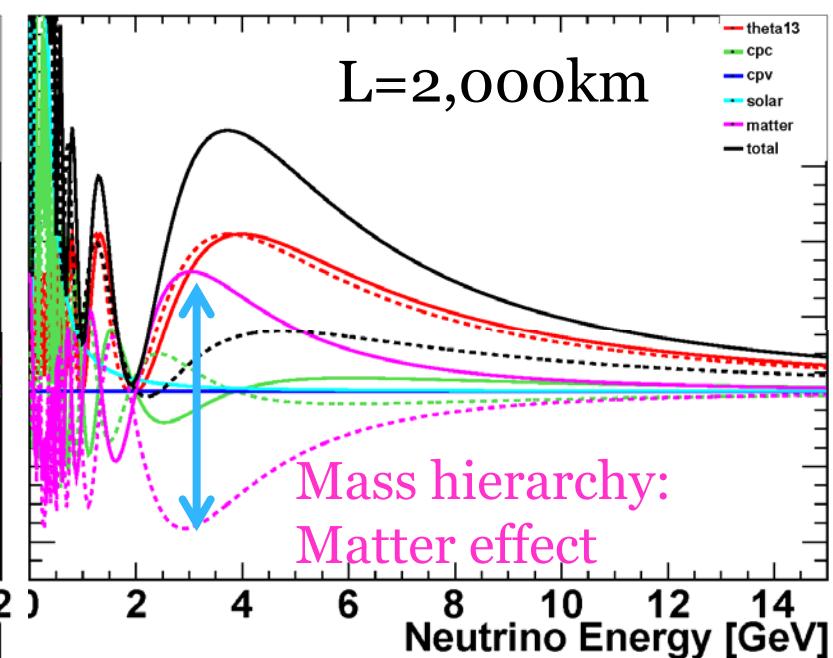
## $\nu_\mu \rightarrow \nu_e$ oscillation probability

- $\Delta m_{23}^2 = 2.4 \text{e-3 eV}^2$
- $\Delta m_{12}^2 = 7.59 \text{e-5 eV}^2$
- $\theta_{12} = 34^\circ$
- $\theta_{23} = 45^\circ$
- $\sin^2 2\theta_{13} = 0.084$
- w/ matter effect ( $\rho=2.8 \text{ g/cm}^3$ )



$\theta_{13}$   
CPC  
CPV  
Solar  
Total

— Normal hierarchy  
- - - Inverted hierarchy



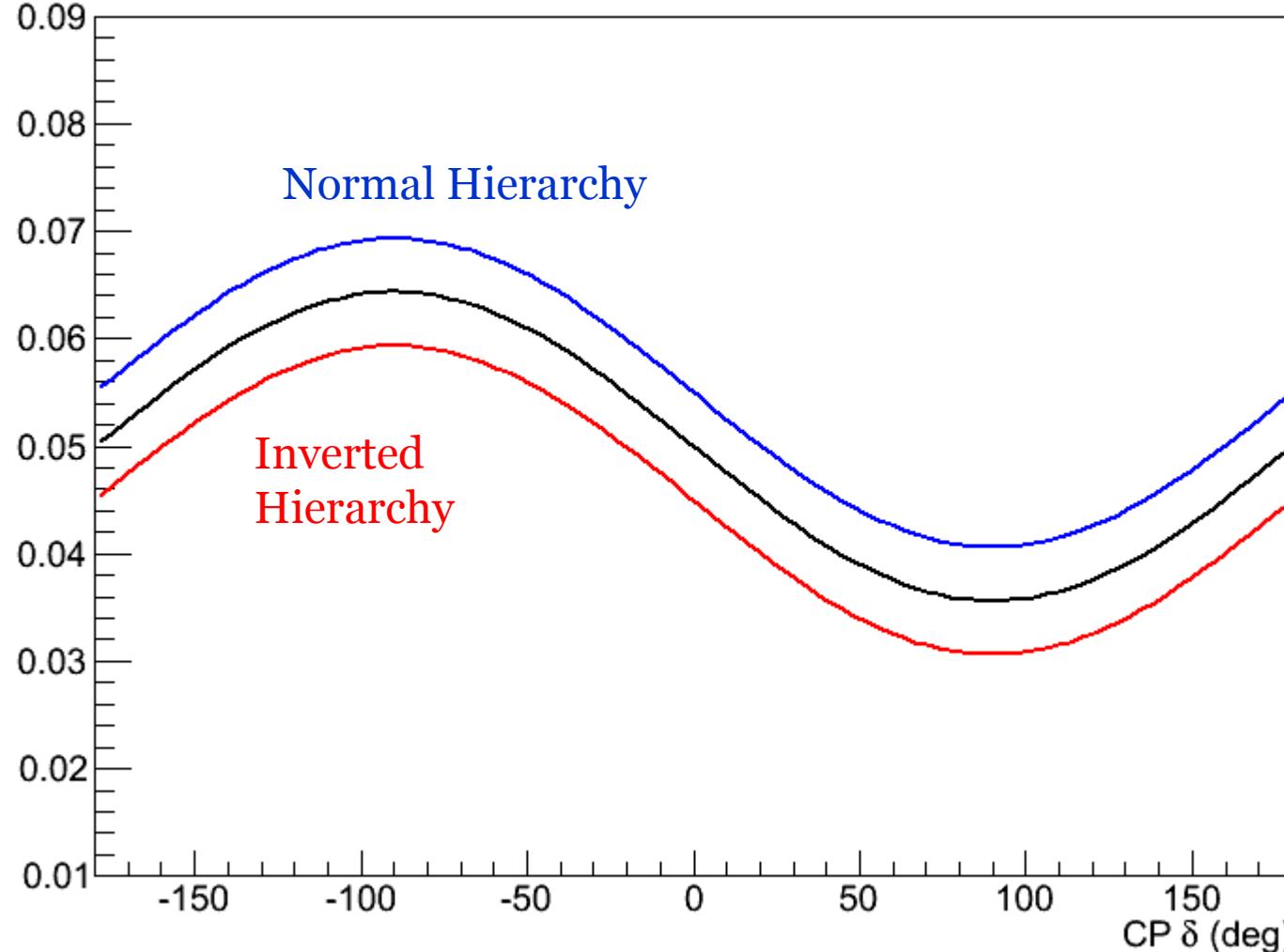
Mass hierarchy:  
Matter effect

# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

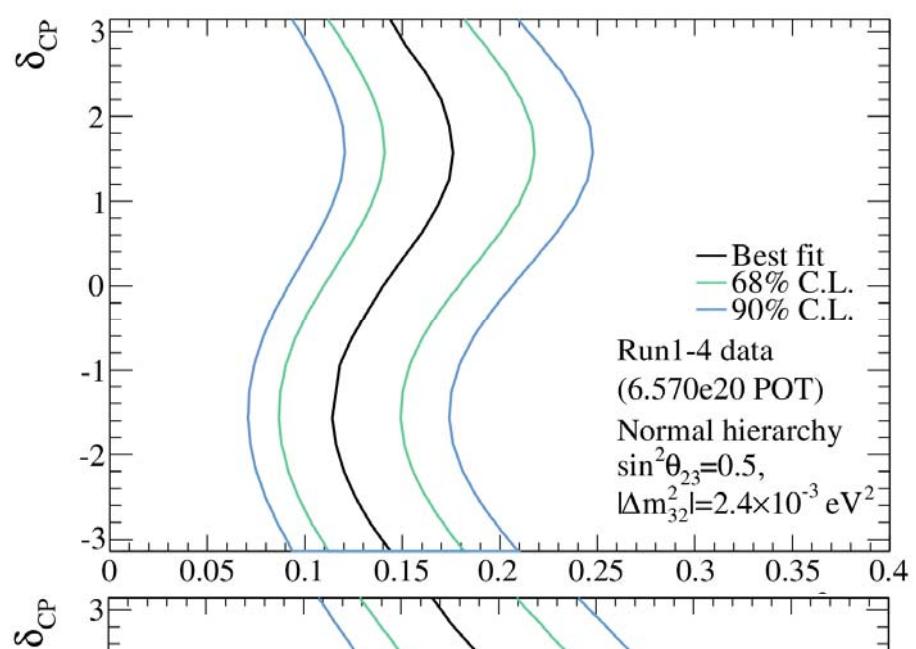
$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect

$P(\nu_\mu \rightarrow \nu_e)$

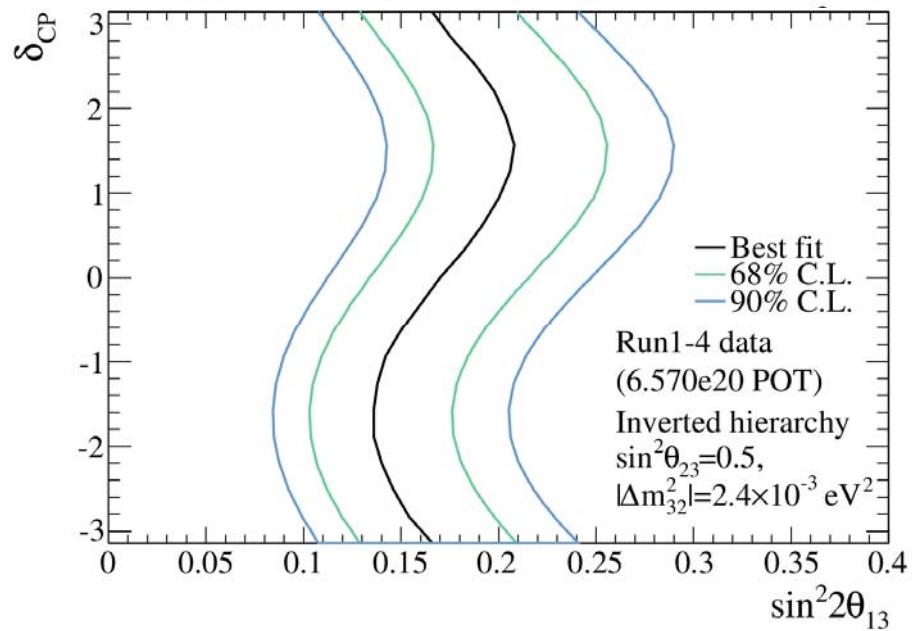


# T2K Run1-4 allowed region of $\sin^2 2\theta_{13}$ for various $\delta_{CP}$ values

Normal Mass Hierarchy

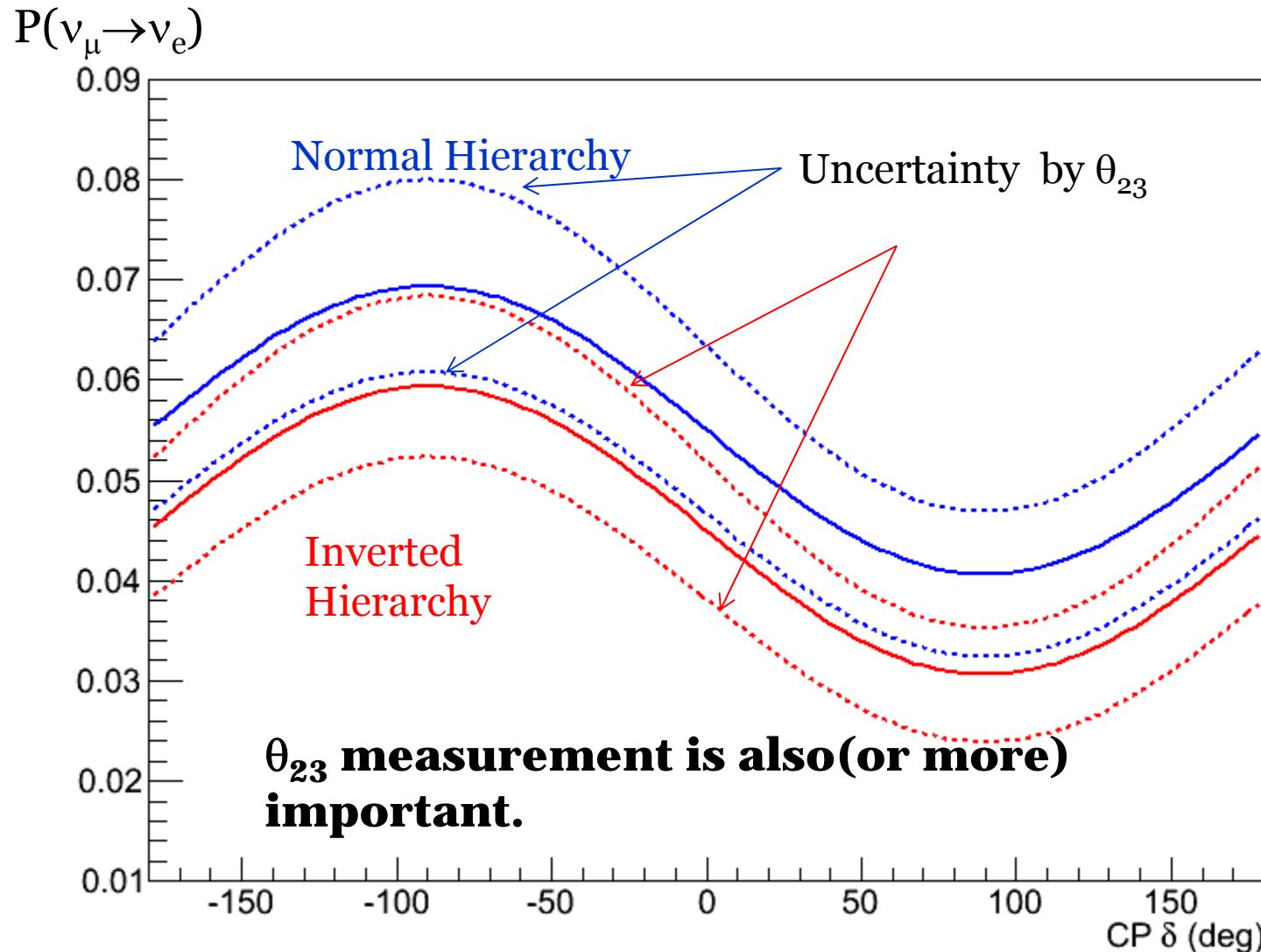


Inverted Mass Hierarchy



**NOTE:**  
w/o reactor info. on  $\sin^2 2\theta_{13}$   
 $\sin^2 2\theta_{23}$  is fixed 1.

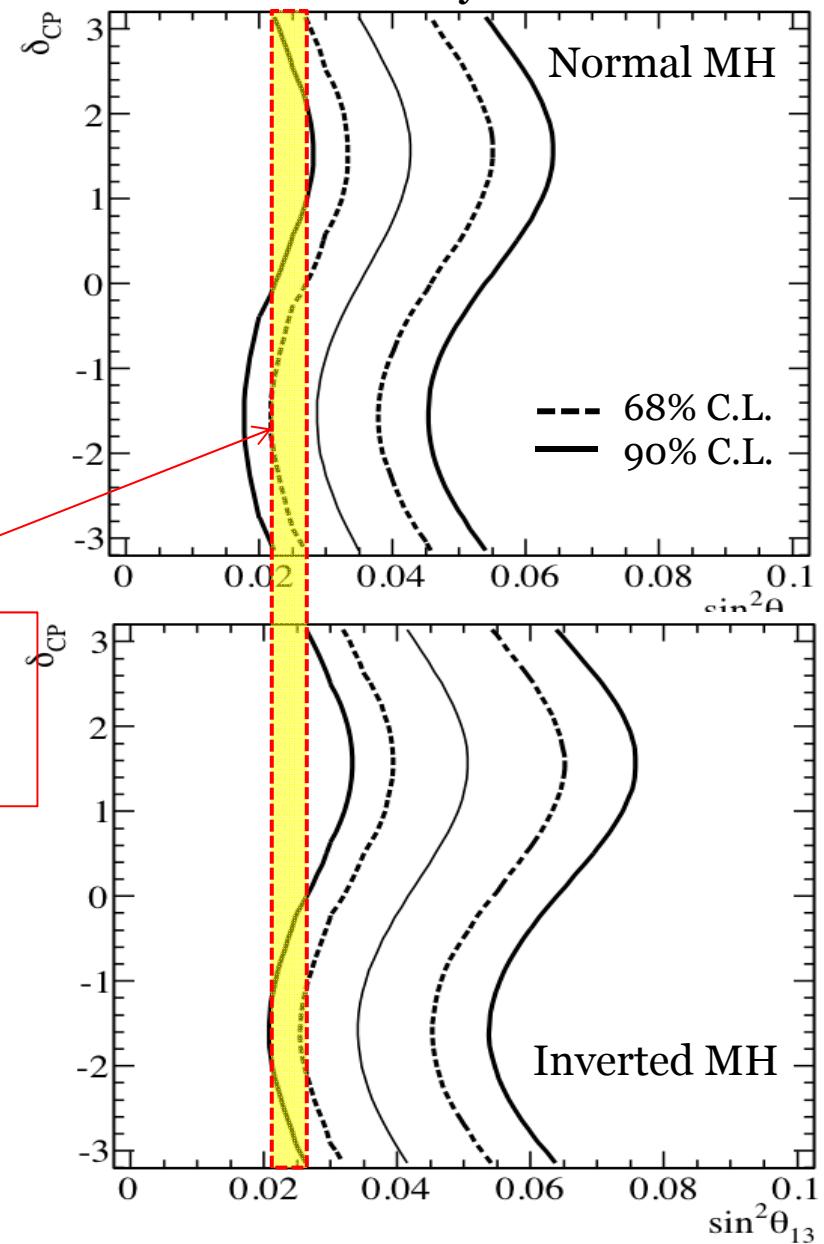
# Actually, w/ $\theta_{23}$ uncertainty



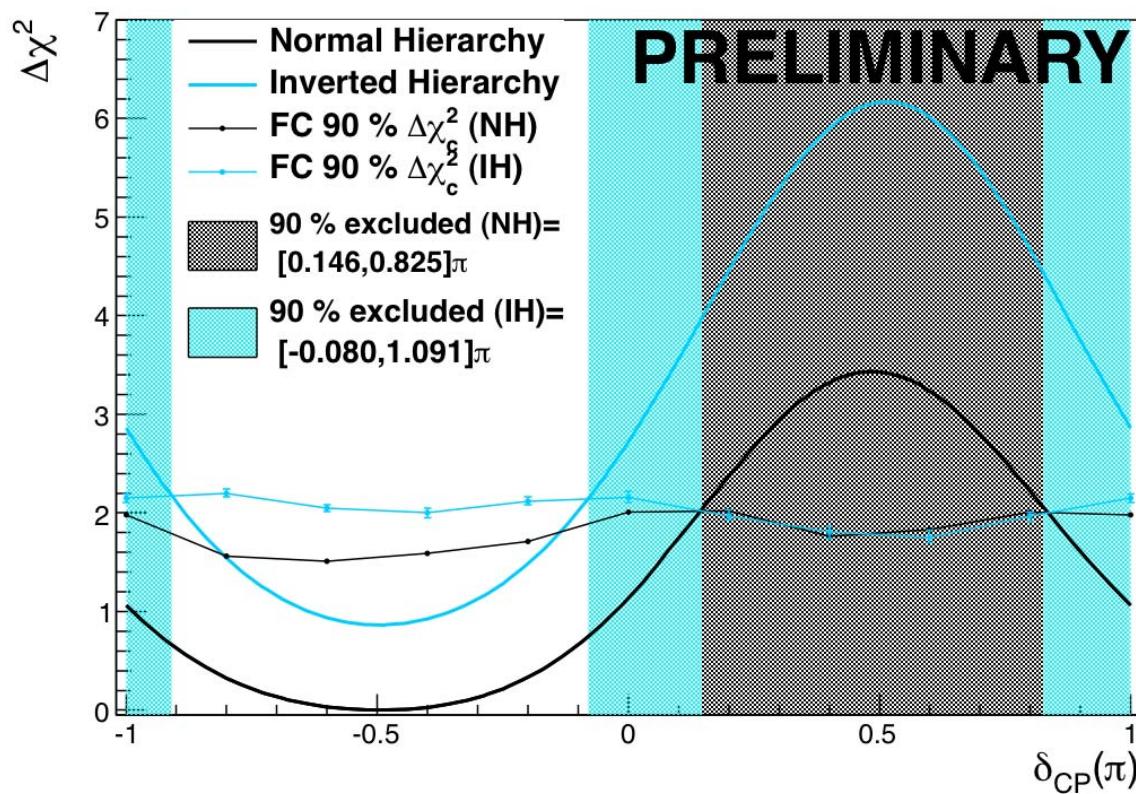
Fit  $\nu_\mu$  sample and  $\nu_e$  sample simultaneously by  $\sin^2\theta_{23}, \sin^2\theta_{13}, \delta$  and  $\Delta m^2$

Now uncertainty of  $\sin^2\theta_{23}$  is included.

68% allowed region from reactor measurement (PDG2013)



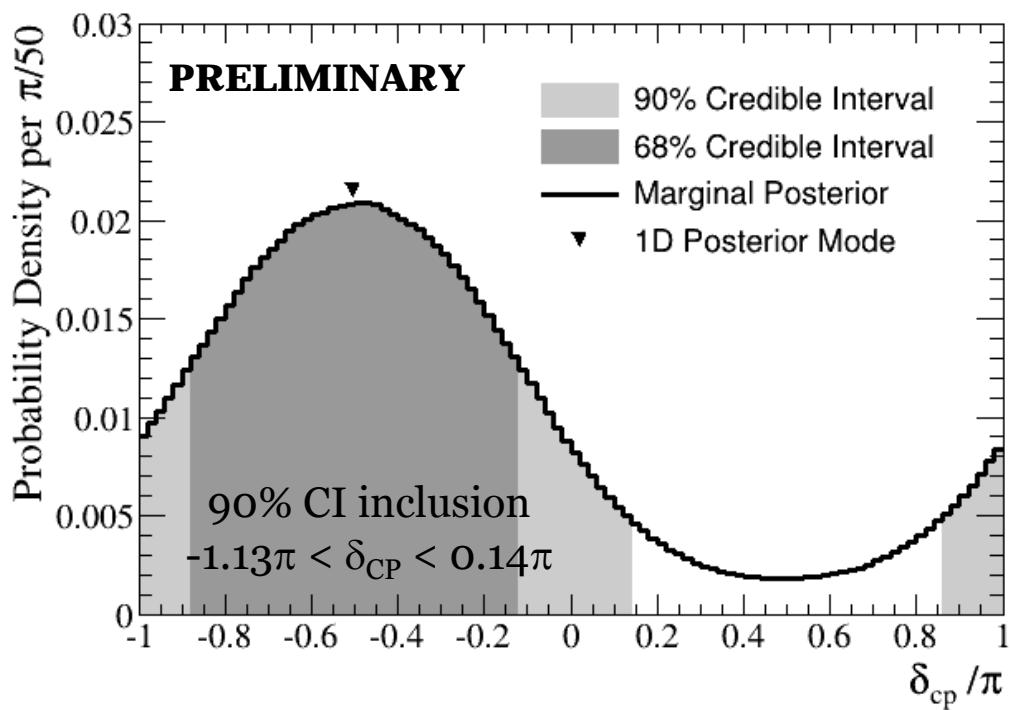
# Get allowed region for $\delta$ by combining T2K data and reactor data



# Another way of expressing result

## Bayesian posterior probability

Assume  $P(NH) = P(IH) = 0.5$



posterior probability for  $\theta_{23}$  octants and mass hierarchies

(%)	NH	IH	Sum
$\sin^2\theta_{23} \leq 0.5$	18	8	26
$\sin^2\theta_{23} > 0.5$	50	24	74
<b>Sum</b>	<b>68</b>	<b>32</b>	

PRELIMINARY

From now on,  
T2K & NOvA

# T2K full stat(7.8E21 POT) sensitivity (7.8E21 POT=750kW x 5e7sec @ 30GeV)

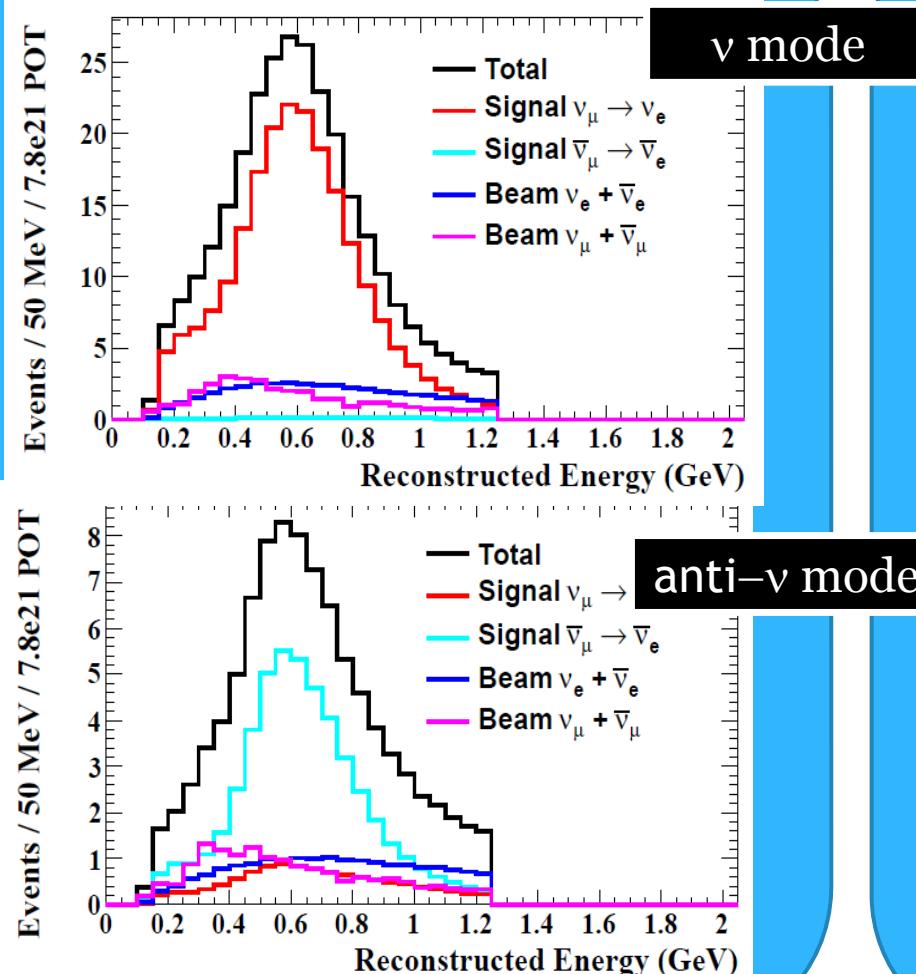
- ✓ Combined 3 Flavor Appearance and Disappearance Fit
- ✓ w/ and w/o reactor results
  - $\sin^2 2\theta_{13} = 0.1 \pm 0.005$ : marginalize by error (=Daya Bay reactor sys. error)
- ✓ w/ and w/o current systematic error
  - Assume same error for anti- $\nu$  mode. +10% overall normalization error.
- ✓  $\nu$ -mode:anti- $\nu$  mode running ratio =50%:50% for case study

# Reconstructed Energy Spectra

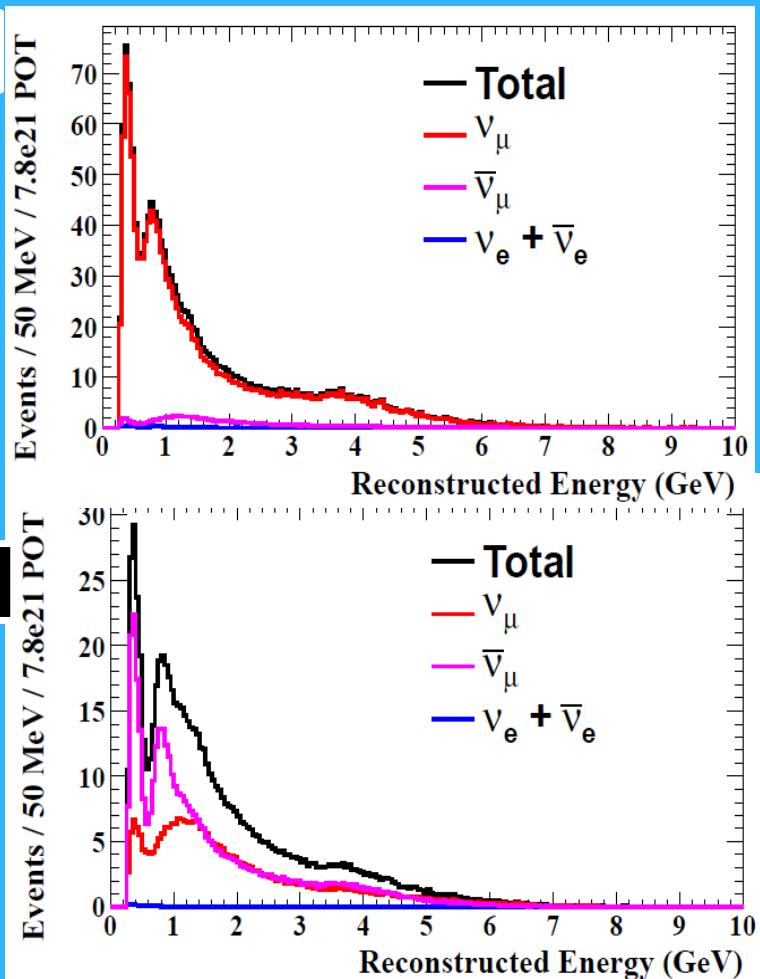
$\delta=0, \sin^2 2\theta_{23}=1$   
 $\Delta m^2=2.4 \times 10^{-7} \text{ eV}^2$

26

$\nu_e$  appearance sample



$\nu_\mu$  disappearance sample



$\nu:\text{anti-}\nu=50\%:50\%$  case

$\nu$  mode 106 signal events, 39 bkg. events

anti- $\nu$  mode 24 signal events, 22 bkg events (5.6 from  $\nu_\mu \rightarrow \nu_e$ )

# T2K full sensitivity

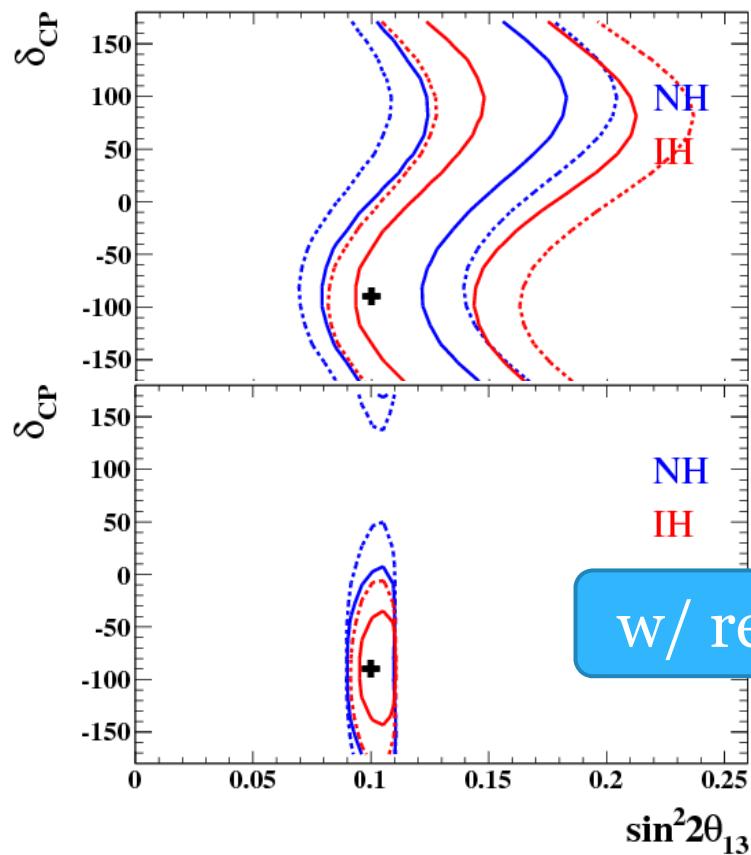
## Expected 90% C.L. allowed region

Running fraction

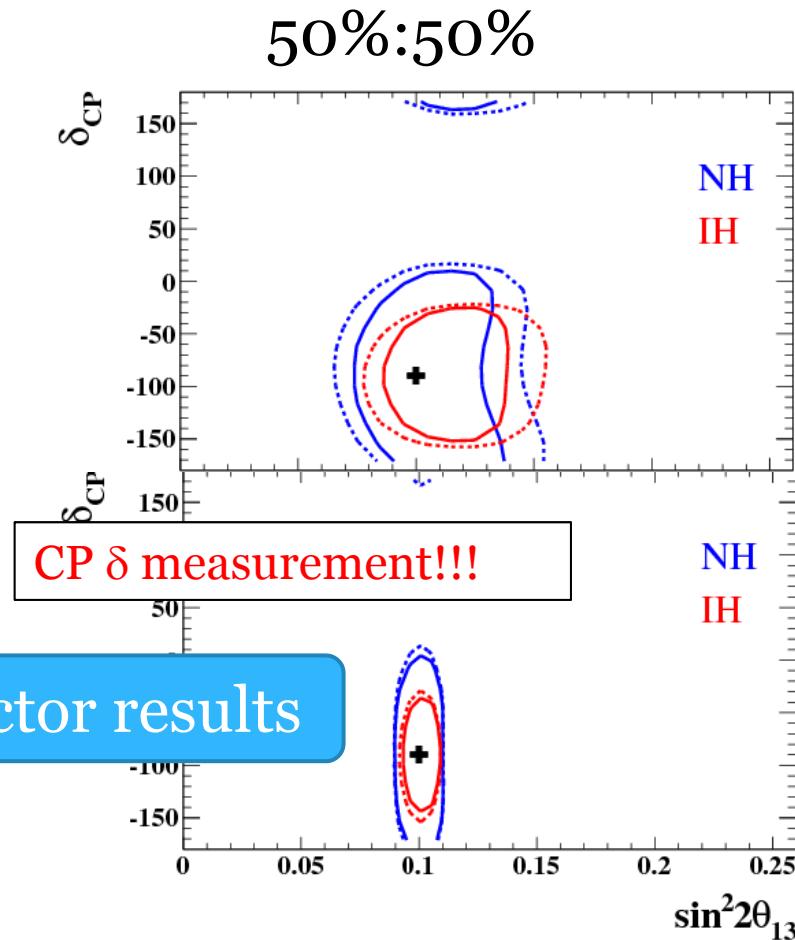
$\nu$  mode:anti- $\nu$  mode = 100%:0%

$\delta_{CP} = -90^\circ$ ,  $\sin^2 2\theta_{23} = 1.0$   
Normal Hierarchy

Allowed region assuming NH or IH  
Solid : w/o systematic error  
Dashed : w/ 2012 systematic error



w/ reactor results



CP δ measurement!!!

Allowed region assuming NH or IH  
 Solid : w/ 2012 systematic error  
 Dashed : w/o systematic error

# Expected 90% C.L. allowed region.

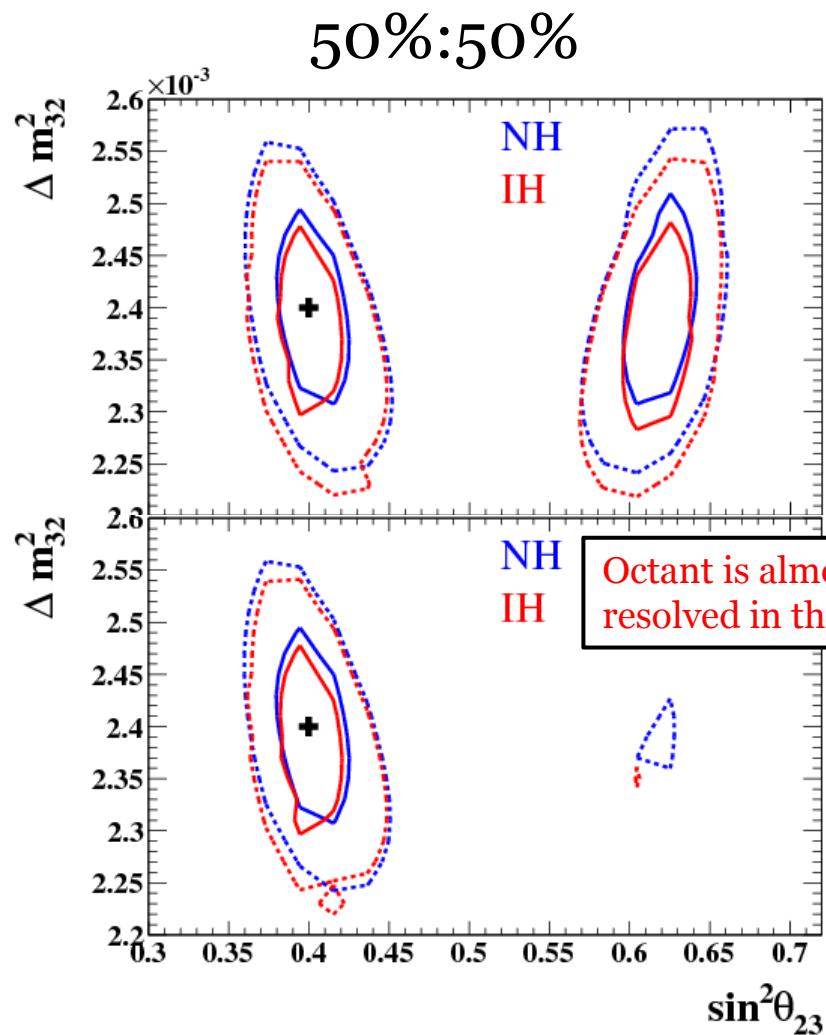
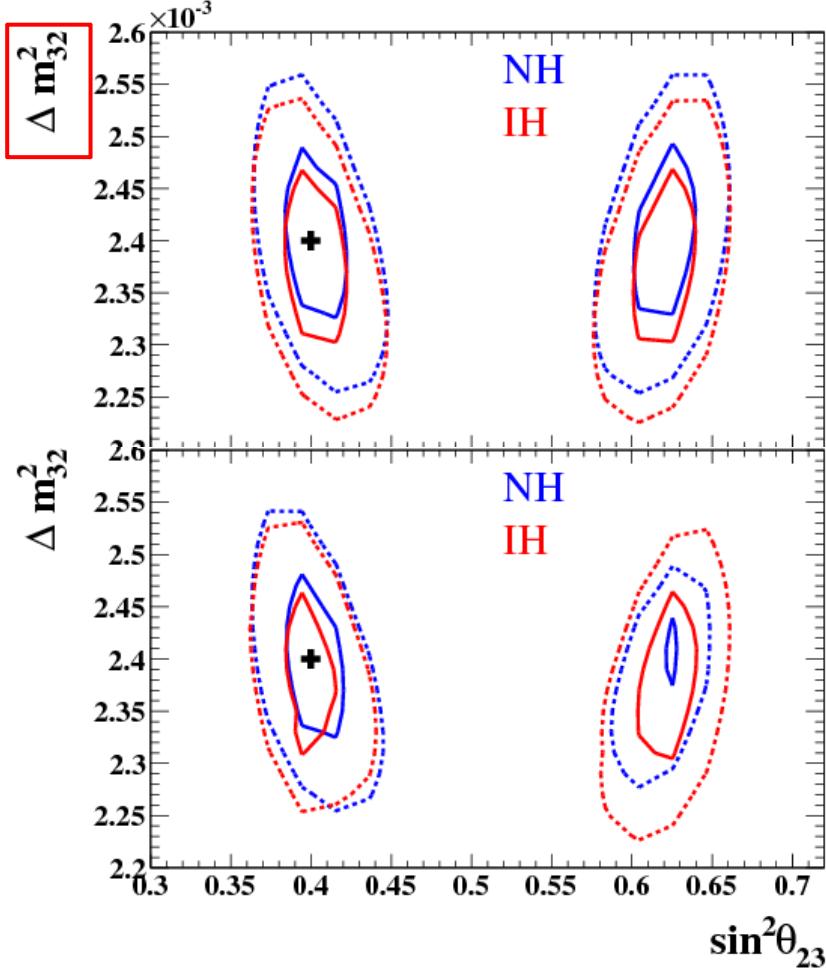
$\theta_{23}$  octant degeneracy,  $\sin^2\theta_{23}=0.4$  case

True  $\delta=0$

\* current 90% limit corresponds to  $\sin^2\theta_{23}=0.39$ .

Running fraction

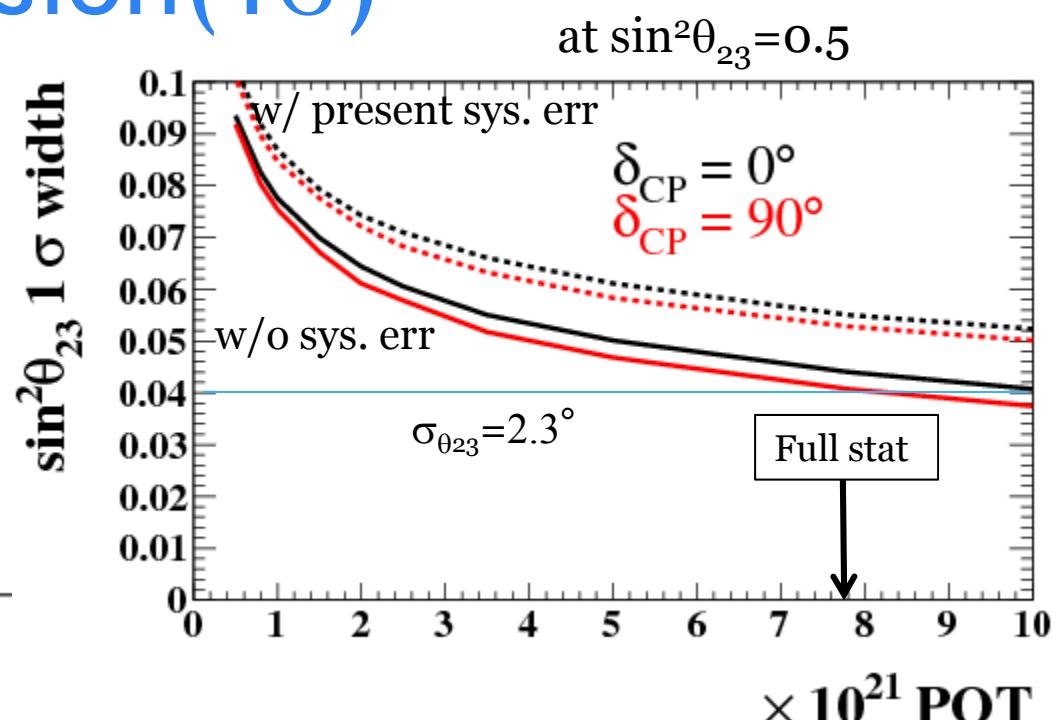
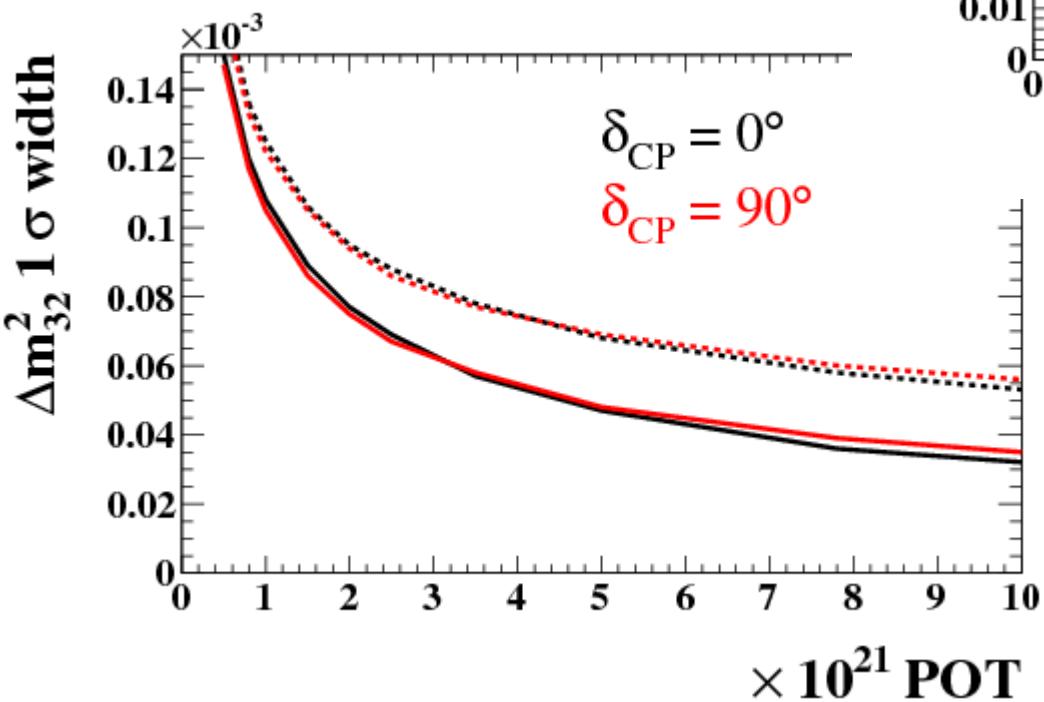
$\nu$  mode:anti- $\nu$  mode = 100%:0%



# Precision( $1\sigma$ )

MH is assumed to be  
unknown(NH)

$\nu$  mode: anti- $\nu$  mode=50%:50%  
w/ reactor result

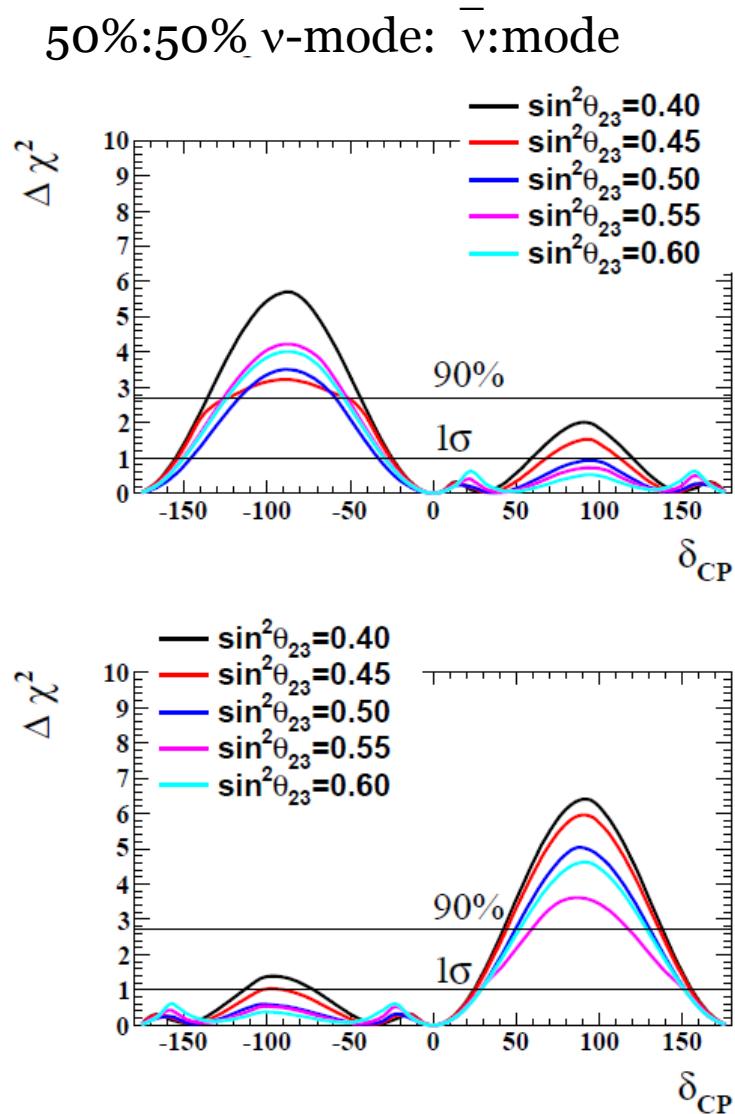
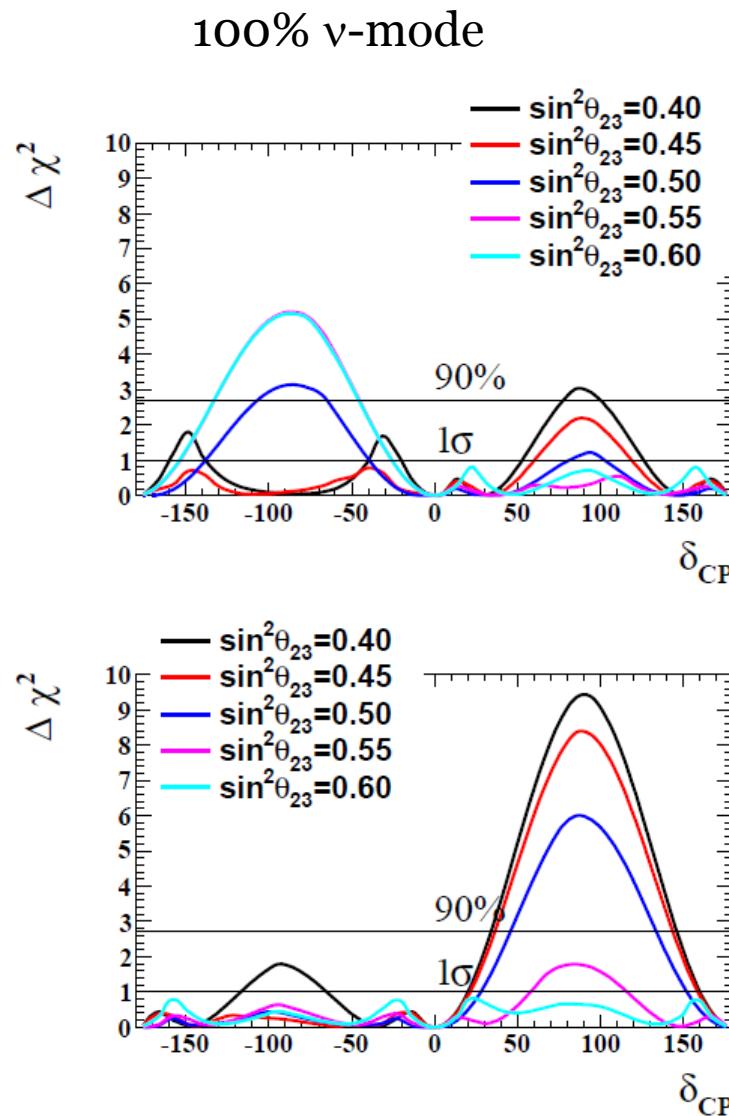


# CP violation ( $\sin\delta \neq 0$ ) sensitivity

NH

NH

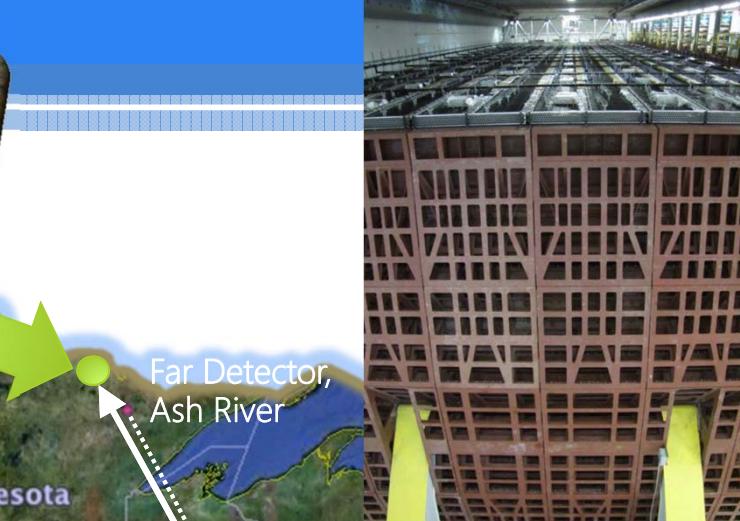
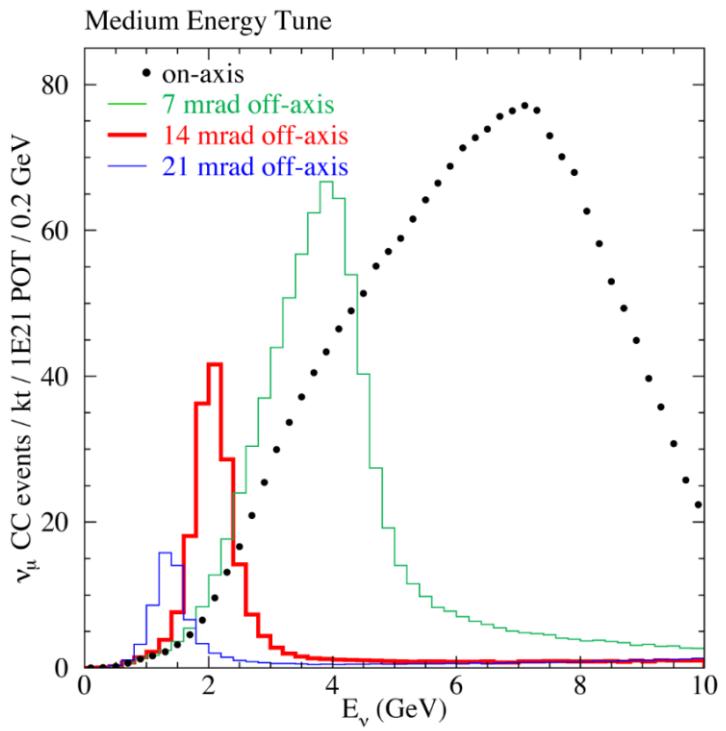
IH



# NOvA

## Ash River Laboratory

Baseline length 810km  
FNAL NuMI off-axis beam  
 $400\text{kW} \rightarrow 700\text{kW}$

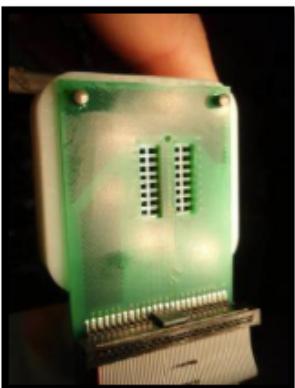


# NO $\nu$ A Detectors

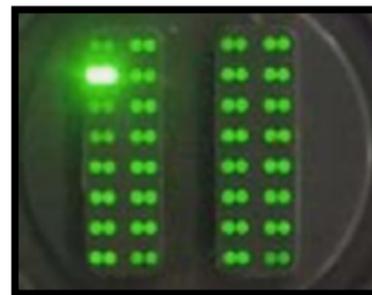
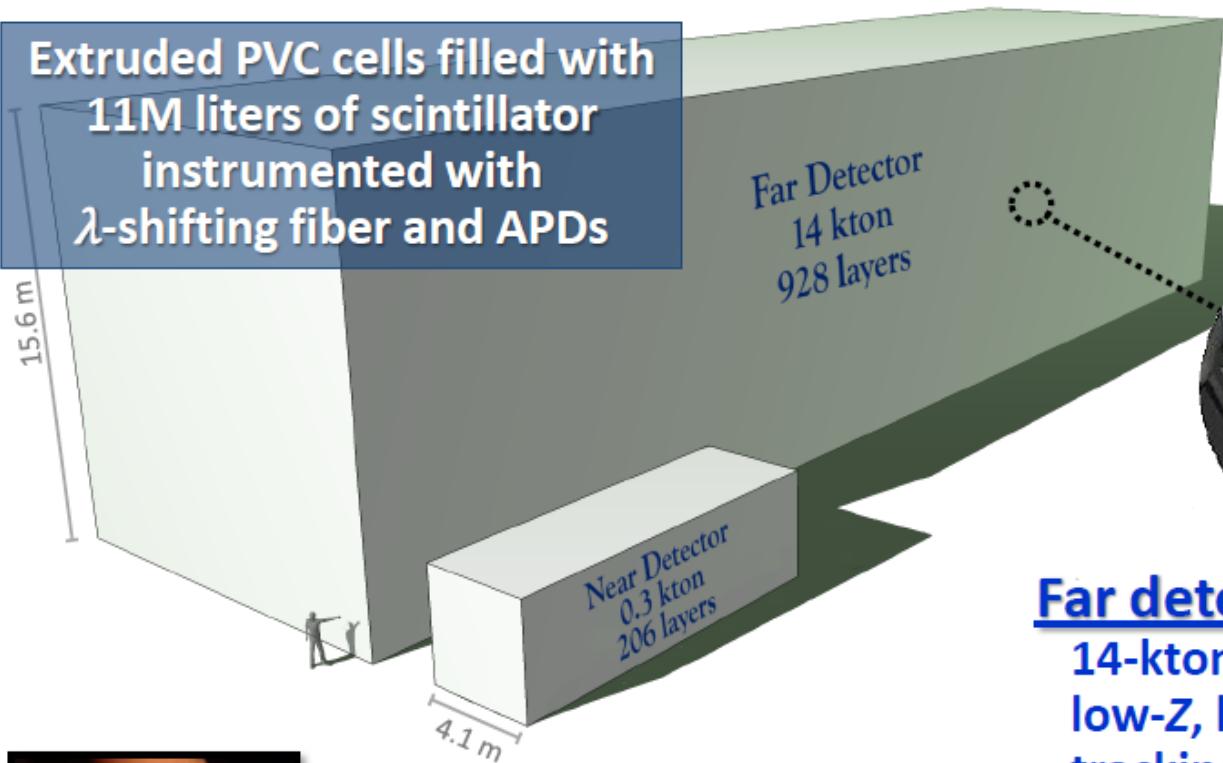
## A NO $\nu$ A cell

Extruded PVC cells filled with  
11M liters of scintillator  
instrumented with  
 $\lambda$ -shifting fiber and APDs

15.6 m

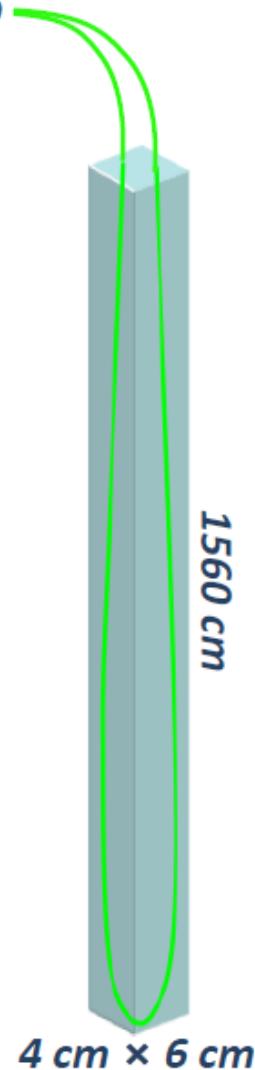


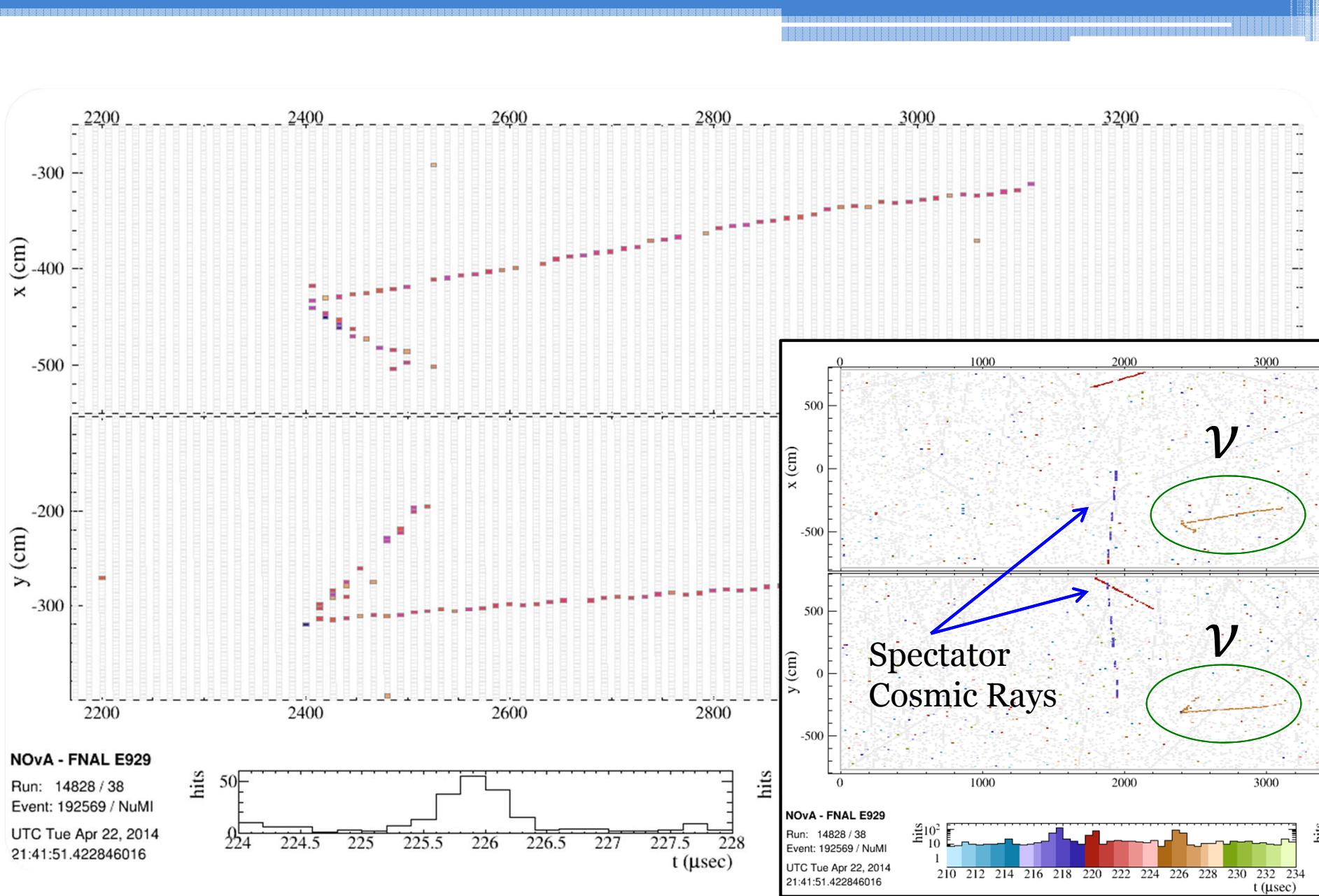
32-pixel APD  
←  
Fiber pairs  
from 32 cells →



**Far detector:**  
14-kton, fine-grained,  
low-Z, highly-active  
tracking calorimeter  
→ 360,000 channels  
→ 77% active by mass

**Near detector:**  
0.3-kton version of  
the same  
→ 18,000 channels





# $\nu_e$ Signal and Background Estimates

Cut	Simulation				Data	All Background
	$\nu_e$ Signal	NC	Beam $\nu_e$	$\nu_\mu$ CC		
All Events	36.7	380	28.1	557	19M	19M
Pre-selection	24.7	83.5	2.9	30.0	56k	56k
Vertex Gap	24.6	81.8	2.9	29.6	55k	55k
$P_T/P$	22.0	59.6	2.6	24.3	1248	1334
Maximum Y	21.2	57.4	2.5	23.0	834	917
Neutral Net	<b>13.9</b>	3.9	1.5	0.7	0.5	<b>6.5</b>
Library Template	<b>14.0</b>	3.5	1.5	1.1	0.9	<b>7.0</b>

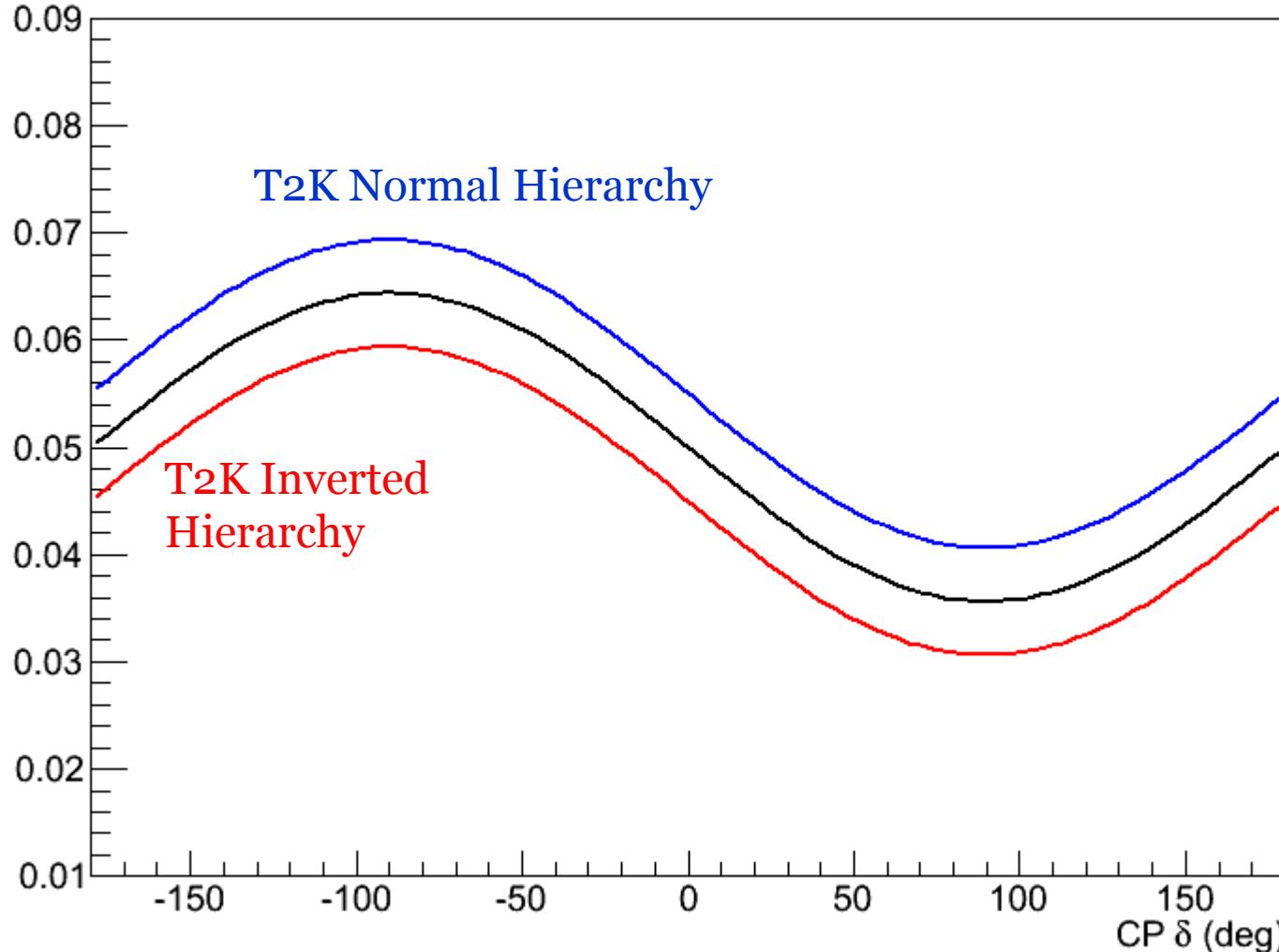
- Exposure  $6 \times 10^{20}$  POT
- 1 yr at design mass/beam power
- 14 kt total detector mass
- Signal estimates are leading order
- Simple oscillation w/o matter effect
- Averaged over hierarchy and  $\delta_{CP}$

# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect

$$P(\nu_\mu \rightarrow \nu_e)$$

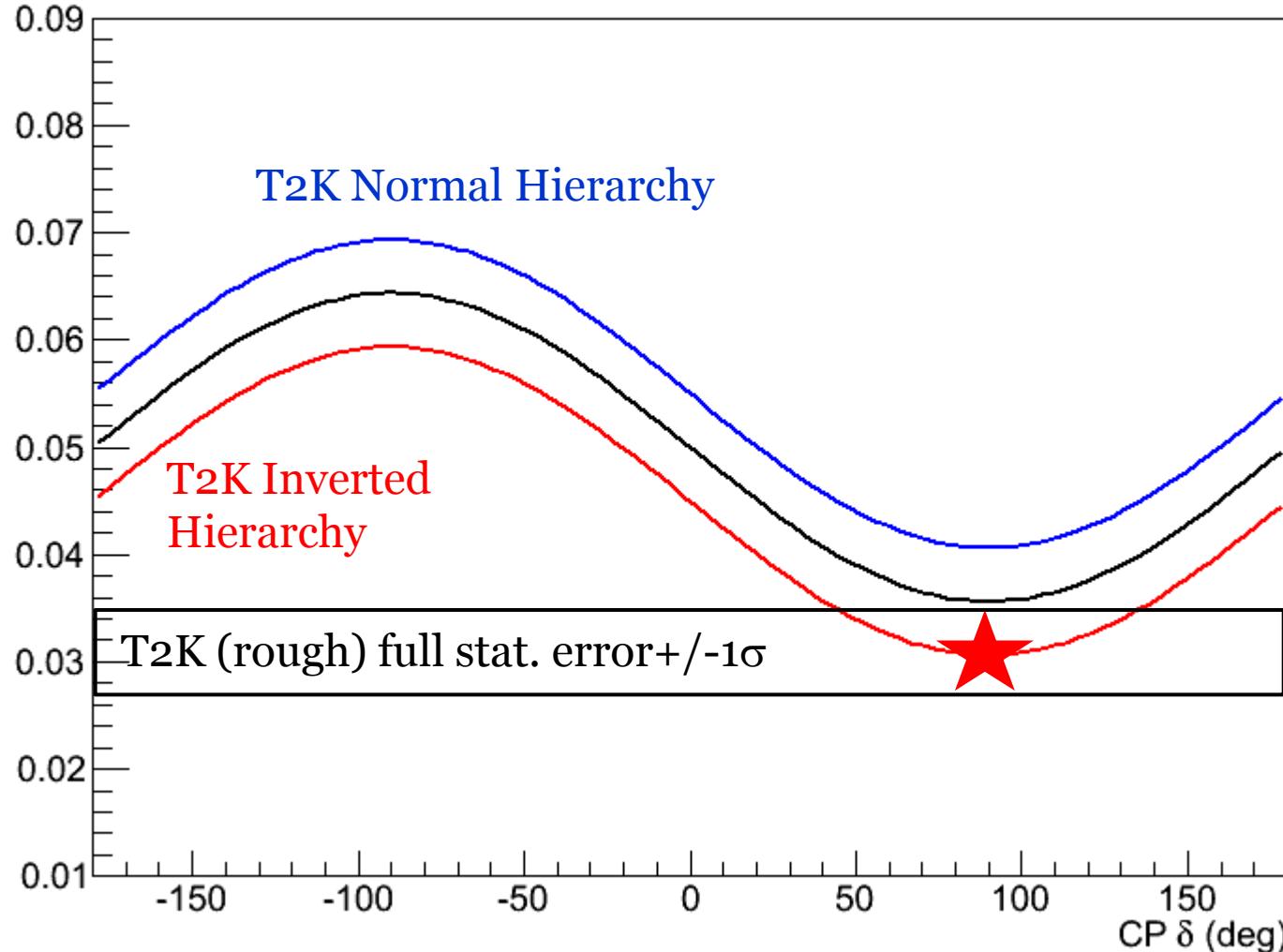


# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect

$$P(\nu_\mu \rightarrow \nu_e)$$

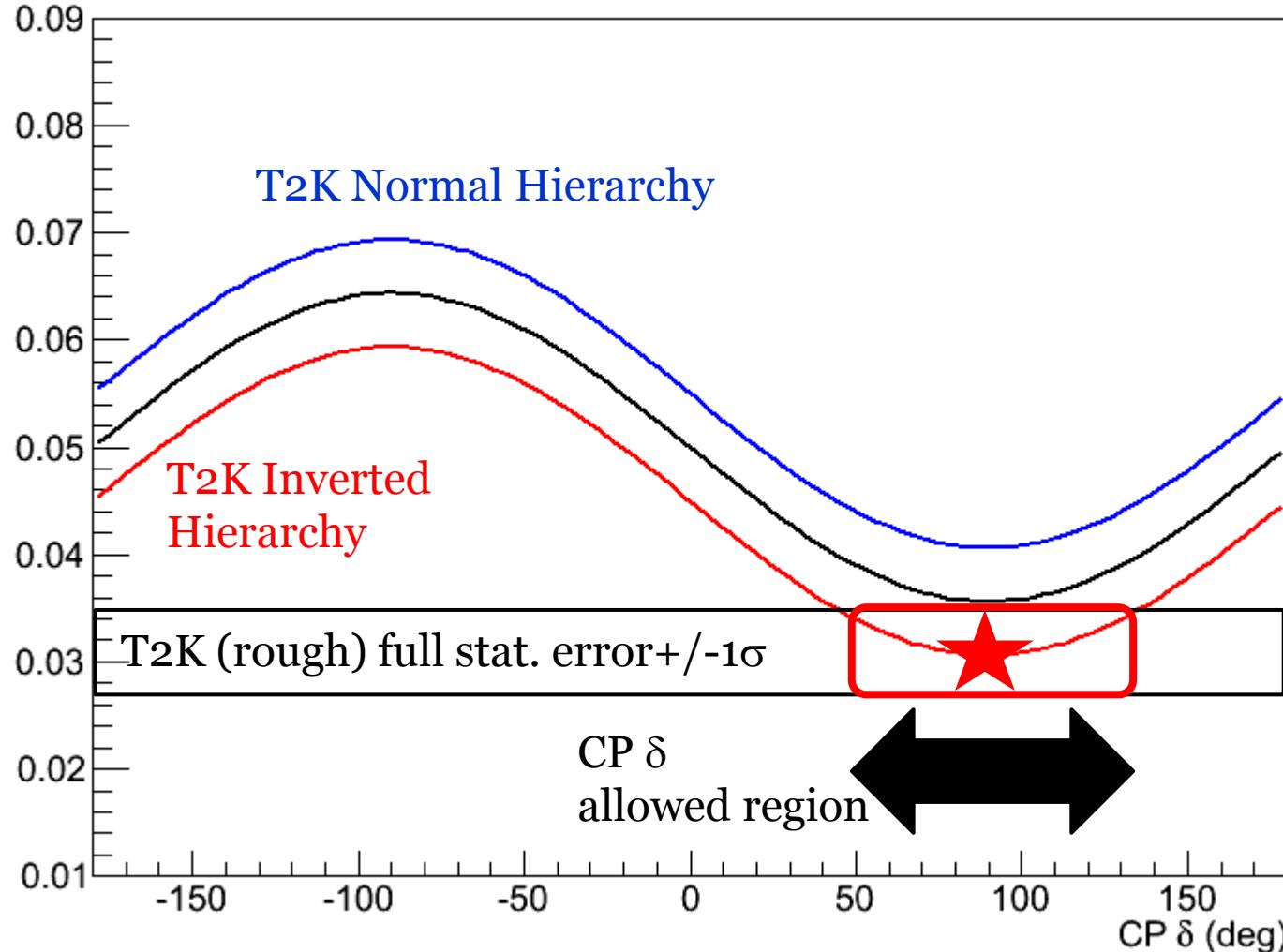


# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect

$$P(\nu_\mu \rightarrow \nu_e)$$

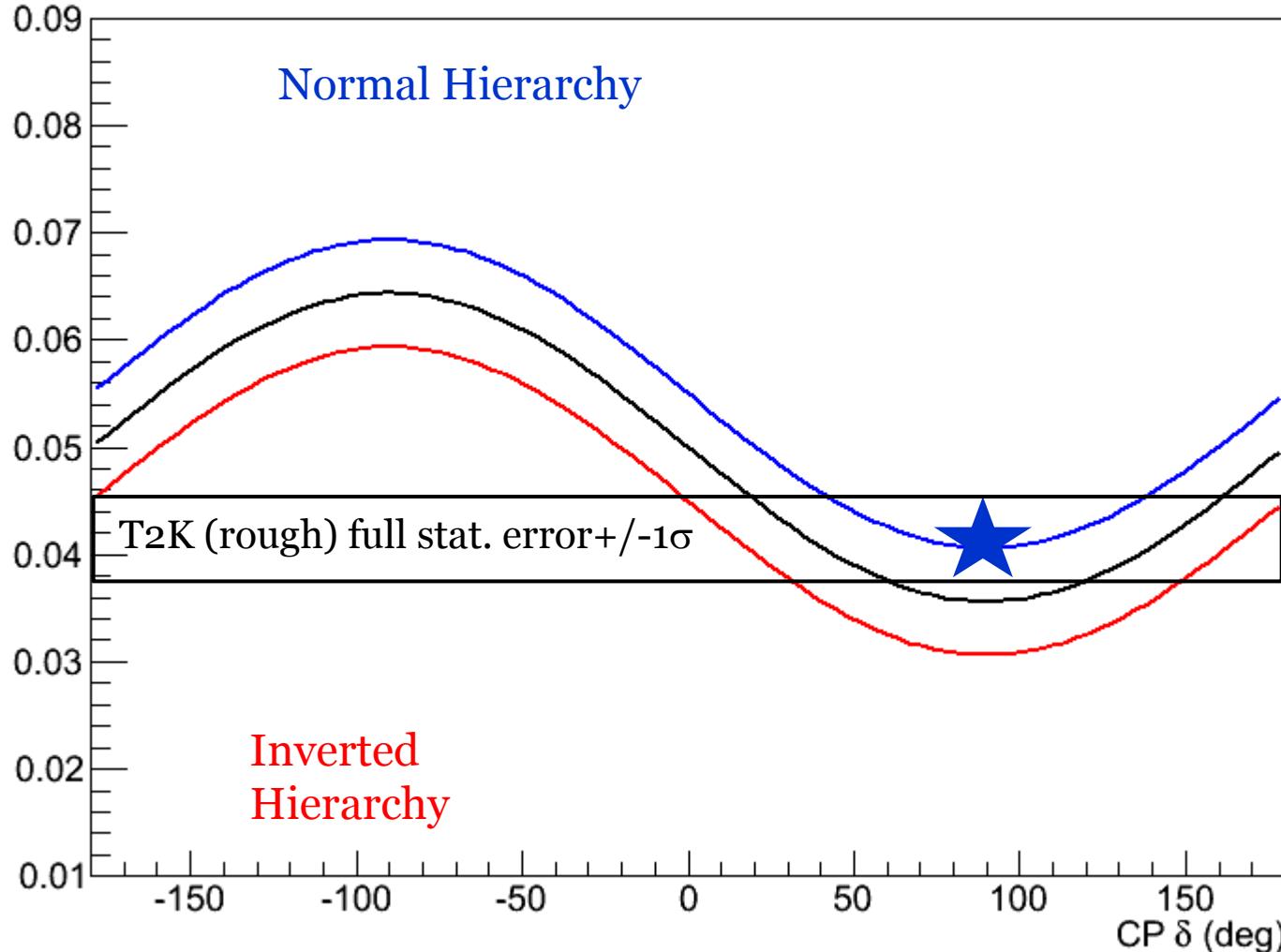


# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect

$P(\nu_\mu \rightarrow \nu_e)$

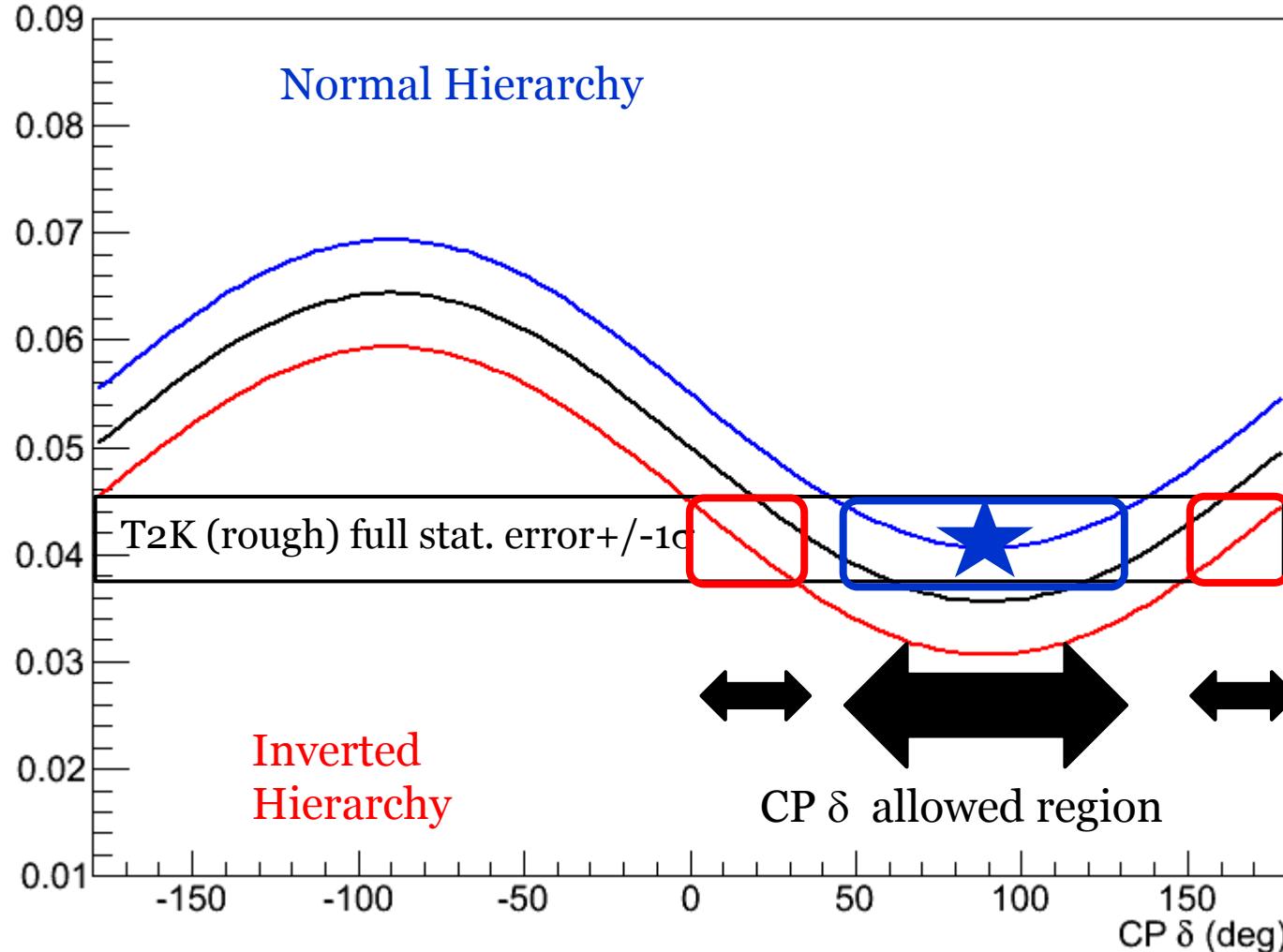


# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect

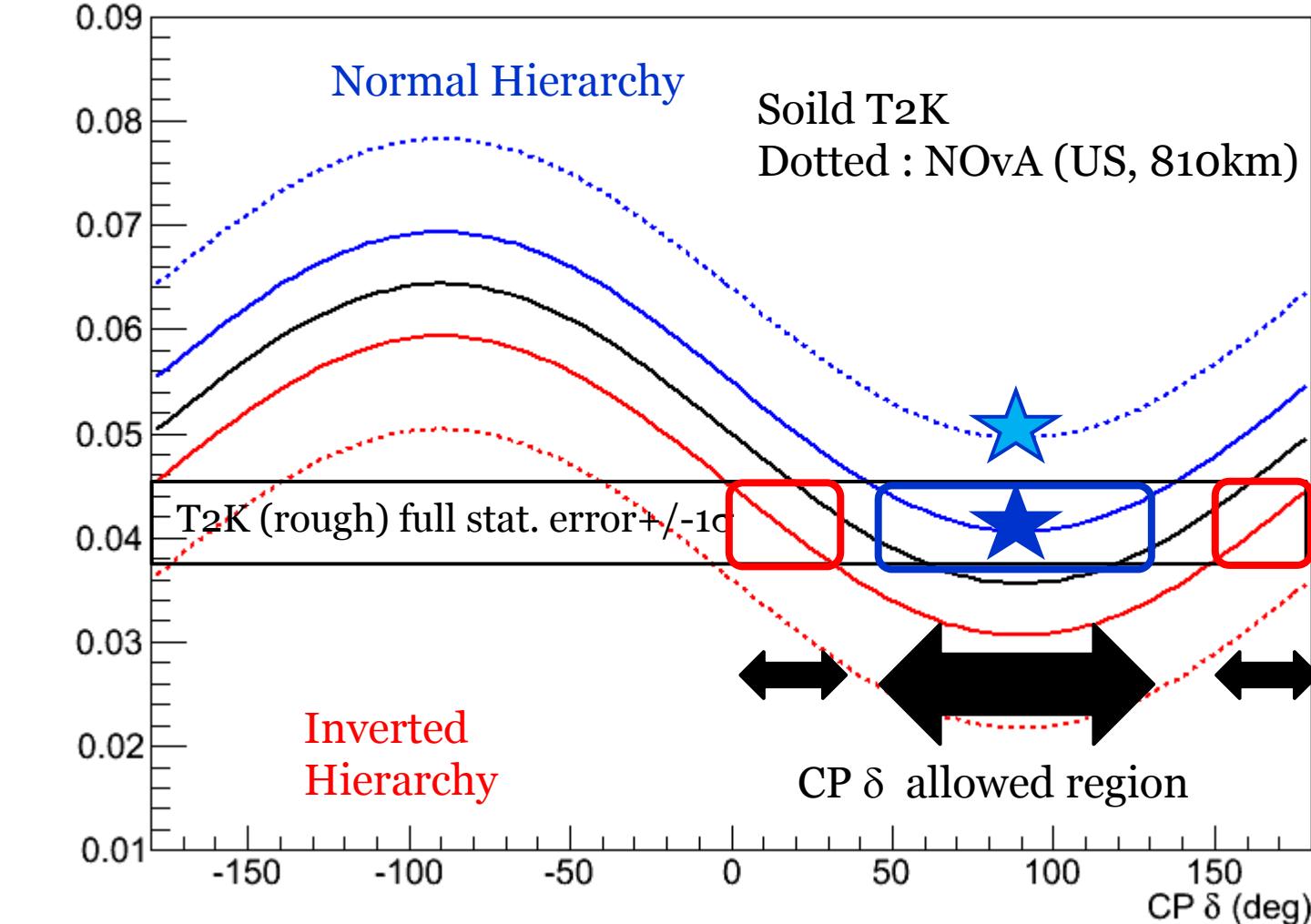
$$P(\nu_\mu \rightarrow \nu_e)$$



# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect  
 $P(\nu_\mu \rightarrow \nu_e)$

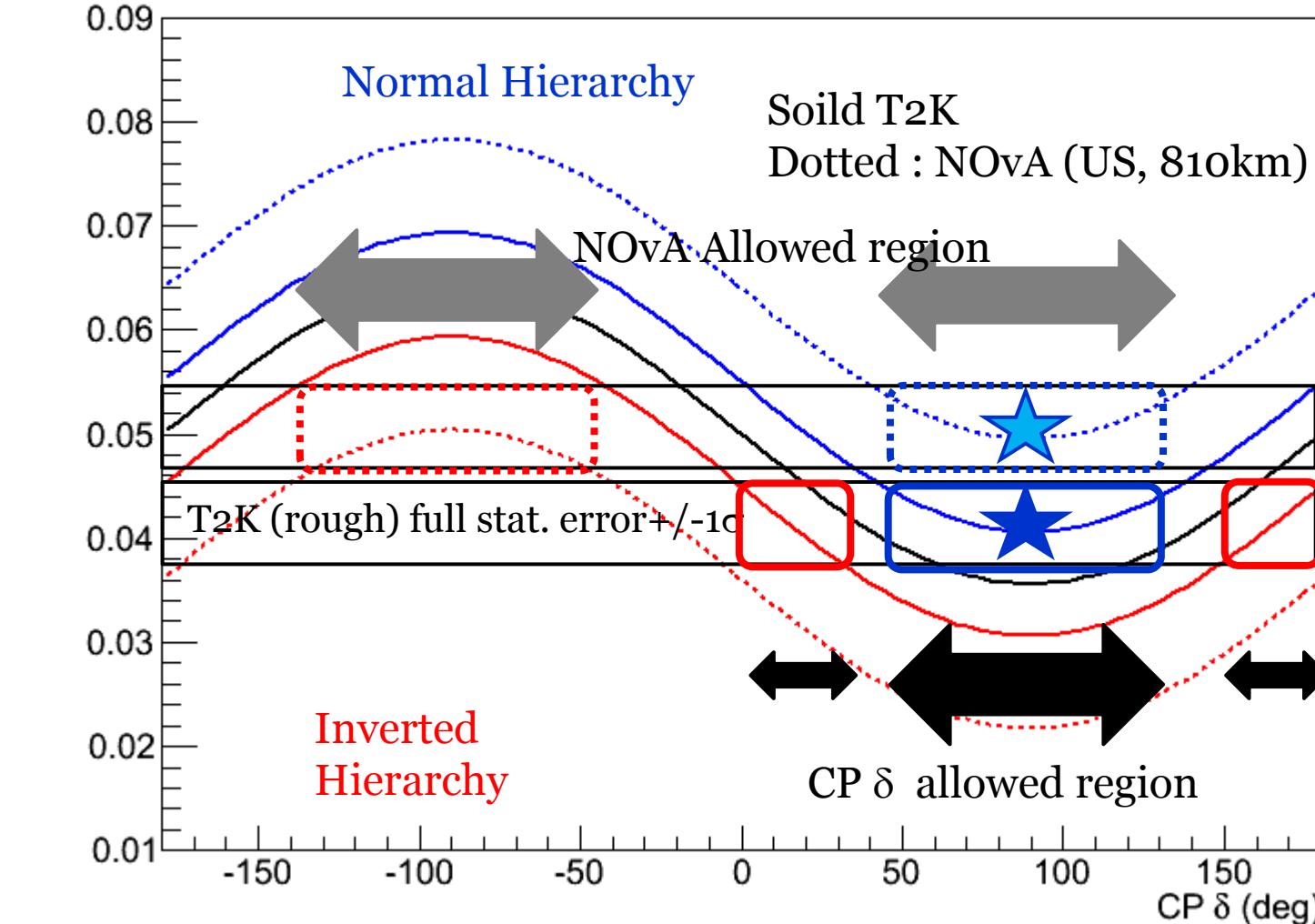


# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect

$P(\nu_\mu \rightarrow \nu_e)$

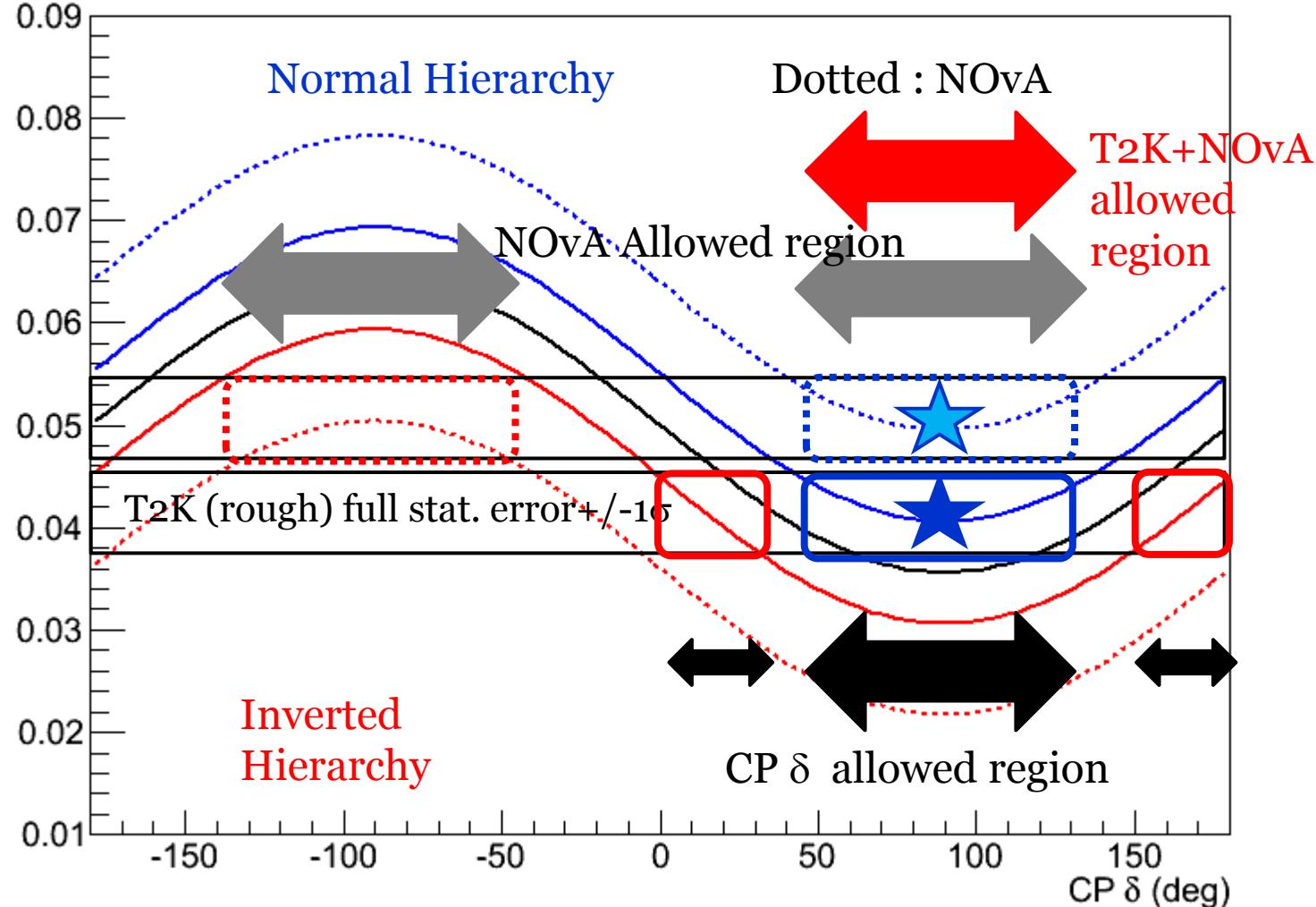


# $\nu_\mu$ to $\nu_e$ oscillation probability

at oscillation maximum

$\sin^2 2\theta_{13} = 0.1$ ,  $\sin^2 2\theta_{23} = 1$ , w/ matter effect

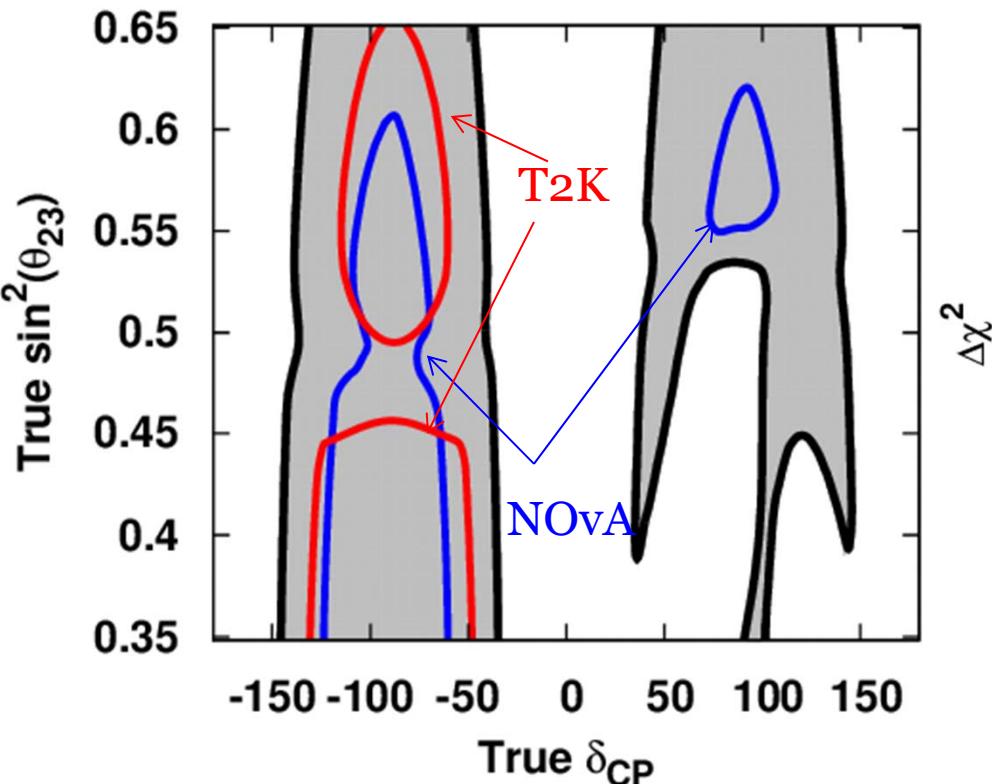
$P(\nu_\mu \rightarrow \nu_e)$



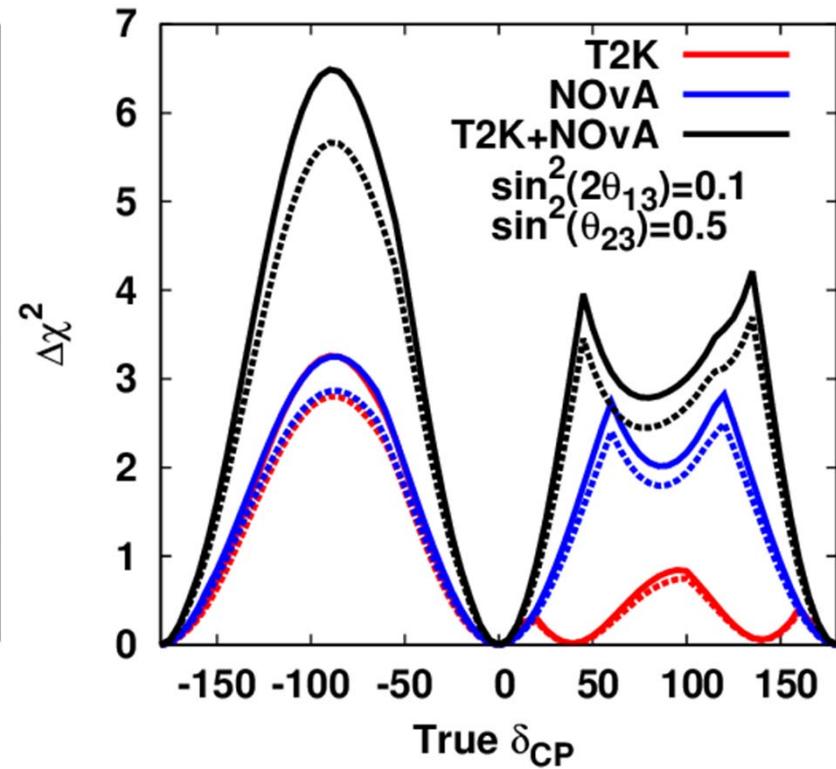
# T2K+NOvA CP violation sensitivity

- Assuming both experiments run 50%  $\nu$ -mode, 50% anti- $\nu$  mode.
- with 5% normalization uncertainty on signal and 10% normalization uncertainty on background.
- Shown is NH case.

solid : w/o sys. error  
dash: w/ sys. error



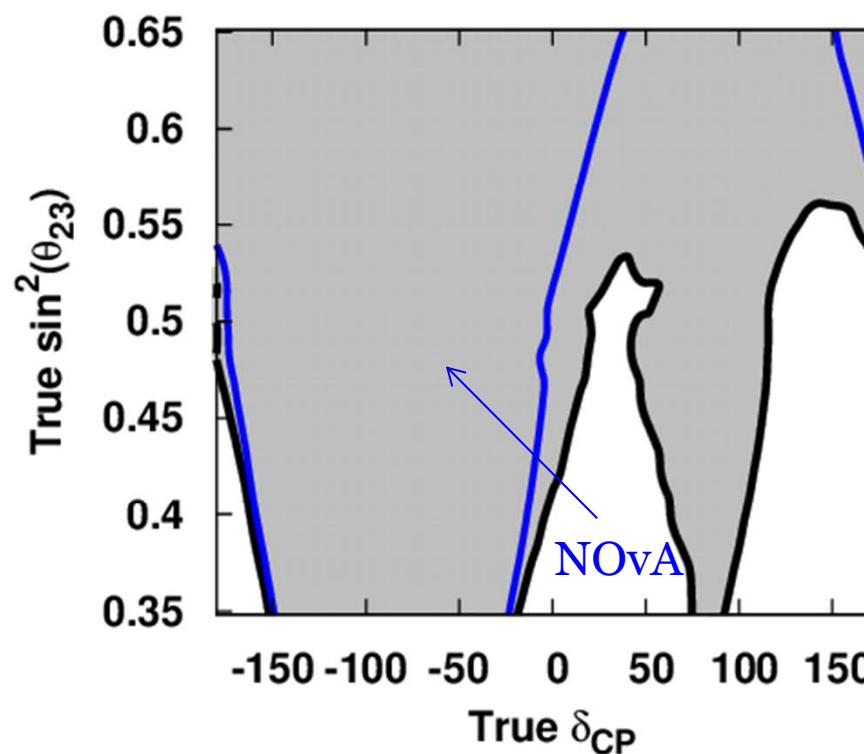
Region where  $\sin \delta = 0$  (CP conservation)  
can be excluded by 90% C.L.



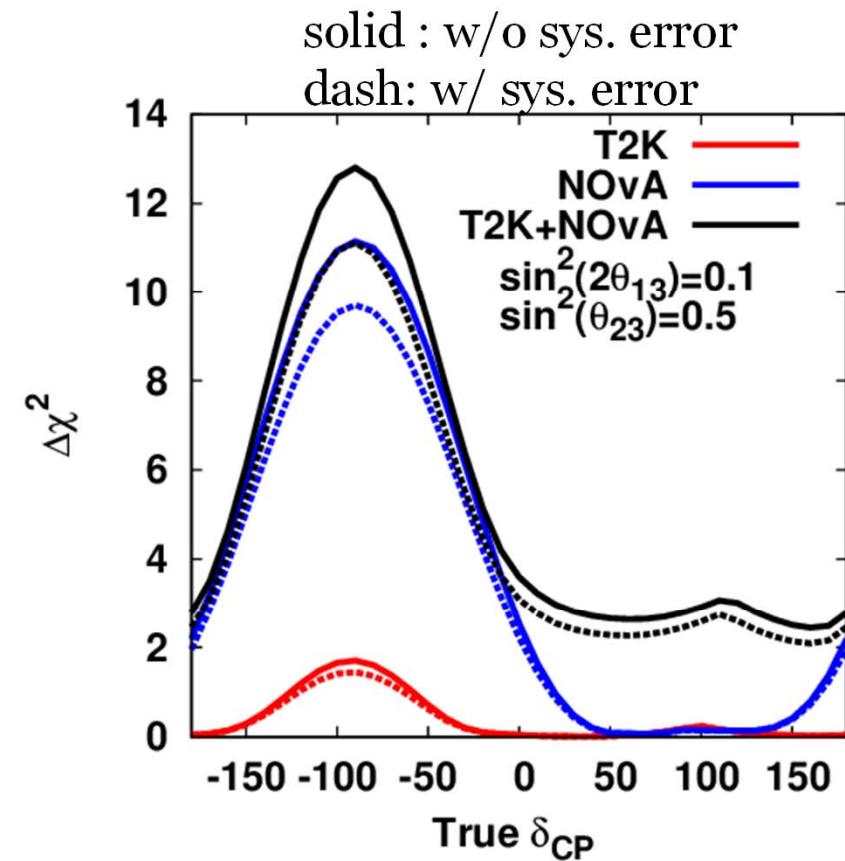
sensitivity to non-zero CP-violating term

# T2K+NOvA CP violation sensitivity

- Assuming both experiments run 50%  $\nu$ -mode, 50% anti- $\nu$  mode.
- with 5% normalization uncertainty on signal and 10% normalization uncertainty on background.
- Shown is NH case.



Region where MH can be distinguished by 90% C.L.



sensitivity to resolve MH

# Target of neutrino oscillation experiments

Mixing matrix for

Unknown

CP phase

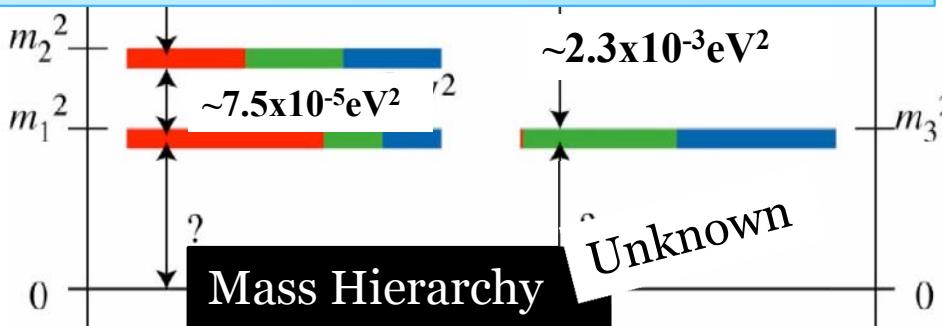
KEY to understand the origin of  
matter dominant universe

Combination of results  
from reactor experiments,

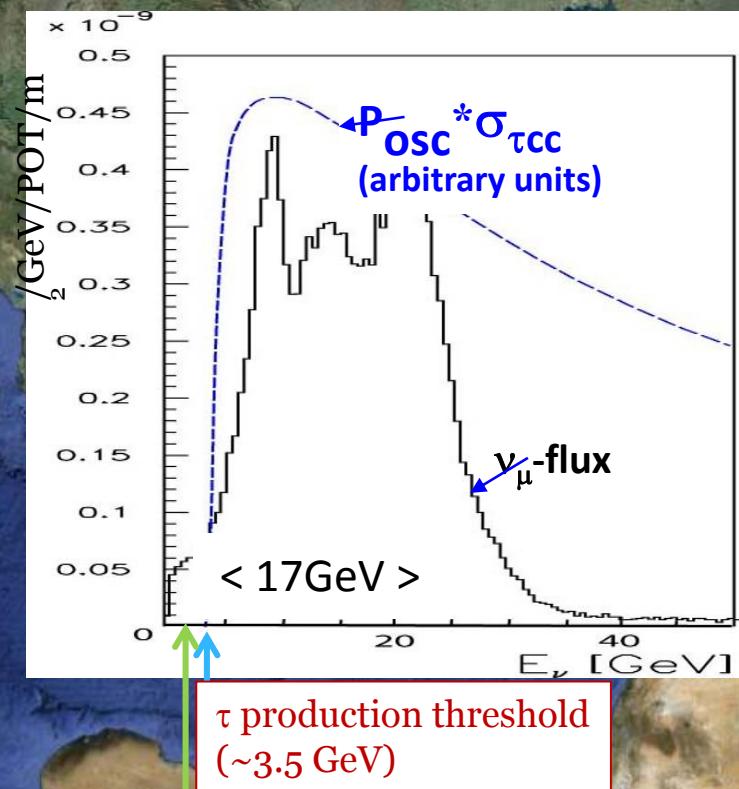
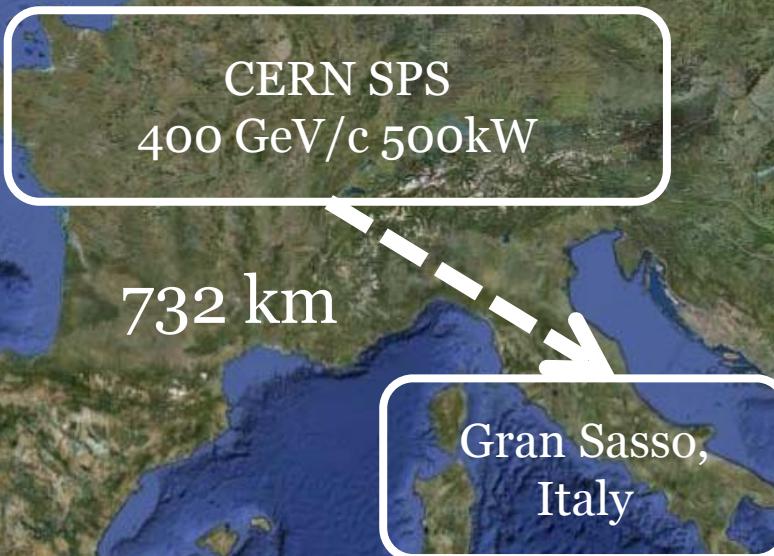
T2K and NOvA

Will enhance the sensitivity  
For these physics

Big Impact  
on ov double- $\beta$  decay search  
(hence on Majorana  $\nu$   
confirmation)

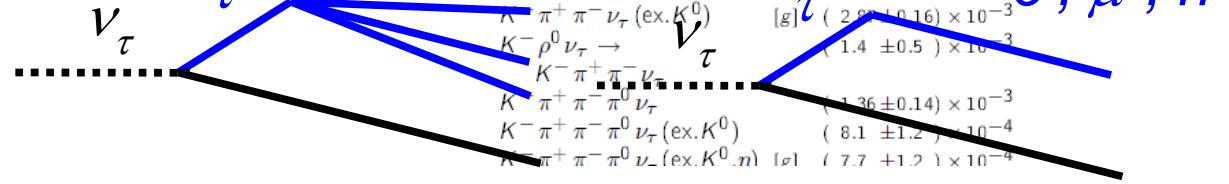


# Has $\nu_\mu$ really oscillate to $\nu_\tau$ ? OPERA experiment



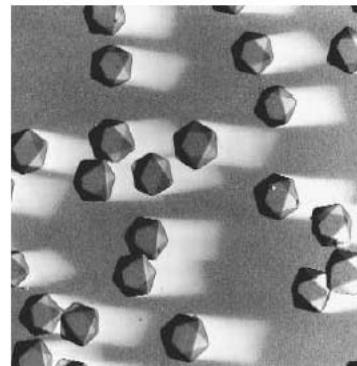
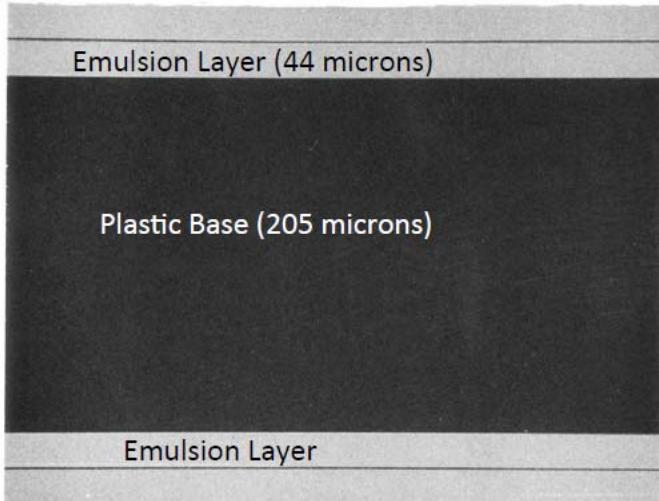
# $\nu_\tau$ – Appearance Experiment

$\tau^-$ DECAY MODES	Fraction ( $\Gamma_i/\Gamma$ )	Confid.	Sca
<b>Modes with one charged particle</b>			
particle $^- \geq 0$ neutrals $\geq 0 K^0 \nu_\tau$ (85.36 $\pm$ 0.08) %			
particle $^- \geq 0$ neutrals $\geq 0 K^0 \nu_\tau$ (84.72 $\pm$ 0.08) %			
$\mu^- \bar{\nu}_\mu \nu_\tau$ [g] (17.36 $\pm$ 0.05) %			
$\mu^- \bar{\nu}_\mu \nu_\tau \gamma$ ✓ Tau lifetime: $2.9 \times 10^{-313}$ s			
$e^- \bar{\nu}_e \nu_\tau$ [g] (17.85 $\pm$ 0.05) %			
$e^- \bar{\nu}_e \nu_\tau \gamma$ [g] (2.75 $\pm$ 0.18) %			
$h^- \geq 0 K_L^0 \nu_\tau$ [g] (12.0 $\pm$ 0.1) %			
$h^- \nu_\tau$ (11.61 $\pm$ 0.06) %			
$\pi^- \nu_\tau$ (1.97 $\pm$ 0.07) %			
$K^- \nu_\tau$ [g] (1.9 $\pm$ 0.2) %			
$h^- \geq 1$ neutrals $\nu_\tau$ (37.06 $\pm$ 0.10) %			
$h^- \geq 1 \pi^0 \nu_\tau$ (25.94 $\pm$ 0.09) %			
$\pi^- \pi^0 \nu_\tau$ VERY high resolution (and large mass) detectors			
$\pi^- \pi^0$ non- $\rho(770)$ $\nu_\tau$ (3.0 $\pm$ 3.2) $\times 10^{-3}$			
$K^- \pi^0 \nu_\tau$ [g] (4.29 $\pm$ 0.15) $\times 10^{-3}$			
$h^- \geq 2 \pi^0 \nu_\tau$ (10.85 $\pm$ 0.12) %			
$h^- 2\pi^0$ Approved for $2.25 \times 10^{19}$ protons on target			
$h^- 2\pi^0 \nu_\tau$ (ex. $K^0$ ) (9.35 $\pm$ 0.11) %			
$\pi^- 2\pi^0$ (ex. $K^0$ ) < $9 \times 10^{-3}$			
$\pi^- 2\pi^0 \nu_\tau$ (ex. $K^0 \mu$ ) < $9 \times 10^{-3}$			
Identification by the characteristic 'kink' on the decay point using			
<b>Modes with three charged particles</b>			
$h^- h^- h^+ \geq 0$ neutrals $\geq 0 K_L^0 \nu_\tau$ (15.19 $\pm$ 0.08) %			
$h^- h^- h^+ \geq 0$ neutrals $\nu_\tau$ (14.56 $\pm$ 0.08) %			
(ex. $K_S^0 \rightarrow \pi^+ \pi^-$ ) ("3-prong")			
$h^- h^- h^+ \nu_\tau$ (9.80 $\pm$ 0.08) %			
$h^- h^- h^+ \nu_\tau$ (ex. $K^0$ ) (9.46 $\pm$ 0.07) %			
$h^- h^- h^+ \nu_\tau$ (ex. $K^0, \omega$ ) (9.42 $\pm$ 0.07) %			
$\pi^- \pi^+ \pi^- \nu_\tau$ (9.32 $\pm$ 0.07) %			
$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ ) (9.03 $\pm$ 0.06) %			
$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ ) < 2.4 %			
$\pi^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ ) [g] (9.00 $\pm$ 0.06) %			
$b^- b^- b^+ \geq 1$ neutrals $\nu_\tau$ (5.38 $\pm$ 0.07) %			
$b^- b^- b^+ \geq 1$ neutrals $\nu_\tau$ (ex. $K^0$ ) (5.08 $\pm$ 0.06) %			
$h^- h^- h^+ \pi^0 \nu_\tau$ (4.75 $\pm$ 0.06) %			
$h^- h^- h^+ \pi^0 \nu_\tau$ (ex. $K^0, \omega$ ) (2.79 $\pm$ 0.08) %			
$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (4.61 $\pm$ 0.06) %			
$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0$ ) (4.48 $\pm$ 0.06) %			
$\pi^- \pi^+ \pi^- \pi^0 \nu_\tau$ (ex. $K^0, \omega$ ) [g] (2.70 $\pm$ 0.08) %			
$h^- h^- h^+ \geq 2 \pi^0 \nu_\tau$ (ex. $K^0$ ) (5.18 $\pm$ 0.33) $\times 10^{-3}$			
$h^- h^- h^+ \geq 2 \pi^0 \nu_\tau$ (ex. $K^0, \omega$ ) (5.06 $\pm$ 0.32) $\times 10^{-3}$			
$h^- h^- h^+ \geq 2 \pi^0 \nu_\tau$ (ex. $K^0, \omega, \eta$ ) [g] (4.95 $\pm$ 0.32) $\times 10^{-3}$			
$h^- h^- h^+ \geq 3 \pi^0 \nu_\tau$ [g] (10 $\pm$ 4) $\times 10^{-4}$			
$h^- h^- h^+ \geq 0$ neutrals $\nu_\tau$ (2.3 $\pm$ 0.7) $\times 10^{-4}$			
$K^- h^- h^+ \geq 0$ neutrals $\nu_\tau$ (6.24 $\pm$ 0.24) $\times 10^{-3}$			
$K^- h^- h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K^0$ ) (4.27 $\pm$ 0.20) $\times 10^{-3}$			
$K^- h^- h^+ \geq 0$ neutrals $\nu_\tau$ (ex. $K^0, \omega$ ) (8.7 $\pm$ 1.2) $\times 10^{-4}$			
$K^- \pi^+ \pi^- \geq 0$ neutrals $\nu_\tau$ (4.78 $\pm$ 0.21) $\times 10^{-3}$			
$K^- \pi^+ \pi^- \geq 0$ neutrals $\nu_\tau$ (3.68 $\pm$ 0.20) $\times 10^{-3}$			
$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ ) (-3.42 $\pm$ 0.17) $\times 10^{-3}$			
$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ ) (2.8 $\pm$ 0.16) $\times 10^{-3}$			
$K^- \rho^0 \nu_\tau \rightarrow$ (1.4 $\pm$ 0.5) $\times 10^{-3}$			
$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \eta$ ) (-1.36 $\pm$ 0.14) $\times 10^{-3}$			
$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0$ ) (8.1 $\pm$ 1.2) $\times 10^{-4}$			
$K^- \pi^+ \pi^- \nu_\tau$ (ex. $K^0, \omega$ ) [g] (7.7 $\pm$ 1.2) $\times 10^{-4}$			

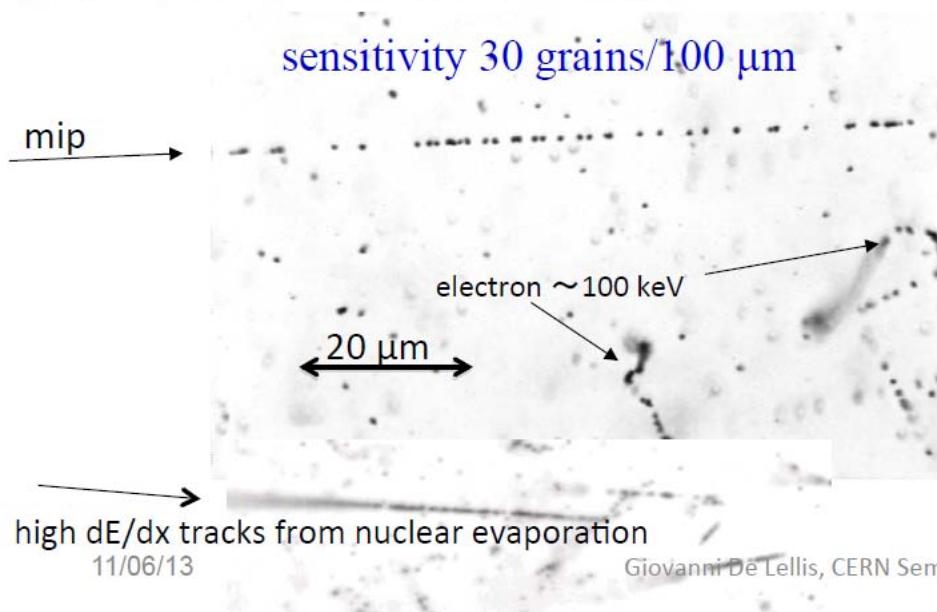


# OPERA using Nuclear emulsion

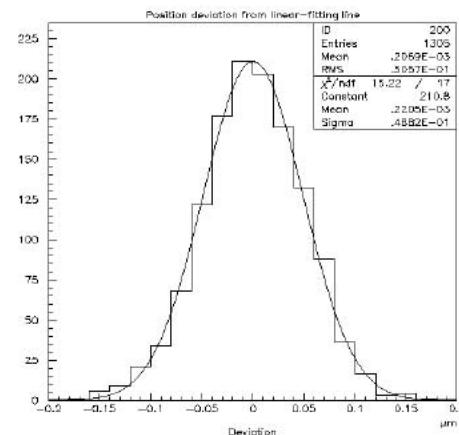
INDUSTRIAL EMULSION FILMS BY FUJI FILM



**basic detector: AgBr crystal,**  
size = 0.2 micron  
detection eff.= 0.16/crystal  
 $10^{13}$  "detectors" per film



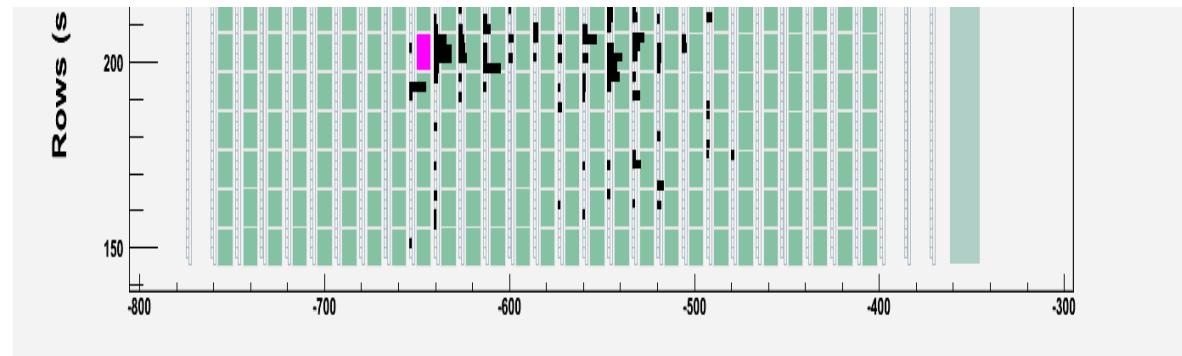
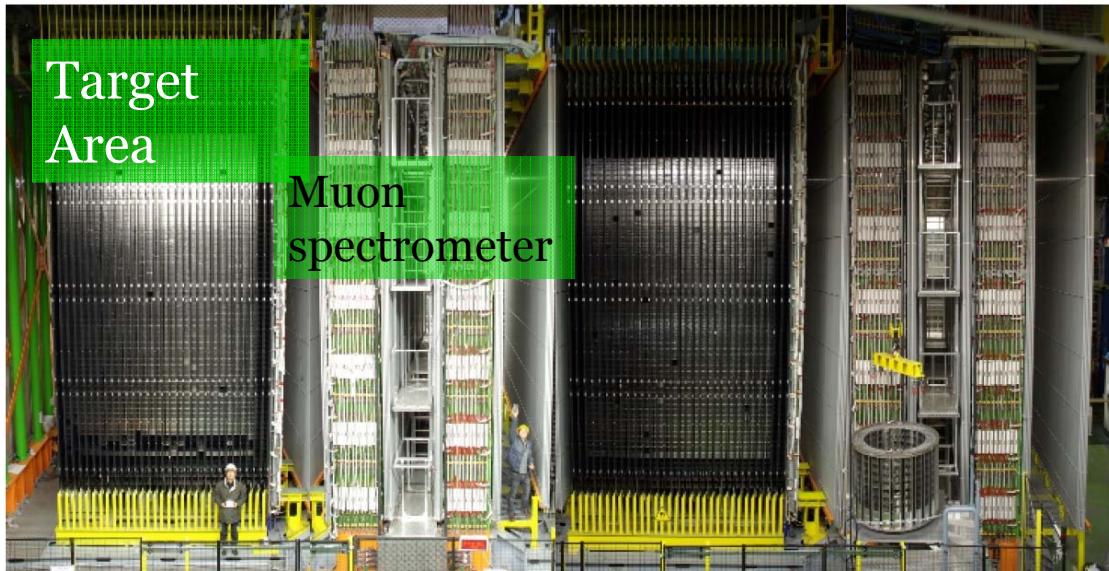
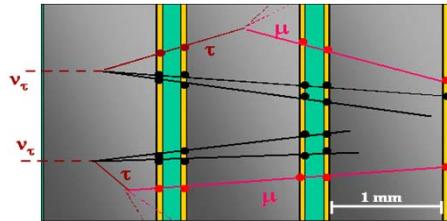
intrinsic resolution: 50 nm  
deviation from linear-fit line. (2D)



# OPERA, 1.2kton emulsion detector, since 2008

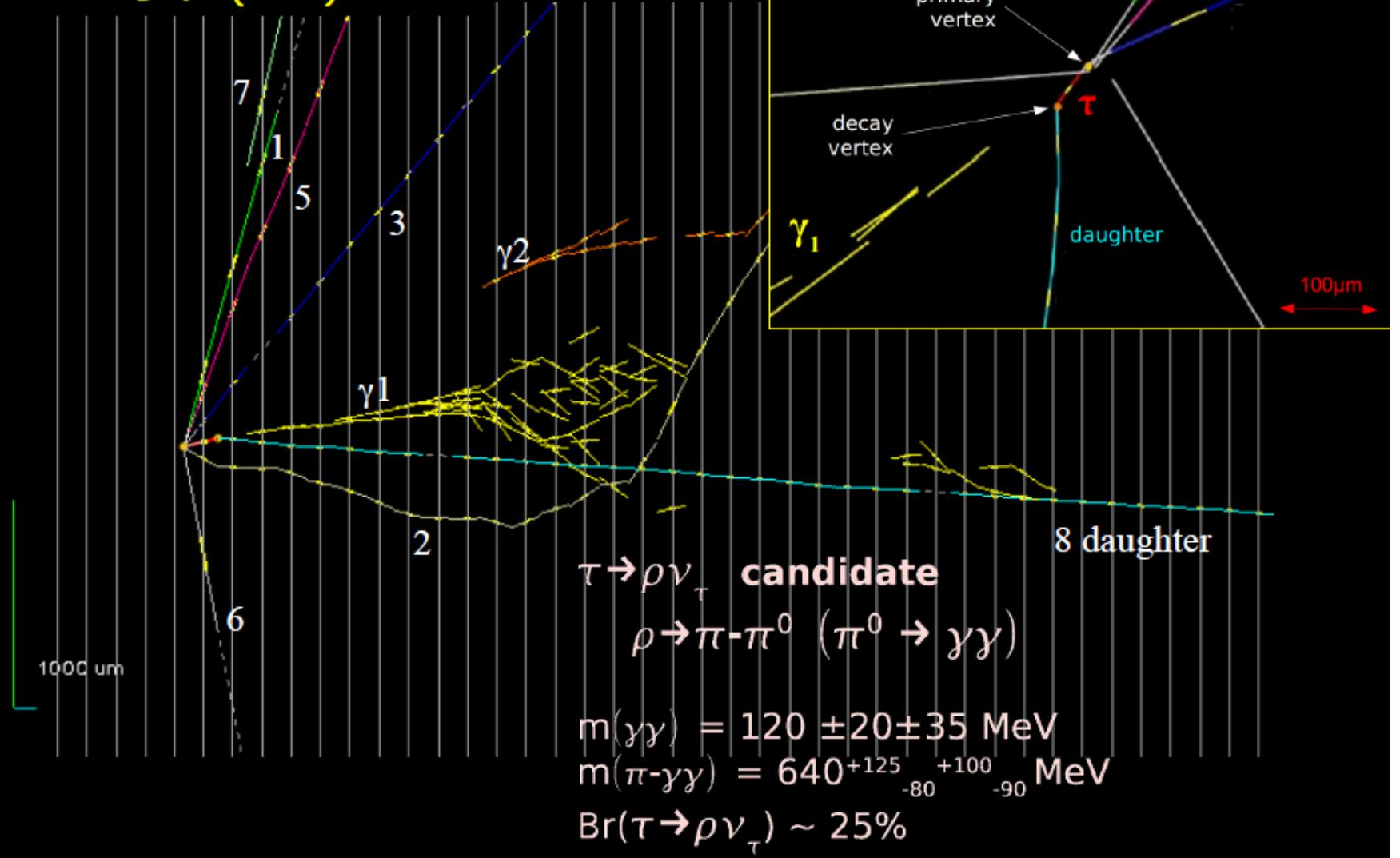
146621 bricks, each 8.3kg

- 56 (1mm) Pb sheets
- 57 (300 $\mu$ m) emulsion layers
- 2 (300 $\mu$ m) changeable sheets (CS)



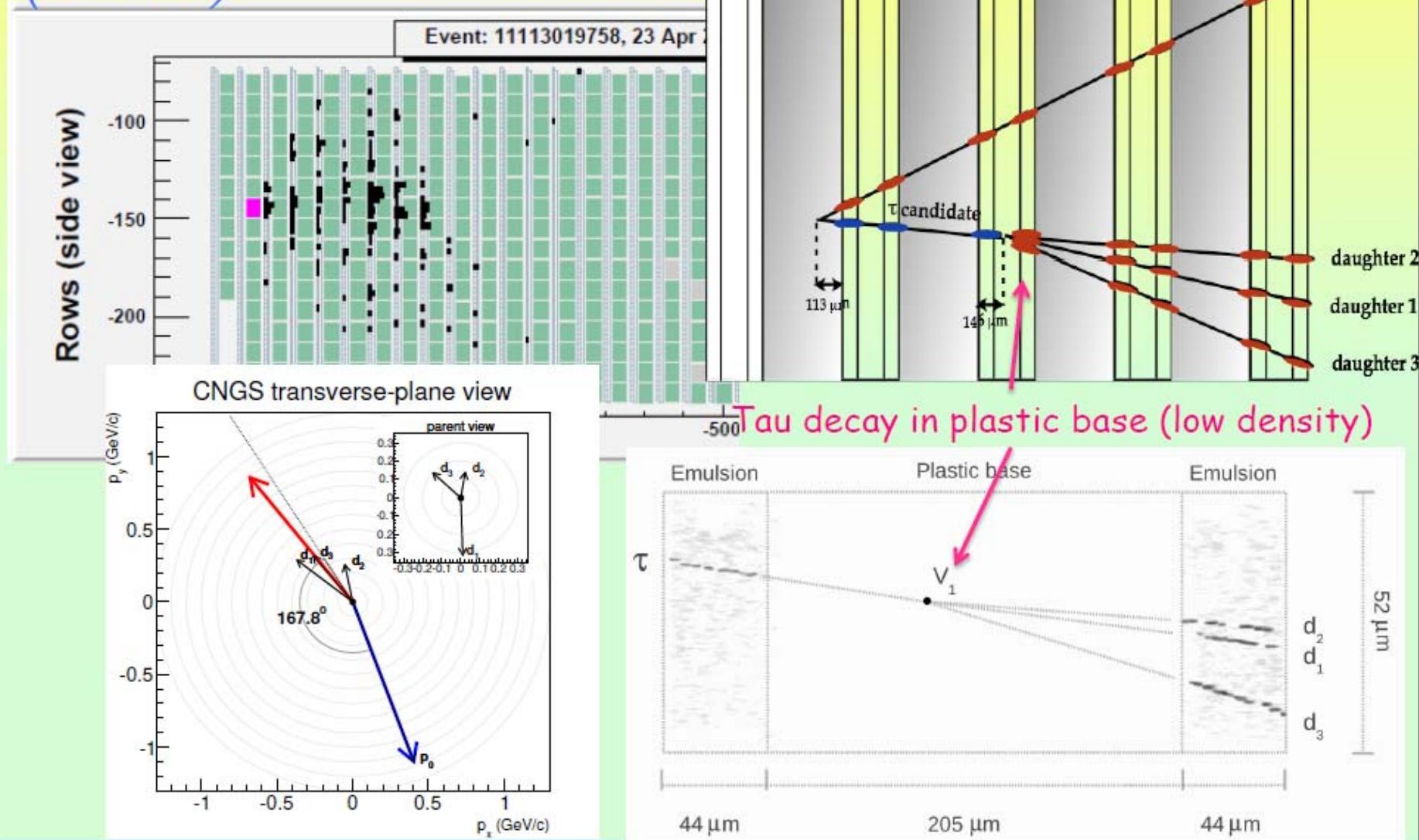
# 1<sup>st</sup> $\nu_\tau$ candidate ( $\tau \rightarrow h$ ) (2010)

Phys. Lett. B691 (2010) 138



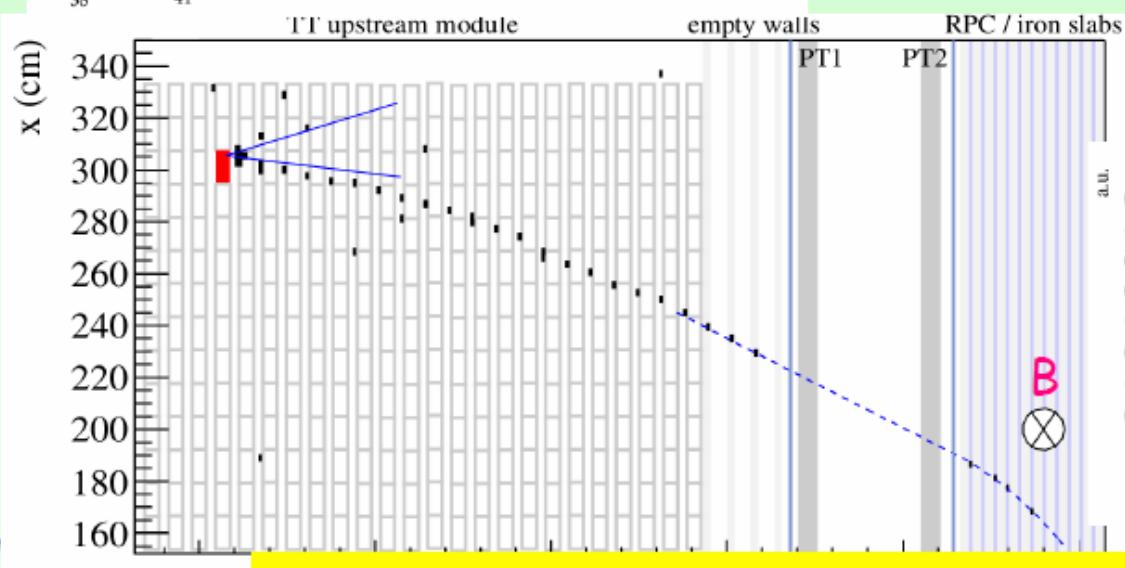
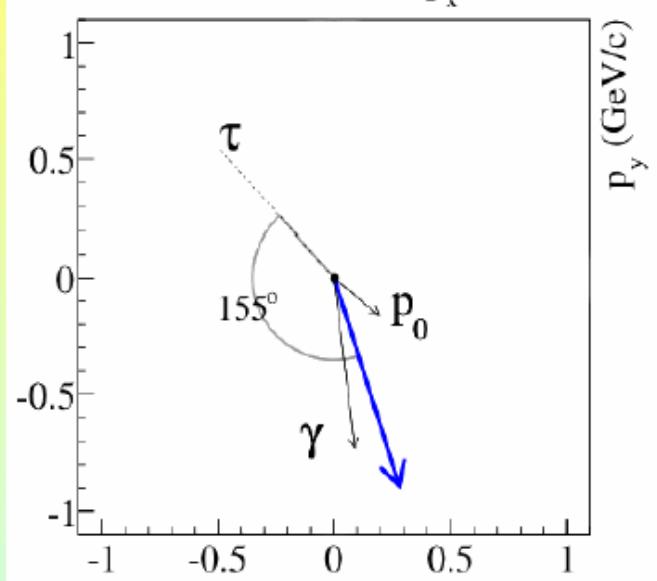
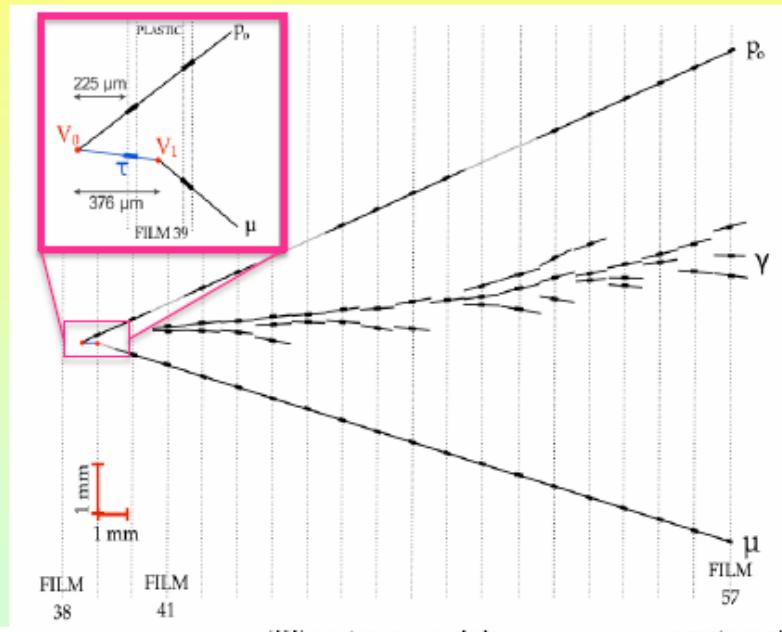
# 2<sup>nd</sup> $\nu_\tau$ candidate (2012) ( $\tau \rightarrow 3h$ )

JHEP 11 (2013) 036

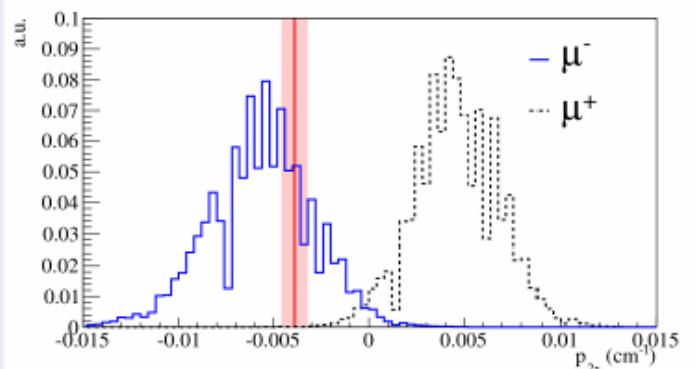


# 3<sup>rd</sup> $\nu_\tau$ candidate ( $\tau \rightarrow \mu$ ) (2013)

PHYSICAL REVIEW D 89 (2014) 051102(R)



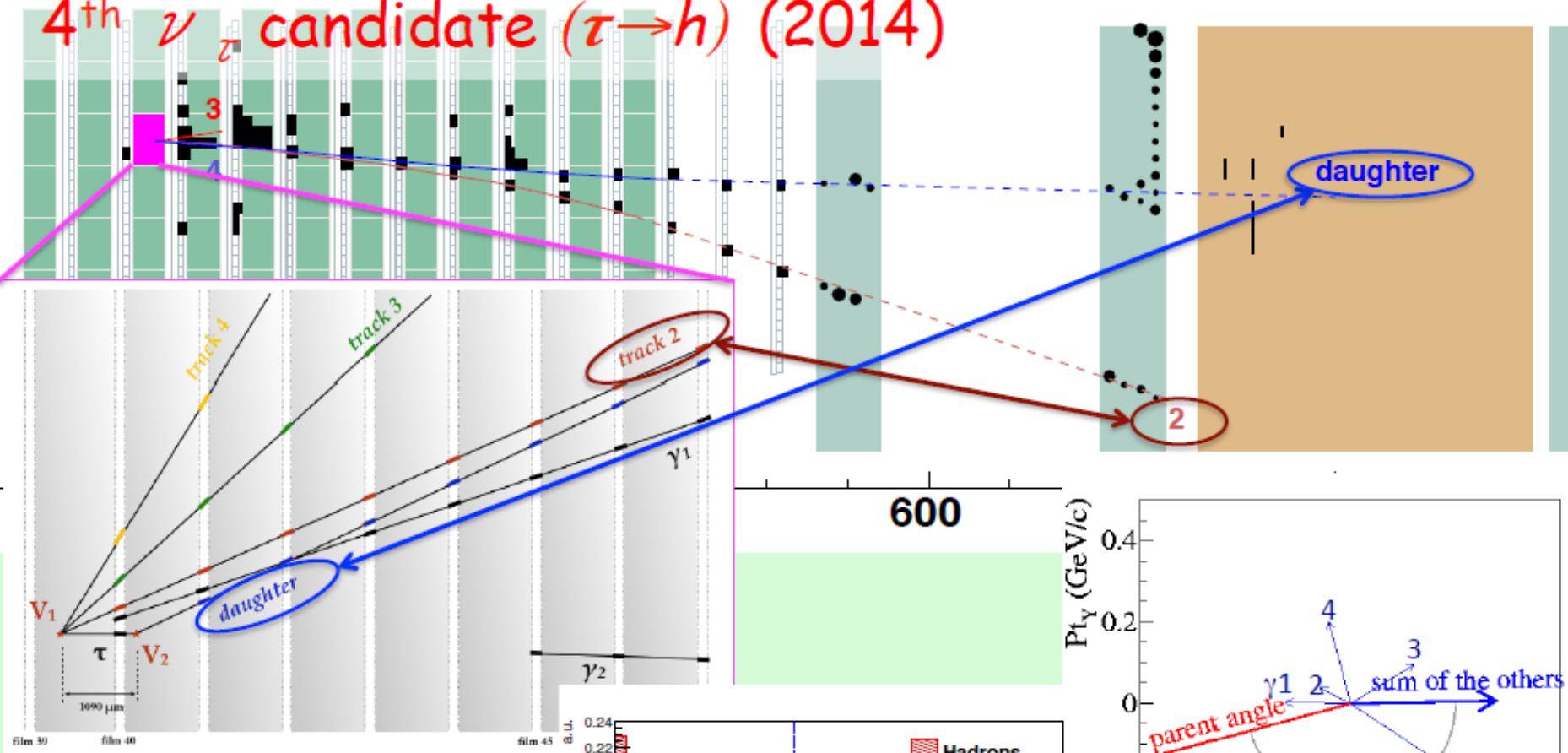
Negative muon measured in the muon spectrometer



Neutrino 2014 First measurement of the lepton charge in appearance mode

10

# 4th $\nu_\tau$ candidate ( $\tau \rightarrow h$ ) (2014)



Track 2 from neutrino interaction vertex,  $p = 1.9 \text{ GeV}$  stopping in first iron slab of the magnet

$$D = \frac{L}{R_{\text{lead}}(p)} \frac{\rho_{\text{average}}}{\rho_{\text{lead}}} = 0.40^{+0.04}_{-0.05}$$

Data sample:

2008/09 : 398 (0 $\mu$  events) + 1553 (1 $\mu$  events)

2010/11/12 : 582 (0 $\mu$  events) + 2153 (1 $\mu$  events)

The expected signal and background is normalized to the number of located events

$$n^{0\mu}(\nu_\tau^{CC}) = \frac{\langle \sigma(\nu_\tau^{CC}) \rangle}{\langle \sigma(\nu_\mu^{CC}) \rangle} \frac{\langle \epsilon^{0\mu}(\nu_\tau^{CC}) \rangle}{\langle \epsilon^{0\mu}(\nu_\tau^{CC}) \rangle + \alpha \langle \epsilon^{0\mu}(\nu_\tau^{NC}) \rangle} n^{0\mu} \quad \alpha = \frac{NC}{CC}$$

Decay channel	Expected signal $\Delta m_{23}^2 = 2.32 \text{ meV}^2$	Total background	Observed
$\tau \rightarrow h$	$0.4 \pm 0.08$	$0.033 \pm 0.006$	2
$\tau \rightarrow 3h$	$0.57 \pm 0.11$	$0.155 \pm 0.03$	1
$\tau \rightarrow \mu$	$0.52 \pm 0.1$	$0.018 \pm 0.007$	1
$\tau \rightarrow e$	$0.61 \pm 0.12$	$0.027 \pm 0.005$	0
<b>Total</b>	$2.1 \pm 0.42$	$0.23 \pm 0.04$	4

Two statistical method :

- Fisher combination of single channel p-value
- Likelihood ratio

p-value =  $1.03 \times 10^{-5}$  of no oscillation

no oscillation excluded  
at  $4.2 \sigma$  CL