

# (Light) Sterile Neutrinos

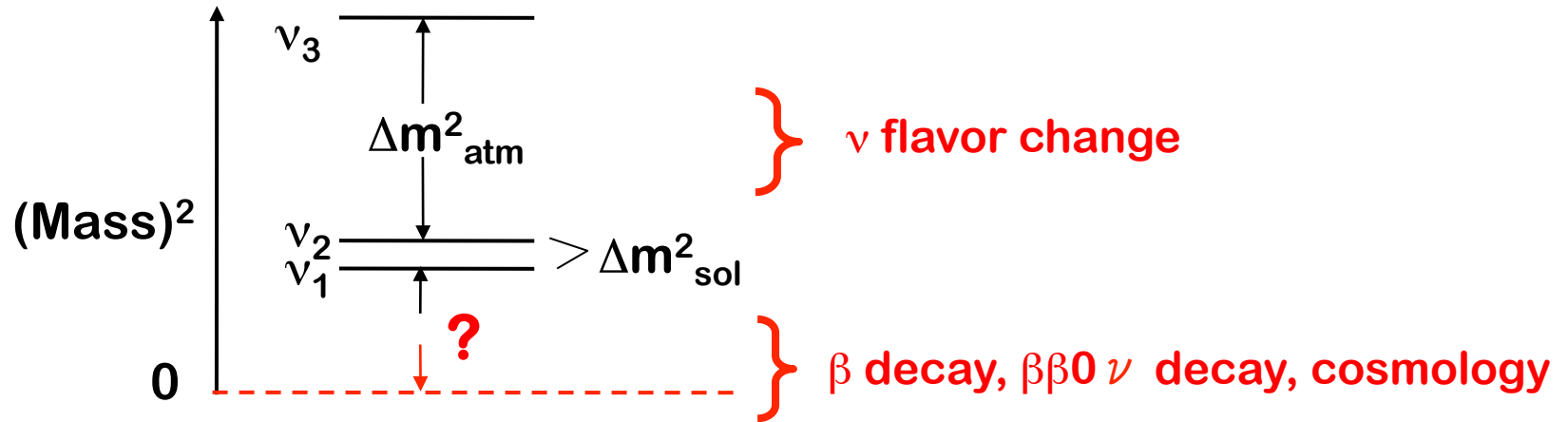


Thierry Lasserre – Saclay & APC

APPEC Meeting  
APC Paris, June 24<sup>th</sup> 2014

# Open questions in $\nu$ physics

- What are the masses of the mass eigenstates  $\nu_i$ ?



- Is the spectral pattern  or ?  $\nu$  behavior in matter,  $\beta\beta 0 \nu$ , osc.

- Is there any conserved Lepton Number (Dirac or Majorana  $\nu$ ) ?  $\beta\beta 0 \nu$

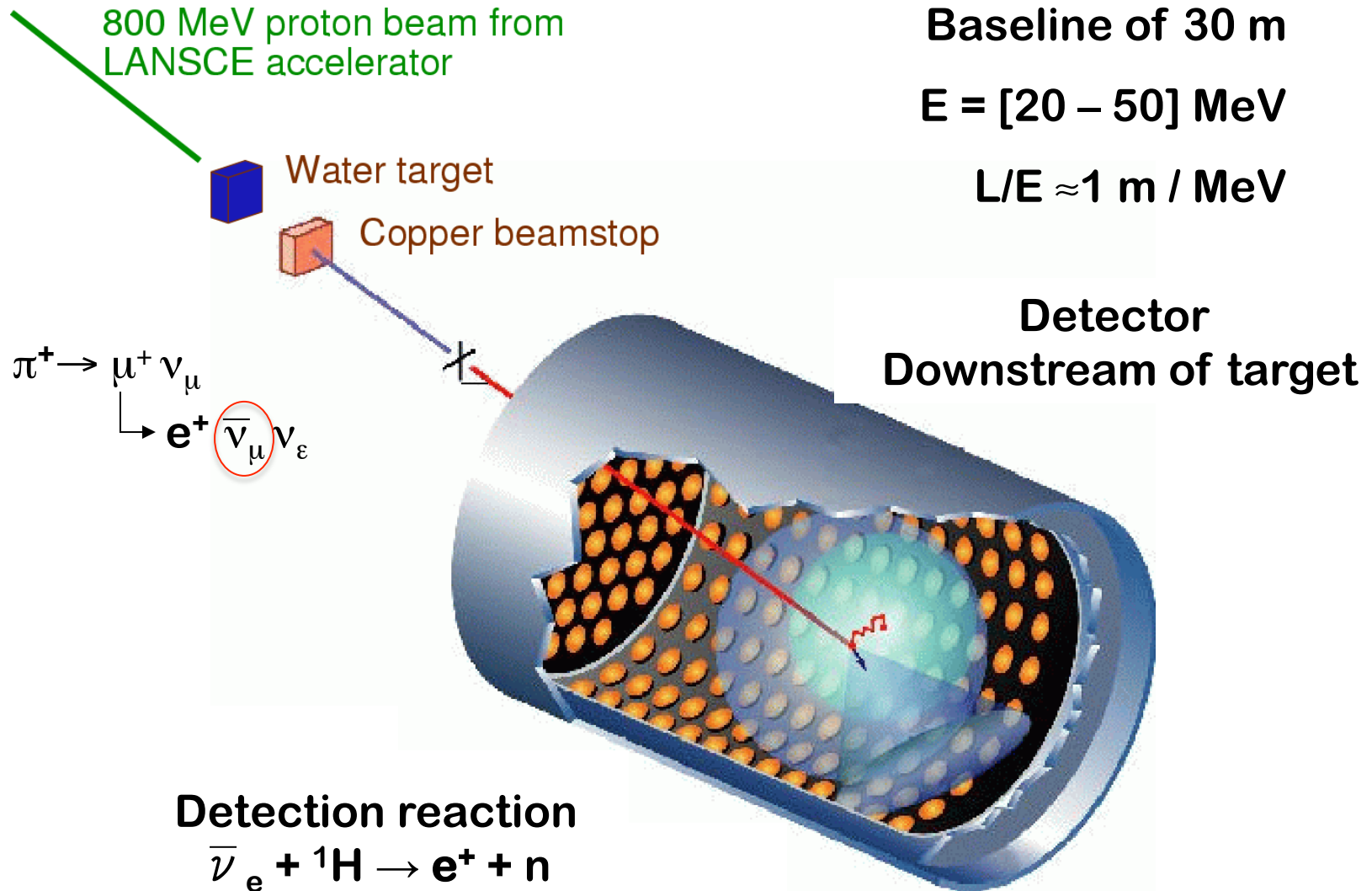
- Precise measurements of the leptonic mixing matrix?
  - Do the behavior of  $\nu$  violate CP?
  - Is leptonic  $CP$  responsible for the matter-antimatter asymmetry?
- }  $\nu$  flavor change

- Are there additional (sterile) neutrino states?  $\nu$  flavor, Astro/Cosmo

# Anomalies or new $\nu$ -oscillation?

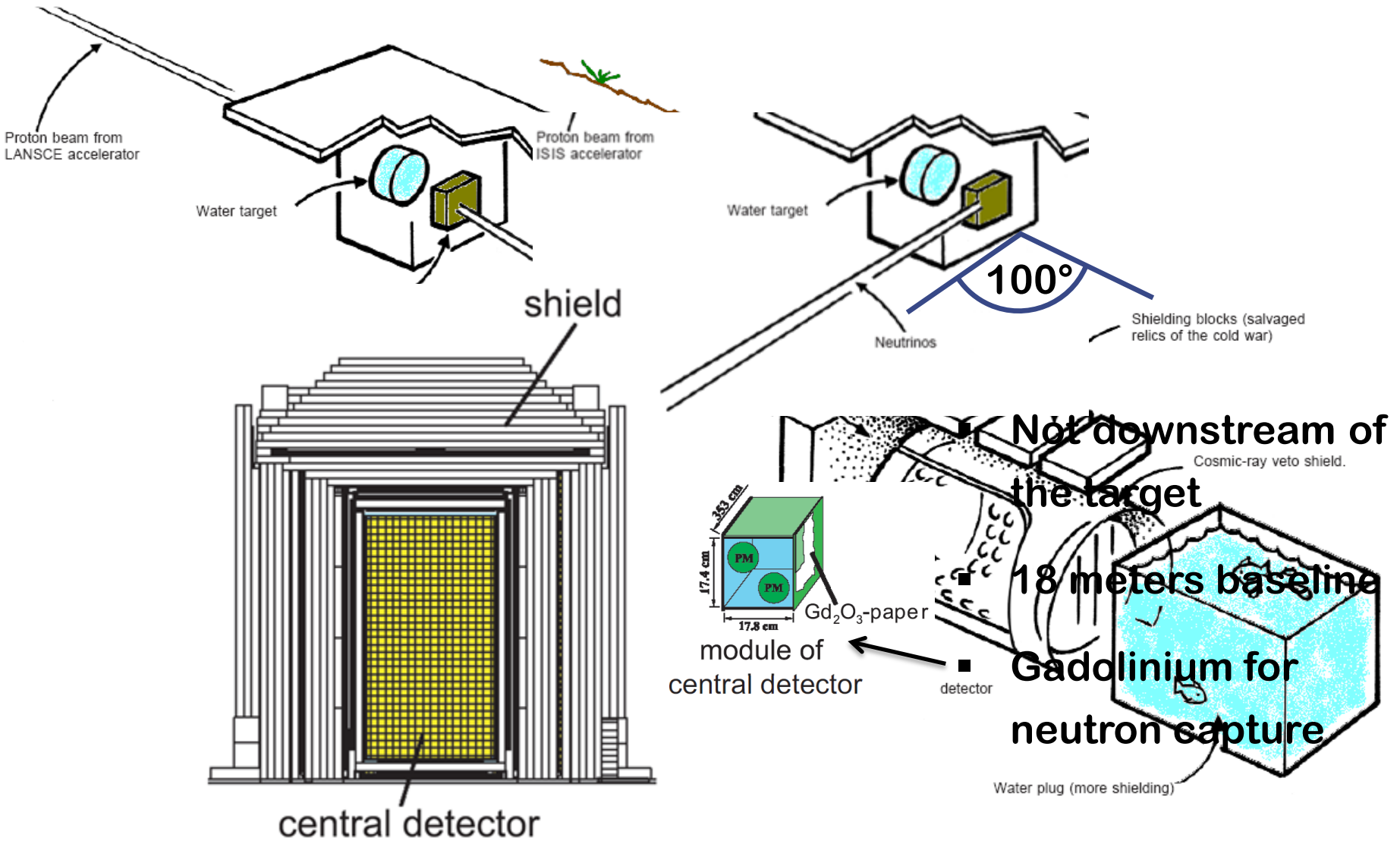
# LSND (stopped $\pi^+$ beam)

Anomaly on the electron antineutrino interaction rate



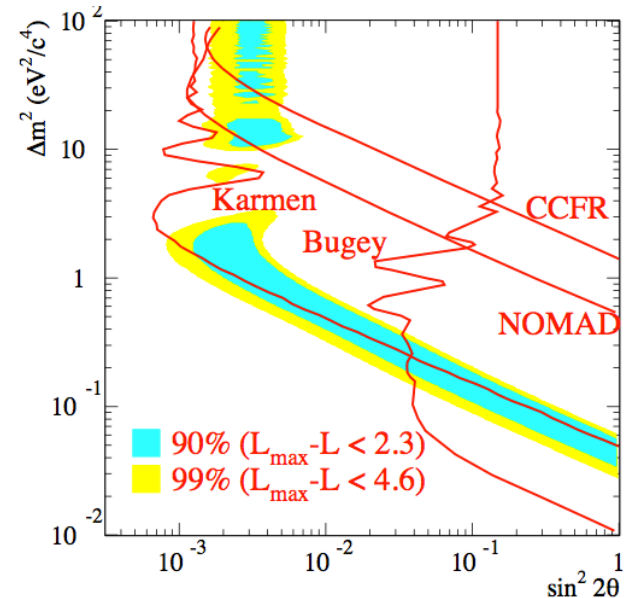
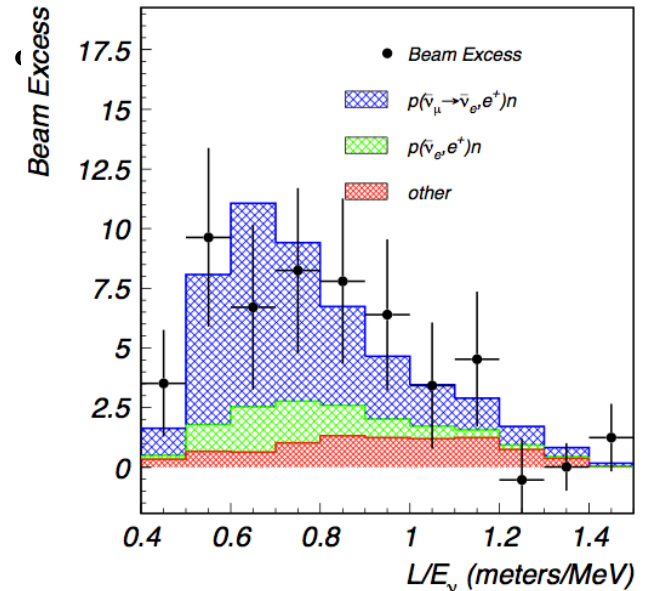
# Karmen (stopped $\pi^+$ beam)

Oscillation not confirmed – exclude part of LSND



# LSND Results

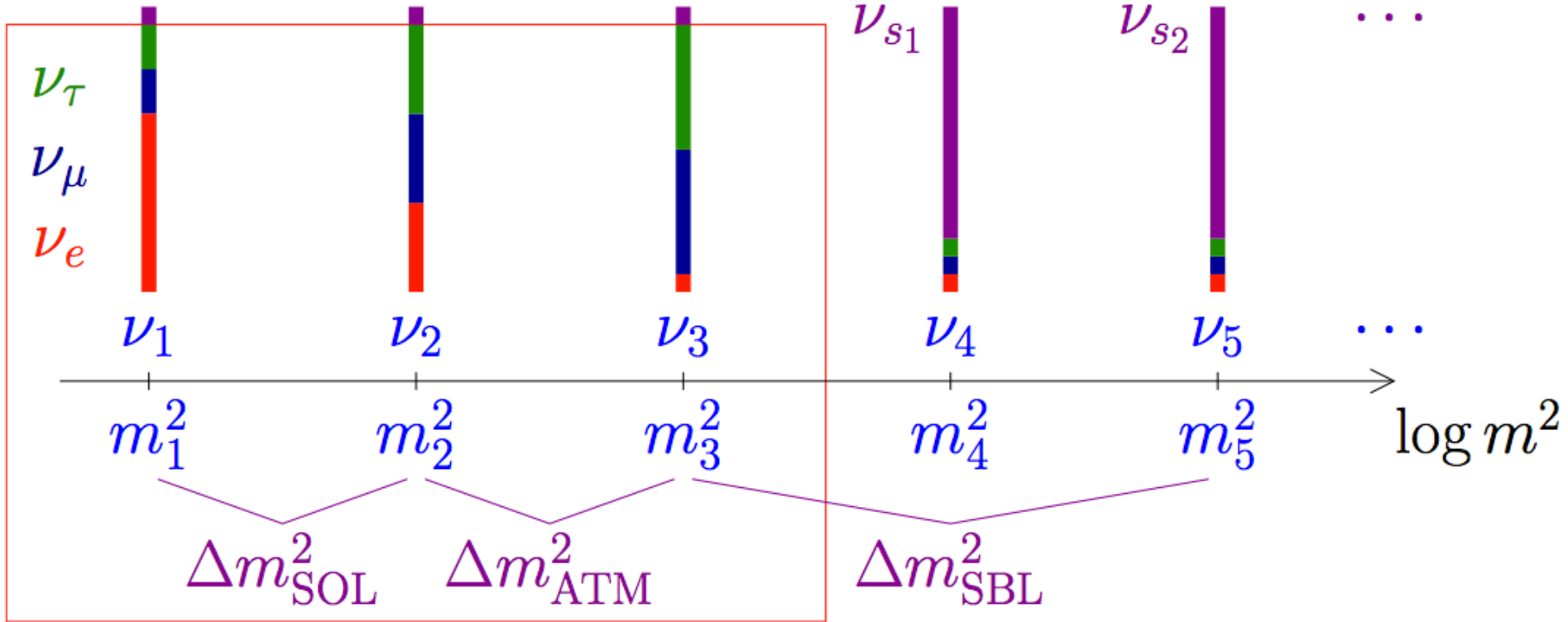
- 1<sup>st</sup> results in 1995
- **Channel:** anti- $\nu_\mu \rightarrow \text{anti-}\nu_e$
- **Detection :** anti- $\nu_e + {}^1\text{H} \rightarrow e^+ + n$
- **Baseline:** 30 m
- **Energy:**  $20 < E \text{ (MeV)} < 50$
- **Status:**
  - anti- $\nu_e$  excess observed  
 $\rightarrow 32.2 \pm 9.4 \pm 2.3 \text{ (3.8}\sigma\text{)}$
  - not confirmed nor ruled out by Karmen
- **$\nu$ -Oscillation interpretation:**
  - $\Delta m^2 > 0.1 \text{ eV}^2 \gg \Delta m_{\text{atm}}^2$
  - Require a 4<sup>th</sup> neutrino state



# The (light) sterile neutrino hypothesis

- Generic extension of SM model
- Add a SM singlet fermion
- Mixing with active  $\nu$ 's

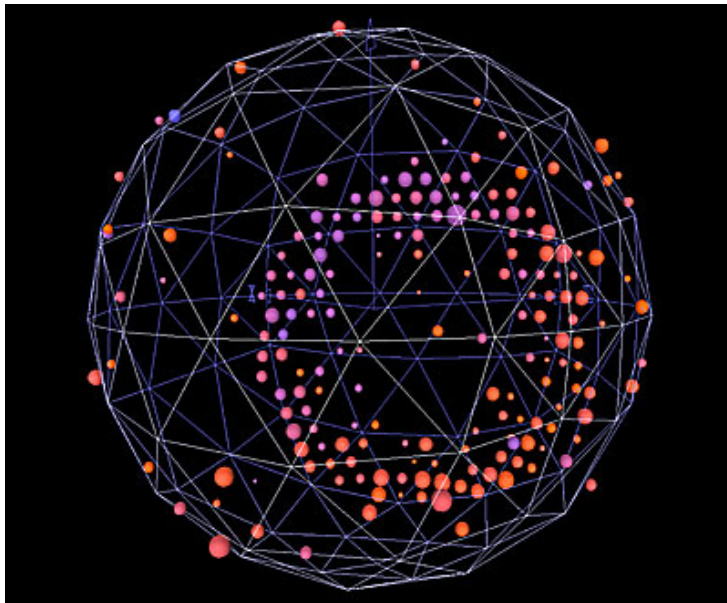
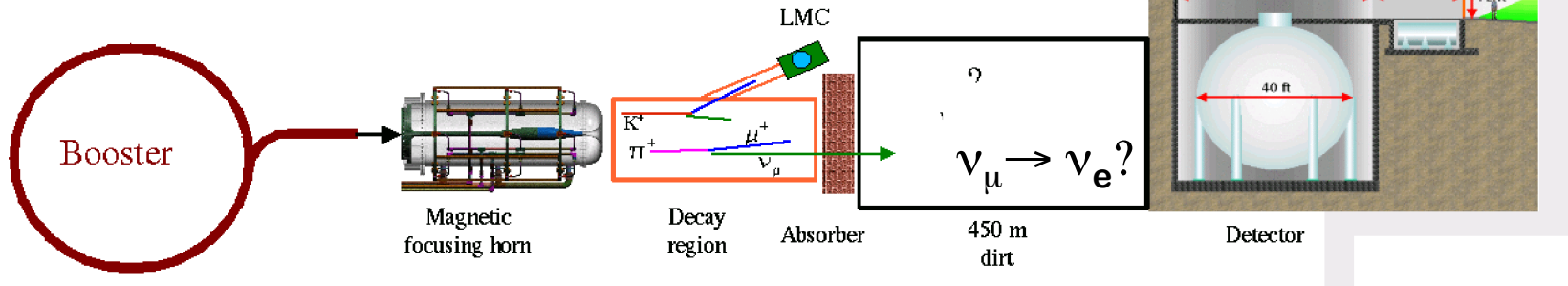
No or tiny SM model interaction



3 $\nu$ -mixing

# MiniBooNE (FNAL)

Primary goal: look for  $\nu_e$  appearance in a  $\nu_\mu$  beam  
 Check the LSND with similar L/E

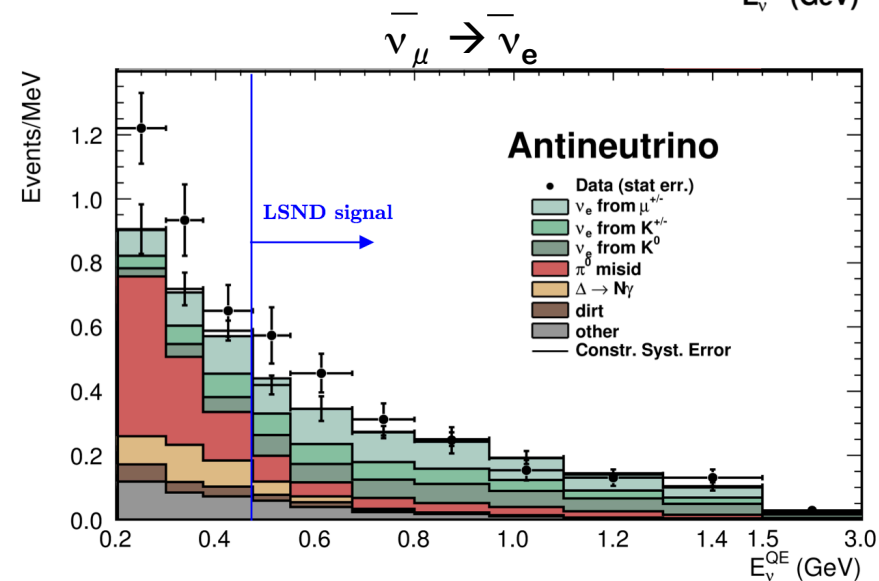
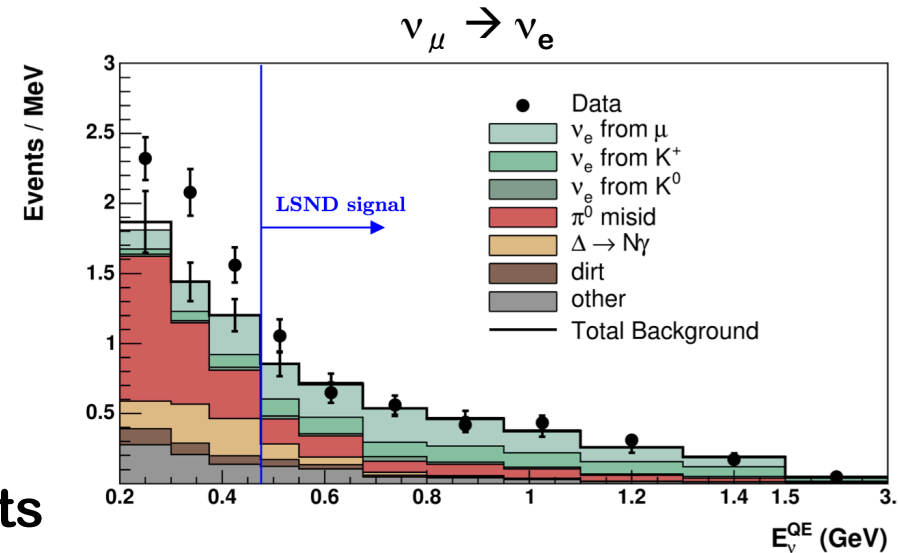


- Beam:  $\pi^+$  ( $\pi^-$ ) decay in flight
- Detection: Cherenkov + scintillation
- L/E  $\approx 1$  m / MeV
  - Baseline: 541 m
  - $200 < E$  (MeV)  $< 3000$
- Statistics:
  - $\nu$  :  $6.46 \times 10^{20}$  POT (2008)
  - $\bar{\nu}$  :  $1.27 \times 10^{20}$  POT (2012)



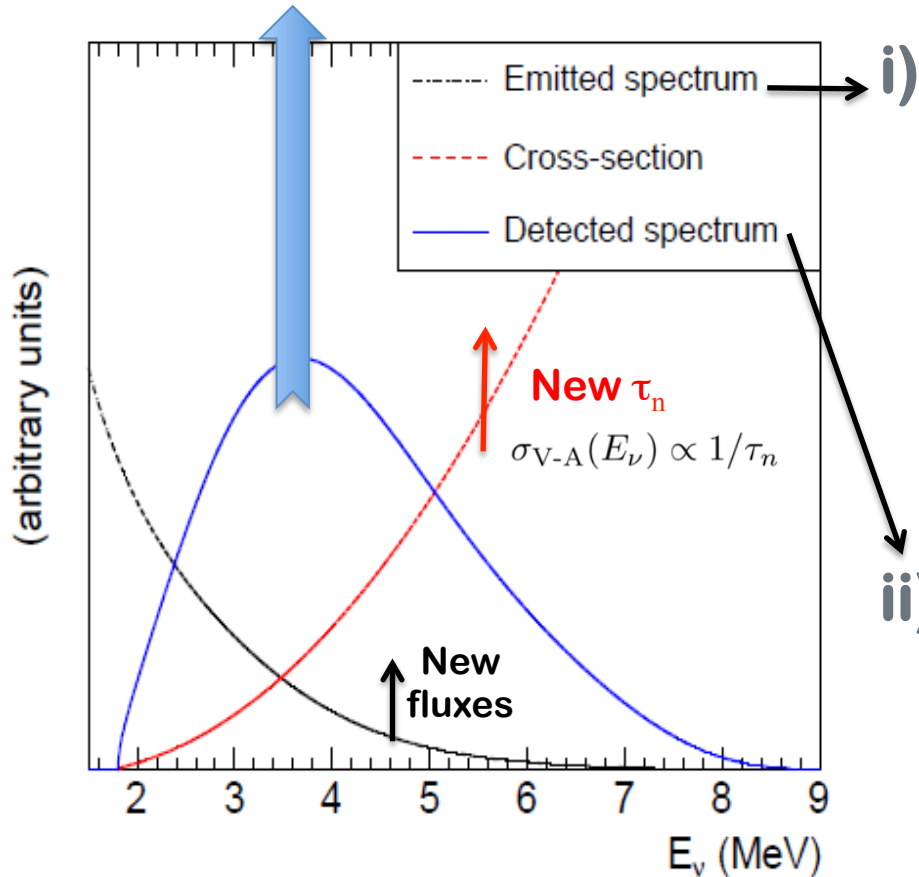
# MiniBooNE Results

- Results published from 2007-12
- **Channel:**  $(\text{anti-})\nu_\mu \rightarrow (\text{anti-})\nu_e$
- **Detection:**  $\nu_e (p)n \rightarrow e p$  (CCQE)
- **Results:**
  - An overall  $3.8 \sigma$  excess of events
  - Mostly at low energy
- **Interpretation:**
  - Backgrounds issue?  
(to be checked by MicroBooNE)
  - 4<sup>th</sup> neutrino? Or more....
- **MiniBooNE was not conclusive checking the LSND anomaly**



# New Reactor $\nu$ -Fluxes

Increased prediction of detected flux by 6.5%



## i) Neutrino Emission:

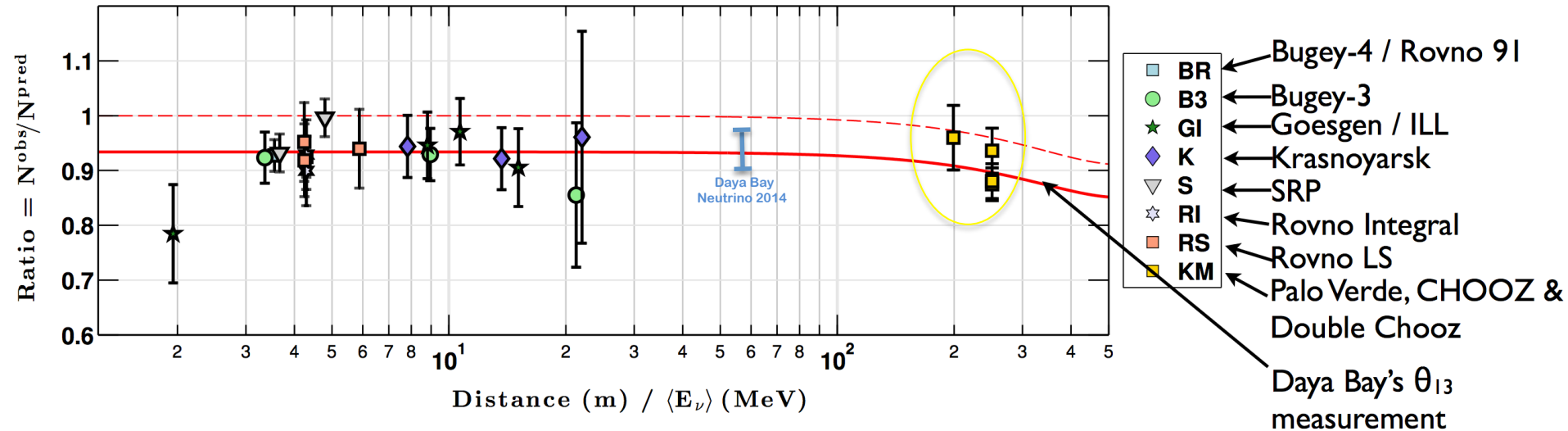
- Improved reactor neutrino spectra  $\rightarrow$  +3.5%
- Accounting for long-lived isotopes in reactors  $\rightarrow$  +1%

## ii) Neutrino Detection:

- Reevaluation of  $\sigma_{IBD} \rightarrow$  +1.5% (evolution of the neutron life time)
- Reanalysis of all SBL experiments

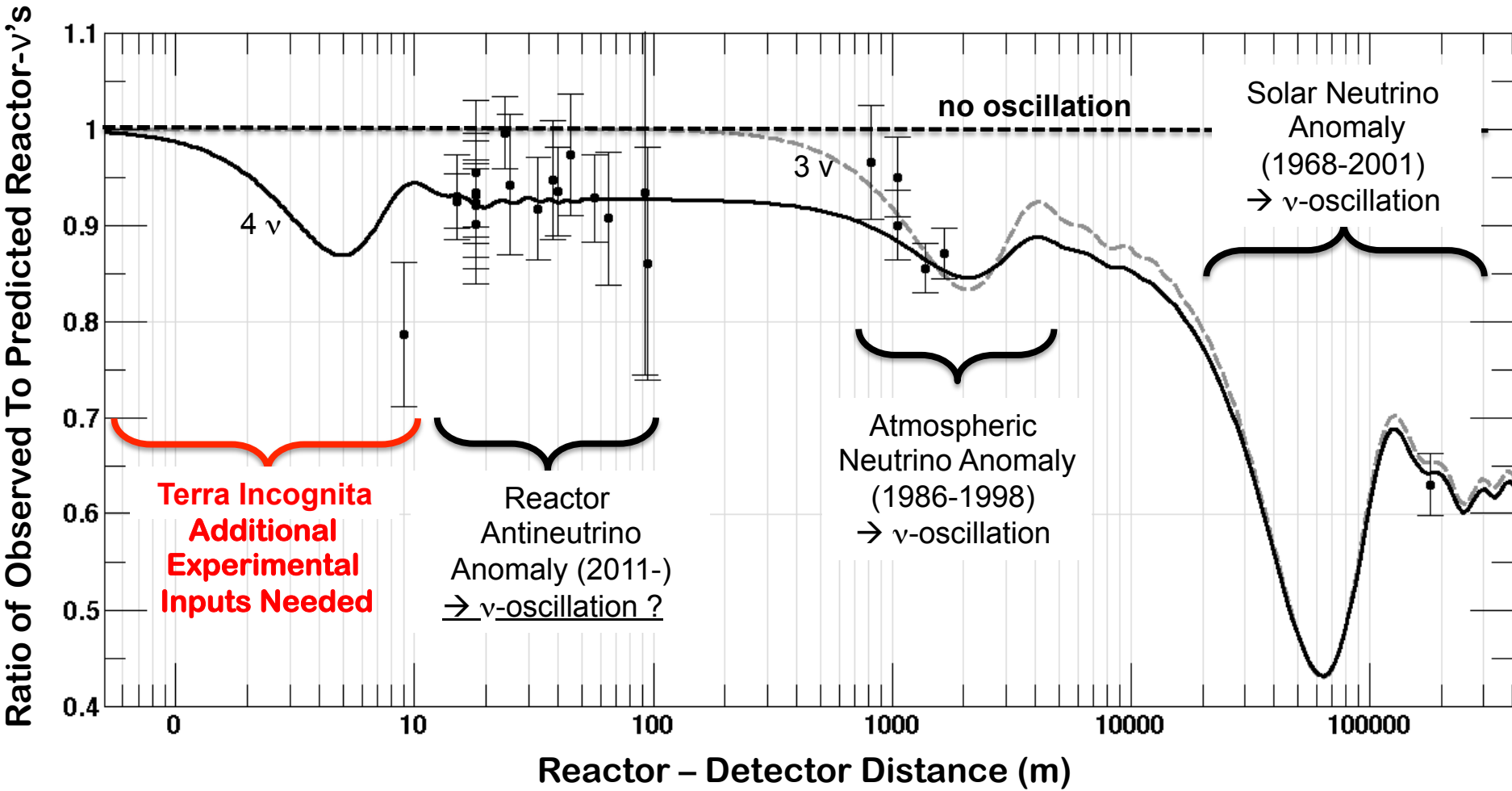
# The Reactor Anomaly

## 2014 Reactor Anomaly Update (new)



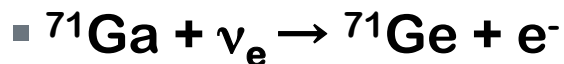
- All known nuclear corrections to  $\beta - \nu$  spectra
- Refined treatment of experimental correlations
- Latest updated neutron mean life ( $\tau_n = 881.5$  s)
- Corrects for a statistical bias (1% shift)
- km-scale baselines (Chooz, DC, PV)
  - correcting for  $\theta_{13}$  deficit from Daya Bay's measured value
- **2014 result:  $\mu = 0.938 \pm 0.023$ ,  $2.7\sigma$  deviation from unity**

# Experimental Artifact or New Physics?



# The Gallium Neutrino Anomaly

- **Test of solar neutrino radiochemical detectors GALLEX and SAGE**

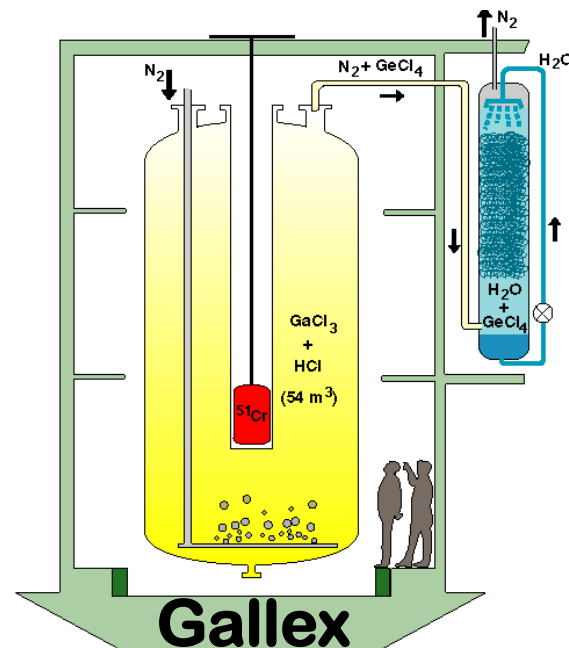
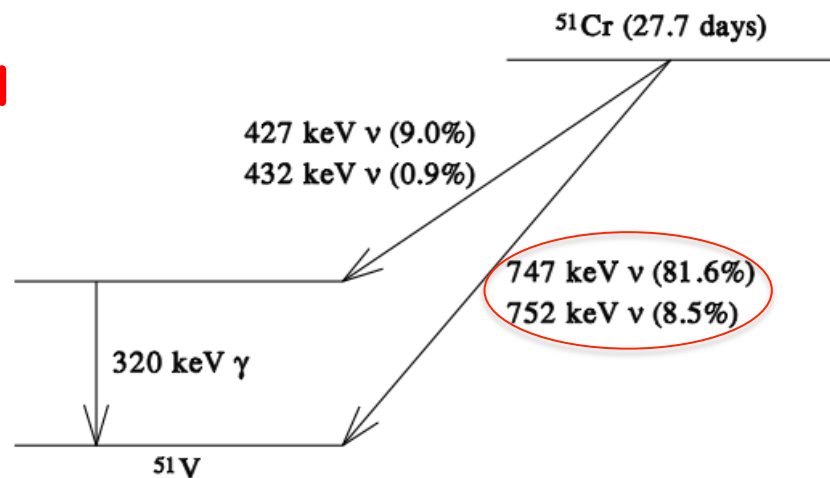


- **4 calibration runs with 0.6 - 2 MCi Electron Capture  $\nu_e$  emitters**

- Gallex,  $\langle L \rangle = 1.9$  m
    - $^{51}\text{Cr}$ , 750 keV
  - Sage,  $\langle L \rangle = 0.6$  m
    - $^{51}\text{Cr}$  &  $^{37}\text{Ar}$  (810 keV)

- **Deficit observed**

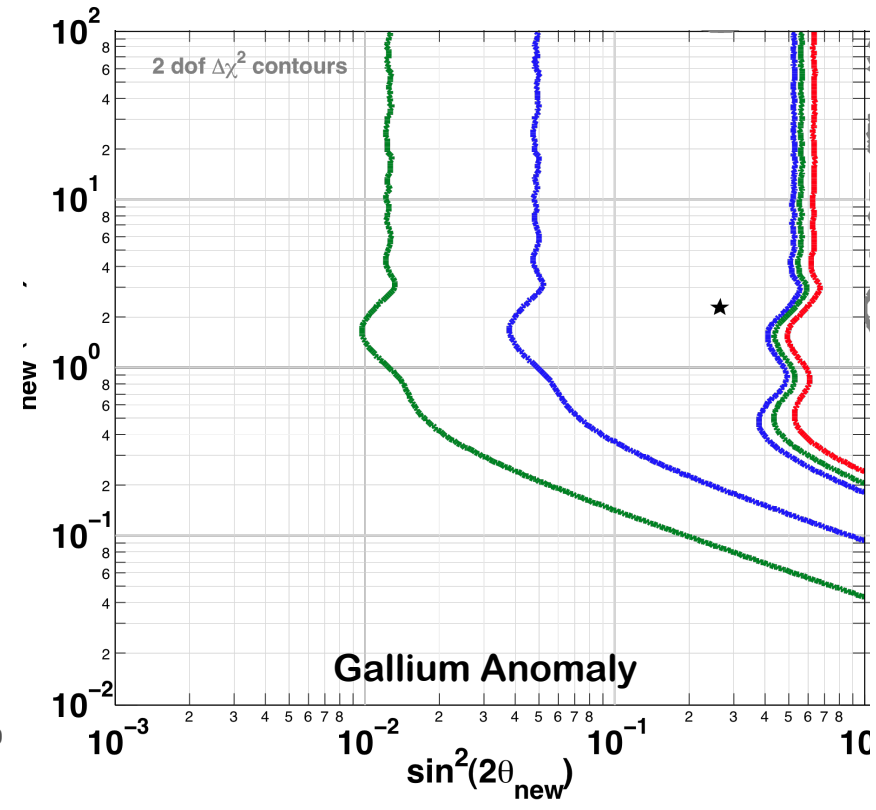
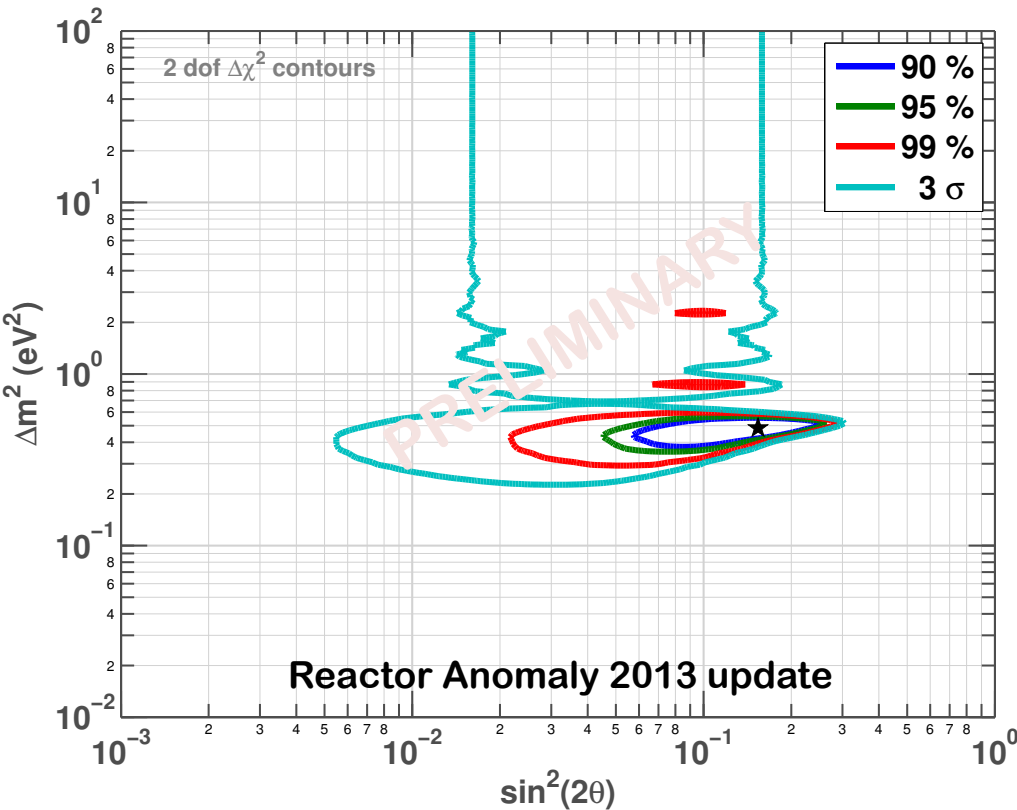
- $3\sigma$  anomaly
  - Supported by new  $^{71}\text{Ga}$  ( $^3\text{He}$ ,  $^3\text{H}$ )  $^{71}\text{Ge}$  cross section measurement



# Sterile Neutrino Interpretation

Fit to  $\nu_e$  and  $\bar{\nu}_e$  disappearance hypothesis (3+1, Okkam razor)

$$\begin{pmatrix} \nu_e \\ \nu_s \end{pmatrix} = \begin{pmatrix} \cos \theta_{\text{new}} & \sin \theta_{\text{new}} \\ -\sin \theta_{\text{new}} & \cos \theta_{\text{new}} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_{\text{new}} \end{pmatrix}, P_{ee} = 1 - \sin^2(2\theta_{\text{new}}) \sin^2\left(\frac{\Delta m_{\text{new}}^2 L}{E}\right)$$



**No-oscillation hypothesis disfavored at >99.9% C.L.**

# Interpreting data as $\nu$ -oscillation

# Anomalous & Regular Results

Anomalous	Source	Type	Signal	Channel	Significance
LSND	Meson Decay-at-Rest	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	<u>Total Rate</u> , Energy	CC	3.8 $\sigma$
MiniBooNE	Meson Decay-in-Flight	$\nu_\mu \rightarrow \nu_e$	<u>Total Rate</u> , Energy	CC	3.8 $\sigma$
Gallium	Electron Capture	$\nu_e$ dis.	<u>Total Rate</u>	CC	2.7-3.0 $\sigma$
Reactor	Beta-decay	$\nu_e$ dis.	<u>Total Rate</u> , Energy	CC	2.7 $\sigma$

Regular	Source	Type	Signal	Channel
KARMEN Icarus/Opera	Meson Decay -at-Rest & Flight	$\nu_\mu \rightarrow \nu_e$	<u>Total Rate</u> , Energy	CC
CDHS/Minos/ MiniBooNE	Meson Decay-in-Flight	$\nu_\mu \rightarrow \nu_\mu$	<u>Total Rate</u> , Energy	CC
Minos	Meson Decay-in-Flight	$\nu_\mu \rightarrow \nu_s$	<u>Total Rate</u>	CC



# Sterile- $\nu$ Phenomenology (3+1)

- $\bar{\nu}_e$  disappearance (Reactor, Gallium, ...)

$$\square \quad P_{ee} = 1 - \sin^2 2\theta_{ee} \sin^2 \frac{\Delta m_{41}^2}{4E} \quad \& \quad \sin^2 2\theta_{ee} = |U_{e4}|^2 (1 - |U_{e4}|^2)$$

- $\bar{\nu}_\mu$  disappearance (CDHS, MiniBOONE, Minos,...)

$$\square \quad P_{\mu\mu} = 1 - \sin^2 2\theta_{\mu\mu} \sin^2 \frac{\Delta m_{41}^2}{4E} \quad \& \quad \sin^2 2\theta_{\mu\mu} = |U_{\mu4}|^2 (1 - |U_{\mu4}|^2)$$

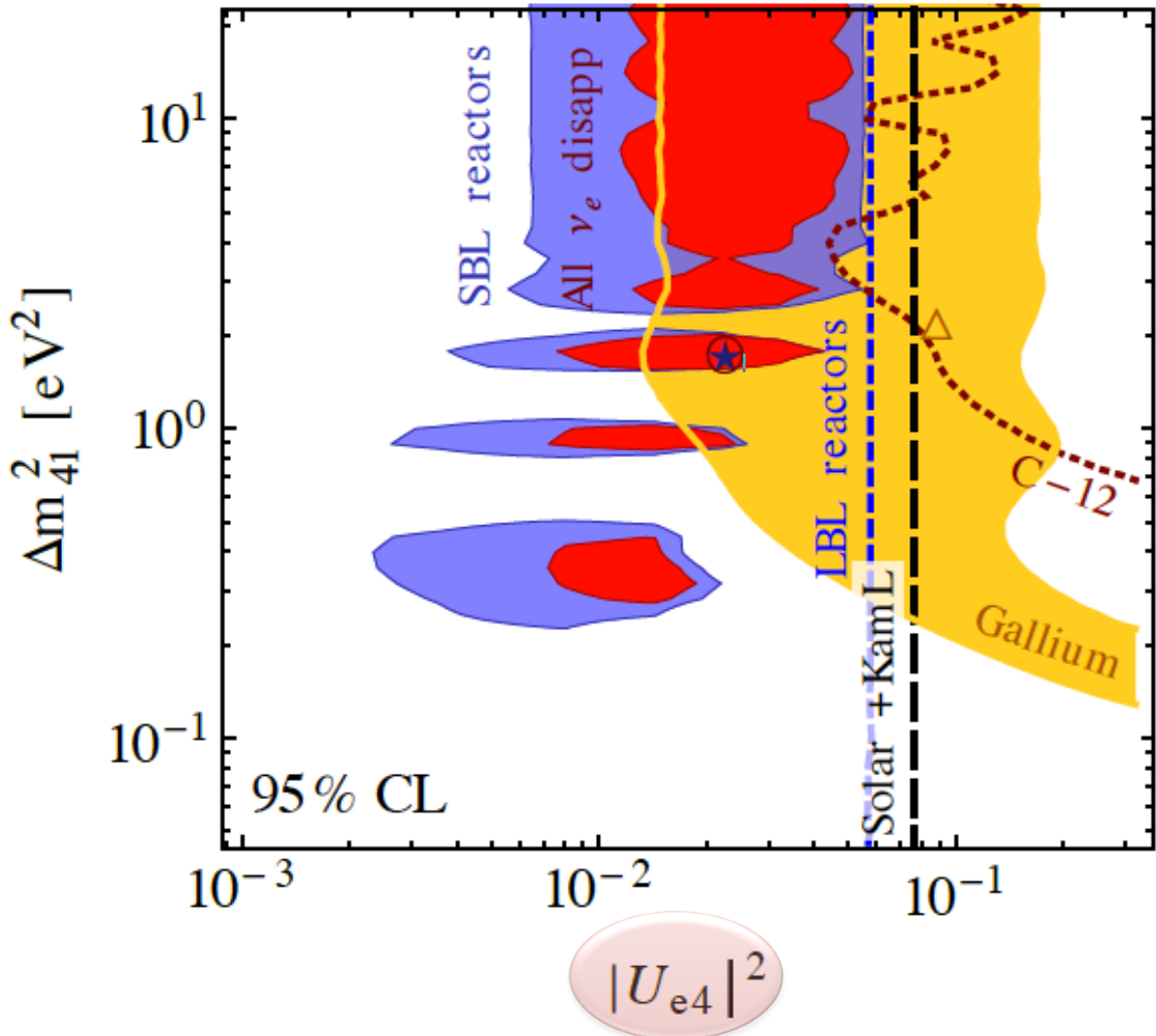
- $\bar{\nu}_e$  appearance (LSND, Karmen, MiniBooNE, Opera, Icarus...)

$$\square \quad P_{\mu e} = 4 \sin^2 2\theta_{\mu e} \sin^2 \frac{\Delta m_{41}^2}{4E} \quad \& \quad \sin^2 2\theta_{\mu e} \approx \frac{1}{4} \sin^2 2\theta_{ee} \sin^2 2\theta_{\mu\mu}$$

$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  appearance requires  $\bar{\nu}_\mu$  &  $\bar{\nu}_e$  disappearance

# $\bar{\nu}_e$ disappearance (3+1 scenario)

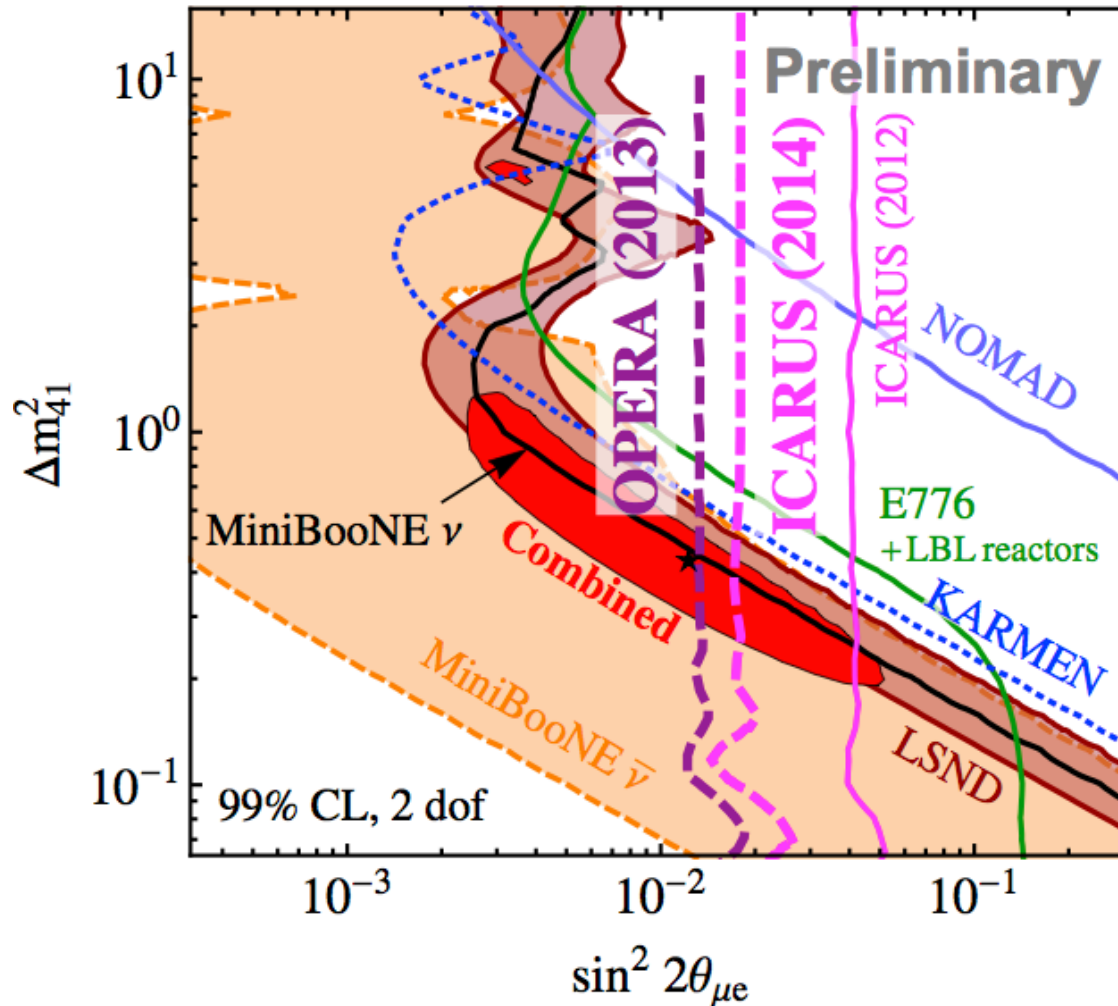
Data consistent with  $\bar{\nu}_e$  disappearance with  $L/E \approx 1$  m/MeV



J. Kopp et al., [arXiv:1303.3011](https://arxiv.org/abs/1303.3011)

# $\bar{\nu}_e$ appearance (3+1 scenario)

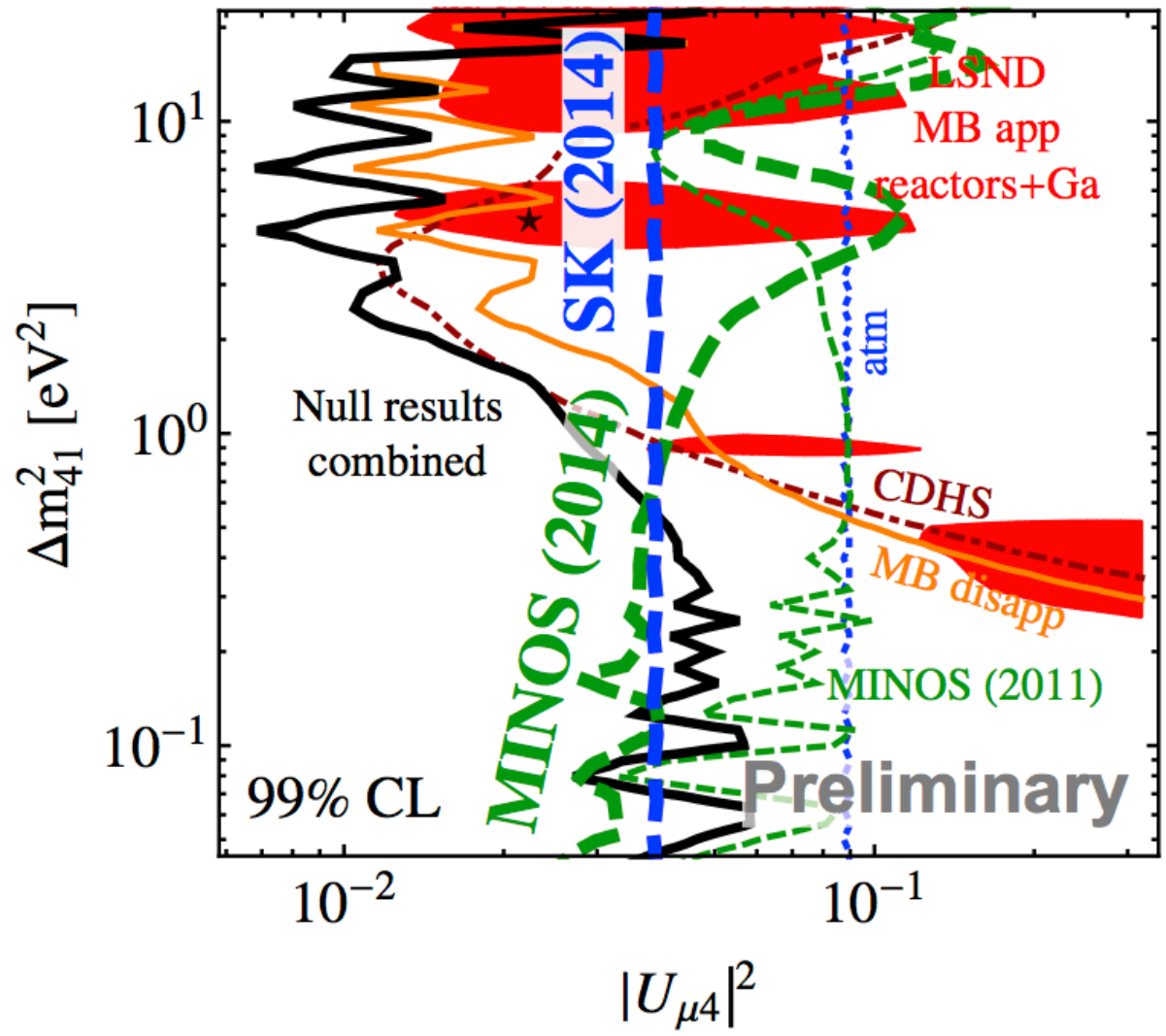
Consistent solution for  $\bar{\nu}_e$  appearance with  $L/E \approx 1$  m/MeV



J. Kopp et al., [arXiv:1303.3011](https://arxiv.org/abs/1303.3011)

# $\bar{\nu}_\mu$ disappearance (3+1 scenario)

No hint for  $\bar{\nu}_\mu$  disappearance with  $L/E \approx 1$  m/MeV

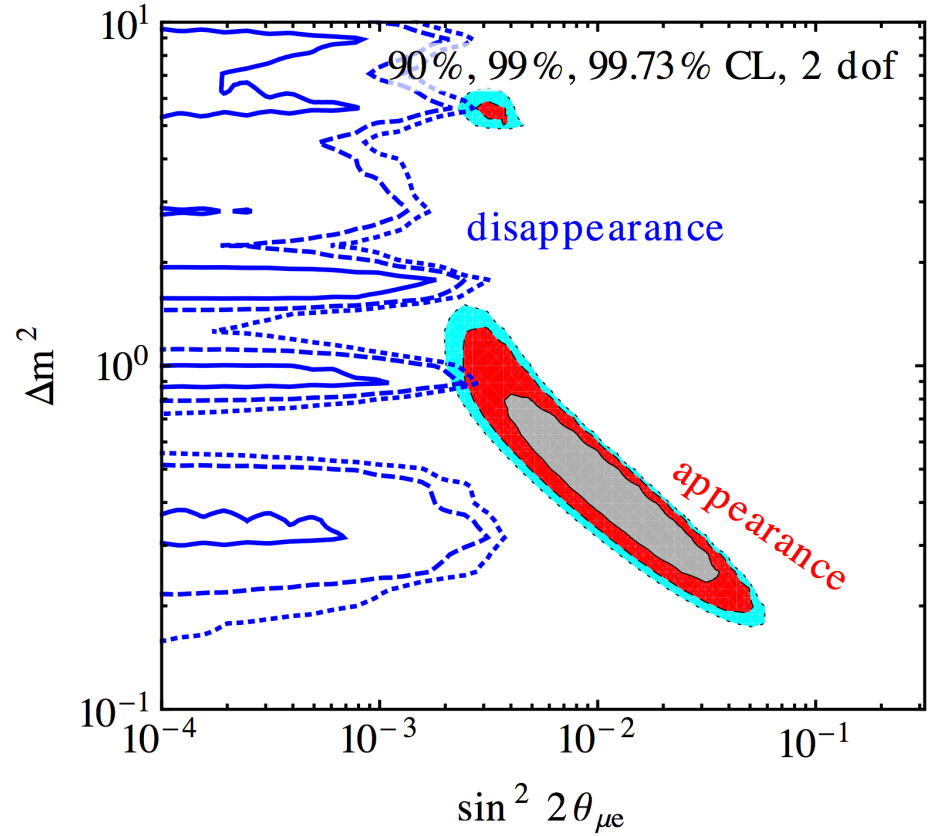


J. Kopp et al., [arXiv:1303.3011](https://arxiv.org/abs/1303.3011)

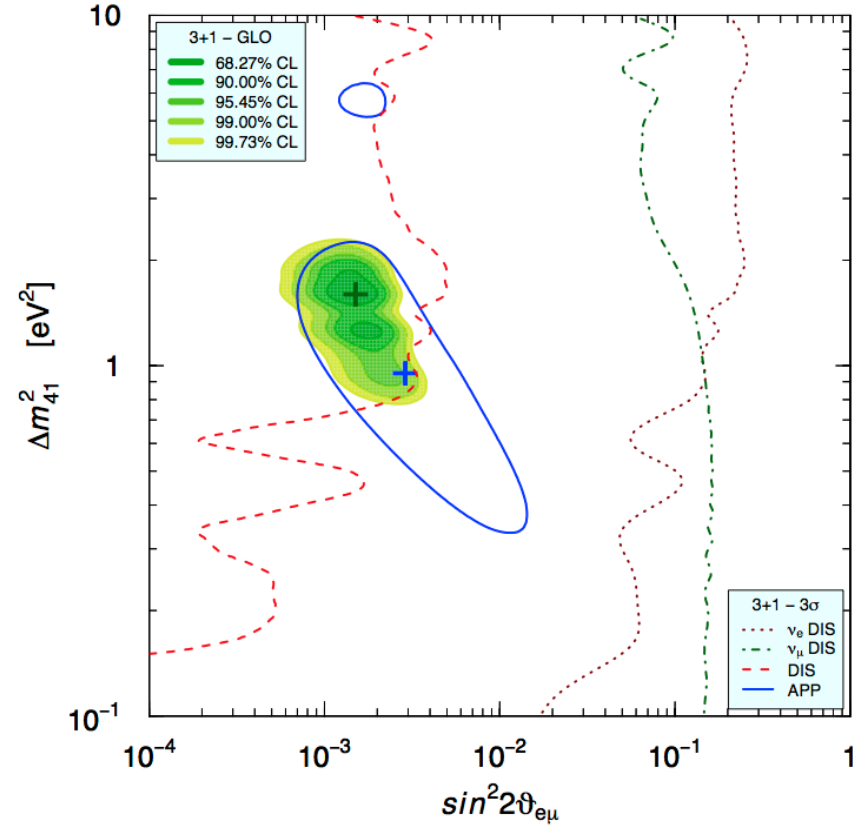
↗ New result MINOS 2014

# Appearance VS Disappearance

J. Kopp et al., arXiv:1303.3011



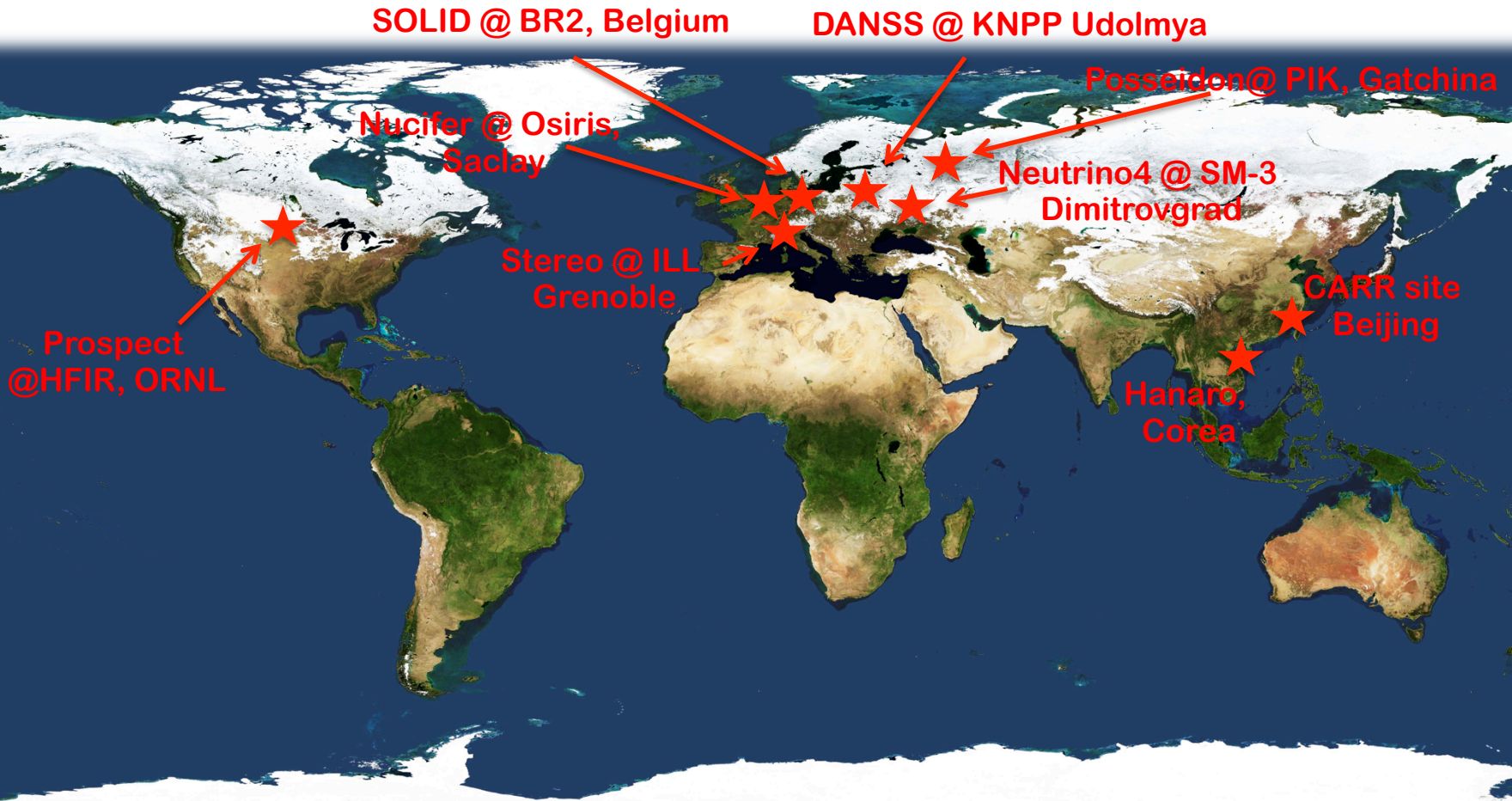
C. Giunti et al., arXiv:1308.5288



**Tension between  $\bar{\nu}_e/\nu_e$  appearance/disappearance and  $\bar{\nu}_\mu/\nu_\mu$  disappearance (3+1 & 3+2 models)**

# Experimental Prospect

# Experimental Prospect: @ Nuclear Reactor



**Test of both reactor & gallium anomalies**

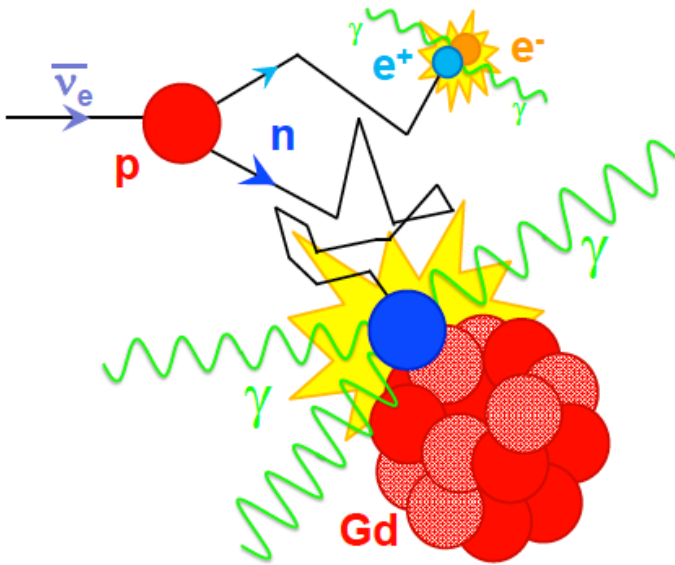
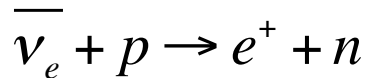
# Testing $(\bar{\nu}_e)$ disappearance anomalies

- Need robust test, beyond the current mean deviation from reactor predicted rate
- **Input from sterile neutrino fits**
  - $\Delta m^2 \approx 0.1-10 \text{ eV}^2 \rightarrow L_{\text{osc}}(\text{m}) = 2.5 \frac{E(\text{MeV})}{\Delta m^2(\text{eV}^2)} \approx 2-10 \text{ m}$
  - $\sin^2(2\theta_{ee}) \approx 0.01-0.15$
- **Experimental specifications**
  - Compact source, <1 meter scale
  - Good vertex and energy resolutions
  - High statistics (few % stat. uncertainty)
  - Few % syst. uncertainty  $\rightarrow$  low backgrounds
- **Search for a new oscillation pattern in E & L completed by normalization information**



# IBD Signal & Backgrounds

## Inverse Beta Decay



**Selective coincidence**  
**e<sup>+</sup> prompt signal & n-capture**

## Background rejection

- **Accidental  $\gamma$ -neutron coincidence**
  - Shielding
  - Segmentation
  - Neutron discrimination
  
- **Fast-n correlated background**
  - Rejection of recoil protons with PSD
  - Cosmic rays induced:
    - Reactor OFF
    - Overburden
  - Reactor induced: a killer!
    - must be negligible

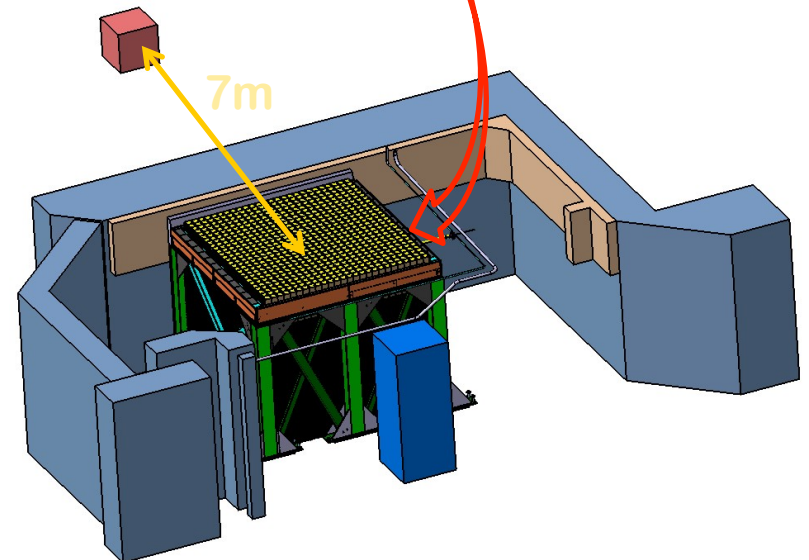
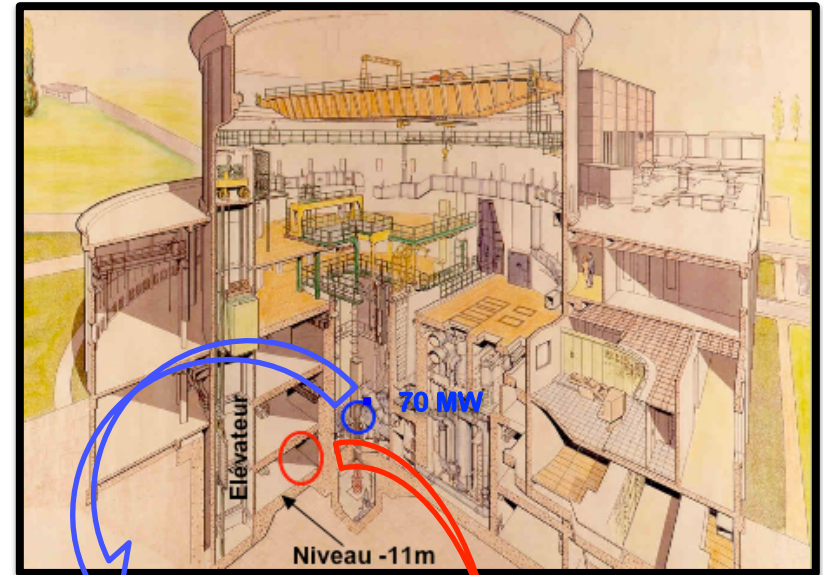
# Reactor v Proposals

Experiment Type	Projects	$P_{Th}$	$M_{det}$	L	Depth
Mature Gd-doped LS detector Technology	Nucifer (FRA)	70 MW	0.7 tons	7 m	Few mwe
	Stéréo (FRA)	50 MW	2 tons	[8-11] m	10 mwe
	Neutrino 4 (RU)	100 MW	2 tons	[6-12] m	Surf.
Highly segmented detector for background reduction	DANSS (RU)	1 GW	1 ton	[10-12] m	50 mwe
	SoLid (UK)	45-80 MW	3 tons	8 m	10 m
Enhanced neutron Tagging					
	Hanaro (KO)	30 MW	0.5 t	6 m	Few mwe
2 detector complex or Moving detector	Prospect	85 MW	-	7m & 18m	Surf.
	China project			-	
	DANSS/Neutrino4				Movable detector

# Nucifer @ OSIRIS (Gd-LS)

Originally Dedicated for non proliferation

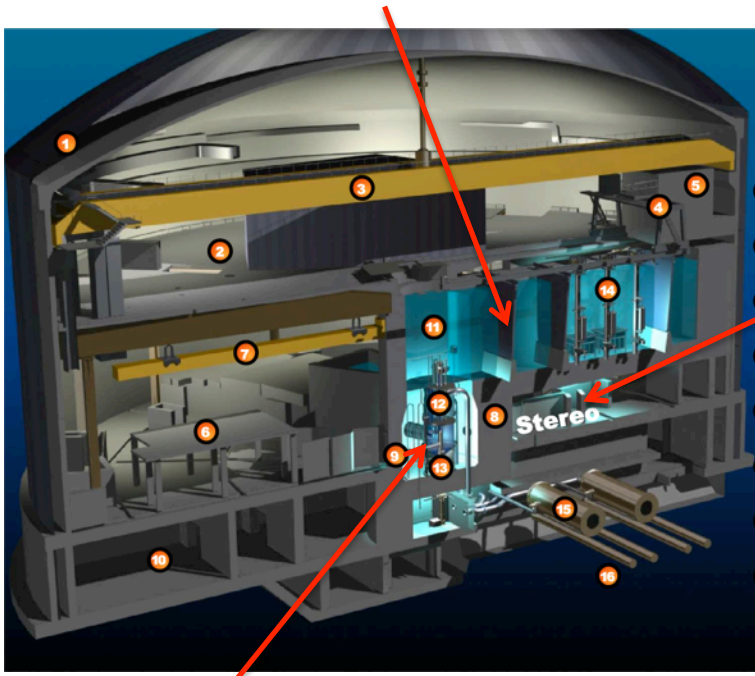
- **Osiris research reactor**
  - At Saclay, France
  - 70 MW, 20%  $^{235}\text{U}$
- **Detector designed for reactor monitoring studies**
  - 850 kg Gd-loaded LS
  - 350 int. expected / day
  - Shallow depth (few mwe)
- **Modest sensitivity to Sterile-v:**
  - Compact core: 60x60x60 cm<sup>3</sup>
  - Short baseline: only 7 m
  - Simple design
  - Challenging Reactor bkg
- **Data taking ongoing**
  - Modest sensitivity to sterile-v



# Stéréo @ ILL (Gd-LS)

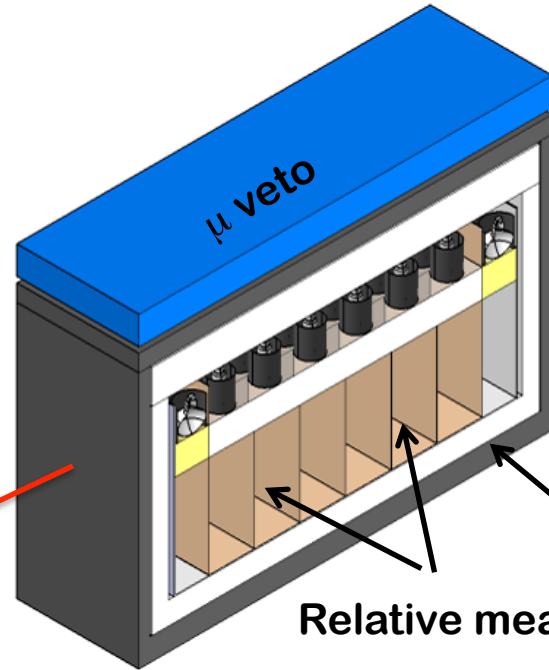
Start Data Taking in 2015

factor 4 attenuation of vertical flux  
from water pool



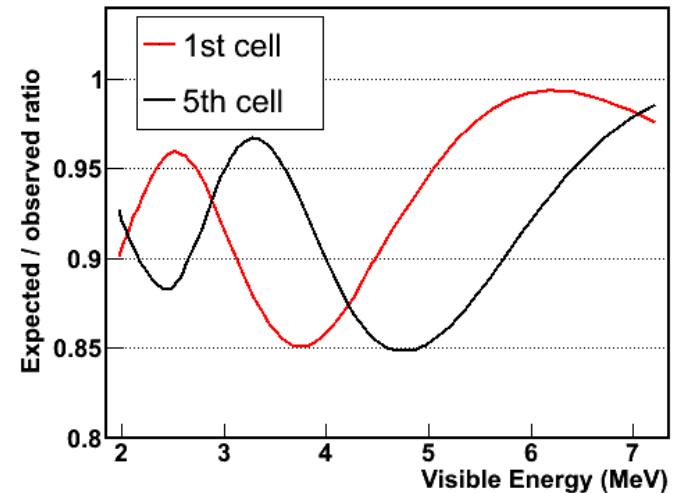
50 MW core  
h=80cm,  $\Phi=40$ cm

[8.5-11] m  
baseline range



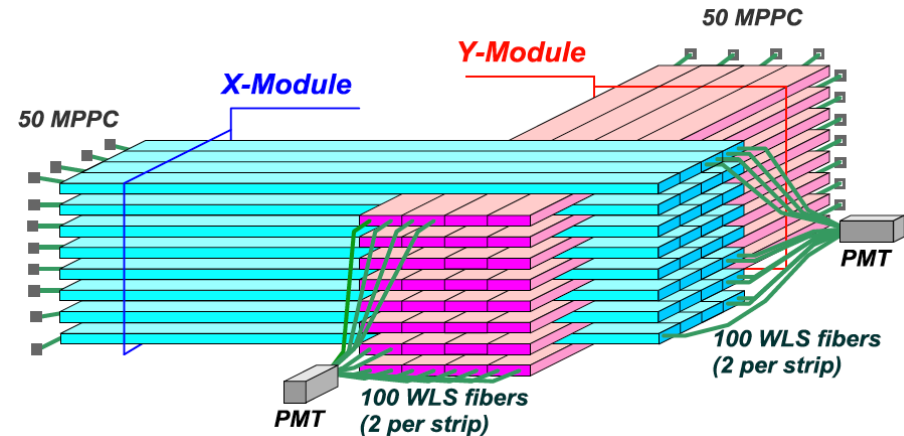
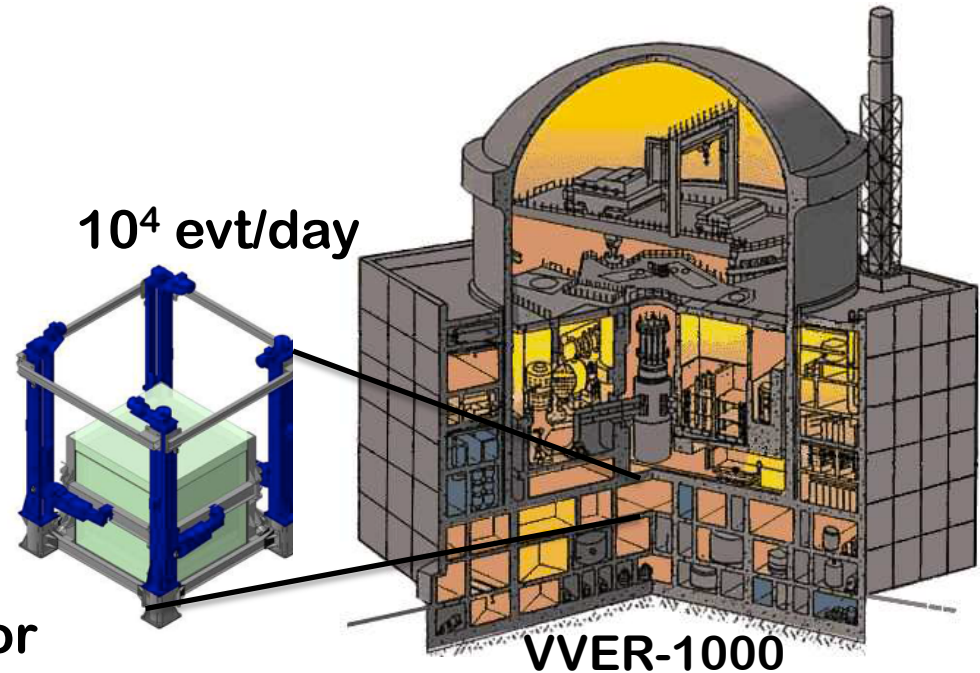
70 tons  $\gamma$  and n  
shielding

Relative measurement in 6 cells



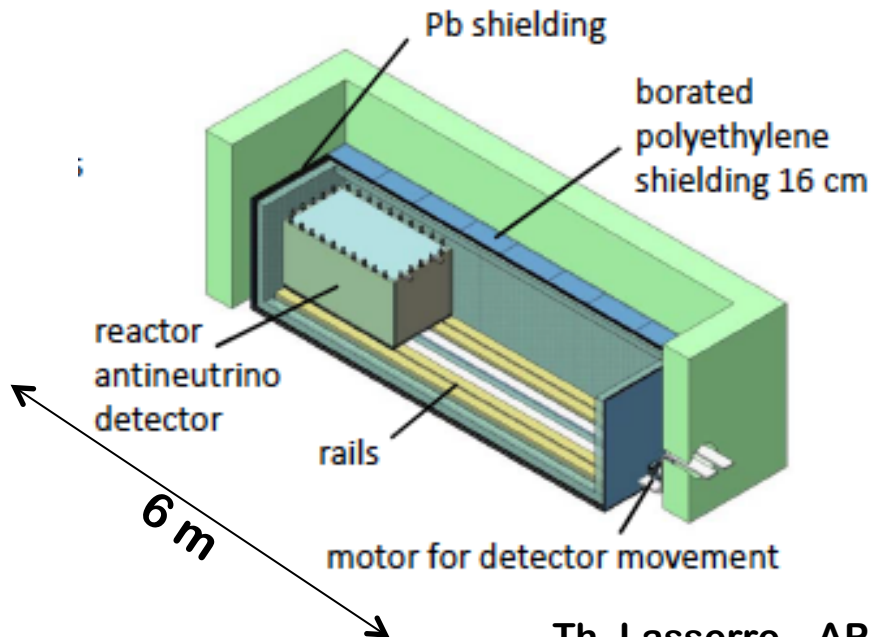
# DANSS @ KNPP (High-Seg)

- 1 GW extended core
- Good overburden
- Vertical motion of the detector (9.7-12.2 m)
- Highly segmented detector  
→ background rejection
- Plastic strips with Gd-loaded interlayer, WLS fibers readout
- Start in 2014/15?



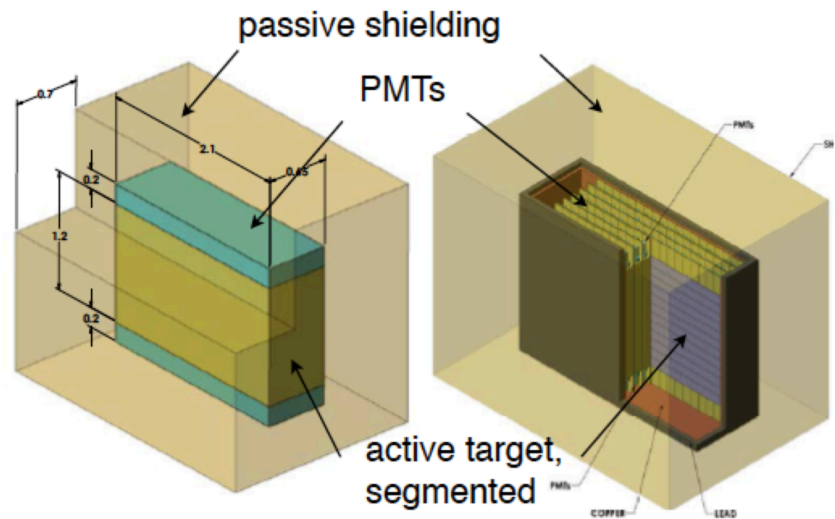
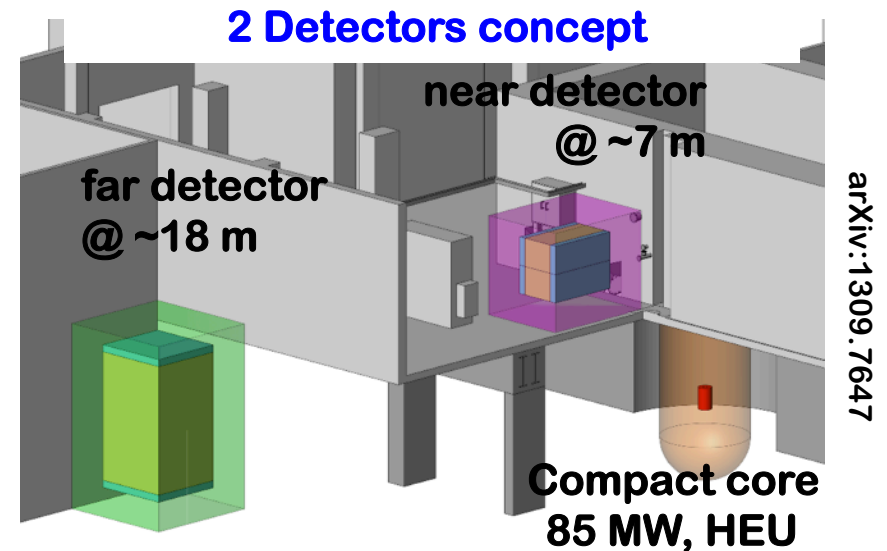
# Neutrino-4 @ SM3 (Gd-LS)

- 2.5 m<sup>3</sup> LS target, 5 section movable detector [6-12] m
- 100 MW compact core
- Detector at Surface
- Status:
  - Proto: On/Off v-data
  - Shielding integrated
  - Start in 2015



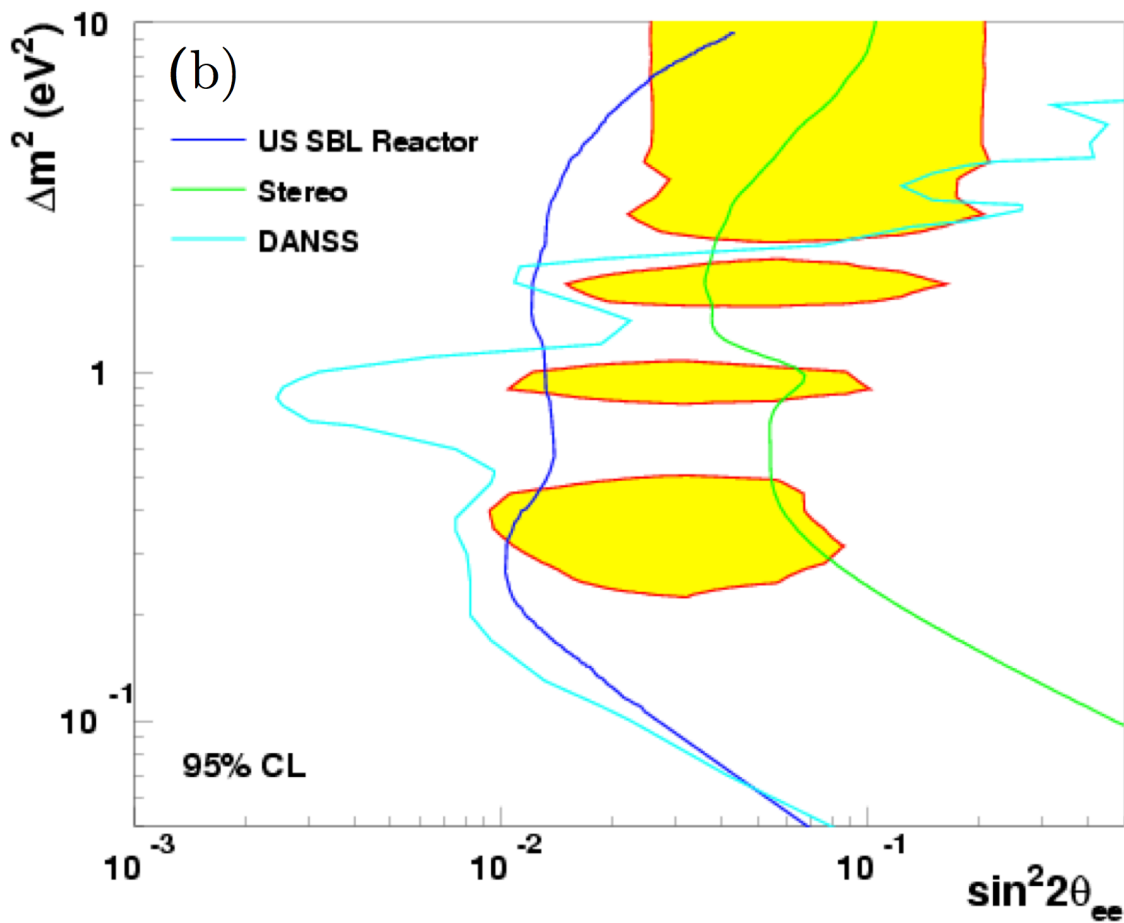
# Prospect @HFIR

- 3 reactor sites
  - HFIR – 85 MW
- 7-18 m baselines
- Surface location
- Detector
  - Segmented
  - $^6\text{Li}$ -doped for n-tagging or Gd-loaded?
- Status:
  - Site characterization
  - R&D
  - Start 2016?



# Reactor Experiment Sensitivity

All current projects have the sensitivity to test the reactor anomaly space of parameters,  $\Delta m^2 > 0.1 \text{ eV}^2$ ,  $\sin^2 2\theta > 0.05$

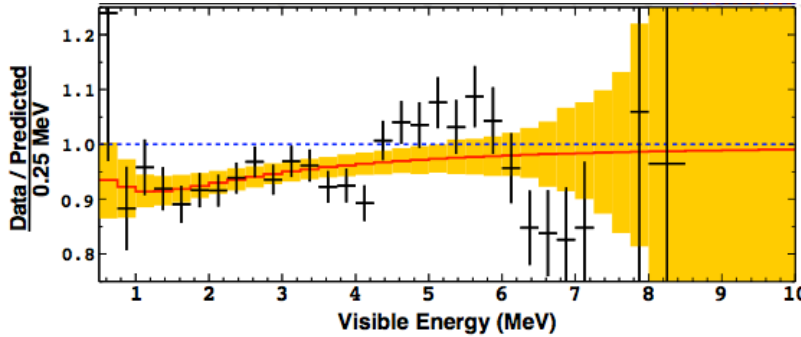




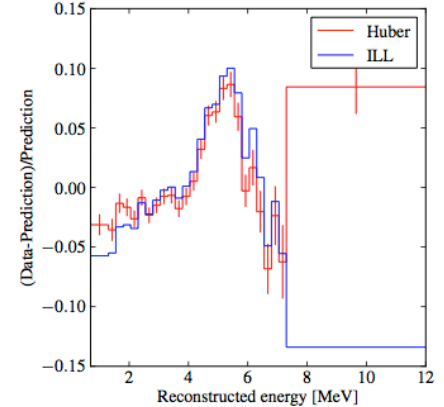
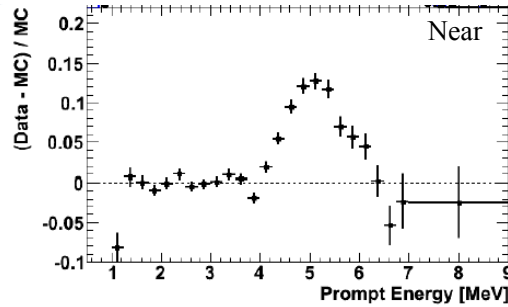
# A 'new' 1-2% n Excess at 5 MeV

DB, PhD, now

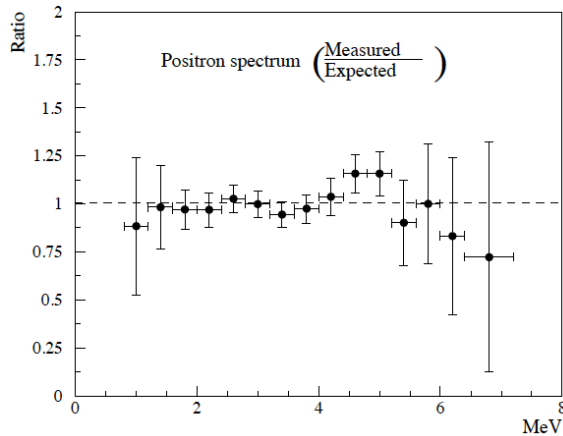
Double Chooz, May 22<sup>nd</sup> 2014



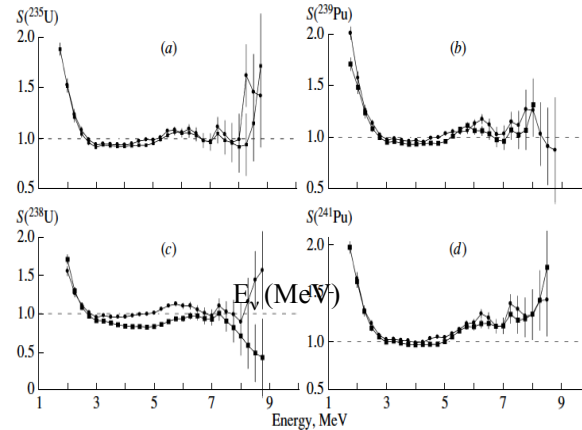
RENO, Neutrino 2014



CHOOZ, PLB466 (1999) 415-430



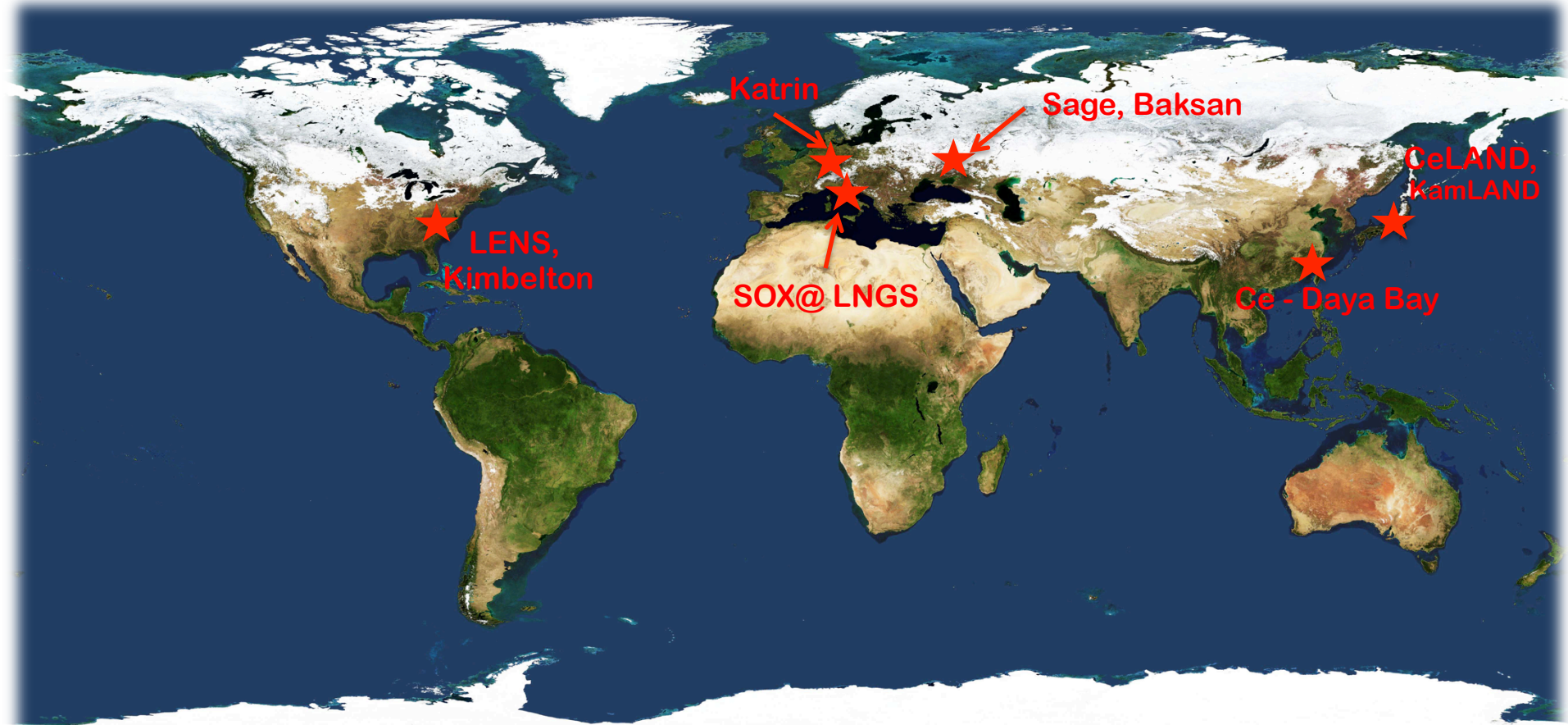
Rovno, arXiv:1207.6956



**Origin to be understood (fluxes or e-ν conversion, new interaction?)**

**Relative measurements: modest impact on sterile-ν search**

# Experimental Program: @ Neutrino Generator



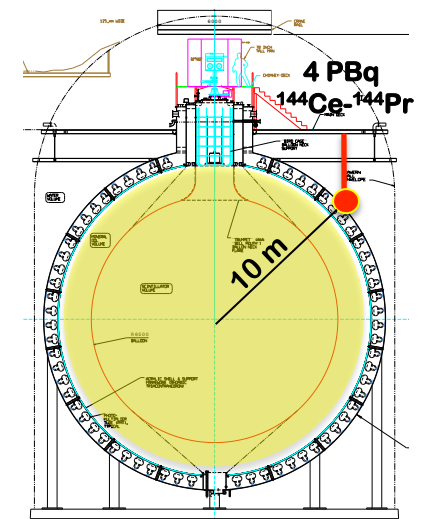
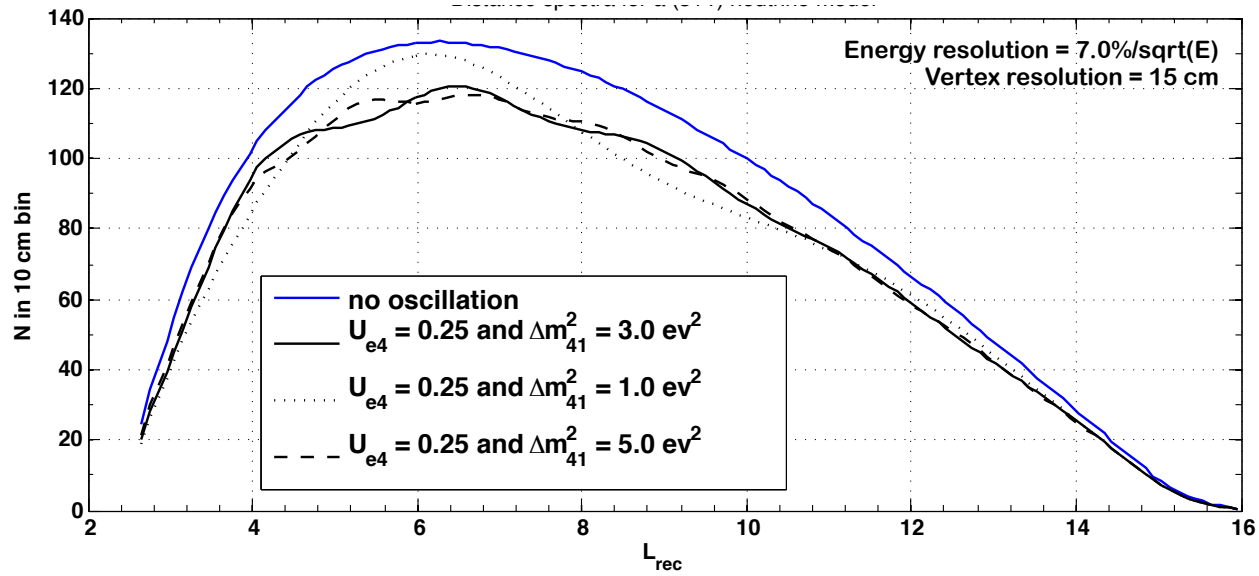
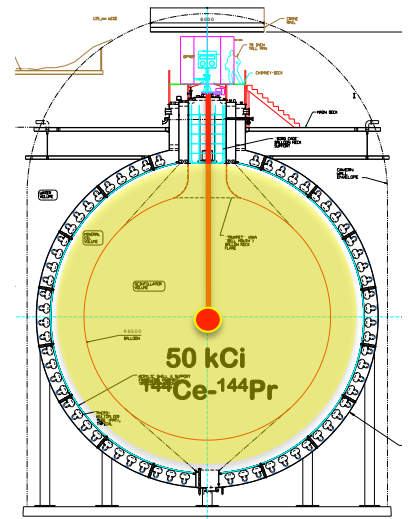
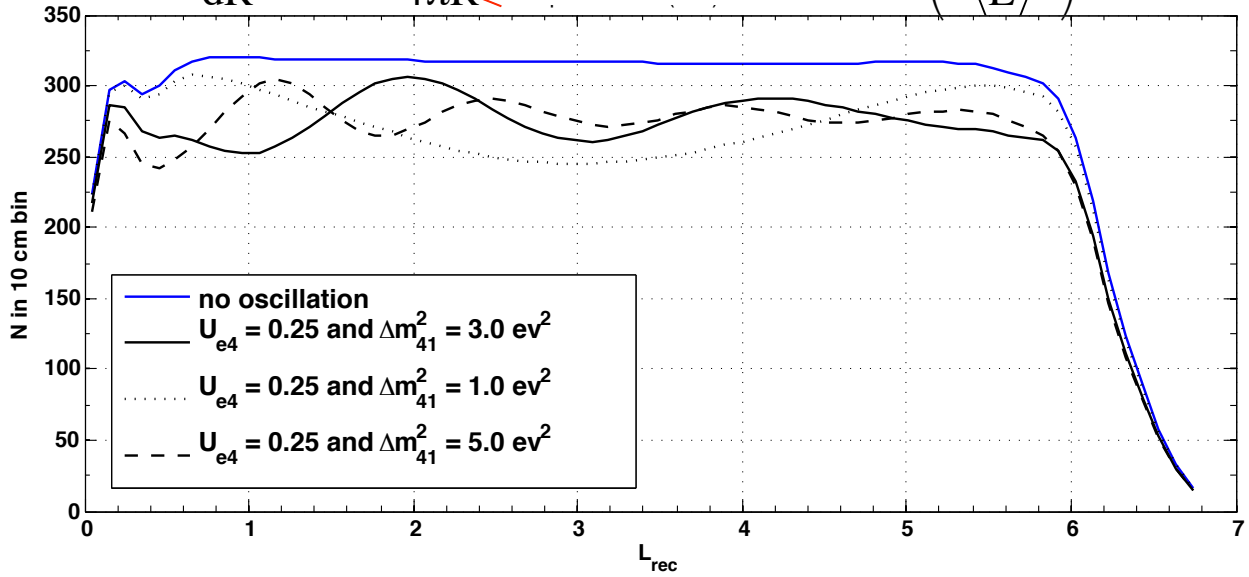
**Test of both reactor & gallium anomalies**

# $\nu$ Generator Proposals

Type	Detection	Background	Isotope	Production	Activity	Projects
$\nu_e$	$\nu_e e \rightarrow \nu_e e$ 5% $E_{res}$ 15cm $R_{res}$  or Radio-chemical	Detector Radioactivity  Solar $\nu$ (irreducible)  $\nu$ generator impurities	$^{51}\text{Cr}$ 0.75 MeV $t_{1/2}=26\text{d}$	$n_{th}$ irradiation in Reactor	>3 MCi	Sage LENS
					>10 MCi	SOX-Cr (SNO+)
			$^{37}\text{Ar}$ 0.8 MeV $t_{1/2}=35\text{d}$	$n_{fast}$ irradiation in Reactor (breeder)	>1 MCi	-
$\bar{\nu}_e$	$\bar{\nu}_e p \rightarrow e^+ n$ $E_{th}=1.8\text{ MeV}$  ( $e^+, n$ )  5% $E_{res}$ 15cm $R_{res}$	reactor $\nu$ , geo $\nu$ ,  $\nu$ generator impurities	$^{144}\text{Ce}$ $E < 3\text{MeV}$ $t_{1/2}=285\text{d}$	spent nuclear fuel reprocessing + REE extraction	75 kCi	CeLAND Ce-SOX
					500 kCi	Daya-Bay
			$^{90}\text{Sr}$ $^{106}\text{Rh}$		-	-
	$^3\text{H} \rightarrow \text{He} e^- \bar{\nu}_e$ EC/ $\beta$ -decay	Kink search	$^3\text{H}$ $E < 18\text{ keV}$	Irradiation in reactors	3 Ci	KATRIN (Mare/Echo)

# Search for $\bar{\nu}_e \rightarrow \bar{\nu}_s$ with $^{51}\text{Cr}/^{144}\text{Ce}$

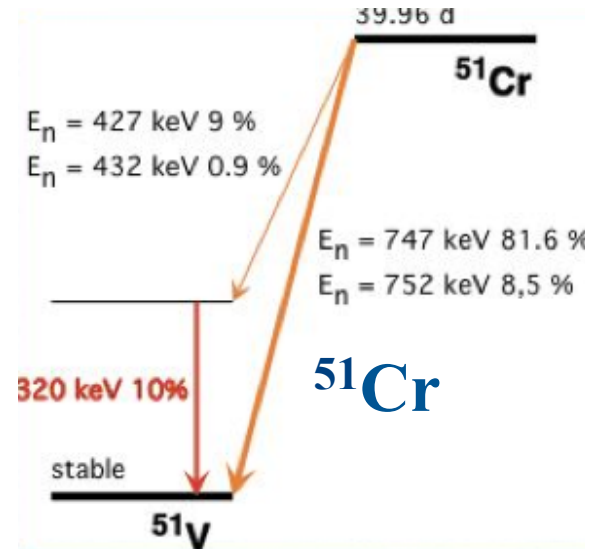
$$\frac{dN}{dR}(R,t) \propto \frac{A(t)}{4\pi R^2} \times \langle \sigma \rangle \times N_p \times \cancel{4\pi R^2} \times P_{ee} \left( \frac{\Delta m^2 R}{\langle E \rangle} \right)$$



# $^{51}\text{Cr}$ neutrino generator

erc

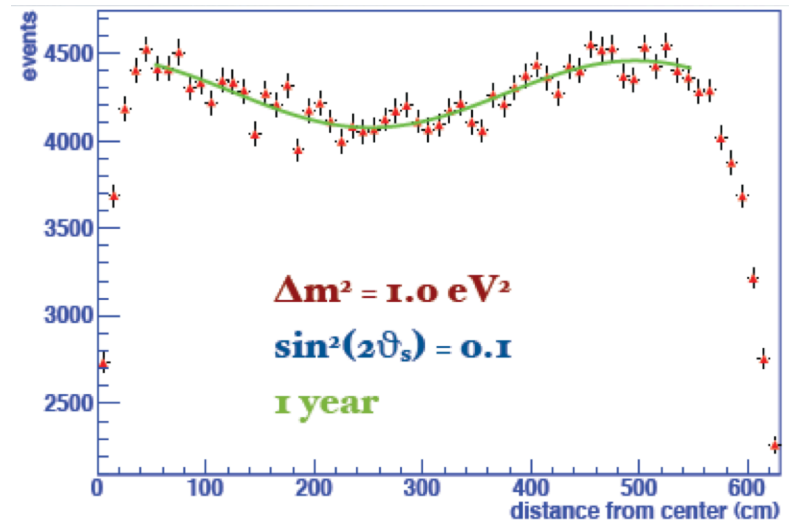
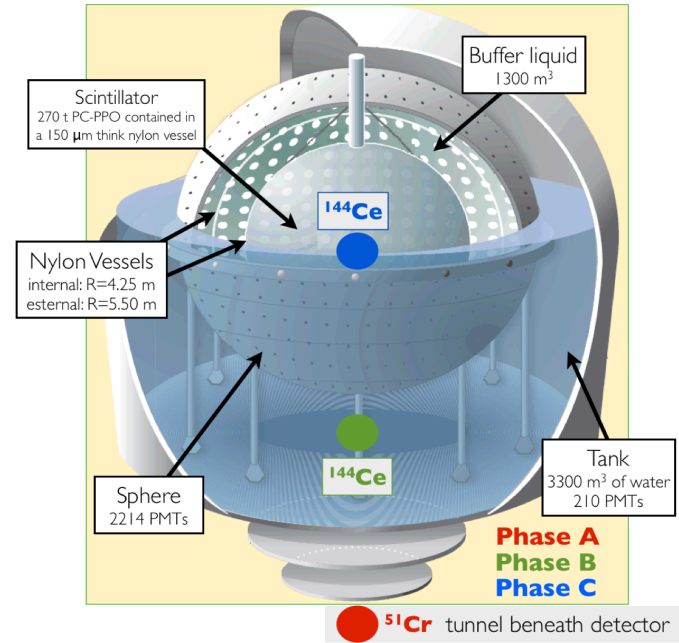
- **$^{51}\text{Cr}$  EC**
  - $E = 0.75 \text{ MeV}$
  - $t_{1/2} = 26 \text{ days}$
  
- **Production** through  $n_{\text{th}}$  irradiation of enriched  $^{50}\text{Cr}$  in a nuclear reactor
  
- **Need 10 MCi  $^{51}\text{Cr}$** 
  - 2 MCi in Gallex/Sage
  
- **Detection:**
  - $^{71}\text{Ga} + \nu_e \rightarrow ^{71}\text{Ge} + e^-$
  - $\nu$  scattering off electrons (SOX)



# $^{51}\text{Cr}$ : SOX (Borexino)

erc

- Re-use **Gallex 36 kg** of enriched chromium
- Production reactors
  - Oak Ridge (US) ?
  - Ludmila (Ru) ?
- Source **8.25 m** from center
- **Detection as for  $^7\text{Be}$  solar  $\nu$** 
  - Well known background
- Status:
  - R&D for irradiation (need 2 x 5 MCi)
- Staged approach: after  $^{144}\text{Ce}$



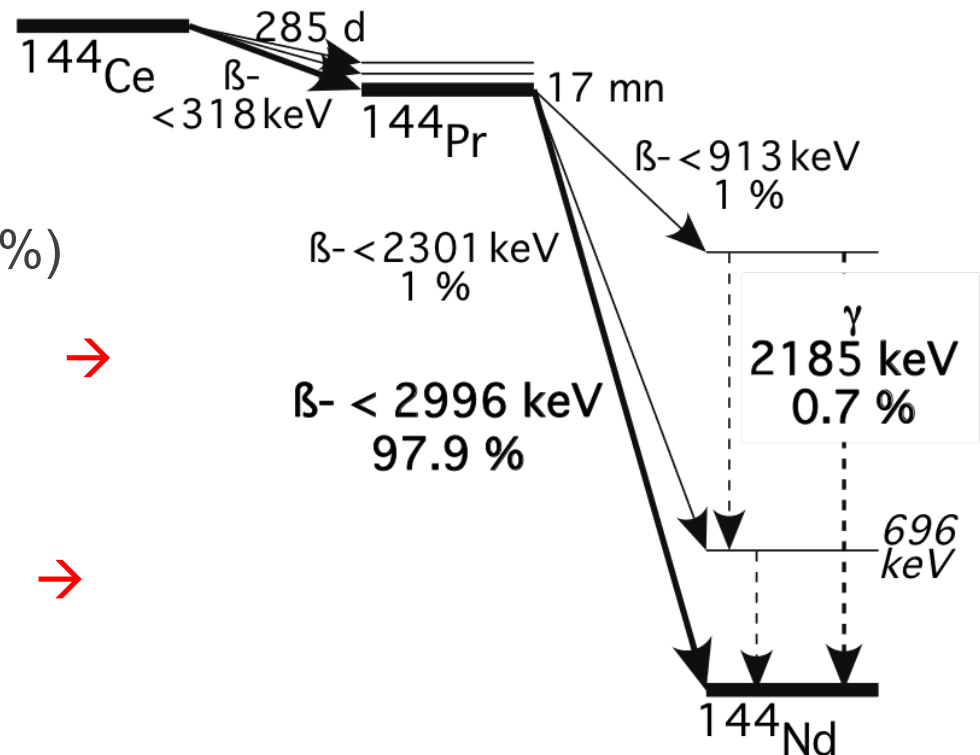
# $^{144}\text{Ce}-^{144}\text{Pr} \bar{\nu}$ generator

erc

- 1<sup>st</sup> Trick:  $\bar{\nu}_e$  source detected via  $\bar{\nu}_e + p \rightarrow e^+ + n$  (Thr=1.8 MeV)
  - High IBD cross section  $\rightarrow$  few PBq activity (100 kCi)
  - $(e^+, n)$  detected in coincidence  $\rightarrow$  Strong background reduction

## 2<sup>nd</sup> Trick: $^{144}\text{Ce}-^{144}\text{Pr}$

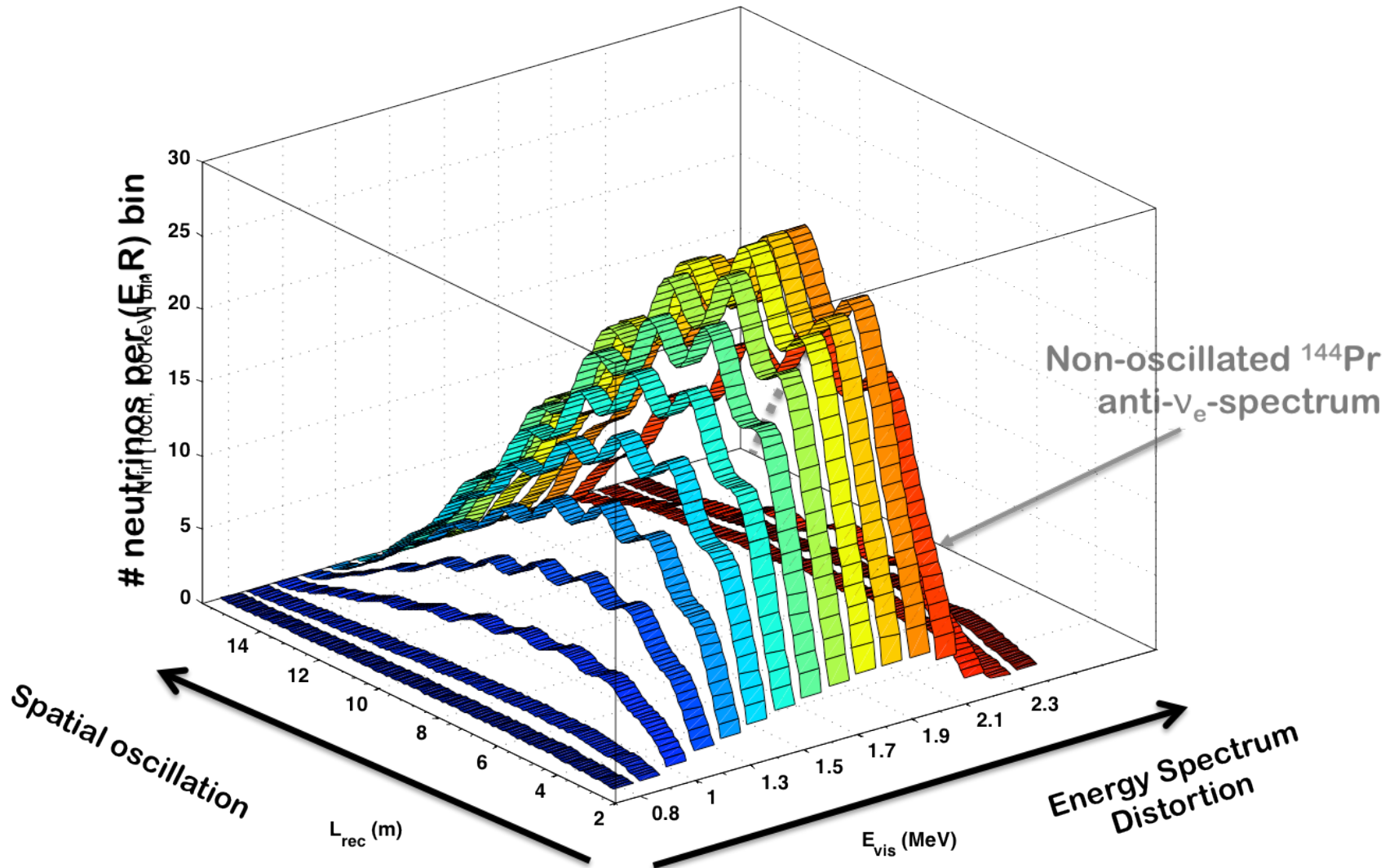
- Abundant fission product (5%)
- $^{144}\text{Ce}$ : long-lived & low- $Q_\beta$   $\rightarrow$  Enough time to produce, transport, use
- $^{144}\text{Pr}$ : short-lived & high- $Q_\beta$   $\rightarrow$   $\bar{\nu}_e$ -emitter above threshold



# $^{144}\text{Ce}-^{144}\text{Pr}$ Signal

75 kCi  $^{144}\text{Ce}-^{144}\text{Pr}$  – 9.3 m from detector center – 1.5 year

2-D reconstructed spectrum for  $U_{e4} = 0.25$  and  $\Delta m_{41}^2 = 3.0 \text{ eV}^2$

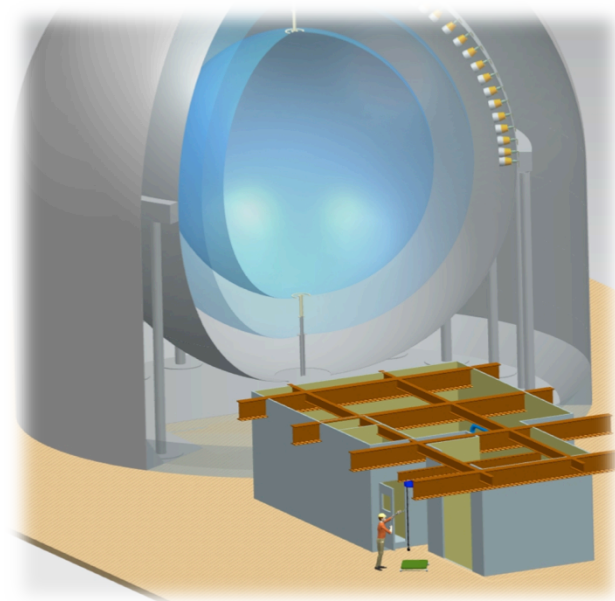
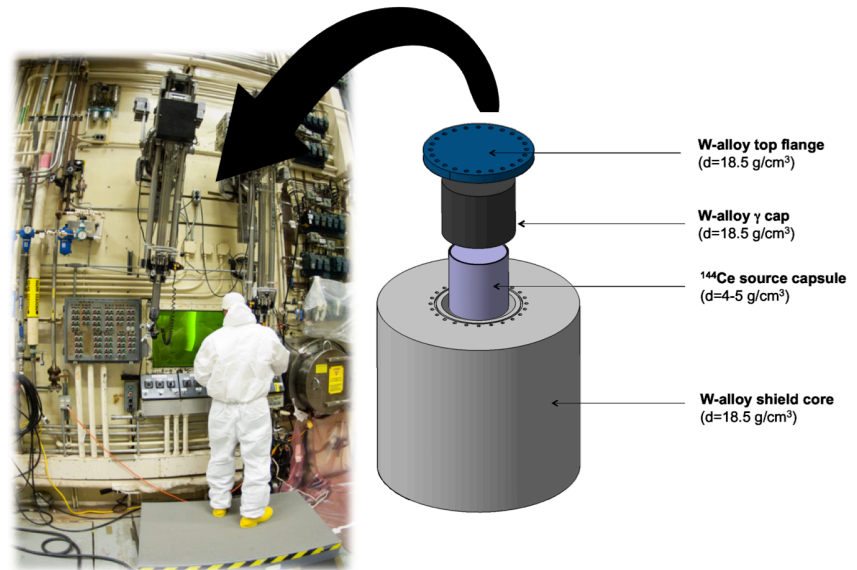




# $^{144}\text{Ce}$ - $^{144}\text{Pr}$ : CeSOX in BX

erc

- 4 PBq of  $^{144}\text{Ce}$ - $^{144}\text{Pr}$  ( $\text{CeO}_2$ )
- **Production feasible at Mayak Facility (RU) in 2014/5 (1 y)**
  - Standard SNF reprocessing
  - Ce extraction through displacement chromatography
- **Need 19 cm tungsten-shield**
- **Borexino being prepared**
  - Tunnel below the detector
  - 8.25 m from center
- **Deployment in 11/2015**
  - 1.5 y data taking
  - 10 000 interactions expected

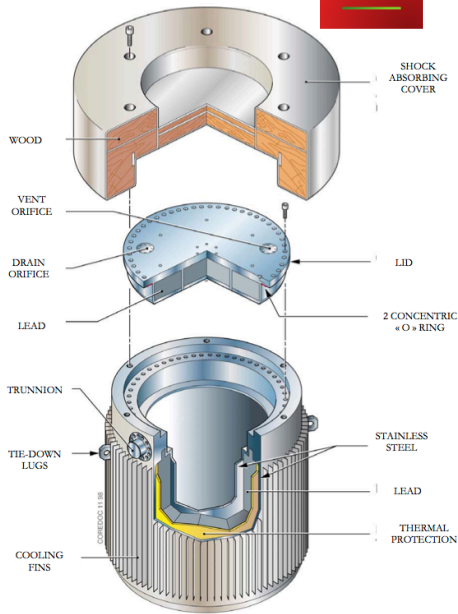


## IAEA rules on Safe Transportation of Radioactive Material

### A) Suitable certified transport container: licensing ongoing

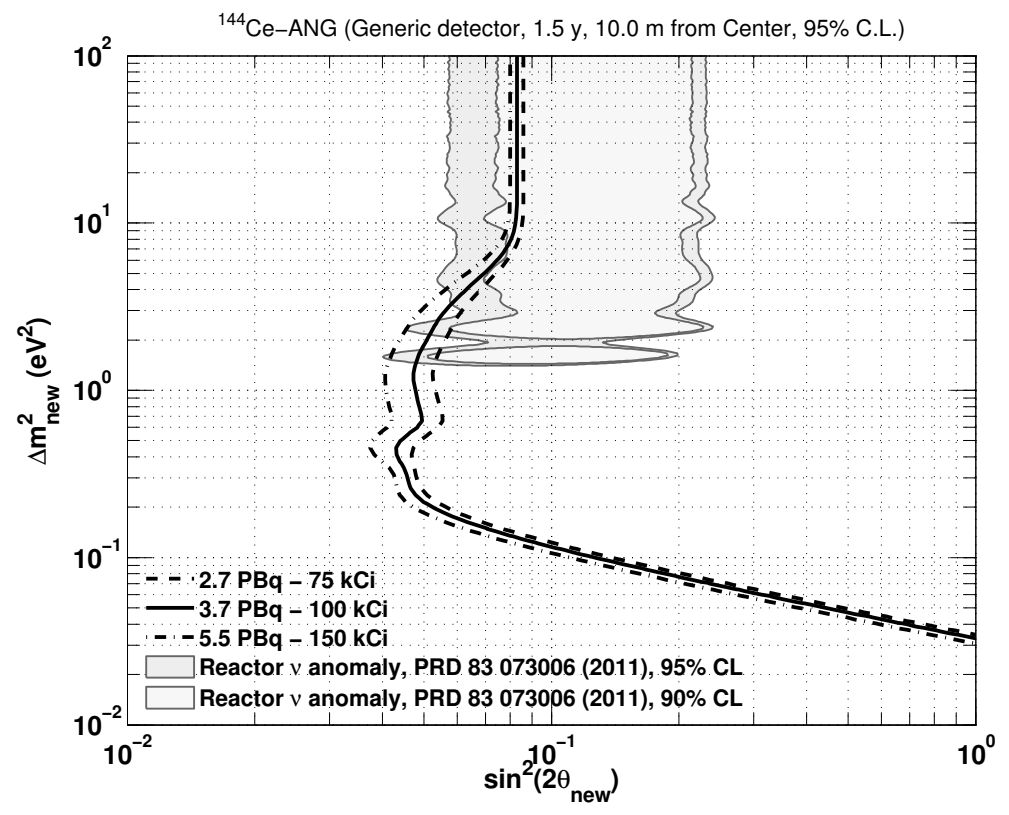
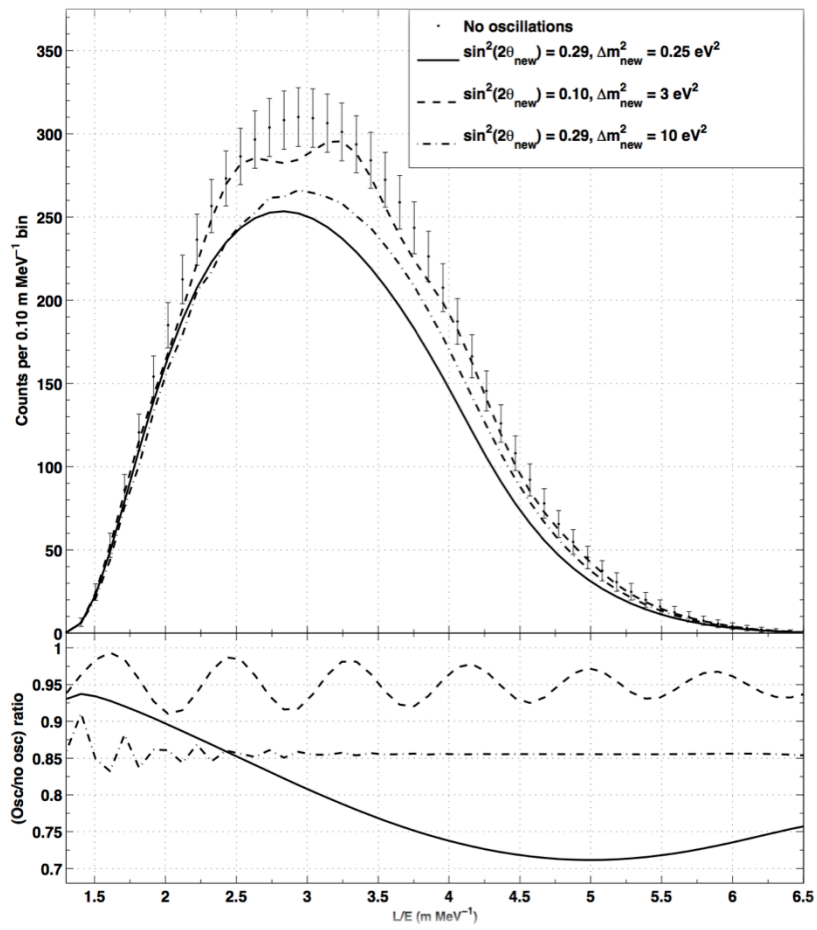
suitable B(U) casks identified

TN - TM - MTR - 23 t



### B) Route: Mayak – St Petersburg – Le Havre – Gran Sasso

# CeSOX sensitivity



# Search for $\nu_s$ with ${}^3\text{H}$ $\beta$ decay

- Source:  ${}^3_1\text{H} \rightarrow {}^3_2\text{He} + e^- + \bar{\nu}_e$

- $\beta$  spectrum shape depends on:

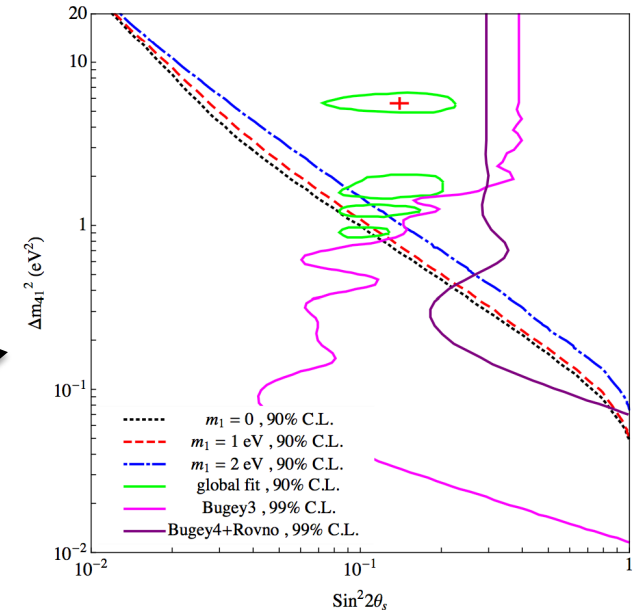
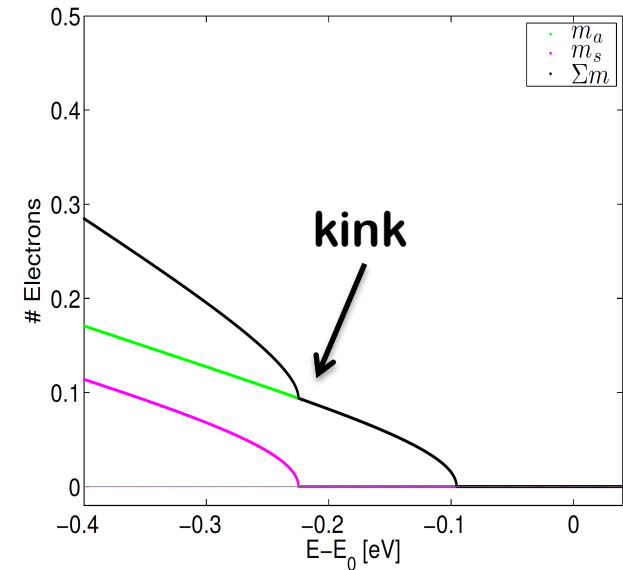
$$\langle m_\beta \rangle = \sqrt{\sum_{1,2,3,\dots} |U_{ei}|^2 m_i^2}$$

- Hypothetical 4<sup>th</sup>  $\nu$  contribution

$$\langle m_\beta \rangle_4 = |U_{e4}| \sqrt{\Delta m_{41}^2}$$

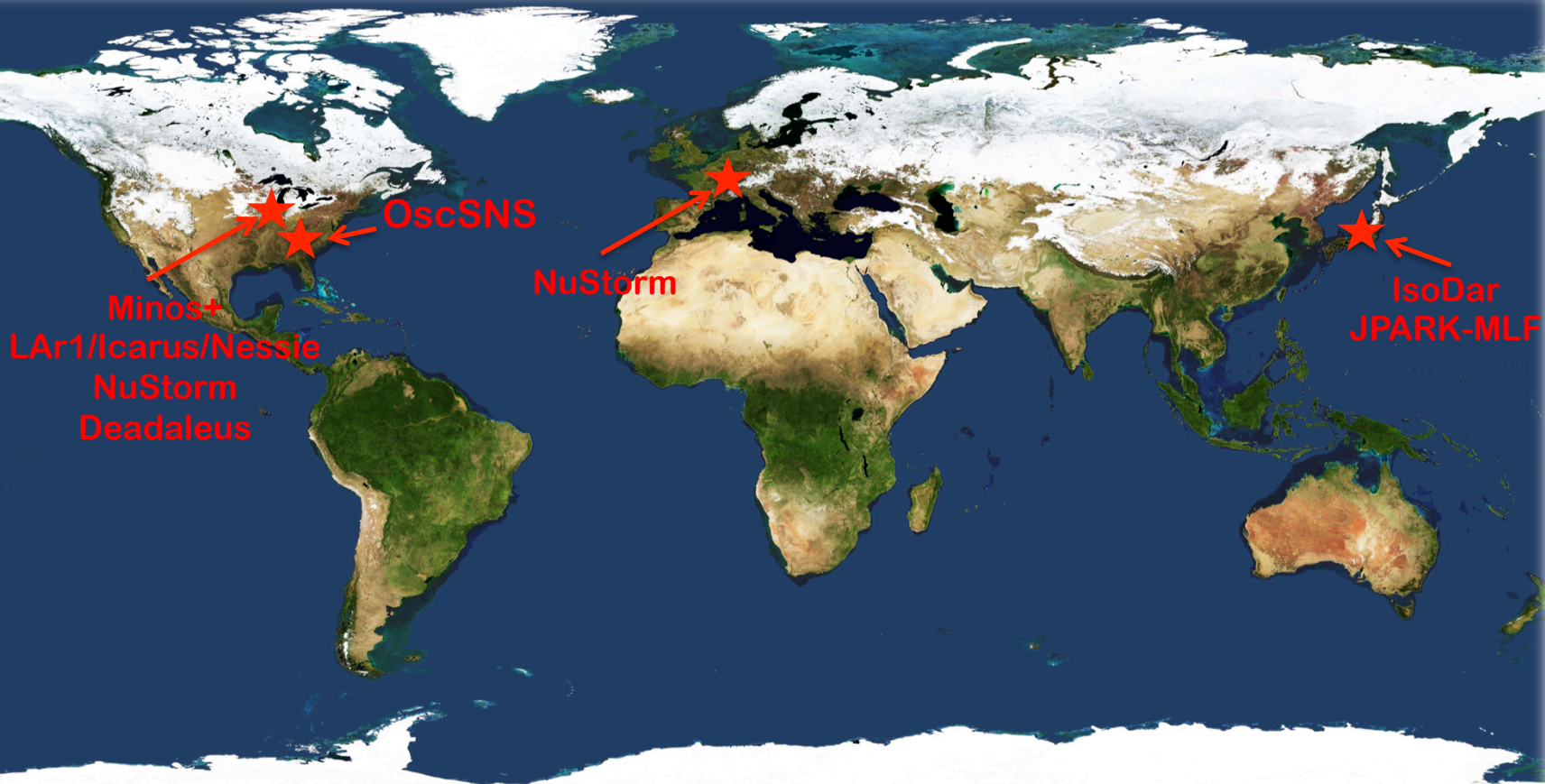
→ Search for a kink few eV below end point

- KATRIN –as designed– can test the  $\nu_e$  disappearance anomalies (sensitivity to be assessed with syst.)



# Experimental Program:

## @ Neutrino Beam



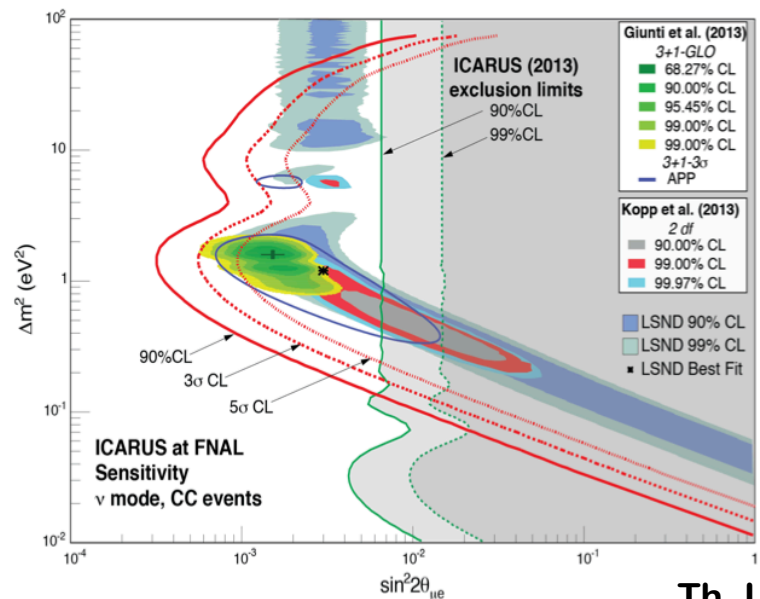
**Test of LSND/MinibooNE/reactor/gallium anomalies**  
**If positive signal, detailed study of sterile- $\nu$  phenomenology**

# $\nu$ Beam Proposals

Type	Source	App. /Dis.	Oscillation Channels	Projects
<b>Isotope Decay at Rest</b>	$p + {}^9\text{Be} \rightarrow {}^8\text{Li} + 2p$ $n + {}^7\text{Li} \rightarrow {}^8\text{Li}$ ${}^8\text{Li} \rightarrow {}^9\text{Be} + e^- + \bar{\nu}_e$	Dis.	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	IsoDAR
<b>Pion (Kaon) Decay at Rest</b>	$\pi^+ \rightarrow \mu^+ \nu_\mu$ $\quad \quad \quad \searrow$ $\quad \quad \quad e^+ \bar{\nu}_\mu \nu_e$	App. & Dis.	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_e \rightarrow \nu_e$	OscSNS, DAE $\delta$ ALUS, KDAR, JPARC-MLF
<b>Pion Decay in Flight</b>	$\pi^+ \rightarrow \mu^+ \nu_\mu$ $\quad \quad \quad \searrow$ $\quad \quad \quad e^+ \bar{\nu}_\mu \nu_e$	App. & Dis.	$\nu_\mu \rightarrow \nu_e$ $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_\mu \rightarrow \nu_\mu$ $\nu_e \rightarrow \nu_e$	MINOS+, MicroBooNE, LAr1kton Icarus/Nessie
<b>Low-E Neutrino Factory</b>	$\mu^+ \rightarrow e^+ \bar{\nu}_\mu \nu_e$ $\mu^- \rightarrow e^- \nu_\mu \bar{\nu}_e$	App. & Dis.	$\nu_e \rightarrow \nu_\mu$ $\bar{\nu}_e \rightarrow \bar{\nu}_\mu$ $\nu_\mu \rightarrow \nu_\mu$ $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$	$\nu$ STORM

# Pion Decay in Flight $\nu$ -sources

- Move T600 from LNGS to CERN for rehauling and then to FNAL
- Build new near detector
- Add magnetic spectrometers (Nessie)



# Pion Decay at Rest $\nu$ -sources

- **High Energy Proton source**

- Each  $\pi^+$  decay

- $\nu_\mu, \nu_e, \bar{\nu}_\mu$

- known E spectrum

- **Detection channels**

- $\nu_e \rightarrow \nu_e$  Disappearance

- $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$  Appearance

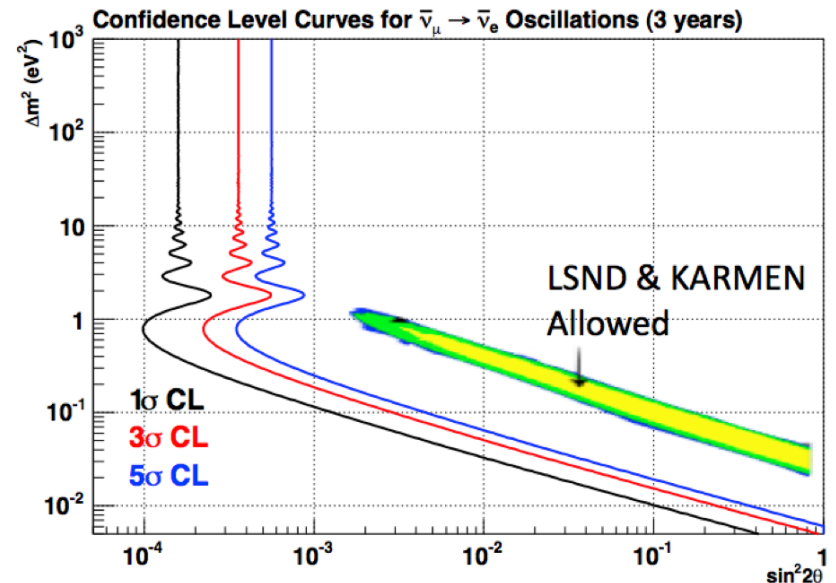
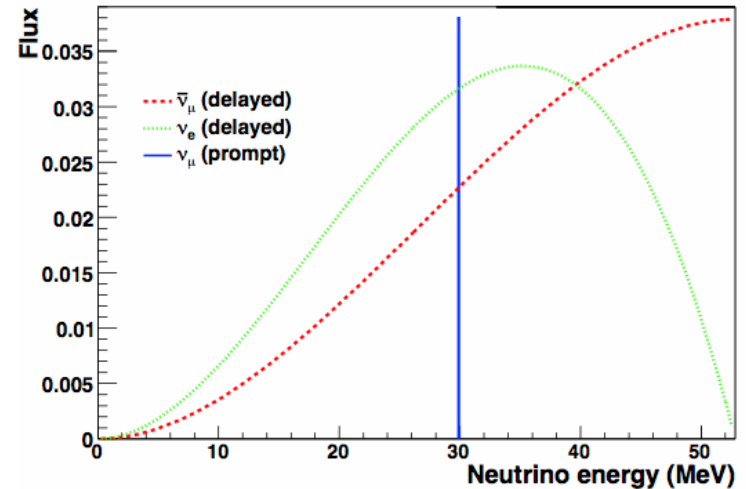
- **Direct Test of LSND**

- **OscSNS (ORNL, 1.4 MW)**

- 800t LS-det @ 60 m

- **JPARC-MLF**

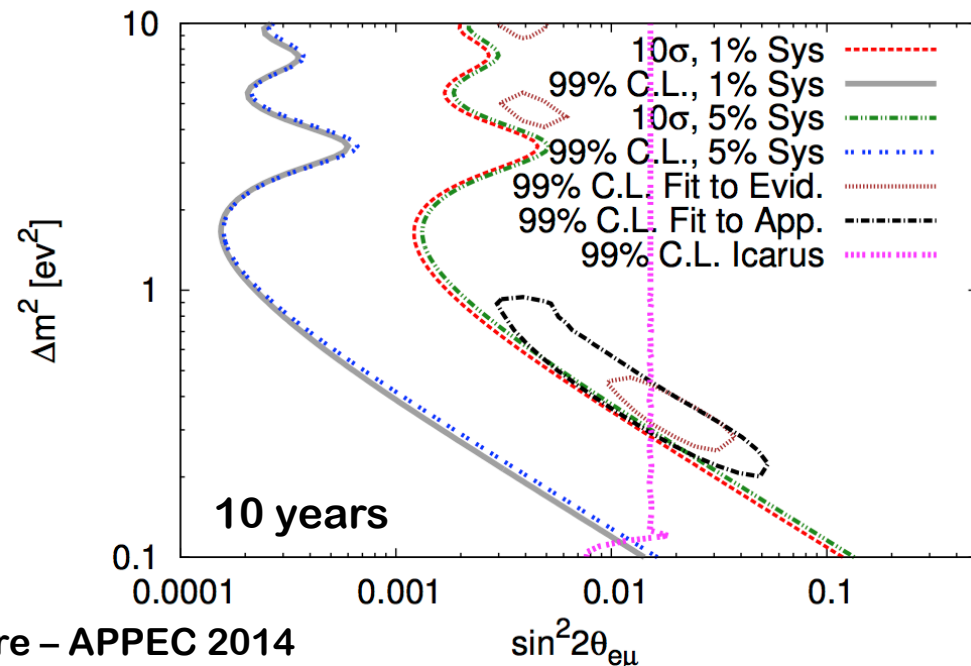
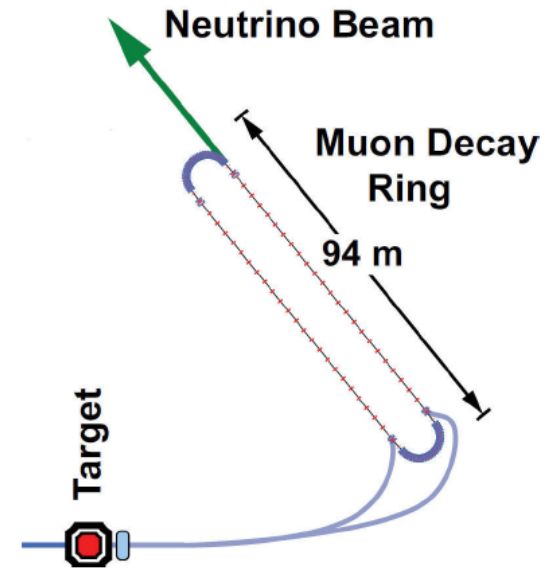
- 2x25ton Gd-LS-det @17 m





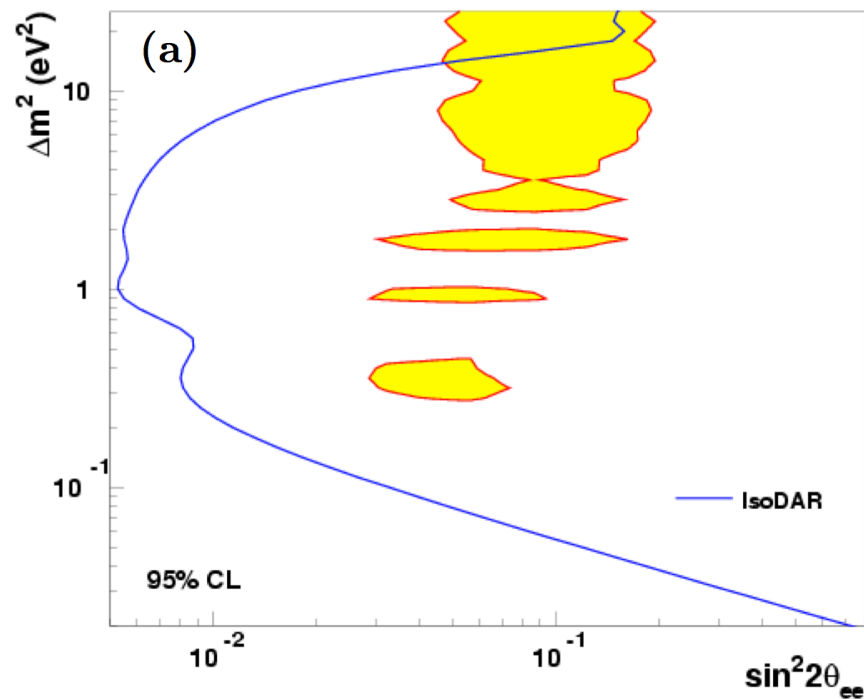
# Muon Decay Rings: $\nu$ -STORM

- **Neutrino Factory Concept**
  - 60 GeV protons on solid target
  - Horn capture and  $\pi$  transfer
  - Muon Decay ring
  
- **APP and DIS channels with:**
  - $(\bar{\nu})_{\mu}, (\bar{\nu})_e$
  
- **1.3 kT-scale Minos-like**
  - 2 km baseline
  - Near detector
  
- **Golden Mode**
  - $(\bar{\nu})_{\mu}$  APP in a  $(\bar{\nu})_e$  beam
  
- **Definitive sterile  $\nu$  search**

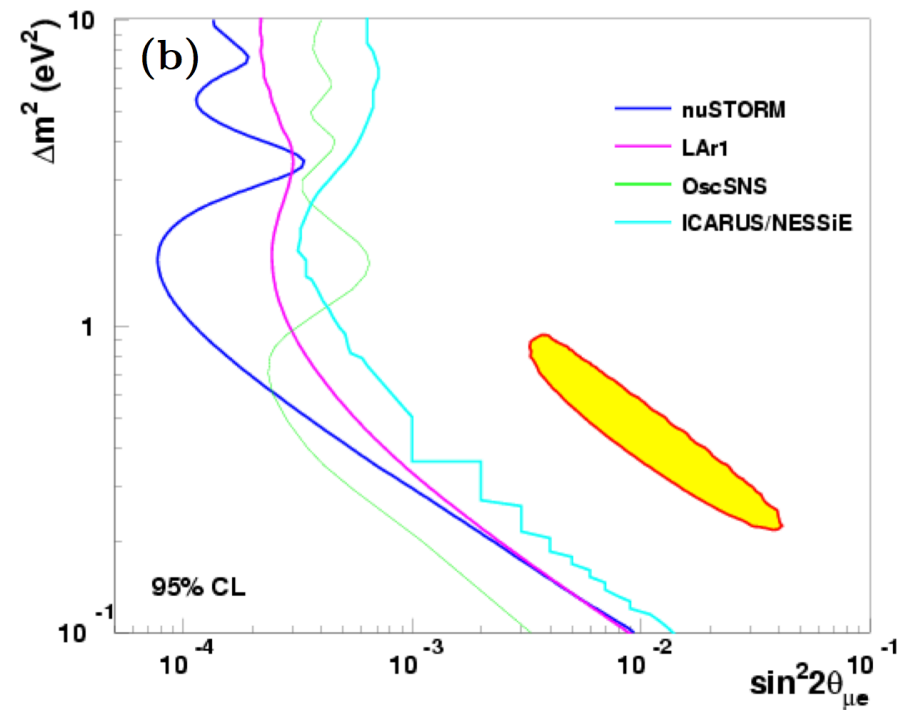


# Beam Experiment Sensitivities

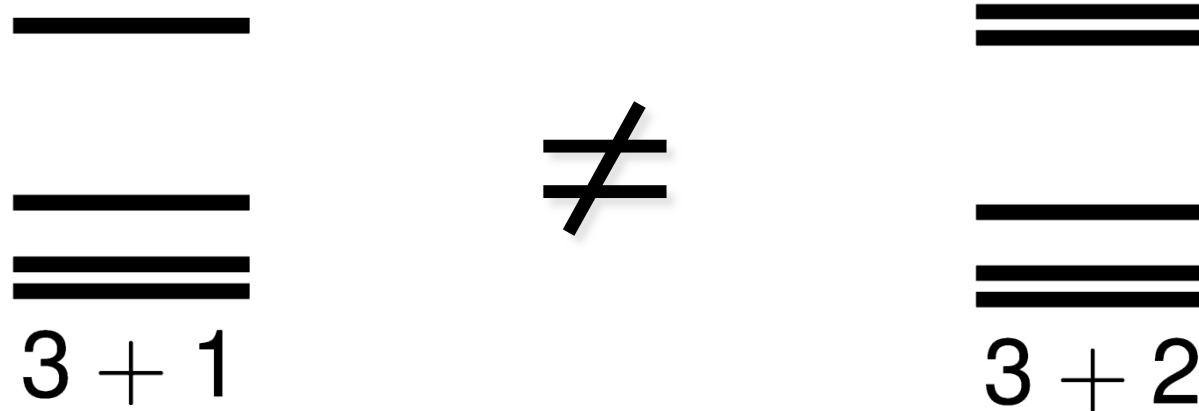
## Disappearance



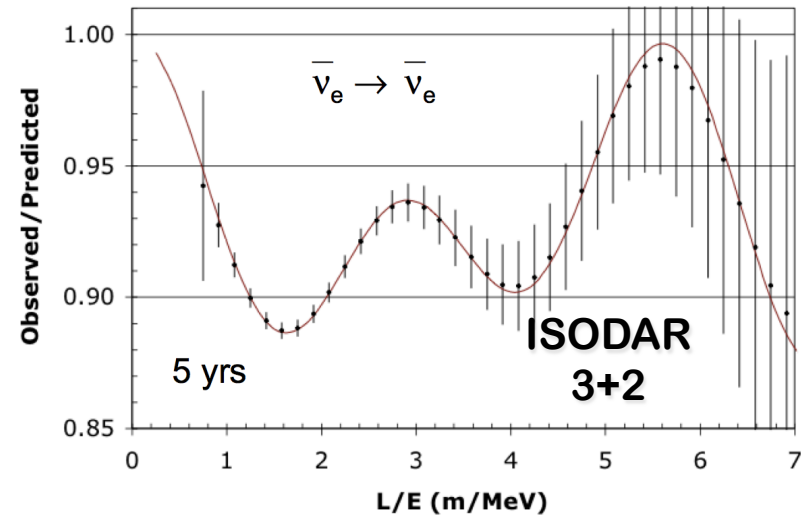
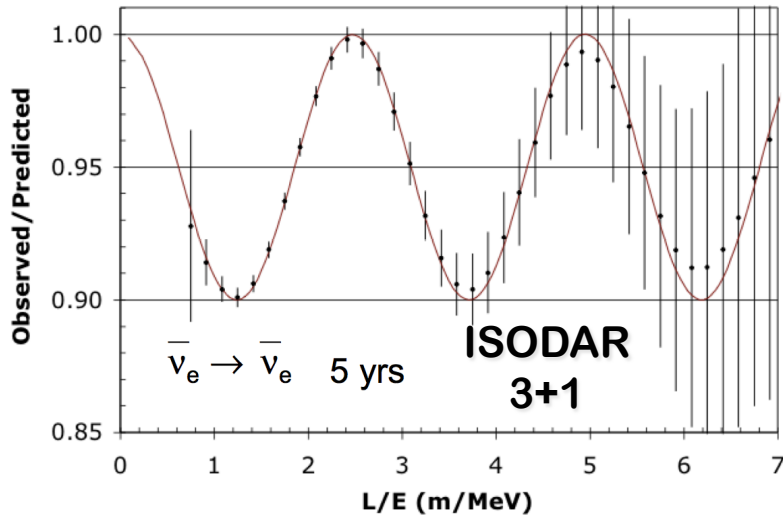
## Appearance



# Isotope Decay at Rest $\nu$ -sources



## Oscillation L/E Waves with High Statistics



# Number of $\nu$ 's From Cosmology

- Constraint sum of neutrino masses  $< 0.5 - 1$  eV
- An excess of non-interacting relativistic energy density  
→ can be interpreted as “extra  $\nu$ ”
- **WMAP 2013 + other observables** →  $N_{\text{eff}} = 3 - 4$
- **Planck:**
  - Planck alone:  $N_{\text{eff}} = 3.36 \pm 0.66$  (95% C.L.)
  - But
    - $H_{0,\text{planck}}$  in tension with  $H_{0,\text{HST}}$
    - Planck + BAO +  $H_0$ :  $N_{\text{eff}} = 3.52 \pm 0.46$  (95% C.L.)
    - Depends on BICEP2 results
- **$N_{\text{eff}} = 4$  mildly disfavored & bound model dependent**
- **Neutrino laboratory experiments should be used as an input for cosmology**

# Conclusion (1)

- **2.7 – 3.8  $\sigma$  anomalies (each) calling for clarification**
  - LSND & MiniBooNE?
  - Gallium Anomaly
  - Reactor Anomaly

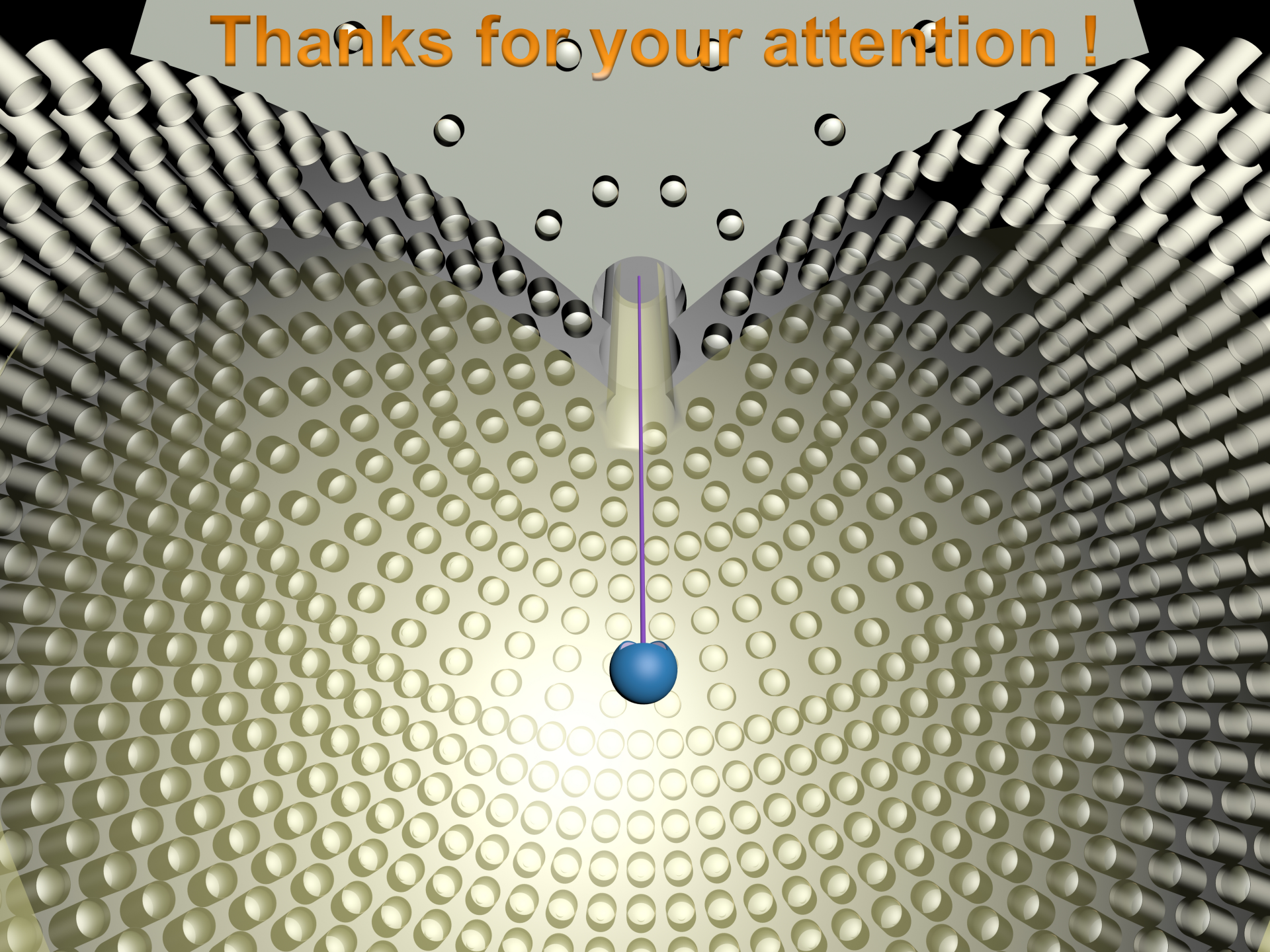
→  $\Delta m^2 \approx eV^2$  Sterile Neutrino? Or Experimental Artifacts?
- **But also negative indications:**
  - No deficit in  $\Delta m^2 \approx eV^2$  muon disappearance
  - Tensions in global fits (APP vs DIS)
- **A definitive probe of this parameter space is necessary → need several experiments**

# Conclusion (2)

- Many proposals with capabilities to unambiguously test  $L/E \approx 1$  m/MeV oscillatory behavior with low backgrounds
- **Reactor Neutrinos**
  - Results within 5 years, Cost :1-10 M\$
  - Background mitigation is challenging
- **Neutrino Generator**
  - Results within 5 years, cost  $\approx 5$  M\$
  - Challenge for the source production and transportation
- **Neutrino 'Beam'**
  - Longer term, cost 10-300 M\$
  - Would allow studying sterile neutrino phenomenology
  - Relevant for Xsection study and R&D for next generation projects
- **Other tests through  $\beta$ -decay and  $(\beta\beta)0\nu$ -decay, Cosmo**

	Primary Channel	Other osc channels	Definitive sterile?	Other physics	Tech R&D?	Cost	Why worry?	Comment
MicroBooNE ( $\pi$ DIF)	$\nu_\mu \rightarrow \nu_e$	$\nu_\mu \rightarrow \nu_\mu$		GeV-scale xsec	Yes	\$20M	tech, cosmics	Exists!
LAr1-ND ( $\pi$ DIF)	$\nu_\mu \rightarrow \nu_e$	$\nu_\mu \rightarrow \nu_\mu$		GeV-scale xsec	Yes	\$13M	tech, cosmics	
ICARUS@FNAL ( $\pi$ DIF)	$\nu_\mu \rightarrow \nu_e$	$\nu_\mu \rightarrow \nu_\mu$		GeV-scale xsec	Yes	Under study	tech, cosmics	
TripleLAr@FNAL ( $\pi$ DIF)	$\nu_\mu \rightarrow \nu_e$ $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	$\nu_\mu \rightarrow \nu_\mu$ $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$	Probably 	GeV-scale xsec	Yes	Under study	tech, cosmics	Work in progress. Anti-nu?
OscSNS ( $\pi, \mu$ DAR)	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	$\nu_e \rightarrow \nu_e$	Yes	Supernova xsec	No	\$20M	intrinsic $\bar{\nu}_e$	
JPARC MLF ( $\pi, \mu, K$ DAR)	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	$\nu_e \rightarrow \nu_e$ $\nu_\mu \rightarrow \nu_e$	Not in phase 1	Supernova and 235 MeV $\nu_\mu$ xsec	No	\$5M	intrinsic $\bar{\nu}_e$	Phase 1
IsoDAR-KamLAND (Isotope DAR)	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	-	Yes	$\bar{\nu}_e e^-$ (electroweak)	Yes	\$30M	timeline, tech	
nuSTORM ( $\mu$ DIF)	$\nu_e \rightarrow \nu_\mu$	$\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$ $\nu_e \rightarrow \nu_e$	Yes	GeV-scale xsec	Yes	\$300M	timeline, tech, cost	P5 says no

Thanks for your attention !





# Projects Overview

Experiment Type	Appearance / Disappearance	Oscillation Channel	Projects
<b>Reactor</b>	Disappearance	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	Nucifer, Stéréo, Scraam, Neutrino-4, DANSS, Poséidon, Solid, ...
<b>Radioactive Source</b>	Disappearance	$\bar{\nu}_e \rightarrow \bar{\nu}_e$ $\nu_e \rightarrow \nu_e$	CeLAND, SOX (Cr & Ce), Sage2, SNO+, LENS-s
<b>Cyclotron</b>	Disappearance	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	IsoDAR
<b>Pion / Kaon Decay-at-Rest</b>	Apparition & Disappearance	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_e \rightarrow \nu_e$	OscSNS, CLEAR, DAEΔALUS, KDAR
<b>Pion Decay-in-Flight (Beam)</b>	Appearance & Disappearance	$\nu_\mu \rightarrow \nu_e$ $\nu_\mu \rightarrow \bar{\nu}_e$ $\nu_\mu \rightarrow \nu_\mu$ $\nu_e \rightarrow \nu_e$	MINOS+, MicroBooNE, LAr1kton+MicroBooNE, Icarus/Nessie@CERN
<b>Low-E Neutrino Factory</b>	Appearance & Disappearance	$\nu_e \rightarrow \nu_\mu$ $\bar{\nu}_e \rightarrow \bar{\nu}_\mu$ $\nu_\mu \rightarrow \nu_\mu$ $\bar{\nu}_e \rightarrow \bar{\nu}_e$	vSTORM@Fermilab

# Reactor v Proposals

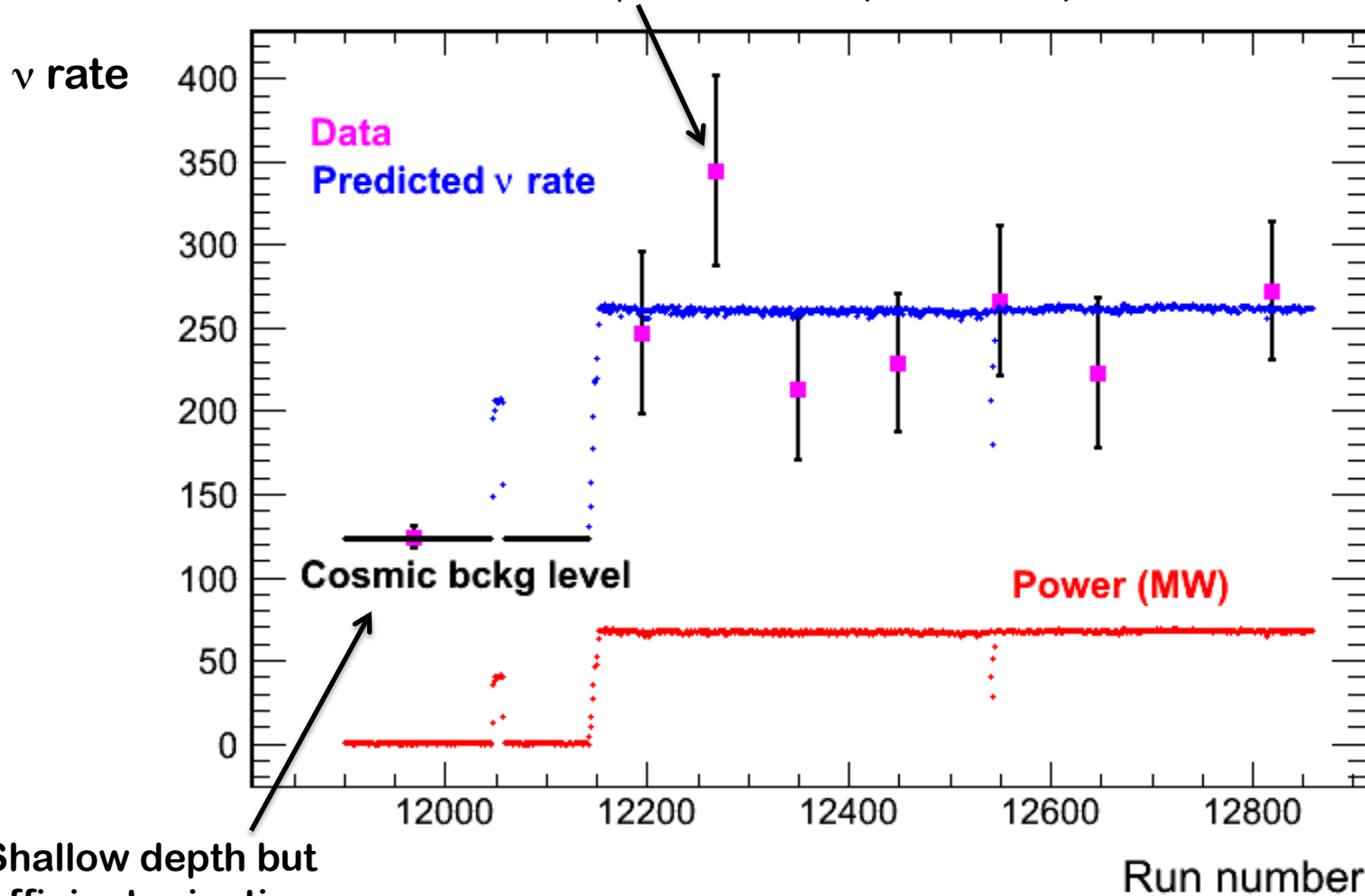
Experiment Type	Experimental Strategy	Projects	Features
Gd-doped LS detector	<ul style="list-style-type: none"> <li>-Clear signature of n-capture by an 8 MeV <math>\gamma</math>-cascade - Mature</li> <li>-High light yield <math>\rightarrow</math> fast n background rejected by PSD</li> <li>-Sensitive to high-E <math>\gamma</math>'s <math>\rightarrow</math> large passive shielding</li> </ul>	Nucifer (FRA)	70 MW / 0.7 t / 7 m / few mwe
		Stéréo (FRA)	50 MW / 2 t / [8-11] m / 10 mwe
		Neutrino 4 (RU)	100 MW / 2 t / [6-12] m / surf.
Highly segmented detector	<ul style="list-style-type: none"> <li>Improved bkg rejection: <ul style="list-style-type: none"> <li>- Vertex correlation between prompt and delayed</li> <li>- Topology of E depositions: <ul style="list-style-type: none"> <li>e = compact track</li> <li><math>\gamma</math> = longer interaction length</li> </ul> </li> </ul> </li> </ul>	DANSS (RU)	1 GW / 1 t / [10-12] m / 50 mwe
		SoLid (UK)	45-80 MW / 3 t / 8 m / 10 mwe
Enhanced neutron Tagging	<ul style="list-style-type: none"> <li>-Unique signature of neutron capture with Li-doped LS/PS</li> </ul> ${}^6\text{Li} + n \rightarrow \alpha + t$	Hanaro (KO)	30 MW / 0.5 t / 6 m / few mwe
Extended Baseline Complex	<ul style="list-style-type: none"> <li>-Better sensitivity to lower <math>\Delta m^2</math></li> <li>-Larger volume and/or Longer running time</li> </ul>	US project	20-120 MW / 4m-15m / surf.
		China project	
		DANSS/Neutrino4	Movable detector

# Reactor ν Proposals

Experiment Type	Experimental Strategy
<p><b>Mature Gd-doped LS detector Technology</b></p>	<ul style="list-style-type: none"> <li>- Clear signature of n-capture (8 MeV <math>\gamma</math>-cascade)</li> <li>- High light yield <math>\rightarrow</math> fast n background rejection by PSD</li> <li>- But sensitive to high-E <math>\gamma</math>'s <math>\rightarrow</math> need large passive shielding</li> </ul>
<p><b>Highly segmented detector for background reduction</b></p>	<ul style="list-style-type: none"> <li>- Vertex correlation between prompt and delayed</li> <li>- Topology of E depositions:               <ul style="list-style-type: none"> <li>e <math>\rightarrow</math> compact track</li> <li><math>\gamma</math> <math>\rightarrow</math> longer interaction length</li> </ul> </li> </ul>
<p><b>Enhanced neutron Tagging</b></p>	<ul style="list-style-type: none"> <li>- Unique signature of neutron capture with Li-doped LS/PS</li> </ul> ${}^6\text{Li} + n \rightarrow \alpha + t$
<p><b>2 detector complex or Moving detector</b></p>	<ul style="list-style-type: none"> <li>- Better sensitivity to lower <math>\Delta m^2</math></li> <li>- But Need larger volume and/or longer running time</li> </ul>

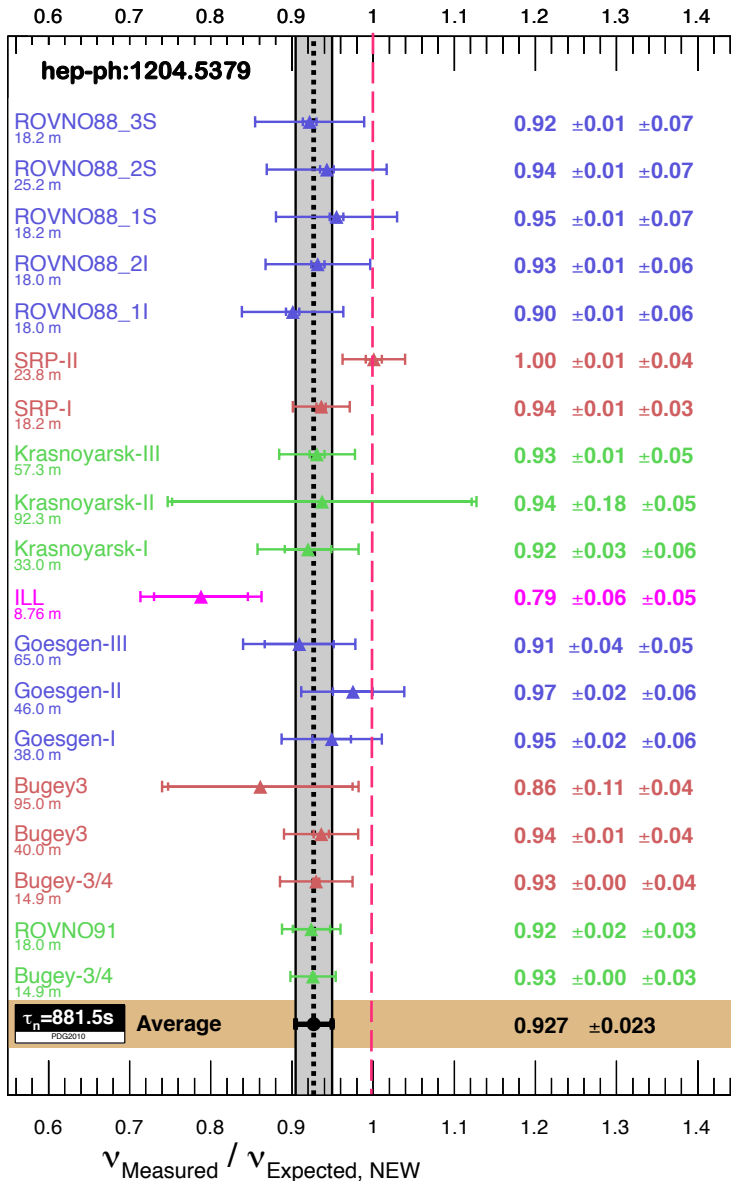
# Nucifer: First Neutrino Run

- No reactor induced fast neutrons
- but need further  $\gamma$  attenuation (lead, 4 cm) for sterile  $\nu$  search



Shallow depth but  
efficient rejection  
of cosmic bkg

# Reactor Antineutrino Anomaly



- 19 Short Baseline Experiments (L<100m)
- Observables: ratios of observed event rate to predicted rate of events

## 2011 results

- Average:  $\mu = 0.943 \pm 0.023$
- 98.6 % C.L. deviation from  $\mu = 1$

## 2012 results

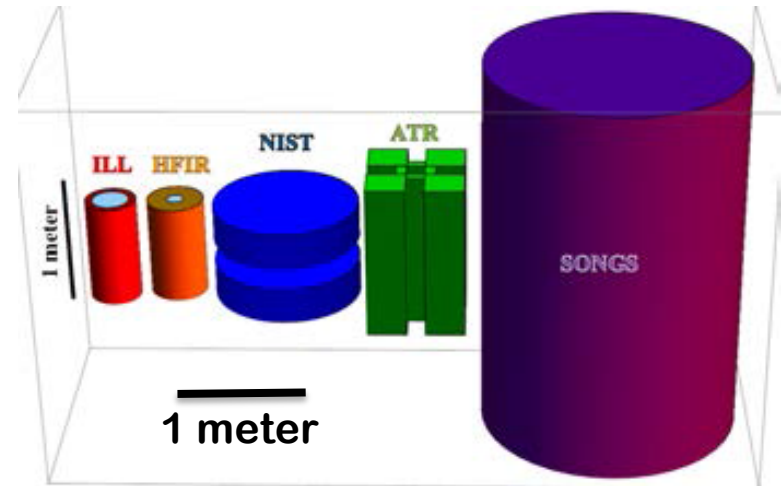
- Average  $\mu = 0.927 \pm 0.023$
- 99.7 % C.L. deviation from  $\mu = 1$

## 2013: update: refined analysis

# New SBL reactor experiments

- **Compact reactor core**
  - No oscillation smearing
- **High statistics (few 100 evts/day/t)**
  - High Power (10-3000 MW)
  - Short baselines (5-50 m)
- **Highly enriched fuel**
  - A singly  $^{235}\text{U}$  fission spectrum component
- **Reactor ON/OFF periods**
  - Moderate overburden compensated by accurate measurement of the cosmogenic bkg
- **But challenging reactor-induced backgrounds ( $\gamma$  and  $n$ )**
  - Need comprehensive site characterization

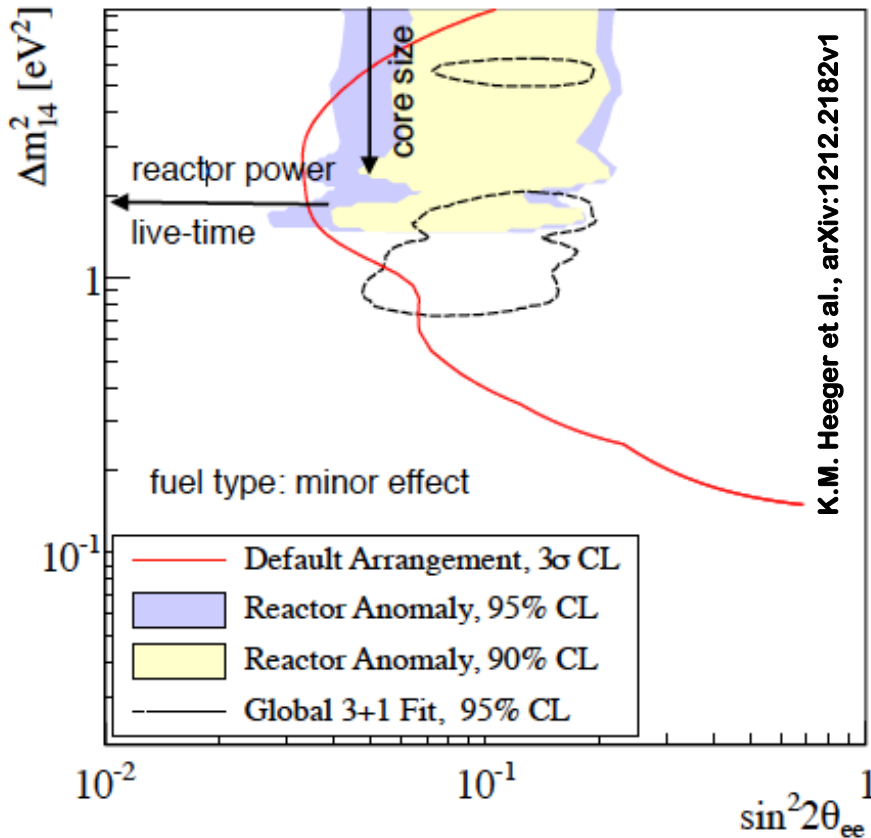
Typical reactor core sizes



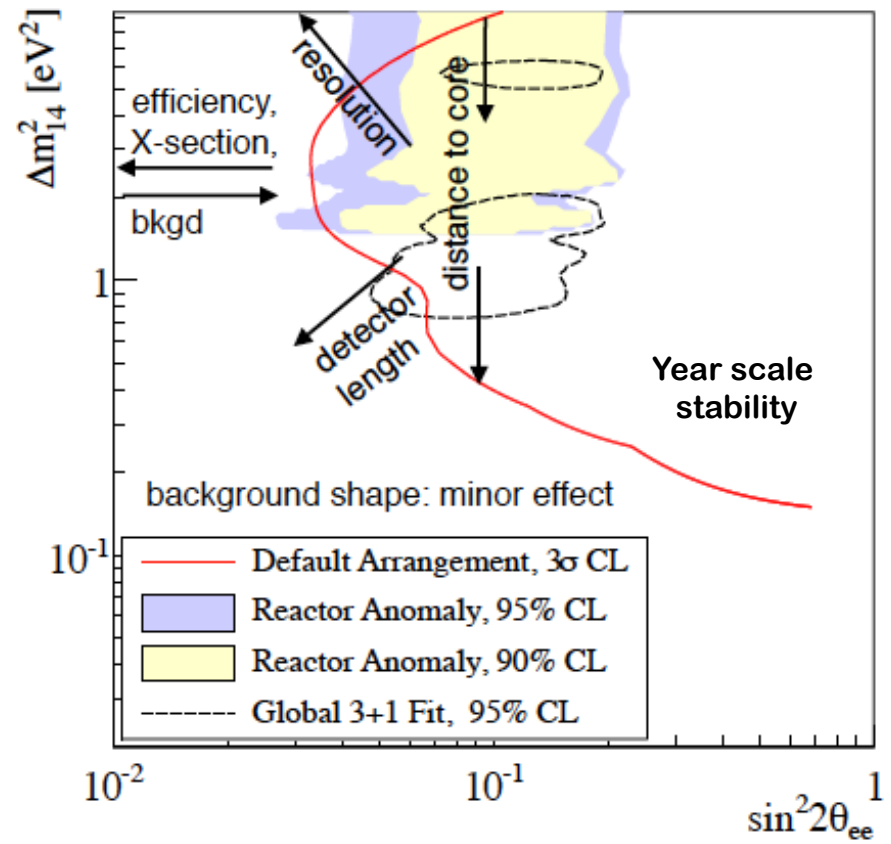
# Influence of Source/Detector Parameters

All current project have the sensitivity to test the reactor anomaly space of parameters,  $\Delta m^2 > 0.1$ ,  $\sin^2 2\theta_{ee} > 0.05$

## Source

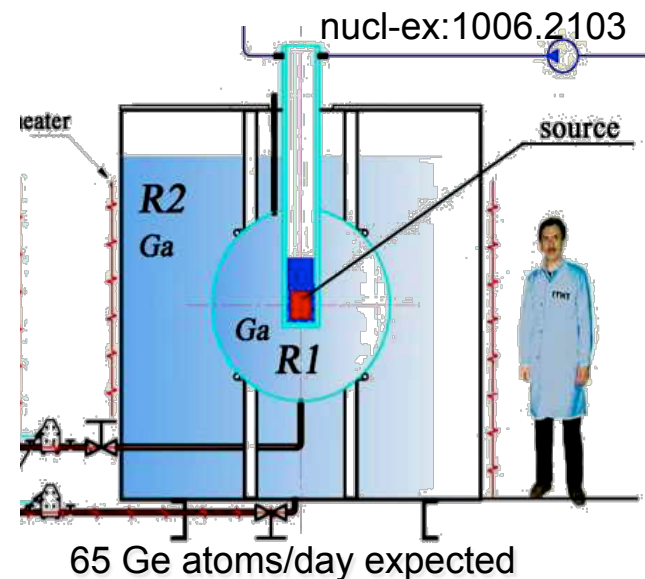


## Detector



# $^{51}\text{Cr}$ : SAGE 2-Zone (Sage)

- $^{51}\text{Cr}$  Source:
  - Enrichment of 3.5 kg  $^{50}\text{Cr}$  (97%, 2014)
  - Irradiation to reach **3 MCi (2015?)** at research reactor SM-3
  
- **2-layer detector in Baksan**
  - Inside a new dual Metallic Ga Target
  - Zone 1: 8t - Zone 2: 42 t metal Ga
  - SAGE procedures well understood
  - Not sensitive to  $\gamma$ -ray background
  
- **Observable**
  - Ratio of  $\nu_e$  capture rates to predicted rate in inner (R1) and outer zone (R2)
  - Ratio  $R_2/R_1$





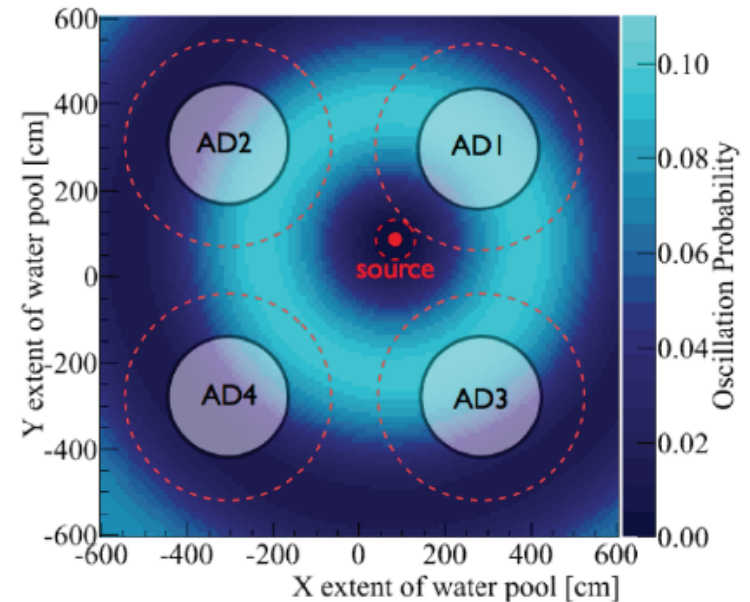
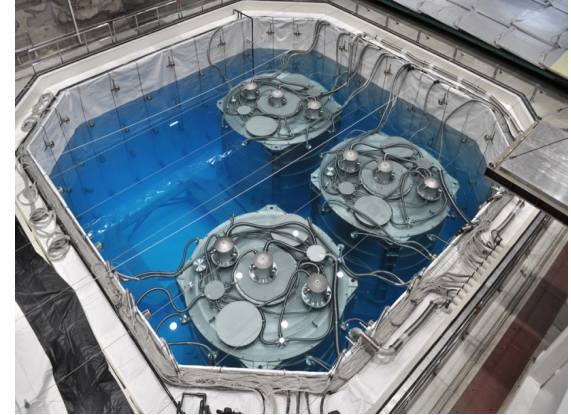
# 500 kCi $^{144}\text{Ce}$ - $^{144}\text{Pr}$ in Daya Bay

- 500 kCi of  $^{144}\text{Ce}$  in the water pool of the Daya Bay far hall

- Baseline range: 1.5 - 8 m
- Energy range: 1.8 - 3 MeV
- 35 000 IBD events/per year
- 'Easy' to deploy

- Ongoing discussion for  $^{144}\text{Ce}$  recovery with LLNL

- Multiple source location to probe sterile oscillations



# 500 kCi $^{144}\text{Ce}$ - $^{144}\text{Pr}$ in Daya Bay

- Specific oscillation pattern through simulation
- Water + 50 cm W-shielding
  - $\gamma$ 's attenuation
- Must subtract reactor neutrino 'background'
  - well-known to <1% from near detectors
- Sterile neutrino oscillations with mass  $>1\text{eV}$  can be tested

