



# System Test for the ATLAS pixel detector

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On behalf of the  
ATLAS-Pixel Collaboration



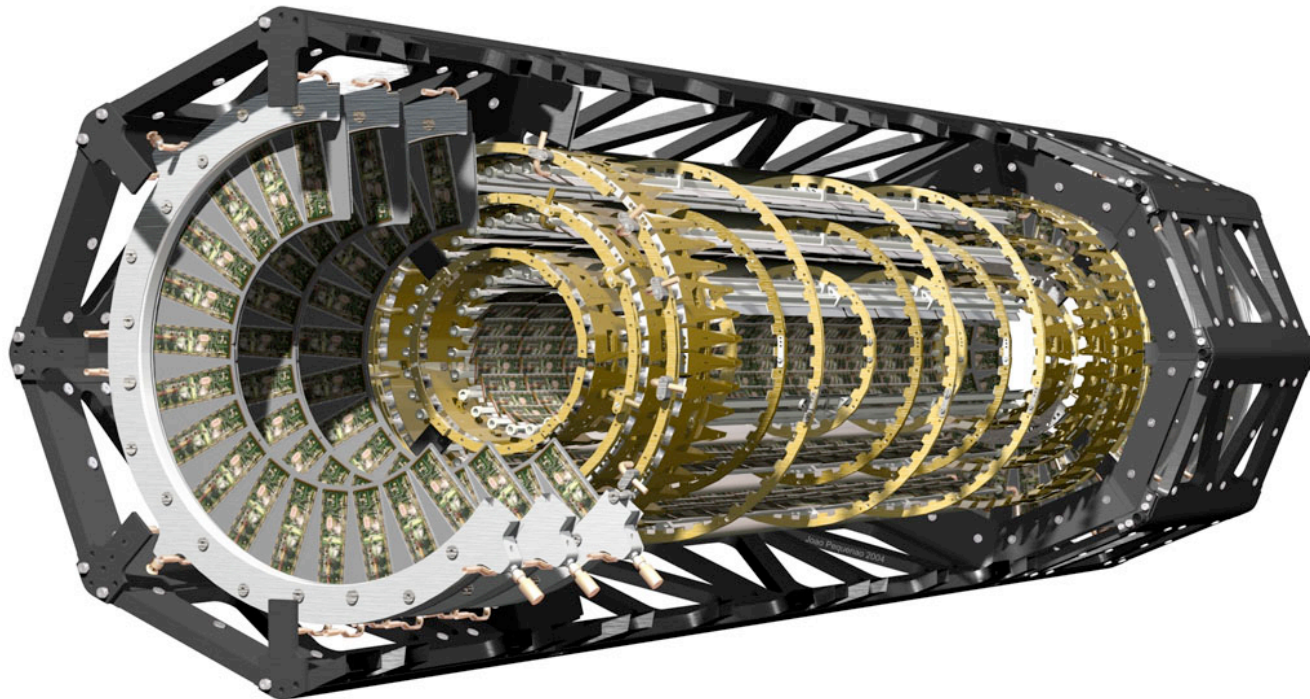
# Outline

- ATLAS Pixel Detector Overview
- System Test Setup
- Calibration runs
- Data taking runs
- Cosmics data analysis
- Conclusions

# ATLAS Pixel Detector

## Pixel Detector design constraints from LHC timing / event-environment physics, ATLAS design

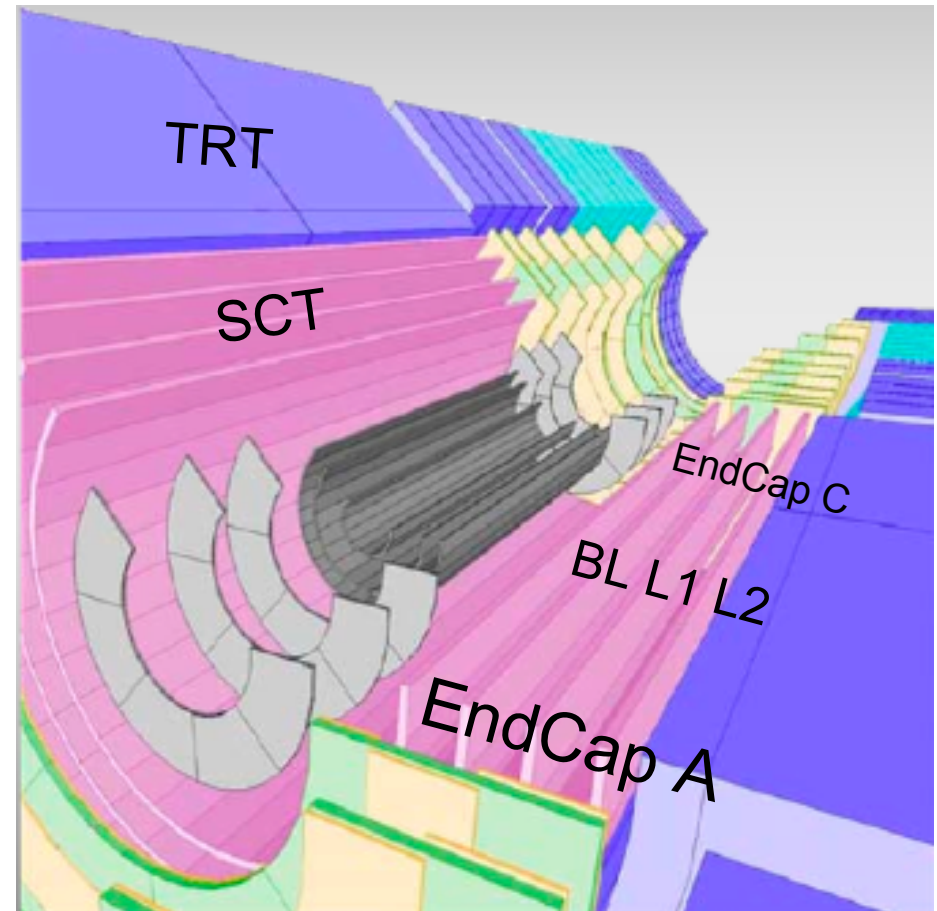
- Tracking in high multiplicity environment (22 interactions  $\sim$  1200 tracks per BC)  $\rightarrow$  high granularity
- Good impact parameter resolution ( $\sigma(d_0) \sim 12 \mu\text{m}$ )  $\rightarrow$  high granularity + low mass
- Distinguishing hits 25ns apart  $\rightarrow$  fast preamp rise time
- 3.2  $\mu\text{s}$  trigger latency (LVL2)  $\rightarrow$  on detector buffering of hits
- High radiation dose  $10^{15} n_{\text{eq}}/\text{cm}^2$  (niel)  $\rightarrow$  low temp., high radiation tolerance  
60 Mrad



# ATLAS Pixel Detector

## Some numbers:

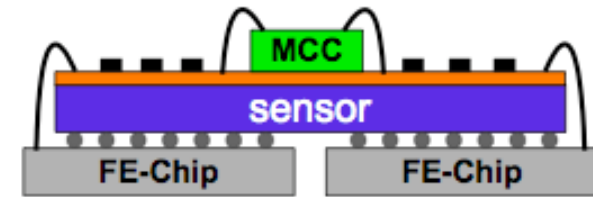
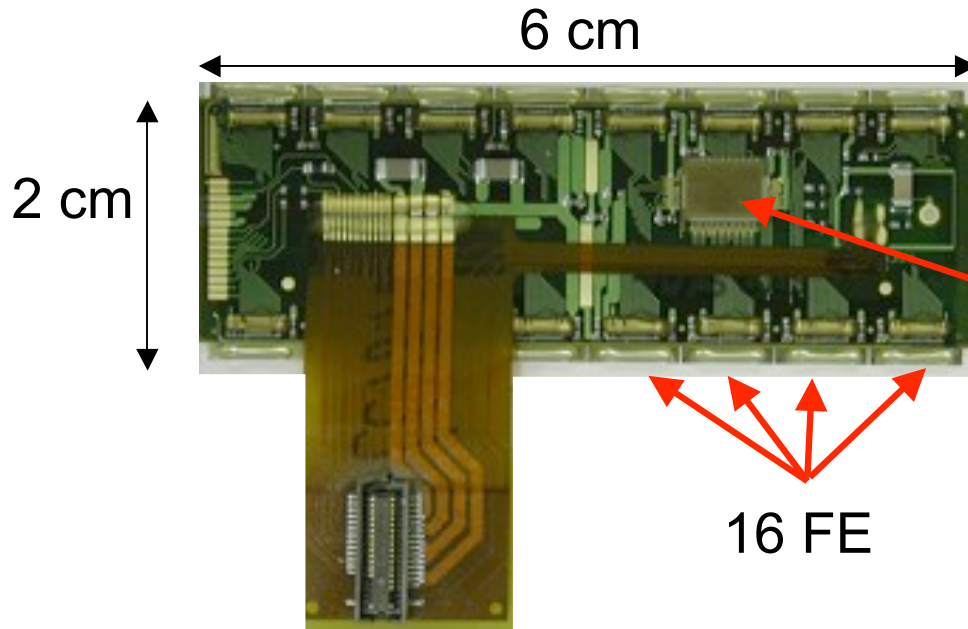
- 3 space points for  $|\eta| < 2.5$
  - 3 barrel layers at  $r = 5, 9, 12$  cm
  - 2 endcaps with 3 layers
  - 80 million channels
  - Area  $1.8 \text{ m}^2$
  - Total mass is  $10\% X_0$  normal to beam
- 
- $\sim 10$  kW operating power in active volume
  - Detector will operate at  $\sim -7^\circ\text{C}$
  - Cooling integrated in local support structure
  - $\text{C}_3\text{F}_8$  evaporative cooling



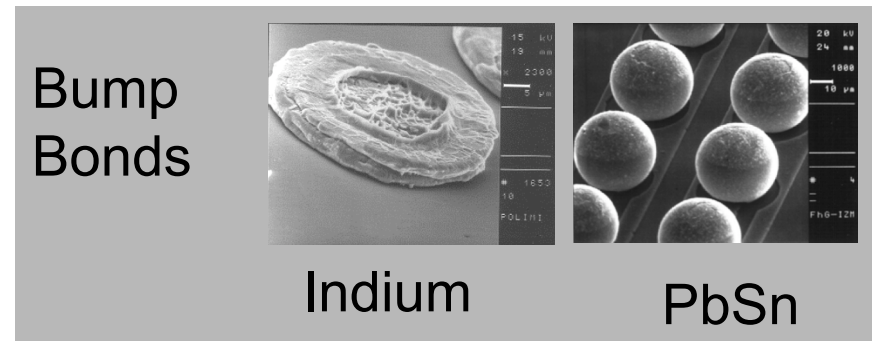
Will be first large-scale active pixel device in operation



# ATLAS Pixel Detector



Module Controller Chip

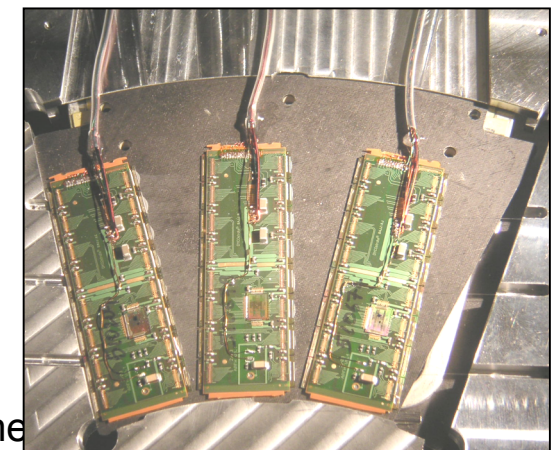


Pixel =  $50 \times 400 \mu\text{m}^2$

2880 pixels x 16 FE = 46080 pixels

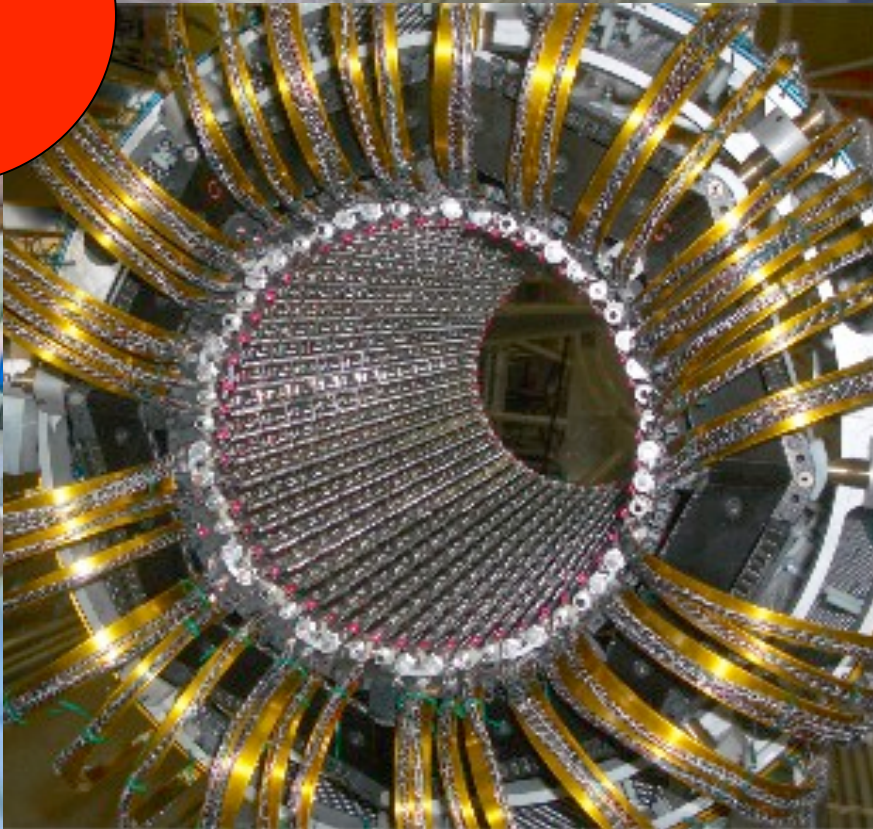
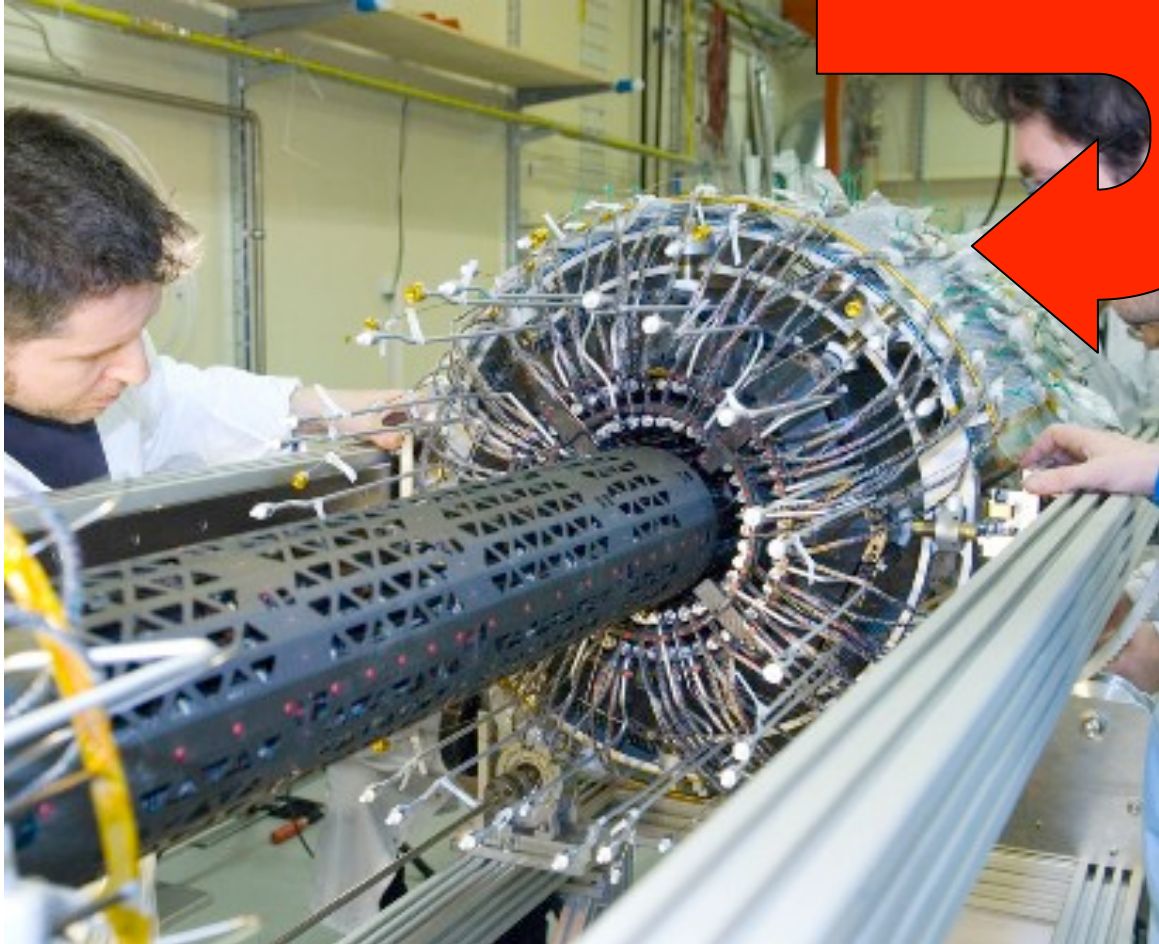
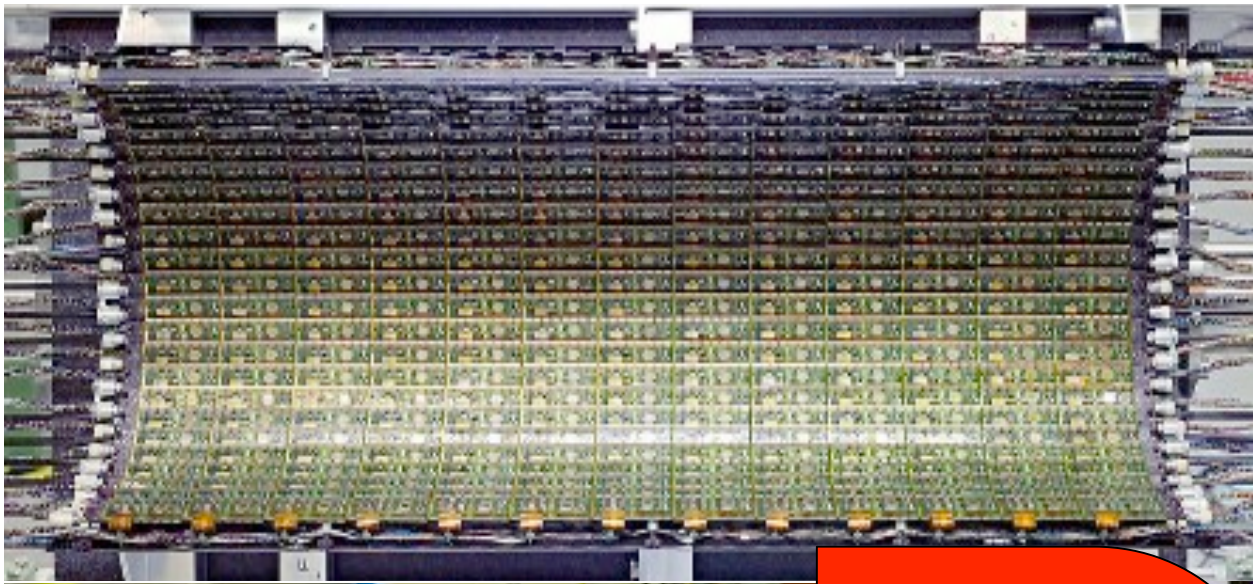
1744 modules x 46080 pixels ~ 80 million channels

Barrels and Disks mount identical modules



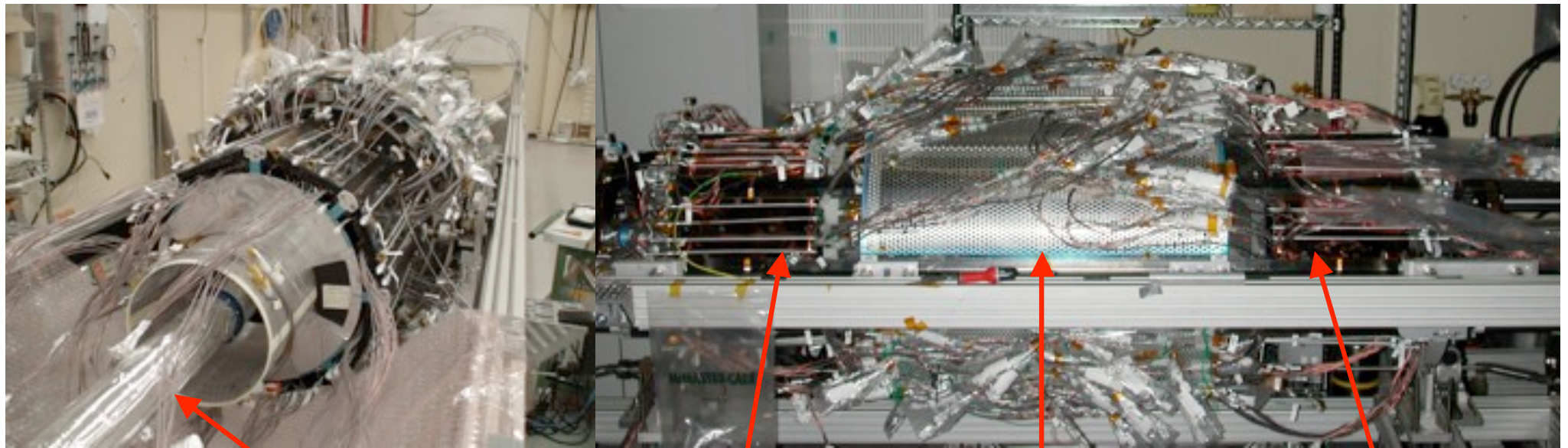
strumente







# ATLAS Pixel Detector



Be beam pipe

Endcap C

Barrel

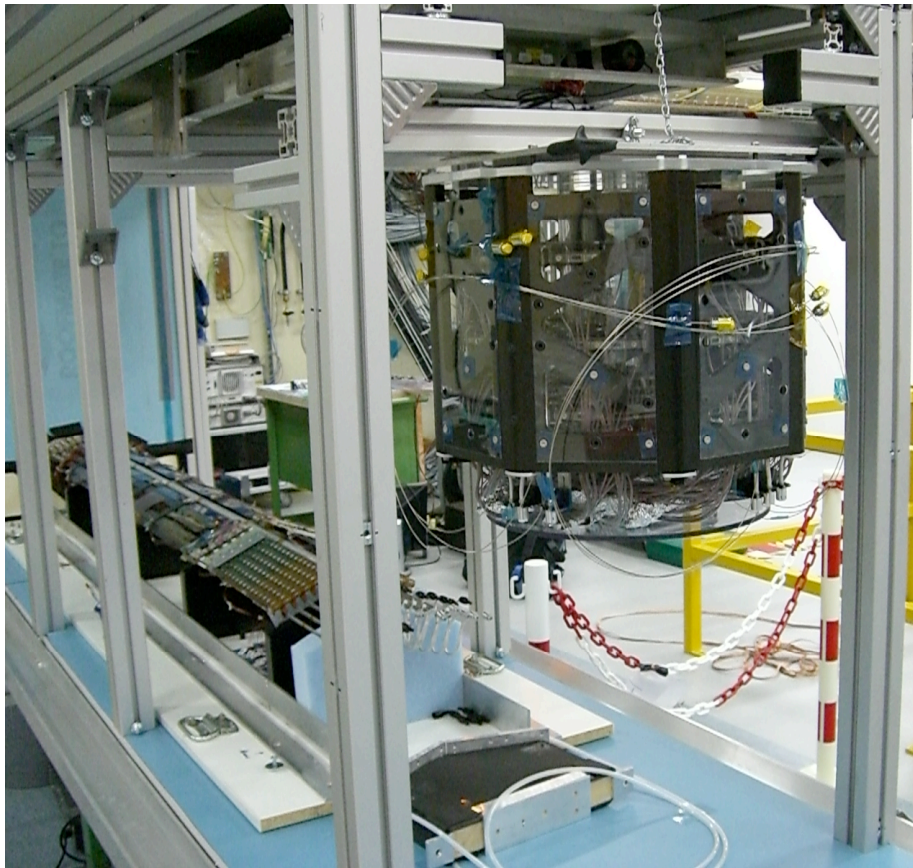
Endcap A

- Integration is well advanced
- EndCaps and Barrels mounted on the beam pipe
- Next steps: service quarter panels integration and connectivity test
- Ready to lower the detector in early June

# System Test: Oct 06- Jan 07

Verify the performance of the detector using **production items**

Test the **complete** infrastructure (HW+SW) on **10% of the entire detector**



## End Cap A :144 modules

### Optical Readout:

24 optoboards (LVDS ↔ optical converters)

### Power Supplies:

24 HV ISEG modules

4 LV Wiener power supplies (Digital + Analog)

24 LV Regulators Channels (custom)

12 SC-OLinks (Optolinks custom power supplies)

### Interlock System monitoring:

Temperatures (modules, optoboards, regulators)

Currents (modules, optoboards)

### ReadOut:

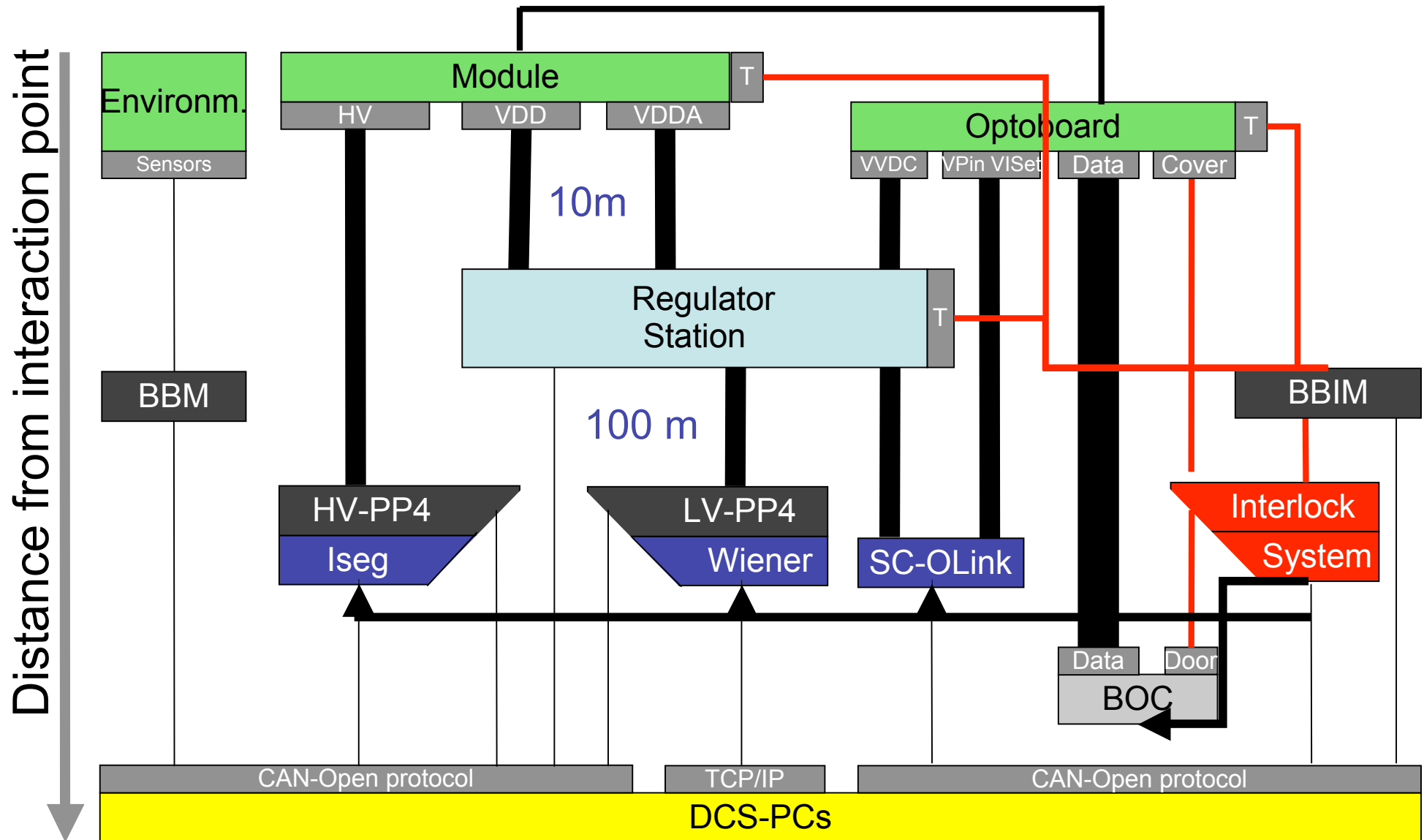
16 Read-Out-Drivers

16 Back-Of-Crate

(opto/electrical conversion + routing)



# Detector Control System (DCS)



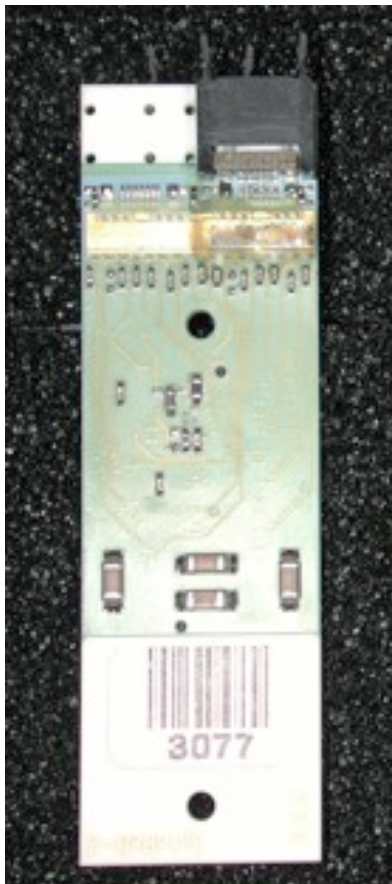
# Services Performance

- The complete services chain has been tested
- Intense debug has been done on many different items giving precious feed back
- Big step forward in DCS software  
(Detector Status, Finite State Machine)

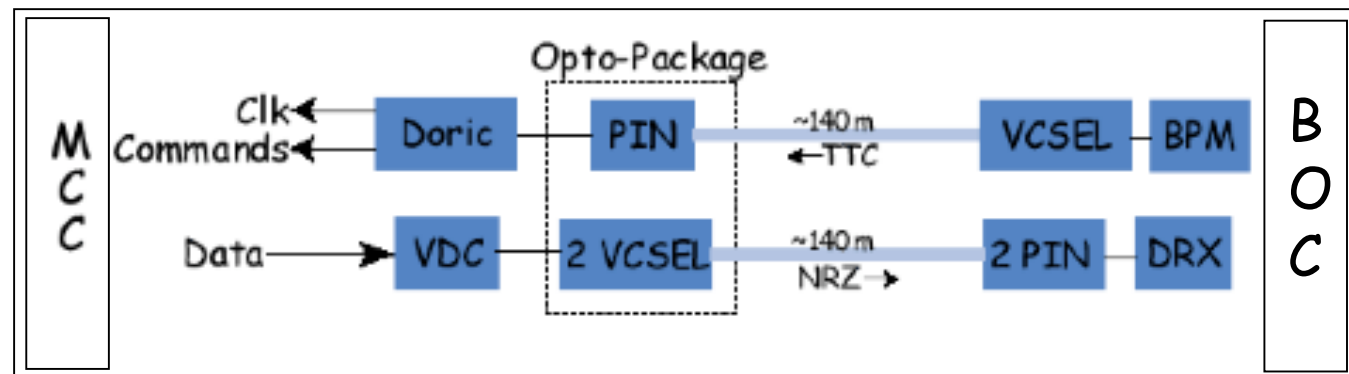
All services are fulfilling the detector requirements



# OptoLinks



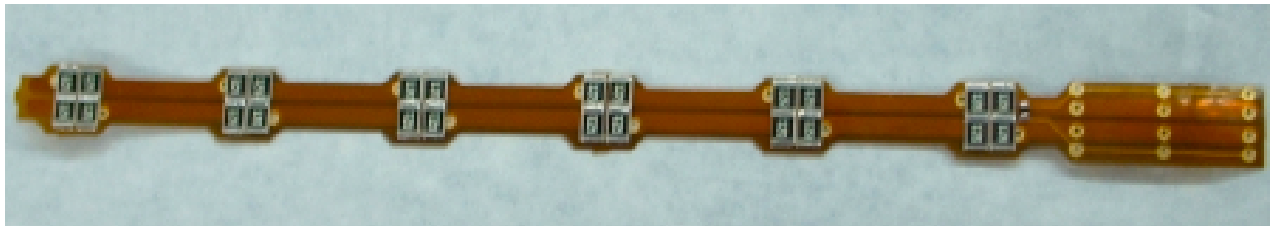
- Optoboards: LVDS signals  $\leftrightarrow$  optical
- Radiation Hard
- Different data bandwidth: 40 Mbit Layer2  
80 Mbit Layer1, disks  
160 Mbit B-layer
- Signals: Clock, Data IN, Data OUT



Needs matching between Optoboards and BOC:  
0/1 discriminating threshold, delays

# OptoLinks Tuning

- We observed that the light output and the channel to channel light spread depends significantly on the optoboard temperature, resulting in non-tunable channels below 5°C
- Increasing driving currents (VISet) does not fix the problem
- We observed that all optoboards are well behaving at ~20°C
- Optoboards have been equipped with heaters
- Thermal management: PID controller for cooling loops and heaters
- Heaters prototype produced and tested

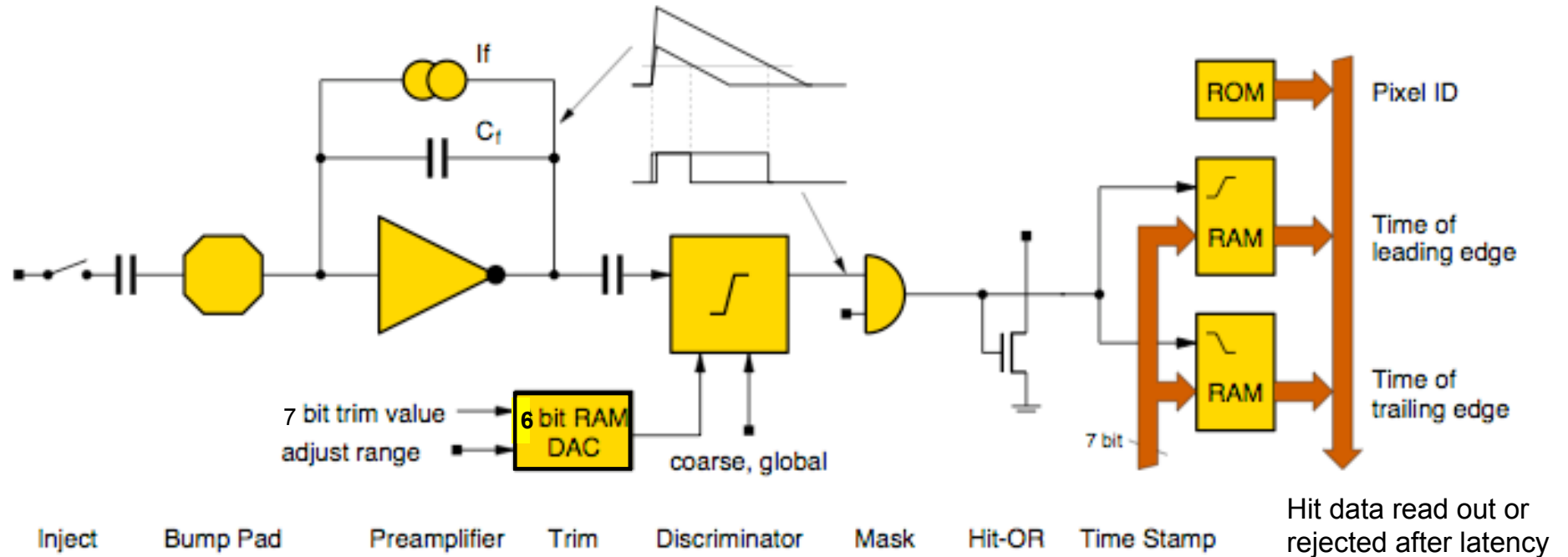


100% tunable channels in the System Test



# Calibration Runs: Module tuning

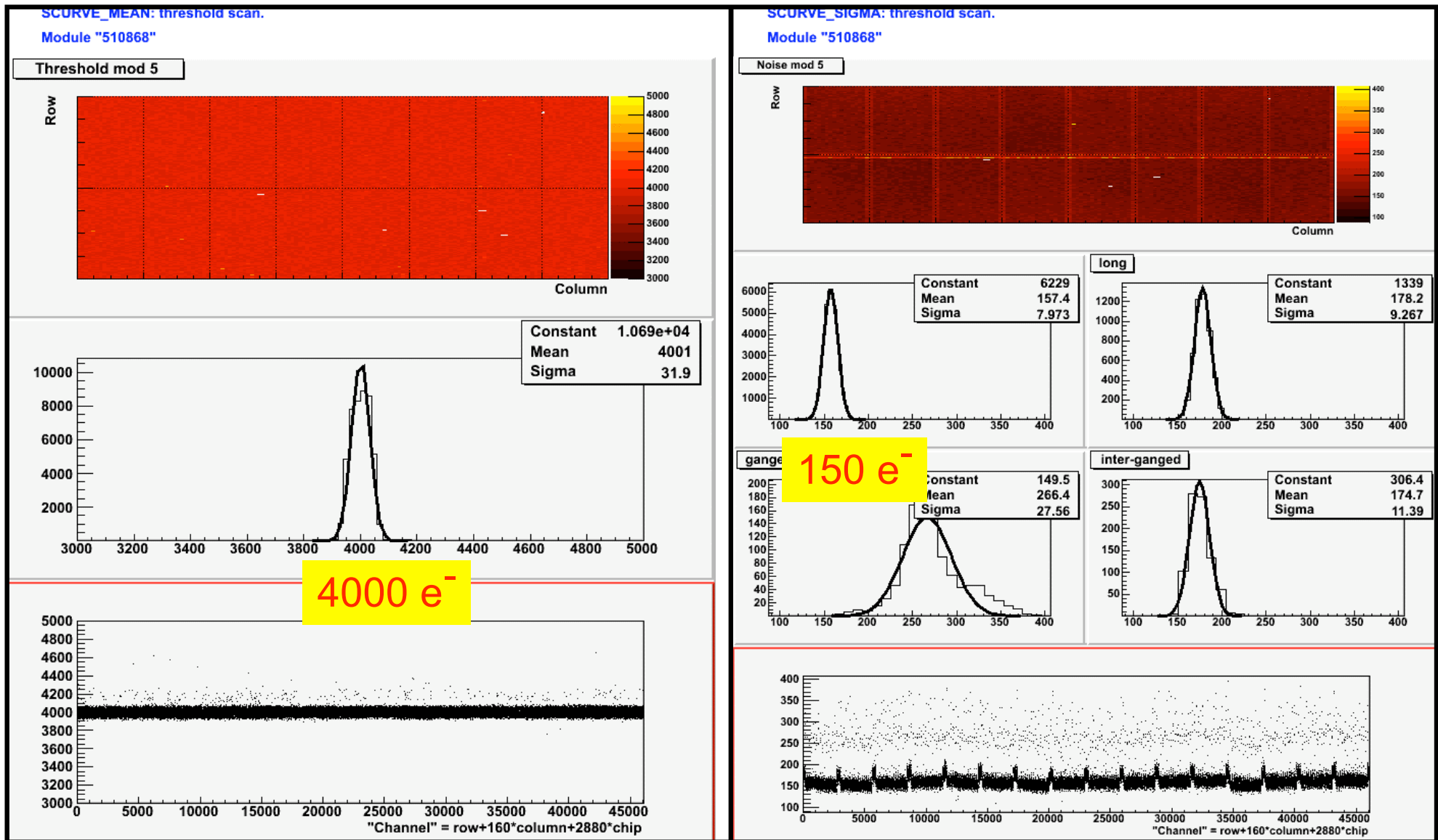
## Front End Schematics



FE chip can be tuned using several DACs among which:

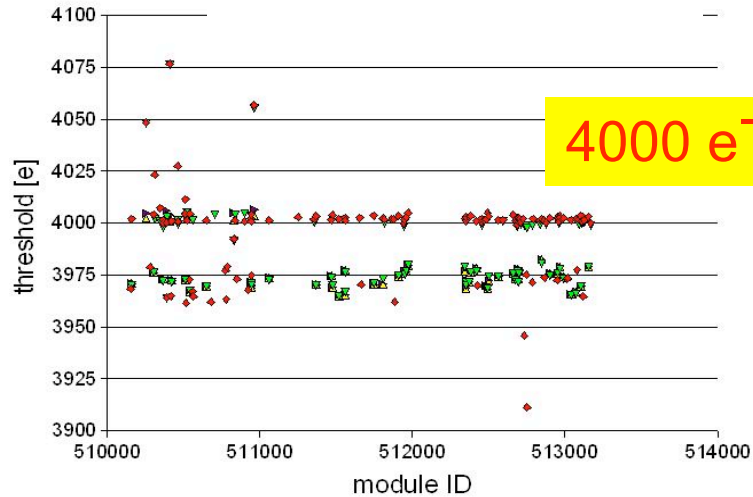
- Coarse and fine tune Discriminator Threshold
- PreAmp Feedback current: tune Time over Threshold (TOT)

# Tuning Performance



# Tuning Performance

## Tuned Thresholds

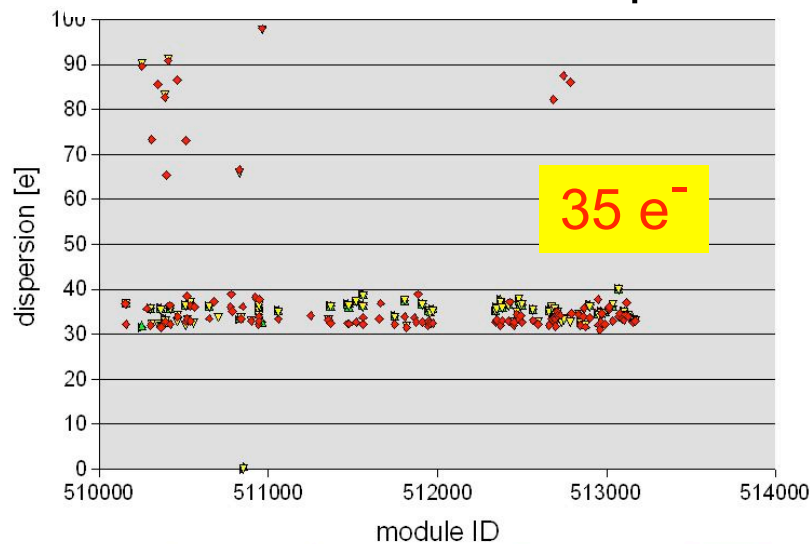


MIP: ~ 20000 e<sup>-</sup> in 250 um silicon  
~ 13000 e<sup>-</sup> including charge sharing  
~ 6000 after 10 years at the LHC

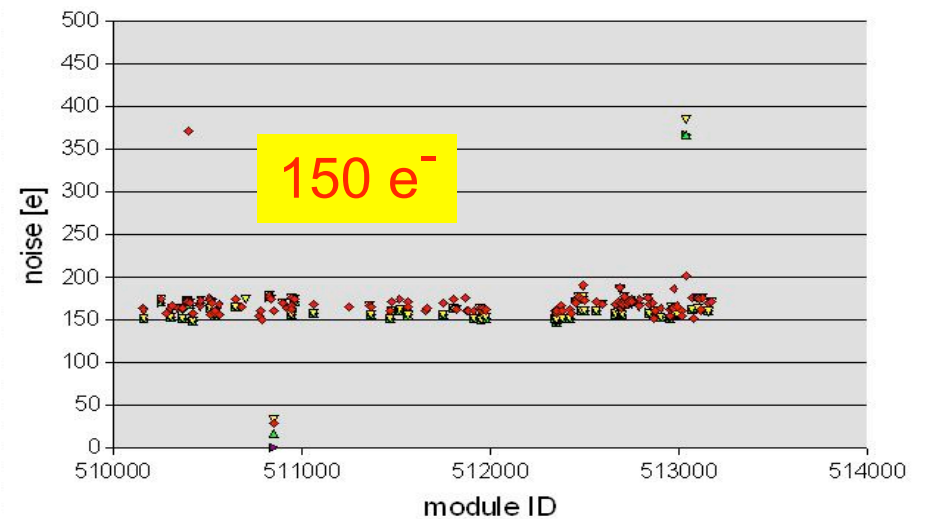
Typical Noise ~ 150 e<sup>-</sup>  
~ 200 e<sup>-</sup> after 10 years at the LHC

**S/N > 30 throughout the LHC lifetime**

## Tuned Thresholds dispersion



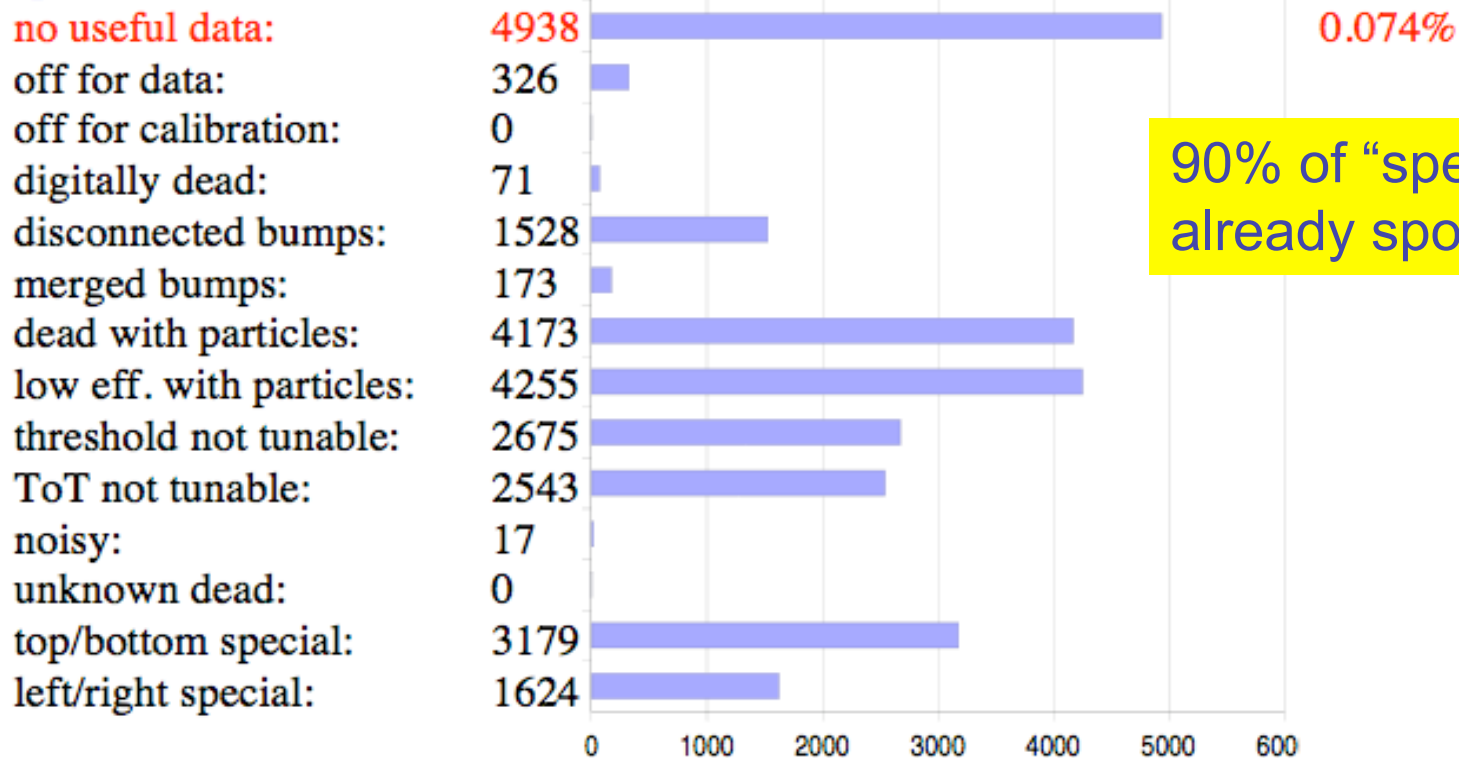
## Noise





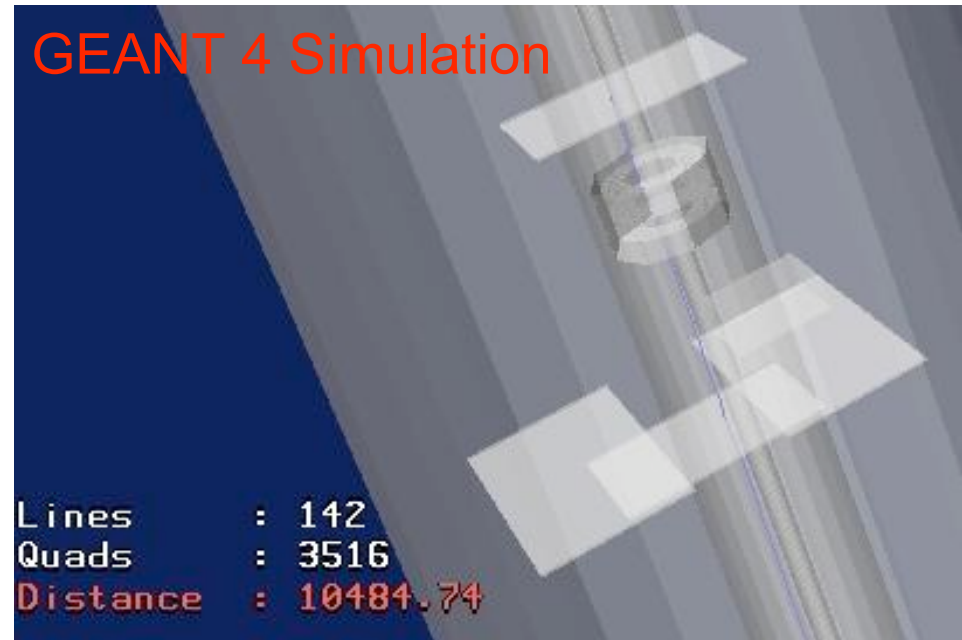
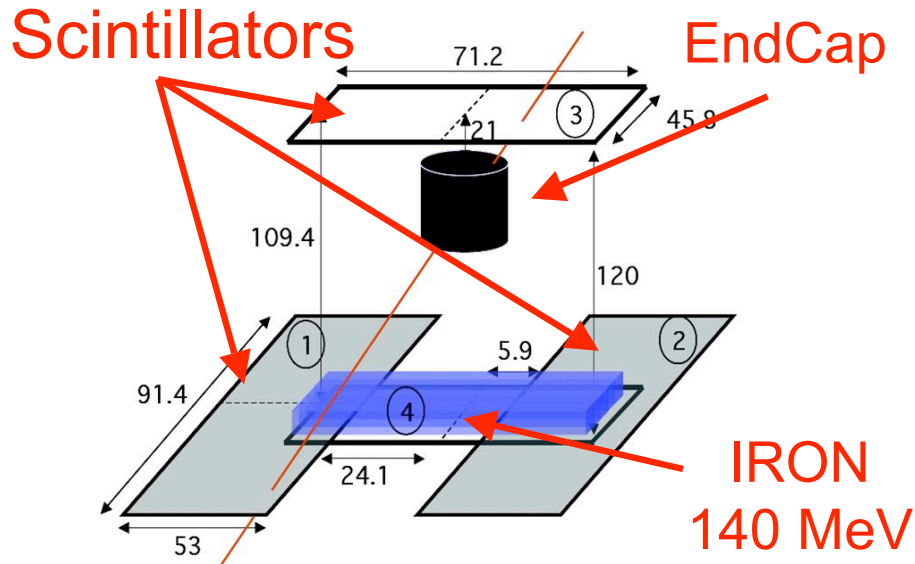
# Random Trigger Data

## “Special Pixels”



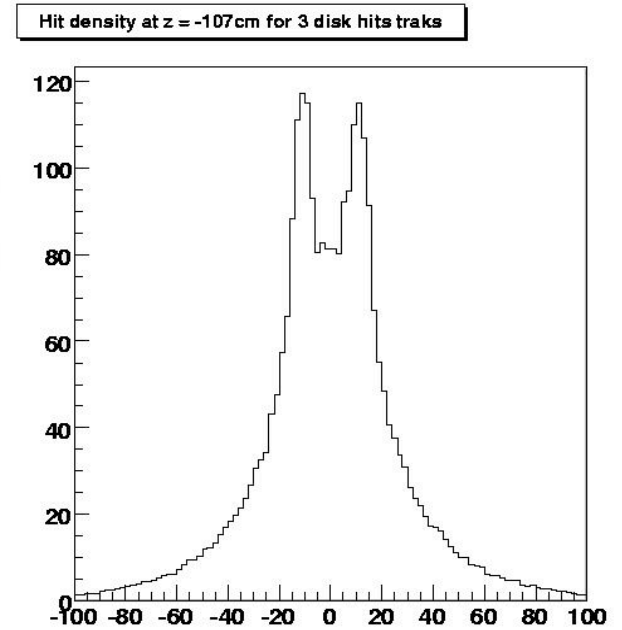
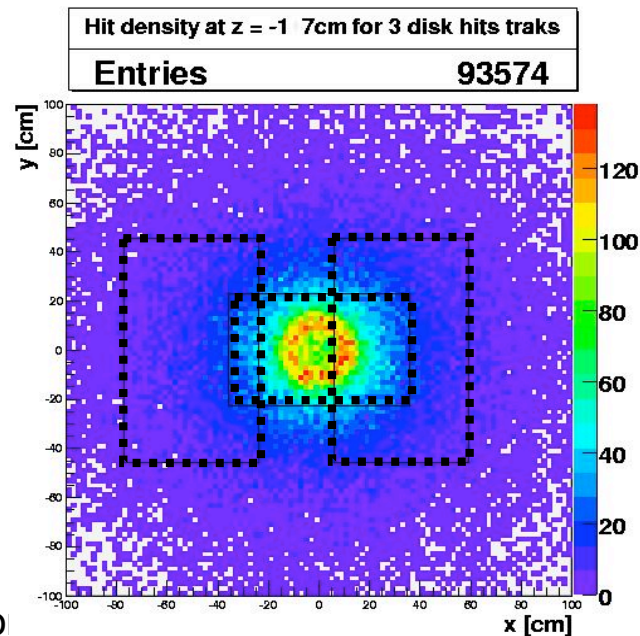
30  $10^6$  noise events analysed:  
~  $2 \cdot 10^{-7}$  averaged occupancy  
~ 90% of the noisy pixels are classified as “special”

# Data Taking Runs: Setup



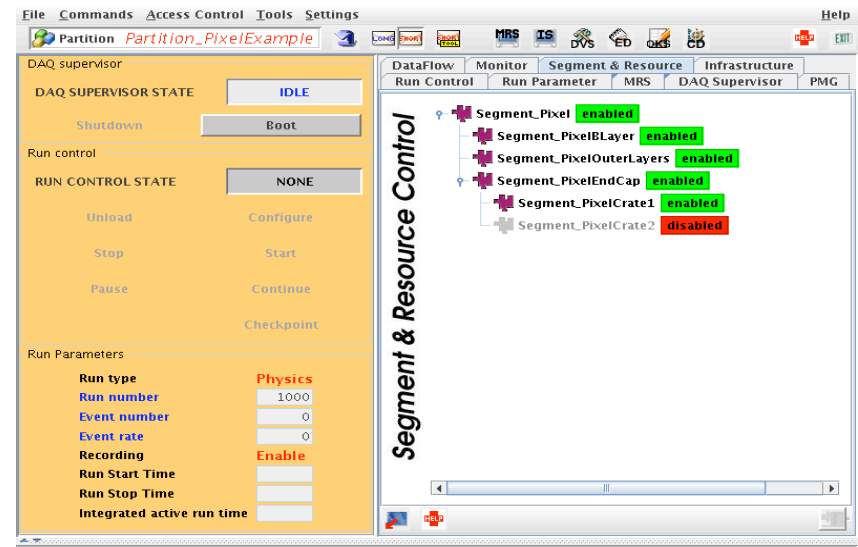
**29 modules disabled  
out of 144**

**Measured count = 15.7 Hz  
Full simulation ~ 18 Hz  
~10% reco efficiency**



# Data Taking Runs

Cosmics data taking represents a realistic test bench to test the complete data acquisition chain before lowering the detector in the pit.



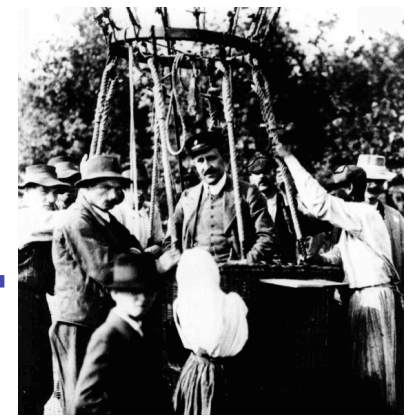
The main Pixel Software application is the **crate run controller** developed in the framework of the **common ATLAS DAQ infrastructure**.

The detector is described as a **hierarchy of segments and partitions**. Each segment corresponds to a **controller application** each implementing a common set of state transition.

In particular the controllers managing VME crates are implemented in the **ROD Crate DAQ (RCD)** framework, providing interfaces to the DAQ software infrastructure (configuration DB, messages, OH...)

The detector specific developers must provide plugins to control the HW components (RODs, TIM...) and the RCD propagates the state transitions to the plugins

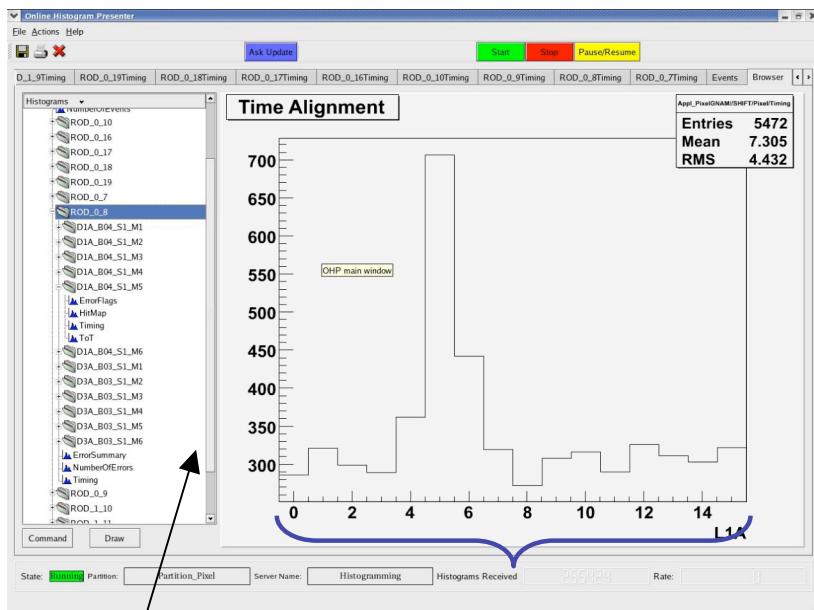
# First Evidence of Cosmics...



Modules Timing

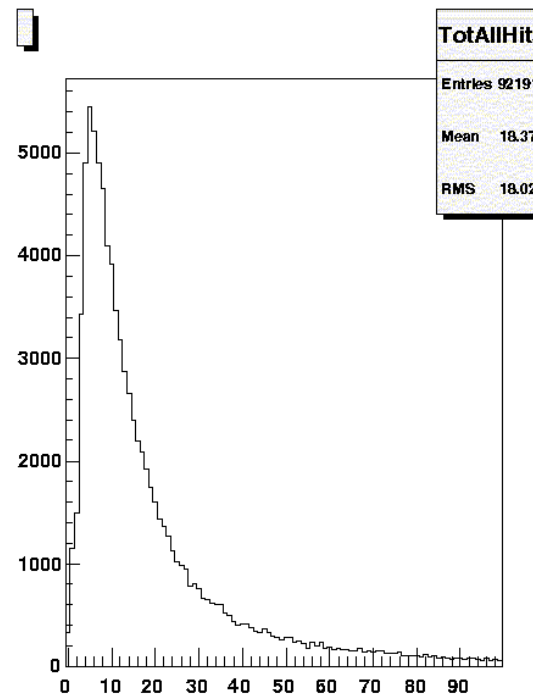
Random Triggers

Cosmics Triggers

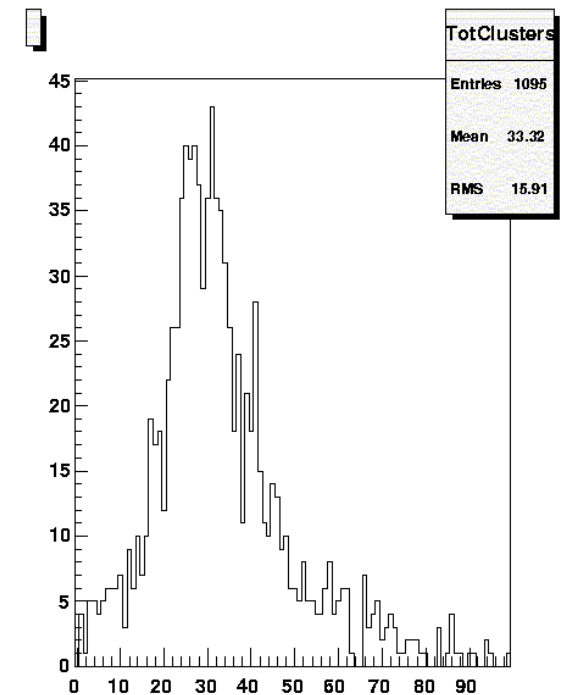


16 consecutive L1A trigger

Online monitoring



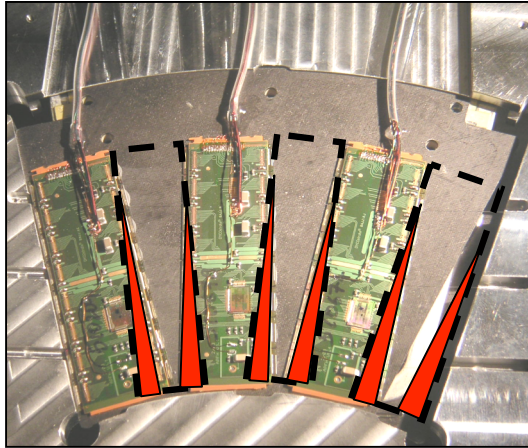
TOT units



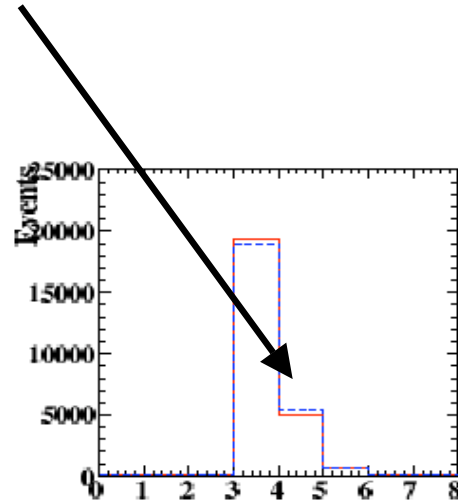
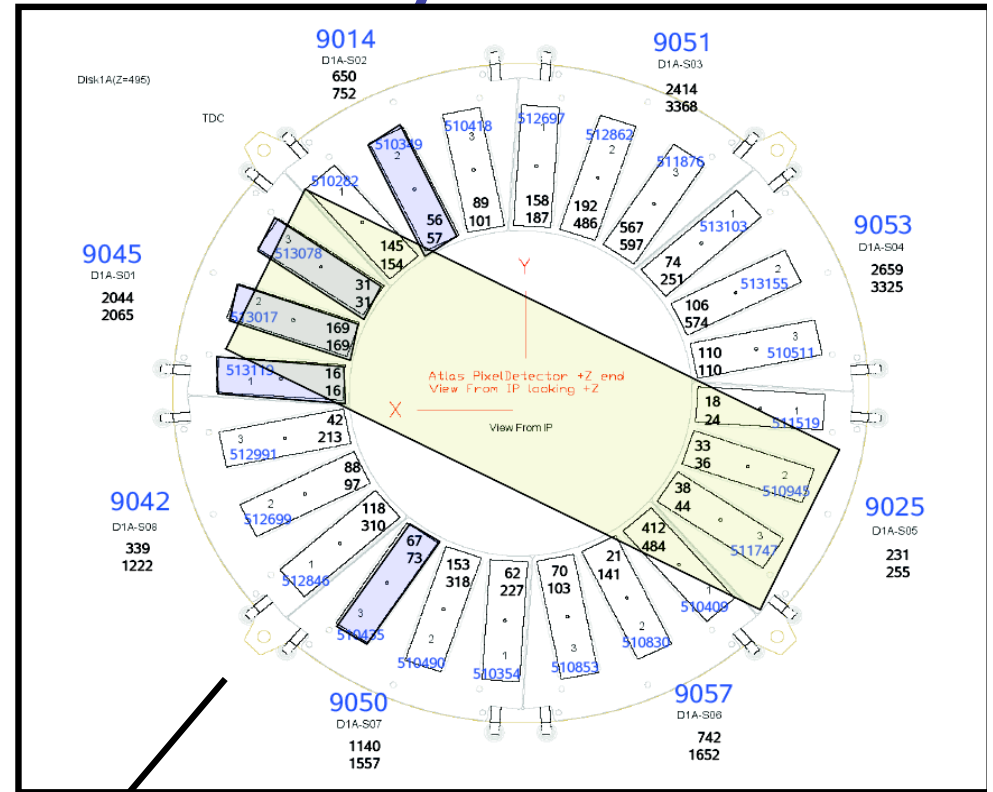
TOT units



# Cosmics Data Analysis

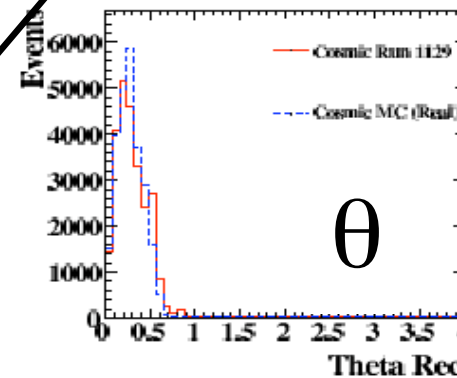
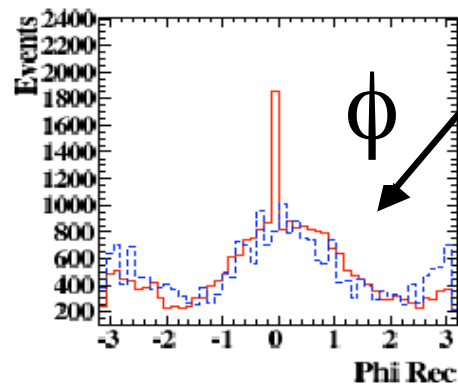


Overlap regions



#clusters on track

— MC  
— Data



# Cosmics Data Analysis

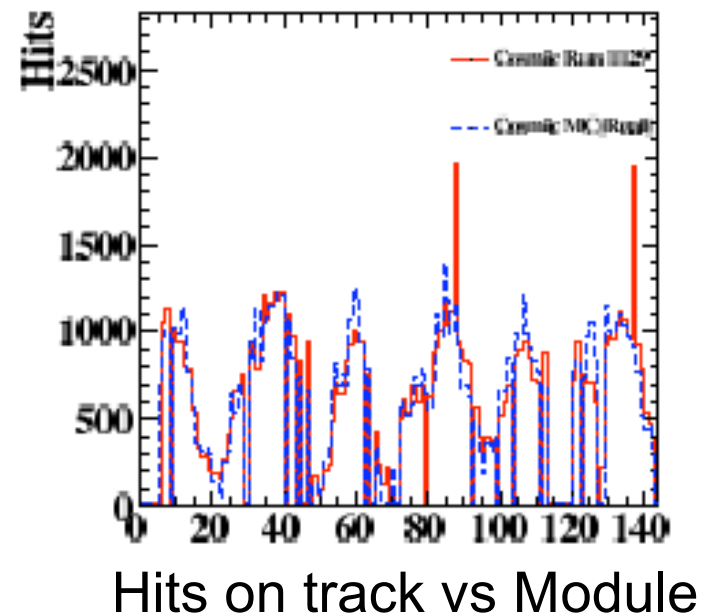
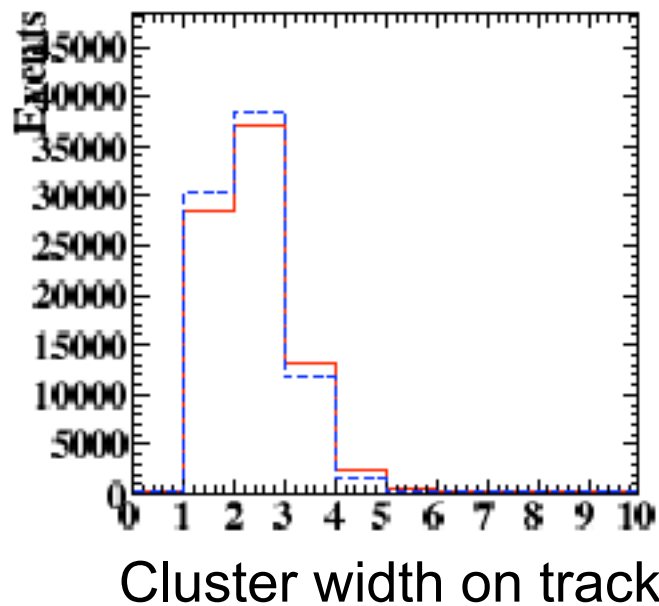
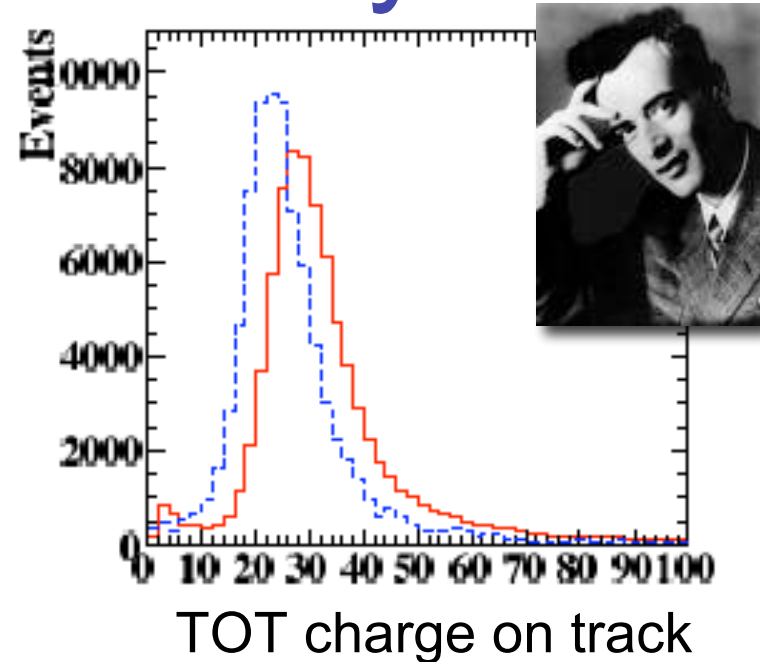
TOT distribution:

MC tuned on Test Beam

Correct shape: TOT calibration OK

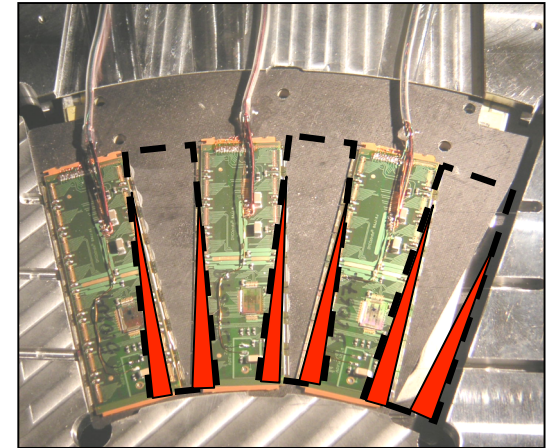
Shift under investigation probably  
different threshold

— MC  
— Data



# First Alignment Studies

Use the overlap between adjacent modules to derive the relative alignment between modules.



There are > 10% overlap between adjacent modules with  $dz = 4.2$  mm. The extrapolation in  $z$  produces a negligible error ( $< 1 \mu\text{m}$  in  $x$  and  $< 5 \mu\text{m}$  in  $y$ )

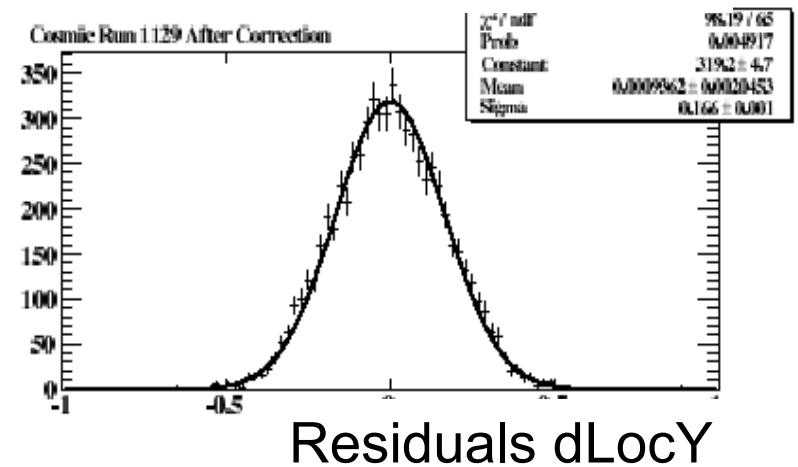
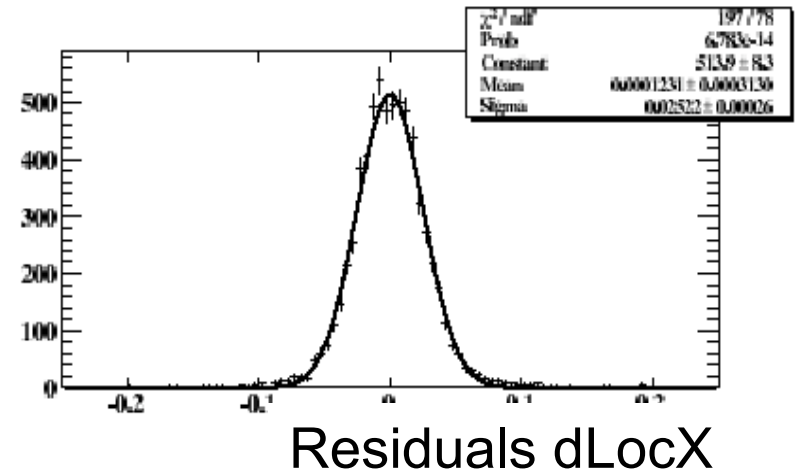
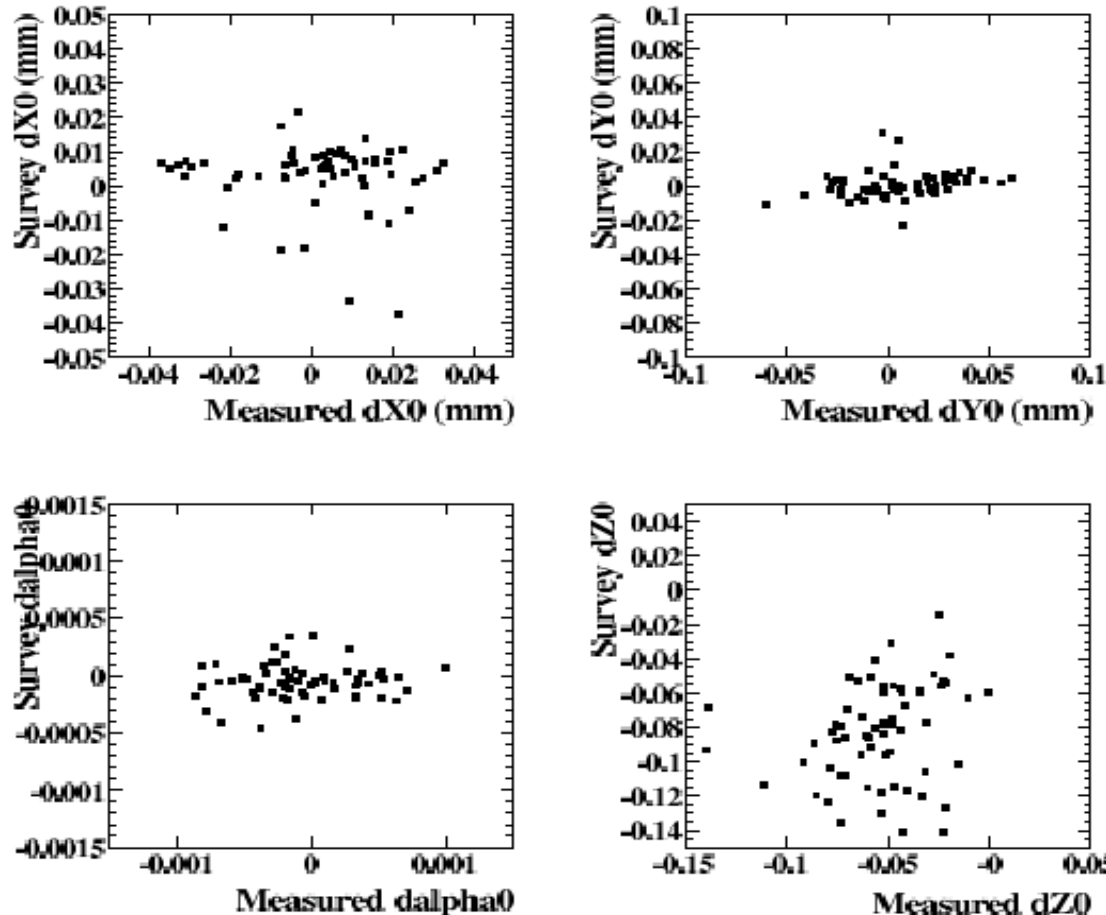
The module is described as a rigid body in the disk plane:

- Shift in  $X_0$  in local  $X$
- Shift in  $Y_0$  in local  $Y$
- Shift in  $Z_0$  in local  $Z$
- Rotation  $\alpha_0$  along  $Z$  axis



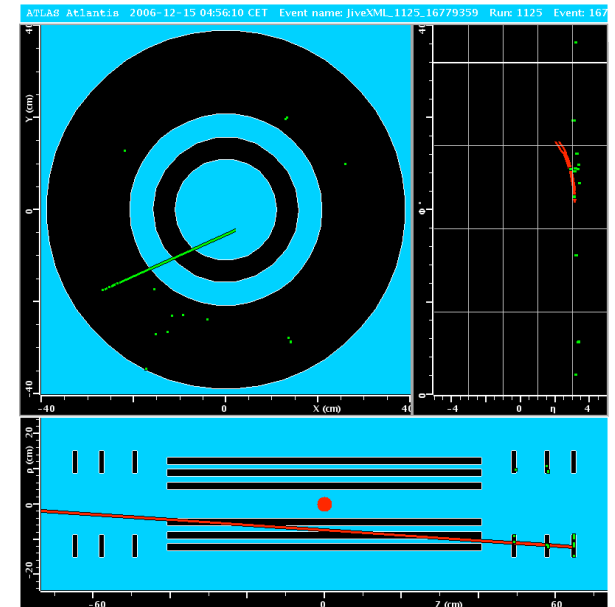
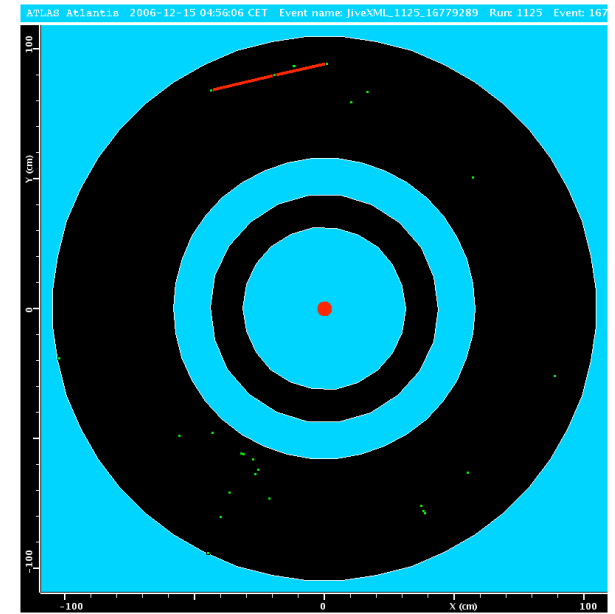
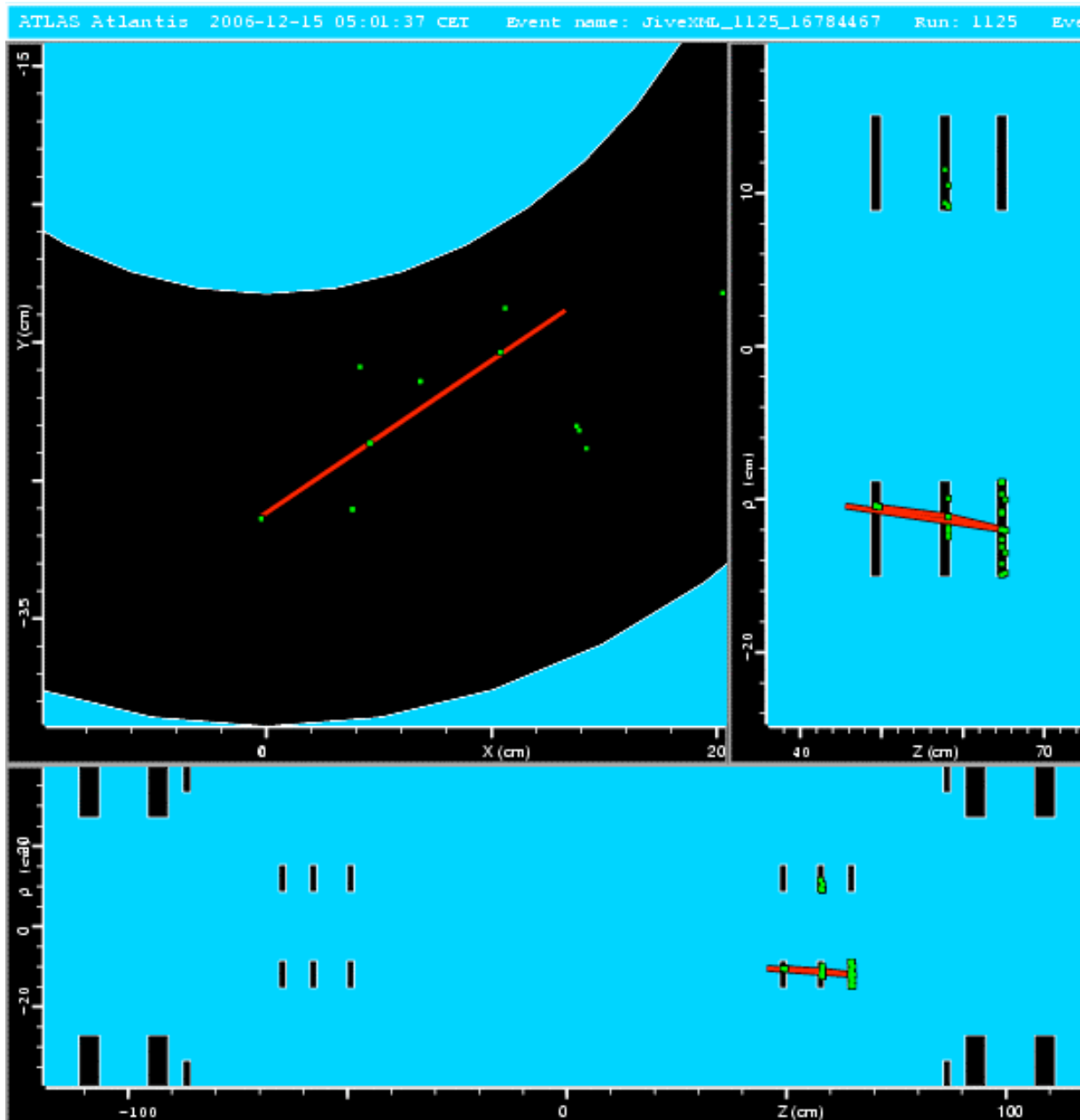
# First Alignment Studies

Survey Alignment vs Cosmics Alignment  
For modules with more than 50 hits in the overlapping regions



**Resolution**  
X : 17.8  $\mu\text{m}$  (MC 15.8  $\mu\text{m}$ )  
Y : 117  $\mu\text{m}$  (MC 117  $\mu\text{m}$ )

# ATLANTIS event display



# Conclusions

## System Test:

- The system test on 10% of the detector is completed
- Intense debug has been done on many different items giving precious feed back
- DCS, calibration and data taking software are well advanced
- The preliminary studies on cosmics data confirm MC expectations

The dedicated effort of many persons made the system test a **SUCCESS**

## Detector status:

- Integration is well advanced
- EndCaps and Barrels are mounted on the beam pipe
- Next steps: service quarter panels integration and connectivity test

**We will be ready to lower the detector in early June**



# BACKUP

# 40 Mbit optolink tuning

Phase space:

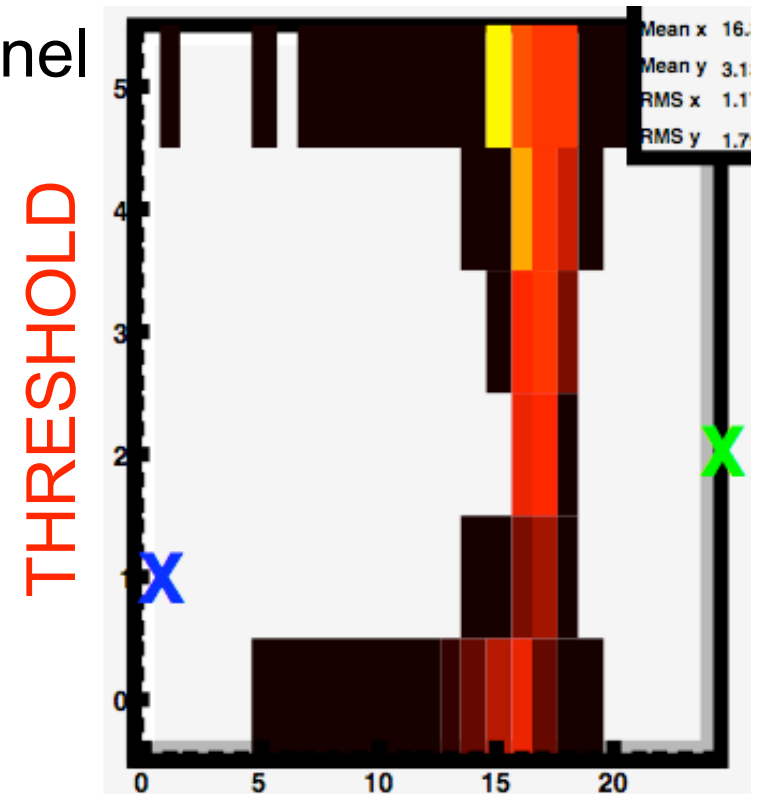
[VSet]x[Threshold]x[RXdelay]x[MSR]x[temperature]

Procedure:

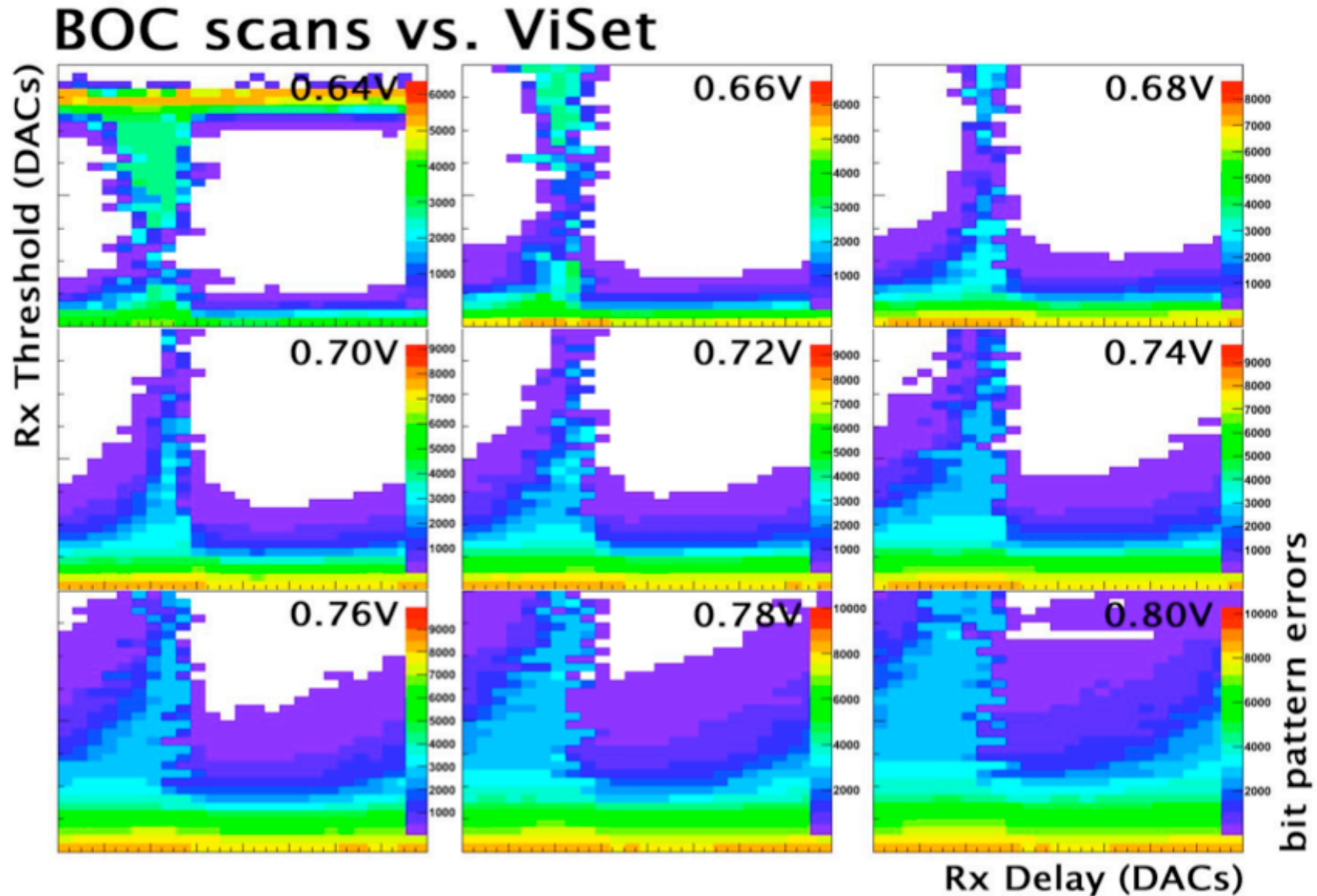
- THR\_RXdel\_scan at different VSet -> find best VSet  
VSet one value /optoboard looking at EFR area
- Find the best THR/Rxdel for each channel

Two scanning engines:

- Pseudo-Random  
(sensitive to slow turn on effects)
- Periodic

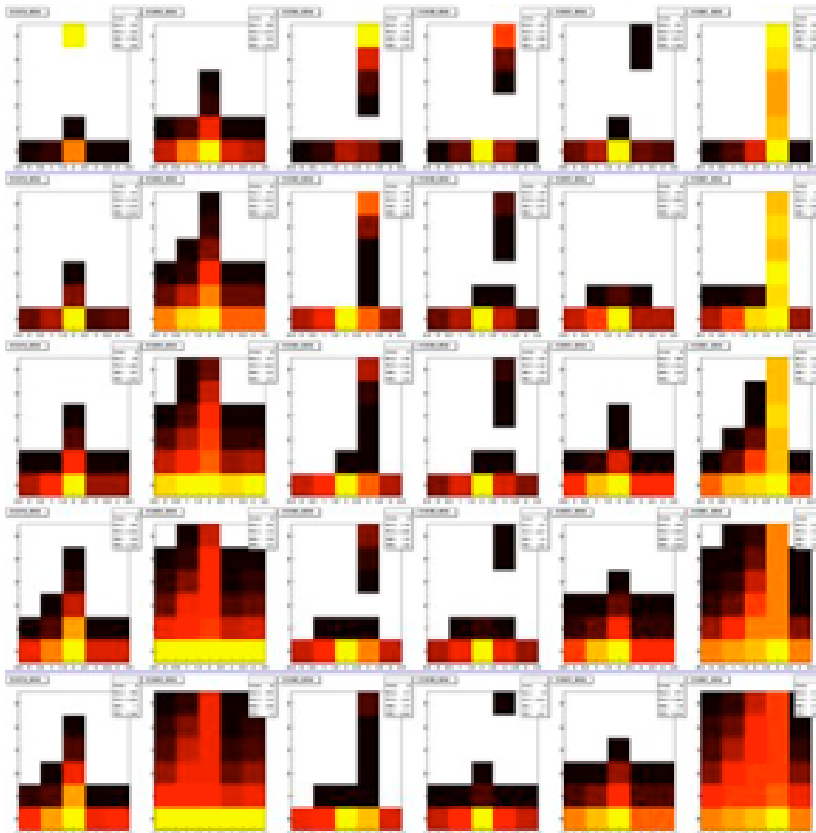


# ViSet BER dependence

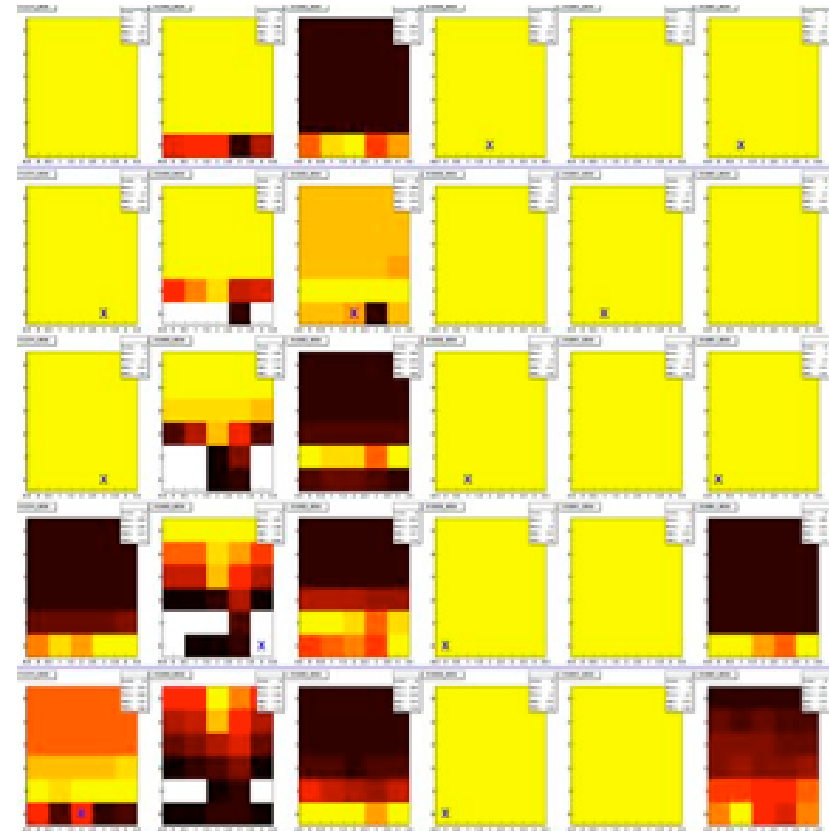




# Temperature dependence

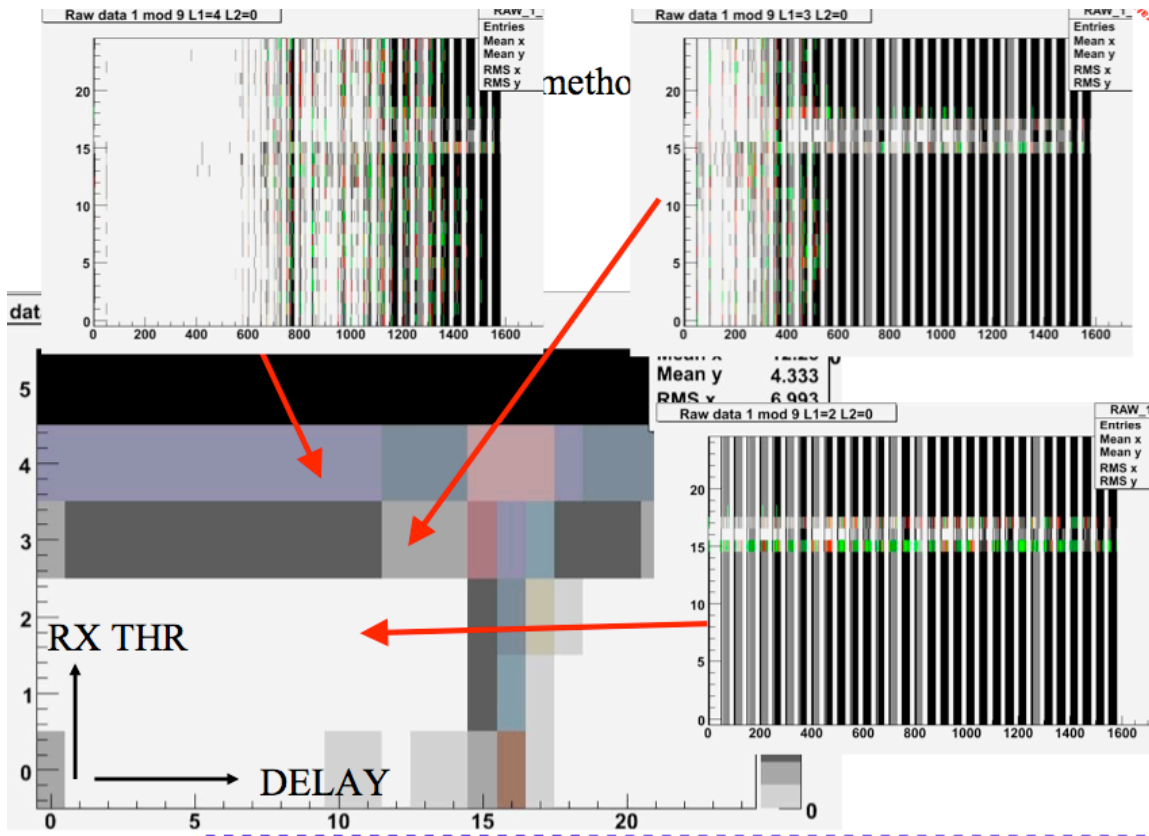


T = + 10 C



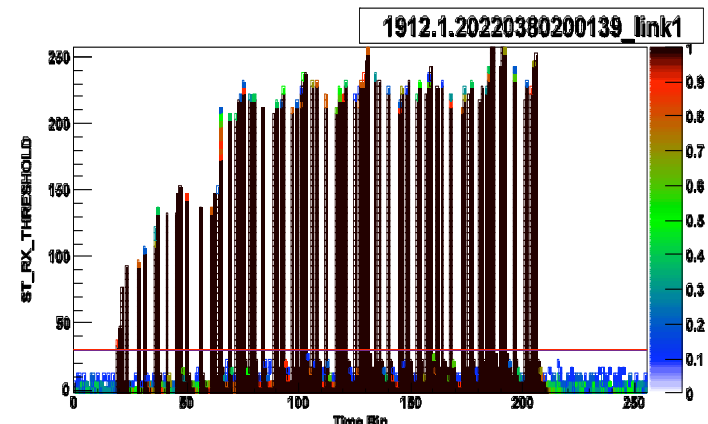
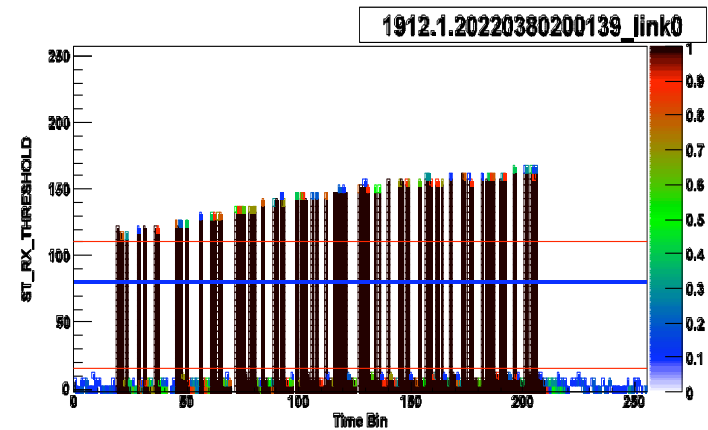
T = -10 C

# VCSEL Slow Turn On



Transmission BER

RX amplitude



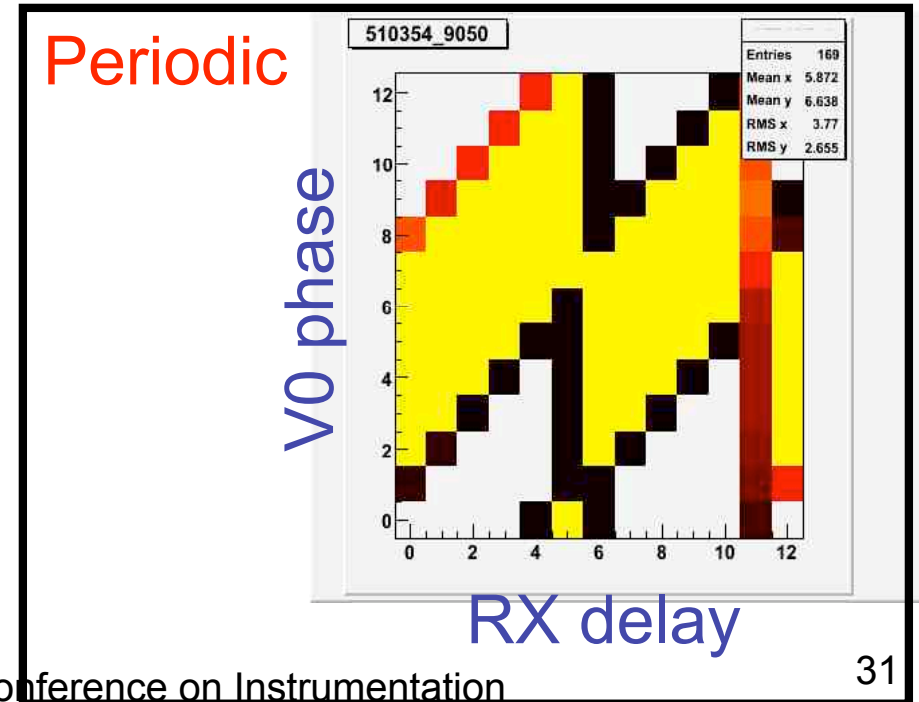
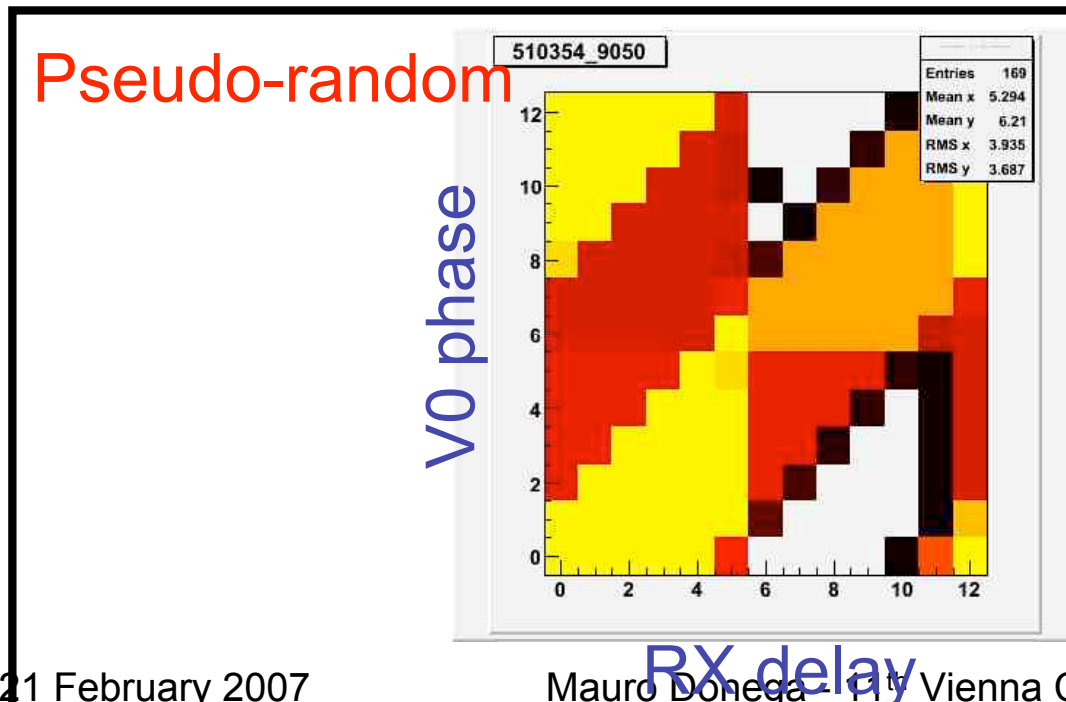
# 80 Mbit optolink tuning

Phase space:

[V0]x[Threshold]x[RXdelay]x[V0]x[MSR]x[temperature]

Procedure:

- Fix best V0Set/threshold (40 Mbit tuning)
- Scan V0\_RXdelay find the optimal point
- V0=0 is typically the optimal value (V0 is one value /BOC)





# Example of problematic channels at 80 Mbit

