

Probing beyond the Standard Model at Low Energy

Tim Chupp
University of Michigan
Ann Arbor



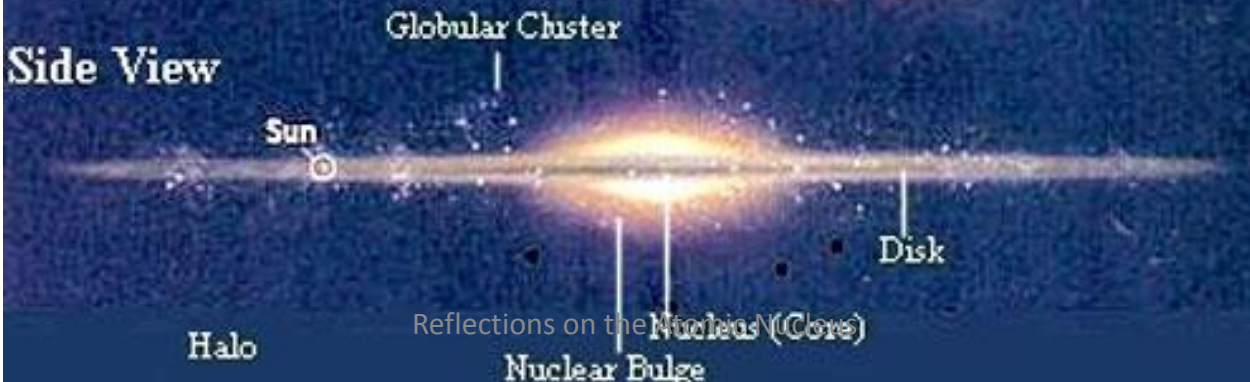


Face-on View

Rotation



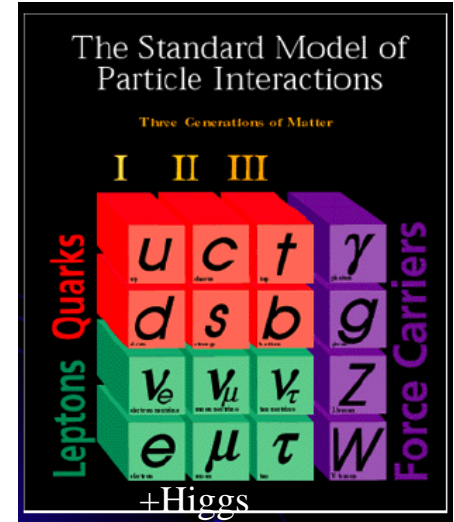
Side View



Reflections on the Milky Way

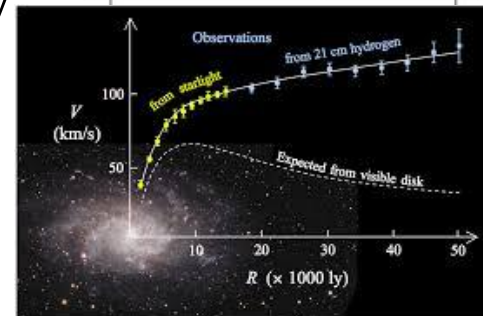
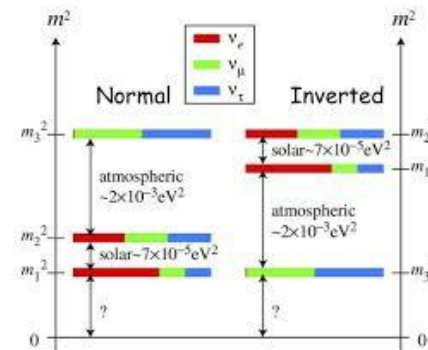
What we know

- The Standard Model of particles and forces
 - QFT with radiative corrections
 - Massive neutrinos
- Gravity weak and strong (GR) => dynamics (we can send a satellite to Pluto!)
 - Dark matter/Dark energy
- The universe is MATTER dominant



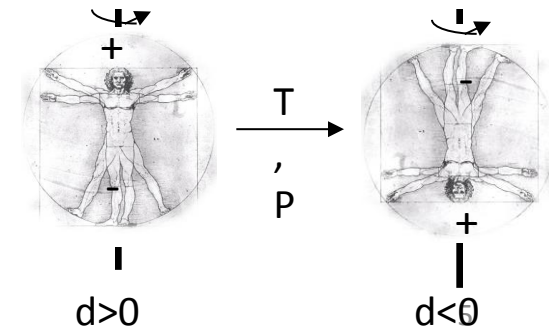
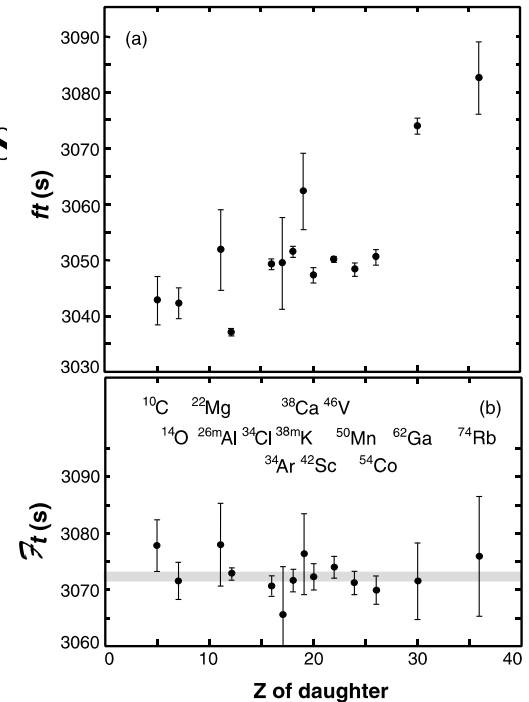
What we don't know

- How neutrinos acquire mass
 - Dirac/Higgs ($m_\nu < 10^{-6} m_e$; $m_u \sim m_d \sim 10^{-6} m_t$)
- What is *THE* dark matter/what is dark energy
- What made the universe MATTER dominant
 - Baryogenesis



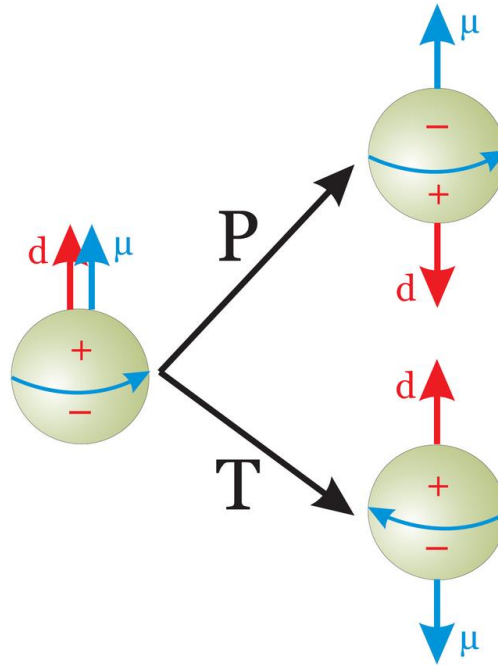
Where to look

- High energy, rare decays, exotic processes (astro)
- Muon $g-2$
 - Precise SM predictions (with uncertainties)
 - 3.7σ difference of experiment and theory
- Beta-decay: neutron, nuclei
 - Probe BSM through T violation
 - Measure CKM element V_{ud}
- Permanent electric dipole moments
 - SM predicts EDMs \ll current sensitivity
 - Note θ_{QCD} contribution (inexplicably small)



EDMs

$$\vec{d} = \int \vec{r}(\rho_+(\vec{r}) - \rho_-(\vec{r}))dV = d\vec{J}$$



Put this in E and B fields

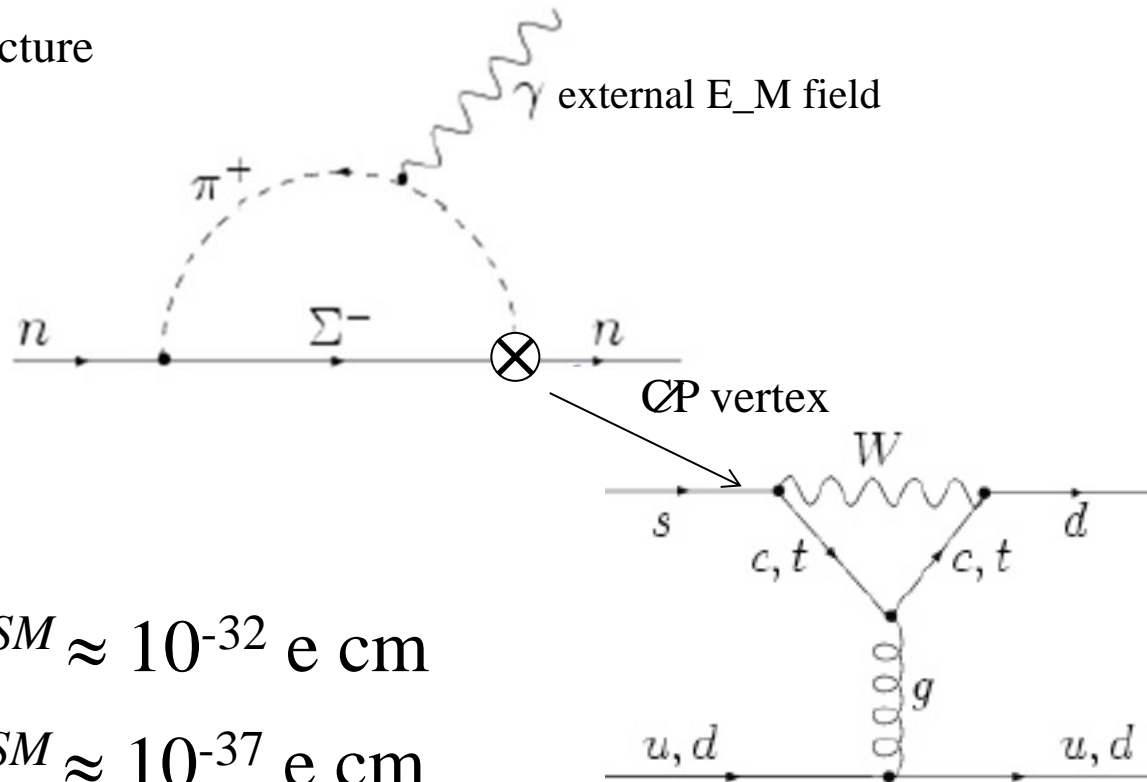
$$H = -\vec{m} \times \vec{B} - \vec{d} \times \vec{E} = -\underbrace{m\vec{J} \times \vec{B}}_{\substack{P_e T_e \\ \perp}} - \underbrace{d\vec{J} \times \vec{E}}_{\substack{P_o T_o \\ \perp}} \quad \not\subset \text{CP}$$

Standard-model EDMs are small

Vanish at 2-loops for quarks and 3-loops for leptons

Khriplovich, Zhitnitsky (1982), McKellar et al., (1987)

Pion-nucleon picture

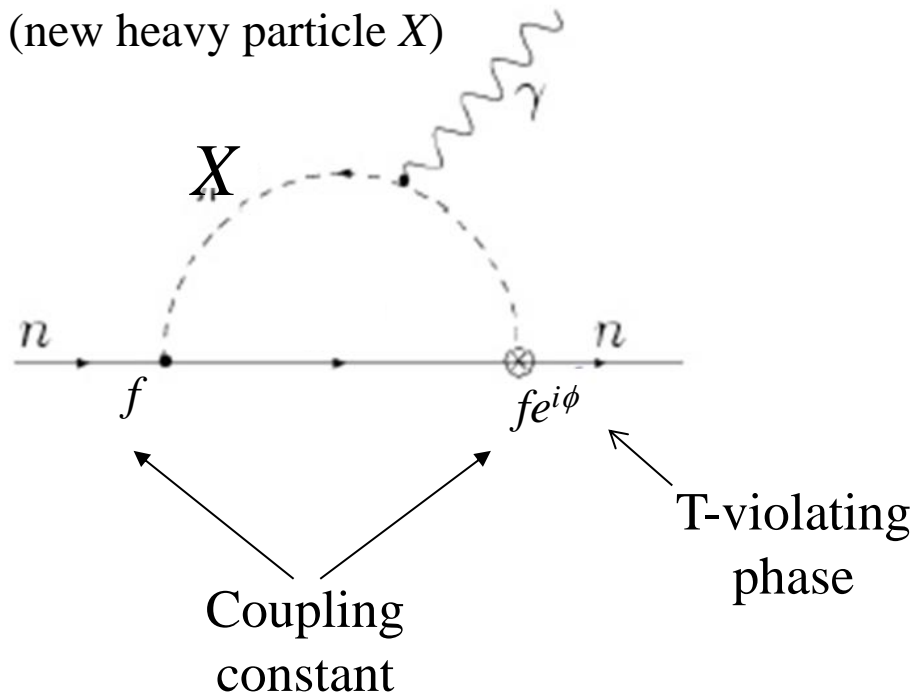


$$d_n^{SM} \approx 10^{-32} \text{ e cm}$$

$$d_e^{SM} \approx 10^{-37} \text{ e cm}$$

$$d_\mu^{SM} \approx 10^{-34} \text{ e cm}$$

EDMs probe TeV-scale “new” physics



$$\mu \approx \frac{e\hbar}{2m} \quad \left(\alpha = \frac{e^2}{\hbar c}\right)$$

$$\frac{d}{m} \gg f^{2N} \frac{m_q}{m_X} \frac{\ddot{\theta}^2}{\theta} \sin f \gg 1$$

$\gg 10^{-14}$ $d_n \sim 10^{-26}$ e-cm

$$m_X \gg m_q \sqrt{10^{14} a^N}$$

loops

$m_x \sim 1$ TeV - LHC scale
or ϕ is small

Baryon Asymmetry

$\cancel{CP} \longrightarrow$ Baryon Asymmetry \longrightarrow NEW PHYSICS (BSMP)

Fact: There is more matter than antimatter

$$n_p \gg n_{\bar{p}} \quad \eta = \frac{n_p - n_{\bar{p}}}{n_p + n_{\bar{p}}} \gg \text{few} \times 10^{-10}$$

(WMAP/PLANCK, $[^4\text{He}]$, ...)

How?

A) Initial condition

B) Evolution from $\eta=0$

1) Baryon number violation

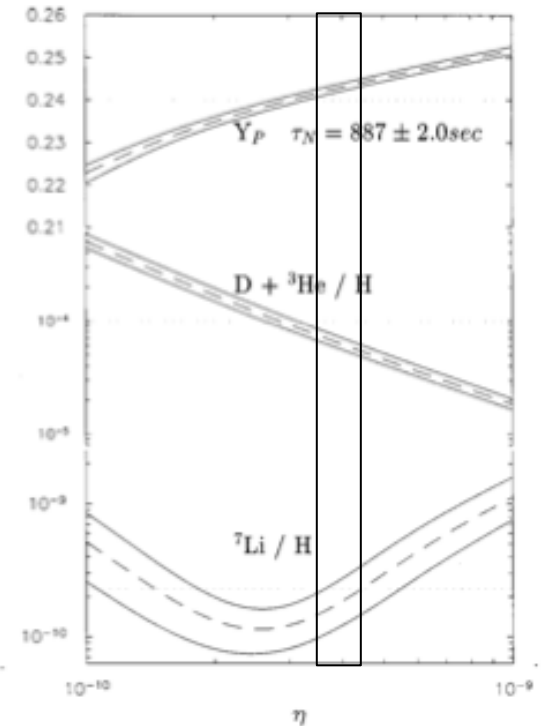
2) CP Violation

3) Rapid expansion (non-equilibrium)



A. Sakharov

Nobel Peace Prize 1975



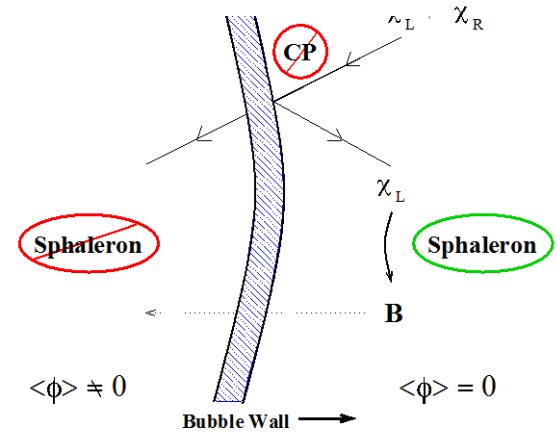
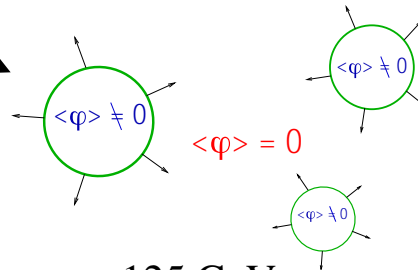
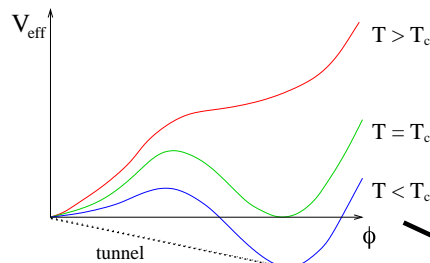
Another possibility: CP violation in neutrinos + “seesaw”

Electroweak Baryogenesis

Kuzmin, Rubakov, Shaposhnikov 87; Cohen, Kaplan, Nelson 90&95

1. First-order EW PT produces expanding bubbles.
2. C and CP violation near the bubble wall induce asymmetries.
3. Electroweak physics (sphalerons) convert this to a baryons

$$V(H, T) = -\frac{1}{2}(\mu^2 - \frac{1}{2}T^2)H^2 - \gamma TH^3 + \frac{\lambda}{4}H^4$$



DOESN'T WORK:

1. The EW PT is not first order for $m_h = 125$ GeV.
Kajantie, Laine, Rummukainen, Shaposhnikov 98

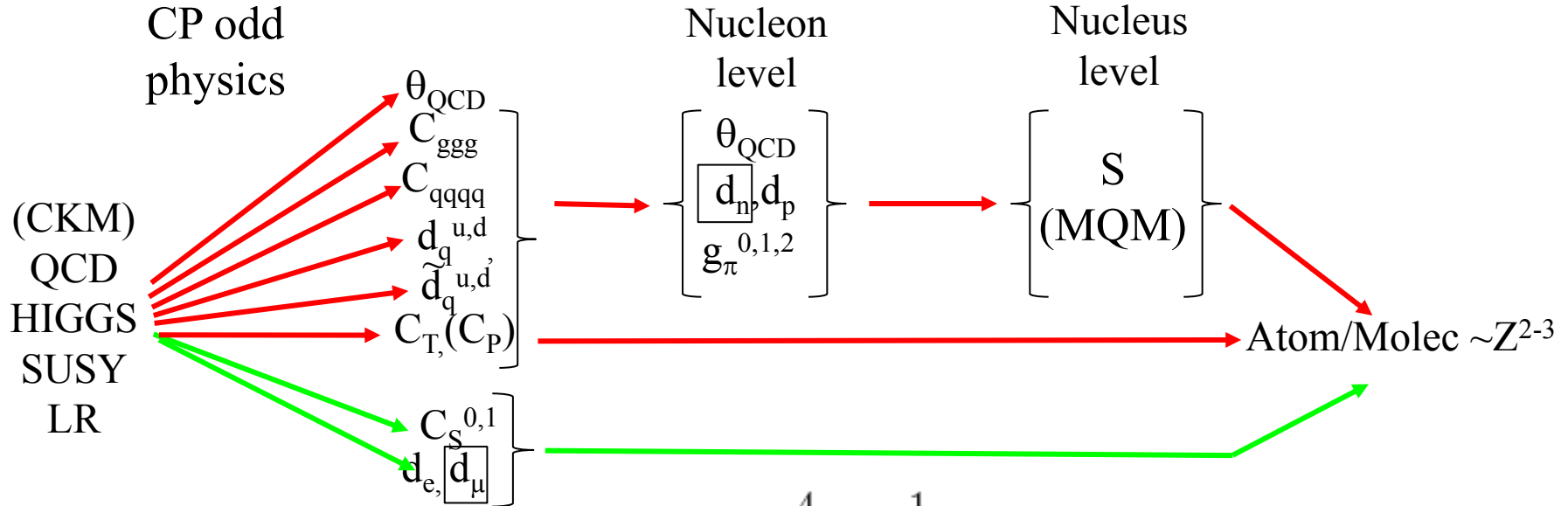
2. Not enough effective CP violation.

Gavela, Hernandez, Orloff, Pene'94; Huet + Sather '95

From D. Morrissey

EDM's

TC, MJ Ramsey Musolf Phys. Rev. C **91** 035502 (2015)
 Upcoming Review: TC, Fierlinger, Ramsey-Musolf, Singh



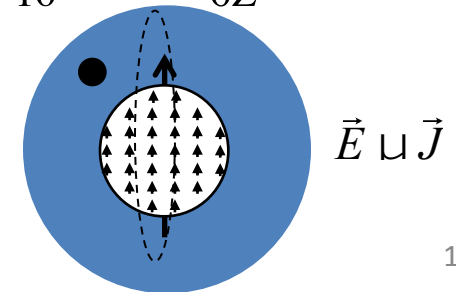
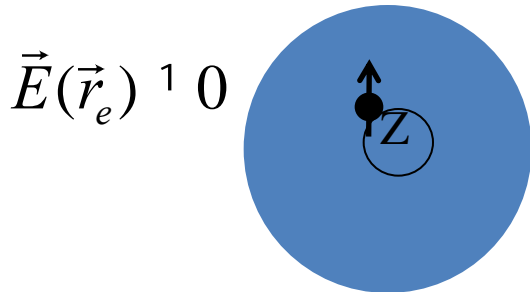
$$d_n = 3.6 \times 10^{-16} \theta_{QCD} e\text{-cm} + \left(\frac{4}{3}d_d - \frac{1}{3}d_u\right) + (1.1\bar{d}_d + 0.55\bar{d}_u)$$

$$d_A = \kappa_S S(\theta_{QCD}, g_\pi) + (k_T C_T + k_S C_S) + \eta_e d_e + \text{h.o. (MQM)}$$

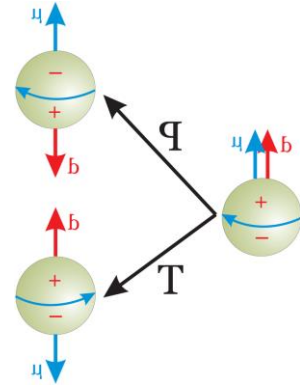
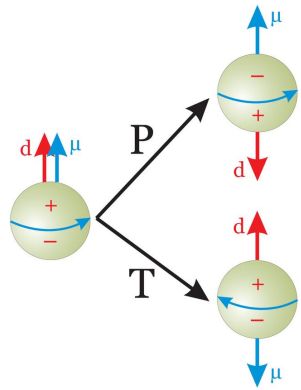
Paramagnetic atoms ($\vec{L} \times \vec{S}$ coupling)
 Cs, Tl, YbF, ThO

Diamagnetic atoms: Schiff moment

Xe, Hg, TlF
$$\vec{S} = \frac{1}{10} \langle r^2 \vec{r}_p \rangle - \frac{1}{6Z} \langle r^2 \rangle \langle \vec{r}_p \rangle$$



Pioneers - experiment



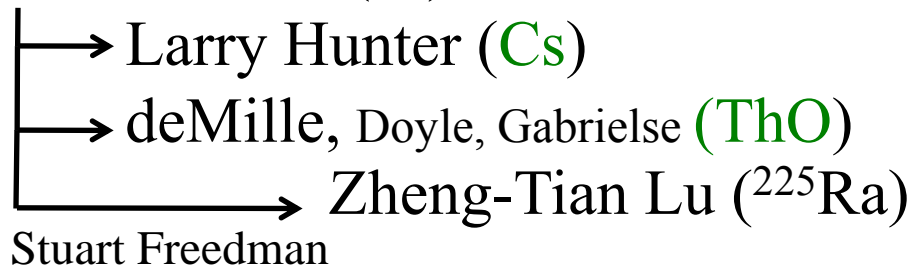
Normal Ramsey & Ed Purcel (neutron)



Pat (P.G.H.) Sandars (Cs , Xe^m , TIF)



Gene Commins (Tl)



System	Year/ref	Result
Paramagnetic systems		
Cs	1989 [33]	$d_A = (-1.8 \pm 6.9) \times 10^{-24}$ e-cm $d_e = (-1.5 \pm 5.6) \times 10^{-26}$ e-cm
Tl	2002 [9]	$d_A = (-4.0 \pm 4.3) \times 10^{-25}$ e-cm $d_e = (-6.9 \pm 7.4) \times 10^{-28}$ e-cm
YbF	2011 [8]	$d_e = (-2.4 \pm 5.9) \times 10^{-28}$ e-cm
ThO	2014 [7]	$\omega^{NE} = 2.6 \pm 5.8$ mrad/s $d_e = (-2.1 \pm 4.5) \times 10^{-29}$ e-cm $C_S = (-1.3 \pm 3.0) \times 10^{-9}$
Diamagnetic systems		
^{199}Hg	2006 [5]	$d_A = (0.49 \pm 1.5) \times 10^{-29}$ e-cm
^{129}Xe	2001 [34]	$d_A = (0.7 \pm 3) \times 10^{-27}$ e-cm
TlF	2000 [35]	$d = (-1.7 \pm 2.9) \times 10^{-23}$ e-cm
neutron	2006 [4]	$d_n = (0.2 \pm 1.7) \times 10^{-26}$ e-cm

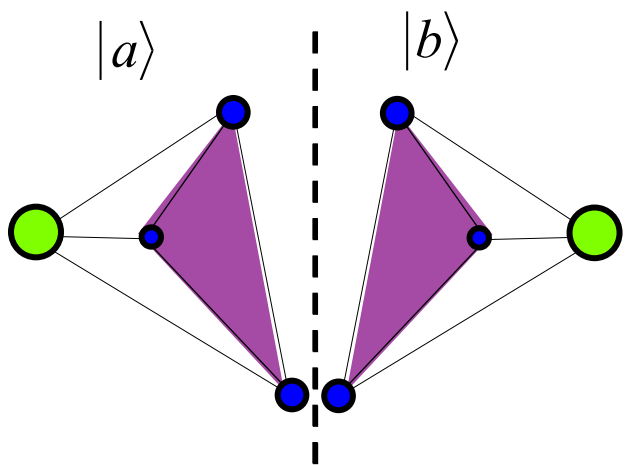
Octupole Enhancements

Intrinsic (body-frame) moment

Polarizability

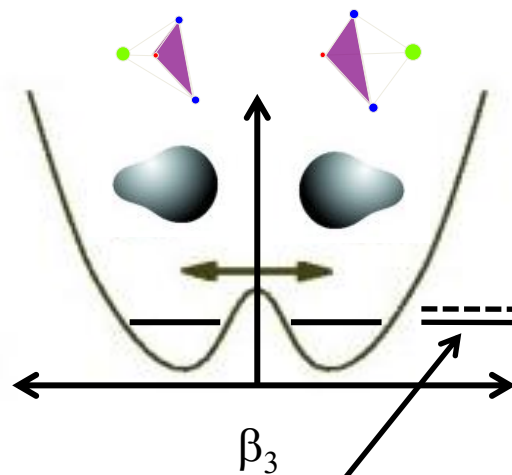
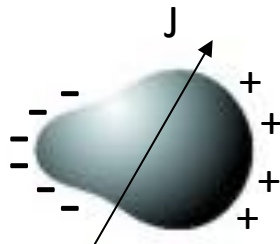
NH₃ (see Feynman vol 3.)

Reflection Symmetry



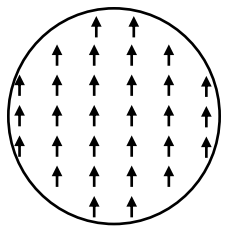
$$|y_+\rangle = \frac{1}{\sqrt{2}} (|a\rangle + |b\rangle)$$

$$|y_-\rangle = \frac{1}{\sqrt{2}} (|a\rangle - |b\rangle)$$



Small splitting (tunnel frequency)
Large electric polarizability

$$\vec{S} = \frac{1}{10} \langle r^2 \vec{r}_p \rangle - \frac{1}{6} Z \langle r^2 \rangle \langle \vec{r}_p \rangle$$



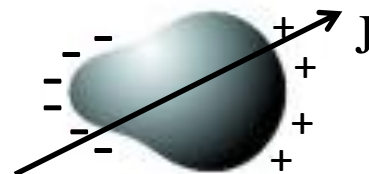
$$\vec{E} \propto \vec{J}$$

$$S \mu \frac{\langle + | hr^2 \cos q | - \rangle}{E_+ - E_-} \gg \frac{hb_2 b_3^2 A^{2/3} r_0^3}{E_+ - E_-}$$

Nuclei with Octupole Deformation/Vibration

(Haxton & Henley; Auerbach, Flambaum, Spevak; Engel et al., Hayes & Friar, etc.)

$$S \mu \frac{\langle + | \hbar r^2 \cos q | - \rangle}{E_+ - E_-} \gg \frac{\hbar b_2 b_3^2 A^{2/3} r_0^3}{E_+ - E_-}$$



	^{223}Rn	^{223}Ra	^{225}Ra	^{223}Fr	^{129}Xe	^{199}Hg
$t_{1/2}$	23.2 m	11.4 d	14.9 d	22 m		
I	7/2	3/2	1/2	3/2	1/2	1/2
ΔE th (keV)	37*	170	47	75		
ΔE exp (keV)	-	50.2	55.2	160.5		
$10^{11} S$ (e-fm ³)	375	150	115	185	0.6	-0.75
$10^{28} d_A$ (e-cm)	1250	1250	940	1050	0.3	2.1

$$\eta_{qq} = 3.75 \times 10^{-4}$$

Ref: Dzuba PRA66, 012111 (2002) - Uncertainties of 50%

*Based on Woods-Saxon Potential

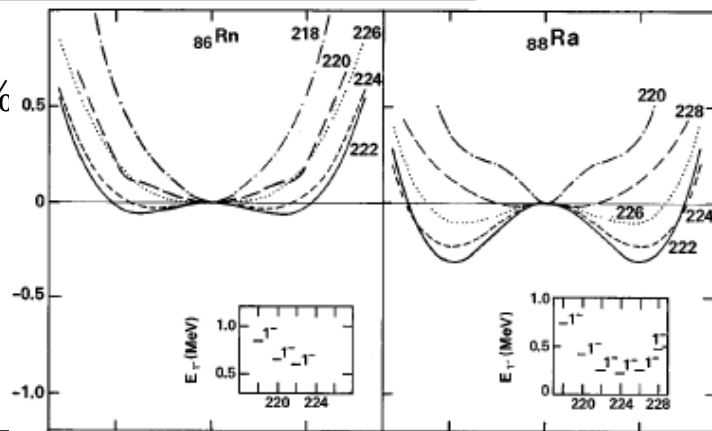
† Nilsson Potential Prediction is 137 keV

NOTES:

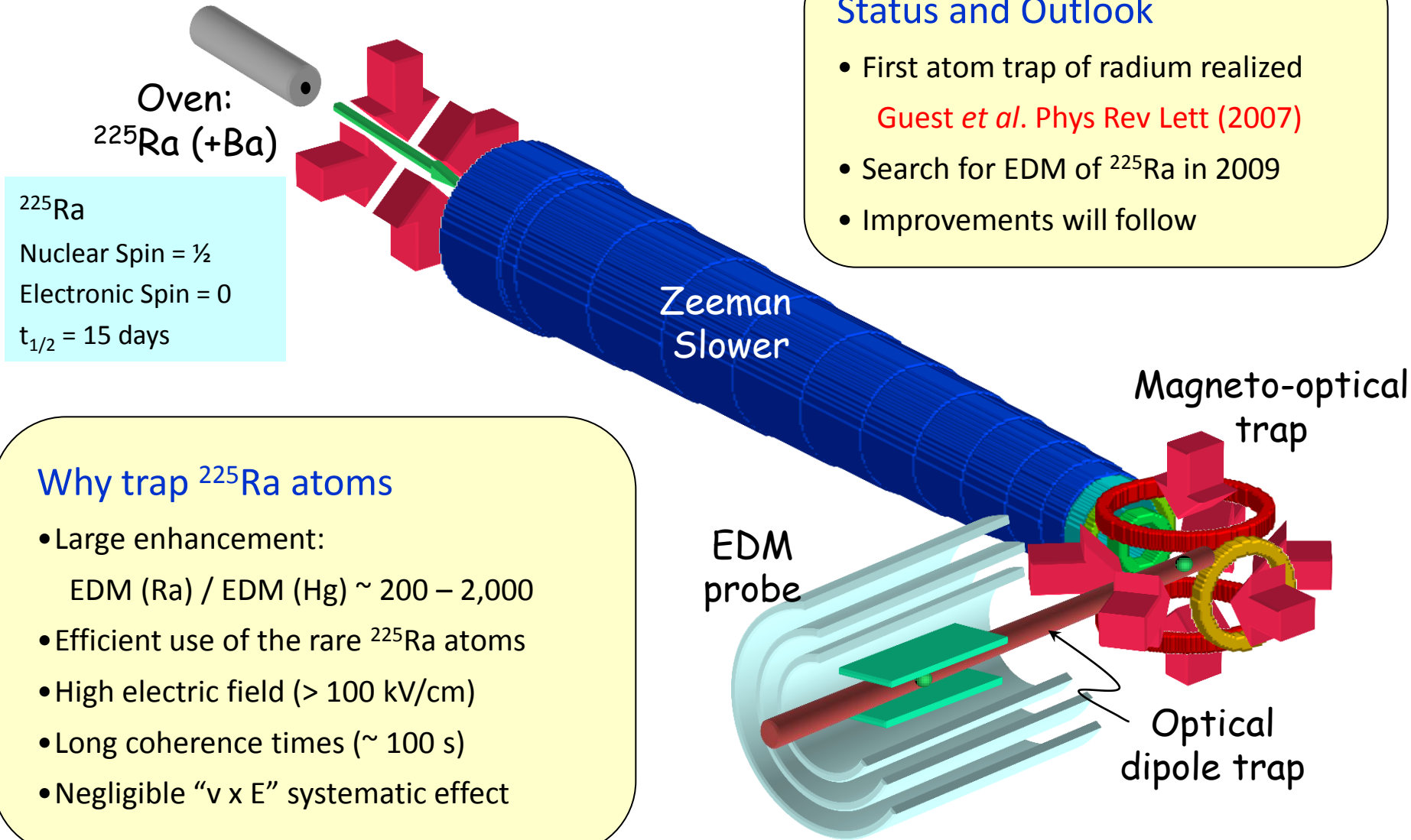
Octupole Enhancements

Engel et al. agree with Flambaum et al.

Even octupole vibrations enhance S (Engel, Flambaum & Zelevinsky)



Search for EDM of ^{225}Ra at Argonne



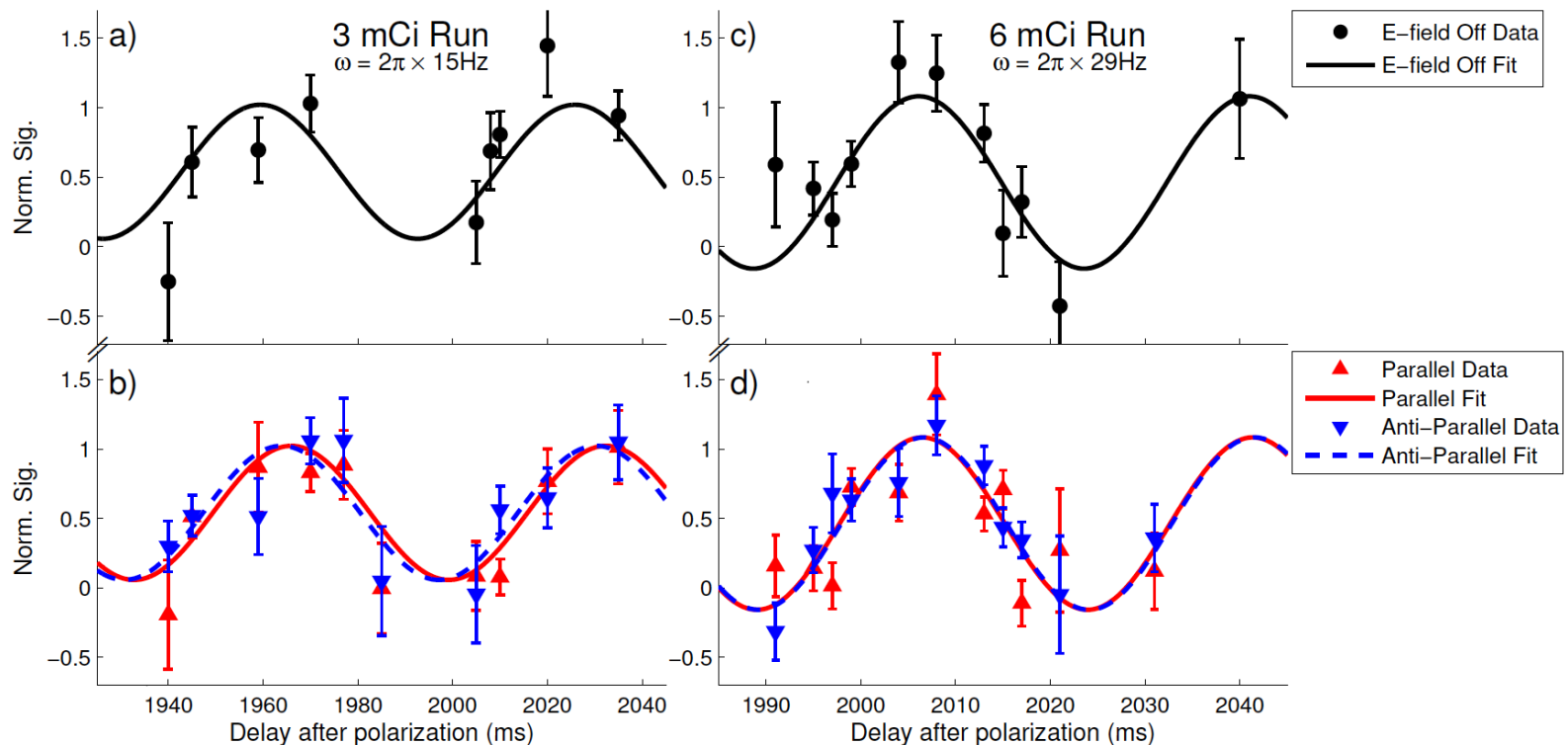
Status and Outlook

- First atom trap of radium realized
Guest et al. Phys Rev Lett (2007)
- Search for EDM of ^{225}Ra in 2009
- Improvements will follow

Why trap ^{225}Ra atoms

- Large enhancement:
EDM (Ra) / EDM (Hg) $\sim 200 - 2,000$
- Efficient use of the rare ^{225}Ra atoms
- High electric field (> 100 kV/cm)
- Long coherence times (~ 100 s)
- Negligible " $\mathbf{v} \times \mathbf{E}$ " systematic effect

First Ra-225 EDM Measurement



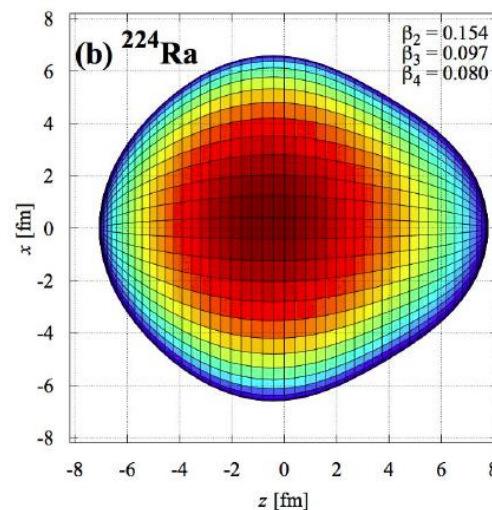
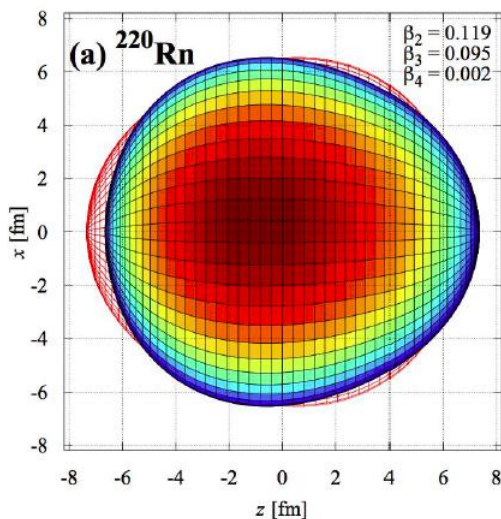
Phys. Rev. Lett. 114, 233002: $|d(\text{Ra-225})| < 5 \times 10^{-22} e \text{ cm}$ (95%)

- all systematic effects estimated to be $< 10^{-25} e \text{ cm}$ (goal)
- first EDM measurement made in a laser trap
- first EDM measurement of an octupole-deformed species

^{221}Rn Enhancement



L. Gaffney
P. Butler
M. Scheck
et al.

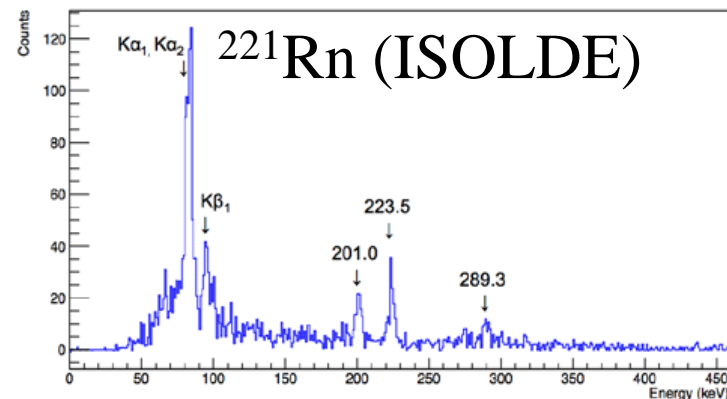


$$S_{\mu} \propto \frac{hb_2 b_3^2 A^{2/3} r_0^3}{E_+ - E_-}$$

$$\frac{S_{Rn}}{S_{Hg}} = \frac{S_{Ra}}{S_{Hg}} \frac{S_{Rn}}{S_{Ra}} \approx 1000 \frac{\beta_2}{\beta_2} \frac{\beta_3^2}{\beta_3^2} \frac{\Delta E_{Ra}}{\Delta E_{Rn}} \approx 50 - 100$$

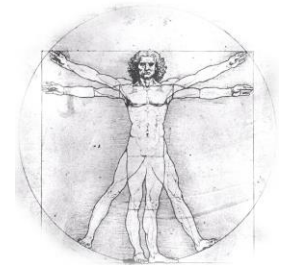
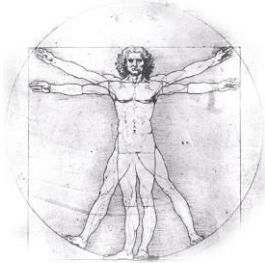
500-1000
(J. Engel et al.)

50 keV
400 keV



^{223}Rn : TBD

Radon-EDM Experiment



TRIUMF E929

Spokesmen: Timothy Chupp & Carl Svensson

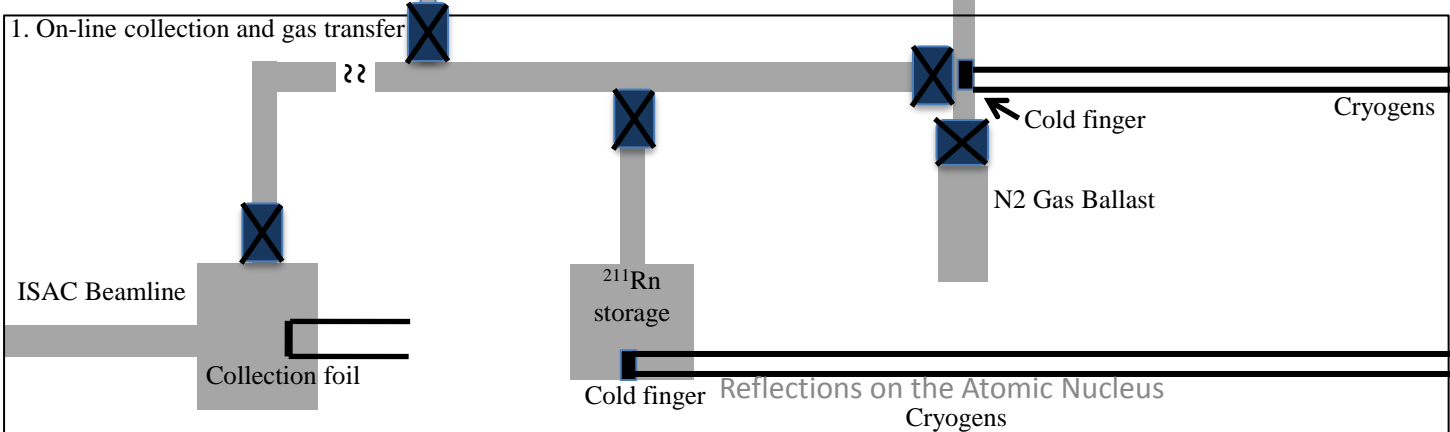
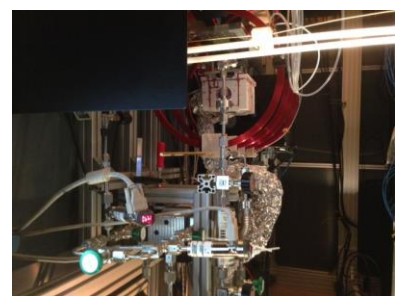
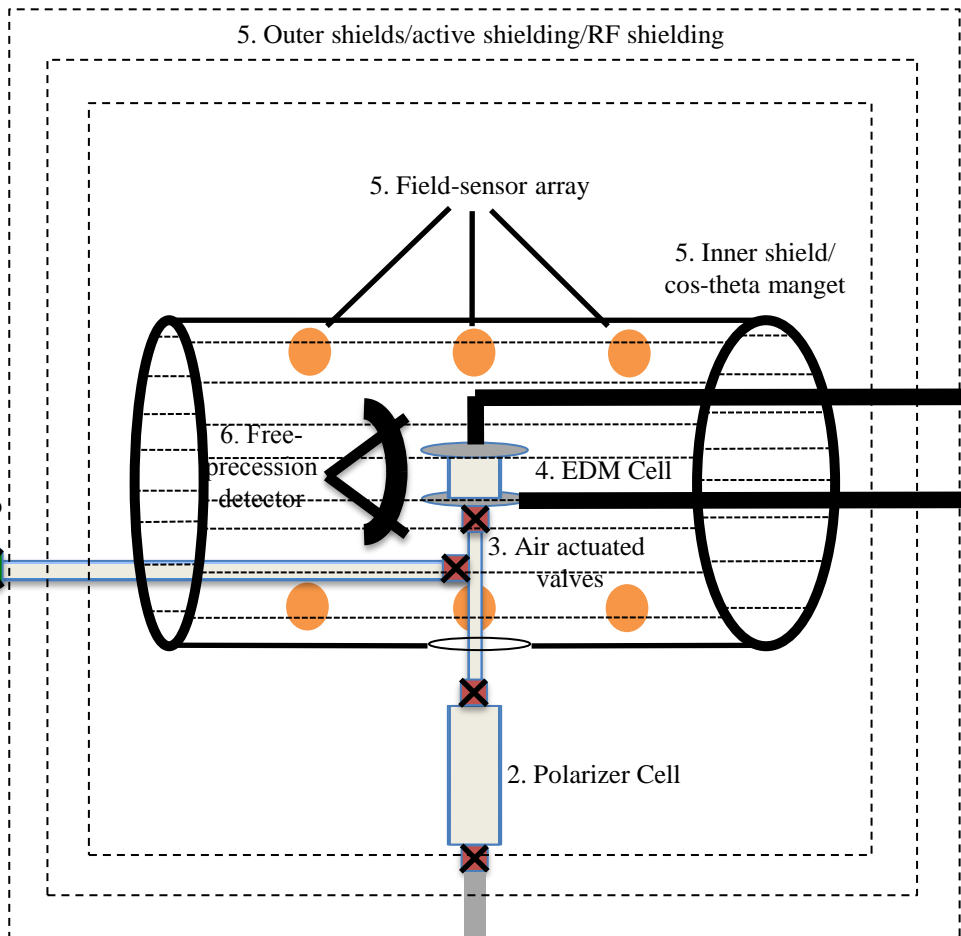
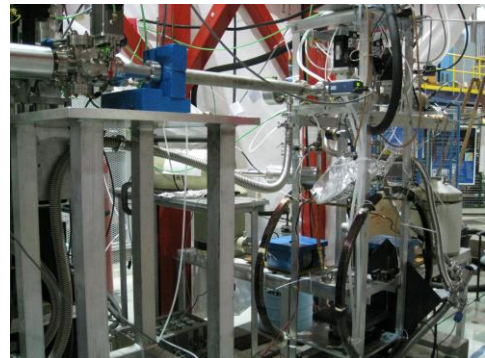


E-929 Collaboration (Guelph, Michigan, SFU, TRIUMF)
TRIUMF

Canada's National Laboratory for Particle and Nuclear Physics

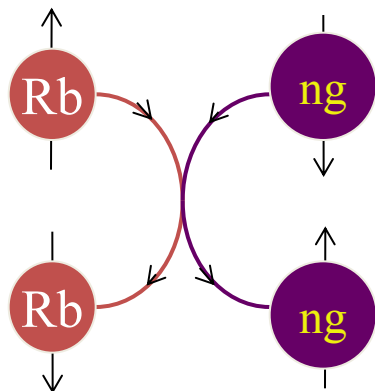
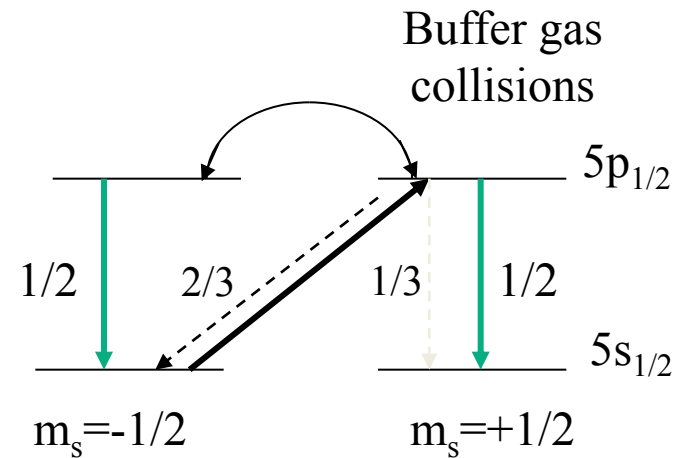
Funding: NSF-Focus Center, DOE, NRC (TRIUMF), NSERC

20-30 min half-life

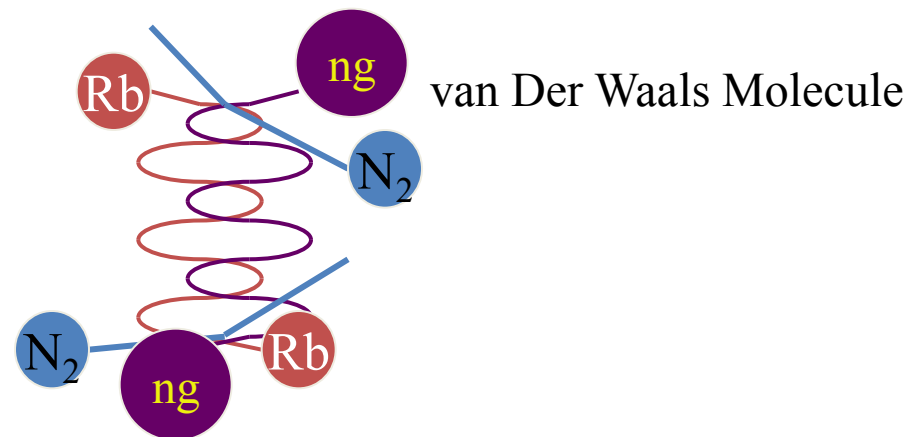


Spin-Exchange Optical Pumping

- Optically pump the Rb with circularly polarized laser light.
- Spin-exchange collisions transfer the polarization to the ^3He , ^{129}Xe , radon nuclei.



Binary Collision:
 $\tau \sim 10^{-12}$ sec.



Nuclear Orientation of Radon Isotopes by Spin-Exchange Optical Pumping

M. Kitano,^(a) F. P. Calaprice, M. L. Pitt, J. Clayhold, W. Happer, M. Kadar-Kallen, and M. Musolf

E_γ (keV)	Spin sequence	Anisotropy R	$R - 1$ (%)
337	$(\frac{1}{2}^-) - (\frac{5}{2}^-)$	0.903(14)	-9.7 ± 1.4
408	$(\frac{3}{2}^-) - \frac{9}{2}^-$	1.009(7)	$+0.9 \pm 0.7$
689	$\frac{5}{2}, \frac{7}{2}^- - \frac{5}{2}^-$	1.079(22)	$+7.9 \pm 2.2$
745	$(\frac{1}{2}^-) - \frac{9}{2}^-$	1.129(14)	$+12.9 \pm 1.4$

Polarization and relaxation of radon

E. R. Tardiff,¹ J. A. Behr,³ T. E. Chupp,¹ K. Gulyuz,⁴ R. S. Lefferts,⁴ W. Lorenzon,² S. R. Nuss-Warren,¹ M. R. Pearson,³ N. Pietralla,⁴ G. Rainovski,⁴ J. F. Sell,⁴ and G. D. Sprouse⁴

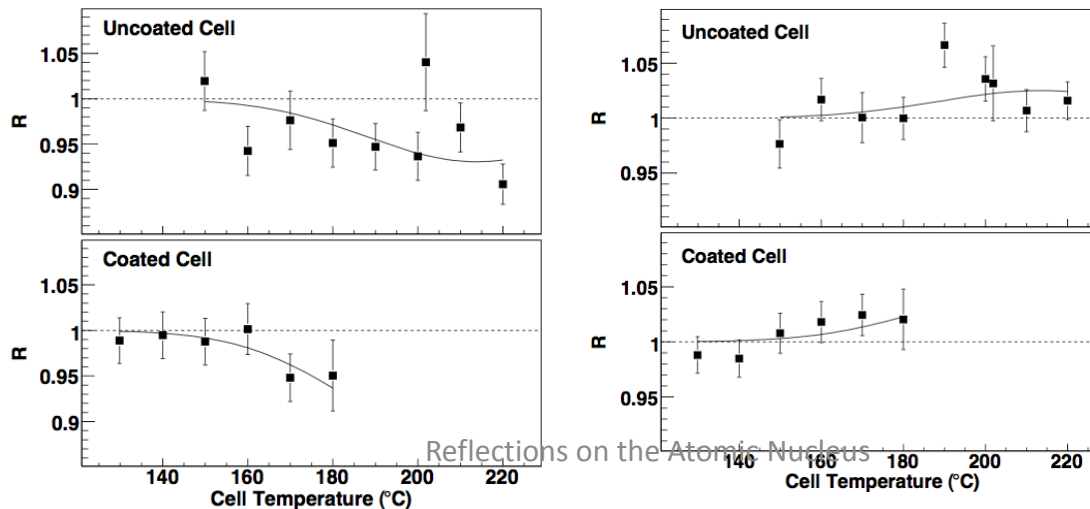
¹FOCUS Center, University of Michigan Physics Department, 450 Church St., Ann Arbor 48109-1040, USA

²University of Michigan Physics Department, 450 Church St., Ann Arbor 48109-1040, USA

³TRIUMF, 4004 Westbrook Mall, Vancouver V6T 2A3, Canada

⁴SUNY Stony Brook Department of Physics and Astronomy, Stony Brook 11794-3800, USA

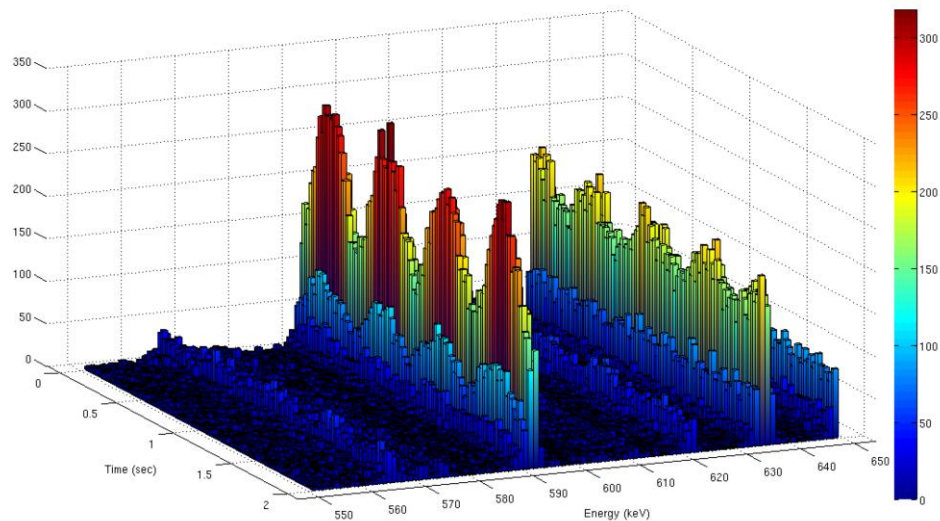
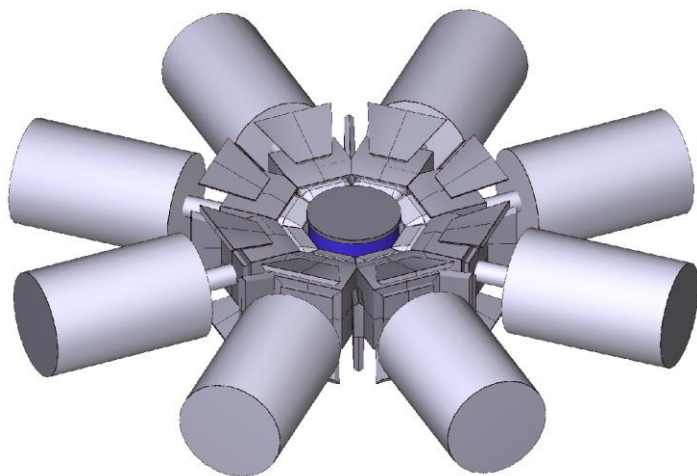
(Dated: December 6, 2006)



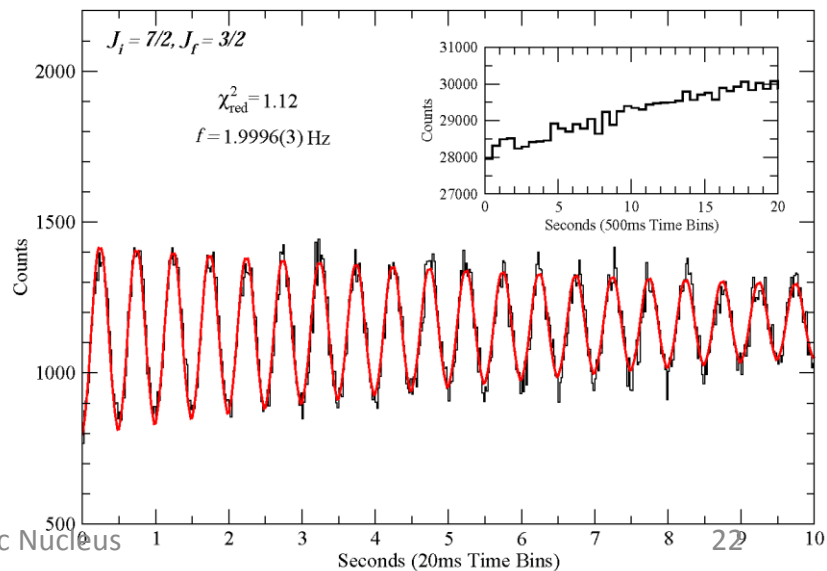
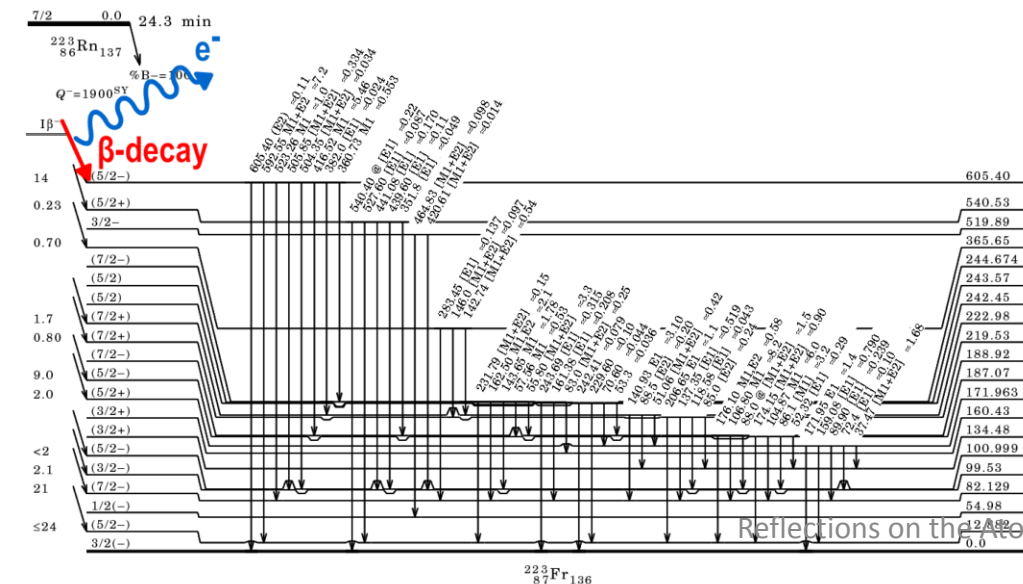
Fit for Γ_2 ($T_a=300^\circ\text{K}$):
 0.05 Hz (uncoated);
 0.03 Hz (coated)s
 Use $2.5 \times 10^{-21} \text{ cm}^2$

Genat-4 simulations by Evan Rand

γ -ray energy-time matrix from the β decay of 1.2 billion ^{223}Rn nuclei from an initial 8×10^{10} nuclei located in the EDM cell surrounded by a ring of eight GRIFFIN detectors in the forward position.

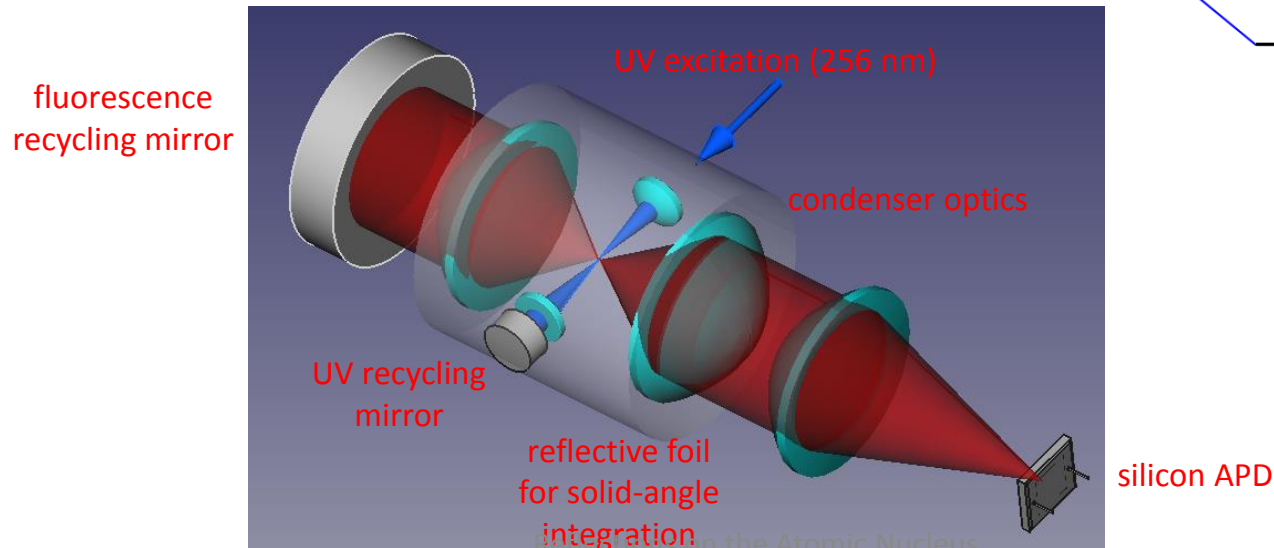
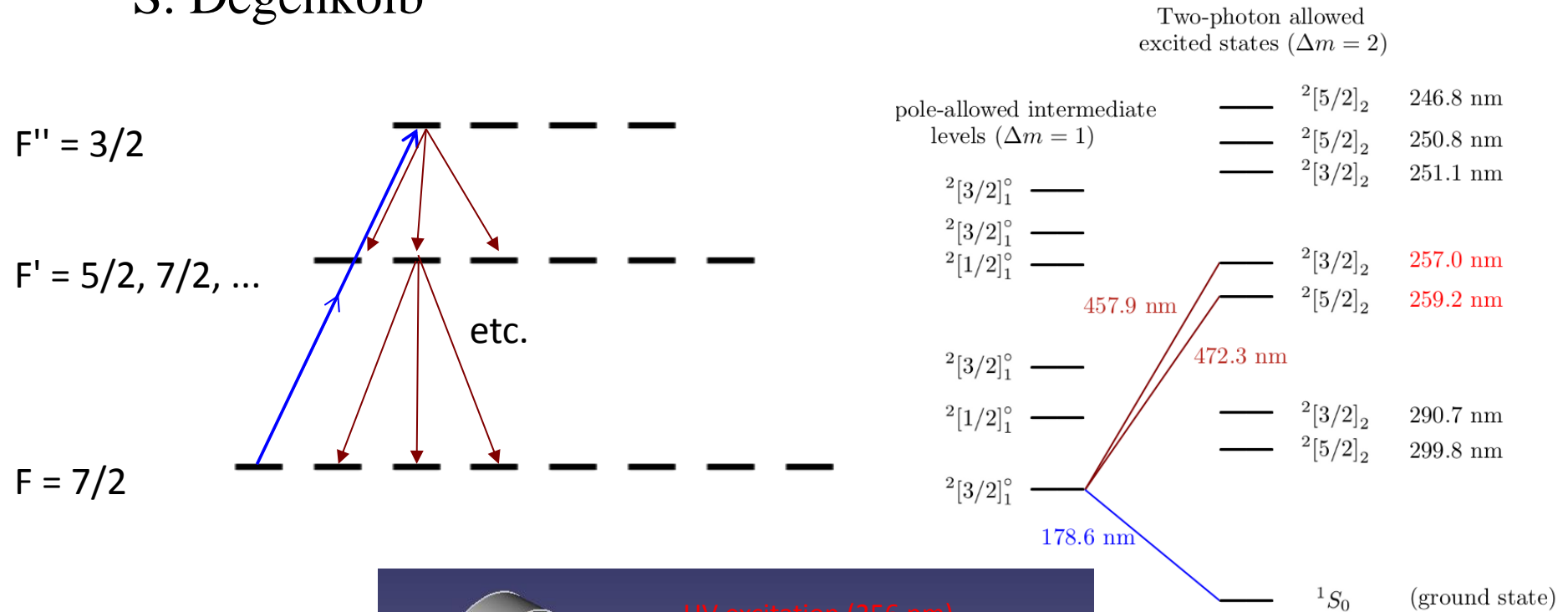


Known Level Structure of ^{223}Fr - Nuclear Data Sheets (2001)



Two-photon magnetometry with $^{221/223}\text{Rn}$ ($J=7/2$)

S. Degenkolb



Radon-EDM Prospects

Global analysis: TC, Ramsey Musolf PRC **91** 035502 (2015)

Goal $\sim 10^{-26}$ e-cm

Facility	TRIUMF-ISAC	FRIB(^{223}Th)
Rate	$2.5 \times 10^7 \text{ s}^{-1}$	$1 \times 10^9 \text{ s}^{-1}$
# atoms	3.5×10^{10}	1.4×10^{12}
σ_{EDM} (100 d)	$2 \times 10^{-27} \text{ e-cm}$	$3 \times 10^{-28} \text{ e-cm}$
^{199}Hg equivalent	$4 \times 10^{-28/29} \text{ e-cm}$	$6 \times 10^{-29/30} \text{ e-cm}$

Assumptions: $E=10 \text{ kV/cm}$, $T_2=15 \text{ s}$, $A=0.2$, 25% duty factor

$$\sigma_d \approx \frac{1}{2E} \frac{\hbar}{AT_2} \frac{1}{\sqrt{N_\gamma}}$$

Summary

EDMs probe TeV-scale “new” physics

\cancel{CP} \longrightarrow Baryon Asymmetry \longrightarrow NEW PHYSICS (BSMP)

Measurements in NEW SYSTEMS are essential

Octupole collectivity enhances Schiff moments:

^{225}Ra and $^{221/223}\text{Rn}$ underway $10^{-25/26}$ e-cm

THANK YOU!

and

Happy *transition(s)* PETER!