

Low  $q^2$   $b \rightarrow sll$  and interplay with radiative

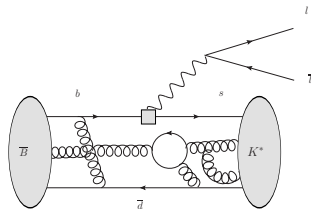
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University of California, San Diego  
Johannes Gutenberg Universität–Mainz

Implications of LHCb measurements and future prospects  
Oct. 15-17th, 2014

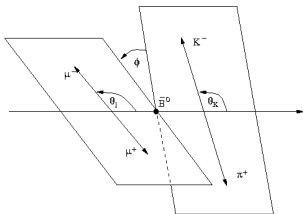


$$\bar{B} \rightarrow \bar{K}^* l^+ l^-$$



Expt.	~# events
CDF	100 <a href="#">PRL106(2011)161801</a>
BaBar	150 <a href="#">PRD86(2012)032012</a>
Belle	200 <a href="#">PRL103(2009)171801</a>
CMS	400 <a href="#">PLB727(2013)77</a>
ATLAS	500 <a href="#">arXiv:1310.4213</a>
LHCb ( $\mu$ )	1000 (1 fb <sup>-1</sup> ) <a href="#">JHEP 1308 (2013) 131</a>
LHCb (e)	128 ([0.0004, 1] GeV <sup>2</sup> ) <a href="#">M Borsato (LHCb)</a>

### ● 4-body decay

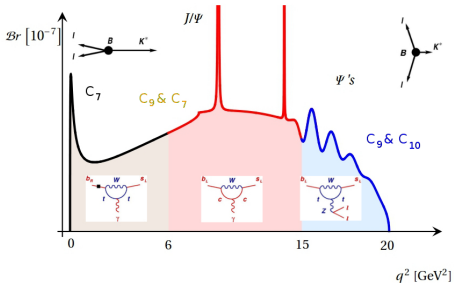


$$\frac{d^{(4)}\Gamma}{dq^2 d(\cos \theta_l) d(\cos \theta_k) d\phi} = \frac{9}{32\pi} (I_1^S \sin^2 \theta_k + I_1^C \cos^2 \theta_k$$

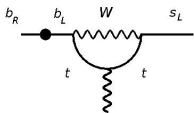
$$+ (I_2^S \sin^2 \theta_k + I_2^C \cos^2 \theta_k) \cos 2\theta_l + I_3 \sin^2 \theta_k \sin^2 \theta_l \cos 2\phi$$

$$+ I_4 \sin 2\theta_k \sin 2\theta_l \cos \phi + I_5 \sin 2\theta_k \sin \theta_l \cos \phi + I_6 \sin^2 \theta_k \cos \theta_l$$

$$+ I_7 \sin 2\theta_k \sin \theta_l \sin \phi + I_8 \sin 2\theta_k \sin 2\theta_l \sin \phi + I_9 \sin^2 \theta_k \sin^2 \theta_l \sin 2\phi)$$



- SM  $\bar{B} \rightarrow \bar{K}^* \gamma$  mostly produces left-handed photons!



$$\mathcal{O}_7 = \frac{e}{4\pi^2} \hat{m}_b \bar{s} \sigma_{\mu\nu} P_R b F^{\mu\nu}$$

### Chirally-flipped operators

$$\mathcal{O}'_7 = \frac{e}{4\pi^2} \hat{m}_b \bar{s} \sigma_{\mu\nu} P_L b F^{\mu\nu}$$

## Connecting theory to experiment: The helicity amplitudes

- Helicity amplitudes  $\lambda = \pm 1, 0$

$$H_V(\lambda) = -iN \left\{ C_9 \tilde{V}_{L\lambda} - \frac{m_B^2}{q^2} \left[ \frac{2\hat{m}_b}{m_B} C_{7\gamma} \tilde{T}_{L\lambda} - 16\pi^2 h_\lambda \right] \right\},$$

$$H_A(\lambda) = -iN C_{10} \tilde{V}_{L\lambda}, \quad H_P = iN \frac{2m_l \hat{m}_b}{q^2} C_{10} \left( \tilde{S}_L + \frac{m_s}{m_b} \tilde{S}_R \right)$$

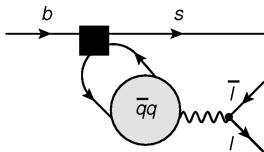
$C_9$  is exposed to various hadronic backgrounds

- Hadronic form factors**

Helicity basis: Best suited for the analysis of power corrections

Bharucha et al. JHEP 1009 (2010) 090, Jäeger and JMC JHEP1305(2013)043

- Non-local contribution**



$$h_\lambda \propto \int d^4y e^{iq \cdot y} \langle \bar{K}^* | J^{\text{em, had}, \mu}(y) \mathcal{H}^{\text{had}}(0) | \bar{B} \rangle \epsilon_\mu^*$$

Especially sensitive to  $c\bar{c}$  contributions!

## Form Factors at large recoil

- **Heavy-quark** and **large-recoil** ( $K^*$ ) limit only **2** independent “soft form factors”

$$T_+ = V_+ = 0, \quad T_- = V_- = \frac{2E}{m_B} \xi_\perp, \quad T_0 = V_0 = S = \frac{E}{m_{K^*}} \xi_\parallel$$

Dugan *et al.* PLB255(1991)583, Charles *et al.* PRD60(1999)014001

- The observable  $P'_5$  Matias *et al.*'12

$$P'_5 = \frac{I_5}{2\sqrt{-I_{2s}I_{2c}}} = \frac{\text{Re}[(H_V^- - H_V^+)H_A^{0*} + (H_A^- - H_A^+)H_V^{0*}]}{\sqrt{(|H_V^0|^2 + |H_A^0|^2)(|H_V^+|^2 + |H_V^-|^2 + |H_A^+|^2 + |H_A^-|^2)}}$$

- *Rationale* behind  $P'$  basis: Ignore in first app.  $\alpha_s$  corrections and  $h_\lambda$

$$H_V^0 \propto \xi_\parallel, \quad H_V^- \propto \xi_\perp, \quad H_V^+ \sim 0$$

$$P'_5 \simeq \frac{C_{10} (C_{9,\perp} + C_{9,\parallel})}{\sqrt{(C_{9,\parallel}^2 + C_{10}^2)(C_{9,\perp}^2 + C_{10}^2)}}, \quad \begin{cases} C_{9,\perp} = C_9^{\text{eff}}(q^2) + \frac{2m_b m_B}{q^2} C_7^{\text{eff}} \\ C_{9,\parallel} = C_9^{\text{eff}}(q^2) + \frac{2m_b E}{q^2} C_7^{\text{eff}} \end{cases}$$

$P'_5$  “hadronic independent” at  $\mathcal{O}(\alpha_s^0, (\frac{\Lambda}{E})^0)$

- $\alpha_s$  corrections can be computed to any order in QCDf or SCET

Beneke *et al.* NPB592(2001)3, NPB612(2001)25, NPB685(2004)249, Bauer *et al.* PRD63(2001)114020

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$P_i^{(\prime)}$  are sensitive to power-corrections!

- Model-independent parameterization (10% p.c.'s)
- Constrained by **exact relations** or **experimental data**

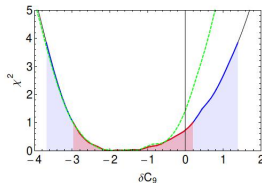
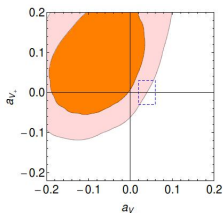
# The significance of the $P_5'$ anomaly

LHCb PRL'13, Descotes-Genon *et al.*'13, Altmannshofer *et al.*...

- The anomaly could be *partly* accommodated in the SM with p.c.'s

$$H_V^- \sim \left\{ C_9(V_-^{\text{QCDf}} + a_{V_-}) - \frac{m_B^2}{q^2} \left[ \frac{2 \hat{m}_b}{m_B} C_{7\gamma} T_-^{\text{QCD}} - 16\pi^2 h_- \right] \right\}$$

- We fit all the  $P_i'$  observables in the bin  $[1,6] \text{ GeV}^2$  using the “**R-Fit method**”  
Profile likelihood over QCD parameters PRELIMINARY Jäger and JMC, to appear



- Similar conclusions were drawn from a bayesian analysis

Beaujean *et al.* arXiv:1310.2478, JHEP1208(2012)030

- ▶ Sizable power corrections (scale-factor method)

- LCSR form factor results (Ball *et al.*'05) suggest values for p.c. parameters that imply a higher significance (blue box)

## $P_1$ and $P_3^{CP}$ and sensitivity to $C_7'$

- At very low  $q^2$  amplitude dominated by  $H_V(\pm)$  (photon pole)

$$H_V(\pm) \simeq iN \frac{2 m_b m_B}{q^2} (C_{7\gamma} T_{\pm} - C'_{7\gamma} T_{\mp}) - 16\pi^2 h_{\pm}$$

- $H_{A,V}^+$  suppressed at factorisable level Burdman et al.'01

• **Exact relation:**  $T_+(0) = 0 \implies T_+(q^2) \sim \mathcal{O}(\Lambda/E) \times q^2/m_b^2$

• **Helicity hierarchy:** Contributions to  $h_+$  are  $(\Lambda/m_B)^2$  Jäger and JMC'12

- Super-clean observables:

$$P_1 \equiv A_T^{(2)} = 2 \frac{l_3 + \bar{l}_3}{\Gamma_T}, \quad P_3^{CP} = -2 \frac{l_9 - \bar{l}_9}{\Gamma_T}$$

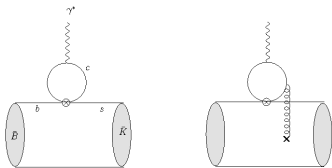
$$l_3 = \frac{-2 \operatorname{Re}(H_V^+ H_V^{-*} + H_A^+ H_A^{-*})}{|H_V^+|^2 + |H_V^-|^2 + |H_A^+|^2 + |H_A^-|^2} \simeq \frac{2 \operatorname{Re}(C_7 C_7')}{|C_7|^2 + |C_7'|^2} \quad l_9 \simeq \frac{2 \operatorname{Im}(C_7 C_7')}{|C_7|^2 + |C_7'|^2}$$

- ▶ Null tests of the SM with very good accuracy at very low  $q^2$  Jäger and JMC'12
- ▶ Sensitive to right-handed currents BSM Melikhov et al.'98, Kruger et al.'05, Becirevic et al.'11
- Only the  $CP$ -combinations  $l_9 - \bar{l}_9$  and  $l_3 + \bar{l}_3$  are sensitive to New Physics!

**Measure  $P_3^{CP}$  !!**

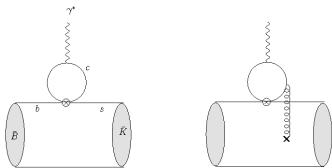


## Double power suppression of $h_+$ at low $q^2$



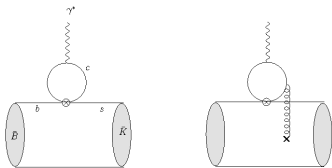
- **QCDF:** Can be computed at leading-power perturbatively in  $\alpha_s$   
One finds  $h_{+|c\bar{c},\text{QCDSF}} = 0$ . At subleading powers factorization breaks down
- Long-distance in **light-cone OPE + SRs:** As large as  $\Lambda/m_b \sim 10\%$  *Khodjamirian et al.'10*
- Detailed investigation shows  $h_{+|c\bar{c},\text{LD}} \sim \mathcal{O}(\frac{\Lambda}{m_b}) h_{-|c\bar{c},\text{LD}}$ !! *Jäger and JMC'12*
- Power corrections from light quarks double-CKM suppressed but “resonate”
  - ▶ **Binned contribution** is very small *Jäger and JMC'12*
  - ▶ **Helicity suppression:**  $h_{+|lq,\text{LD}} \sim \mathcal{O}(\Lambda/m_B)^2$  *Kagan '04*
- Contributions of  $\mathcal{O}_8$  to  $h_\lambda$  estimated to be very small *Dimou et al.'12, Khodjamirian '12*

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## The electronic mode

- The electronic modes have the end-point very close to  $q^2 = 0$

- ▶ Theoretically cleaner: Boosted  $BR$  at low  $q^2$

$$BR_{[0.0004, 1]}^e = 31_{-11}^{+15} 10^{-8}, \quad BR_{[0.1, 0.98]}^\mu = 9.5_{-3.5}^{+5.2} 10^{-8}$$

- ▶ Enhanced sensitivity to  $C_7^{(\prime)}$

- **Experimentally?**  $m_\ell = 0$  is a very good approximation

- Predictions for the electronic mode:

$F_L$	$P_1$	$P_2$	$P_3^{\text{CP}} [10^{-4}]$
$0.07_{-0.04}^{+0.09}$	$0.030_{-0.041}^{+0.045}$	$-0.054_{-0.012}^{+0.015}$	$0.1_{-0.6}^{+0.7}$
$P_4'$	$P_5'$	$P_6'$	$P_8'$
$0.18_{-0.07}^{+0.06}$	$0.52_{-0.11}^{+0.10}$	$0.05_{-0.06}^{+0.07}$	$0.01_{-0.08}^{+0.08}$

PRELIMINARY Jäger and JMC, to appear

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- ▶ Enhanced sensitivity to  $C_7^{(\prime)}$
- **Experimentally?**  $m_\ell = 0$  is a very good approximation
- Errors in  $P_1$  and  $P_3^{CP}$  will be statistically dominated!

	QCDF	Fact. p.c.'s	Non-fact. p.c.'s
$P_1$	0.008	0.009	0.028
$P_3^{CP} [10^{-4}]$	0.3	0.1	0.3

PRELIMINARY Jäger and JMC, to appear

## Interplay with the radiative decays

- The radiative modes are sensitive to  $C'_7$

E. Kou, previous talk

$$\begin{aligned} \mathcal{A}(\bar{B} \rightarrow V(\lambda)\gamma(\lambda)) &= \lim_{q^2 \rightarrow 0} \frac{q^2}{e} H_V(q^2 = 0; \lambda) \\ &= \frac{iNm_B^2}{e} \left[ \frac{2\hat{m}_b}{m_B} (C_{7\gamma} T_\lambda(0) - C'_{7\gamma} T_{-\lambda})(0) - 16\pi^2 h_\lambda(q^2 = 0) \right] \end{aligned}$$

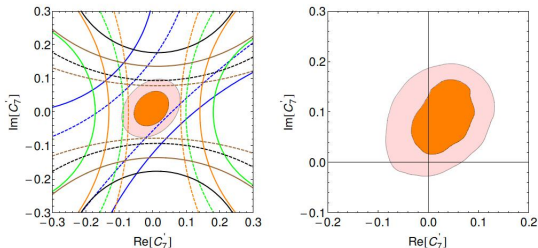
- $BR$  depends on a form factor  $T_1(0)$
- $BR$  has real photons in final state: Sensitivity goes as  $|C'_{7\gamma}|^2$  (also with  $B \rightarrow X_s\gamma$ )
- **Interference and clean observable:** Time-dependent  $CP$  asymmetries

$$S_{K^*\gamma} = \frac{2 \operatorname{Im} [e^{-2i\beta} (H_V^- (\bar{H}_V^-)^* + H_V^+ (\bar{H}_V^+)^*)]}{|H_V^+|^2 + |H_V^-|^2 + |\bar{H}_V^+|^2 + |\bar{H}_V^-|^2}$$

Atwood *et al.* '97, Ball *et al.* '06

- We predict  $S_{K^*\gamma} = -0.02^{+0.016}_{-0.023} 10^{-2}$  Jäger and JMC, to appear while  $S_{K^*\gamma}^{\text{HFAG}} = -0.16 \pm 0.22$

## Prospects and current bounds on $C_7'$



PRELIMINARY Jäger and JMC, to appear

$$S \simeq \frac{2\text{Im}(e^{-2i\beta} C_7 C_7')}{|C_7|^2 + |C_7'|^2}, \quad P_1 \simeq \frac{2\text{Re}(C_7 C_7')}{|C_7|^2 + |C_7'|^2}, \quad P_3^{\text{CP}} \simeq \frac{2\text{Im}(C_7 C_7')}{|C_7|^2 + |C_7'|^2}$$

- **Left:** Ideal plot assuming  $\sigma_{P_i} = 0.25$  for  $\mu$  and  $e$  modes ( $1$  and  $2\sigma$ )
- **Right:** Profile likelihood to current data (slight tension driven by  $A_9^\mu$ !)
- $B \rightarrow K * \{mumu, ee\}$  provide excellent theoretically clean window on  $C_7'$
- With radiative they form a complete system to determine  $C_7$  and  $C_7'$

# Conclusions

- **Current status**
- Observables in the semileptonic decays generally exhibit a sensitivity to  $\Lambda/m_b$

At low  $q^2$   $P_1$  and  $P_3^{CP}$  provide (theoretical) sensitivity to  $C_7'$  as small as  $\mathcal{O}(\Lambda/m_b)^2 \sim 0.01$

**Implications:**  $B \rightarrow K^* \ell \ell$  and radiative decays provide a complete set of clean observables to constrain  $C_7$  and  $C_7'$

Jäger and JMC, to appear

- **Future and prospects**
- New  $e$  and  $\mu$  data eagerly waited for:
  - Frequentist approach including QCD parameters in  $R$ -fit scheme
    - ▶ Improved bounds in the  $C_7^{(\prime)}$  planes
    - ▶ Finer binning could allow to overconstrain parameter space (WCs+QCD)
    - ▶ Assessment of the possible low  $q^2$  anomalies



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## Obtention of error bands and comparison

- A standard method for modelling power corrections

Egede *et al.* JHEP 1010 (2010) 056

Introduce a scale factor  $\zeta_i$  per amplitude, e.g.

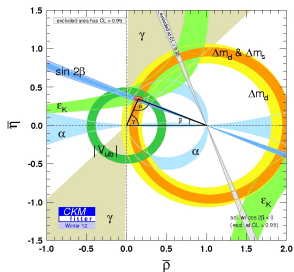
$$H_V(\lambda) \mapsto \zeta_{i,\lambda} \left\{ C_9 \tilde{V}_{L\lambda} - \frac{m_B^2}{q^2} \left[ \frac{2 \hat{m}_b}{m_B} C_{7\gamma} \tilde{T}_{L\lambda} - 16\pi^2 h_\lambda \right] \right\}$$

- Run a Montecarlo over  $\zeta_i$  and other uncertainties and quote 67% interval (th. 1- $\sigma$ )
- Add  $\sigma_{\text{th}}$  and  $\sigma_{\text{expt}}$  in quadratures and perform conventional  $\chi^2$  analysis

### Two possible issues

- 1  $\zeta_i$  can miss interference between power corrections in FFs or  $h_\lambda$
- 2 Is the treatment of theoretical error as experimental adequate?

## Obtention of error bands and comparison



We use the *Rfit* method

Method employed by **CKMfitter** for treating hadronic uncertainties

Höcker *et al.* EPJC21(2001)225

$$\chi^2(\vec{y}_{ew}, \vec{y}_{QCD}) = \left( \frac{x_{\text{exp},i} - x_{\text{th},i}(\vec{y}_{ew}, \vec{y}_{QCD})}{\sigma_{\text{exp}}} \right)^2, \quad \text{if } y_{QCD,i} \in [\bar{y}_i - \sigma_i, \bar{y}_i + \sigma_i] \quad \forall i$$

$$\chi^2(\vec{y}_{ew}, \vec{y}_{QCD}) = \infty, \quad \text{otherwise}$$

- Minimize  $\chi^2$  scanning  $\vec{y}_{QCD}$  by Montecarlo using flat PDFs
- **Our error intervals:** maximum spread of results resulting from Montecarlos

## Form factors and PCs

$$F(q^2) = F^\infty(q^2) + a_F + b_F q^2/m_B^2 + \mathcal{O}([q^2/m_B^2]^2). \quad (1)$$

Reference form factors ( $V_0$  and  $V_-$  could be used in their stead):

$$\xi_\perp(q^2) \equiv T_1(q^2) \simeq \frac{m_B}{2E} T_-(q^2), \quad \xi_\parallel(q^2) \equiv S(q^2),$$

$$\xi_\perp(0) = T_-(0) = 0.31(4), \quad \xi_\parallel(0) = 0.31(6),$$

Modified HQ/LE scalings:

$$\xi_X(q^2) = \xi_X(0) \left( \frac{1}{1 - q^2/m_B^2} \right)^{2+\alpha_X}, \quad X = \perp, \parallel.$$

$$|\alpha_\perp^{\max}| = 0.2, \quad |\alpha_\parallel^{\max}| = 0.7.$$

Model independent constraints:

$$a_{T_+} = 0, \quad a_{V_0} = a_S.$$

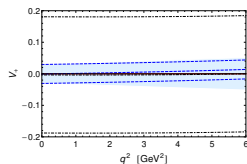
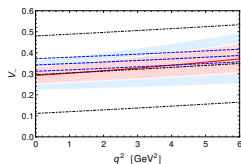
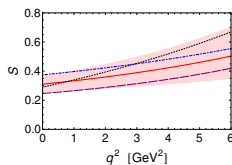
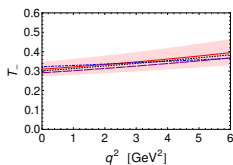
$a_{T_-} = a_S = 0$  and  $b_{T_-} = b_S = 0$  effectively absorbed in  $\xi_{\perp,\parallel}(0)$  and  $\alpha_{\perp,\parallel}$

## Form factors and PCs (cont ...)

Power corrections estimated bt power counting  $\Lambda/m_b \sim 10\%$

$$|a_F^{\text{max,pc}}| \simeq 0.03,$$

$$|b_F^{\text{max,pc}}| \simeq 0.10.$$



**Table :** Results for the bin  $[1, 6] \text{ GeV}^2$  in the SM for a selection of observables and using different schemes for the estimation of the theoretical uncertainties. In the last column we show the experimental data.

	Max. Spread	Max. spread ( $V_-$ and $V_0$ )	$1\sigma$ gaussian	Expt.
$P_1$	$-0.02^{+0.21}_{-0.23}$	$-0.02^{+0.20}_{-0.21}$	$-0.02^{+0.11}_{-0.11}$	0.15(0.4)
$P_2$	$-0.20^{+0.44}_{-0.35}$	$-0.19^{+0.37}_{-0.32}$	$-0.18^{+0.19}_{-0.18}$	-0.66(23)
$P'_5$	$-0.27^{+0.48}_{-0.38}$	$-0.28^{+0.38}_{-0.36}$	$-0.27^{+0.20}_{-0.18}$	0.21(21)