

# The LAT tracker performance and first results after launch

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**On behalf of GLAST/LAT**  
**collaboration**  
**INFN Bari**



**17th International Workshop on Vertex detectors**  
**July 28–August 1, 2008, Utö Island, Sweden**

# The GLAST/LAT collaboration



## France

IN2P3  
CEA/Saclay



## Italy

INFN  
ASI  
INAF



## Japan

Hiroshima University  
ISAS/JAXA  
RIKEN  
Tokyo Institute of Technology



## Sweden

Royal Institute of Technology (KTH)  
Stockholm University



## United States

Stanford University (SLAC and HEPL/Physics)  
University of California at Santa Cruz - Santa Cruz Institute for Particle Physics  
Goddard Space Flight Center  
Naval Research Laboratory  
Sonoma State University  
Ohio State University  
University of Washington

**Principal Investigator:**  
**Peter Michelson (Stanford University)**

~270 Members  
(~90 Affiliated Scientists, 37 Postdocs,  
and 48 Graduate Students)

**construction managed by**  
**Stanford Linear Accelerator Center**  
**(SLAC), Stanford University**

# Outline

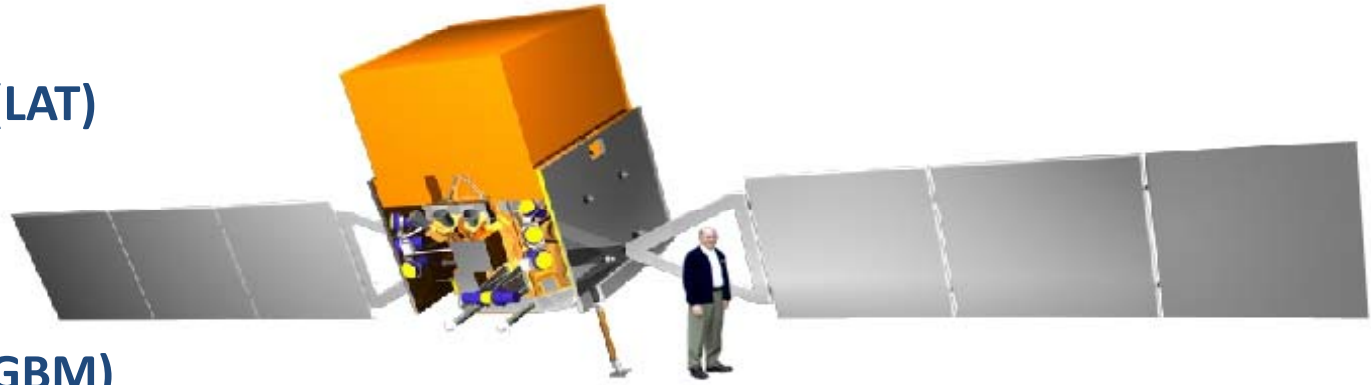
- GLAST/LAT
  - Overview
- LAT Tracker performance
  - Flight Data !!!
- Conclusions

# The GLAST mission

## Large Area Telescope (LAT)

20 MeV - 300 GeV

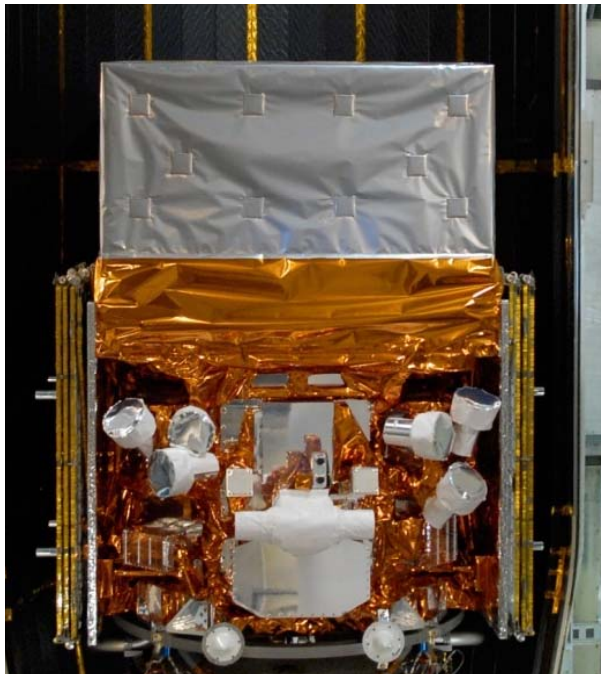
>2.5 sr FoV



## GLAST Burst Monitor (GBM)

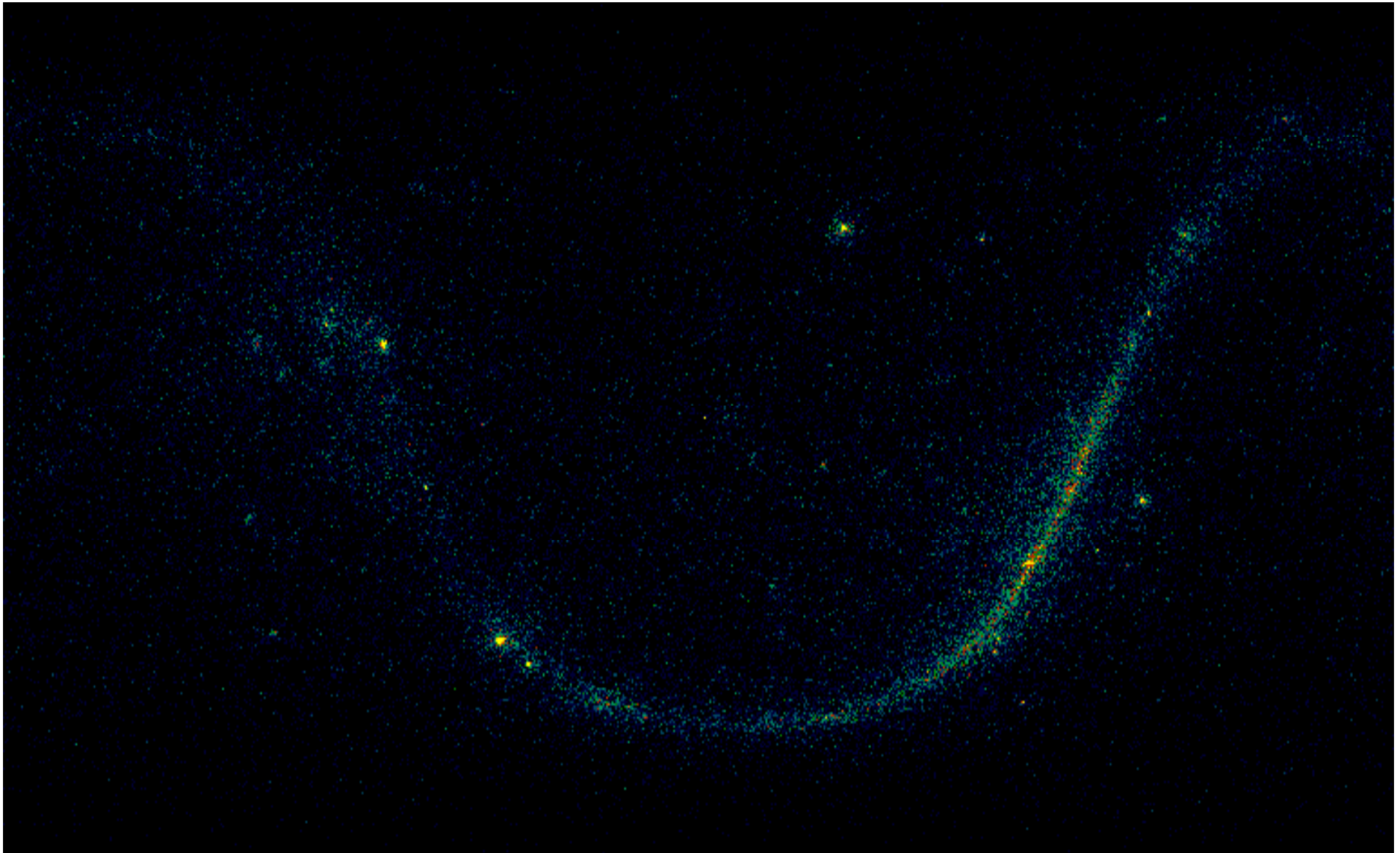
8 keV – 30 MeV

9 sr FoV



- **Huge field of view (FOV)**
  - LAT: 20% of the sky at any instant; in sky survey mode, expose all parts of sky for ~30 minutes every 3 hours.
  - GBM: whole unocculted sky at any time.
- **Huge energy range**, including largely unexplored band 10 GeV - 100 GeV
- **Will transform the HE gamma-ray catalog:**
  - by more than one order of magnitude in # point sources
  - by more than one order of magnitude in timing resolution
  - spatially extended sources
  - sub-arcmin localizations (source-dependent)

# The gamma sky is really variable





# The LAT Instrument

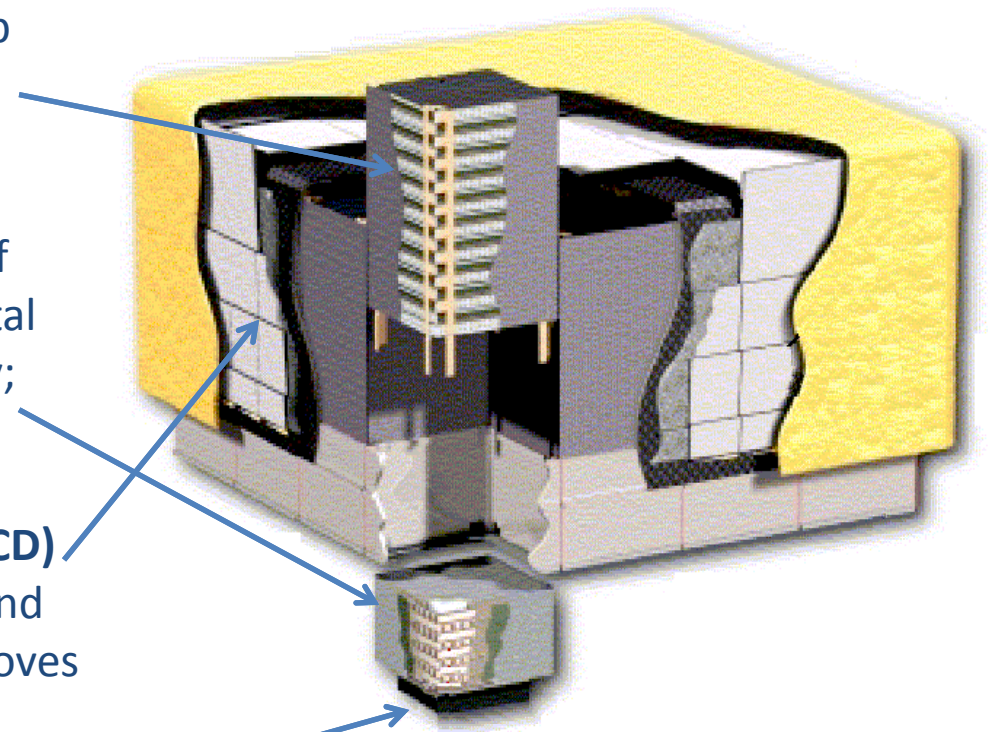
## (a pair-conversion gamma telescope)

- **Precision Si-strip Tracker (TKR)** 16 Towers. 18 XY tracking planes. Single-sided silicon strip detectors Measure the photon direction; gamma ID.

- **Hodoscopic CsI Calorimeter(CAL)** Array of 1536 CsI(Tl) crystals in 8 layers (8.6 X0). Total mass 1.5 tons. Measure the photon energy; image the shower.

- **Segmented Anticoincidence Detector (ACD)** 89 plastic scintillator tiles. Reject background of charged cosmic rays; segmentation removes self-veto effects at high energy.

- **Electronics System** Includes flexible, robust hardware trigger and software filters.



# Tracker Design Overview

- **16 tower modules**

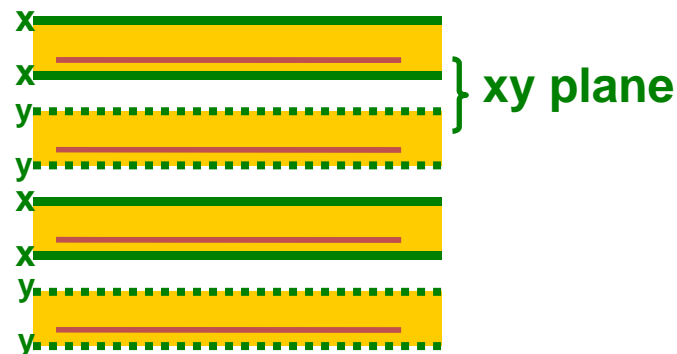
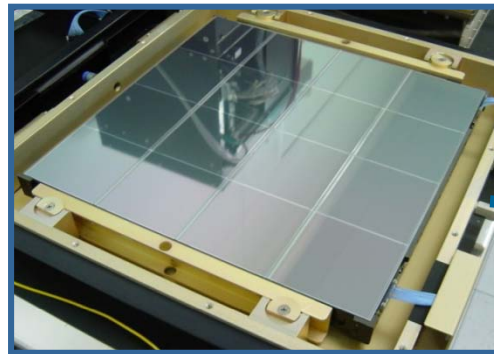
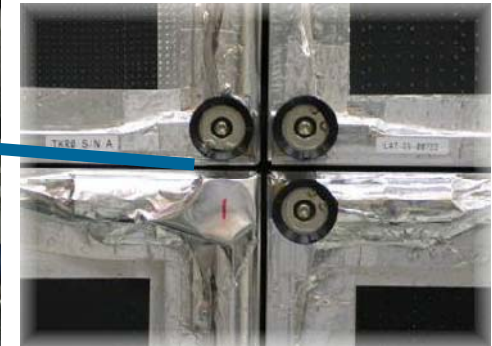
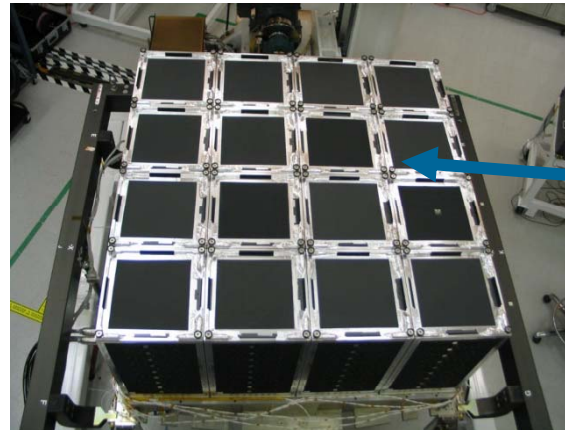
- $37 \times 37 \text{ cm}^2$  active cross section/layer
- 2 mm inter-tower separation in order to minimize the inactive area
- **$70 \text{ m}^2$  of Si (in space!!!)**
- 11500 SSD  $8.95 \times 8.95 \text{ cm}^2$ ,
- 384 strips - 880,000 channels
- $440 \text{ }\mu\text{m}$  thick
- $228 \text{ }\mu\text{m}$  strip pitch:

- **18 xy layers per tower**

- 19 “trays” structure
  - 12 with  $2.7\% X_0 W$
  - 4 with  $18\% X_0 W$
  - 3 with no converter foils
- every tray is rotated by  $90^\circ$
- W foils are followed by
  - x,y plane of detectors
  - 2mm gap between

- **Electronics on sides of trays:**

- Minimize gap between towers



# Readout Electronics Architecture

This a **space experiment**: Emphasis on compactness, minimum of wiring, redundancy, very low power consumption:

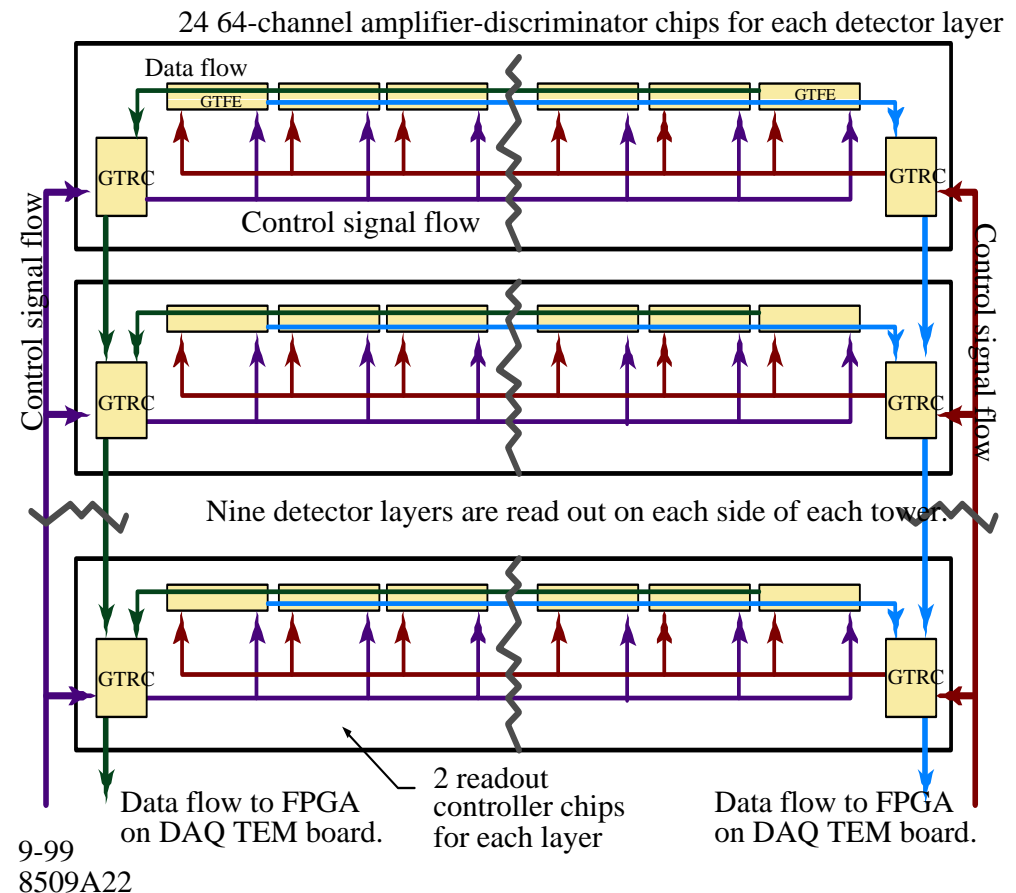
Any single component (GTFE, GTRC, cable) can fail without affecting the other.

## GTFE (GLAST Tracker Front-end Electronics) ASIC

- Preamplifier - shaper - discriminator
- One threshold DAC and one calibration DAC per chip.
- 64 channels per chip, 24 chips per MCM.
- Noise:  $\sim 1500$  e for 4 SSD ladder.
- Gain:  $\sim 100$  mV/fC.
- Peaking time:  $1.5 \mu\text{s}$ .
- $0.1$  mW/channel.

## GTRC (GLAST Tracker Readout Controller)

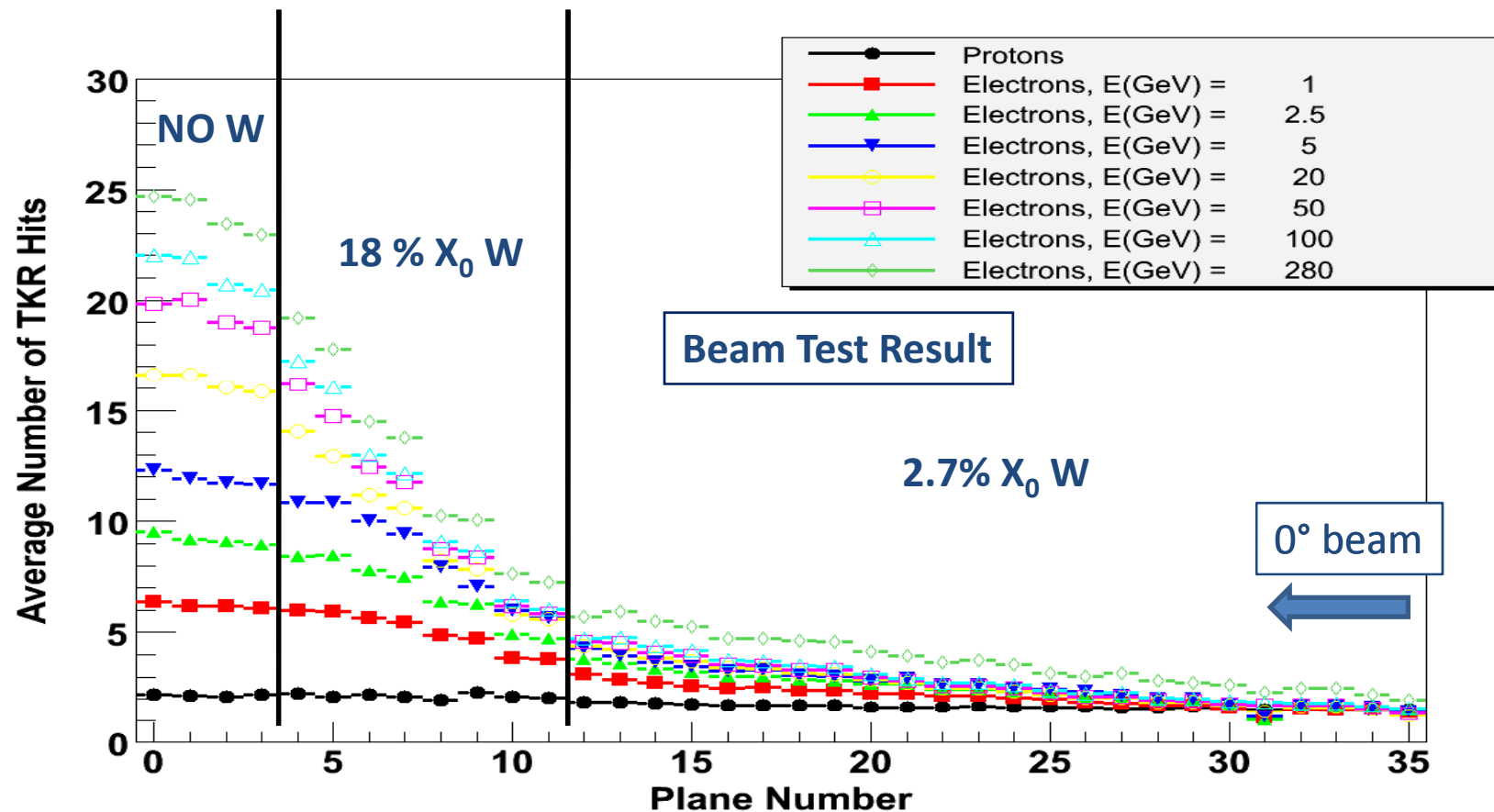
- 9 GTRC per cable.
- Communication between 24 GTFE and back-end electronics.
- TOT measurement from layer-OR trigger signal





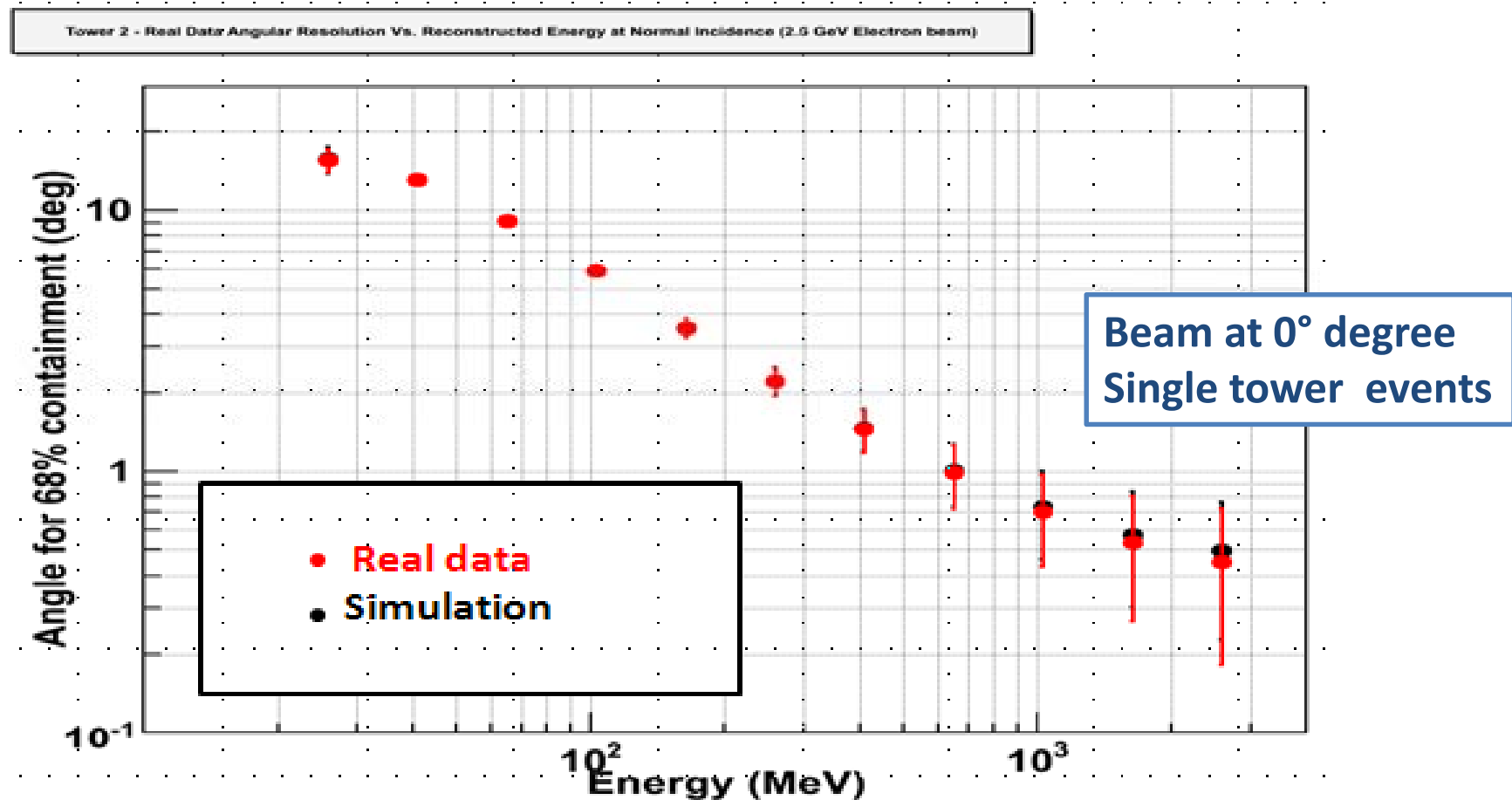
# The tracker isn't light....

- The tracker-converter design is a compromise between very thin converter foils, good for **angular resolution**, versus thick foils, good for **effective area**.
- The hits multiplicity increases as the incident particle crosses the tracker. The main increase is (for normal incidence) in the bottom x-y planes where the thick W converters are located . The data are from beam test campaign in summer 2006



## ...but has a good angular resolution for gammas

- The Angular Resolution shown has been evaluated from Beam Test Data (gammas produced by bremsstrahlung between beam electrons and the upstream materials)
- It has been evaluated with respect to the nominal electron beam direction



# LAUNCH!!!



- Launch from Cape Canaveral Air Station 11<sup>th</sup> June 2008 at 12:05PM EDT

- Circular orbit, 565 km altitude (96 min period), 25.6 deg inclination



# Tracker milestones

- **23<sup>rd</sup> June LAT Turn On**

- Conservative SSD voltage (80 V)
- Nominal threshold loaded (0.28MIP)
- Everything was fine. The noise occupancy was very low . The trigger rate was as expected

- **30<sup>th</sup> June Tracker at nominal values**

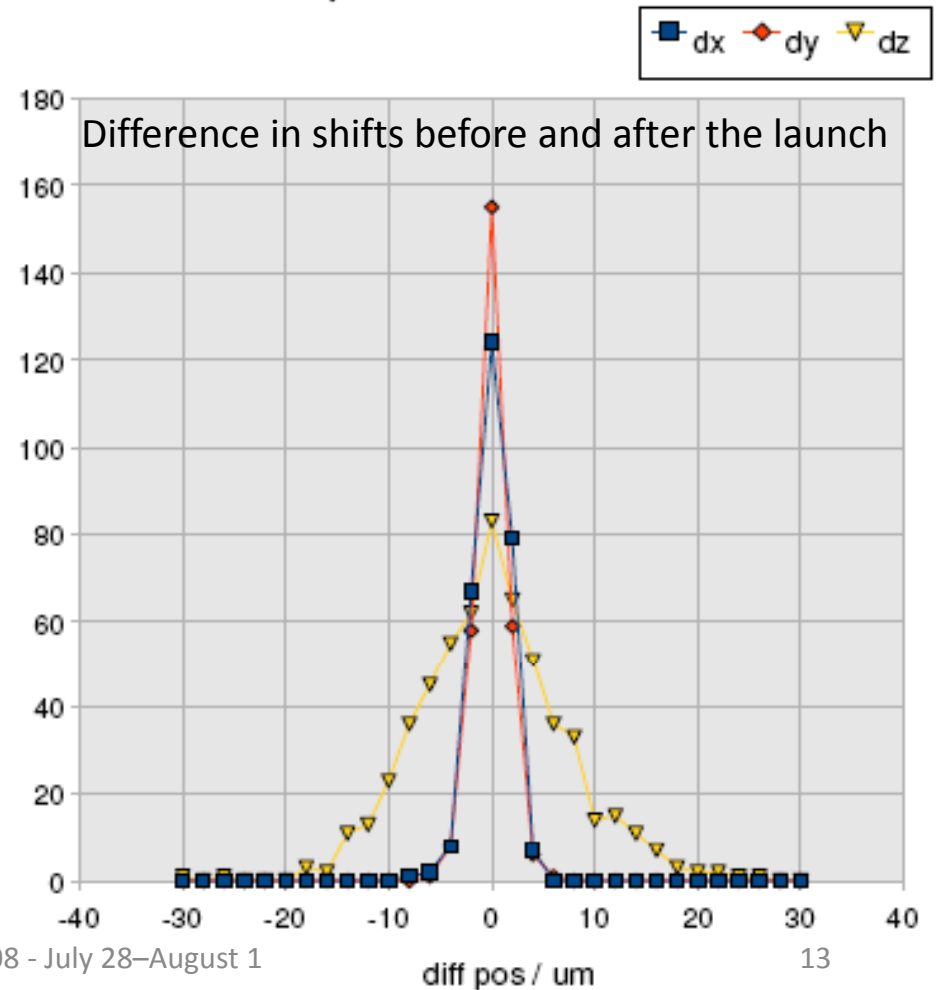
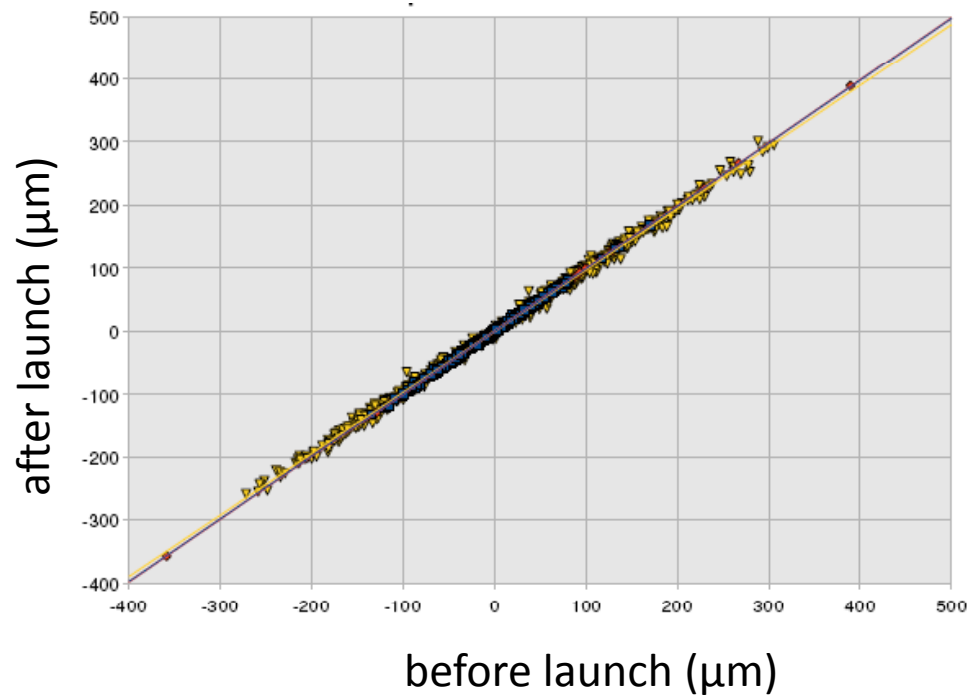
- SSD voltage raised to 105V (nominal value). An increase of 0.16% in hit efficiency as been observed
- No need to load new threshold (the on orbit calibration gave the same result of ground calibration)

- **After 7 days from the Turn ON (and only 23 after the launch) we work in space with the nominal configurations !!!!**

# Well aligned after the launch

## Translations

- GLAST underwent **4g acceleration** during the launch.
- We have verified the alignment of the planes in a tower before and after the launch using crossing particles
- No major shifts one respect to the other



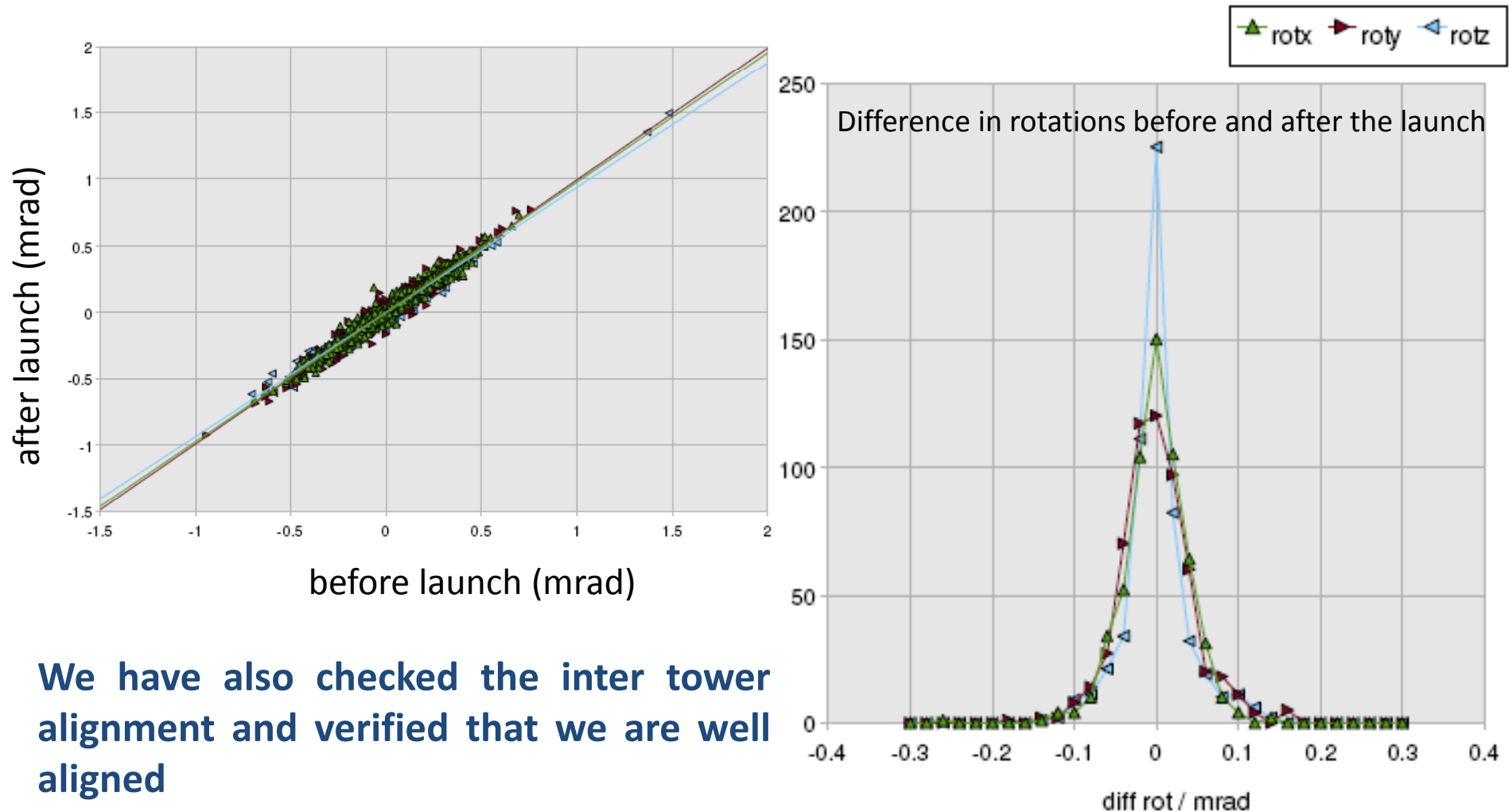
Thanks: M.Kuss (INFN-Pi)

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# ... and Rotations

- No mayor rotations of the planes one respect to the other



**We have also checked the inter tower alignment and verified that we are well aligned**

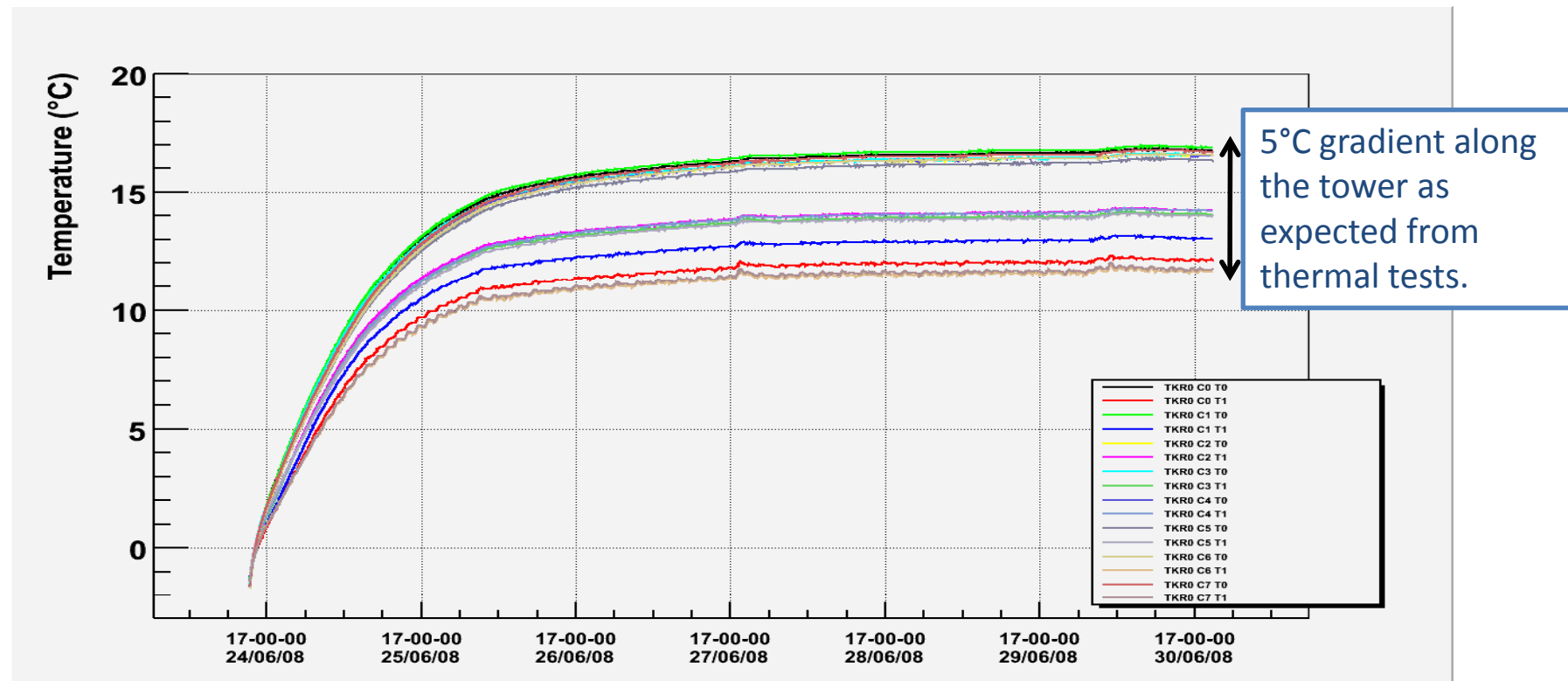
Thanks: M.Kuss (INFN-Pi)

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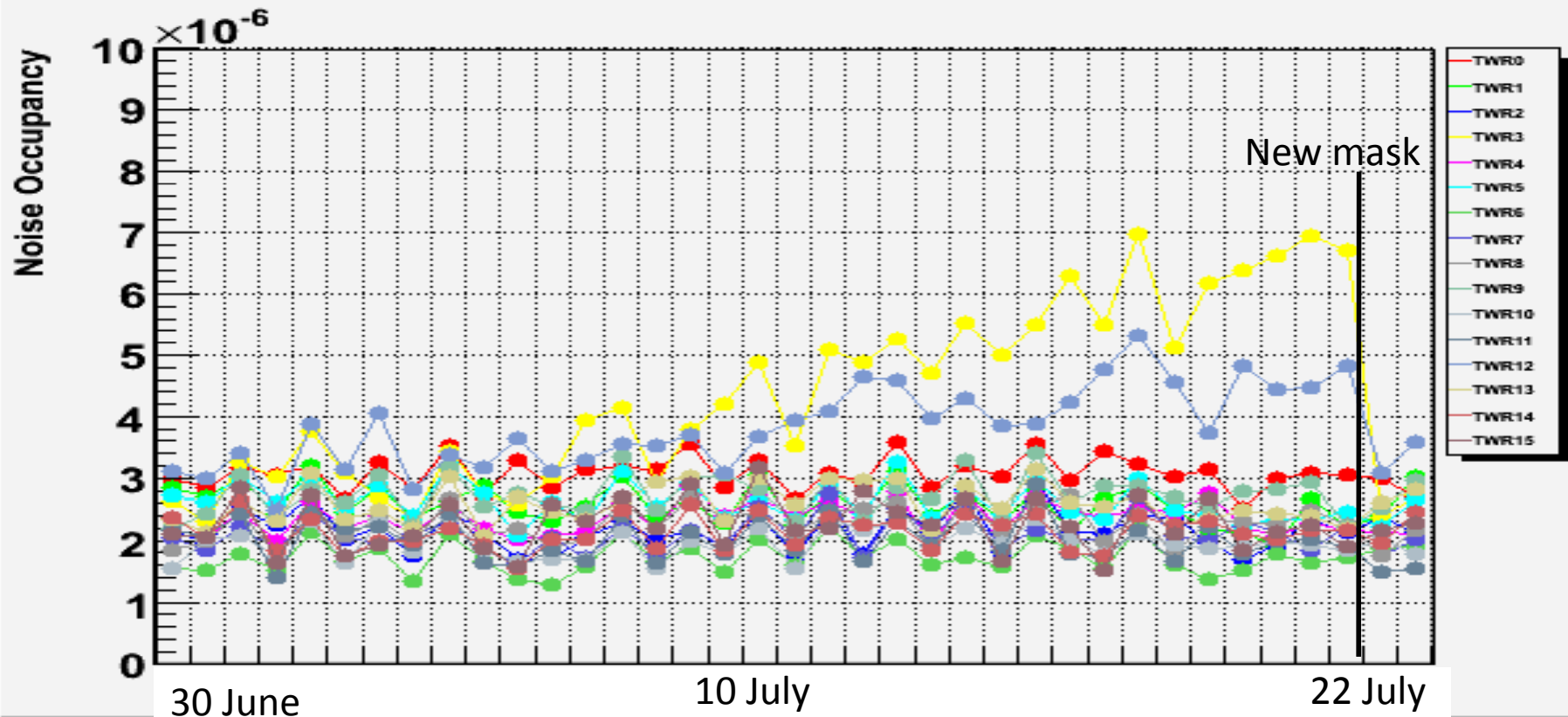
# Temperature is under control

- The spacecraft temperatures are under tight control by an efficient thermal system
- The TKR temperature reach a steady value in a week from turn on
- In sky survey mode only small fluctuations are observed ( $<0.5^{\circ}\text{C}$ ). In pointed mode there is an overall increase of  $2^{\circ}\text{C}$ .
- The tracker is completely insensitive to this kind of variations (no effect on noise)
- The only temperature requirement is to keep the temperature  $< 25^{\circ}\text{C}$  to avoid some strips to flare ( 0.05% on average)

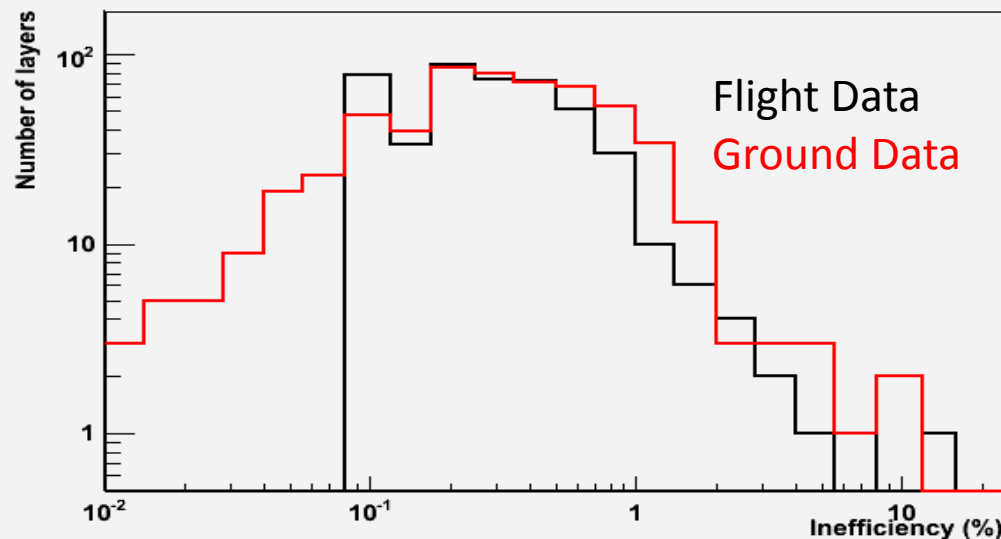
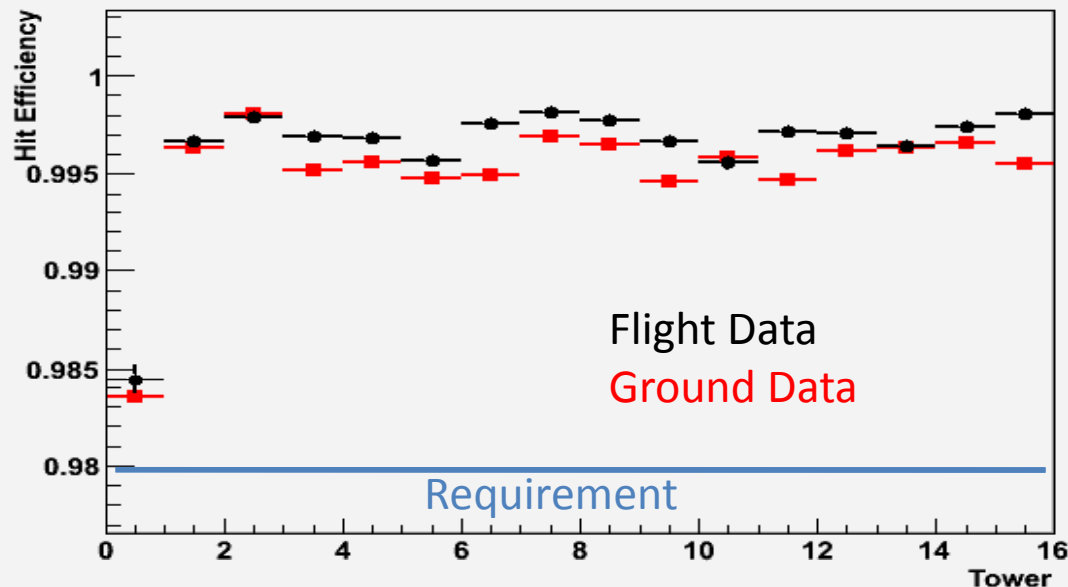


# Very low Noise Occupancy

- The **Noise Occupancy** is defined as the average fraction of channels above threshold at any snapshot in time. We have evaluated it using the on board periodic trigger.
- The on orbit noise occupancy is almost an **order of magnitude lower** than the requirement  $5 \times 10^{-5}$  (averaged over all towers). It is very good for long term operation!!!



# Few hits missed !!!

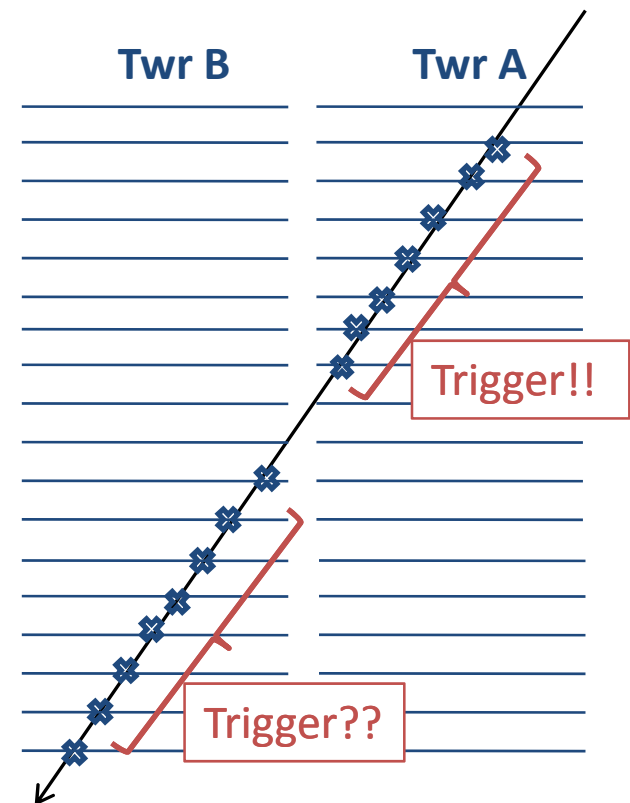
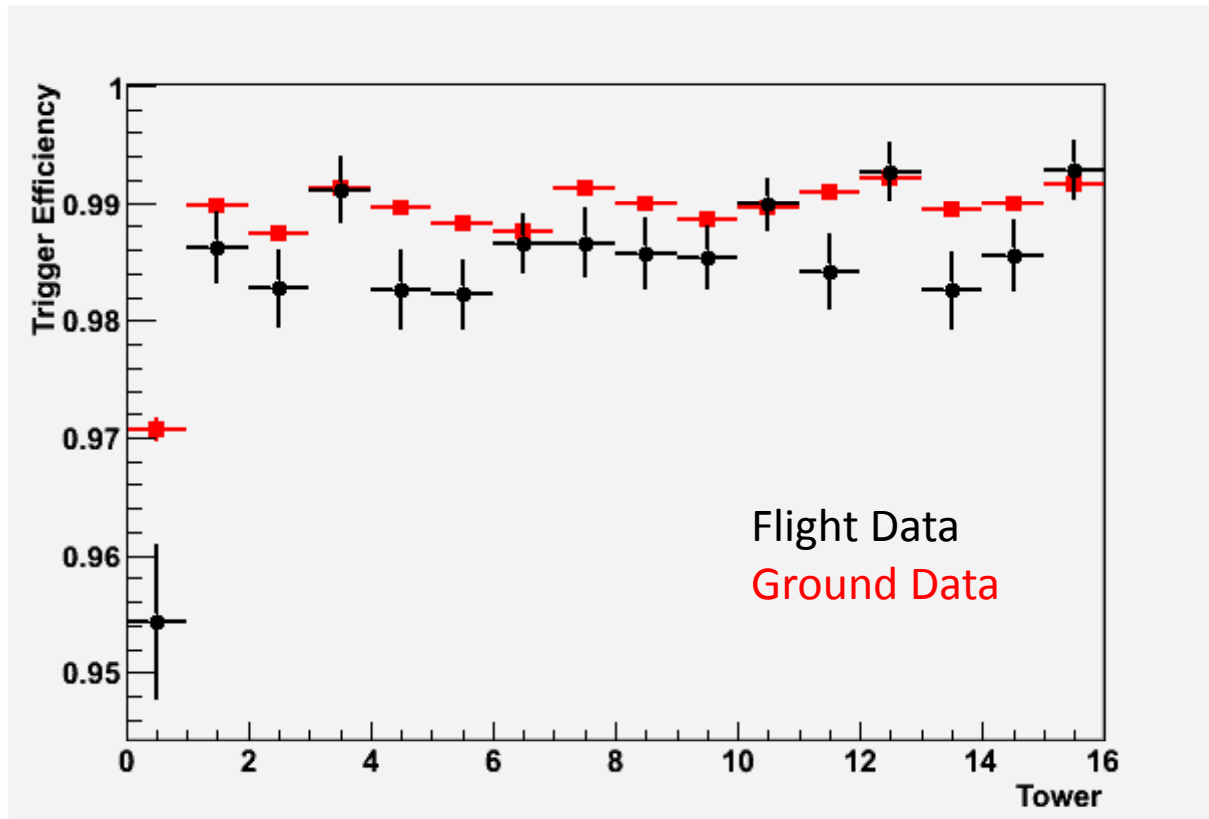


- An high **hit efficiency** is crucial for photon direction especially at low energy in order to measure the first two points after the conversion. The requirements is **> 98%**
- The hit efficiency is evaluated looking for missing hits along the reconstructed track ( $\pm 3.0\text{mm}$  confidential zone)
- The dead zone are taken into account in this procedure (geometrical inefficiency  $\approx 4\%$ ). The hit inefficiency is mainly due to dead strips.
- The overall efficiency is very high **> 99%** for 15 towers
- Only 24 layers (4% of the total) show an hit inefficiency **> 1%**
- Due to the low statistic at the moment we cannot measure inefficiency **< 0.1%**

# High Trigger Efficiency

- A tower trigger is set when there are 6 consecutive hits ( $3x + 3y$ )
- To evaluate the trigger efficiency we select events with a particle crossing two neighbor towers. The first (TwrA) is the one that fires the trigger. We then search in TwrB if there are  $3x+3y$  hits and, if it is the case, we search for a trigger request from TwrB.

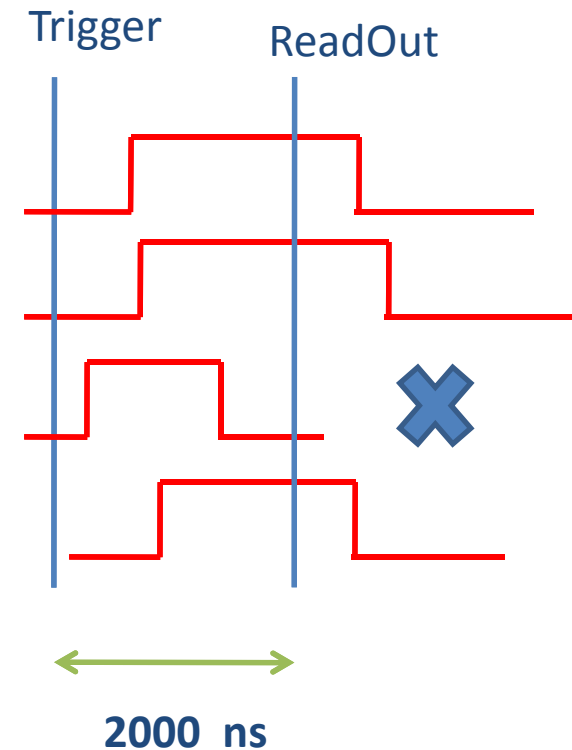
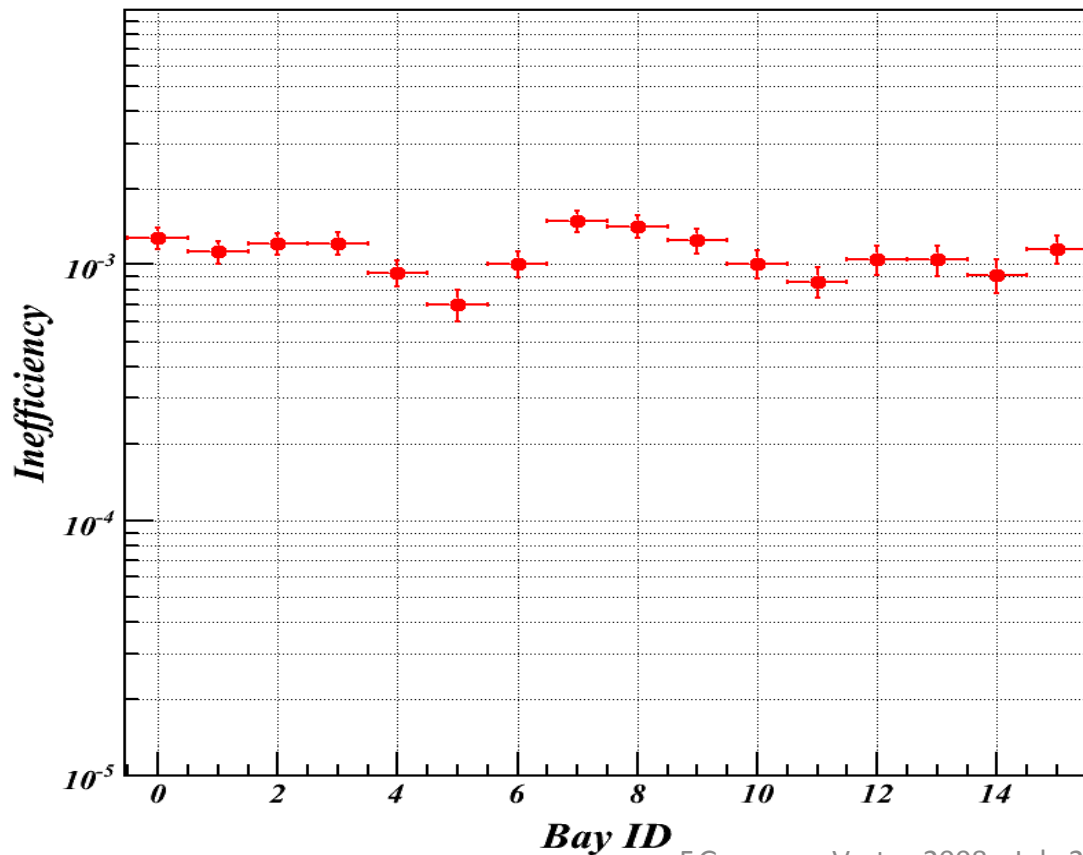
The required hit efficiency is 90%





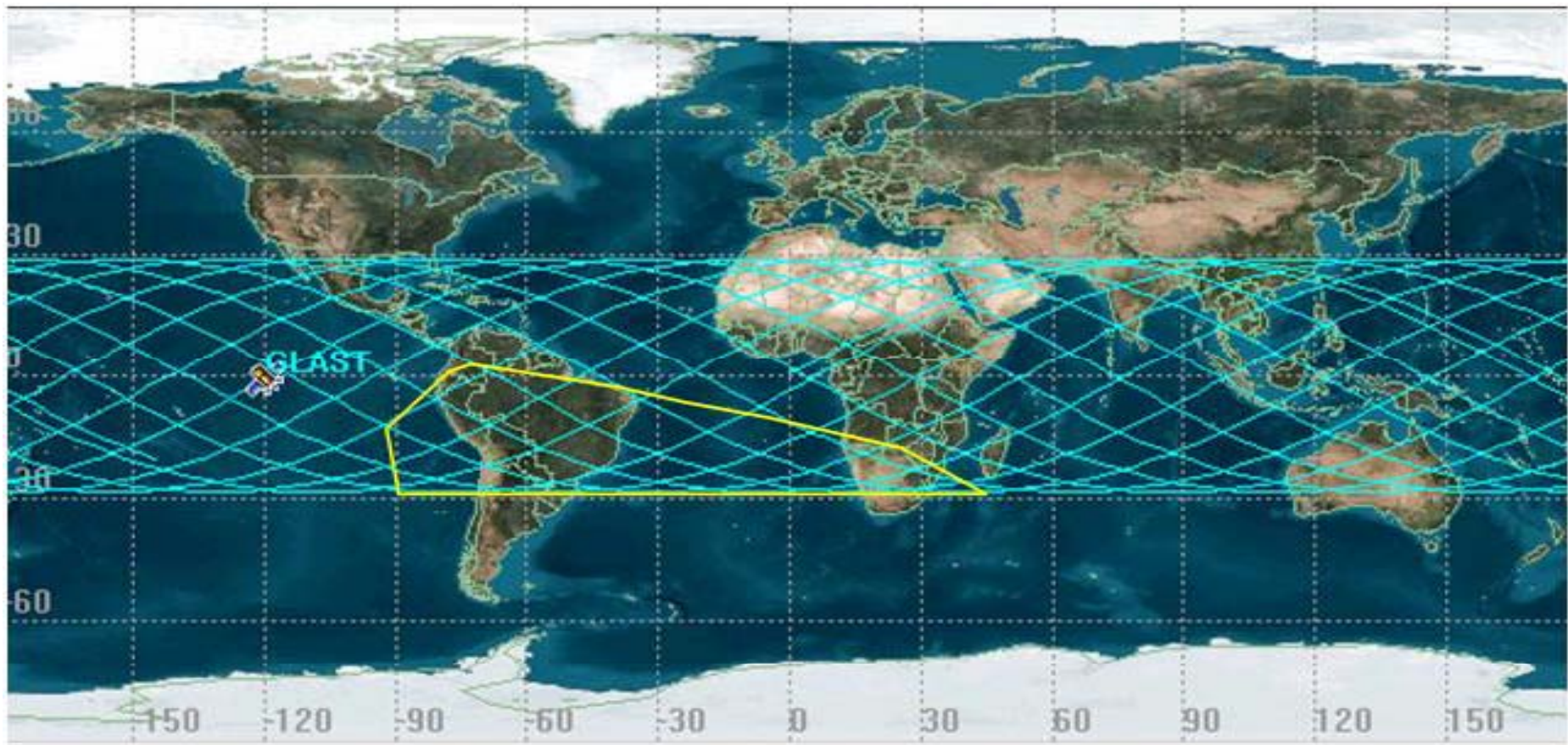
# Very good read out timing

- The **hit capture inefficiency** is evaluated counting all the events in which we miss a ToT in a plane that belongs to a triggering tower
- The hit capture inefficiency is  $\approx 0.1\%$
- The read out timing is really good and we miss very few hits !!!!



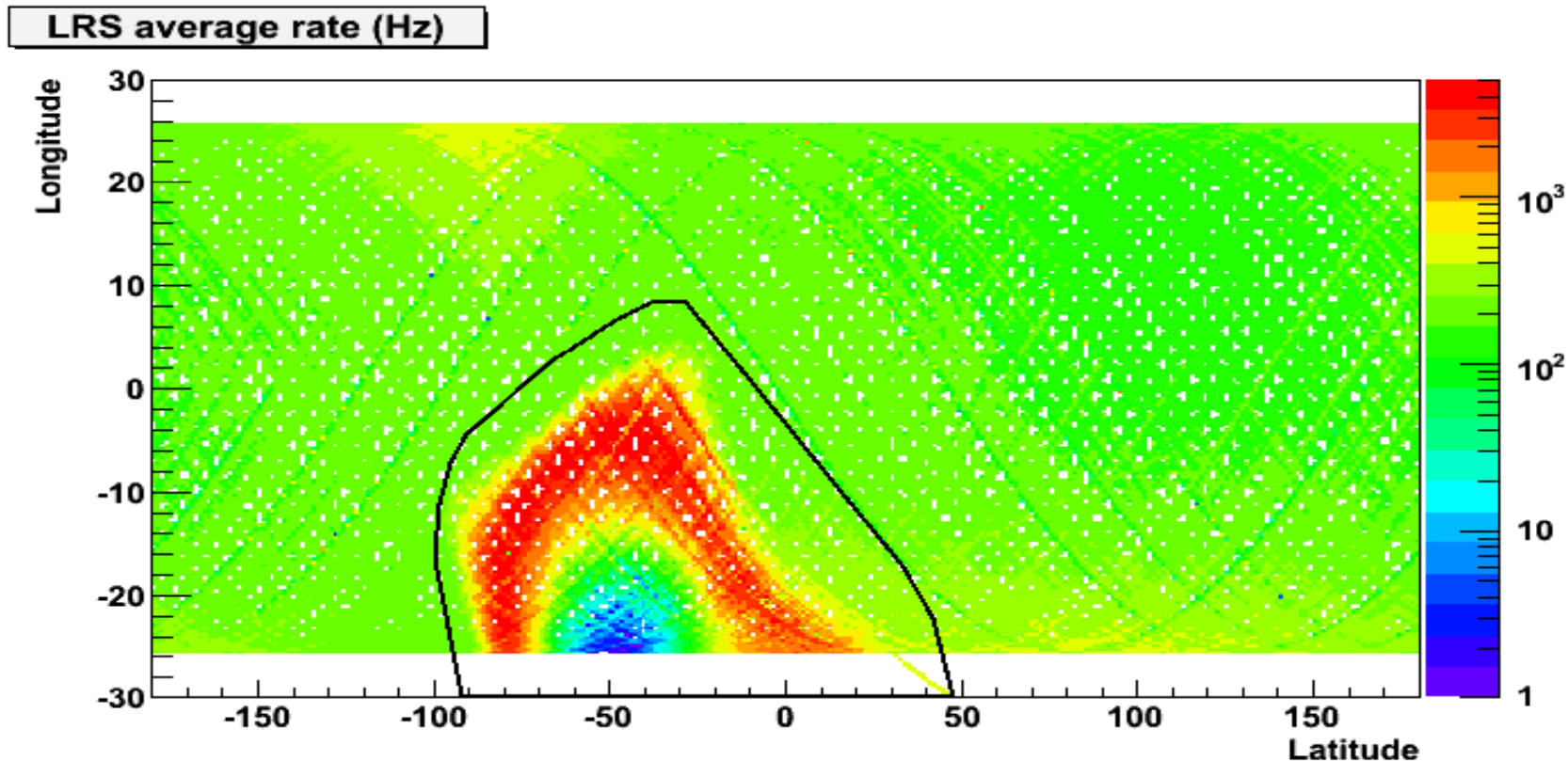
# GLAST in the SAA

- The **South Atlantic Anomaly** (or SAA) is the region where Earth's inner van Allen radiation belt makes its closest approach to the planet's surface. Thus, for a given altitude, the radiation intensity is greater within this region than elsewhere.
- When GLAST in the SAA we don't take physics data. The rate of charged particle is too high to be efficiently filtered out, moreover the high rate can cause saturation effects in the readout electronics.
- With the actual SSA definition GLAST spend  $\approx 18\%$  of the time in it.



# The SAA as seen by the tracker

- There are 4 counters for each tower that register the triggers inside the SAA. They are used to build a map of the SAA in order to study its boundary.
- The map is averaged over all towers and all counters.
- The hole right in the middle of the SAA for the Tracker LRS counters is due to the saturation of the GTFE preamplifier.
- We are working to **reduce the boundary** in order to increase the total live time. The goal is to reduce the time in the SAA to 13% of the total.



Thanks: L.Baldini (INFN-Pi)

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# Conclusion

- GLAST has been on orbit for 47 days now!!! (29<sup>th</sup> July)
- The TKR performance are being monitored and everything behaves as expected
  - It underwent 4g acceleration during the launch
  - It is operating in vacuum (thermosphere  $P < 10^{-5}$  Torr) at  $\approx 15^{\circ}\text{C}$
  - It has worked fine since the turn on without any problem !!!
- We are still in the commissioning phase (-12 days)
  - We are working on fine tuning calibrations
  - We are looking in more details at the TKR performance
- **Stay tuned for the first science results**