




Leda and the Swan, copy by [Cesare da Sesto](#) after a lost original by [Leonardo](#), 1515–1520, 

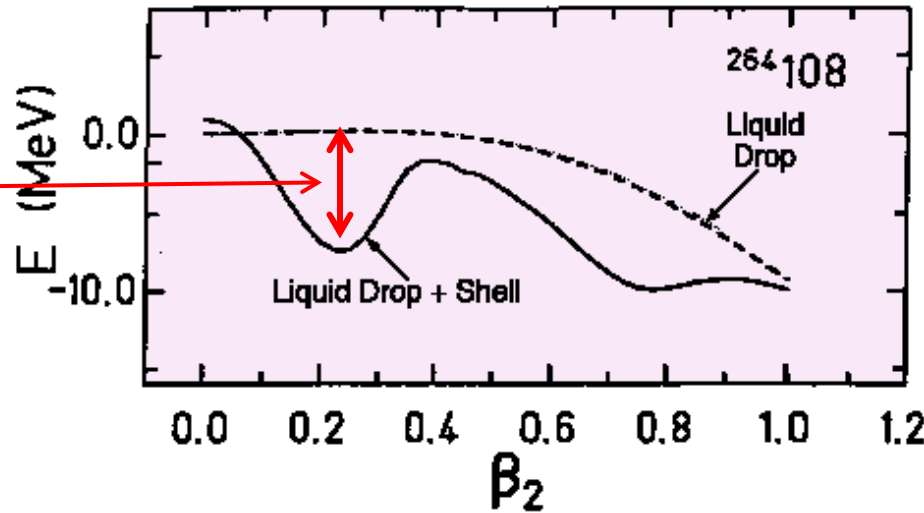


Impact of superheavy element research on theory

Teng Lek Khoo
Argonne National Laboratory

- Nuclei at three simultaneous limits: Z , I , E^*
- Models anchored on nuclei where data available for parameter fits.
- Calculations of superheavy elements (SHE) -- & superdeformed nuclei -- represent true predictions, rather than descriptions.

Why SHN do exist?



$$E = E(\text{LD}) + E(\text{shell}) + E(\text{pair}) + E(\text{deformation})$$

$$T_{\text{sf}}(\text{exp})/T_{\text{sf}}(\text{LD}) > 10^{13}$$

SHN exist only because E_{shell} provides extra binding, creating a

fission barrier, $B_f = E_{\text{saddle}} - E_{\text{gs}}$.

E_{shell} arises from gaps in the **single-particle spectrum**.

All properties can, in principle, be calculated once the spectrum is established.

E_{shell} , B_f , **single-particle spectrum**: each & every one necessary for existence of SHN, **essential** aspects of the field.

Shell energy E_{shell} at high spin I

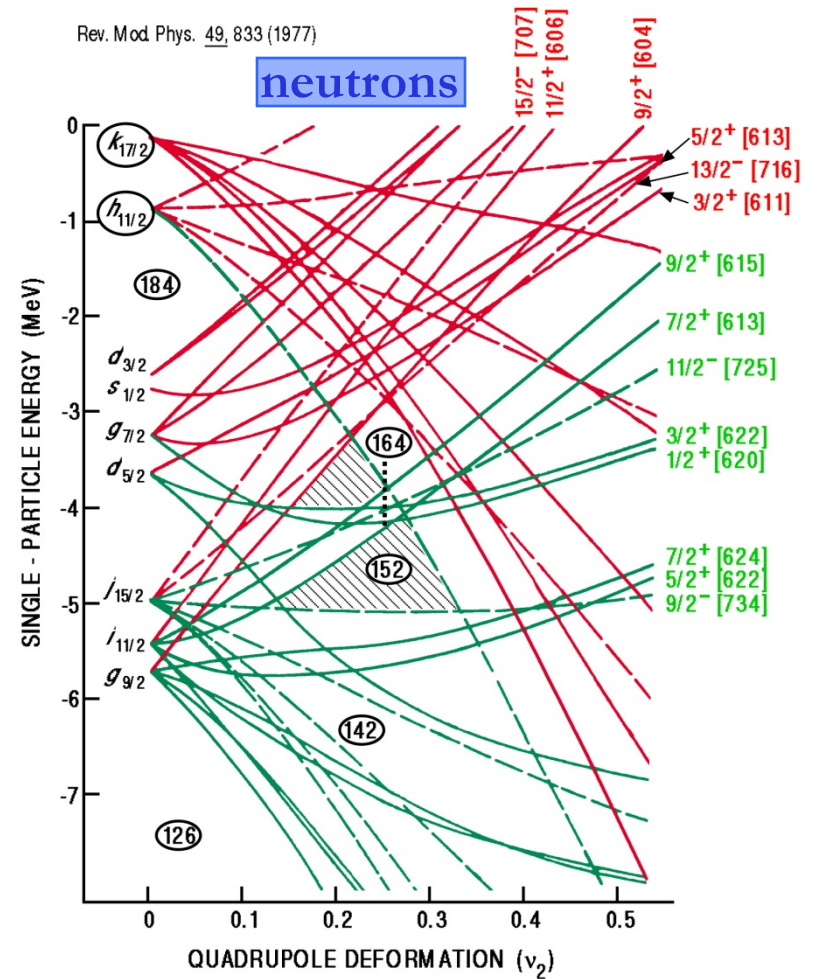
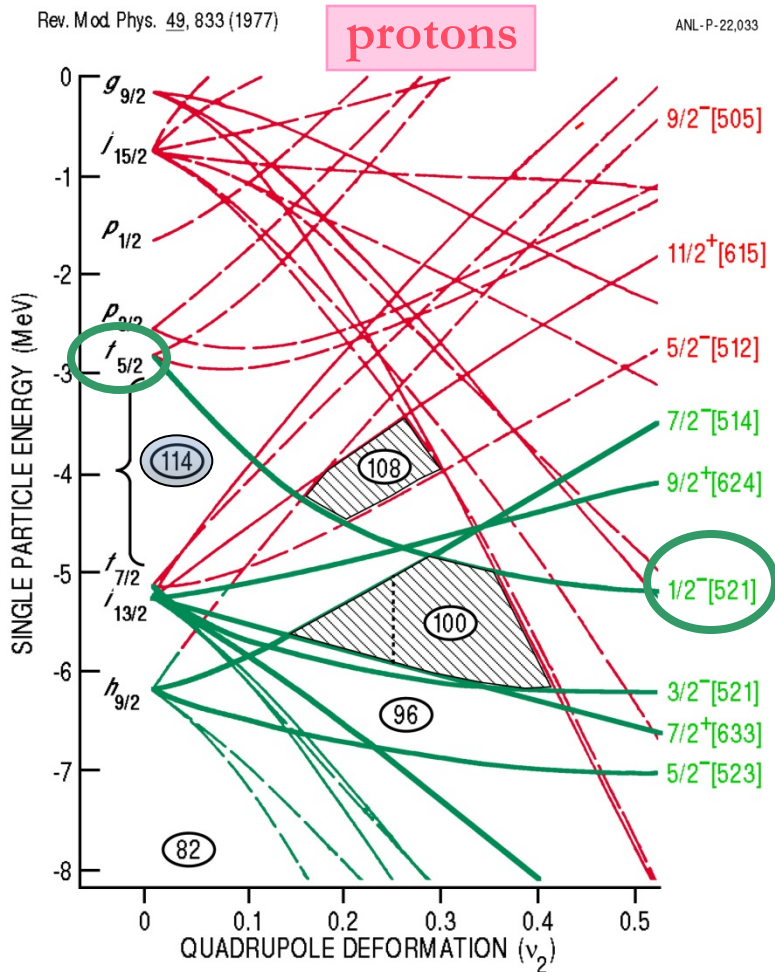
- Search for ever heavier superheavy elements: an exciting (& glamorous) frontier in nuclear physics.
- The underlying physics: *how E_{shell} persists at large Z , i.e. $E_{shell}(Z)$. Magic numbers of SHN?*
- $E_{shell}(I,Z,N) \rightarrow$ more insight.

Effect of **deformation & rotation** on orbital energy

- **Deformation & rotation** lower sub-states of high-lying j shells towards Fermi level.
- Shells become **accessible** which otherwise would not be (if nucleus were spherical).

Proton and neutron single-particle energies (Woods-Saxon potential)

ANL-P-22,262



R. Chasman et al., Rev. Mod. Phys. 49, 833 (1977)

T. L. Khoo, Reflections 2015

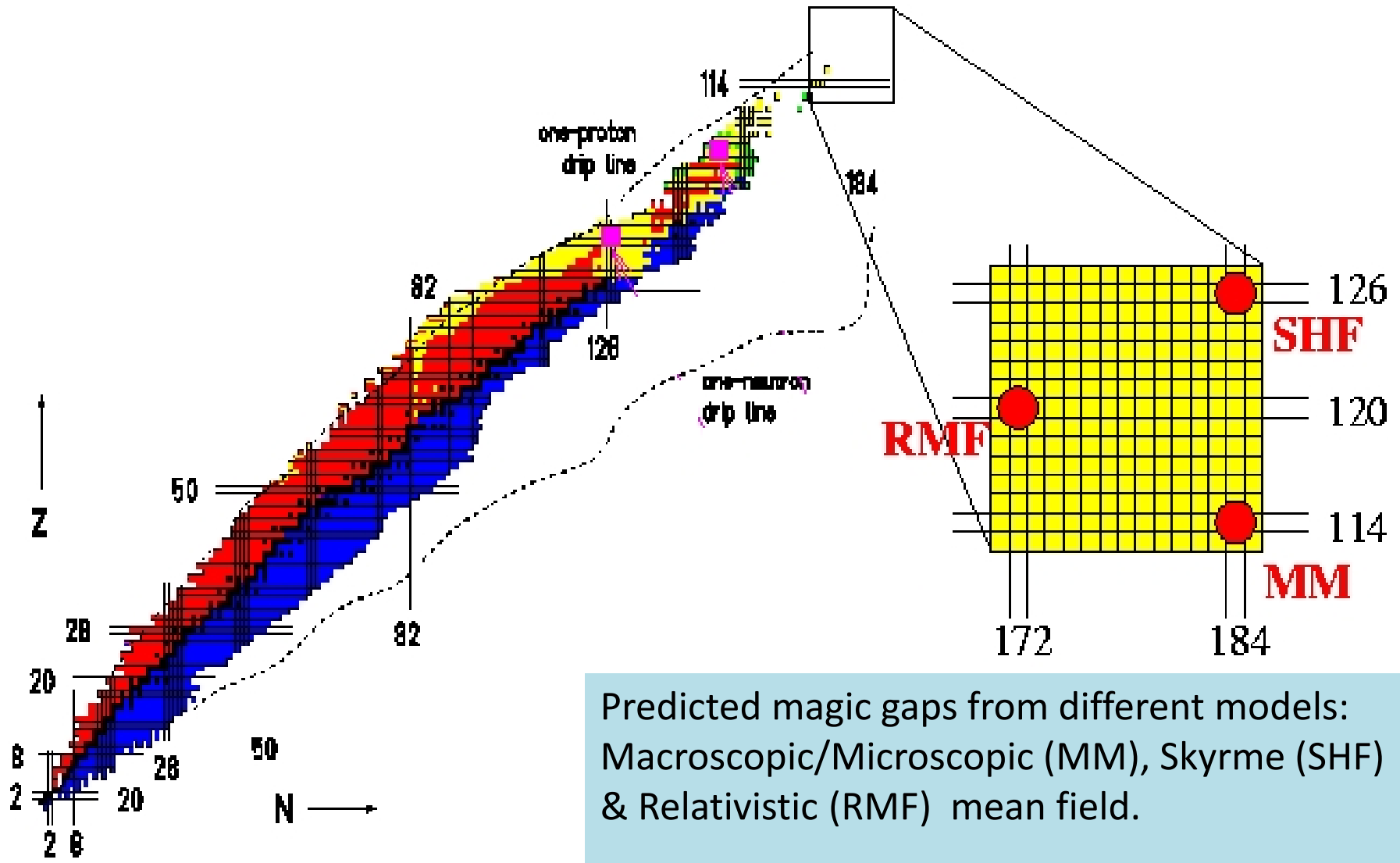
Spectroscopy, mass and reaction data

- Combine **spectroscopic, mass and reaction data** to develop a more complete, consistent description.
- Especially when data is sparse, as for SHN.
- **SPE** from spectroscopy, mass.
- Entry distribution \leftrightarrow Fission barrier B_f , formation mechanism.
- $E_{\text{shell}}(I)$ from mass & yrast line.

1. Single-particle energies (SPE) of SHE:

- Shell energy originates from single-particle spectra, particularly from gaps.
- **SPE represent the core from which all follows: existence, mass, shell energy, fission barrier.**
- Accurate predictions require accurate SPE.
- Different models have discrepant SPE.
- Can predictions be accurate without accurate SPE?

Where are the magic gaps for superheavy nuclei?



Predicted magic gaps from different models:
 Macroscopic/Microscopic (MM), Skyrme (SHF)
 & Relativistic (RMF) mean field.

SPE must be different; therefore testable.

Single-particle energies (SPE) of SHN:

decisive test of theory

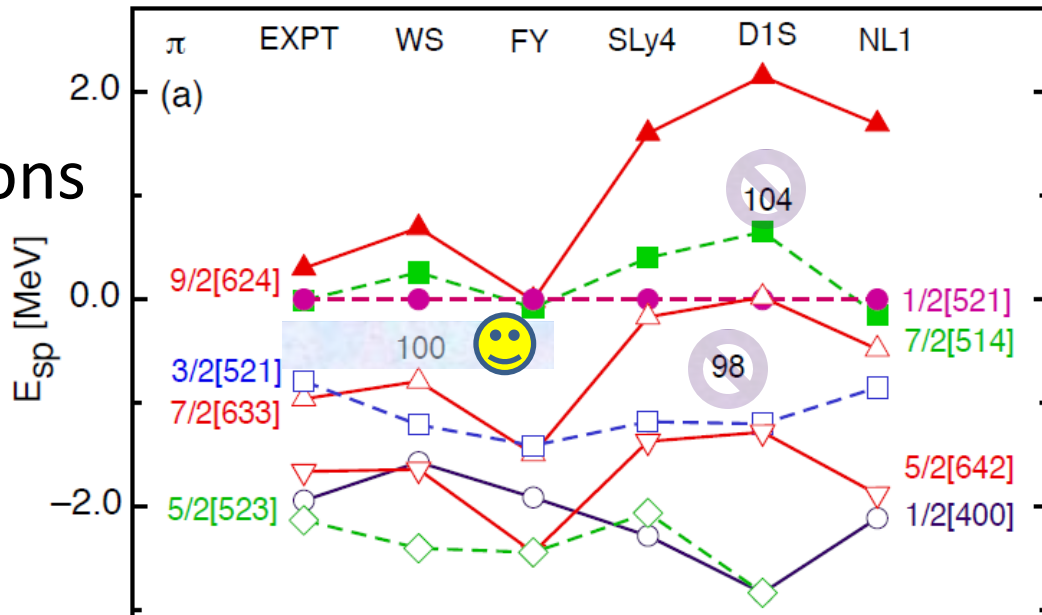
Extract a set of **SPE**, which give **E(1qp)** equal to measured energies in $^{247,249}\text{Bk}$, ^{251}Es & ^{247}Cm , ^{251}Cf .

- Trial set of SPE.
- Calculate E(1qp), using LN procedure for pairing.
- Iterate until E(1qp) reproduce experiment.

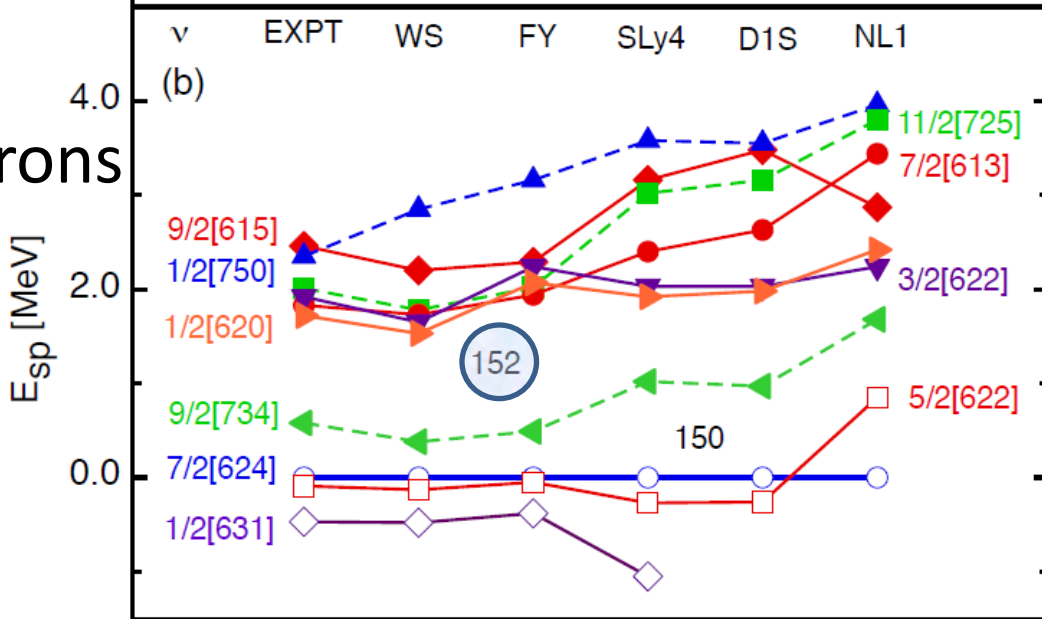
E(2qp) provide complementary tests.

“EXPT” Woods Saxon Density Funct Th

protons



neutrons



“SPE” (deduced from E_{1qp})
Expt. vs. theory

Spectrum directly reflected in E_{1qp} , E_{2qp} & masses.

Degeneracies in SPE \rightarrow
 $\Delta E_{1qp} < 9, 55$ keV.

Gaps at $Z=100, N=152 \rightarrow$
largest E_{shell} at ^{252}Fm ;

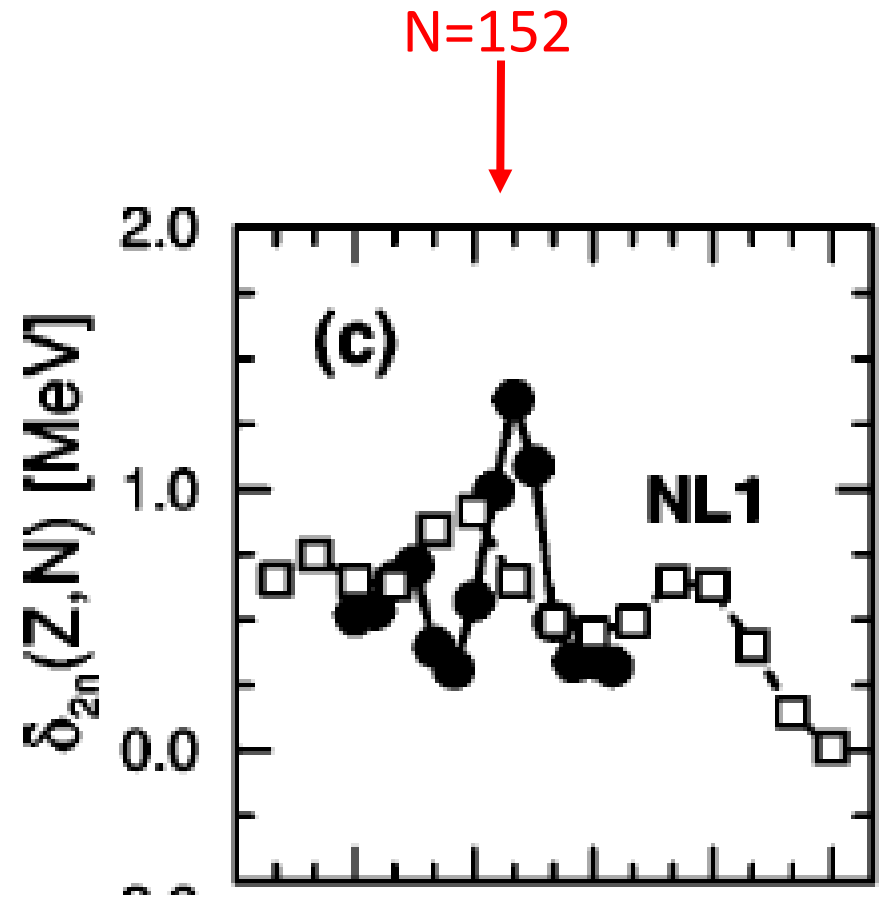
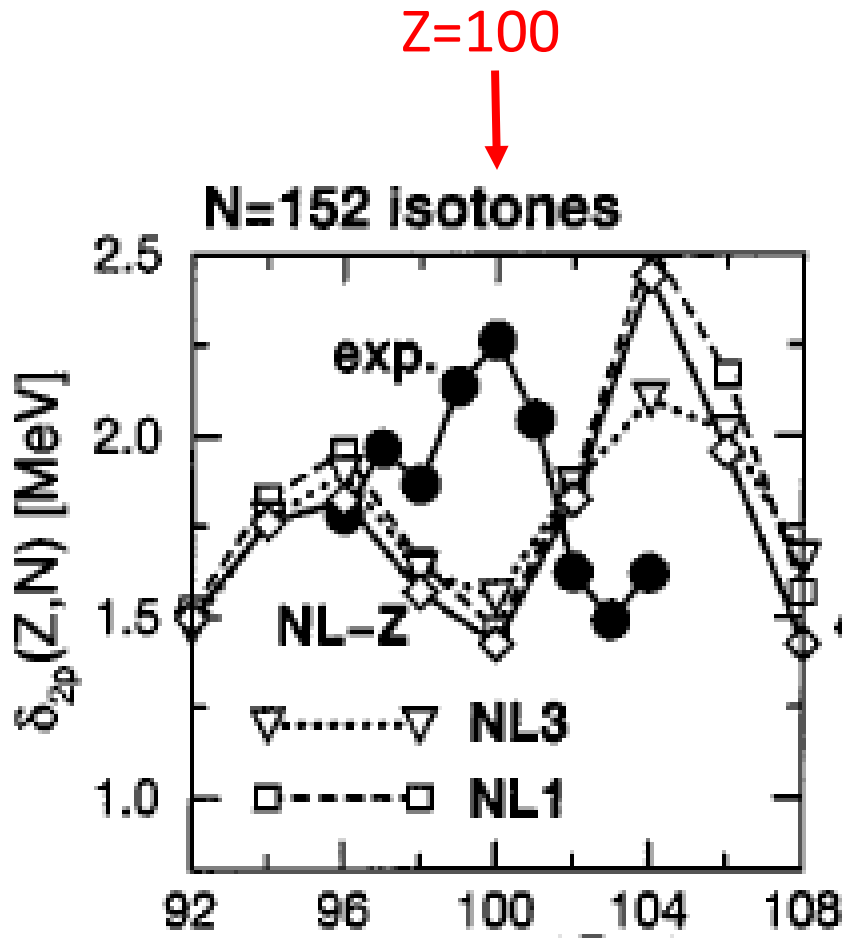
also reflected in $\delta S_{2n}, \delta S_{2p}$.

Errors in $E_{1qp,2qp}$ (DFT) traced to
too high $E(\text{high-j})$; ΔE up to 1.4 MeV!

Spectroscopy, mass

Single-particle spectra directly manifest in binding energy, hence mass & δS_{2n} , δS_{2p} .

Also governs B_f .



Afanasvej et al, PR **C67**, 024309 (2003)

Comments on theory

- DFT^{a)} require improved effective interactions.
- Predict gaps at $Z = \underline{120, 126}$, but give wrong deformed gaps.
- Woods-Saxon (WS) potential gives reasonable (best) spe. Gap at $Z=114$.
- High-K 2-qp isomers establish configurations & support above conclusions.

2. Fission Barrier B_f

B_f in ^{254}No . How does it test theories that predict properties of SHE?

- ^{254}No shell-stabilized, like SHE.
- **Dominant contribution to B_f from shell energy, which depends on single-particle spectrum (spe).**
- Theoretical approaches have different spe, hence, different B_f . (Mic-mac: 6.76 MeV; DFT: 6 – 12.5 MeV.)

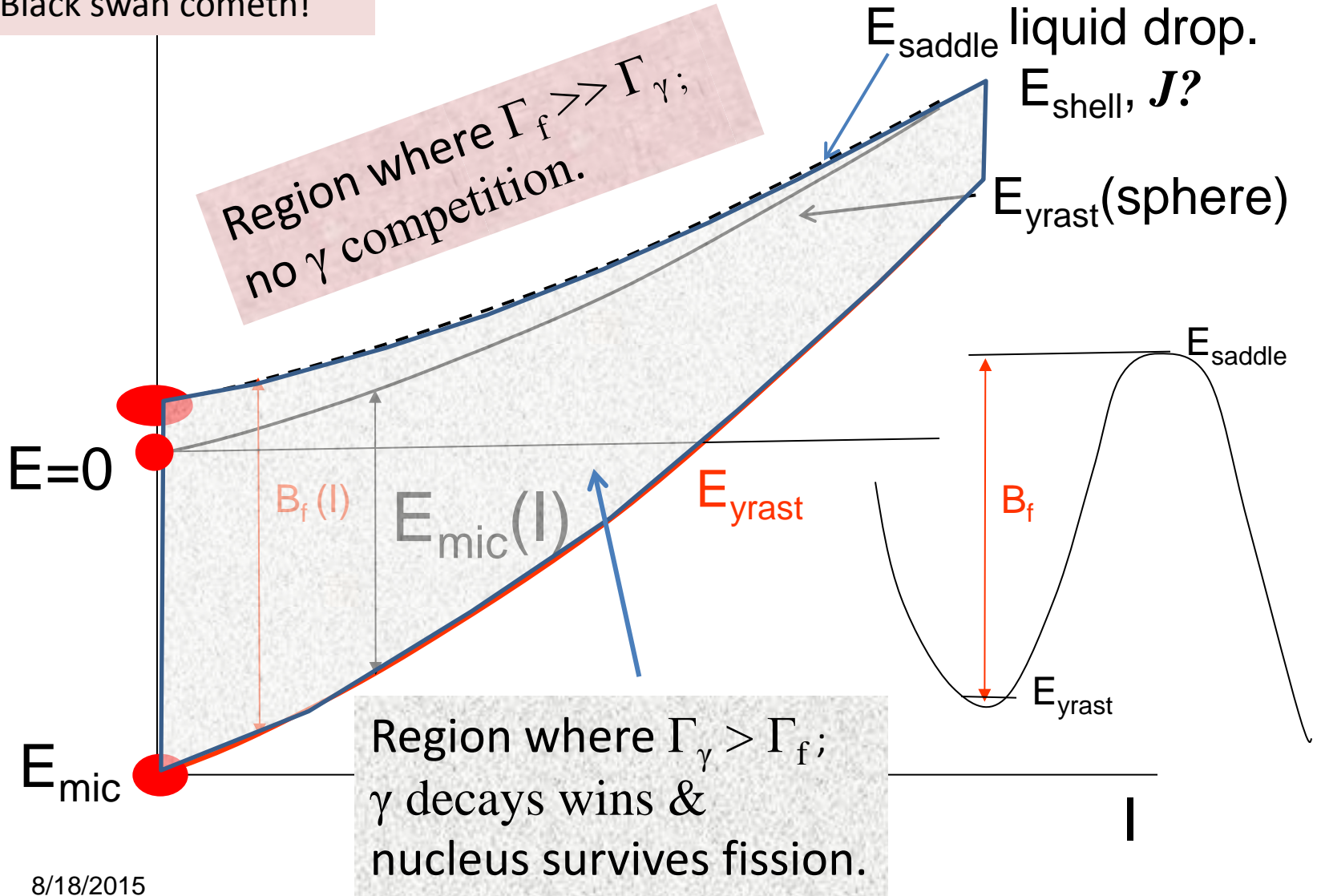
- Classical & most direct method: **fission threshold** after transfer reactions.
- Up to Cf (Z=98). **No results for heavier Z for >40 y** due to lack of stable target, so we **need a new method**.
- $\Gamma_\gamma / (\Gamma_\gamma + \Gamma_f) \rightarrow 0$ at **fission threshold**, below S_n , (where only γ & fission channels open).
- **γ calorimetry with Gammasphere \rightarrow entry distribution** (starting points for γ decay to ground state)
- **E_{cutoff}^* of entry distribution $\rightarrow E_{\text{saddle}}(I), B_f(I)$.**
- **$B_f(I \sim 15) \sim 6.6$ (5) MeV; $B_f(I=0) \sim 6.0$ (9) MeV.**

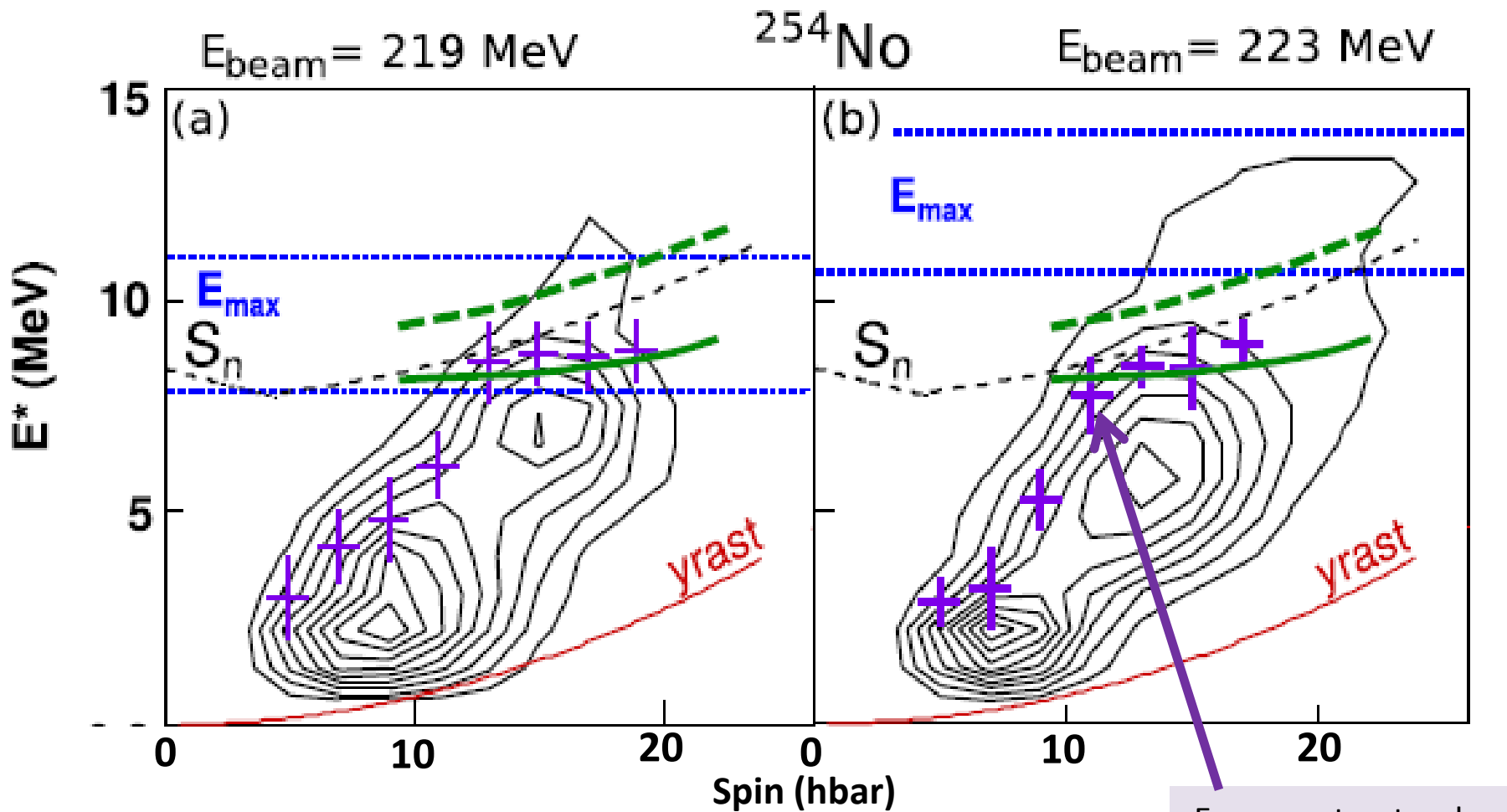
G. Henning et al., PRL 113, 262505 (2014)

Fission barrier $B_f(I)$ by measuring limits of γ -emitting region

Black swan cometh!

Region where $\Gamma_f \gg \Gamma_\gamma$;
no γ competition.





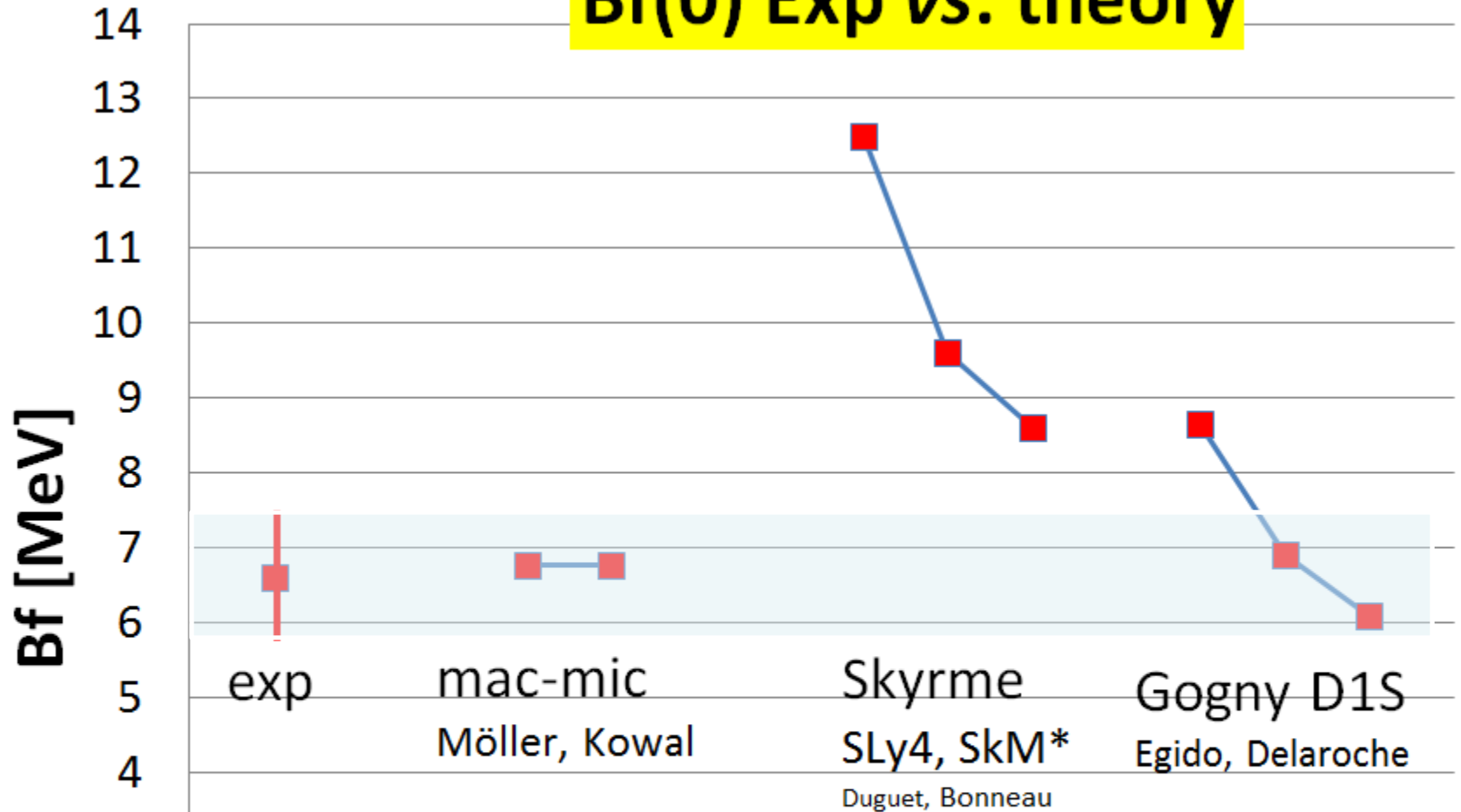
γ calorimetry with Gammasphere \rightarrow entry distribution

Measure $E_{\text{saddle}}(I)$, $B_f(I)$ by inspecting where, in excited states of ^{254}No , fission dominates γ decay and truncates the entry distribution.

$$B_f(I \sim 15) = 6.0(5) \text{ MeV.}$$

G. Henning et al., PRL 113, 262505 (2014)

Bf(0) Exp vs. theory



SHN formed at high spin.
 No yield at $l = 0$.
 If $E_{\text{shell}}(l) \sim 0$, no SHE.

$$\sigma(l) \text{ } ^{254}\text{No}$$

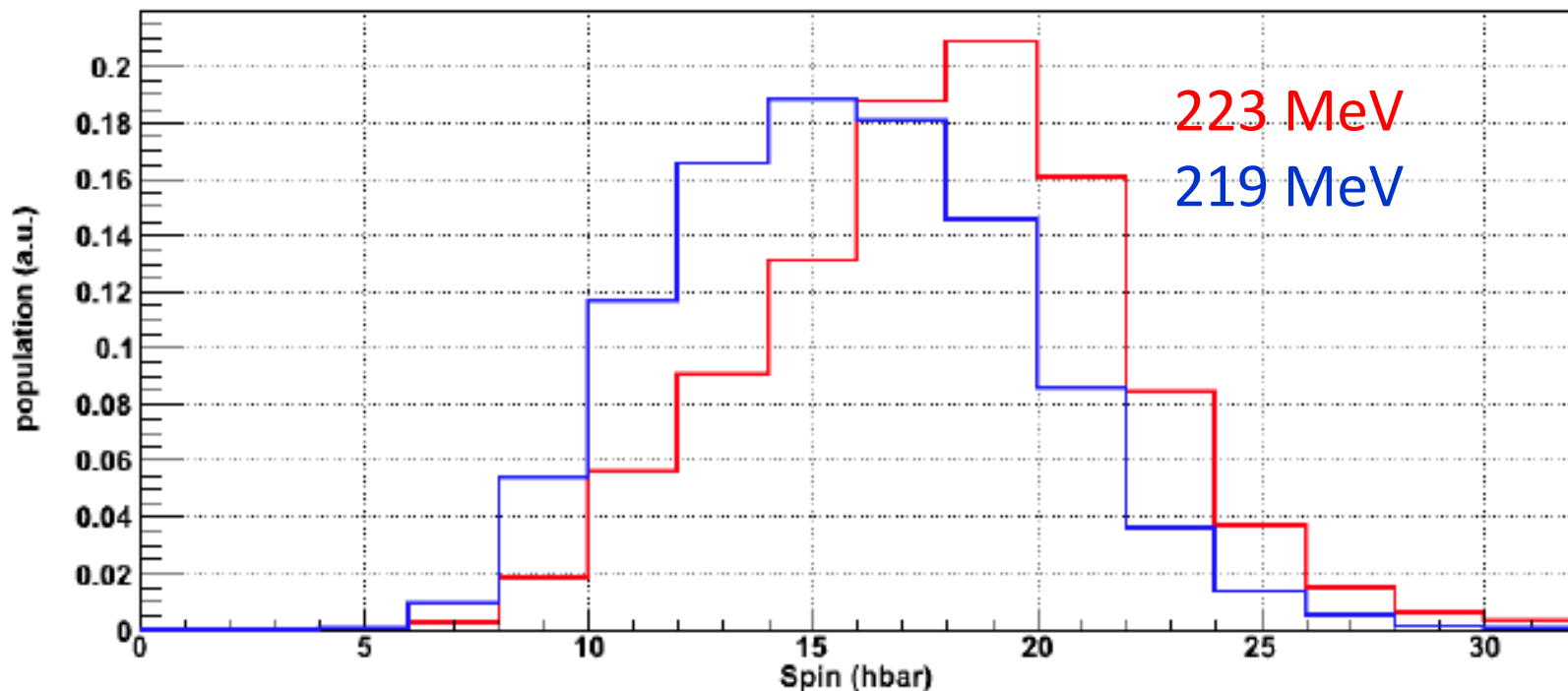


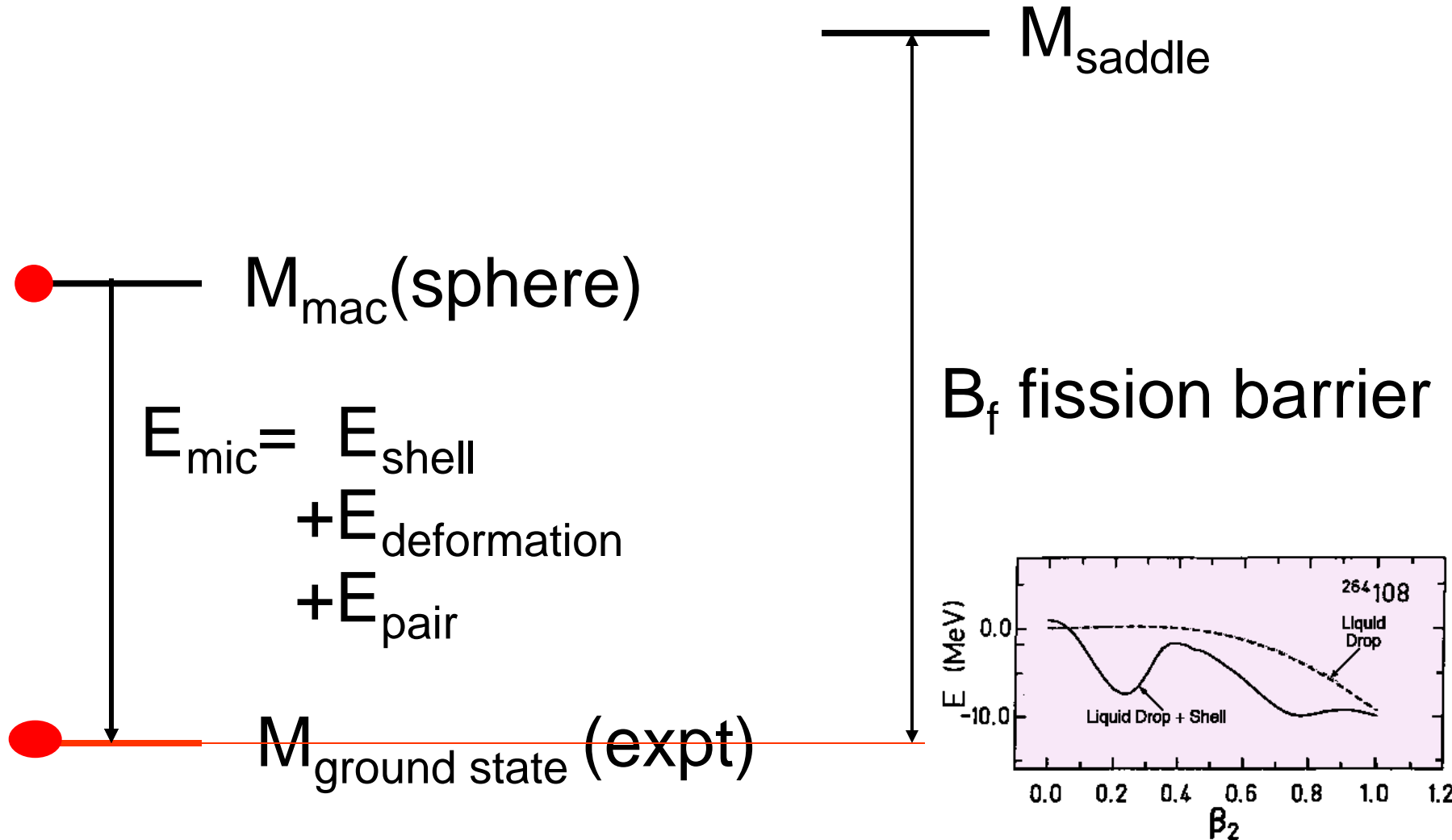
Figure 6.4: Spin distribution profile for $E_{\text{Beam}} = 219 \text{ MeV}$ recoil gated (blue) and $E_{\text{Beam}} = 223 \text{ MeV}$ recoil gated (red).

structure/reaction

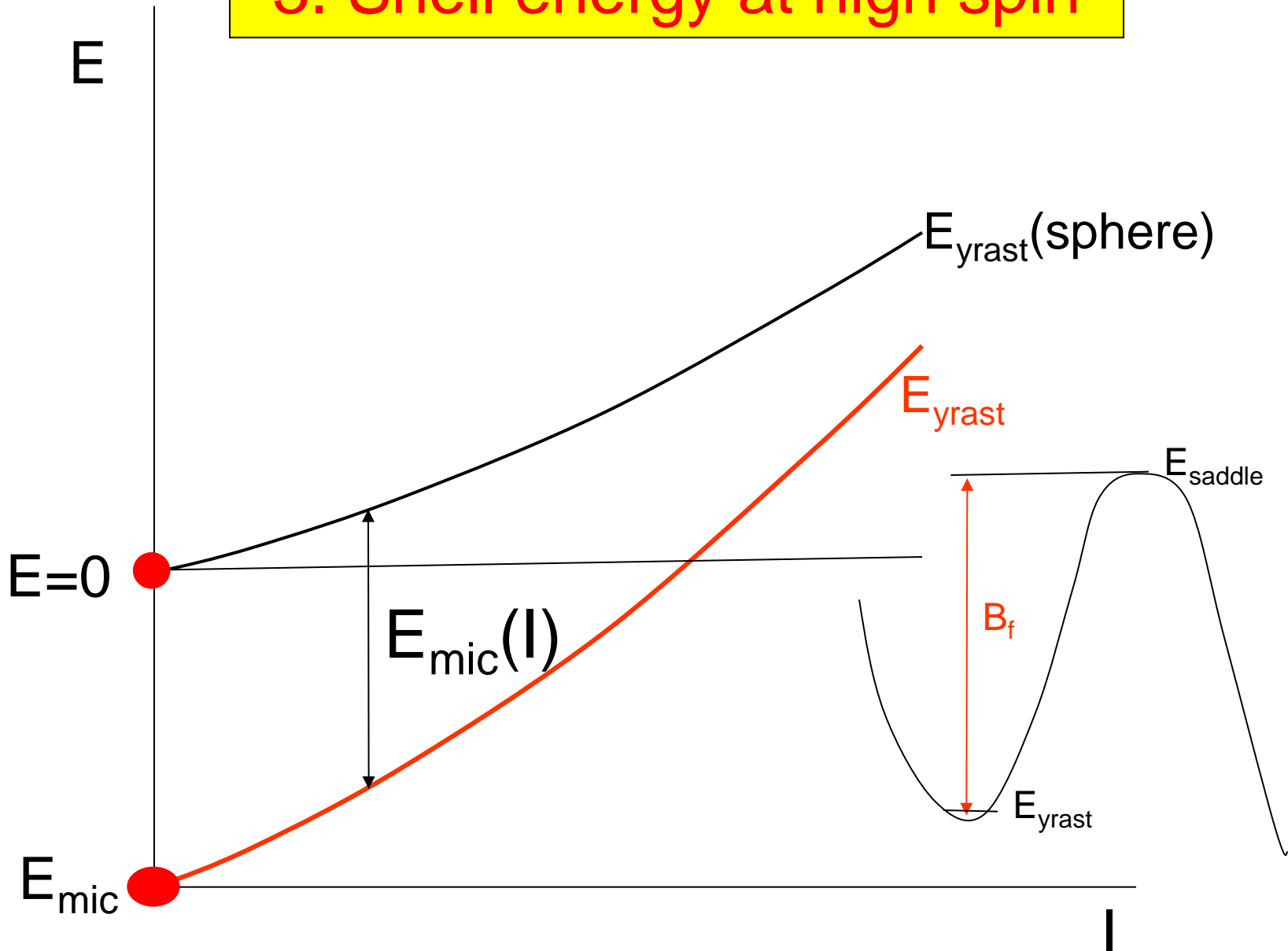
$B_f, \sigma(l) \leftrightarrow$ formation mechanism.

G. Henning et al., Ph. D. thesis, U. Paris Sud

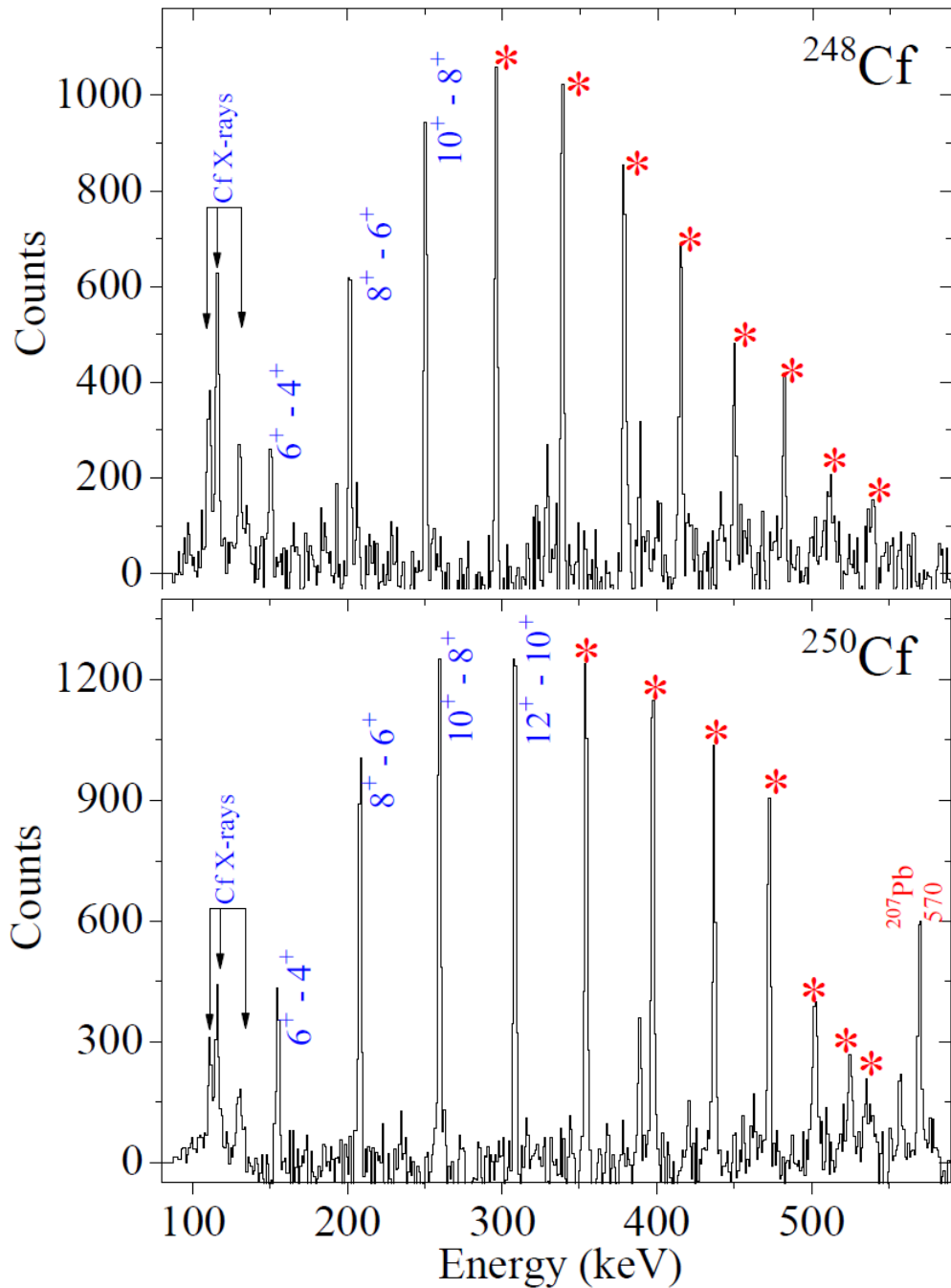
Ground-state shell energy



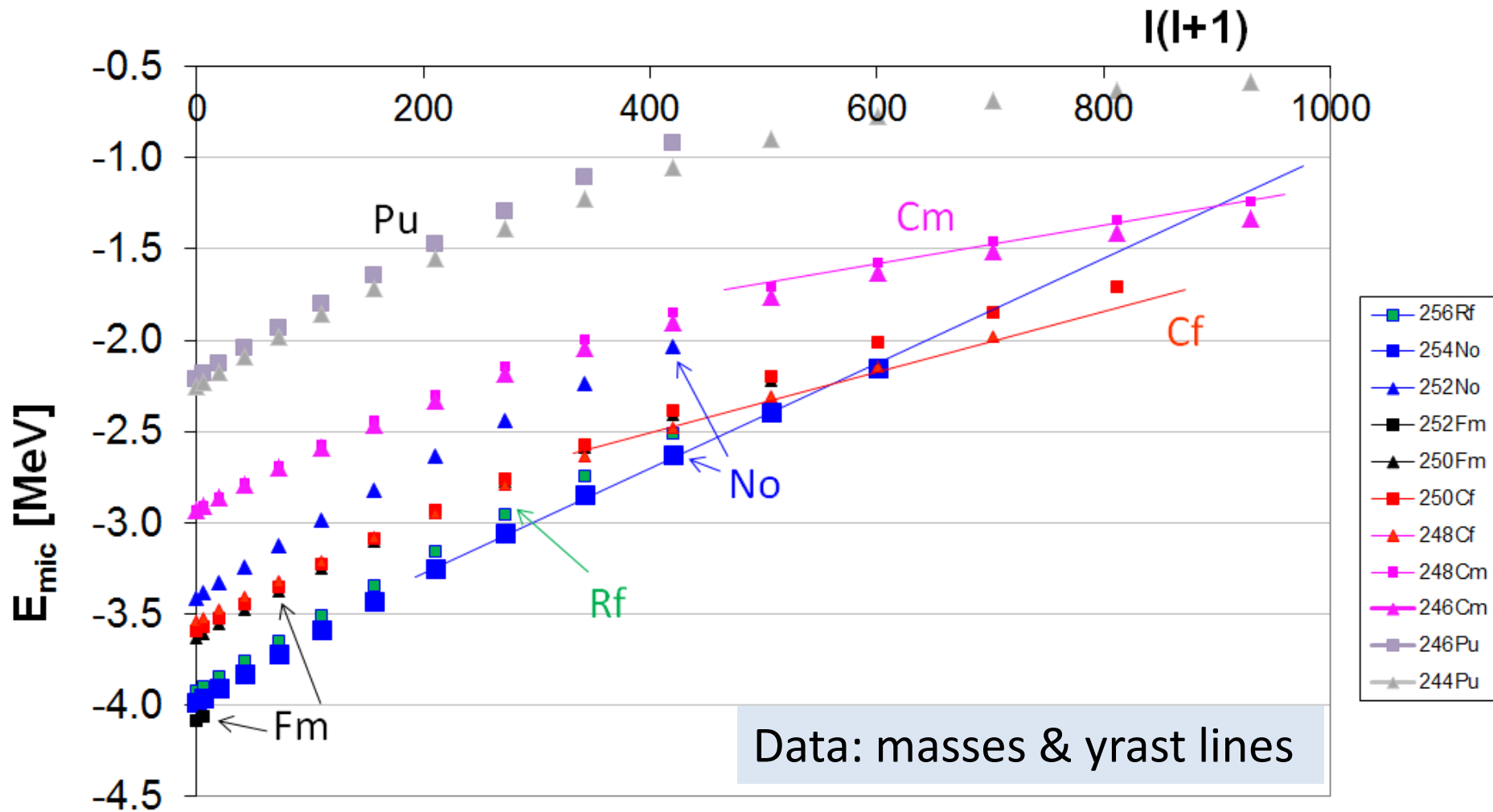
3. Shell energy at high spin



High spin states ($I \leq 32$)
 in $^{248,250}\text{Cf}$ from transfer
 reactions in $^{208}\text{Pb} + ^{249}\text{Cf}$



S. Hota et al., U. Mass Lowell,
 Ph. D. thesis, $^{208}\text{Pb} + ^{249}\text{Cf}$



- $E_{shell}(0)$ largest for ^{252}Fm due to gaps at $Z=100$, $N=152$
- E_{mic} decreases with I .
- Rate of decrease depends on moment of inertia, backbend.
- For $I \geq 24$, E_{shell} in ^{248}Cf is more negative than in ^{254}No .

- No calculation yet which simultaneously describes g.s. and yrast mass at high spin.

Summary

Theoretical description of SHN

A good model should describe **all** properties:

- single-particle spectra (proton and neutron gaps),
- $E_{\text{shell}}(I)$,
- masses of ground & saddle states,
- fission barriers.

Microscopic-macroscopic models (with WS or folded Yukawa potentials) almost satisfy this criterion.

DFT is limited by the hunt for a good effective interaction.

SHN robust & *populated predominantly at high spin.*

collaborators

**G. Henning^{1,6}, S. Hota³, A. Lopez-Martens⁶, I. Ahmad¹,
P. Chowdhury³, T.L. Khoo¹, F.G. Kondev¹, D. Seweryniak¹,
M. Asai⁴, B.B. Back¹, P. A. Butler⁸, M.P. Carpenter¹, C.N. Davids¹,
J. P. Greene¹, P. T. Greenlees⁵, K. Hauschild⁶, A. Heinz⁷, R.-D. Herzberg⁸,
R.V.F. Janssens¹, D.G. Jenkins², G.D. Jones⁸, S. Ketelhut⁵,
T. Lauritsen¹, C.J. Lister¹, E. McCutchan¹, T. Nakatsukasa⁹,
P. Papadakis⁸, D. Peterson¹, J. Qian⁷, P. Reiter, A. Robinson²,
S.K. Tandel³, U.S. Tandel³, X.F. Wang^{1,11}, S.F. Zhu¹**

¹Argonne National Laboratory, Argonne, IL, USA, ²University of York, Heslington, York, UK, ³University of Massachusetts Lowell, Lowell, MA, USA, ⁴Japan Atomic Energy Agency (JAEA), Japan, ⁵ University of Jyväskylä, Finland, ⁶CSNSM, IN2P3-CNRS, Orsay, France, ⁷WNSL, Yale University, New Haven, CT, USA, ⁸University of Liverpool, Liverpool, U.K., ⁹RIKEN Nishina Center, Japan, ¹⁰University of Maryland, College Park, MD, USA, ¹¹University of Notre Dame, Notre Dame, IN, USA

The end!

- Thank you for your attention.
- Thank ye.
- Diolch i chi.
- Go raibh maith agaibh.

Supported by U.S. Department of Energy, Office of Nuclear Physics, under Contract No. DE-AC02-06CH11357.