

# Beta-strength, decay heat and anti-neutrino spectra from Total Absorption Spectroscopy

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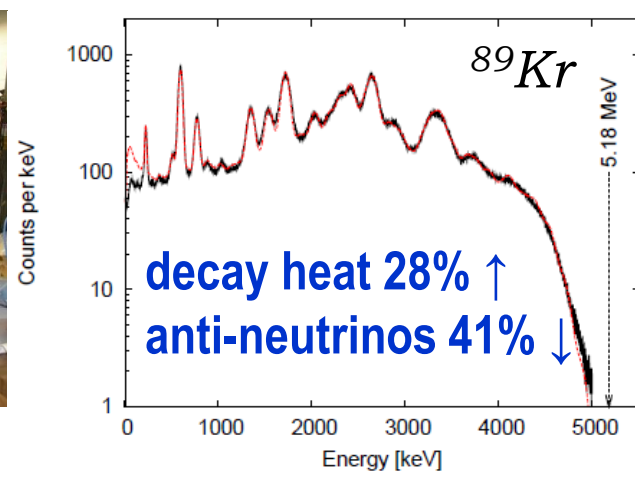
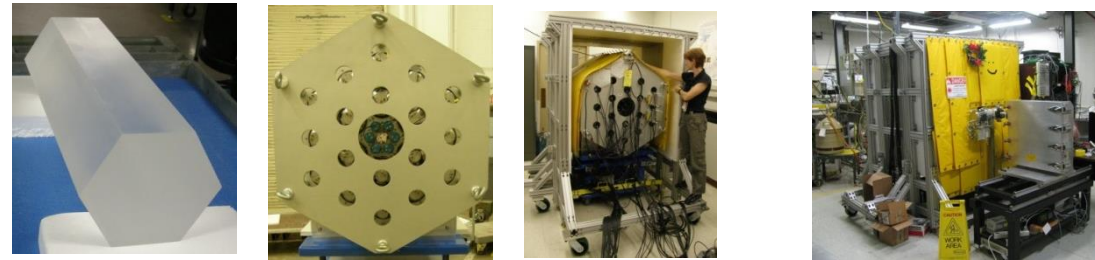
**A.Fijałkowska, M.Wolińska-Cichocka, B.C. Rasco, K. C. Goetz**

*M. Karny, R.K. Grzywacz, C.J. Gross, E. Zganjar, J.W. Johnson, N. Brewer*

*J. Matta, K. Miernik, D. Miller, M. Madurga, S. Paulauskas, S. Padgett,*

*D. Stracener, C. Jost, R. Goans, E. Spejewski, J. C. Batchelder, L. Cartegni,*

*M. Al-Shudifat, J. H. Hamilton and A.V. Ramayya*



An example of nuclear structure studies truly interesting for nuclear energy and fundamental neutrino physics

# Outline

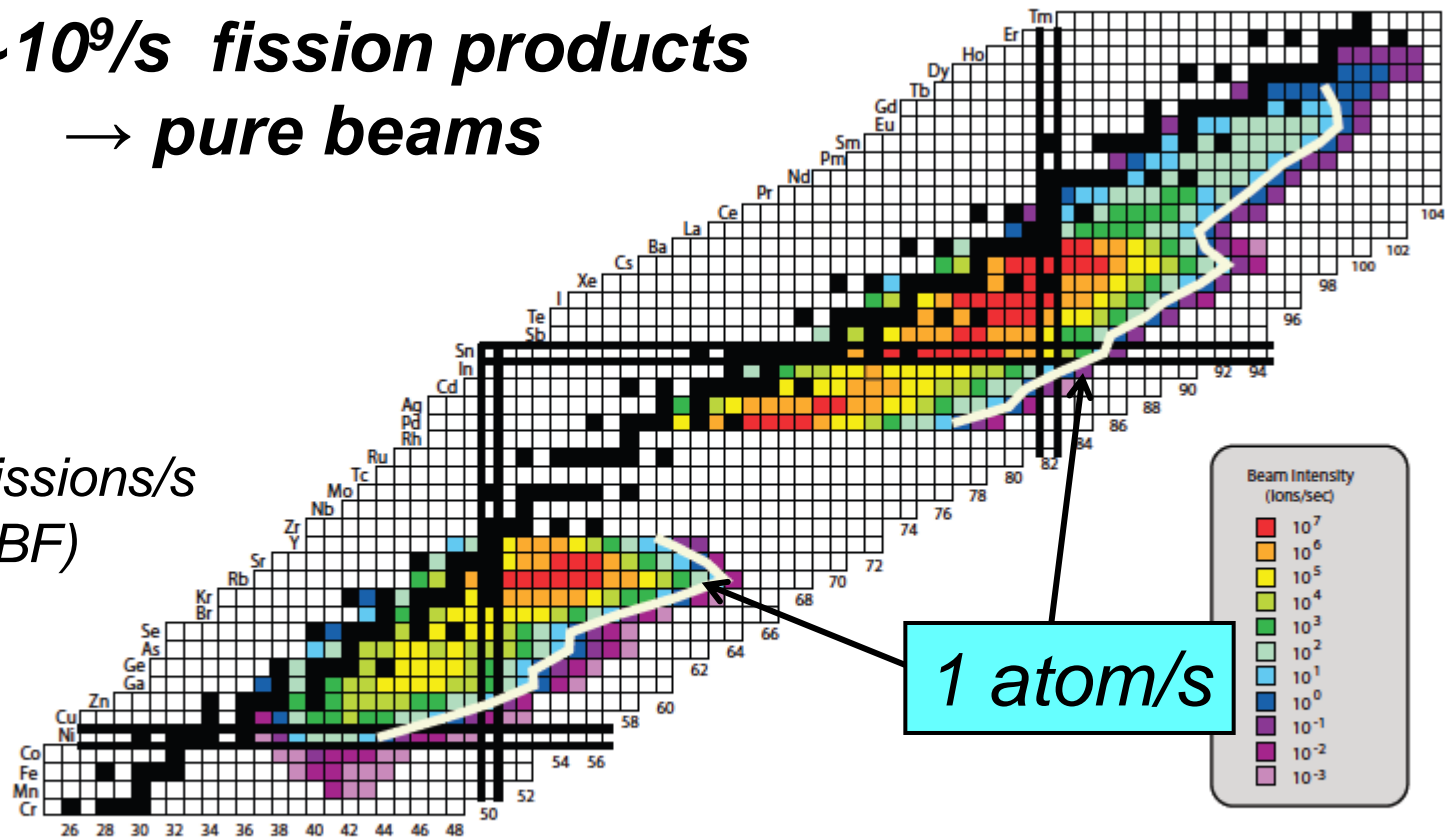
1. Physics of beta decay of neutron-rich nuclei (fission fragment)
2. Results on  $\beta$ -strength in fission fragments, and their impact
  - decay heat and reactor anti-neutrino anomaly
  - spectrum of high-energy reactor anti-neutrinos



**Proton-induced fission of  $^{238}\text{U}$  creates many neutron-rich nuclei, with a similar yield pattern as in a nuclear reactor**

**$\sim 10^9/s$  fission products  
→ pure beams**

*(we had  $10^{11}$  fissions/s at the HRIBF)*

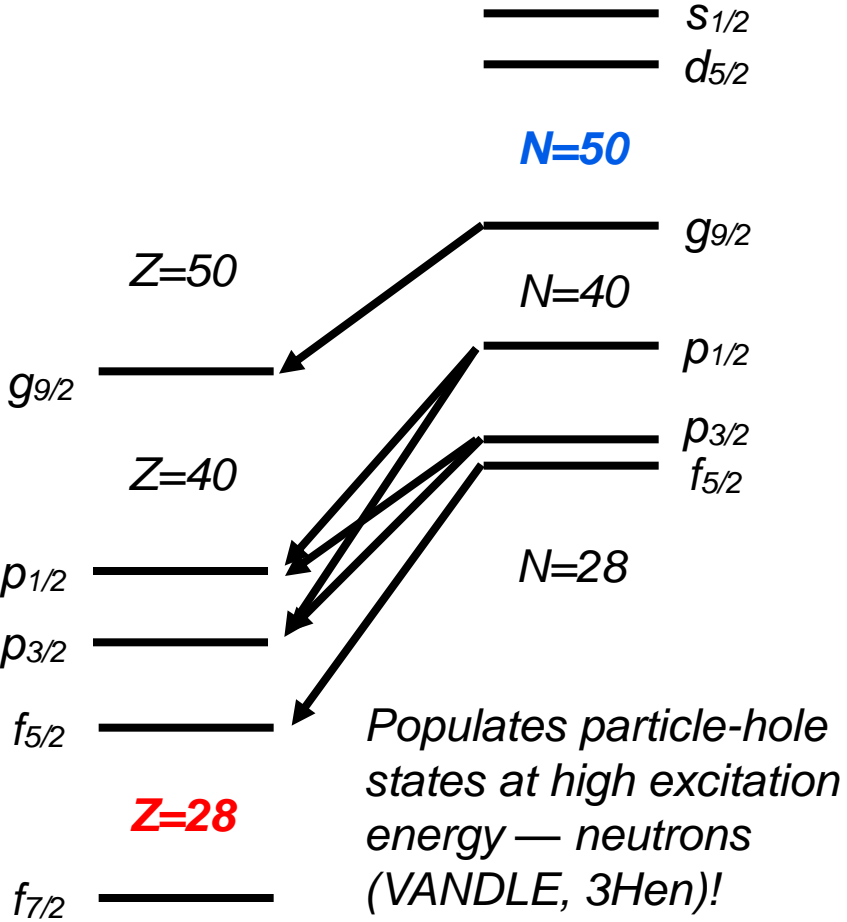


*The chart of nuclides indicates by color the expected intensity of fission fragments delivered to the measuring end-station with 100 nA of 46 MeV protons on a thick  $\text{UC}_x$  target.*

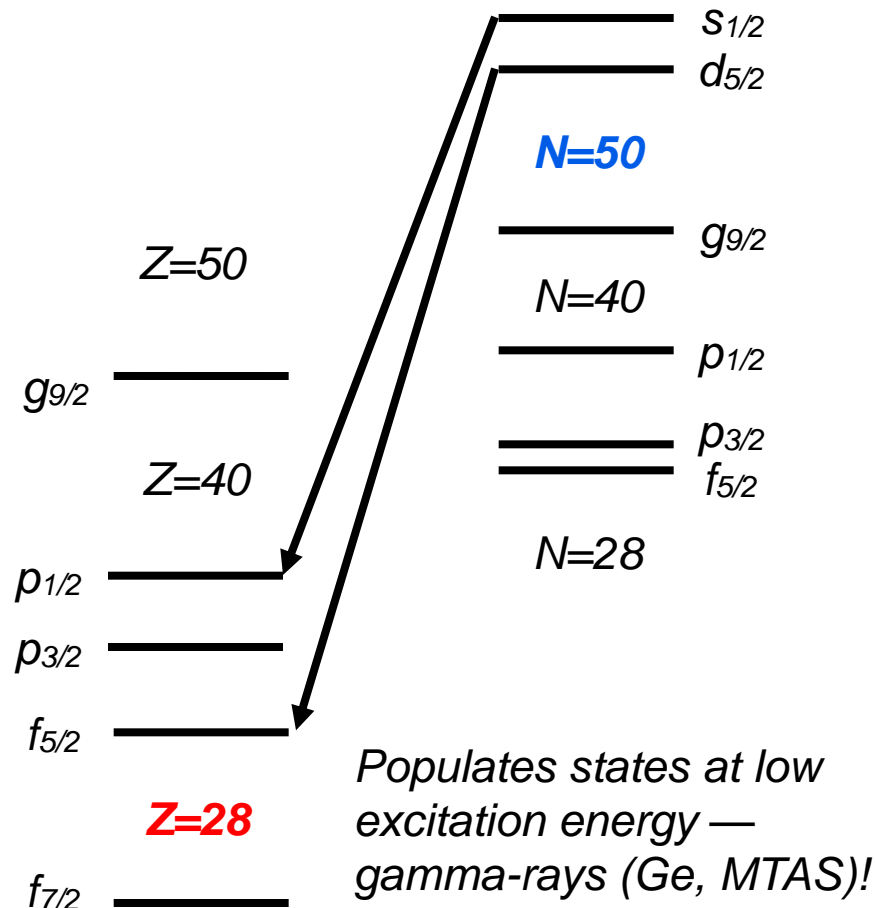
**Nuclear structure,  $r$ -process nucleosynthesis, reactor applications and reactor anti-neutrino physics**

**Exotic n-rich nuclei with large  $Q_\beta \rightarrow$  competition between allowed and forbidden  $\beta$ -decays**

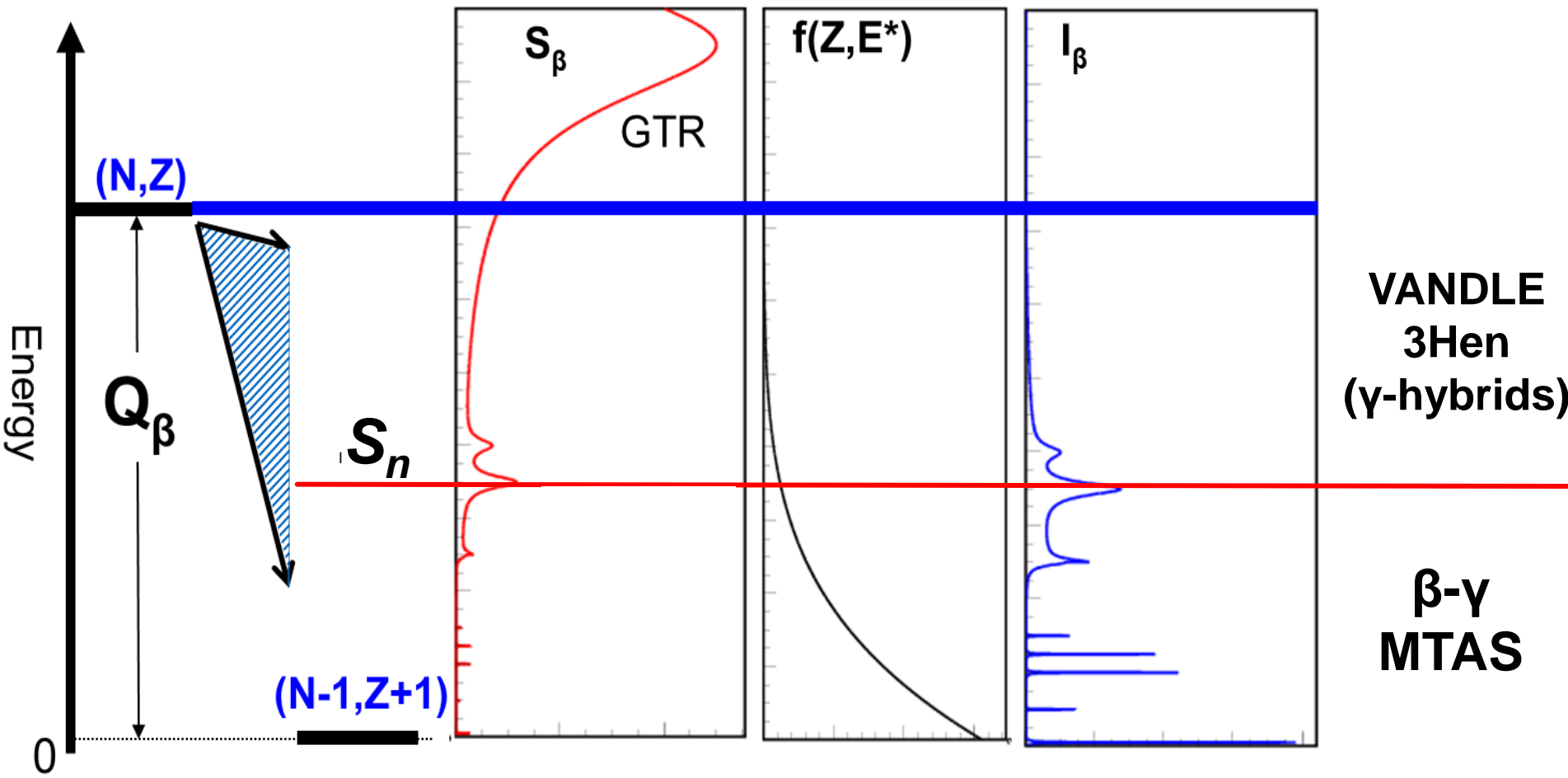
*Gamow-Teller transitions*  
 large strength, small beta energy  
 no parity change  
 no angular momentum change  
 spin change 0, 1



*Forbidden transitions*  
 small strength, large beta energy  
 parity change  
 angular momentum change  
 spin change 0, 1, 2



# Complex $\beta$ -decay of a very neutron-rich nucleus (like $^{91}\text{Br}$ )



# Complex $\beta$ -decays

N-RICH PARENT (Z,N)

“Pandemonium”

- \*  $\beta$ -transitions (mostly Gamow-Teller) are feeding highly excited states,
- \* these many, weak  $\beta$ -transitions are followed by the cascades of  $\gamma$ -transitions in the daughter nucleus,
- \* these weak  $\gamma$ -transitions are very difficult to impossible to detect with radiation detectors with low efficiency

## Total absorption $\gamma$ -spectroscopy

- \* to determine **true  $\beta$ -feeding** and resulting  $\gamma$ -decay patterns (nuclear structure),
- \* to determine **“decay heat”** released by radioactive nuclei produced during a nuclear fuel cycle
- \* to determine **anti-neutrino spectrum** associated with  $\beta$ -decay of fission products,  $Y \rightarrow X^* + e^- + \bar{\nu}$

$\beta$  - transitions

$\gamma$  - transitions

DAUGHTER (Z+1, N-1)

The true picture of the neutron-rich parent nucleus (Z,N), with many weak  $\beta$ -transitions and following low intensity  $\gamma$ -transitions.

**J. Hardy et al., Physics Letters 71 B, 307, 1977**

A. Algora et al., PRL 105, 202501, 2010

K. P. Rykaczewski, Viewpoint in Physics 3, 94, 2010

K. P. Rykaczewski, 2013 McGraw-Hill Yearbook of Science and Technology, p.9

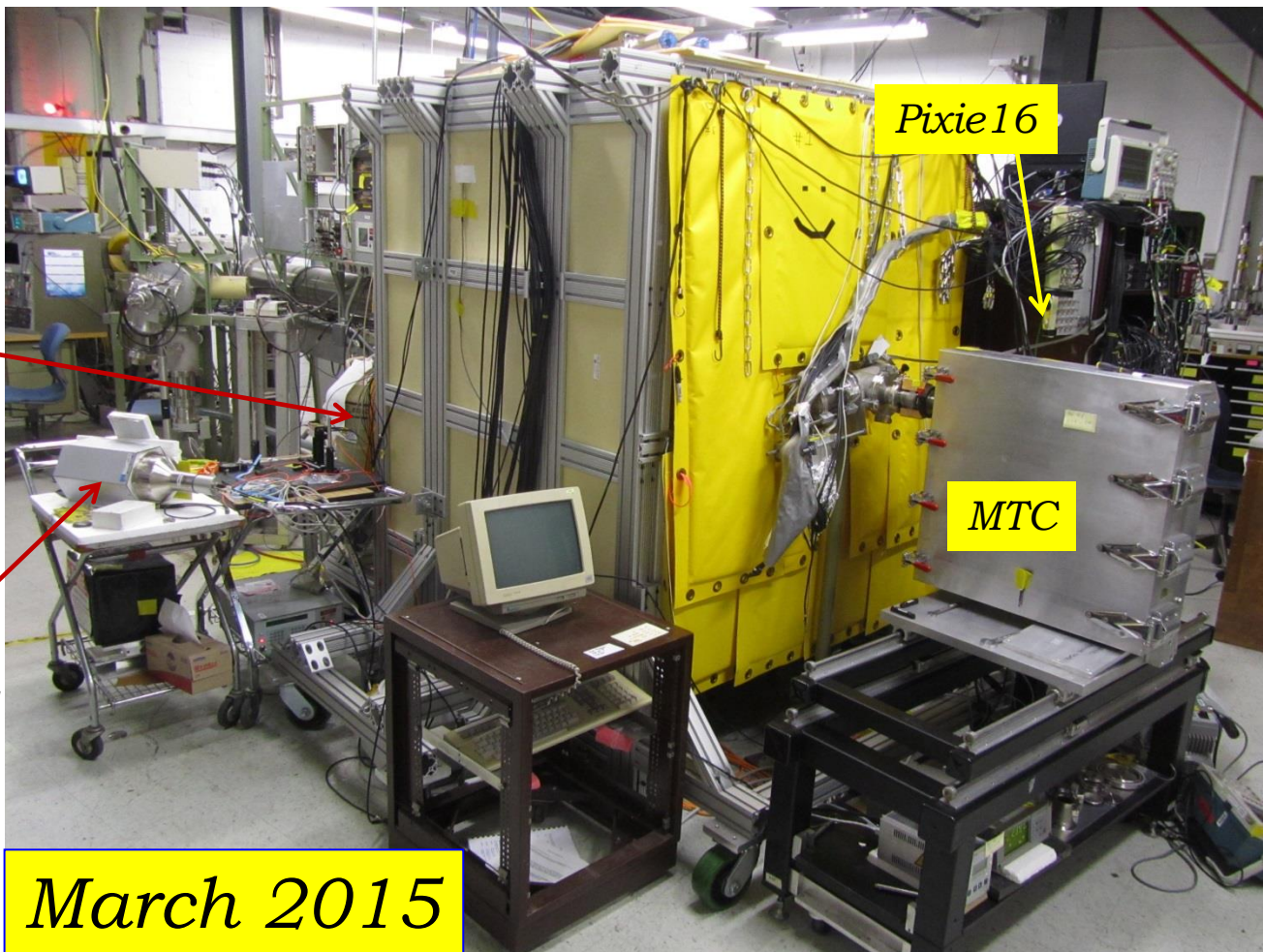
# Unique “state-of-art” MTAS on-line to the mass separator OLTF

*Data ACQ  
digital pulse processing  
(R. Grzywacz et al.)*

*Ge detector  
for monitoring  
collection point*

*Reference NaI(Tl)  
crystal with pulsed blue  
laser light for monitoring  
amplification changes.  
Laser light was split  
between reference and  
central/inner ring  
detectors.*

*K.C. Goetz, R. Grzywacz  
(UTK)*

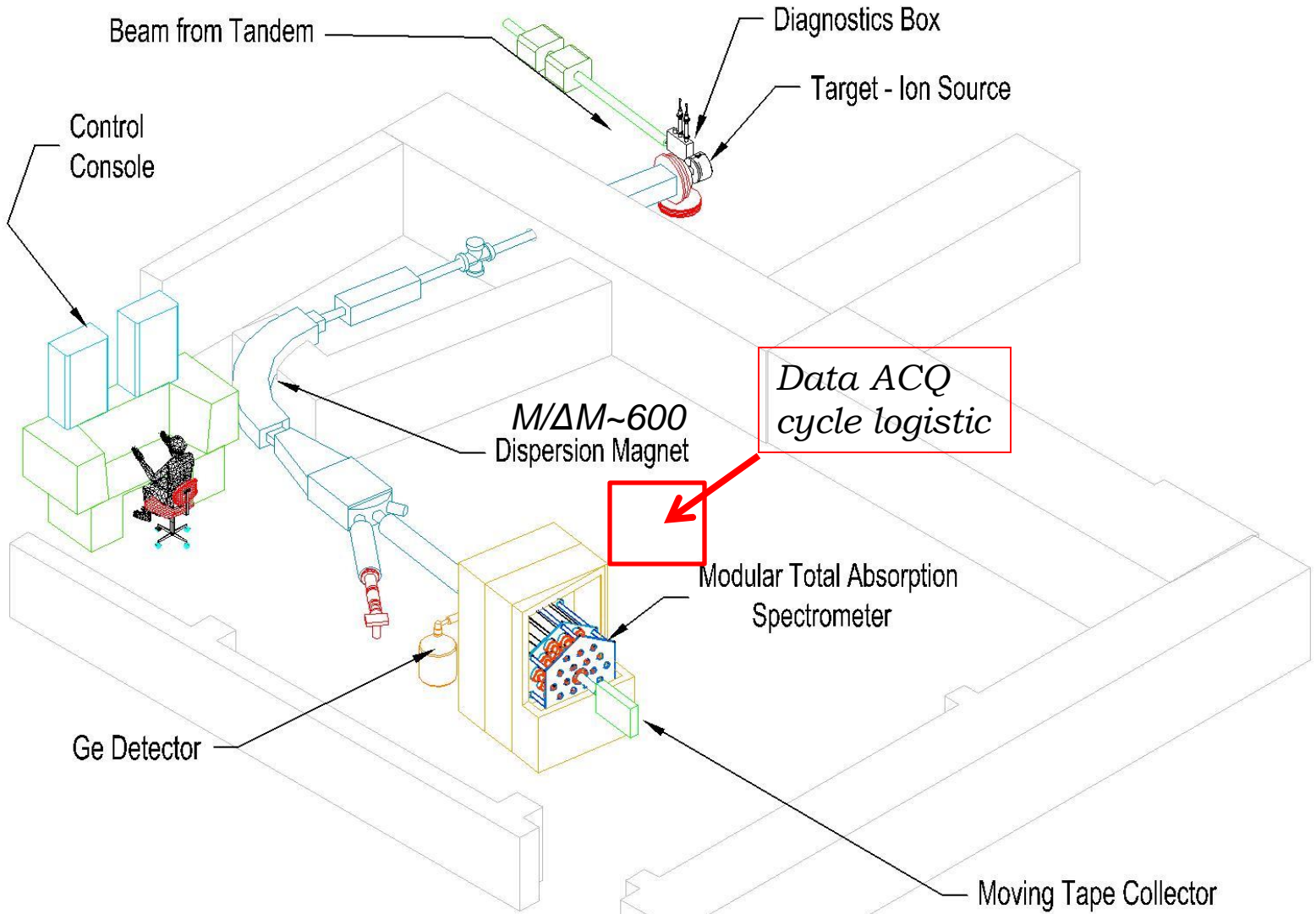


**March 2015**

*Multi-point temperature monitor, MTC and OLTF cycle control (C.J. Gross)*

**Many months of calibrations and perfecting the response function  
e.g., Rasco et al., NIM 788, 139, 2015; Wolinska et al, NDS 120, 22, 2014**

~ 50nA protons



**On-Line Test Facility OLTF = mass separator on-line to ORNL's Tandem  
(last on-line mass separator in the US)**



# Decays measured with MTAS in January 2012 and March 2015

Y 87 79.8 h	Y 88 106.65 d	Y 89 100	Y 90 64.00 h	Y 91 58.51 d	Y 92 3.54 h	Y 93 10.18 h	Y 94 18.7 m	Y 95 10.3 m	Y 96 5.34 s	Y 97 3.75 s	Y 98 548 ms
Sr 86 9.86	Sr 87 7.00	Sr 88 82.58	Sr 89 50.53 d	Sr 90 28.79 y	Sr 91 9.63 h	Sr 92 2.66 h	Sr 93 7.423 m	Sr 94 75.3 s	Sr 95 v 23.90 s	Sr 96 1.07 s	Sr 97 429 ms
Rb 85 72.17	Rb 86 18.642 d	Rb 87 27.83	Rb 88 17.78 m	Rb 89 15.15 m	Rb 90 2 v 2.6 m	Rb 91 58.4 s	Rb 92 2 v 4.492 s	Rb 93 v 5.84 s	Rb 94 v 2.702 s	Rb 95 377.5 ms	Rb 96 203 ms
Kr 84 57.00	Kr 85 10.776 y	Kr 86 17.30	Kr 87 76.3 m	Kr 88 2.84 h	Kr 89 1 3.18 m	Kr 90 1 32.32 s	Kr 91 v 8.57 s	Kr 92 1.840 s	Kr 93 1.286 s	Kr 94 210 ms	Kr 95 114 ms
Br 83 2.40 h	Br 84 31.80 m	Br 85 2.90 m	Br 86 1 v 55.1 s	Br 87 1 55.65 s	Br 88 1 v 16.36 s	Br 89 v 4.40 s	Br 90 1.910 s	Br 91 0.64 s	Br 92 343 ms	Br 93 102 ms	Br 94 70 ms
Se 82 8.73	Se 83 22.3 m	Se 84 3.1 m	Se 85 33 s	Se 86 14.1 s	Se 87 5.8 s	Se 88 1.53 s	Se 89 410 ms	Se 90 >300 ns	Se 91 270 ms	Se 92 100 ms	Se 93 50 ms



Priority 1 for decay heat simulation according to NEA 2007 assessment (5)



Decay among 20+ most important ones for high energy reactor anti-neutrino spectra, Sonzogni et al., PR C 2015 (9)

# Decays measured with MTAS in January 2012 and March 2015

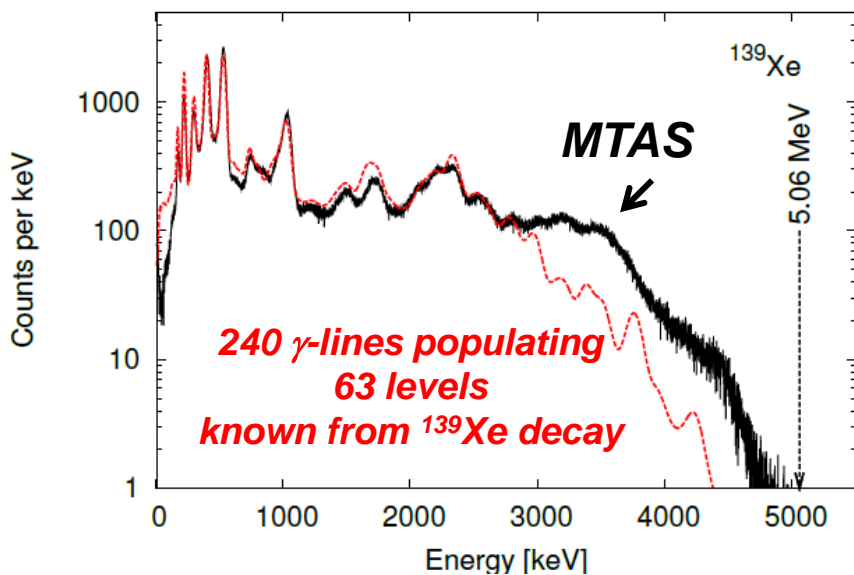
La 139 99.910	La 140 1.6781 d	La 141 3.92 h	La 142 92.6 m	La 143 14.3 m	La 144 40.9 s	La 145 24.8 s	La 146 6.27 s	La 147 4.015 s	La 148 1.26 s
Ba 138 71.698	Ba 139 83.06 m	Ba 140 12.752 d	Ba 141 18.27 m	Ba 142 10.7 m	Ba 143 14.5 s	Ba 144 11.5 s	Ba 145 4.31 s	Ba 146 2.22 s	Ba 147 893 ms
Cs 137 30.1671 y	Cs 138 32.2 m	Cs 139 9.27 m	Cs 140 63.7 s <span style="color: yellow;">v</span>	Cs 141 24.94 s	Cs 142 1.689 s <span style="color: yellow;">3 v</span>	Cs 143 1.791 s	Cs 144 994 ms	Cs 145 582 ms	Cs 146 323 ms
Xe 136 8.87	Xe 137 3.83 m <span style="color: yellow;">1</span>	Xe 138 14.08 m <span style="color: yellow;">v</span>	Xe 139 39.68 s <span style="color: yellow;">1</span>	Xe 140 13.60 s <span style="color: yellow;">1</span>	Xe 141 1.73 s	Xe 142 1.22 s	Xe 143 511 ms	Xe 144 388 ms	Xe 145 188 ms
I 135 6.61 h	I 136 45s 84s <span style="color: yellow;">1 1</span>	I 137 24.2 s <span style="color: yellow;">1</span>	I 138 6.4 s	I 139 2.29 s	I 140 860 ms	I 141 430 ms	I 142 ~200 ms	I 143 100 ms	I 144 50 ms
Te 134 41.8 m	Te 135 18.6 s	Te 136 17.5 s	Te 137 2.49 s	Te 138 1.4 s	Te 139 >300 ns	Te 140 300 ms	Te 141 100 ms	Te 142 50 ms	
Sb 133 2.5 m	Sb 134 780 ms	Sb 135 1.68 s	Sb 136 923 ms	Sb 137 450 ms	Sb 138 500 ms	Sb 139 300 ms			

6 1 3 v

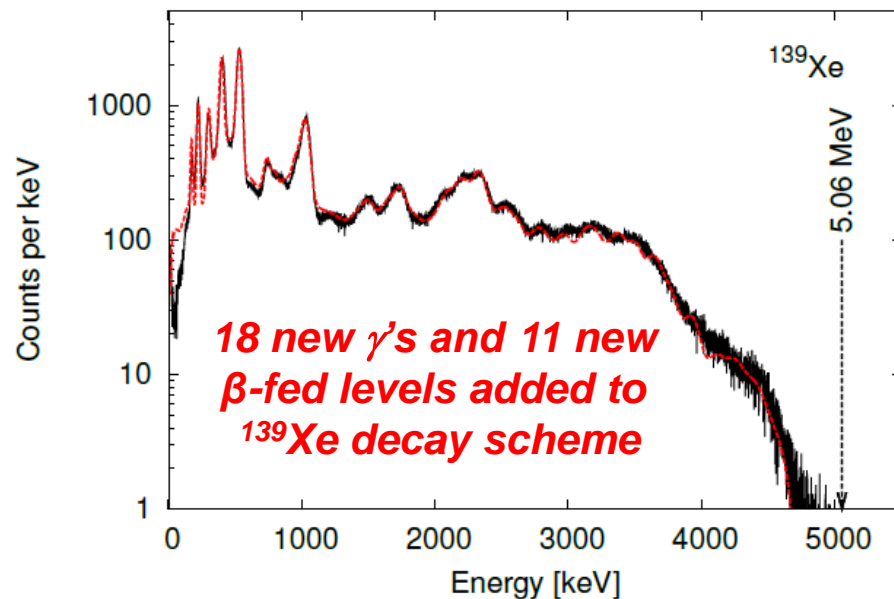
# Priority "1" $^{139}\text{Xe} \rightarrow ^{139}\text{Cs}$ decay ( $T_{1/2} = 39.7 \text{ s}$ )

cumulative yield of  $^{139}\text{Xe}$  in  $n_{th} + ^{235}\text{U}$  fission is about 5%

MTAS full spectrum ENSDF decay scheme



MTAS full spectrum - evaluated



## MTAS result:

average **gamma** energy release per  $^{139}\text{Xe}$  decay increased from 935 keV to 1370 keV

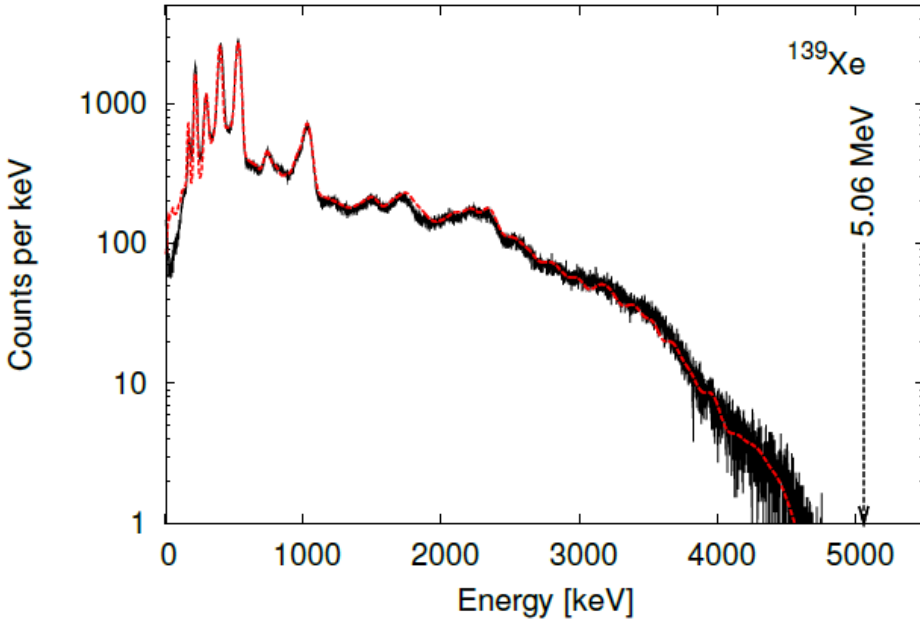
**47 % increase**

average **beta** energy release per  $^{139}\text{Xe}$  decay decreased from 1774 keV to 1573 keV

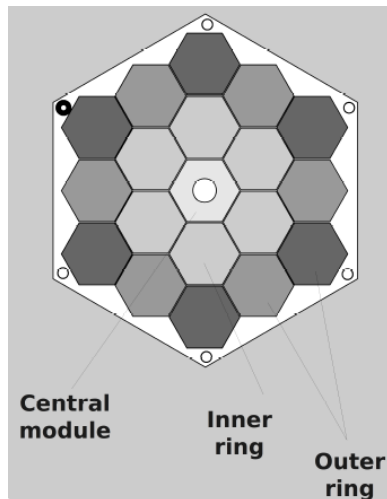
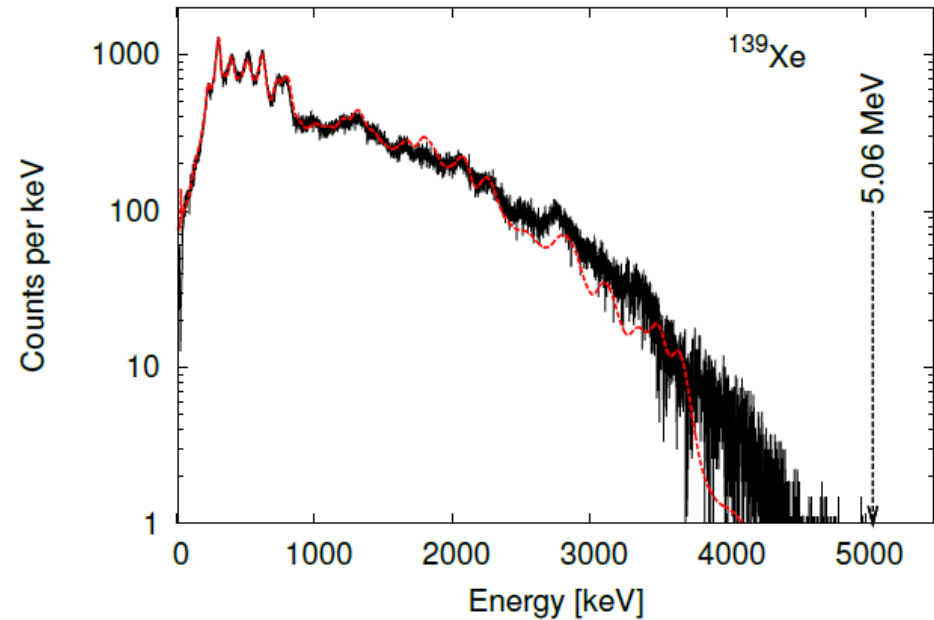
**13 % decrease**

# Measured and modified $^{139}\text{Xe}$ decay in central and inner rings of MTAS

MTAS Central detector only



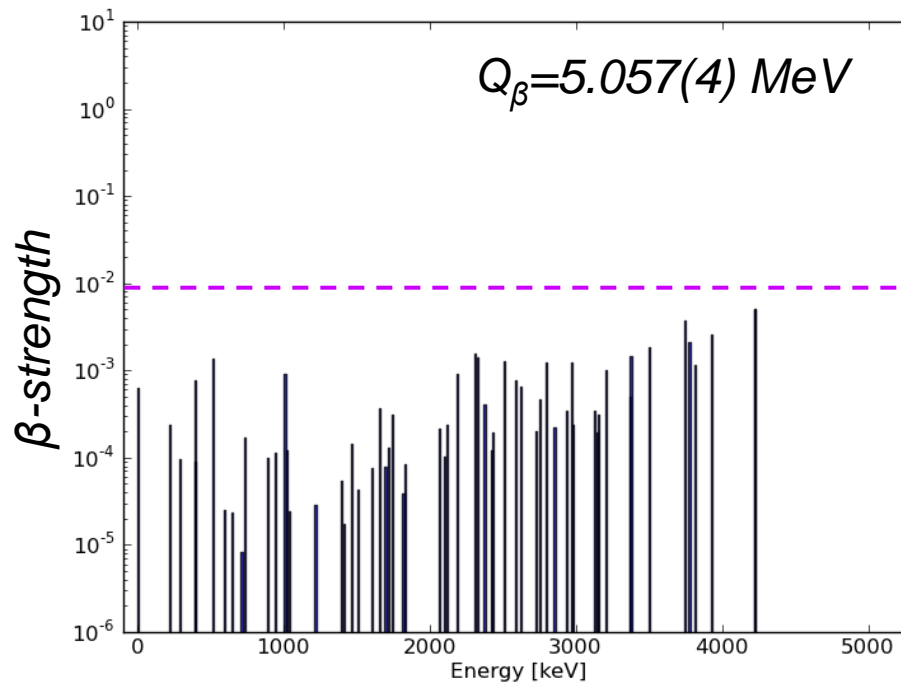
MTAS I-ring detectors



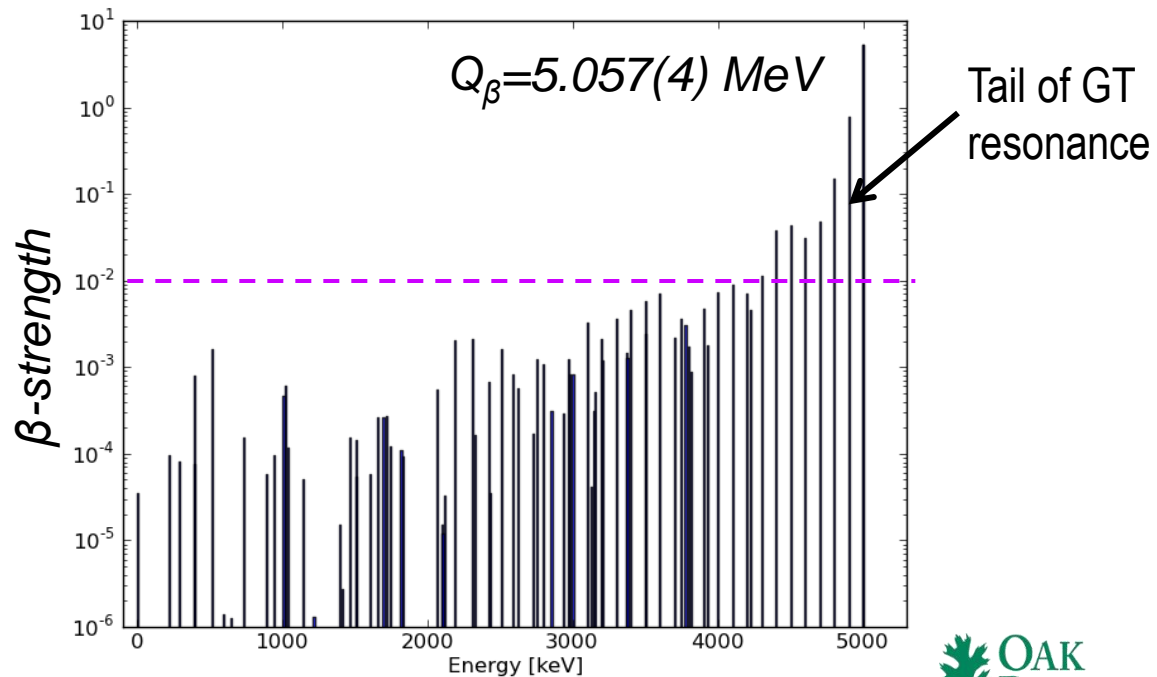
*Modular construction of MTAS  
helps to validate results !*

see also *B.C. Rasco et al., ARIS 2014 proceedings*

**$\beta$ -strength in  $^{139}\text{Xe}$  decay**  
**ENSDF**



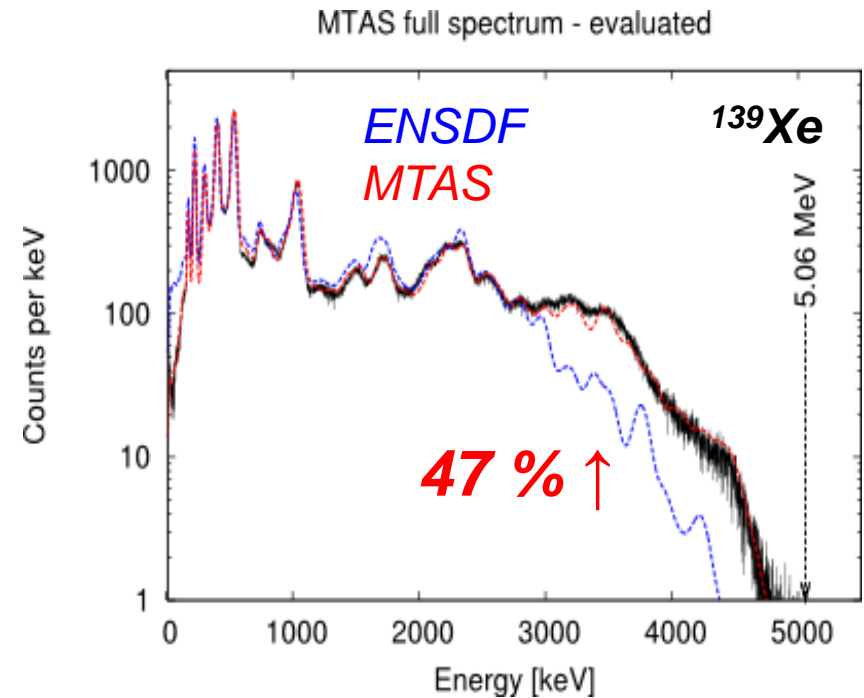
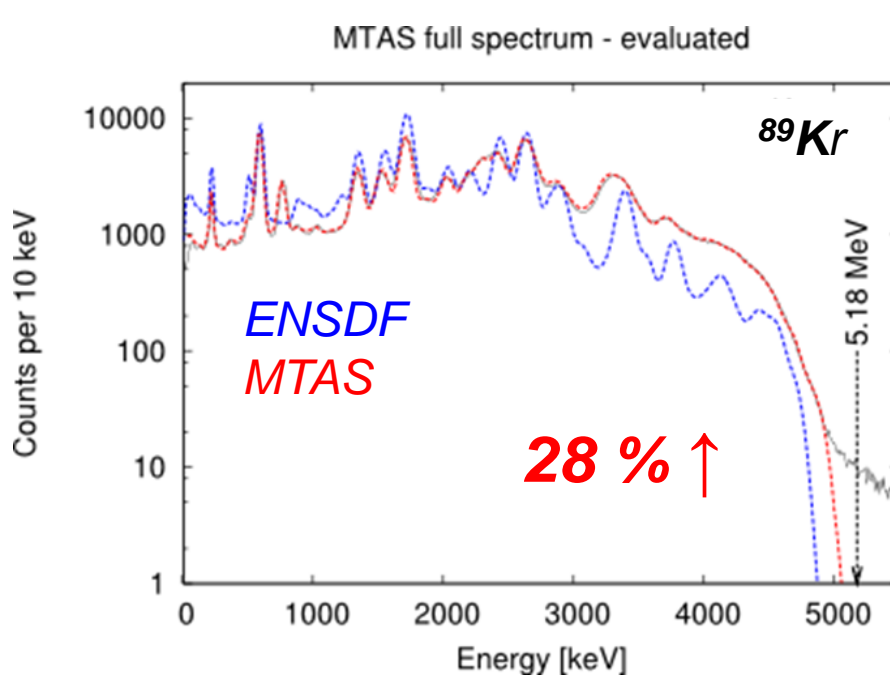
**$\beta$ -strength in  $^{139}\text{Xe}$  decay**  
**MTAS**



**A. Fijałkowska**  
**PhD thesis 2015**

# Decay heat $\langle E_\nu \rangle$ in $^{89}\text{Kr}$ and $^{139}\text{Xe}$ studied with MTAS

production in  $n_{th} + ^{235}\text{U}$ :  $^{89}\text{Kr} \sim 4.5\%$  and  $^{139}\text{Xe} \sim 5\%$



**MTAS results for  $^{89}\text{Kr}$  and  $^{139}\text{Xe}$  decays point to much larger  $\langle E_\nu \rangle$  values**

$^{89}\text{Kr}$ : 1801 keV (ENSDF) to 2304 keV (MTAS),  $^{139}\text{Xe}$ : from 935 keV to 1370 keV.

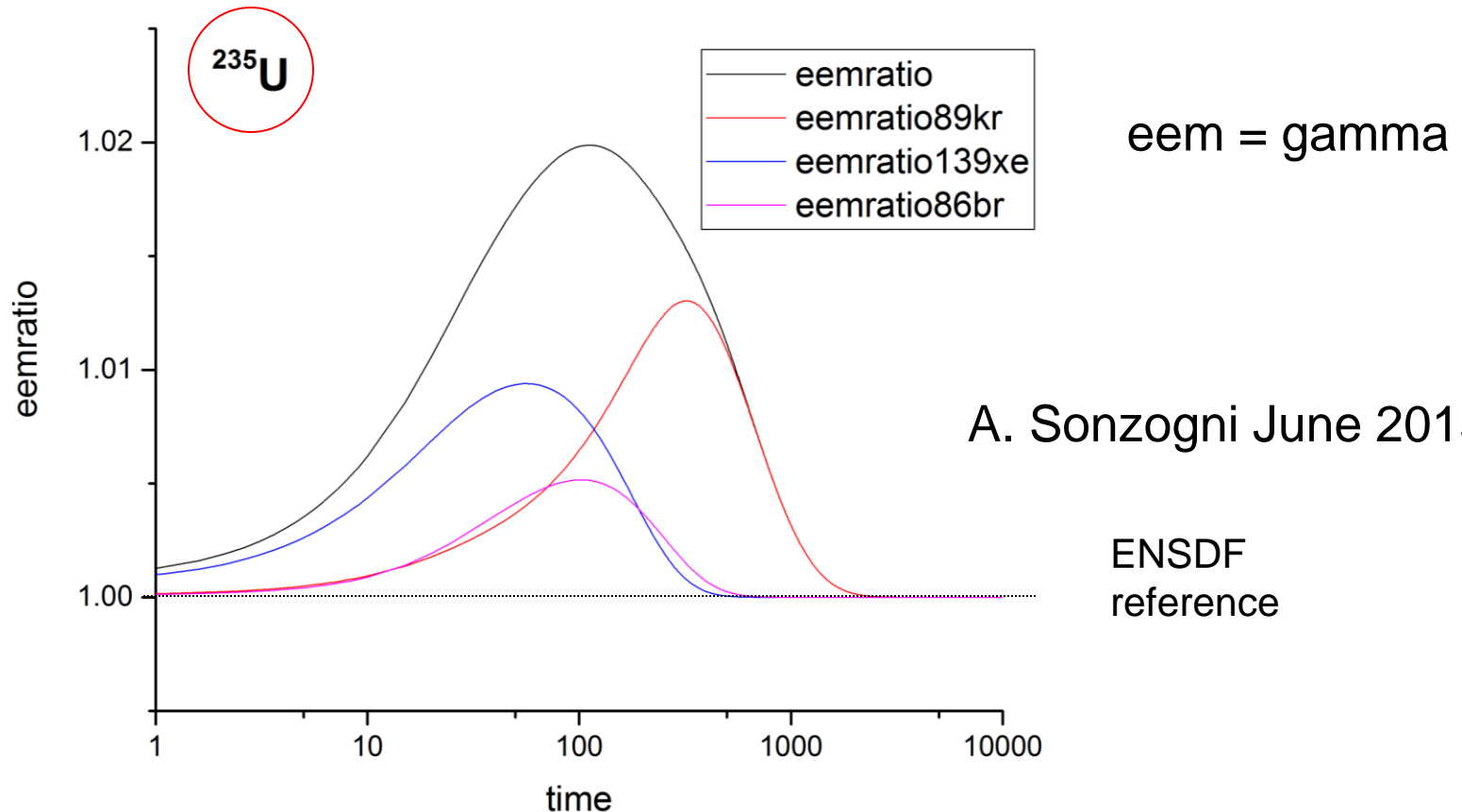
**Respectively, average beta and anti-neutrino energy go down**

$^{89}\text{Kr}$ :  $\langle E_\beta \rangle = 1462$  keV (ENSDF) to  $\langle E_\beta \rangle = 1222$  keV (MTAS),

$^{139}\text{Xe}$ : 1774 keV to 1573 keV

# Impact of $^{86}\text{Br}$ , $^{89}\text{Kr}$ and $^{139}\text{Xe}$ on total decay heat pattern

Gamma decay heat ratio: (MTAS and ENSDF)/ENSDF



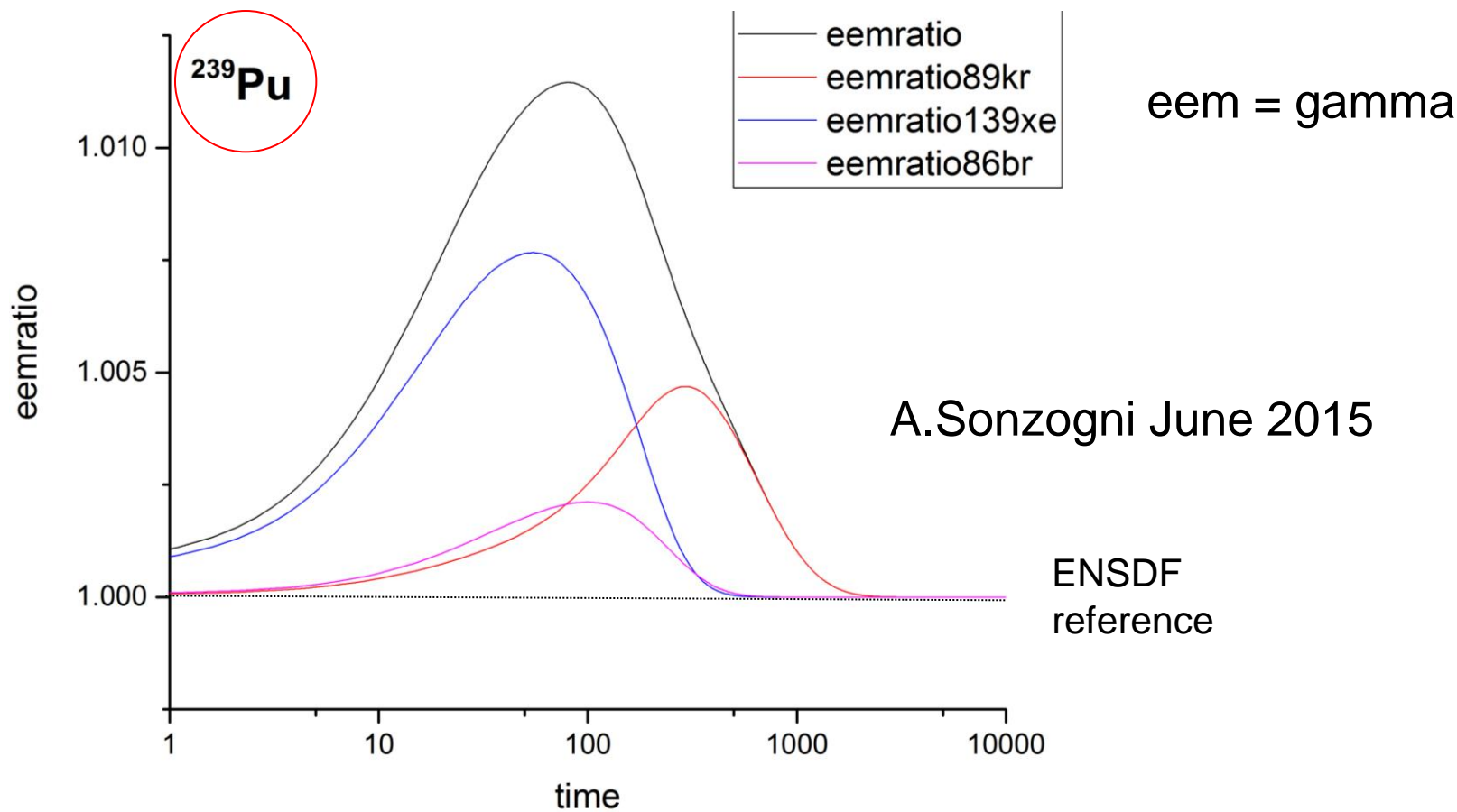
A. Sonzogni June 2015

ENSDF  
reference

For  $^{235}\text{U}$  component of nuclear fuel, three out of 11 priority “1” decays measured with MTAS result in up to 2 % change per time unit in total decay heat between 5 s and 1000 s after fission

# Impact of $^{86}\text{Br}$ , $^{89}\text{Kr}$ and $^{139}\text{Xe}$ on total decay heat pattern

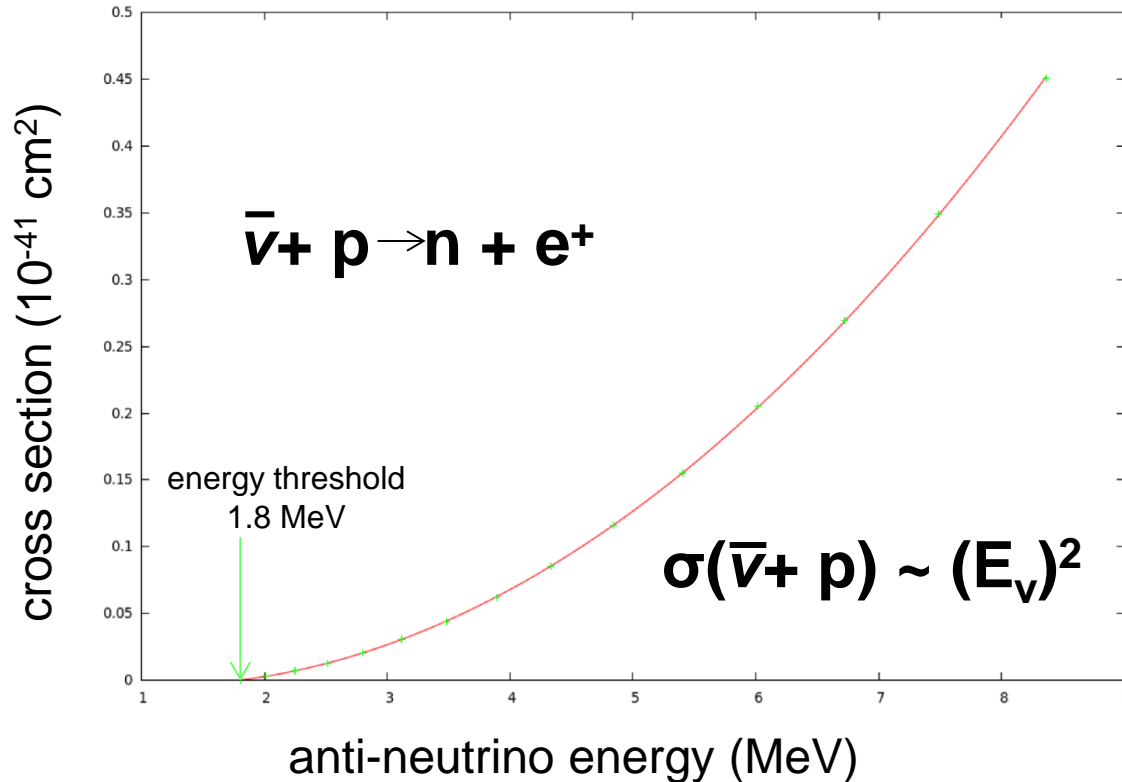
Gamma decay heat ratio: (MTAS and ENSDF)/ENSDF



For  $^{239}\text{Pu}$  component of nuclear fuel, three out of 11 priority “1” decays measured with MTAS result in up to 1 % change per time unit in total decay heat between 5 s and 1000 s



# Anti-neutrino interactions with protons



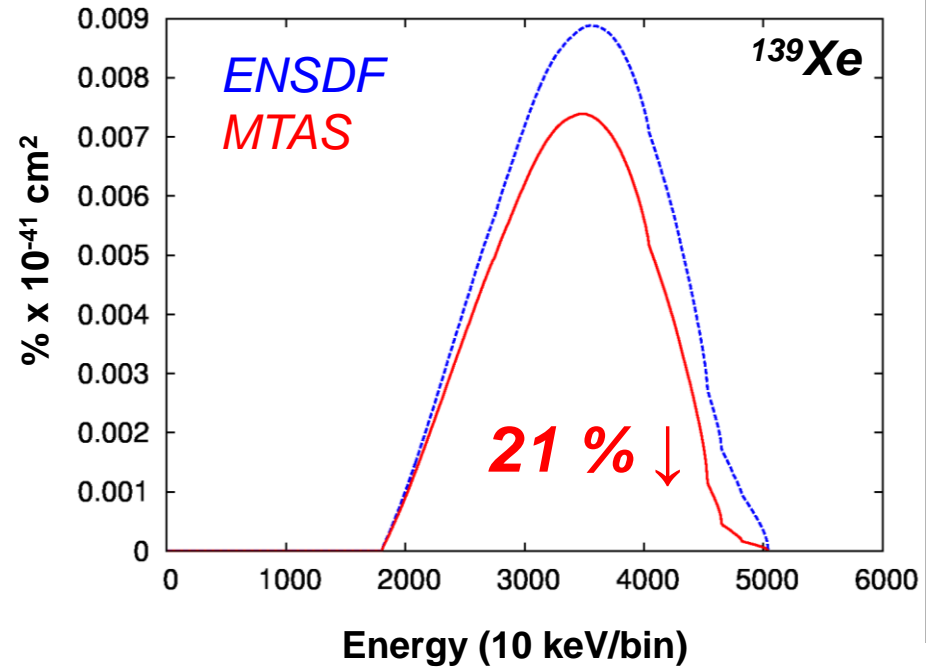
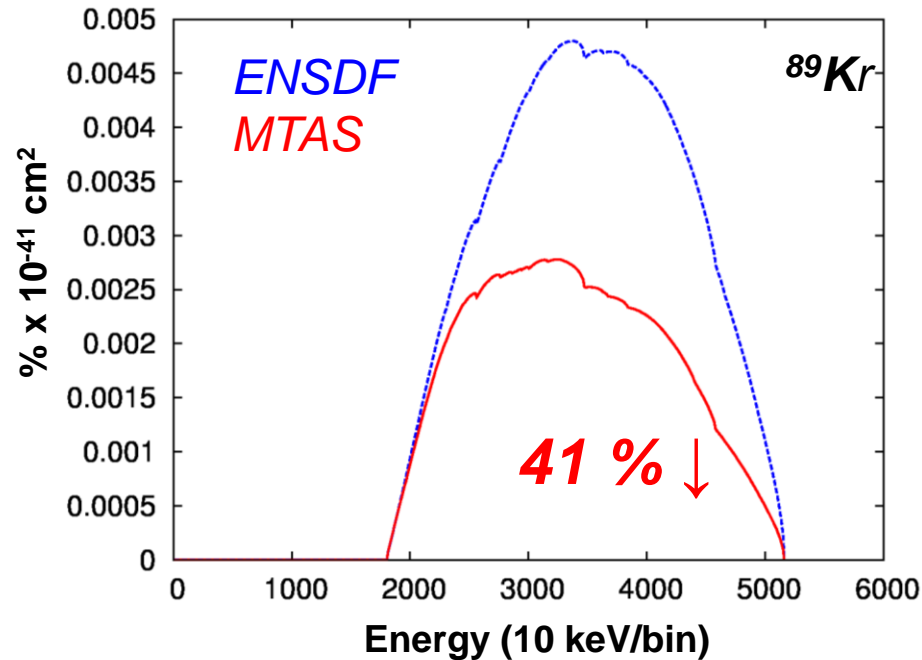
High energy anti-neutrinos have higher cross section for the interactions with protons, i.e., have higher probability to be removed from  $\bar{\nu}$  beam.

For anti-neutrinos with energies below 1.8 MeV energy threshold: no interactions on protons.

**A. Strumia and F. Vissani, PL B 564, 42, 2003**

# Interacting anti-neutrinos emitted in $^{89}\text{Kr}$ and $^{139}\text{Xe}$ decay

Number of interactions of  $^{89}\text{Kr}$  and  $^{139}\text{Xe}$  anti-neutrinos per 10 keV energy bin (in  $\% \times 10^{-41} \text{ cm}^2$  units).



reactor antineutrino anomaly defined as observed/expected signals

$94.3(23) \% \longrightarrow \sim 6 \% \text{ missing}$

Mention et al., PR D 83, 2011

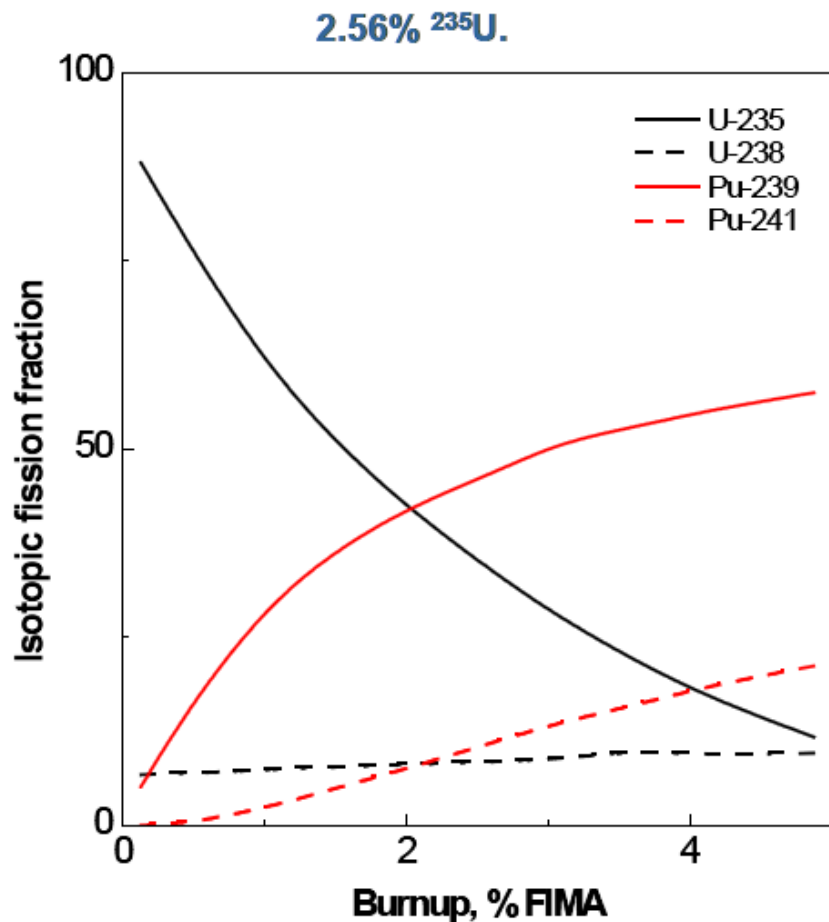
# Impact - reactor anti-neutrino physics

MTAS results for  $^{89}\text{Kr}$  and  $^{139}\text{Xe}$  decays account for

**3 to 2 of the “missing 6 %” difference**

in the reactor anti-neutrino anomaly, depending on a burn-up phase.

Here, a reference number of interacting reactor anti-neutrinos is calculated using ENSDF

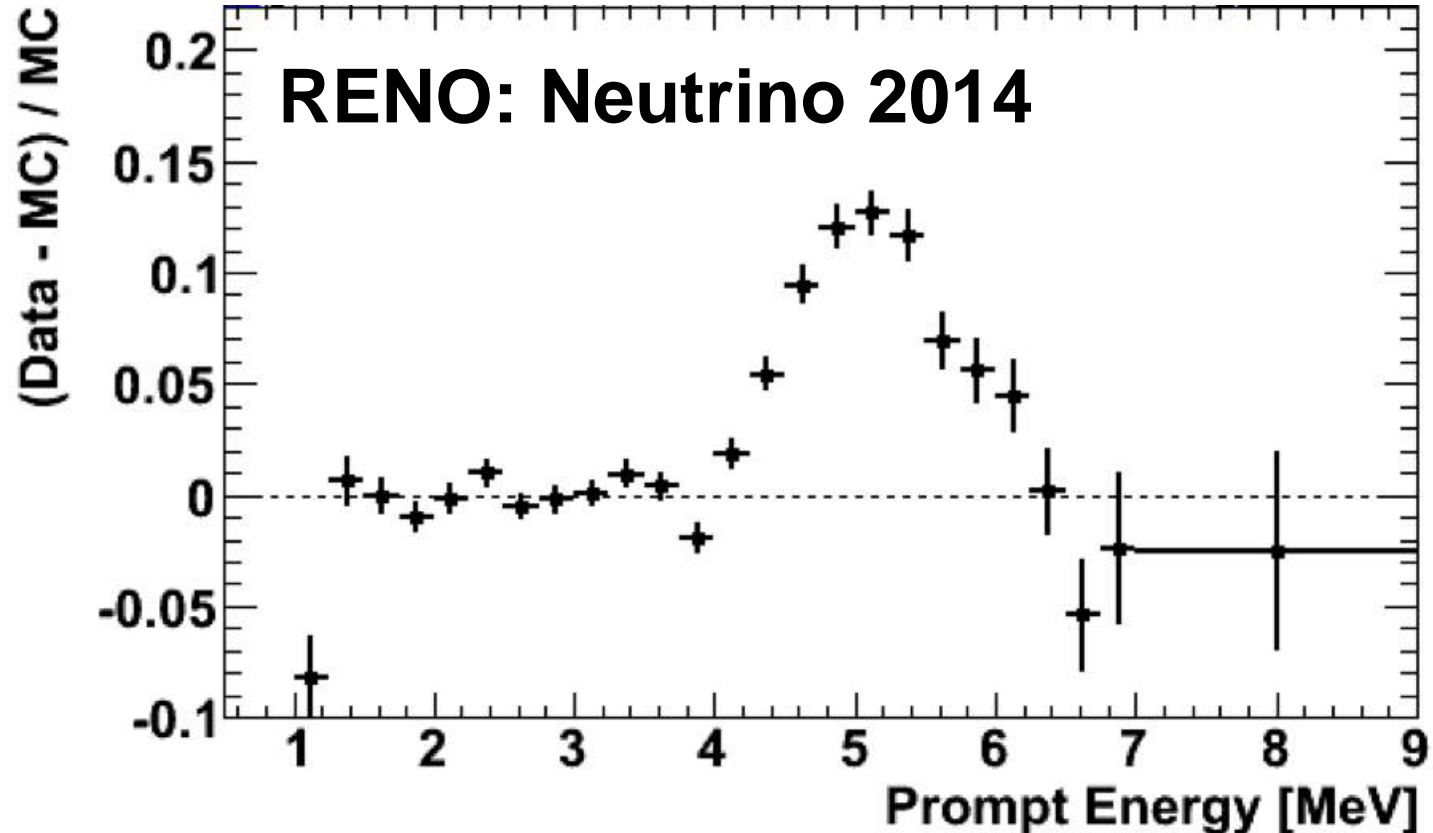


from Ian Gauld, 2014

ORNL Nuclear Reactor Science Group  
(ORIGEN for SCALE )

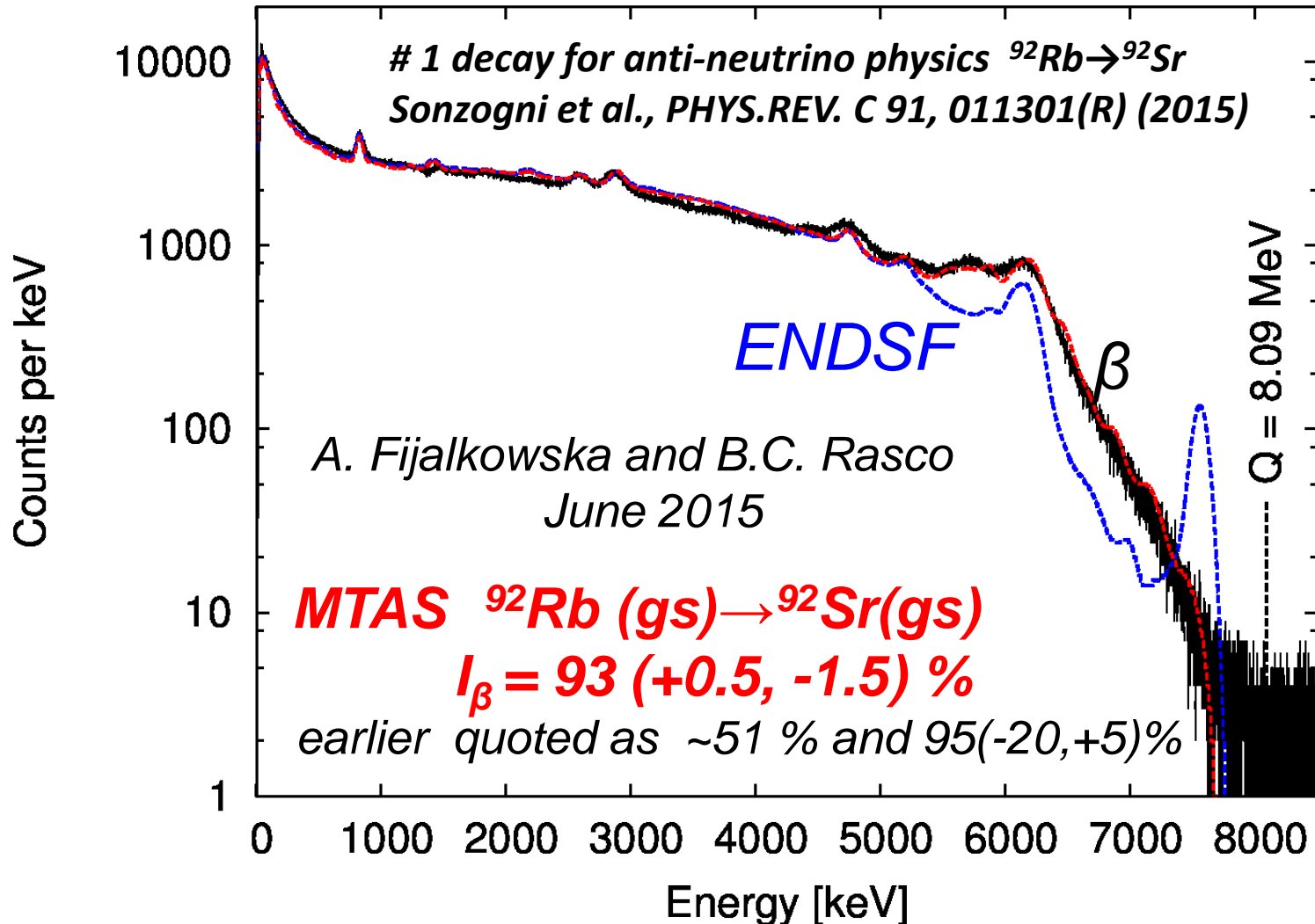
# Recent $\bar{\nu}_e$ energy measurements (RENO, Daya Bay, Double Chooz) also disagree with existing models

(e.g., Dan Dwyer, NDNCA meeting, May 2015)



Anti-neutrino energy = Prompt Energy + 0.8 MeV

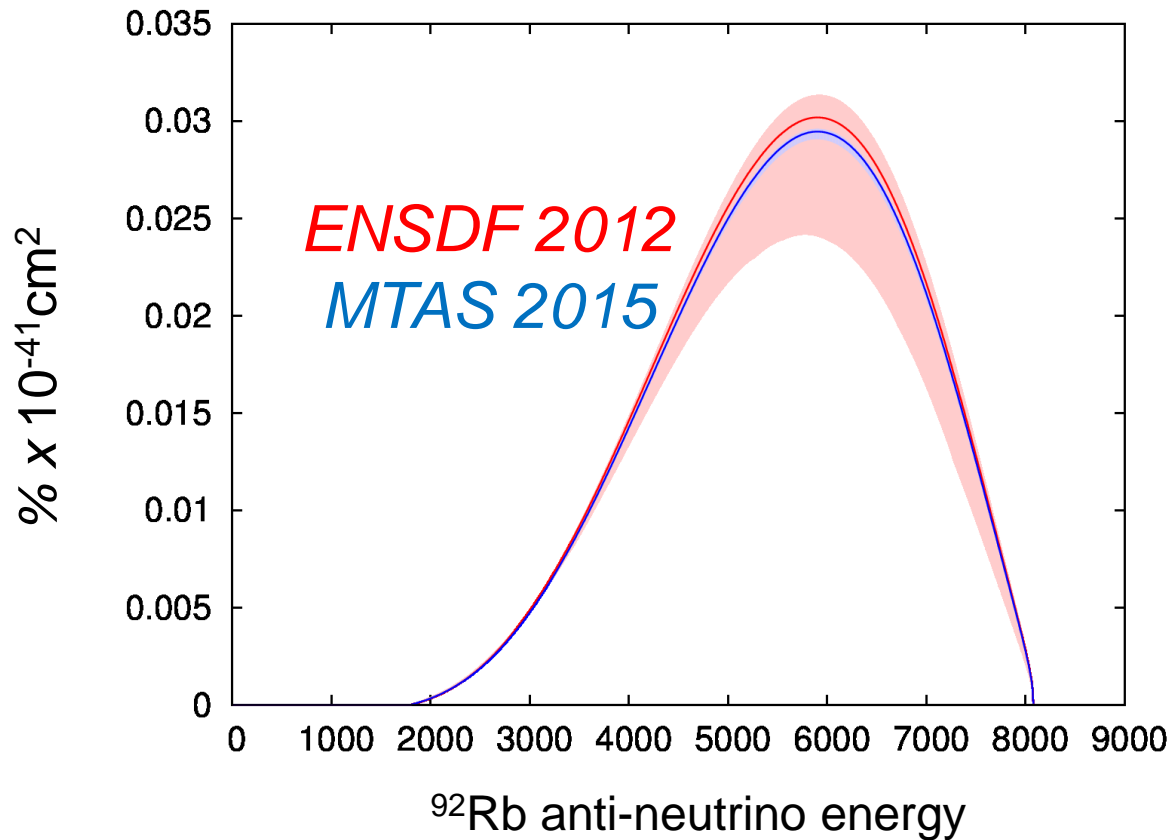
# $^{92}\text{Rb}$ decay measured with Tandem-OLTF-MTAS (March 2015)



$12 \cdot 10^6$  counts in  $\beta$ -gated MTAS total gamma absorption spectrum (counts per keV)

# Impact of MTAS measurement of $^{92}\text{Rb}$ gs-gs branching ratio on a number of anti-neutrinos interacting with matter

Aleksandra Fijałkowska and Charlie B. Rasco, July 2015



ENSDF 1992:	51(2)%	→ interacting antineutrinos	6.95(-0.14,+0.28) % x $10^{-41}\text{cm}^2$
ENSDF 2012:	95.2(-20,+5)%	→ interacting antineutrinos	10.05(-1.77,+0.23) % x $10^{-41}\text{cm}^2$
MTAS 2015:	93(-1.5,+0.5)%	→ interacting anti-neutrinos	9.81(-0.12,+0.04) % x $10^{-41}\text{cm}^2$

# Summary

We got important results on decay properties in fission products, in particular on  $\beta$ -strength function and its consequences:

- decay heat (11 priority “1” cases out of initial list of 27 from NEA 2007)
- reactor anti-neutrino interactions with matter (“anomaly”, as above)
- reactor anti-neutrino energy spectra (“bump”, 12 out of 20+ apparent contributors)
- branching ratios and high energies in  $\beta$ n-emission
- new decay schemes and single-particle level properties (> 20)
- new nuclear structure model analyses (R.K. Grzywacz, I.N. Borzov)