

High-energy resummation effects in Mueller-Navelet jet production at the LHC

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Orsay

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in collaboration with

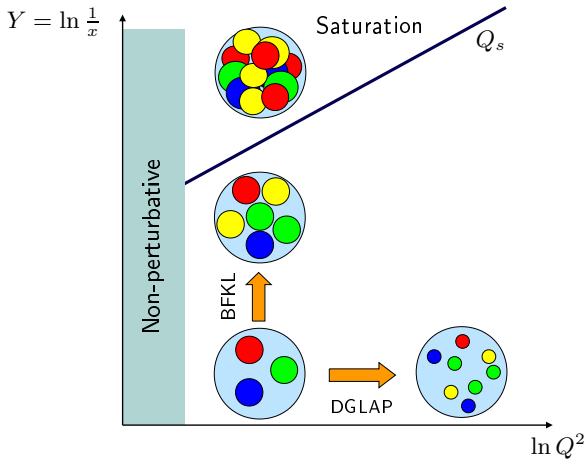
D. Colferai (Florence U. & INFN, Florence), B. Ducloué (Jyväskylä U.),
F. Schwennsen (DESY), S. Szymanowski (NCBJ, Warsaw)

D. Colferai, F. Schwennsen, L. Szymanowski, S. W., JHEP 1012 (2010) 026 [arXiv:1002.1365 [hep-ph]]

B. Ducloué, L. Szymanowski, S. W., JHEP 1305 (2013) 096 [arXiv:1302.7012 [hep-ph]]

B. Ducloué, L. Szymanowski, S. W., PRL 112 (2014) 082003 [arXiv:1309.3229 [hep-ph]]

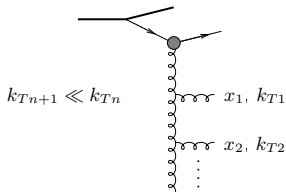
The different regimes of QCD



Resummation in QCD: DGLAP vs BFKL

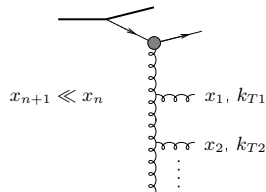
Small values of α_s (perturbation theory applies if there is a hard scale) can be compensated by large logarithmic enhancements.

DGLAP

strong ordering in k_T

$$\sum (\alpha_s \ln Q^2)^n$$

BFKL

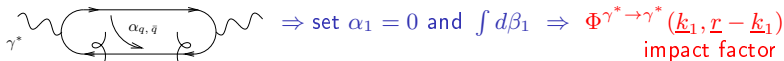
strong ordering in x

$$\sum (\alpha_s \ln s)^n$$

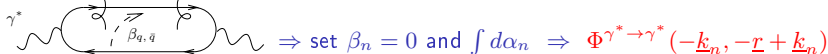
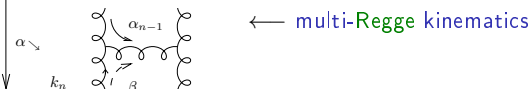
When \sqrt{s} becomes very large, it is expected that a BFKL description is needed to get accurate predictions

Opening the boxes: Impact representation $\gamma^* \gamma^* \rightarrow \gamma^* \gamma^*$ as an example

- Sudakov decomposition: $k_i = \alpha_i p_1 + \beta_i p_2 + k_{\perp i}$ ($p_1^2 = p_2^2 = 0$, $2p_1 \cdot p_2 = s$)
- write $d^4 k_i = \frac{s}{2} d\alpha_i d\beta_i d^2 k_{\perp i}$ ($\underline{k} = \text{Eucl.} \leftrightarrow k_{\perp} = \text{Mink.}$)
- t -channel gluons have non-sense polarizations at large s : $\epsilon_{NS}^{up/down} = \frac{2}{s} p_{2/1}$



$$\mathcal{M} = \frac{is}{(2\pi)^2} \int \frac{d^2 \underline{k}}{\underline{k}^2} \Phi^{up}(\underline{k}, \underline{r} - \underline{k}) \int \frac{d^2 \underline{k}'}{\underline{k}'^2} \Phi^{down}(-\underline{k}', -\underline{r} + \underline{k}') \\ \times \int_{\delta - i\infty}^{\delta + i\infty} \frac{d\omega}{2\pi i} \left(\frac{s}{s_0}\right)^\omega G_\omega(\underline{k}, \underline{k}', \underline{r})$$



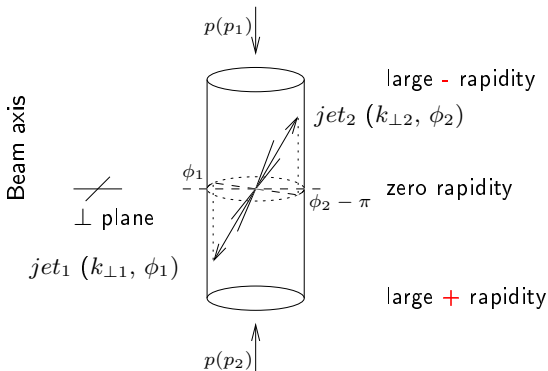
Higher order corrections

- Higher order corrections to BFKL kernel are known at NLL order (Lipatov Fadin; Camici, Ciafaloni), now for arbitrary impact parameter $\alpha_S \sum_n (\alpha_S \ln s)^n$ resummation
- impact factors are known in some cases at NLL
 - $\gamma^* \rightarrow \gamma^*$ at $t = 0$ (Bartels, Colferai, Gieseke, Kyrieleis, Qiao; Balitski, Chirilli)
 - forward jet production (Bartels, Colferai, Vacca; Caporale, Ivanov, Murdaca, Papa, Perri; Chachamis, Hentschinski, Madrigal, Sabio Vera)
 - inclusive production of a pair of hadrons separated by a large interval of rapidity (Ivanov, Papa)
 - $\gamma_L^* \rightarrow \rho_L$ in the forward limit (Ivanov, Kotsky, Papa)

Mueller-Navelet jets: Basics

Mueller-Navelet jets

- Consider two jets (hadrons flying within a narrow cone) **separated by a large rapidity**, i.e. each of them almost fly in the direction of the hadron “close” to it, and with very similar transverse momenta
- Pure LO *collinear* treatment: these two jets should be emitted **back to back** at leading order: $\Delta\phi - \pi = 0$ ($\Delta\phi = \phi_1 - \phi_2 =$ relative azimuthal angle) and $k_{\perp 1} = k_{\perp 2}$. No phase space for (untagged) emission between them



Master formulas

 k_T -factorized differential cross section

$$\frac{d\sigma}{d|\mathbf{k}_{J1}| d|\mathbf{k}_{J2}| dy_{J1} dy_{J2}} = \int d\phi_{J1} d\phi_{J2} \int d^2\mathbf{k}_1 d^2\mathbf{k}_2$$

$$\times \Phi(\mathbf{k}_{J1}, x_{J1}, -\mathbf{k}_1)$$

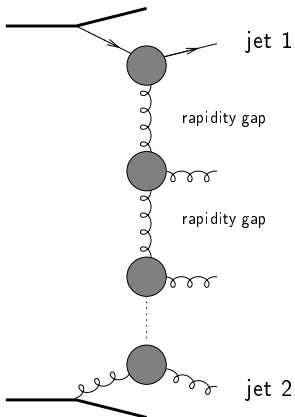
$$\times G(\mathbf{k}_1, \mathbf{k}_2, \hat{s})$$

$$\times \Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2)$$

with $\Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2) = \int dx_2 f(x_2) V(\mathbf{k}_2, x_2)$ $f \equiv \text{PDF}$ $x_J = \frac{|\mathbf{k}_J|}{\sqrt{s}} e^{y_J}$

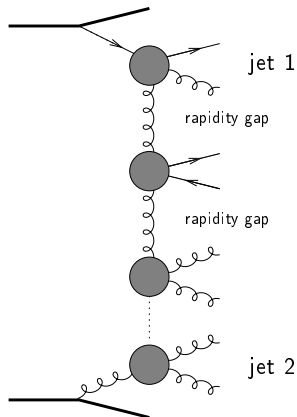
Mueller-Navelet jets: LL vs NLL

LL BFKL



$$\sum(\alpha_s \ln s)^n$$

NLL BFKL



$$\sum(\alpha_s \ln s)^n + \alpha_s \sum(\alpha_s \ln s)^n$$

Results for a symmetric configuration

In the following we show results for

- $\sqrt{s} = 7 \text{ TeV}$
- $35 \text{ GeV} < |\mathbf{k}_{J1}|, |\mathbf{k}_{J2}| < 60 \text{ GeV}$
- $0 < |y_1|, |y_2| < 4.7$

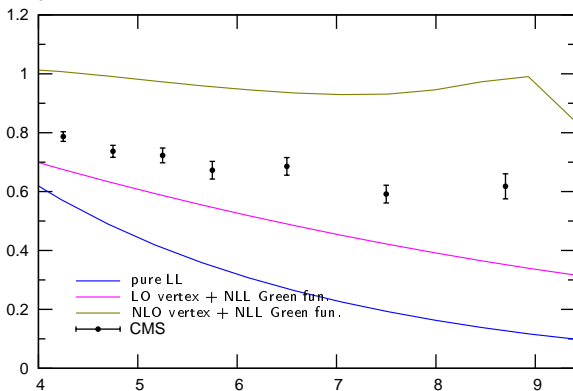
These cuts allow us to compare our predictions with the first experimental data on azimuthal correlations of **Mueller-Navelet** jets at the LHC presented by the **CMS** collaboration (CMS-PAS-FSQ-12-002)

note: unlike experiments we have to set an upper cut on $|\mathbf{k}_{J1}|$ and $|\mathbf{k}_{J2}|$. We have checked that our results do not depend on this cut significantly.

Results: azimuthal correlations

Azimuthal correlation $\langle \cos \varphi \rangle$

$$\frac{c_1}{c_0} = \langle \cos \varphi \rangle \equiv \langle \cos(\phi_{J1} - \phi_{J2} - \pi) \rangle$$

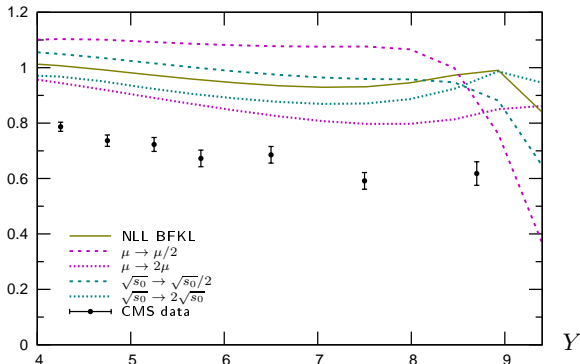


The NLO corrections to the jet vertex lead to a large increase of the correlation

Results: azimuthal correlations

Azimuthal correlation $\langle \cos \varphi \rangle$

$$\langle \cos \varphi \rangle \equiv \langle \cos(\phi_{J1} - \phi_{J2} - \pi) \rangle$$



$$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$$

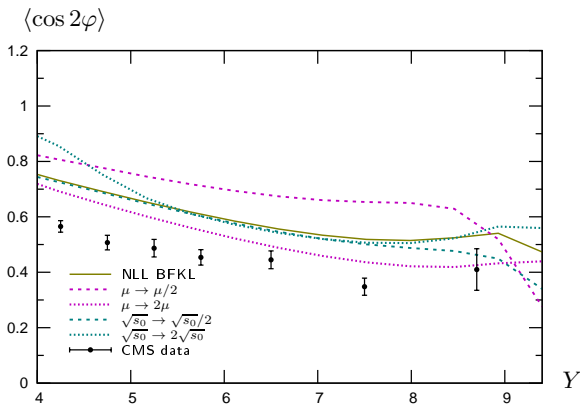
$$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$$

$$0 < |y_1| < 4.7$$

$$0 < |y_2| < 4.7$$

- NLL BFKL predicts a too small decorrelation
- The NLL BFKL calculation is still rather dependent on the scales, especially the renormalization / factorization scale

Results: azimuthal correlations

Azimuthal correlation $\langle \cos 2\varphi \rangle$ 

$$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$$

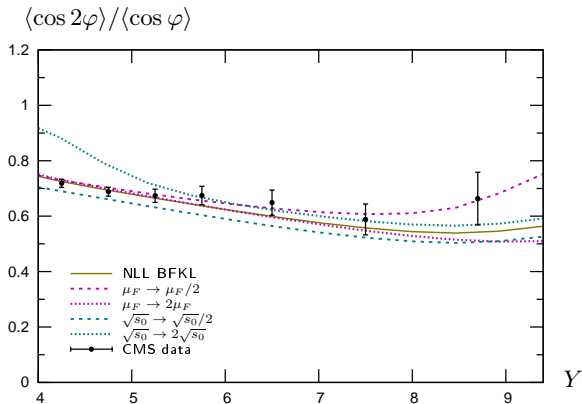
$$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$$

$$0 < |y_1| < 4.7$$

$$0 < |y_2| < 4.7$$

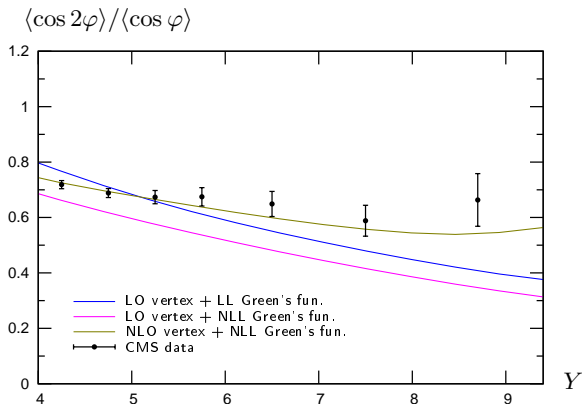
- The agreement with data is a little better for $\langle \cos 2\varphi \rangle$ but still not very good
- This observable is also very sensitive to the scales

Results: azimuthal correlations

Azimuthal correlation $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  $35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$ $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$ $0 < |y_1| < 4.7$ $0 < |y_2| < 4.7$

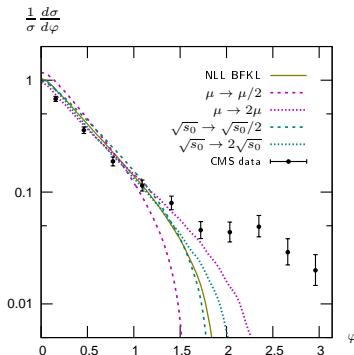
- This observable is more stable with respect to the scales than the previous ones
- The agreement with data is good across the whole Y range

Results: azimuthal correlations

Azimuthal correlation $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  $35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$ $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$ $0 < |y_1| < 4.7$ $0 < |y_2| < 4.7$

It is necessary to include the NLO corrections to the jet vertex to reproduce the behavior of the data at large Y

Results: azimuthal distribution

Azimuthal distribution (integrated over $6 < Y < 9.4$)

$$\frac{1}{\sigma} \frac{d\sigma}{d\varphi} = \frac{1}{2\pi} \left\{ 1 + 2 \sum_{n=1}^{\infty} \cos(n\varphi) \langle \cos(n\varphi) \rangle \right\}.$$

- Our calculation predicts a too large value of $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$ for $\varphi \lesssim \frac{\pi}{2}$ and a too small value for $\varphi \gtrsim \frac{\pi}{2}$
- It is not possible to describe the data even when varying the scales by a factor of 2

Results: limitations

- The agreement of our calculation with the data for $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ is good and quite stable with respect to the scales
- The agreement for $\langle \cos n\varphi \rangle$ and $\frac{1}{\sigma} \frac{d\sigma}{d\varphi}$ is not very good and very sensitive to the choice of the renormalization scale μ_R
- An all-order calculation would be independent of the choice of μ_R . This feature is lost if we truncate the perturbative series
 ⇒ How to choose the renormalization scale?
 'Natural scale': sometimes the typical momenta in a loop diagram are different from the natural scale of the process

We decided to use the **Brody-Lepage-Mackenzie (BLM)** procedure to fix the renormalization scale

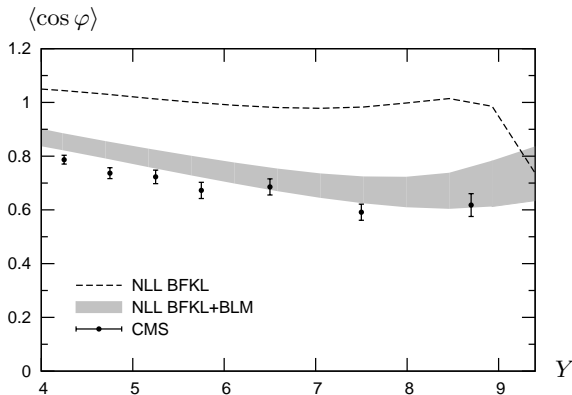
The BLM renormalization scale fixing procedure

The **Brodsky-Lepage-Mackenzie (BLM)** procedure resums the self-energy corrections to the gluon propagator at one loop into the running coupling.

First attempts to apply **BLM** scale fixing to **BFKL** processes lead to problematic results. **Brodsky, Fadin, Kim, Lipatov and Pivovarov** suggested that one should first go to a physical renormalization scheme like MOM and then apply the 'traditional' **BLM** procedure, i.e. identify the β_0 dependent part and choose μ_R such that it vanishes.

We followed this prescription for the full amplitude at NLL.

Results with BLM

Azimuthal correlation $\langle \cos \varphi \rangle$ 

$$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$$

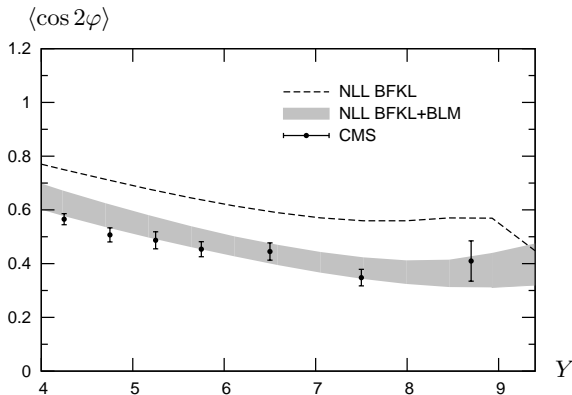
$$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$$

$$0 < |y_1| < 4.7$$

$$0 < |y_2| < 4.7$$

Using the BLM scale setting, the agreement with data becomes much better

Results with BLM

Azimuthal correlation $\langle \cos 2\varphi \rangle$ 

$$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$$

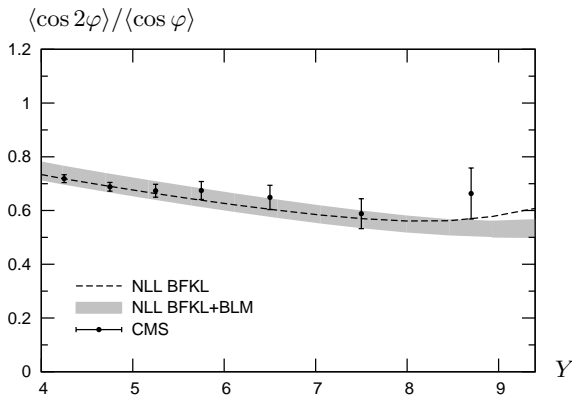
$$35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$$

$$0 < |y_1| < 4.7$$

$$0 < |y_2| < 4.7$$

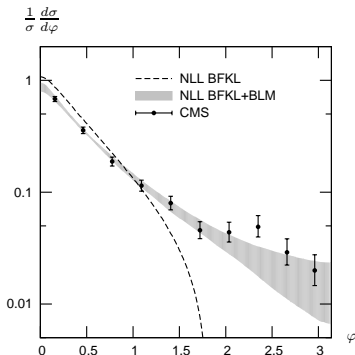
Using the BLM scale setting, the agreement with data becomes much better.

Results with BLM

Azimuthal correlation $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$  $35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$ $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$ $0 < |y_1| < 4.7$ $0 < |y_2| < 4.7$

Because it is much less dependent on the scales, the observable $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ is almost not affected by the BLM procedure and is still in good agreement with the data.

Results with BLM

Azimuthal distribution (integrated over $6 < Y < 9.4$)

With the BLM scale setting the azimuthal distribution is in good agreement with the data across the full φ range.

Comparison with fixed-order

Using the **BLM** scale setting:

- The agreement $\langle \cos n\varphi \rangle$ with the data becomes much better
- The agreement for $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ is still good and unchanged as this observable is weakly dependent on μ_R
- The azimuthal distribution is in much better agreement with the data

But the configuration chosen by **CMS** with $\mathbf{k}_{J_{\min 1}} = \mathbf{k}_{J_{\min 2}}$ does not allow us to compare with a **fixed-order** $\mathcal{O}(\alpha_s^3)$ treatment (i.e. without resummation)

These calculations are unstable when $\mathbf{k}_{J_{\min 1}} = \mathbf{k}_{J_{\min 2}}$ because the cancellation of some divergencies is difficult to obtain numerically

Comparison with fixed-order

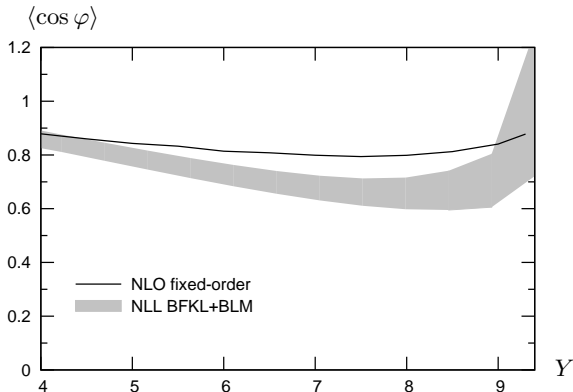
Results for an asymmetric configuration

In this section we choose the cuts as

- $35 \text{ GeV} < |\mathbf{k}_{J1}|, |\mathbf{k}_{J2}| < 60 \text{ GeV}$
- $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$
- $0 < |y_1|, |y_2| < 4.7$

and we compare our results with the NLO fixed-order code Dijet ([Aurenche, Basu, Fontannaz](#)) in the same configuration

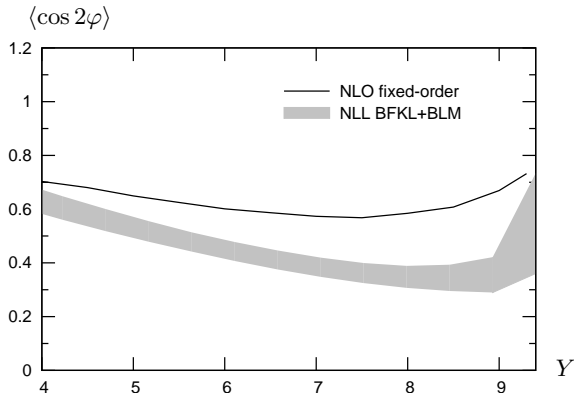
Comparison with fixed-order

Azimuthal correlation $\langle \cos \varphi \rangle$ 

$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$
 $0 < |y_1| < 4.7$
 $0 < |y_2| < 4.7$

The NLO fixed-order and NLL BFKL+BLM calculations are very close

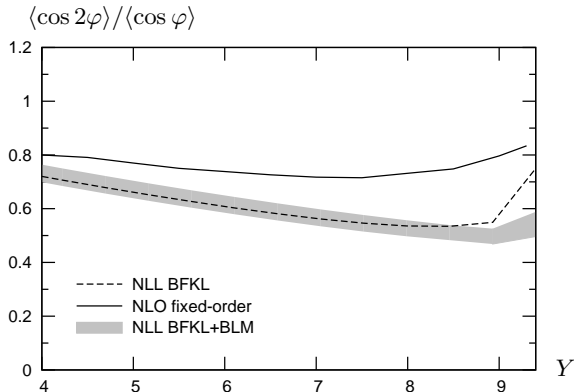
Comparison with fixed-order

Azimuthal correlation $\langle \cos 2\varphi \rangle$ 

$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$
 $0 < |y_1| < 4.7$
 $0 < |y_2| < 4.7$

The BLM procedure leads to a sizable difference between NLO fixed-order and NLL BFKL+BLM.

Comparison with fixed-order

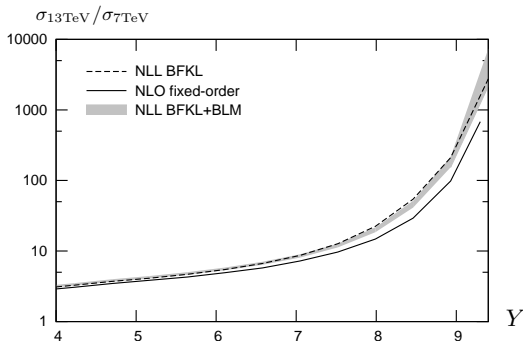
Azimuthal correlation $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ 

$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$
 $0 < |y_1| < 4.7$
 $0 < |y_2| < 4.7$

Using **BLM** or not, there is a **sizable difference** between **BFKL** and fixed-order.

Comparison with fixed-order

Cross section: 13 TeV vs. 7 TeV



$35 \text{ GeV} < |\mathbf{k}_{J1}| < 60 \text{ GeV}$
 $35 \text{ GeV} < |\mathbf{k}_{J2}| < 60 \text{ GeV}$
 $50 \text{ GeV} < \text{Max}(|\mathbf{k}_{J1}|, |\mathbf{k}_{J2}|)$
 $0 < |y_1| < 4.7$
 $0 < |y_2| < 4.7$

- In a **BFKL** treatment, a **strong rise of the cross section with increasing energy** is expected.
- This rise is faster than in a fixed-order treatment

Energy-momentum conservation

- It is necessary to have $\mathbf{k}_{J_{\min 1}} \neq \mathbf{k}_{J_{\min 2}}$ for comparison with fixed order calculations but this can be problematic for **BFKL** because of energy-momentum conservation
- There is no strict energy-momentum conservation in **BFKL**
- This was studied at LO by **Del Duca and Schmidt**. They introduced an effective rapidity Y_{eff} defined as

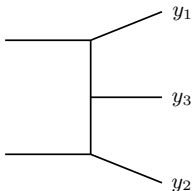
$$Y_{\text{eff}} \equiv Y \frac{\sigma^{2 \rightarrow 3}}{\sigma^{\text{BFKL}, \mathcal{O}(\alpha_s^3)}}$$

- When one replaces Y by Y_{eff} in the expression of σ^{BFKL} and truncates to $\mathcal{O}(\alpha_s^3)$, the exact $2 \rightarrow 3$ result is obtained

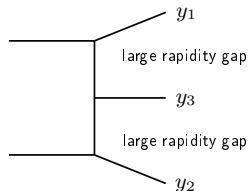
Energy-momentum conservation

We follow the idea of **Del Duca and Schmidt**, adding the NLO jet vertex contribution:

exact $2 \rightarrow 3$

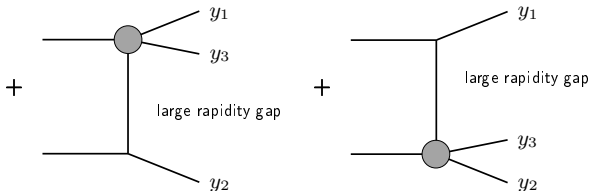


BFKL



one emission from the Green's function + LO jet vertex

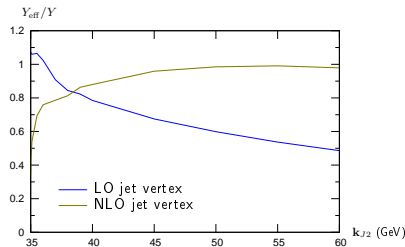
we have to take into account these additional $\mathcal{O}(\alpha_s^3)$ contributions:



no emission from the Green's function + NLO jet vertex

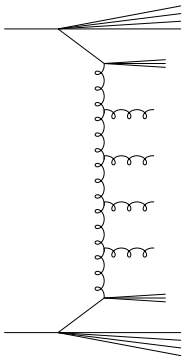
Energy-momentum conservation

Variation of Y_{eff}/Y as a function of k_{J2} for fixed $k_{J1} = 35$ GeV (with $\sqrt{s} = 7$ TeV, $Y = 8$):



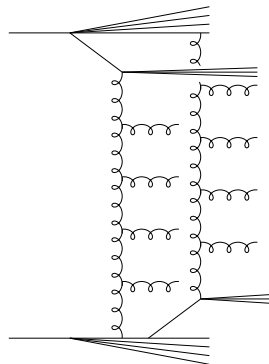
- With the **LO** jet vertex, Y_{eff} is much smaller than Y when k_{J1} and k_{J2} are significantly different
- This is the region important for comparison with fixed order calculations
- The improvement coming from the **NLO** jet vertex is very large in this region
- For $k_{J1} = 35$ GeV and $k_{J2} = 50$ GeV, typical of the values we used for comparison with fixed order, we get $\frac{Y_{\text{eff}}}{Y} \simeq 0.98$ at NLO vs. ~ 0.6 at LO

Can Mueller-Navelet jets be a manifestation of multiparton interactions?



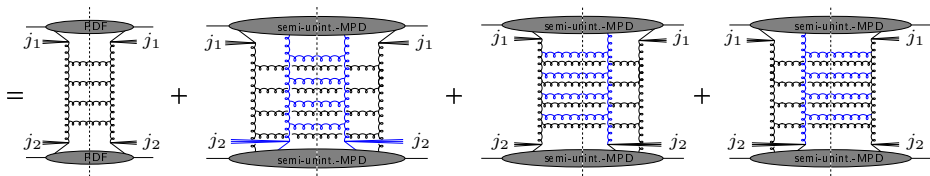
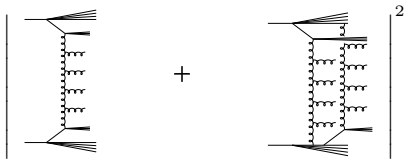
MN jets in the single partonic model

+



MN jets in MPI

Can Mueller-Navelet jets be a manifestation of multiparton interactions?



single \mathbb{P} ladder

two \mathbb{P} ladders

interferences

scaling: $s^{\alpha_{\mathbb{P}}}$

(??) $s^{2\alpha_{\mathbb{P}}}$

??

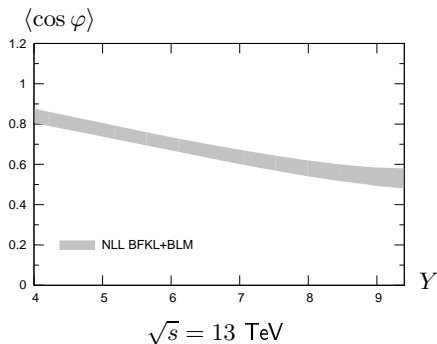
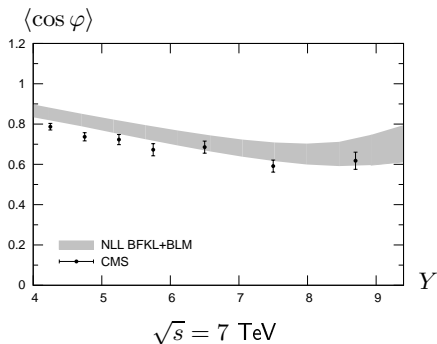
- The twist counting is not easy for MPI kinds of contributions at small x
- $k_{\perp 1,2}$ are not integrated \Rightarrow MPI may be competitive, and enhanced by small- x resummation
- Interference terms are not governed by **BJKP** (this is not a fully interacting 3-reggeons system) (for **BJKP**, $\alpha_{\mathbb{P}} < 1 \Rightarrow$ suppressed)

Conclusions

- We studied **Mueller-Navelet** jets at full (vertex + Green's function) **NLL BFKL** accuracy and compared our results with the first data from the **LHC**
- The agreement with **CMS** data at 7 TeV is greatly improved by using the **BLM** scale fixing procedure
- $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ is almost not affected by **BLM** and shows a clear difference between **NLO fixed-order** and **NLL BFKL** in an **asymmetric configuration**
Energy-momentum conservation seems to be less severely violated with the NLO jet vertex
- We did the same analysis at 13 TeV:
 - Azimuthal decorrelations don't show a very different behavior at 13 TeV compared to 7 TeV
 - **NLL BFKL** predicts a stronger rise of the cross section with increasing energy than a **NLO fixed-order** calculation**A measurement of the cross section at $\sqrt{s} = 7$ or 8 TeV would be needed to test this**

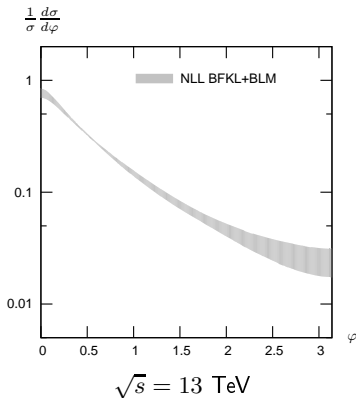
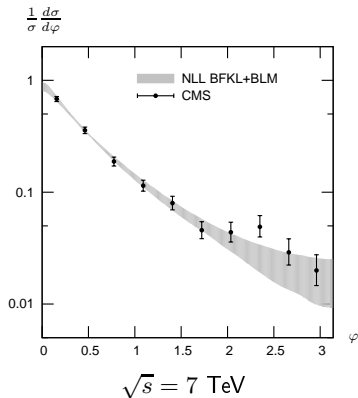
Backup

Azimuthal correlation $\langle \cos \varphi \rangle$



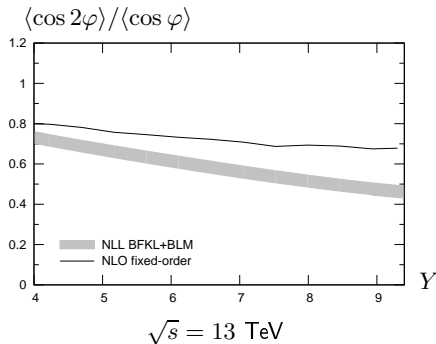
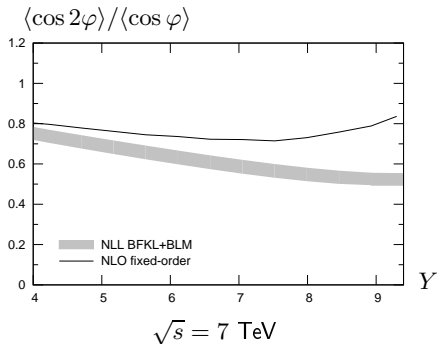
The behavior is similar at 13 TeV and at 7 TeV

Azimuthal distribution (integrated over $6 < Y < 9.4$)



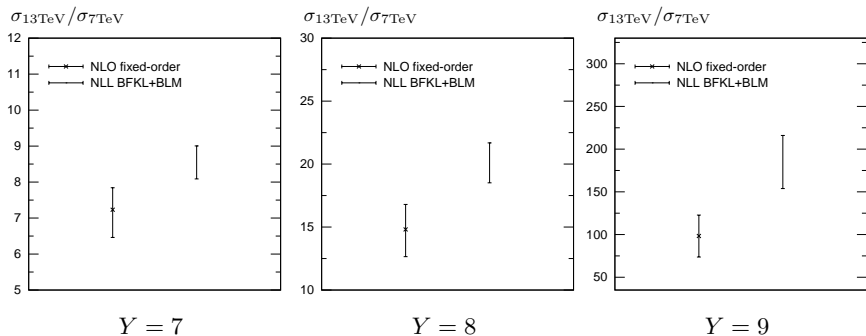
The behavior is similar at 13 TeV and at 7 TeV

Azimuthal correlation $\langle \cos 2\varphi \rangle / \langle \cos \varphi \rangle$ (asymmetric configuration)



The difference between BFKL and fixed-order is smaller at 13 TeV than at 7 TeV

Cross section



It is useful to define the coefficients \mathcal{C}_n as

$$\mathcal{C}_n \equiv \int d\phi_{J1} d\phi_{J2} \cos(n(\phi_{J1} - \phi_{J2} - \pi)) \\ \times \int d^2\mathbf{k}_1 d^2\mathbf{k}_2 \Phi(\mathbf{k}_{J1}, x_{J1}, -\mathbf{k}_1) G(\mathbf{k}_1, \mathbf{k}_2, \hat{s}) \Phi(\mathbf{k}_{J2}, x_{J2}, \mathbf{k}_2)$$

- $n = 0 \implies$ differential cross-section

$$\mathcal{C}_0 = \frac{d\sigma}{d|\mathbf{k}_{J1}| d|\mathbf{k}_{J2}| dy_{J1} dy_{J2}}$$

- $n > 0 \implies$ azimuthal decorrelation

$$\frac{\mathcal{C}_n}{\mathcal{C}_0} = \langle \cos(n(\phi_{J,1} - \phi_{J,2} - \pi)) \rangle \equiv \langle \cos(n\varphi) \rangle$$

- sum over $n \implies$ azimuthal distribution

$$\frac{1}{\sigma} \frac{d\sigma}{d\varphi} = \frac{1}{2\pi} \left\{ 1 + 2 \sum_{n=1}^{\infty} \cos(n\varphi) \langle \cos(n\varphi) \rangle \right\}$$