

General Relativity Theory is a set of non-linear partial differential field equations for the tensor metric field $g_{\mu\nu}$

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$$R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} = \kappa T_{\mu\nu} \quad \text{and with } \Gamma_{\mu\nu}^{\alpha} = \frac{1}{2} g^{\alpha\beta} \left(\frac{\partial g_{\mu\beta}}{\partial x^{\nu}} + \frac{\partial g_{\nu\beta}}{\partial x^{\mu}} - \frac{\partial g_{\mu\nu}}{\partial x^{\beta}} \right)$$

$$R_{\mu\nu} = \frac{\partial^2}{\partial x^{\mu} \partial x^{\nu}} \log(g^{1/2}) - \frac{\partial}{\partial x^{\lambda}} \Gamma_{\mu\nu}^{\lambda} + \Gamma_{\mu\alpha}^{\beta} \Gamma_{\nu\beta}^{\alpha} - \Gamma_{\mu\nu}^{\alpha} \frac{\partial}{\partial x^{\alpha}} \log(g^{1/2}) \quad R = R_{\mu\nu} g^{\mu\nu}$$

Except for a few special cases of Symmetric Sources $T_{\mu\nu}$ solution is usually by iterative perturbation theory to successive non-linear approximation. Finding the general N - Body equations of motion is typically done by similar process.

I develop and describe a different method for finding the N - Body gravitational equations of motion to arbitrary order using

iterative linear algebraic equations

based on enforcing

symmetry and invariance properties of observational gravity.

Properties of Gravity

which yield **iterative algebraic equations** for obtaining the n th order $(1/c^{2n})$ N - body potentials in terms of the previously determined potentials up to the $n - 1$ order :

1. Exterior Effacement. A local gravitational system's Lagrangian is unaffected by distribution of distant spectator matter in the universe (when expressed in local proper coordinates).

2. Interior Effacement. A spherical **composite body** shows just a single mass - energy parameter throughout

its appearances in N - body Lagrangian : $m_A \Rightarrow M_A = \sum_a m_a \left(1 + \frac{v_a^2}{2c^2} - \sum_{a'} \frac{G m_{a'}}{2c^2 r_{aa'}} + \dots \right)$ and same for \vec{J}

3. Time Dilation : A local dynamical gravitational system shows identical dynamics with respect to dilated time variable $d\tau^* = \sqrt{1 - w^2 / c^2} dt$ when all system velocities are boosted by common \vec{w} .

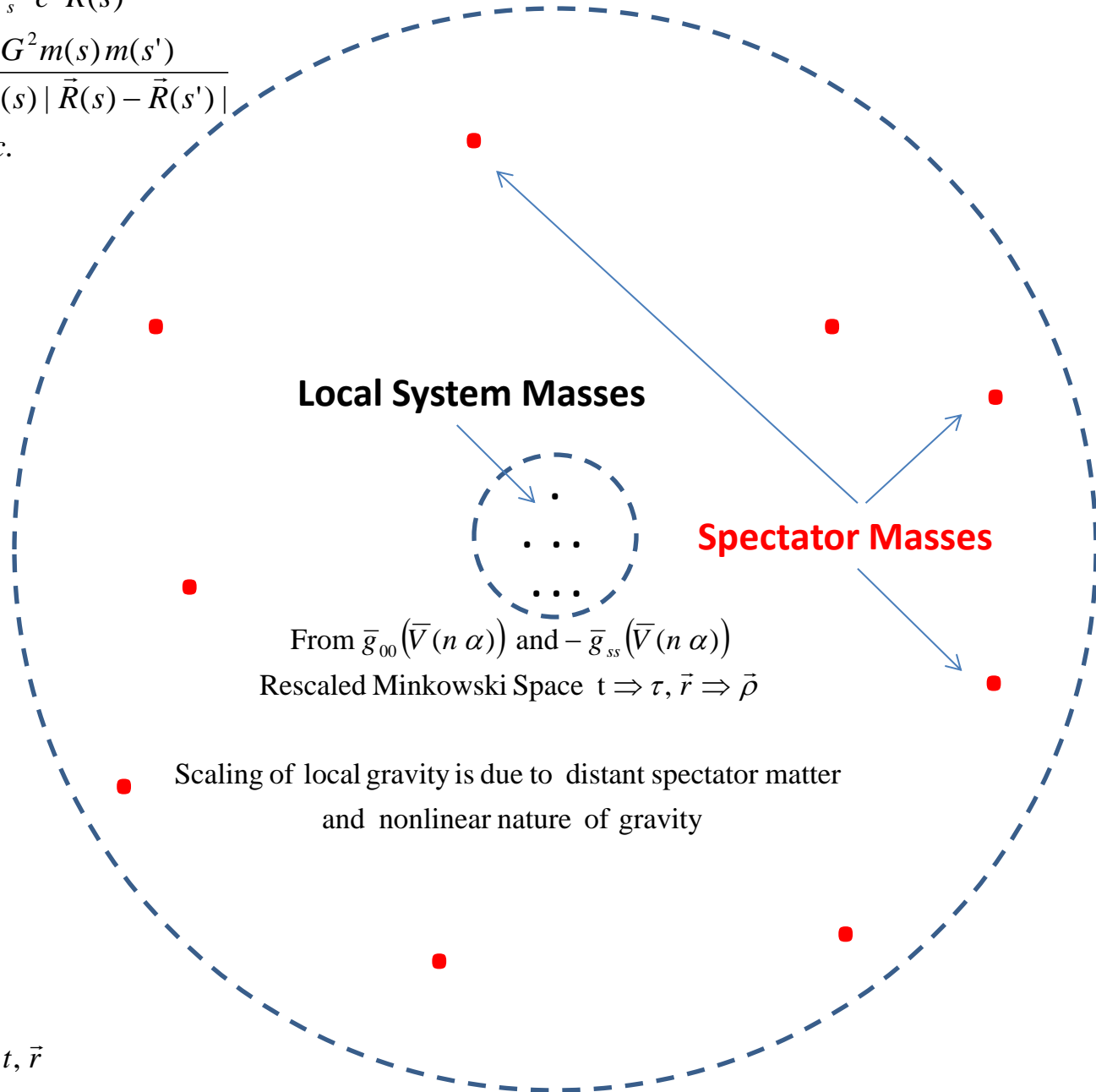
4. Lorentz Contraction. A static configuration of bodies requires that a moving configuration of those bodies must be Lorentz contracted along direction of boost \vec{w} .

Scaling due to Exterior "Spectator" Matter

$$\bar{V}(1,1) = U = \sum_s \frac{G m(s)}{c^2 R(s)}$$

$$\bar{V}(2,2) = W = \sum_{s,s'} \frac{G^2 m(s) m(s')}{c^4 R(s) |\vec{R}(s) - \vec{R}(s')|}$$

....etc.



Minkowski Space t, \vec{r}

Spectator
Bodies

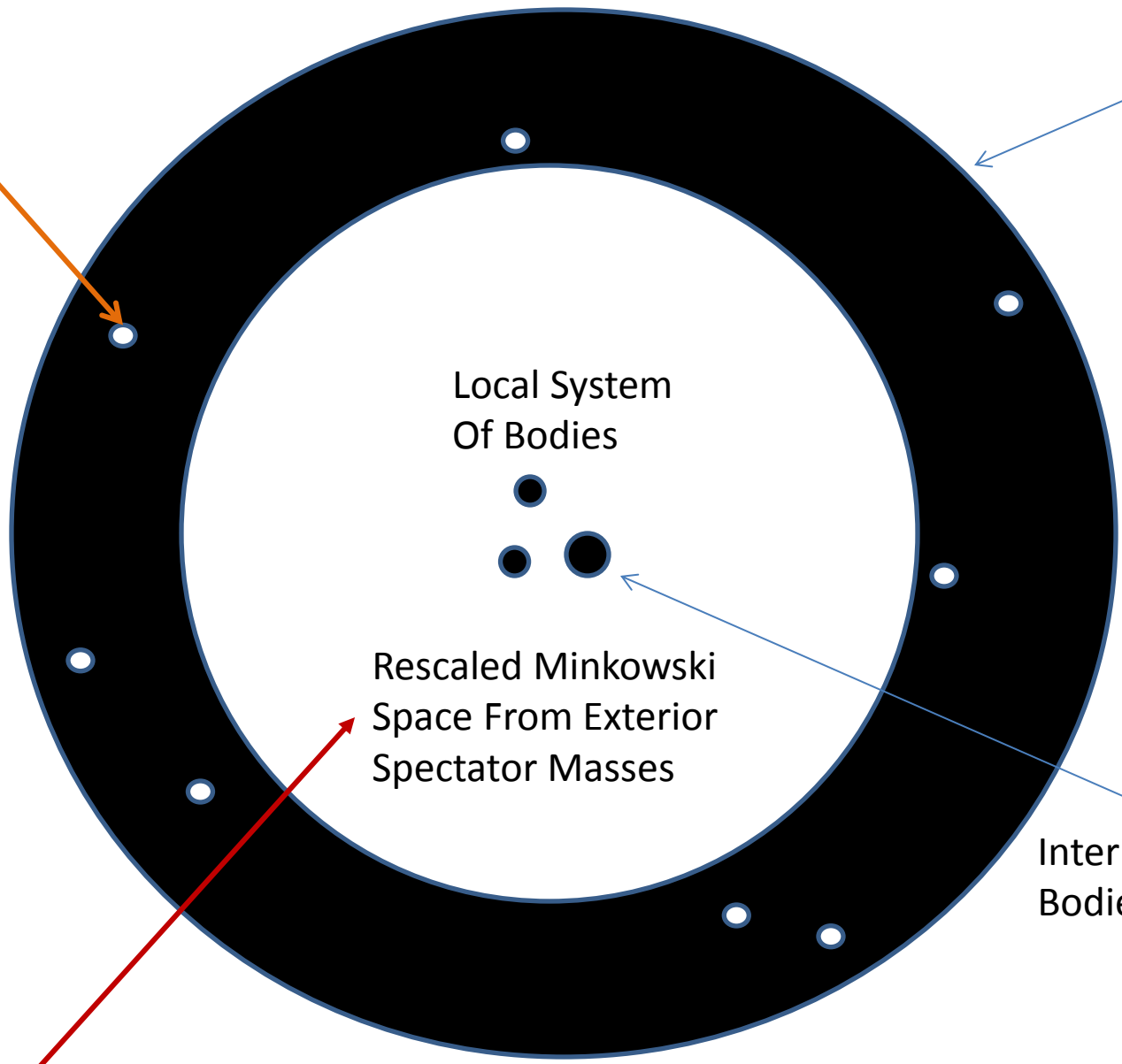
Outside world
Is effaced.

Local System
Of Bodies

Rescaled Minkowski
Space From Exterior
Spectator Masses

Interior of Composite
Bodies are Effaced.

Minkowski Space



Distant **spectator bodies** produce a background spatial metric

$$\bar{g}_{ab} = \bar{g}_{SS} \delta_{ab} + \bar{h}_{ab} \quad a, b = x, y, z$$

g_{SS} is -1 plus the motion - independent set of potentials, while h_{ab} is the motion - dependent and non - diagonal portion of the spacial metric.

$$\text{So } \rho^2 = -\bar{g}_{ab} r^a r^b = -\bar{g}_{SS} r^2 + \bar{h}_{ab} r^a r^b$$

$$\frac{1}{r} = \frac{\sqrt{-\bar{g}_{SS}}}{\rho} \left(1 - \frac{h_{ab} r^a r^b}{\rho^2} \right)^{-1/2}$$

$$\text{and } \frac{1}{r^n} = \frac{(-\bar{g}_{SS})^{n/2}}{\rho^n} + \text{motion - dependent potential additions}$$

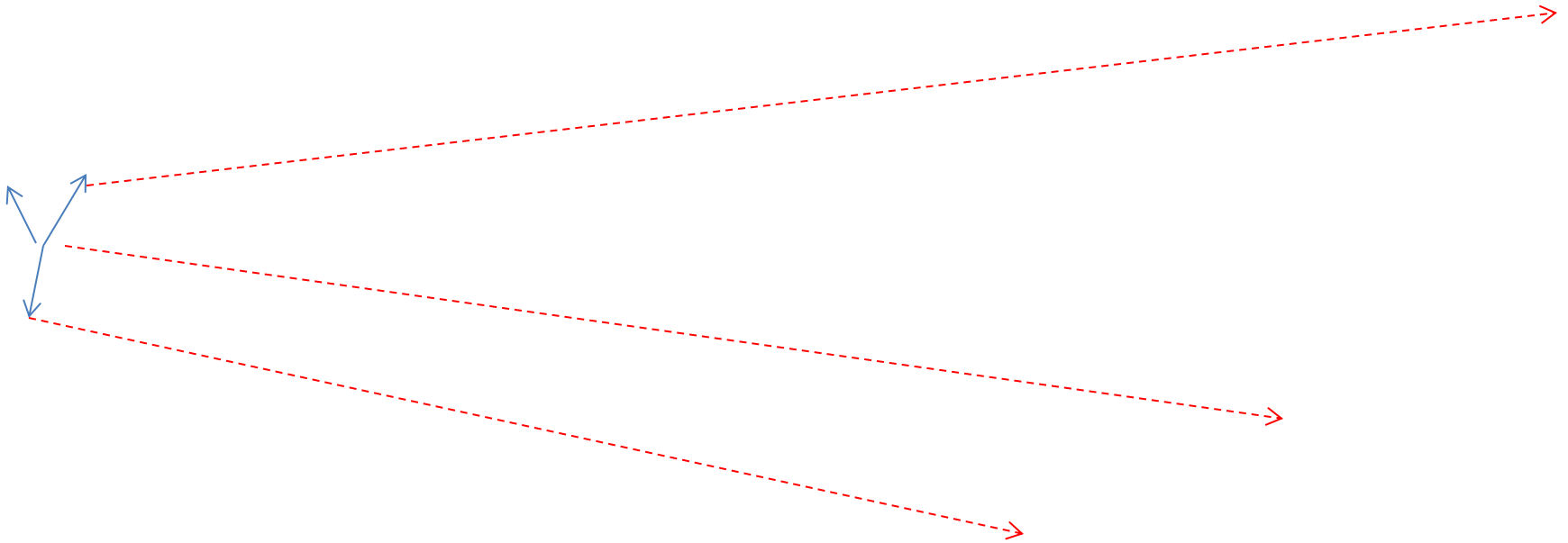
Exterior effacement must hold for the arbitrary distribution of distant spectator bodies with arbitrary velocities, etc.

The purely motion - independent portion of this transformation of local system coordinates can therefore be used by itself as a necessary condition for effacement.

$$\sum_i \frac{G m_i}{c^2 |\vec{r} - \vec{r}_i|} = \sum_{i(\text{local})} \frac{G m_i}{c^2 |\vec{r} - \vec{r}_i|} + \sum_s \frac{G m_s}{c^2 R_s}$$

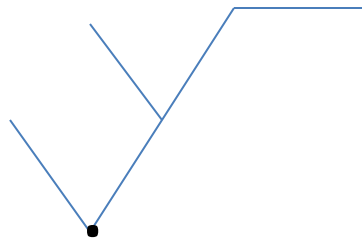
$\sum_s \frac{G m_s}{c^2 R_s}$ approaches 1 as more and more of universe is included.

Non - linearity means spectator scaling of local gravity.

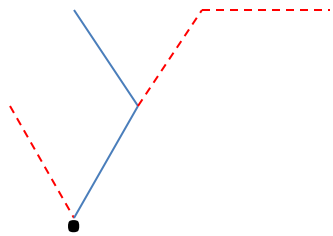


Example of Spectator Scaling of Lower order Local Potential by Higher order Potential

with $r_i = |\vec{r} - \vec{r}_i|$ and $r_{ij} = |\vec{r}_i - \vec{r}_j|$

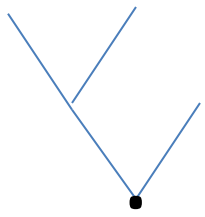


$$= \sum_{ijklq} \frac{G^5 m_i m_j m_k m_l m_q}{c^{10} r_i r_j r_{jk} r_{jl} r_{lq}}$$



$$= \sum_s \frac{G m_s}{c^2 R_s} \sum_{s's''} \frac{G^2 m_{s'} m_{s''}}{c^4 R_{s'} R_{s's''}} \sum_{jk} \frac{G^2 m_j m_k}{c^4 r_j r_{jk}}$$

Spectator Scaling of Lower Order Local Potentials by Higher Order Potentials



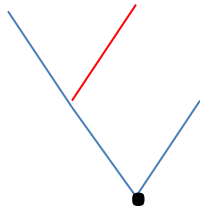
$$= \sum_{ijkl} \frac{G^4 m_i m_j m_k m_l}{c^8 r_i r_j r_{jk} r_{jl}}$$



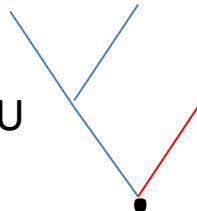
Find all reductions of this Potential into lower order Potentials times spectator scaling potentials



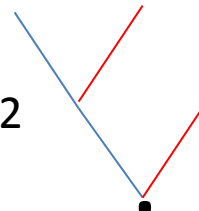
2U



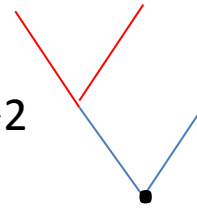
U



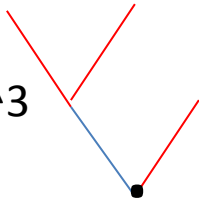
2U²



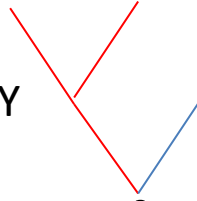
U²



U³



Y



field point \vec{r}

Spectator Potentials in red

$$U = \sum_s \frac{G m_s}{c^2 R_s}$$

$$Y = \sum_{ss's''} \frac{G^3 m_s m_{s'} m_{s''}}{c^6 R_s R_{ss'} R_{ss''}}$$

Goal is to find the numerical coefficients $\kappa(n \alpha)$ in :

$$-g_{SS}(\vec{r}) = 1 + \sum_{n \alpha} \kappa(n \alpha) V(n \alpha \vec{r} \dots) \quad \text{and} \quad -\bar{g}_{SS} = 1 + \sum_{n \alpha} \kappa(n \alpha) \bar{V}(n \alpha)$$

Useful Properties : $\bar{V}(n \alpha) \bar{V}(n' \alpha') = \bar{V}(n + n' \alpha'')$

$\mathbf{c}(n \alpha; n' \alpha' \beta)$ is the integer number of ways (including 0) that

$$V(n \alpha \vec{r} \dots) \text{ factors into } V(n' \alpha' \vec{r} \dots) \bar{V}(n - n' \beta)$$

Number of Potential Types at Each Order : 1, 2, 4, 9, 20, 48, 115, 286,

Fundamental **iterative algebraic equations** for Exterior Effacement are then

$$\sum_{\alpha} \mathbf{c}(n \alpha; n' \alpha' \beta) \kappa(n \alpha) = \left[\frac{\partial}{\partial \bar{V}(n - n' \beta)} (-\bar{g}_{SS})^{1-n'/2} \right] \kappa(n' \alpha')$$

for each $n' < n$, each α' , each β

Then $g_{00}(\vec{r}) = 1 + \sum_{n \alpha} \kappa^*(n \alpha) V(n \alpha \vec{r} \dots)$

$$\sum_{\alpha} \mathbf{c}(n \alpha; n' \alpha' \beta)^* \kappa^*(n \alpha) = \left[\frac{\partial}{\partial \bar{V}(n - n' \beta)} (\bar{g}_{00}) (-\bar{g}_{SS})^{-n'/2} \right] \kappa^*(n' \alpha')$$

The Spatial Metric g_{ab} Scaling

$-g_{ab} = \delta_{ab} \left(1 + \sum_{n \geq 1} V(n, \alpha \bar{r} \dots) \right)$ + non - diagonal motion - dependent potentials. $V(n, \alpha \bar{r} \dots)$ is type α and order m^n

$-\bar{g}_{ss} = \left(1 + \sum_{\alpha n=11}^{\infty} \bar{V}(n, \alpha) \right)$ yields local system's spatial background metric potential; $V(n, \alpha \bar{r} \dots) \Rightarrow \bar{V}(n, \alpha)$ by all $\bar{r}_i \rightarrow \bar{R}_s$ and $\bar{r} = 0$

$\bar{V}(n, \alpha)$ being spectator potentials, each $V(n, \alpha \bar{r} \dots)$ factors into $V(n', \alpha' \bar{r} \dots) \bar{V}(n-n', \alpha'')$ for all $n' \leq n, \alpha', \alpha''$

Sum of each potential's scaled contributions must be : $|\bar{g}_{ss}|^{1-n/2} V(n, \alpha \bar{r} \dots) \Rightarrow V(n, \alpha \bar{\rho} \dots)$ for $\bar{\rho}_i = |\bar{g}_{ss}|^{1/2} \bar{r}_i$

Spatial metric potentials to order m^5 are shown below as derived from Exterior Effacement algebraic scaling conditions outlined above.

$$r_i = |\bar{r} - \bar{r}_i| \quad r_{ij} = |\bar{r}_i - \bar{r}_j| \dots \quad \text{in units } G = c = 1$$

$$\begin{aligned}
 -g_{ss} = & 1 + 2 \sum_i \frac{m_i}{r_i} + \left(\frac{3}{2} \sum_{ij} \frac{m_i m_j}{r_i r_j} - \sum_{ij} \frac{m_i m_j}{r_i r_{ij}} \right) + \left(\frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_i r_j r_k} - \frac{3}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_i r_j r_{jk}} + \frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_i r_{ij} r_{ik}} + \frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_i r_{ij} r_{jk}} \right) \\
 & + \left(\frac{1}{16} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_j r_k r_l} - \frac{3}{4} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_j r_k r_{kl}} + \frac{1}{2} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_j r_{jk} r_{jl}} + \frac{5}{8} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_j r_{ik} r_{jl}} \right) \\
 & + \left(\frac{3}{4} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_j r_{jk} r_{kl}} - \frac{1}{8} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_{ij} r_{ik} r_{il}} - \frac{3}{4} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_{ij} r_{ik} r_{kl}} - \frac{1}{8} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_{ij} r_{jk} r_{jl}} - \frac{1}{4} \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_i r_{ij} r_{jk} r_{kl}} \right) \\
 & + \left(0 \sum_{ijklq} \frac{m_i m_j m_k m_l m_q}{r_i r_j r_k r_l r_q} - \frac{1}{8} \sum_{ijklq} \frac{m_i m_j m_k m_l m_q}{r_i r_j r_k r_l r_{iq}} + \dots \left(c(5,20) = \frac{1}{8} \right) \sum_{ijklq} \frac{m_i m_j m_k m_l m_q}{r_i r_{ij} r_{jk} r_{kl} r_{iq}} \right) + \dots \text{order} \left(\frac{m}{r} \right)^6
 \end{aligned}$$

For single source GR gives : $-g_{ss} = \left(1 + \frac{m}{2r} \right)^4$ a finite polynomial ending in term $\frac{1}{16} \left(\frac{m}{r} \right)^4$

N - Body Lagrangian Iterative Algebraic Equations for Enforcing Exterior Effacement

$$I = \int L dt \quad L = -\sum_i m_i \sqrt{1 - v_i^2 / c^2} + \sum \lambda(\alpha n n' n'' \dots) V(\alpha n n' n'' \dots \vec{r}_{ij} \dots \vec{v}_i \dots)$$

$V(\alpha n n' n'' \dots \vec{r}_{ij} \dots \vec{v}_i \dots)$ proportional to $1/r^n, v^n, a^n \dots$

So Needed Spectator Scaling Factors are: $\left[(\bar{g}_{00})^{(1-n')/2-n'' \dots} (-\bar{g}_{ss})^{(-n+n'+n'' \dots)/2} \right]$

$$c^*(\alpha n n' n'' \dots; \alpha' n^* n' n'' \dots \beta) = \text{number of ways } V(\alpha n n' n'' \dots \vec{r}_{ij} \dots \vec{v}_i \dots) \text{ factors into } V(\alpha' n^* n' n'' \dots \vec{r}_{ij} \dots \vec{v}_i \dots) \bar{V}(n - n^* \beta)$$

Spectator Scaling iterative algebraic equations for N - body Lagrangian coefficients $\lambda(\alpha n n' n'' \dots)$ are then :

$$\sum_{\alpha} c^*(\alpha n n' n'' \dots; \alpha' n^* n' n'' \dots \beta) \lambda(\alpha n n' n'' \dots) = \left[\frac{\partial}{\partial \bar{V}(n - n^* \beta)} \left((\bar{g}_{00})^{(1-n')/2-n'' \dots} (-\bar{g}_{ss})^{(-n+n'+n'' \dots)/2} \right) \right] \lambda(\alpha' n^* n' \dots)$$

For each $n^* < n$, each type α' , each $n', n'' \dots$, and each spectator potential type β

$$\frac{G^4 m_i m_j m_k m_l m_q}{c^6 r_{ij} r_{ik} r_{kl} r_{kq}} = \begin{matrix} \swarrow \searrow \\ \swarrow \searrow \\ \swarrow \searrow \\ \swarrow \searrow \end{matrix} \Rightarrow \begin{matrix} U & 2U & 2U^2 & U^2 \\ W & 2UW & Y & U^3 \end{matrix}$$

A potential scales at all lower orders

Spectator Potentials $U = \sum_s \frac{G m_s}{c^2 R_s} \quad W = \sum_{ss'} \frac{G^2 m_s m_{s'}}{c^4 R_s R_{ss'}} \quad Y = \sum_{ss's''} \frac{G^3 m_s m_{s'} m_{s''}}{c^6 R_s R_{ss'} R_{ss''}}$

Lagrangian Motion-Independent Potentials

Newtonian



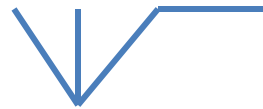
1PN



2PN



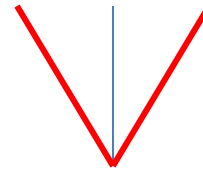
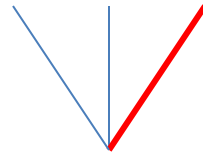
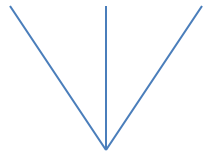
3PN



4PN

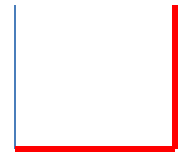
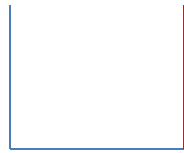


$$\frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{il}} \rightarrow 3U \frac{m_i m_j m_k}{r_{ij} r_{ik}} + 3U^2 \frac{m_i m_j}{r_{ij}}$$



m

$$\frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{kl}} \rightarrow 2U \frac{m_i m_j m_k}{r_{ij} r_{ik}} + U^2 \frac{m_i m_j}{r_{ij}} + 2W \frac{m_i m_j}{r_{ij}}$$



$$\text{Diagram} \rightarrow 5U \text{ Diagram} + 10U^2 \text{ Diagram} + 10U^3 \text{ Diagram} + 5U^4 \text{ Diagram}$$

$$\text{Diagram} \rightarrow 4U \text{ Diagram} + 2U^2 \text{ Diagram} + 4U^2 \text{ Diagram} + 4U^3 \text{ Diagram} + U^4 \text{ Diagram} + 4UY \text{ Diagram}$$

$$\text{Diagram} \rightarrow U \text{ Diagram} + 3U \text{ Diagram} + W \text{ Diagram} + 3W \text{ Diagram} + 3U^2 \text{ Diagram} + 4U^3 \text{ Diagram} + U^4 \text{ Diagram} + 3UW \text{ Diagram} + 3U^2W \text{ Diagram}$$

$$\text{Diagram} \rightarrow 2U \text{ Diagram} + (2W+U^2) \text{ Diagram} + 2UW \text{ Diagram} + W^2+2Z \text{ Diagram} + 2X \text{ Diagram} + 2UX \text{ Diagram}$$

$$\text{Diagram} \rightarrow U \text{ Diagram} + 2U \text{ Diagram} + U^2 \text{ Diagram} + 2W \text{ Diagram} + U^3 \text{ Diagram} + W^2 \text{ Diagram} + 4UW \text{ Diagram} + 2U^2W \text{ Diagram}$$

$$\text{Diagram} \rightarrow U \text{ Diagram} + 2U \text{ Diagram} + W \text{ Diagram} + 3U^2 \text{ Diagram} + X \text{ Diagram} + 2UX \text{ Diagram} + 2UX \text{ Diagram} + YU \text{ Diagram} + U^3 \text{ Diagram}$$

$$s \text{ Diagram} = U \quad s \text{ Diagram} \text{ --- } s' \text{ Diagram} = W \quad s \text{ Diagram} \text{ --- } s' \text{ Diagram} \text{ --- } s'' \text{ Diagram} = X \quad s' \text{ Diagram} \text{ --- } s \text{ Diagram} \text{ --- } s'' \text{ Diagram} = Y \quad s \text{ Diagram} \text{ --- } s' \text{ Diagram} \text{ --- } s'' \text{ Diagram} \text{ --- } s''' \text{ Diagram} = Z$$

The Motion - independent N - Body Lagrangian Potentials

From Iterative Algebraic Equations Enforcing Exterior Effacement

$$\begin{aligned}
 L = & \frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r_{ij}} - \frac{1}{2} \sum_{ijk} \frac{G^2 m_i m_j m_k}{c^2 r_{ij} r_{ik}} + \left(\frac{1}{4} \sum_{ijkl} \frac{G^3 m_i m_j m_k m_l}{c^4 r_{ij} r_{ik} r_{il}} + \frac{3}{8} \sum_{ijkl} \frac{G^3 m_i m_j m_k m_l}{c^4 r_{ij} r_{jk} r_{kl}} \right) \\
 & - \left(\frac{1}{8} \sum_{ijkl} \frac{G^4 m_i m_j m_k m_l m_q}{c^6 r_{ij} r_{ik} r_{il} r_{iq}} + \frac{1}{2} \sum_{ijkl} \frac{G^4 m_i m_j m_k m_l m_q}{c^6 r_{ij} r_{ik} r_{il} r_{lq}} + \frac{1}{4} \sum_{ijkl} \frac{G^4 m_i m_j m_k m_l m_q}{c^6 r_{ij} r_{jk} r_{kl} r_{lq}} \right) \\
 & + \left(\frac{1}{16} \sum_{ijklqp} \frac{G^5 m_i m_j m_k m_l m_q m_p}{c^8 r_{ij} r_{ik} r_{il} r_{iq} r_{ip}} + \frac{5}{16} \sum_{ijklqp} \frac{G^5 m_i m_j m_k m_l m_q m_p}{c^8 r_{ij} r_{ik} r_{il} r_{iq} r_{qp}} + \frac{5}{16} \sum_{ijklqp} \frac{G^5 m_i m_j m_k m_l m_q m_p}{c^8 r_{ij} r_{ik} r_{il} r_{kq} r_{lp}} \right. \\
 & \left. + \frac{5}{32} \sum_{ijklqp} \frac{G^5 m_i m_j m_k m_l m_q m_p}{c^8 r_{ij} r_{ik} r_{kl} r_{jq} r_{qp}} + \frac{5}{16} \sum_{ijklqp} \frac{G^5 m_i m_j m_k m_l m_q m_p}{c^8 r_{ij} r_{ik} r_{il} r_{lq} r_{qp}} + \frac{5}{32} \sum_{ijklqp} \frac{G^5 m_i m_j m_k m_l m_q m_p}{c^8 r_{ij} r_{jk} r_{kl} r_{lq} r_{qp}} \right) \\
 & + \text{order } m^7
 \end{aligned}$$

$$L = 1/2 \quad | \quad -1/2 \quad \triangle \quad +1/4 \quad \triangle \triangle \quad +3/8 \quad \square$$

$$-1/8 \quad \triangle \triangle \triangle \quad -1/4 \quad \square \quad -1/2 \quad \triangle \triangle \triangle$$

$$+1/16 \quad \triangle \triangle \triangle \triangle \quad +5/32 \quad \triangle \triangle \triangle \quad +5/16 \quad \triangle \triangle \triangle \triangle \quad +5/32 \quad \square \square$$

$$+5/16 \quad \triangle \triangle \triangle \triangle \quad +5/16 \quad \triangle \triangle \triangle \triangle \quad + \dots$$

$$(g_{00})^{1/2} = \frac{1 - m/2r}{1 + m/2r} = 1 - \frac{m}{r} + \frac{1}{2} \left(\frac{m}{r} \right)^2 - \frac{1}{4} \left(\frac{m}{r} \right)^3 + \frac{1}{8} \left(\frac{m}{r} \right)^4 - \frac{1}{16} \left(\frac{m}{r} \right)^5 + \dots$$

Some Lowest Order Examples of Required Scaling Fulfilled

$$L(1PN) = \sum_i m_i \left(-c^2 + \frac{1}{2} v_i^2 + \frac{1}{8c^2} v_i^4 \right) + \frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r_{ij}} \left(1 + \frac{1}{2c^2} (3v_{ij}^2 - \vec{v}_i \cdot \vec{v}_j - \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_j \cdot v_j) \right) - \frac{1}{2} \sum_{ijk} \frac{G^2 m_i m_j m_k}{c^2 r_{ij} r_{ik}} + \dots$$

$$E = \sum_i m_i \left(c^2 + \frac{1}{2} v_i^2 + \frac{3}{8c^2} v_i^4 \right) - \frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r_{ij}} \left(1 - \frac{1}{2c^2} (3v_{ij}^2 - \vec{v}_i \cdot \vec{v}_j - \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_j \cdot v_j) \right) + \frac{1}{2} \sum_{ijk} \frac{G^2 m_i m_j m_k}{c^2 r_{ij} r_{ik}} + \dots$$

$$\vec{P} = \sum_i \frac{\partial L}{\partial \vec{v}_i} \quad \vec{J} = \sum_i \vec{r}_i \times \frac{\partial L}{\partial \vec{v}_i}$$

Exterior Effacement upon scaling by Spectators' Potential $U = \sum_s Gm_s / c^2 R_s$ proper coordinates are: $d\tau = (1-U) dt$ $\vec{\rho}_i = (1+U) \vec{r}_i$

and

$$L dt \Rightarrow \left(\sum_i m_i \left(-c^2 [1-U] + \frac{1}{2} v_i^2 [1+3U] \right) + \frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r_{ij}} [1-2U] \right) dt = \left(\sum_i m_i \left(-c^2 + \frac{1}{2} u_i^2 \right) + \frac{1}{2} \sum_{ij} \frac{G m_i m_j}{\rho_{ij}} \right) d\tau \quad \text{Note : } \mathbf{G} \Rightarrow \mathbf{G}$$

Time Dilation on giving all bodies of a planar dynamical system a perpendicular boost: $\vec{v}_i \Rightarrow \vec{v}_i + \vec{w}$ and $d\tau^* = \sqrt{1-w^2/c^2} dt$ yields another scaling invariance of Newtonian Lagrangian:

$$L dt \Rightarrow \left(\sum_i m_i \left(-c^2 [1-\mathbf{w}^2/2c^2] + \frac{1}{2} v_i^2 [1+\mathbf{w}^2/2c^2] \right) + \frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r_{ij}} [1-\mathbf{w}^2/2c^2] \right) dt = \left(\sum_i m_i \left(-c^2 + \frac{1}{2} u_i^2 \right) + \frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r_{ij}} \right) d\tau^*$$

Lorentz Contraction

A rigid object has atoms which interact with each other

via potentials U_β and are at rest at sites \vec{r}_a

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$$\vec{\nabla}_a \sum_\beta U_\beta(\vec{r}_i) = 0 \quad \text{for all atom sites } \vec{r}_a$$

$$\vec{r}'_i = \vec{r}_i + \vec{r}_i \cdot \hat{w} \hat{w} (\sqrt{1 - w^2/c^2} - 1) + \vec{w} t \quad \text{so} \quad \vec{r}'_{ij} = \vec{r}_{ij} + \vec{r}_{ij} \cdot \hat{w} \hat{w} (\sqrt{1 - w^2/c^2} - 1)$$



w

• • • • •
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$$\vec{\nabla}_a \sum_\beta U_\beta(\vec{r}'_i, \vec{w}) = 0 \quad \text{so } (\vec{w} \cdot \vec{r}_{ij})^n \text{ terms must vanish for all powers } n.$$

$$\frac{G}{4c^2} \sum_{ij} \frac{m_i m_j}{r_{ij}} (3v_{ij}^2 - \vec{v}_i \cdot \vec{v}_j - \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j) \Rightarrow -\frac{G}{4c^2} \sum_{ij} \frac{m_i m_j}{r_{ij}} (w^2 + [\vec{w} \cdot \hat{r}_{ij}]^2)$$

$$\frac{G}{2} \sum_{ij} \frac{m_i m_j}{r'_{ij}} = \frac{G}{2} \sum_{ij} \frac{m_i m_j}{r_{ij} \sqrt{1 - [\vec{w} \cdot \hat{r}_{ij}]^2 / c^2}} \Rightarrow \frac{G}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} (1 + [\vec{w} \cdot \hat{r}_{ij}]^2 / 2c^2 + \dots)$$

So the $(\vec{w} \cdot \vec{r}_{ij})^2 / c^2$ term is canceled. This must be achieved to all orders in $1/c^2$

Time Dilation Scaling

$$\frac{G}{4c^2} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(3v_{ij}^2 - \vec{v}_i \cdot \vec{v}_j - \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j \right) \Rightarrow -\frac{G w^2}{4c^2} \sum_{ij} \frac{m_i m_j}{r_{ij}}$$

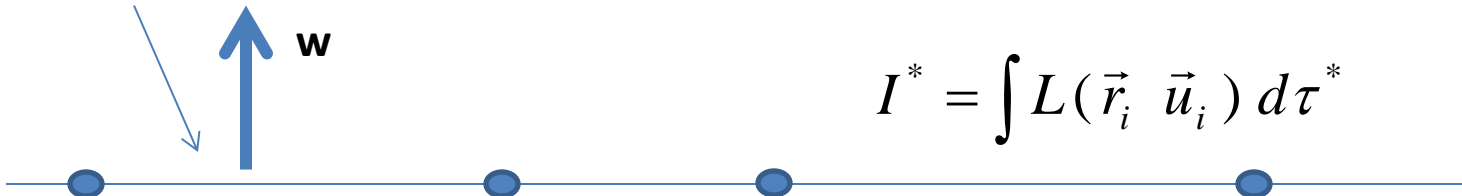
with $\vec{v}_i \Rightarrow \vec{v}_i + \vec{w}$ \vec{w} perpendicular to planar dynamics; all $\vec{w} \cdot \vec{v}_i = 0$

So
$$\frac{G}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} dt \Rightarrow \frac{G}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(1 - \frac{w^2}{2c^2} + \dots \right) dt \Rightarrow \frac{G}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} d\tau^*$$

$$I = \int L(\vec{r}_i, \vec{v}_i) dt$$

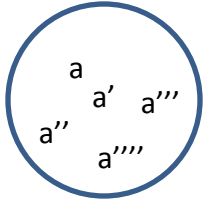


For all planar but general N-body dynamical systems



$$I^* = \int L(\vec{r}_i, \vec{u}_i) d\tau^*$$

$$d\tau^* = \sqrt{1 - w^2 / c^2} dt \quad \vec{u}_i = \vec{v}_i / \sqrt{1 - w^2 / c^2}$$



Composite Body Model is Gas of Atoms $a, a', a'' \dots$
with Thermal Motion and Gravitational interactions.

Interior Effacement : From Lagrangian expansion there is energy expansion $E = \sum_i \vec{v}_i \cdot \partial L / \partial \vec{v}_i - L + \dots$

At Rest Composite Body's Inertial Mass $M_A = E_A / c^2 = \sum_a m_a \left(1 + \frac{v_a^2}{2c^2} - \sum_{a'} \frac{G m_{a'}}{2c^2 r_{aa'}} + \frac{1}{c^4} \dots + \dots \right)$

Then two composite bodies A and B at rest and R_{AB} apart must have total energy :

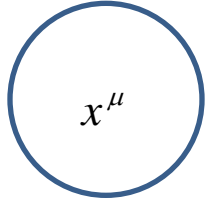
$$E_{AB} = \sum_a m_a \left(1 + \frac{u_a^2}{2c^2} - \sum_{a'} \frac{G m_{a'}}{2c^2 \rho_{aa'}} + \frac{1}{c^4} \dots + \dots \right) c^2 + \sum_b m_b \left(1 + \frac{u_b^2}{2c^2} - \sum_{a'} \frac{G m_{b'}}{2c^2 \rho_{bb''}} + \frac{1}{c^4} \dots + \dots \right) c^2$$

$$- \frac{G}{R_{AB}} \sum_a m_a \left(1 + \frac{u_a^2}{2c^2} - \sum_{a'} \frac{G m_{a'}}{2c^2 \rho_{aa'}} + \frac{1}{c^4} \dots + \dots \right) \sum_b m_b \left(1 + \frac{u_b^2}{2c^2} - \sum_{a'} \frac{G m_{b'}}{2c^2 \rho_{bb''}} + \frac{1}{c^4} \dots + \dots \right) + \dots$$

in which each body's proper coordinates are rescaled by presence of other body :

with $\rho_{aa''} = (-g_{SS}(B))^{1/2} r_{aa'}$, $\vec{u}_a = (-g_{SS}(B))^{1/2} (g_{00}(B))^{-1/2} \vec{v}_a$ etc.

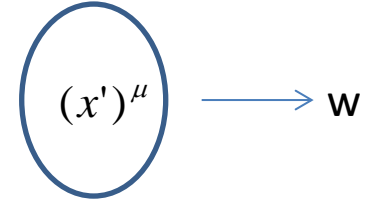
Interior Effacement



$$\vec{r}' - \vec{r}' \cdot \hat{w} \hat{w} = \vec{r} - \vec{r} \cdot \hat{w} \hat{w}$$

$$\vec{r}' \cdot \hat{w} \hat{w} = \frac{1}{\sqrt{1 - w^2/c^2}} (\vec{r} \cdot \hat{w} \hat{w} + \vec{w} t)$$

$$t' = \frac{1}{\sqrt{1 - w^2/c^2}} \left(t + \left(\frac{\vec{r} \cdot \hat{w}}{c^2} \right) \right)$$



$$\vec{v}' = \left((\vec{v} - \vec{v} \cdot \hat{w} \hat{w}) \sqrt{1 - w^2/c^2} + (\vec{w} + \vec{v} \cdot \hat{w} \hat{w}) \right) / \left(1 + \vec{w} \cdot \vec{v} / c^2 \right) - \vec{w} \vec{w} \cdot \vec{r} / c^2 + \dots$$

$$\vec{r}'_{ij} \cdot \vec{r}'_{ij} = r_{ij}^2 - (\vec{r}_{ij} \cdot \vec{w})^2 / c^2$$

$$E(N) = -\frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r'_{ij}} = -\frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r_{ij}} \left(1 + \frac{1}{2} \frac{(\hat{r}_{ij} \cdot \vec{w})^2}{c^2} + \dots \right)$$

$$E(N) = \frac{1}{2} \sum_i m_i (\vec{v}'_i)^2 = \frac{1}{2} \sum_i m_i \left(v_i^2 (1 - w^2/c^2) - 3(\vec{v}_{ij} \cdot \vec{w})^2 / c^2 \right) + \frac{1}{2} \sum_{ij} \frac{G m_i m_j}{r'_{ij}} \frac{(\vec{w} \cdot \hat{r}_{ij})^2}{c^2} + \dots$$

$$E(1PN) = \frac{3}{8} \sum_i m_i (\vec{v}'_i)^4 / c^2 = \frac{3}{8} \sum_i m_i v_i^4 / c^2 + \frac{3}{4} \sum_i m_i v_i^2 w^2 / c^2 + \frac{3}{2} \sum_i m_i (\vec{v}_i \cdot \vec{w})^2 / c^2 + \dots$$

$$E(1PN) = \frac{1}{4} \sum_{ij} \frac{G m_i m_j}{c^2 r'_{ij}} (3v_{ij}^2 - \vec{v}_i \cdot \vec{v}_j - \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j) = -\frac{1}{4} \sum_{ij} \frac{G m_i m_j}{c^2 r_{ij}} (w^2 + (\vec{w} \cdot \hat{r}_{ij})^2)$$

$$E = \sum_i m_i \left(1 + \frac{v_i^2}{2c^2} - \sum_j \frac{G m_j}{2c^2 r_{ij}} + \dots \right) \left(c^2 + \frac{1}{2} w^2 + \dots \right) \quad \text{So } \mathbf{M}_G = \mathbf{M}_I$$

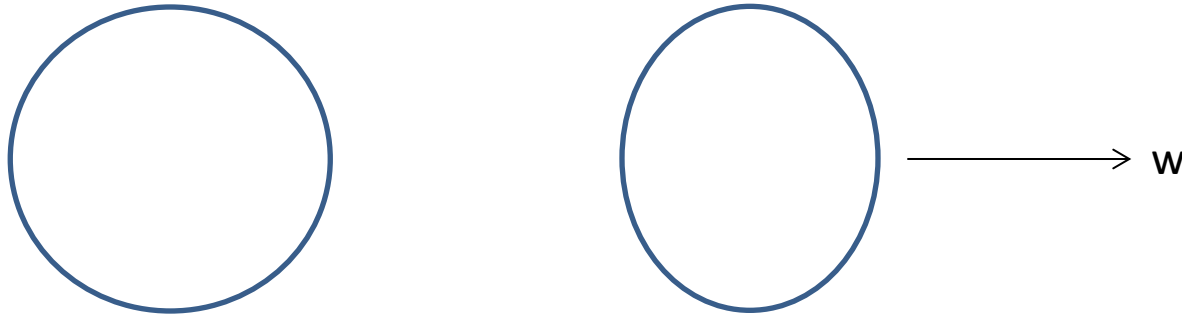
Summary of Interior Effacement Conditions used to Fix Lagrangian Expansion

$$E = \sum_i m_i \left(1 + \frac{1}{2} \frac{v_i^2}{c^2} - \frac{1}{2} \sum_j \frac{G m_j}{c^2 r_{ij}} + \frac{3}{8} \frac{v_i^4}{c^4} + \frac{1}{2} \sum_{jk} \frac{G^2 m_j m_k}{c^4 r_{ij} r_{ik}} + \frac{1}{4} \sum_j \frac{G m_j}{c^4 r_{ij}} (3v_{ij}^2 - \vec{v}_i \cdot \vec{v}_j - \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j) + \dots \right) \left(c^2 + \frac{1}{2} w^2 + \dots \right)$$

and

$$E = \sum_i m_i \left(1 + \frac{1}{2} \frac{v_i^2}{c^2} - \frac{1}{2} \sum_j \frac{G m_j}{c^2 r_{ij}} + \dots \right) \left(c^2 + \frac{1}{2} w^2 + \frac{3}{8} \frac{w^4}{c^2} + \dots \right)$$

for a moving composite body must be obtained, with the expression for body mass expressed in its rest frame, proper coordinates.



For two composites at rest :
$$E = - \frac{G}{R} \sum_a m_a \left(1 + \frac{1}{2} \frac{v_a^2}{c^2} - \frac{1}{2} \sum_j \frac{G m_{a'}}{c^2 r_{aa'}} + \dots \right) \sum_b m_b \left(1 + \frac{1}{2} \frac{v_b^2}{c^2} - \frac{1}{2} \sum_j \frac{G m_{b'}}{c^2 r_{bb'}} + \dots \right)$$

and if two composite bodies at rest are given rotational angular momentum :

$$E = - \frac{G (\vec{J}_A \cdot \vec{J}_B - 3 \vec{J}_A \cdot \hat{R} \hat{R} \cdot \vec{J}_B)}{c^2 R^3} \quad \text{with} \quad \vec{J}_A = \sum_a m_a \vec{r}_a \times \vec{v}_a \quad \text{etc.}$$

But at next order conserved
$$\vec{J}_A = \sum_a m_a \vec{r}_a \times \left(\left[1 + \frac{v_a^2}{2c^2} + 3 \sum_{a'} \frac{G m_{a'}}{c^2 r_{aa'}} \right] \vec{v}_a - \sum_{a'} \frac{G m_{a'}}{2c^2 r_{aa'}} [7 + \hat{r}_{aa'} \hat{r}_{aa'} \cdot] \vec{v}_{a'} \right)$$
 so ??

1/c⁴ order N-Body Lagrangian from Iterative Algebraic Equations

$$\begin{aligned}
 L = & \sum_i m_i \left(-c^2 + \frac{1}{2} v_i^2 + \frac{1}{8} v_i^4 + \frac{1}{16} v_i^6 \right) + \frac{1}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(1 + \frac{3}{2} v_{ij}^2 - \frac{1}{2} [\vec{v}_i \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j] - \sum_k \frac{m_k}{r_{ik}} \right) \\
 & + \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(\begin{aligned} & \frac{7}{16} v_{ij}^4 - \frac{1}{4} v_i^2 v_j^2 + \frac{3}{16} (\vec{v}_i \cdot \vec{v}_j)^2 + \frac{5}{8} (\vec{v}_i \cdot \vec{v}_j) v_{ij}^2 \\ & \left(+ \frac{9}{32} v_{ij}^2 - \frac{1}{8} \vec{v}_i \cdot \vec{v}_j + \frac{3}{32} ((\vec{v}_i \cdot \hat{r}_{ij})^2 + (\hat{r}_{ij} \cdot \vec{v}_j)^2) \right) \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j \\ & + \frac{11}{32} (\vec{a}_i \cdot \vec{v}_j \vec{v}_j - \vec{a}_j \cdot \vec{v}_i \vec{v}_i + 2 \vec{a}_i \cdot \vec{v}_j \vec{v}_i - 2 \vec{a}_j \cdot \vec{v}_i \vec{v}_j) \end{aligned} \right) \\
 & + \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} \left(2v_i^2 - \frac{1}{2} [v_j^2 + v_k^2] - \frac{7}{4} \vec{v}_i \cdot [\vec{v}_j + \vec{v}_k] + \frac{11}{4} \vec{v}_j \cdot \vec{v}_k - \frac{1}{4} \vec{v}_i \cdot [\hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j + \hat{r}_{ik} \hat{r}_{ik} \cdot \vec{v}_k] \right) \\
 & + \frac{1}{4} \sum_{ijk} \frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{il}} + \frac{3}{8} \sum_{ijk} \frac{m_i m_j m_k m_l}{r_{ij} r_{jk} r_{kl}} \quad \boxed{G = c = 1}
 \end{aligned}$$

plus a collection of potentials with one undetermined coefficient ξ

$$+ \xi \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(\begin{aligned} & ((\hat{r}_{ij} \cdot \vec{v}_j)^2 \vec{a}_i - (\vec{v}_i \cdot \hat{r}_{ij})^2 \vec{a}_j - 2 \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j \vec{a}_{ij}) \cdot \vec{r}_{ij} \\ & + (3(\vec{v}_j \cdot \hat{r}_{ij})^2 - 2v_{ij}^2) \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j + (v_j^2 - \vec{v}_i \cdot \vec{v}_j)(\vec{v}_i \cdot \hat{r}_{ij})^2 + (v_i^2 - \vec{v}_i \cdot \vec{v}_j)(\vec{v}_j \cdot \hat{r}_{ij})^2 \end{aligned} \right)$$

From the $1/c^2$ Lagrangian there is a spin - spin energy between bodies

$$E = - \frac{G \left(\vec{J}_A \cdot \vec{J}_B - 3 \vec{J}_A \cdot \hat{R} \hat{R} \cdot \vec{J}_B \right)}{c^2 R^3}$$

$$\text{with } \vec{J}_A = \sum_a \vec{r}_a \times \vec{p}_a = \sum_a m_a \vec{r}_a \times \vec{v}_a \text{ etc.}$$

But at $1/c^2$ level

$$\vec{J}_A = \sum_a m_a \vec{r}_a \times \left(\left[1 + \frac{v_a^2}{2c^2} + 3 \sum_{a'} \frac{G m_{a'}}{c^2 r_{aa'}} \right] \vec{v}_a - \sum_{a'} \frac{G m_{a'}}{2c^2 r_{aa'}} [7 + \hat{r}_{aa'} \hat{r}_{aa'} \cdot] \vec{v}_{a'} \right)$$

Interior Effacement requires that at the $1/c^4$ order the original E holds.

$$\begin{aligned}
g_{00} &= 1 - 2 \sum_i \frac{m_i}{r_i} + 2 \sum_{ij} \frac{m_i m_j}{r_i r_j} + 2 \sum_{ij} \frac{m_i m_j}{r_i r_{ij}} - 3 \sum_i \frac{m_i}{r_i} v_i^2 + \dots \\
- g_{ab} &= \left(1 + 2 \sum_i \frac{m_i}{r_i} + \frac{3}{2} \sum_{ij} \frac{m_i m_j}{r_i r_j} - \sum_{ij} \frac{m_i m_j}{r_i r_{ij}} + \sum_i \frac{m_i}{r_i} [\vec{v}_i \cdot \hat{r}_i]^2 + \dots \right) \delta_{ab} - 2 \sum_i \frac{m_i}{r_i} (\vec{v}_i)_a (\vec{v}_i)_b + \dots \\
(g_{0x}, g_{0y}, g_{0z}) &\equiv \vec{h} = \frac{7}{2} \sum_i \frac{m_i}{r_i} \vec{v}_i + \frac{1}{2} \sum_i \frac{m_i}{r_i} \vec{v}_i \cdot \hat{r}_i \hat{r}_i + \dots
\end{aligned}$$

Spectators' interior metric field components

$$\begin{aligned}
g_{00} &= 1 - 2U + 2U^2 + 2W - 3K + \dots \\
- g_{ab} &= \left(1 + 2U + \frac{3}{2}U^2 - W + K^* + \dots \right) \delta_{ab} - 2K_{ab} + \\
(g_{0x}, g_{0y}, g_{0z}) &= \vec{h} = \frac{7}{2} \vec{S} + \frac{1}{2} \vec{S}^* + \dots
\end{aligned}$$

wanting just the scaling factors : $d\tau = (g_{00})^{1/2} dt$ $\vec{\rho} = (-g_{ss})^{1/2} \vec{r}$, so $K_{ab} \rightarrow \frac{1}{3} K \delta_{ab}$ Also : $\vec{S} = 0 = \vec{S}^*$

$$\begin{aligned}
U &= \sum_s \frac{m_s}{R_s} \quad [U] = \sum_s \frac{m_s}{R_s} \hat{R}_s \hat{R}_s \rightarrow \frac{1}{3} U \delta_{ab} \quad W = \sum_{ss'} \frac{m_s m_{s'}}{R_s R_{s'}} \quad K = \sum_s \frac{m_s}{R_s} v_s^2 \quad K^* = \sum_s \frac{m_s}{R_s} (\vec{v}_s \cdot \hat{R}_s)^2 \quad K_{ab} = \sum_s \frac{m_s}{R_s} (\vec{v}_s)_a (\vec{v}_s)_b \\
\vec{S} &= \sum_s \frac{m_s}{R_s} \vec{v}_s \quad \vec{S}^* = \sum_s \frac{m_s}{R_s} \vec{v}_s \cdot \hat{R}_s \hat{R}_s \quad [[U]] = \sum_s \frac{m_s}{R_s} \hat{R}_s \hat{R}_s \hat{R}_s \hat{R}_s \rightarrow \frac{1}{15} U (\delta_{ab} \delta_{cd} + \delta_{ac} \delta_{bd} + \delta_{ad} \delta_{bc})
\end{aligned}$$

Spectator Scaling

$$\left((g_{00})^a (-g_{ss})^b - 1 \right) = \sum_t c_t S_t$$

$S_t = U, W, U^2, UW, \dots$ are different dimensionless gravitational potentials produced by spectator bodies in the time - time and spatial metric components at the interior location and which rescale proper time and proper spatial intervals there. The powers a and b depend on powers of $1/r$ and velocity of bodies for the particular Lagrangian term needing rescaling to interior location proper space - time coordinates.

On the other hand, higher order terms in the Lagrangian can have these same spectator potentials extracted from themselves, thereby producing lower order Lagrangian terms for N body systems in the interior location but multiplied by this or that dimensionless spectator potential P_t .

The sum of these rescaling contributions to any N body Lagrangian term

must reproduce the same $\sum_t c_t P_t$

$$\int L(\vec{r}_i, \vec{v}_i) dt = L(\vec{R}_s, \vec{V}_s) + L(\vec{R}_s, \vec{V}_s, \vec{r}_i, \vec{v}_i) + L(\vec{r}_i, \vec{v}_i) \rightarrow \int L(\vec{\rho}_i, \vec{u}_i) d\tau$$

L for inner bodies after spectator scaling by "mixed" L should reproduce itself.

$$\vec{r}_i \rightarrow \vec{\rho}_i (-g_{ss})^{-1/2} \quad \vec{v}_i \rightarrow \vec{u}_i (-g_{ss})^{-1/2} (g_{00})^{1/2} \quad d\tau = (g_{00})^{1/2} dt$$

Starting point is the 1PN order GR Lagrangian :

$$\begin{aligned} L &= \sum_i m_i \left(-1 + \frac{1}{2} v_i^2 + \frac{1}{8} v_i^4 \right) + \frac{1}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(1 - \frac{1}{2} [\vec{v}_i \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_j \cdot \vec{v}_j] \right) \\ &\quad + \frac{3}{4} \sum_{ij} \frac{m_i m_j}{r_{ij}} v_{ij}^2 - \frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} + \dots \\ &\Rightarrow \frac{1}{2} \sum_i m_i v_i^2 (1 + 3U) + \frac{1}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} (1 - 2U) \\ \text{or } &\Rightarrow \frac{1}{2} \sum_i m_i v_i^2 \left(1 + \frac{1}{2} w^2 \right) + \frac{1}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(1 - \frac{1}{2} w^2 \right) \end{aligned}$$

Total L is assumed superposition of terms proportional to $\left(\frac{G^n}{c^{2n-2+2s}} \right) \frac{m^{n+1}}{r^n} v^{2s}$

A Lagrangian term proportional to $v^{2s} r^{-n}$ needs

spectator scaling factor $(g_{00})^{(1-2s)/2} (-g_{ss})^{(2s-n)/2}$

$$L(GR) = \sum_i m_i \left(-1 + \frac{1}{2} v_i^2 + \frac{1}{8} v_i^4 + \dots \right) + \frac{1}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(1 - \frac{1}{2} [\vec{v}_i \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_j \cdot v_j] + \dots \right) \\ + \frac{3}{4} \sum_{ij} \frac{m_i m_j}{r_{ij}} v_{ij}^2 - \frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} + \dots$$

$$L(LIPPN) = \sum_i m_i \left(-1 + \frac{1}{2} v_i^2 + \frac{1}{8} v_i^4 + \dots \right) + \frac{1}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(1 - \frac{1}{2} [\vec{v}_i \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_j \cdot v_j] + \dots \right) \\ + \frac{2+\gamma}{4} \sum_{ij} \frac{m_i m_j}{r_{ij}} v_{ij}^2 - \frac{2\beta-1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} + \dots \\ \gamma \rightarrow \gamma(\vec{r}, t) \quad \beta \rightarrow \beta(\vec{r}, t)$$

$$L(CBPPN) = \sum_i M(I)_i \left(-1 + \frac{1}{2} v_i^2 + \frac{1}{8} v_i^4 + \dots \right) + \frac{1}{2} \sum_{ij} \frac{\Gamma_{ij}}{r_{ij}} \left(1 - \frac{1}{2} [\vec{v}_i \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_j \cdot v_j] + \dots \right) \\ + \frac{2+\gamma}{4} \sum_{ij} \frac{\Theta_{ij}}{r_{ij}} v_{ij}^2 - \frac{2\beta-1}{2} \sum_{ijk} \frac{A_{ijk}}{r_{ij} r_{ik}} + \dots$$

$$\Gamma_{ij} = G(\vec{r}, t) \left(M(G)_i M(G)_j + \xi U_i U_j + \dots \right) \quad M(G) / M(I) \text{ not necessarily 1, etc.....}$$

$$L = \frac{1}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} - \frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} + s \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{il}} + r \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_{ij} r_{jk} r_{kl}} + \dots$$

$$\sum_{ijkl} \frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{il}} \Rightarrow 3U^2 \sum_{ij} \frac{m_i m_j}{r_{ij}} + 3U \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}}$$

$$\sum_{ijkl} \frac{m_i m_j m_k m_l}{r_{ij} r_{jk} r_{kl}} \Rightarrow (U^2 + 2W) \sum_{ij} \frac{m_i m_j}{r_{ij}} + 2U \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}}$$

U spectator scaling is then :

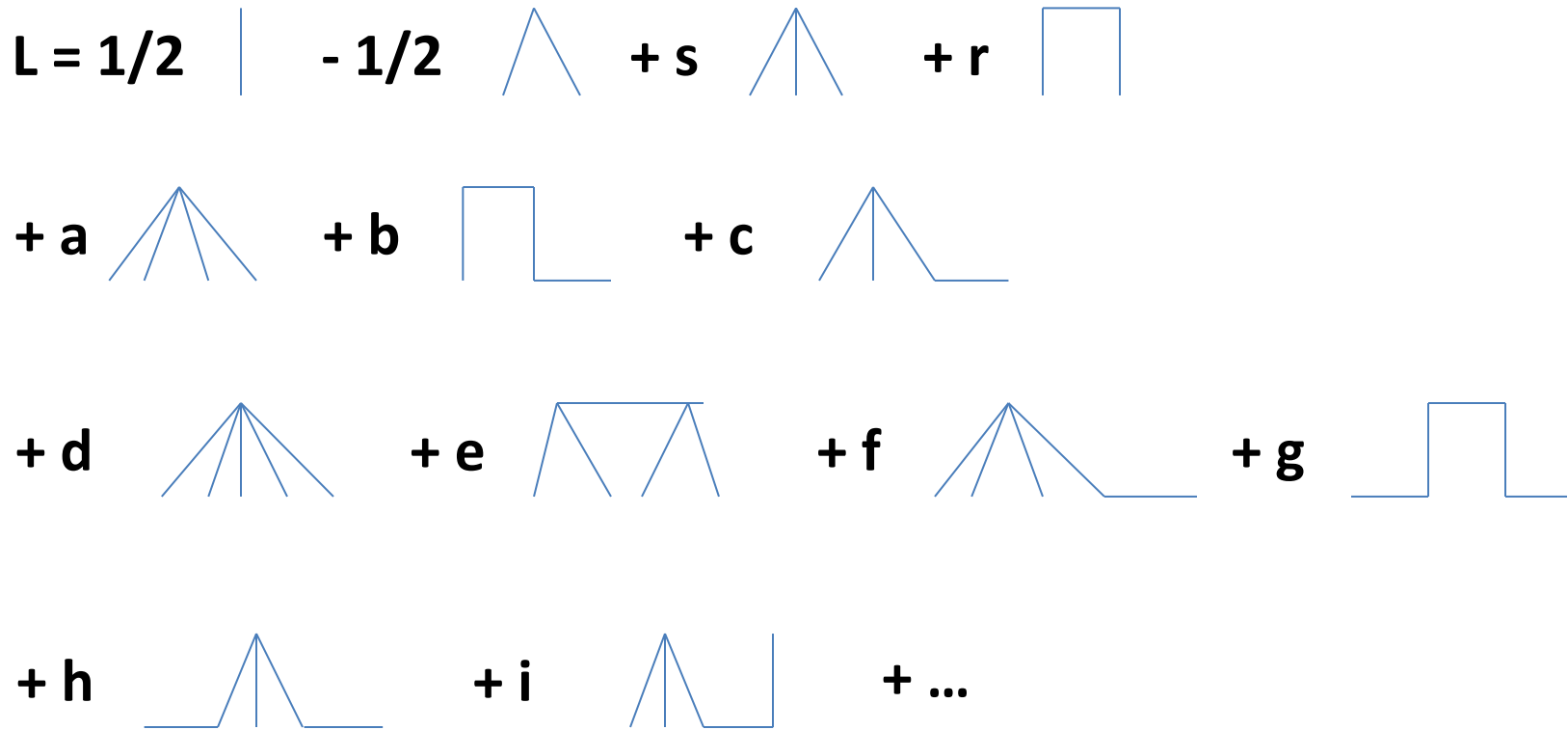
$$-\frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} + 3sU \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} + 2rU \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} = -\frac{1}{2} (g_{00})^{1/2} \sum_{ijk} \frac{m_i m_j m_k}{\rho_{ij} \rho_{ik}}$$

$$\text{with } (g_{00})^{1/2} \cong 1 - U \quad \text{and} \quad \frac{1}{\rho} \cong \frac{(1-U)}{r} \quad \text{so} \quad 3s + 2r = \frac{3}{2}$$

Then U^2 and W scaling of Newtonian potential gives additional linear equations involving s and r

$$\frac{n+2}{2} V(n) = V(n+2, W) \quad \frac{(2n+1)(n+2)}{4} V(n) = V(n+2, U^2)$$

$$2r = \frac{3}{4} \quad 3s + r = \frac{9}{8} \quad \rightarrow \quad s = 1/4 \quad r = 3/8$$



To convert interaction pictures into potentials put in numerator a mass M_i at free ends of lines and at junctions of lines. For each line put in denominator the appropriate inter - mass distance R_{ij} . Add required powers of G and c and sum over all bodies.

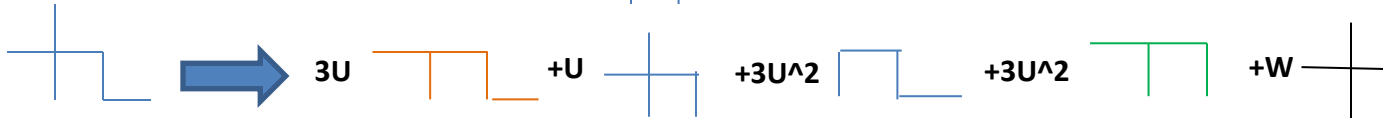
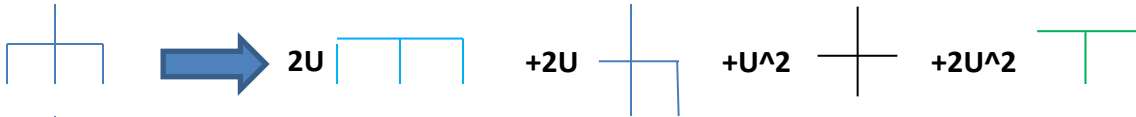
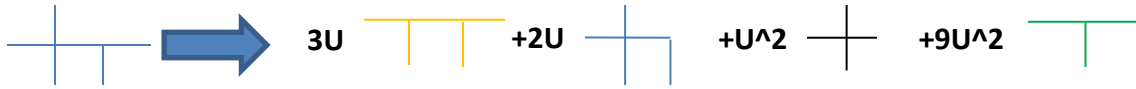
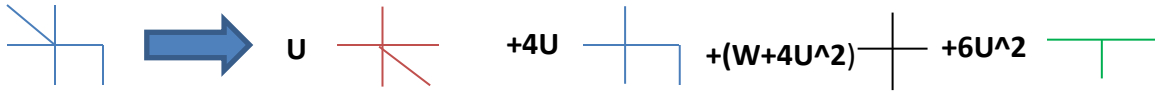
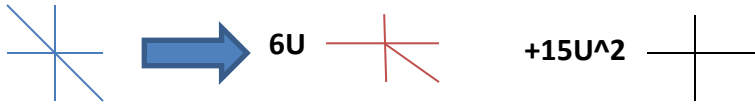
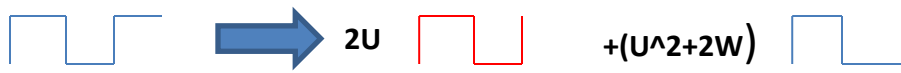
For interaction picture e this produces :





$$V(e) = \sum_{ijklpq} \frac{G^5 M_i M_j M_k M_l M_p M_q}{c^8 R_{ij} R_{ik} R_{il} R_{lp} R_{lq}}$$

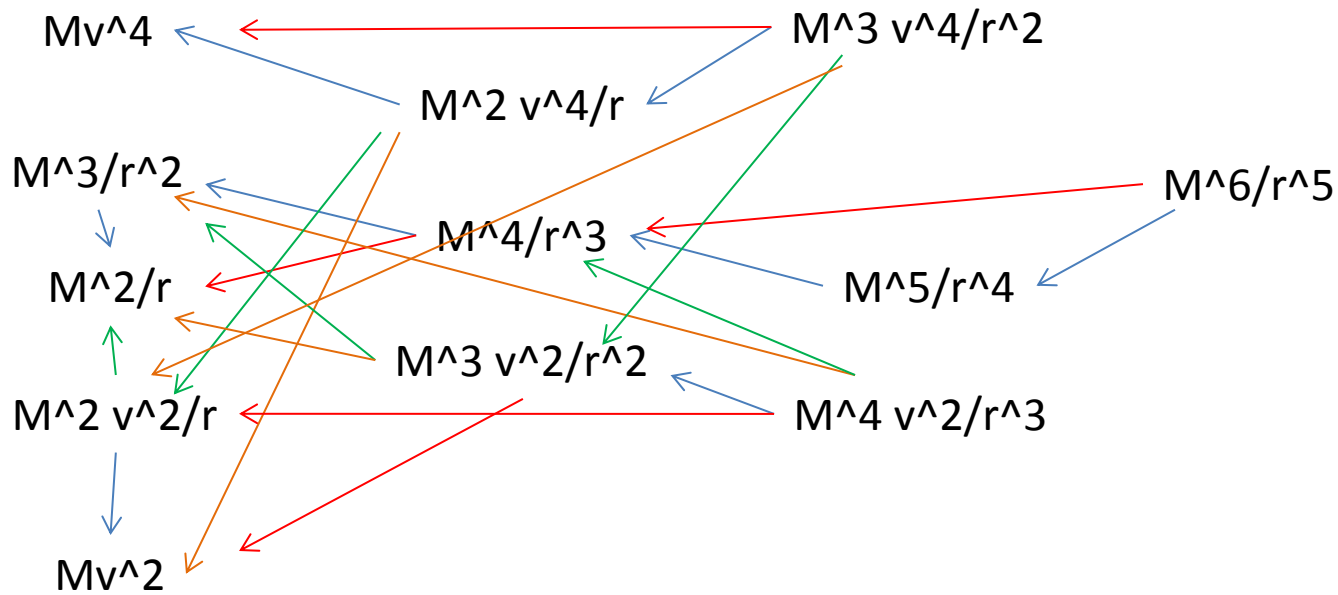
m^7 order(r^6 order) Potentials

Rules for computers to generate all

m^N order Potentials needed!



-  U Scaling
-  U², W Scaling
-  K Scaling
-  w² SR Scaling



Special Relativistic Scaling

Consider any general PLANAR dynamical system with Lagrangian

$$L = -\sum_i m_i (1 - v_i^2)^{1/2} + \sum_{ij\dots} L(\vec{r}_i, \vec{r}_j, \dots, \vec{v}_i, \vec{v}_j, \dots) \text{ and equations of motion :}$$

$$m_i \frac{d}{dt} \frac{\vec{v}_i}{(1 - v_i^2)^{1/2}} = \left(\frac{\partial}{\partial \vec{r}_i} - \frac{d}{dt} \frac{\partial}{\partial \vec{v}_i} \right) L(\vec{r}_i, \vec{v}_i)$$

$$\text{and } \frac{d}{dt} \frac{\vec{v}_i}{(1 - v_i^2)^{1/2}} = (1 - w^2)^{1/2} \frac{d}{d\tau} \frac{\vec{u}_i}{(1 - u_i^2)^{1/2}} \quad \vec{u} = \frac{\vec{v}}{(1 - w^2)^{1/2}}$$

Give each body an identical velocity \vec{w} perpendicular to the planar motion,

convert to proper time in frame moving with plane, $d\tau = (1 - w^2)^{1/2} dt$,

and corresponding proper in - plane velocities $\vec{u}_i = \vec{v}_i / (1 - w^2)^{1/2}$,

and identical planar motions must exist in the proper time variable.

$$L(\vec{r}_i, \vec{v}_i + \vec{w}) = (1 - w^2)^{1/2} L(\vec{r}_i, \vec{v}_i / (1 - w^2)^{1/2}) \quad \left| \text{planar dynamics edge views} \right| \rightarrow \vec{w}$$

Since interaction Lagrangian is sum of terms proportional to various powers

of velocity, v^n , each such term must acquire a scaling factor $(1 - w^2)^{(1-n)/2}$

from the replacements $\vec{v}_i \rightarrow \vec{v}_i + \vec{w}$ in the interaction Lagrangian.

$$\text{Known 1PN } L_0 = \frac{1}{2} \sum_i m_i v_i^2 \left(1 + \frac{1}{4} v_i^2 \right) + \frac{1}{2} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(1 + \frac{3}{2} v_{ij}^2 - \frac{1}{2} [\vec{v}_i \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j] \right) - \frac{1}{2} \sum_{ij} \frac{m_i m_j m_k}{r_{ij} r_{ik}}$$

Then imposing U, U^2 and w^2 scaling, and nullifying \vec{S}^* with last term :

$$L = -\frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} \left(1 - 4v_i^2 + [v_k^2 + v_j^2] + \frac{7}{2} \vec{v}_i \cdot [\vec{v}_j + \vec{v}_k] - \frac{11}{2} \vec{v}_j \cdot \vec{v}_k + \frac{1}{2} [\vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ik} \hat{r}_{ik} \cdot \vec{v}_k] - \frac{1}{2} \vec{v}_k \cdot \hat{r}_{jk} \hat{r}_{jk} \cdot \vec{v}_j \right)$$

or $L = -\frac{1}{2} \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} \left(1 - 2[v_{ij}^2 + v_{ik}^2] + 3v_{jk}^2 - \frac{1}{2} \vec{v}_i \cdot [\vec{v}_j + \vec{v}_k] + \frac{1}{2} \vec{v}_j \cdot \vec{v}_k + \frac{1}{2} [\vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ik} \hat{r}_{ik} \cdot \vec{v}_k] - \frac{1}{2} \vec{v}_k \cdot \hat{r}_{jk} \hat{r}_{jk} \cdot \vec{v}_j \right)$

But scaling term proportional to \vec{S} remains : $\sum_{ij} \frac{m_i m_j}{r_{ij}} (\vec{v}_i + \vec{v}_j) \cdot \vec{S}$

Next order then :

$$+ \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{il}} \left(\begin{aligned} & \frac{1}{4} + a v_i^2 + b [v_j^2 + v_k^2 + v_l^2] + c \vec{v}_i \cdot [\vec{v}_j + \vec{v}_k + \vec{v}_l] + d [\vec{v}_j \cdot \vec{v}_k + \vec{v}_j \cdot \vec{v}_l + \vec{v}_k \cdot \vec{v}_l] \\ & + \lambda [\vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ik} \hat{r}_{ik} \cdot \vec{v}_k + \vec{v}_i \cdot \hat{r}_{il} \hat{r}_{il} \cdot \vec{v}_l] + \mu [\vec{v}_k \cdot \hat{r}_{kj} \hat{r}_{kj} \cdot \vec{v}_j + \vec{v}_l \cdot \hat{r}_{lk} \hat{r}_{lk} \cdot \vec{v}_k + \vec{v}_j \cdot \hat{r}_{jl} \hat{r}_{jl} \cdot \vec{v}_l] \end{aligned} \right)$$

$$+ \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{jl}} \left(\frac{3}{8} + e [v_i^2 + v_j^2] + f [v_k^2 + v_l^2] + g \vec{v}_i \cdot \vec{v}_j + h [\vec{v}_i \cdot \vec{v}_k + \vec{v}_j \cdot \vec{v}_l] + r [\vec{v}_i \cdot \vec{v}_l + \vec{v}_j \cdot \vec{v}_k] + s \vec{v}_k \cdot \vec{v}_l + \dots \right)$$

with $a + 3b + 3c + 3d = -\frac{1}{8}$ $2e + 2f + g + 2h + 2r + s = -\frac{3}{16}$ from w^2 scaling.

$$r_{ij} = |\vec{r}_i - \vec{r}_j|, \quad v_{ij}^2 = (\vec{v}_i - \vec{v}_j)^2, \text{ etc. } \quad G = c = 1$$

Is such a term in L viable?
$$\sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} \left[\vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j + \vec{v}_i \cdot \hat{r}_{ik} \hat{r}_{ik} \cdot \vec{v}_k \right]$$

Doing a spectator reduction with : $\vec{S}^* = \sum_s \frac{G m(s) \vec{v}(s) \cdot \hat{R}(s) \hat{R}(s)}{c^3 R(s)}$ leads to :

$$L = \sum_{ij} \frac{m_i m_j}{r_{ij}} (\vec{v}_j + \vec{v}_i) \cdot \vec{S}^*$$

For our galaxy $|\vec{S}^*|$ is of order 10^{-9} and velocity of moon relative to Earth is order 10^{-5} , making this sidereal perturbation just about detectable in LLR. Perhaps

it is more observable in binary or trinary pulsar systems?

And \vec{S}^* from universe as a whole could be of order 10^{-5} ?

$$-g_{ss} = 1 + 2U + \frac{3}{2}U^2 - W + \dots$$

$$(-g_{ss})^{n/2} = 1 + nU + \frac{n(2n-1)}{4}U^2 - \frac{n}{2}W + \dots \quad \frac{1}{r^n} = (-g_{ss})^{n/2} \frac{1}{\rho^n}$$

$$(g_{00})^{1/2} = 1 - U + \frac{1}{2}U^2 + W + \dots \quad (g_{00})^{1/2} dt = d\tau$$

with $U = \sum_s \frac{G m(s)}{c^2 R(s)}$ and $W = \sum_{s,s'} \frac{G^2 m(s)m(s')}{c^4 R(s)R(ss')}$ are scaling potentials due to outside world bodies.

$$V(n) + U V(n+1, U) + U^2 V(n+2, U^2) + W V(n+2, W) + \dots = (g_{00})^{1/2} (-g_{ss})^{-n/2} V(n)$$

with $V(n)$, $V(n+1, U)$, $V(n+2, U^2)$, $V(n+2, W)$ being Lagrangian potentials with n interbody distances in denominator and $n+1$ body masses in numerator.

$$-(n+1) V(n) = V(n+1, U)$$

$$\frac{n+2}{2} V(n) = V(n+2, W)$$

$$\frac{(2n+1)(n+2)}{4} V(n) = V(n+2, U^2)$$

$$L = \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(a v_i^2 v_j^2 + b (\vec{v}_i \cdot \vec{v}_j)^2 + e \vec{v}_i \cdot \vec{v}_j \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j \right)$$

Under Special Relativity scaling :

$$\rightarrow w^2 \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(a [v_i^2 + v_j^2] + 2b \vec{v}_i \cdot \vec{v}_j + e \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j \right)$$

Scale factor is $\left(1 + \frac{1}{2} w^2 \right)$ so $e = -\frac{1}{8}$ $a = \frac{3}{8}$ $b = -a - \frac{1}{16} = -\frac{7}{16}$

$$L = \frac{1}{16} \sum_{ij} \frac{m_i m_j}{r_{ij}} \left(6 v_i^2 v_j^2 - 7 (\vec{v}_i \cdot \vec{v}_j)^2 - 2 \vec{v}_i \cdot \vec{v}_j \vec{v}_i \cdot \hat{r}_{ij} \hat{r}_{ij} \cdot \vec{v}_j \right)$$

$$\begin{aligned}
L = & -\frac{1}{4} \sum_{ij} \frac{m_i m_j}{r_{ij}} \vec{v}_i \cdot \vec{v}_j - \frac{1}{4} \sum_{ijk} \frac{m_i m_j m_k}{r_{ij} r_{ik}} \left(\vec{v}_i \cdot [\vec{v}_j + \vec{v}_k] - 3\vec{v}_j \cdot \vec{v}_k \right) \\
& + b \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{il}} \left(\vec{v}_i \cdot [\vec{v}_j + \vec{v}_k + \vec{v}_l] + \lambda [\vec{v}_j \cdot \vec{v}_k + \vec{v}_j \cdot \vec{v}_l + \vec{v}_k \cdot \vec{v}_l] \right) \\
& + c \sum_{ijkl} \frac{m_i m_j m_k m_l}{r_{ij} r_{ik} r_{jl}} \left(\vec{v}_i \cdot \vec{v}_j + \eta [\vec{v}_i \cdot \vec{v}_k + \vec{v}_j \cdot \vec{v}_l] + \kappa [\vec{v}_i \cdot \vec{v}_l + \vec{v}_j \cdot \vec{v}_k] + \omega [\vec{v}_l \cdot \vec{v}_k] \right) + \dots \\
& 3b(1 + \lambda) = -\frac{1}{8} \quad c(1 + 2\eta + 2\kappa + \omega) = -\frac{3}{16}
\end{aligned}$$

m^7 order($1/r$ 6 order) Potentials

Rules for computers to generate all

m^N order Potentials needed!

