SO(10) Grand Unification from *M*-theory on a G_2 -manifold

In collaboration with

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- 2 Model building in *M*-Theory
 - The framework
 - The G₂-MSSM
- \bigcirc SO(10) models from *M*-Theory
- 4 Conclusions and future work

Outline

Introduction and Motivation

- Model building in *M*-Theory
 - The framework
 - The G₂-MSSM
- 3 *SO*(10) models from *M*-Theory
 - 4 Conclusions and future work

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- These predictions and constraints are usually within the Minimal Supersymmetric Standard Model (MSSM) setup.
- One is led to search for more general SUSY models. A fruitful guideline for this has been given by String Theory phenomenology.
- Nowadays we understand different String Theories to be limits of an underlying 11 dimensional theory, *M*-Theory.

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- Recent developments in *M*-Theory compactified in *G*₂-manifolds have shown that SUSY breaking and moduli stabilisation provide hierarchal effective field theories.
- In this case we also find all the realistic scenario ingredients: N = 1 SUSY, Chiral fermions, Non-abelian gauge theories, etc.

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- *G*₂ holonomy manifolds do not have continuous symmetries but admit discrete symmetries.
- Witten (2001) showed that such symmetries provide a solution for the doublet-triplet splitting problem.

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This breaks the gauge group to H = {g ∈ G, [g, W] = 0} as Ψ^w transforms under the elements of G that commute with W.

W has a topological meaning and furnishes a representation of π₁(K).
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- This provides a mechanism with a solution for the doublet-triplet splitting problem.



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[Acharya, Kane, Kuflik, Lu (2011)]

Consider the SU(5) GUT with the MSSM spectrum, plus the geometric symmetry Z_n .

Let $\overline{\mathbf{5}}^{w} \supset (H_d, \overline{D})$ be localised along the Wilson line. Under the discrete symmetry The GUT multiplets transform

$$\begin{split} \overline{\mathbf{5}}^{w} &\to \eta^{\omega} \left(\eta^{\delta} H_{d}^{w} \oplus \eta^{\gamma} \overline{D}^{w} \right), \\ \mathbf{5}^{h} &\to \eta^{\chi} \mathbf{5}^{h}, \\ \overline{\mathbf{5}}^{m} &\to \eta^{\tau} \overline{\mathbf{5}}^{m}, \\ \mathbf{10}^{m} &\to \eta^{\sigma} \mathbf{10}^{m}, \end{split}$$

where $\eta \equiv e^{2\pi i/n}$, $2\delta + 3\gamma = 0 \mod n$ (for simplicity we will set $\overline{\mathbf{5}}^w$ to $\omega = 0$).

Description	Coupling	Constraint
Down-type Yukawas	$H_{u}^{h} 10^{m} 10^{m}$	$2\sigma + \chi = 0 \mod n$
Up-type Yukawas	$H_d^w 10^m \overline{5}^m$	$\sigma+\tau+\delta=0 \ \mathrm{mod} \ n$
Neutrino Majorana-masses	$H_d^w H_d^w \overline{5}^m \overline{5}^m$	$2\chi + 2\tau = 0 \mod n$
Colour-triplet mass	$\overline{D}^{w}D^{h}$	$\chi + \gamma = 0 \mod n$

A solution exists while forbidding a tree-level μ -term and dimension four and five proton decay operators.

An effective μ -term is generated by moduli vev

$$K \supset \frac{s}{M_{Pl}}H_uH_d + \text{h.c.},$$

$$\mu = \langle m_{3/2} K_{H_u H_d} - F^k K_{H_u H_d k} \rangle \sim 0.1 m_{3/2} \sim \mathcal{O}(\text{TeV}).$$

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 For 10^w

$$\mathbf{10}^{w} o \eta^{\omega} \left(\eta^{-lpha} \mathcal{H}^{w}_{d} \oplus \eta^{eta} \overline{\mathcal{D}}^{w} \oplus \eta^{lpha} \mathcal{H}^{w}_{u} \oplus \eta^{-eta} \mathcal{D}^{w}
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• In minimal SO(10) the μ -term is contained in

$$W \supset \mathbf{10}^{w} \mathbf{10}^{w} = \mu H_{u} H_{d} + m_{D} D \overline{D},$$

but as $\mathcal{W} \in SO(10)$, both terms are invariant under its action.
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$$W \supset \mathbf{H}_{d}^{T} \cdot \mu_{H} \cdot \mathbf{H}_{u} + \overline{\mathbf{D}}^{T} \cdot M_{D} \cdot \mathbf{D},$$

where μ_H and M_D are two 2 × 2 superpotential mass parameters matrices, $\mathbf{H}_{u,d}^T = (H_{u,d}^w, H_{u,d}^h)$, $\overline{\mathbf{D}}^T = (\overline{D}^w, \overline{D}^h)$, and $\mathbf{D}^T = (D^w, D^h)$.

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• There is no choice of vanishing constraints that leaves one eigenvalue of μ_H light while keeping both masses of M_D heavy.

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where matter irreps transform $\mathbf{16}^m \rightarrow \eta^m \mathbf{16}^m$.

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 - Moduli vevs break the discrete symmetry, leading to to proton-decay interactions.
 - The presence of light coloured states will ruin unification.

• Consider the relevant Kahler potential operators

$$\begin{split} & K \supset \frac{s}{M_{Pl}^2} DQQ + \frac{s}{M_{Pl}^2} De^c u^c + \frac{s}{M_{Pl}^2} DNd^c + \\ & + \frac{s}{M_{Pl}^2} \overline{D}d^c u^c + \frac{s}{M_{Pl}^2} \overline{D}QL. \end{split}$$

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• The effective potential may be calculated a la Giudice-Masiero to be

$$W_{eff} \supset \lambda DQQ + \lambda De^{c}u^{c} + \lambda DNd^{c} + \lambda \overline{D}d^{c}u^{c} + \lambda \overline{D}QL,$$

where

$$\lambda \approx \frac{1}{M_{Pl}^2} \left(\langle s \rangle m_{3/2} + \langle F_s \rangle \right) \sim 10^{-14}$$

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• The D triplet decay rate can also be estimated

$$au_D = \Gamma_D^{-1} \sim \left(\lambda^2 m_D\right)^{-1} \sim 0.1 ~{
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which is consistent with BBN constraint.

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- Using the discrete symmetry, we find a solution where d_X^c , $\overline{d^c}_X$ are split from the rest of the family.
- The resulting spectrum is effectively the same as the MSSM with additional vector-like family.



• Such solution exists if either $\mathbf{16}_X$ or $\mathbf{\overline{16}}_X$ absorb Wilson line phases. Considering

$$\begin{aligned} \mathbf{16}_X \to \eta^{\mathsf{x}} \left(\eta^{-3\gamma} L \oplus \eta^{3\gamma+\delta} e^{\mathbf{c}} \oplus \eta^{3\gamma-\delta} \mathsf{N} \oplus \eta^{-\gamma-\delta} u^{\mathbf{c}} \oplus \right. \\ \oplus \eta^{-\gamma+\delta} d^{\mathbf{c}} \oplus \eta^{\gamma} Q \right). \end{aligned}$$

and $\overline{\mathbf{16}}_X \to \eta^{\overline{x}} \overline{\mathbf{16}}_X$, the splitting condition is given by

$$\overline{d^c}_X d^c_X : x - \gamma + \delta + \overline{x} = 0 \mod n,$$

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- The remaining states of $\mathbf{16}_X, \mathbf{\overline{16}}_X$ get a TeV scale μ -term from moduli vevs.
- Unification scale is found to be $M_{
 m GUT} \sim 10^{16}$ GeV, with $\alpha_u^{-1} \sim 9.6$.

• The tree-level superpotential of our model allowed by the discrete symmetry is

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• One can perform an RGE analysis of the spectrum assuming general values for the mass parameters at the GUT scale.



 The addition of an extra family has other benefits for model building: N_X, N
 Vevs (v_x) break the rank of the gauge group and generate Right-handed neutrinos Majorana masses.

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• R-Parity violating (RPV) interactions through the Kahler terms

$$K_{RPV} \supset \alpha \frac{s}{M_{Pl}^3} \mathbf{16}_X \mathbf{16}^m \mathbf{16}^m \mathbf{16}^m + \beta \frac{s}{M_{Pl}^2} \mathbf{10}^w \mathbf{16}_X \mathbf{16}^m.$$

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 - It's not clear what symmetry breaking mechanism can break the rank.
 - Detailed spectrum and Z boson mass prediction study has not been carried out yet.

Thank you!

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- Tree-level terms in the superpotential are functions of distances between the supermultiplets in *K*.
- The superpotential mass parameters, the μ -terms, values are suppressed Planck masses.
- By considering 5^w and 5^h to be near in K then the D^hD^w mass can be GUT/Planck valued while the Higgses don't have a μ-term ⇒ we split the H and D masses.
- On the other hand, direct gaugino searches tell us $\mu \ge 100$ GeV.
- Fortunately there is a natural way of generating an effective μ -term of $\mathcal{O}(\text{TeV})$ in *M*-Theory!
- Discrete symmetry is a geometric symmetry of the extra dimensions, therefore the moduli fields are naturally charged under it. As moduli acquire vevs they break the discrete symmetry.

Backup slides: The G₂-MSSM

 The moduli vev generate an effective μ-term a la Giudice-Masiero from Kahler potential operators of the form

$$K \supset \frac{s}{M_{Pl}}H_uH_d + \text{h.c.}.$$

From standard supergravity calculations we know that

$$\mu = \langle m_{3/2} K_{H_u H_d} - F^k K_{H_u H_d k} \rangle,$$

which leads to

$$\mu \sim \frac{\langle s \rangle}{M_{Pl}} m_{3/2} + \frac{\langle F_s \rangle}{M_{Pl}}.$$

- Developments in M-Theory have shown that the moduli vevs are approximately $\langle s \rangle \sim 0.1 M_{Pl}$, $\langle F_s \rangle \sim m_{1/2} M_{Pl}$, $\mathbf{m}_{3/2} \sim \mathcal{O}(\mathbf{10} \text{ TeV})$.
- In *M*-Theory the gaugino masses are suppressed, so the F-term is subleading and we conclude

$$u \sim 0.1 m_{3/2} \sim \mathcal{O}(\text{TeV}).$$

Backup slides: SO(10) models from *M*-Theory

- To study the potential mixing we focus on the superpotential contributions to the up-type quarks mass matrix.
- Schematically, $W \supset \overline{U} \cdot M_U \cdot U$, where $U^T = (u_i, \overline{u^c}_X, u_X)$, $\overline{U}^T = (u_i^c, \overline{u}_X, u_X^c)$, with i = 1, 2, 3, and

$$M_{U} = \begin{pmatrix} y_{u}^{ij} H_{u} & \mu_{iX} & \lambda_{iX} \\ \mu_{Xj} & \lambda_{XX} & \mu_{XX} \\ \lambda_{Xj} & \mu_{XX} & \lambda_{XX} \end{pmatrix}$$

where y_u are EWS Yukawas, μ -terms are of order TeV, and λ couplings are of order 10^{-14} .

- In the limit $\lambda_{iX}, \lambda_{Xj}, \lambda_{XX} \to 0$ one finds that the determinant of M_U is independent of μ_{iX}, μ_{Xj} , i.e. of the mixing masses.
- To leading order in λ the mass eigenstates do not mix matter with $\overline{16}_{\chi}.$

Backup slides: SO(10) models from *M*-Theory

• RPV interactions arise from the Kahler potential terms and are effectively described by the superpotential

$$W_{RPV}^{eff} \supset \alpha \lambda \frac{v_X}{M_{Pl}} LLe^c + \alpha \lambda \frac{v_X}{M_{Pl}} QLe^c + \alpha \lambda \frac{v_X}{M_{Pl}} u^c d^c d^c + \beta \lambda v_X LH_u.$$

 Due to the nature of SO(10), neutrinos have the same Dirac mass as the up-type quarks. A realistic τ-neutrino physical mass requires

$$M_{{\sf Majorana}}\gtrsim 10^{14}~{
m GeV}\Rightarrow v_X\sim 10^{16}~{
m GeV}.$$

• The effective strength of the first terms is $\lambda v_X / M_{Pl} \sim O(10^{-16})$ (taking $\alpha \sim O(1)$). The resulting LSP lifetime can be estimated as

$$au_{LSP} \simeq rac{10^{-13}\,{
m sec}}{(v_X/M_{Pl})^2} \left(rac{m_0}{10\,\,{
m TeV}}
ight)^4 \left(rac{100\,\,{
m GeV}}{m_{LSP}}
ight)^5 \sim 10^{-13}\,{
m sec},$$

for $m_0 \sim 10$ TeV, $m_{LSP} \sim 100$ GeV.

Backup slides: SO(10) models from *M*-Theory

• The last term in the effective superpotential

 $\beta \lambda v_x L H_u$,

is constrained by neutrino mass limit exclusions.

• The bilinear contribution to the mixing to be at most of

$$\beta \lambda v_X \lesssim \mathcal{O}(10^{-3} \text{ TeV}),$$

which means leads to the upper bound

$$v_X \lesssim 10^{14}/eta~{
m GeV},$$

and one has to consider there is some suppression by β of order $\mathcal{O}(10^{-2})$.