

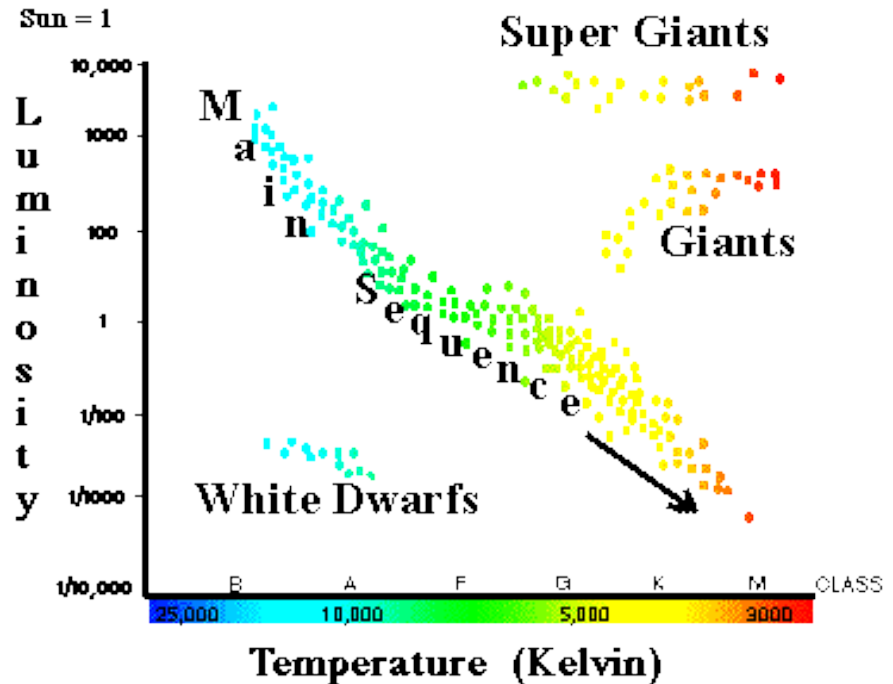
# The life and time of stars

Stellar formation, Stellar evolution, Stellar structure and  
Theoretical stars

# Luminosity, Brightness and H-R diagrams

- ▶ Very useful parameter in measuring and comparing stars.
- ▶ Luminosity/intrinsic luminosity Total power radiated(W)
- ▶ Also apparent brightness,  $b$ , power crossing unit area at the Earth perpendicular to the path of light.

$$b = \frac{L}{4\pi d^2}$$



- ▶ Colour is related to intrinsic luminosity and therefore mass.
- ▶ Standardized brightness of each star.

# Stellar nurseries

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- ▶ Interstellar clouds mostly contain Hydrogen.
- ▶ Higher density regions form clouds, this is where stars form.
- ▶ Much of H is in molecular form.
- ▶ kinetic energy of the gas pressure = potential energy of gravitational force.
- ▶ The cloud undergoes gravitational collapse.



- ▶ The mass is called Jeans Mass.
- ▶ Cloud breaks down become stellar embryo.

# Protostars

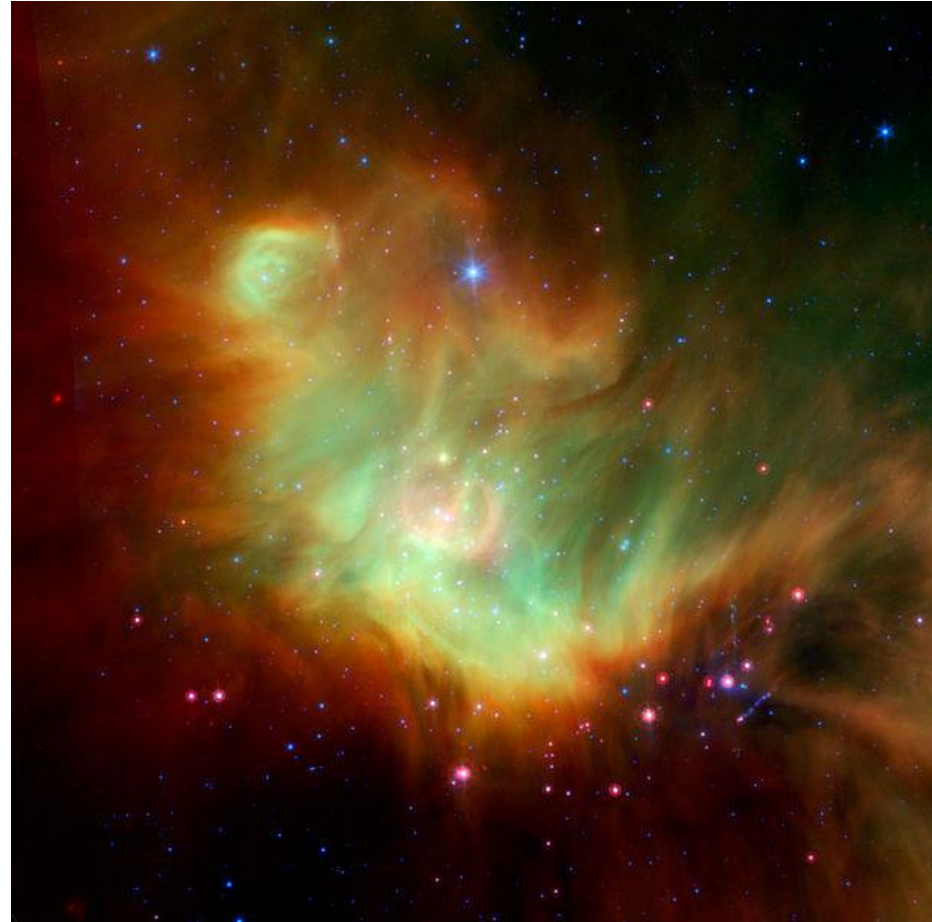
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- ▶ A large mass that forms by contraction out of the gas of a molecular cloud in the interstellar medium.
- ▶ During collapse density of cloud increases towards the centre.
- ▶ At 2000K  $H_2$  dissociates.
- ▶ After a certain density the material is transparent to allow energy to escape.
- ▶ Convection within the star and radiation from the exterior allows for further contraction.
- ▶ This continues until it is hot enough for pressure to support gravitational collapse.

# Observations

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- ▶ Early stages of a stars life cycle can be seen in infrared.
- ▶ Observed in near-IR extinction maps
- ▶ Can only be directly observed in our own galaxy.
- ▶ Distant galaxies, detection is through spectral signature.
- ▶ MWC 349-estimated at 1000 years old
- ▶ VLA 1623- First class 0 Protostars, yet to accumulate majority of mass.



# Role of mass in the life cycle of the star

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- ▶ Different masses of stars form via different mechanisms.
- ▶ Low-mass star formation, is due to gravitational collapse of rotating density enhancements within molecular clouds.
- ▶ Massive stars emit large quantities of radiation which push against infalling material.