

Chiral Magnetic Effect in QED induced by longitudinal photons.

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Introduction.

Objectives and
Motivation.

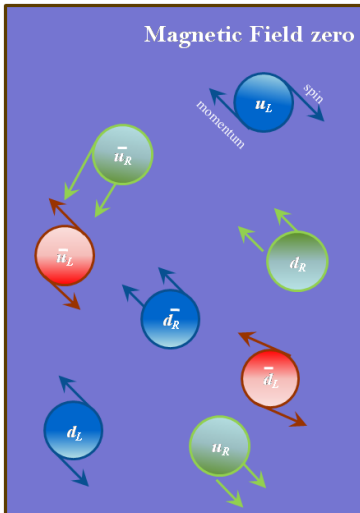
Axial anomaly
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in QED in a
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Chiral current
generation in QED.

Chiral conductivity
in non-static limit.

Axial anomaly in
presence of mass
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Conclusions.



- No magnetic field: No polarization.

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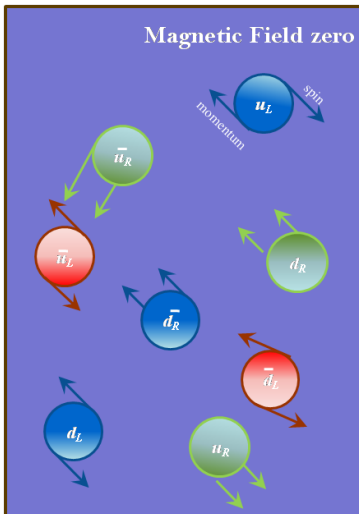
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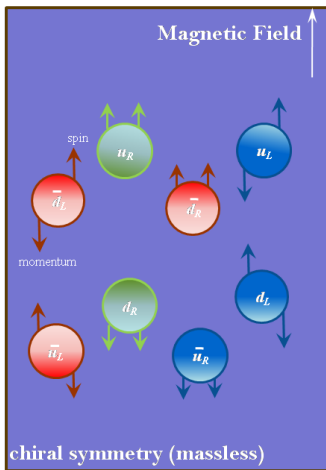
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- A magnetic field will align the spins, depending on their electric charge.
- The momenta of the quarks align along the magnetic field. A quark with right-handed helicity will have momentum opposite to a left-handed one.
- !!!In this way the magnetic field can distinguish between right and left!!!

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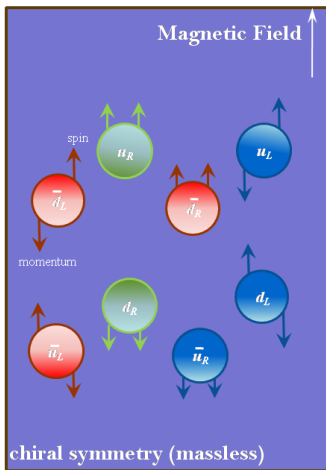
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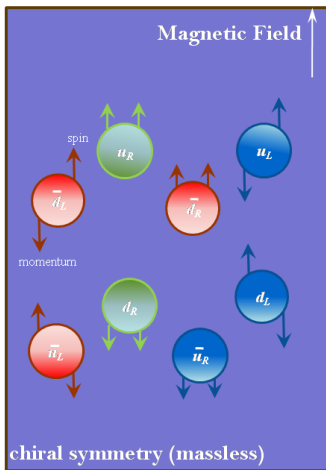
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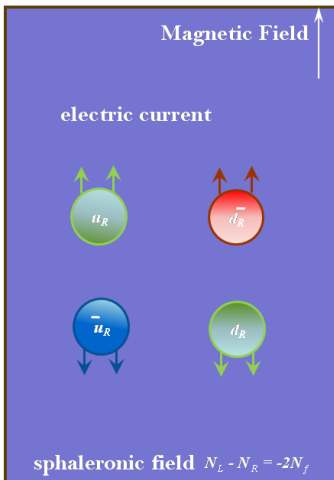
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- What does a magnetic field do with chirality (generated by topological charge)?
- Positively charged particles move parallel to the magnetic field.
- Negatively charged particles move antiparallel to the magnetic field.
- !!An electric current is created along the magnetic field!! \Leftrightarrow The Chiral Magnetic Effect in QCD.

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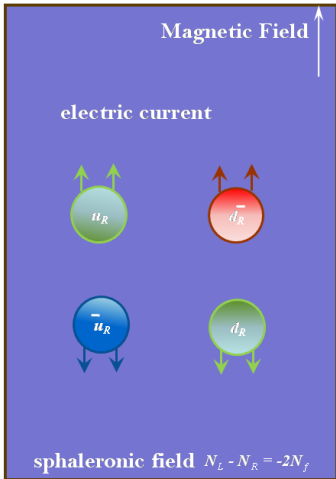
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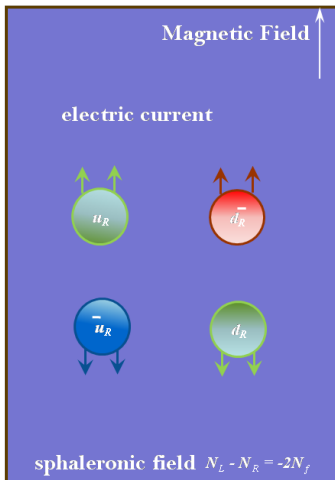
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Objectives.

- Demonstrate the existence of the chiral magnetic effect in an electron-positron magnetized gas in QED. *A pseudo-vector Ohm current is induced by the electric field related to the longitudinal photons in QED propagating parallel to the external magnetic field \mathbf{B} and separating opposite charges of the same helicity.*
- Obtain an anomaly relation in a medium of massive particles in presence of an external magnetic field, that bears some analogy to the Adler-Bell-Jackiw. *The effect is interesting in connection to the QCD chiral magnetic case reported in current literature.*

Motivation.

- Study the process associated to the chiral symmetry in a charged medium in presence of an external field \mathbf{B} .



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Motivation.

- Study the process associated to the chiral symmetry in a charged medium in presence of an external field \mathbf{B} .

The equation of Schwinger-Dyson for the photon:

$$[k^2 g_{\mu\nu} - \Pi_{\mu\nu}(k|A_\mu^{ext})]a^\nu(k) = 0$$

- a_μ (*radiation field*) is a small perturbation added to A_μ^{ext} (*external field*)
- the total external electromagnetic field is $A_\mu^{ext} + a_\mu$ (*the electric field of the wave $E \ll B$*).

The second term takes into account the interaction of the radiation with the external field through the virtual pairs (e^+ , e^-) and the excitation of particles in electron-positron plasma.

The quantum corrections are given for the tensor of polarization $\Pi_{\mu\nu}$.



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In the external constant field B , electrons and positrons move in bound states characterized by energy levels given by:

$$\varepsilon_{n_l, p_3} = \sqrt{p_3^2 + m^2 + |e|B(2n_l + 1 - \text{sgn}(e)s_3)}$$

- $s_3 = \pm 1$ are spin eigenvalues along x_3 .
- $n_l = 0, 1, \dots$ are the Landau quantum numbers. These are two-fold spin degenerate, except for the ground state ε_0 in which $n_l = 0$, and for electrons is $s_3 = -1$ and for positrons $s_3 = 1$.
- Quantum states are also degenerate with regard to the orbit's center coordinates.

For a system in equilibrium at temperature T and chemical potential μ .

Net density of charged particles in the ground state (LLL)

$$N_0 = \frac{eB}{2\pi^2} \int_{-\infty}^0 dp_3 (n_R^e - n_L^p) + \int_0^{\infty} dp_3 (n_L^e - n_R^p)$$

Magnetization in the LLL

$$\mathcal{M}_0 = \frac{e}{4\pi^2} \int_{-\infty}^0 \frac{p_3^2 dp_3}{\epsilon_0} (n_R^e + n_L^p) + \int_0^{\infty} \frac{p_3^2 dp_3}{\epsilon_0} (n_L^e + n_R^p)$$

where $n^{e,p} = [1 + e^{(\epsilon_0 \mp \mu)/T}]^{-1}$ (T is given in energy units) are the electron and positron densities.

- !!!!both terms in the sum of the right side are equal!!!! \Rightarrow !!! equal densities of L,R helicities in the state of equilibrium!!!

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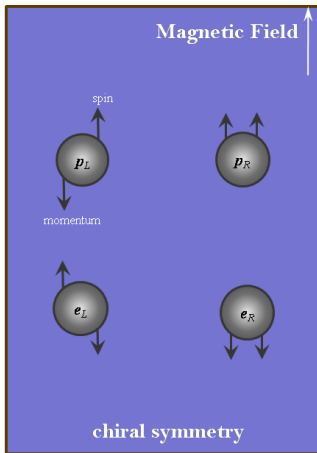
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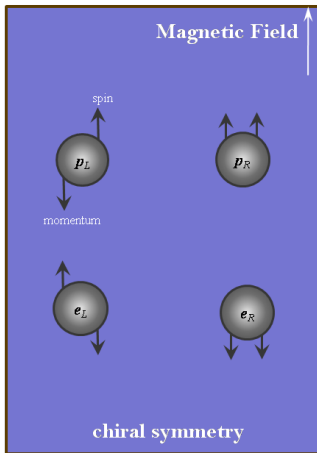
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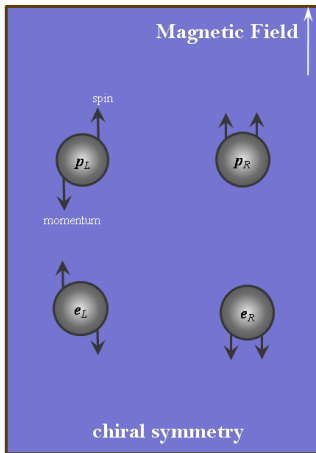
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If we will make here the fundamental assumption that (with $\mu \gtrsim m$): $eB \gg m^2, \mu, T^2 \Rightarrow$ LLL dominant.

The diagonalization of the photon self-energy tensor leads to the equation:

$$\Pi_{\mu\nu} b^{\nu(i)} = \eta_i b_{\mu}^{(i)}$$

having three non-vanishing eigenvalues η_i and three eigenvectors $b_{\mu}^{(i)}$ for $i = 1, 2, 3$, corresponding to three photon propagation modes.

For each mode it is obtained a dispersion law

$$k^2 = \eta_i(k_3, k_{\perp}, \omega, B)$$

$k^2 = k_3^2 + k_{\perp}^2 - \omega^2$, k_3 and k_{\perp} are respectively the components of the photon four-momentum in directions \parallel and \perp to \mathbf{B} , and ω its energy.

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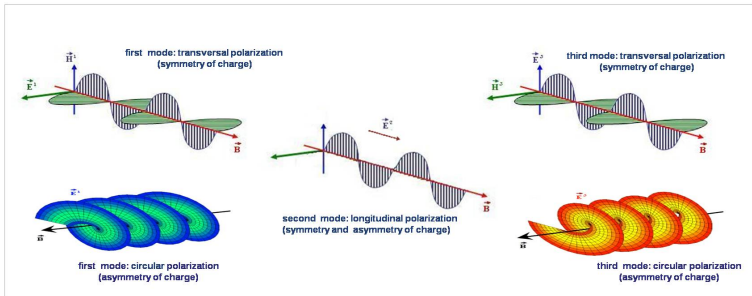
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In a charged medium, for propagation along the field \mathbf{B} , in addition to the two transverse modes, there is a **longitudinally polarized eigenmode** independent of the charge-symmetry.

Propagation parallel to the magnetic field





The longitudinal mode is a pseudo-vector

$$b_{\mu}^{(2)}(k) = R\tilde{F}_{\mu\nu}k^{\nu}$$

- R is an arbitrary constant
- $\tilde{F}_{\mu\nu}$ is the dual of the electromagnetic field tensor $F_{\mu\nu}$

The electric polarization vector associated to longitudinal mode

$$\mathbf{E}^{(2)} = R(k_3^2 - \omega^2)\mathbf{B}$$

!!!! For QED in a charged e^{\pm} medium, the longitudinal mode along \mathbf{B} is not on the light cone, that is $k_3^2 - \omega^2 \neq 0$!!!!

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The electromagnetic current as a function of $A_\mu^{ext} + a_\mu$ depends on the two relativistic invariants:

$$\mathfrak{F} = \frac{1}{4} F^{\mu\nu} F_{\mu\nu} = \frac{1}{2} (B^2 - E^2) \simeq \frac{1}{2} B^2, \quad E \ll B$$

$$\mathfrak{G} = \frac{1}{4} \tilde{F}^{\mu\nu} F_{\mu\nu} = \mathbf{B} \cdot \mathbf{E}$$

Notice that for the case of propagation along \mathbf{B} , the pseudo-scalar $\mathfrak{G} \neq 0$ only for the longitudinal mode.

An expansion in functional series gives:

$$j_\mu(A_\mu^{ext} + a_\mu) = j_\mu(A_\mu^{ext}) + (\delta j_\mu / \delta A_\nu^{ext}) a_\nu + \dots$$

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Its linear term in a_ν is:

$$j_i = \Pi_{i\nu} a_\nu = Y_{ij} E_j, \quad \nu = 1, 2, 3, 4, \quad i, j = 1, 2, 3$$

- $E_j = i(\omega a_j - k_j a_0)$ is the electric field, with $a_4 = ia_0$ and $k_4 = i\omega$.
- $j_\mu(A_\mu^{ext}) = \rho \delta_{\mu 4}$, where ρ is the charge density.
- a_μ is in general a linear combination of the eigenmodes $b_\mu^{(i)}$. Below we particularize to the case in which a_μ is the eigenvector $a_\mu = b_\mu^{(2)}$, for which ($\mathbf{E} \parallel \mathbf{B}$).

The complex conductivity tensor or admittivity is:

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$$j_i = \Pi_{i\nu} b^\nu(2) = s b_i(2)$$

$s = b_\nu^{(2)} \Pi_\rho^\nu b^\rho(2)$ is the eigenvalue of the photon self-energy tensor corresponding to the longitudinal mode.

- !!!!As $b_\nu^{(2)}$ is a pseudo-vector !!!! \Rightarrow !!!For propagation along B the current j_ν is also a pseudo-vector!!!
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The non-conservation of the two-dimensional axial current:

$$k_\mu j_A^\mu = \frac{z_1}{k_4} j_3 \neq 0$$

Here $z_1 = k_3^2 - \omega^2$, we are interested only in the region of real
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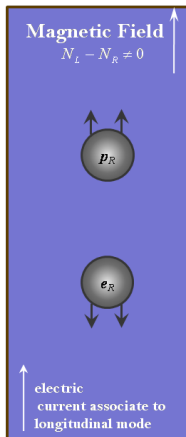
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- The longitudinal mode produces an electric current along \mathbf{B} , in which opposite charged particles, moving in opposite directions, have the same helicity.

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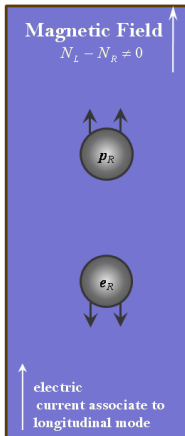
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The current density can be written in the general form as:

$$j_i = \sigma_{ij}^0 E_j + (E \times S)_i$$

$$\sigma_{ij}^0 = \text{Im}[\Pi_{ij}^s]/\omega, \quad S_i = \frac{1}{2} \epsilon^{ijk} \sigma_{jk}^H$$

$$\sigma_{jk}^H = \text{Im}[\Pi_{ij}^A]/\omega$$

- ϵ^{ijk} is the third rank antisymmetric unit tensor
- Π_{ij}^s, Π_{ij}^A are the symmetric and antisymmetric parts of the photon self-energy tensor.

The first term of the current density corresponds to the Ohm (real conductivity) current and the second is the Hall (real conductivity) current.



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Here we will only work with the Ohm current associated to the **longitudinal mode**.

For the specific case of propagation along **B**.

The current density associated to the longitudinal mode can be expressed in the form:

$$j_3 = \sigma_{33}^0 E_3$$
$$\sigma_{33}^0 = \text{Im}[\Pi_{33}]/\omega = -\omega \text{Im}[s]/z_1$$

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Scalar s in the one-loop approximation

$$s = -\frac{e^3 B}{\pi^2} \sum_{n=0}^{\infty} \int_{-\infty}^{\infty} \frac{dp_3}{\varepsilon_q} \left[\frac{\alpha_n \varepsilon_{n,0}^2 (2p_3 k_3 + z_1)}{4z_1 p_3^2 + 4p_3 k_3 z_1 + z_1^2 - 4\omega^2 \varepsilon_{n,0}^2} \right] \cdot [n^p(\varepsilon_q) + n^e(\varepsilon_q) - 1]$$

$$\varepsilon_{n,0} = \sqrt{m^2 + 2eBn}, \quad n = n_l + 1/2 + s_3/2$$

$$\alpha_n = 2 - \delta_{n,0}, \quad q = (n, p_3)$$

The imaginary part of the scalar s

$$\text{Im}[s] = -\frac{e^3 B}{2\pi} \sum_{n=0}^{\infty} \alpha_n \varepsilon_{n,0}^2 S_n$$



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The chiral conductivity at a finite temperature T and density characterized by a chemical potential μ

$$\sigma_{33}^0 = \frac{e^3 B \omega}{2\pi z_1} \sum_{n=0}^{\infty} \alpha_n \varepsilon_{n,0}^2 S_n$$

$$S_n = \{ \Delta N + \theta(-4\varepsilon_{n,0}^2 - z_1) \Delta H \} / \Lambda$$

$$\Lambda = \sqrt{z_1(z_1 + 4\varepsilon_{n,0}^2)}$$

!!!The expression for the chiral conductivity contain both diamagnetic and paramagnetic terms !!!



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$$\Delta N = [N(\varepsilon_r) - N(\varepsilon_r + \omega)]$$

The term ΔN accounts for the **excitation of particles** $[\varepsilon(p_3, n) \rightarrow \varepsilon(p_3 + k_3, n)]$ by increasing their momentum along **B**.

$$\Delta H = [H(-\varepsilon_s) + H(\omega + \varepsilon_s - 2)]$$

The term ΔH accounts for the pair creation (only in the region $z_1 < -4\varepsilon_{n,0}^2$).

$$N = n^e(\varepsilon_r) + n^p(\varepsilon_r), \quad H = n^e(\varepsilon_s) + n^p(\omega - \varepsilon_s)$$

$$\varepsilon_s = (\omega z_1 + |k_3| \Lambda) / 2z_1, \quad \varepsilon_r = (-\omega z_1 + |k_3| \Lambda) / 2z_1, \quad r, s = (n, \omega, k_3)$$



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Taking into account the expression for the four-divergence of the axial current and the equation for the chiral conductivity it is obtained:

Anomaly relation in a medium of massive particles

$$k_{\mu} j_A^{\mu} = a[m\mathbb{A}(m) + \mathbb{C}(m)] \frac{e^2}{2\pi^2} \mathbf{E} \cdot \mathbf{B}$$

$$\mathbb{A}(m) = (2\pi m/e) \sum_{n=0}^{\infty} \alpha_n S_n$$

$$\mathbb{C}(m) = 8\pi B \sum_{n=1}^{\infty} n S_n$$

In the $m \rightarrow 0$ limit, it is obtained: $\mathbb{A}(m \rightarrow 0) \rightarrow 0$ and $\mathbb{C}(m \rightarrow 0) \neq 0$.



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- 2 The chiral conductivity was calculated in the general case of finite temperature and density, and for massive fermions, which might be relevant for astrophysical applications.
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