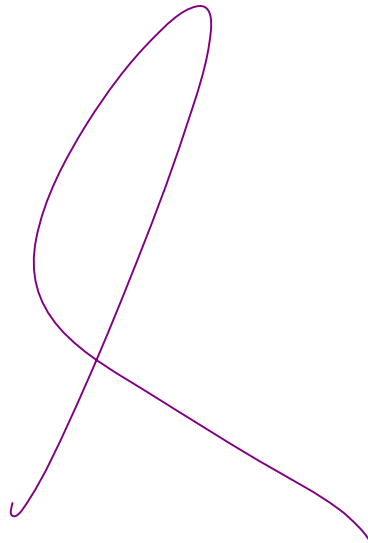


Note Title

3/23/2015



# Spinodal Effects During Inflation

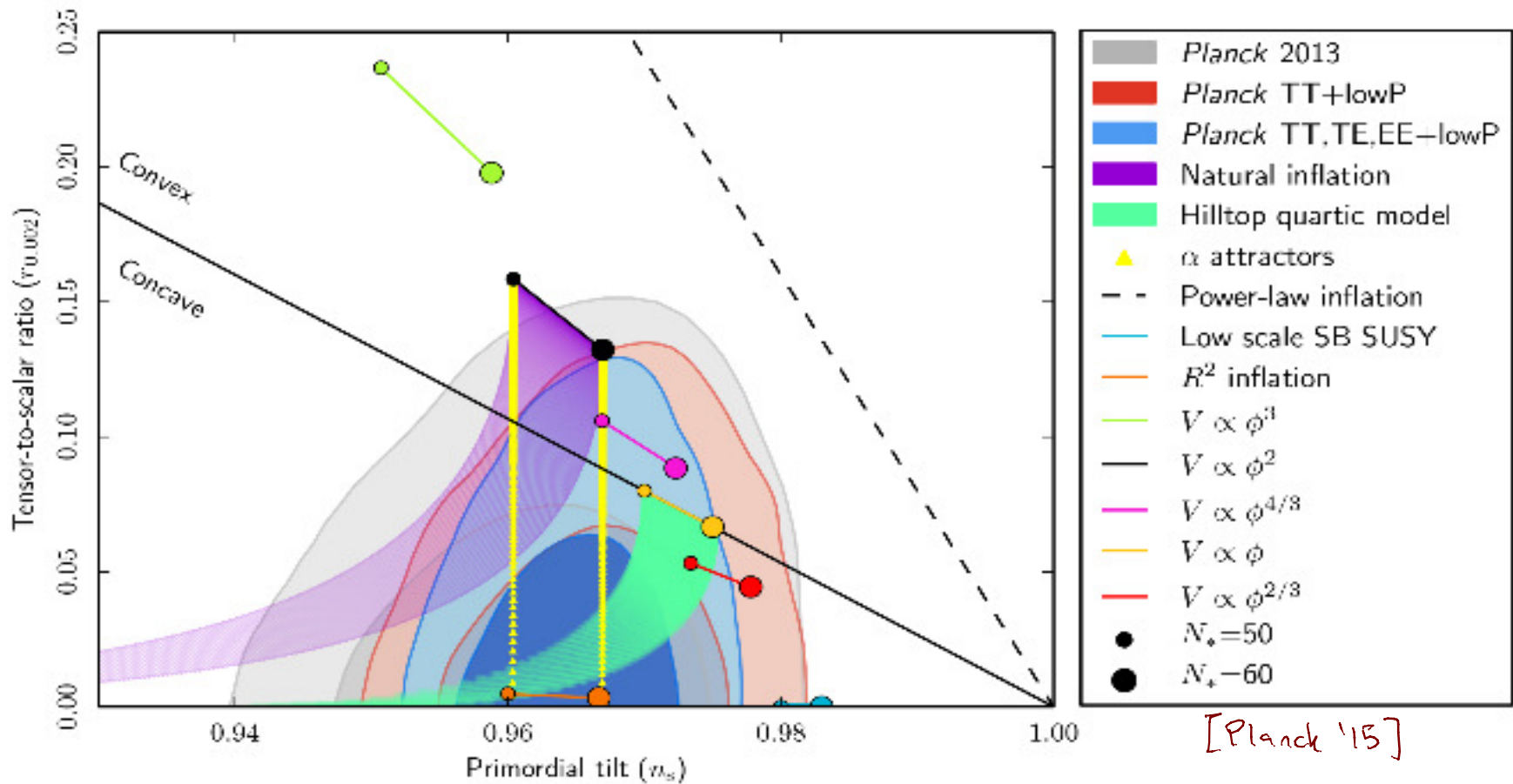
McCullen Sandora

CP<sup>3</sup> Origins

Cosmology & Particle Physics



[ arxiv: 1504.03332 ]  
w/ B. Richard ]



$$r = 8 \frac{V_{,12}}{V_{,2}^2}$$

$$n_s - 1 = -3 \frac{V_{,12}}{V_{,2}^2} + 2 \frac{V_{,11}}{V_{,2}}$$

$$r < .08 \Rightarrow V'' < 0$$

# Outline

↳ Tachyonic Inflaton  $\Rightarrow$  IR instability

↳ Power can grow and influence dynamics

↳ Can encapsulate in a two field model [Cormier, Holman '99]

↳ Diakyon field  $\sigma \leftrightarrow$  long wavelength modes

$$V^H(\bar{\phi}, \sigma) = \int_{\mathbb{R}} d\psi \frac{1}{\sqrt{4\pi}\sigma} e^{-\frac{\psi^2}{2\sigma^2}} V(\bar{\phi} + \psi)$$

↳ Many implications to explore

## Setup

Action  $S(\phi)$ ,  $\phi = \bar{\phi}(t) + \psi$

Expansion history

perturbation spectrum

$$\left. \begin{aligned} \ddot{\bar{\phi}} + 3H\dot{\bar{\phi}} + V'(\bar{\phi}) &= 0 \\ (-\square + V''(\bar{\phi}))\psi &= 0 \end{aligned} \right\} \text{lowest order}$$

However:

$V'' < 0 \Rightarrow$  tachyon  
Long wavelength modes can become large  
can backreact on BOTH equations!

## Path Integral

$$Z = \int \mathcal{D}\psi e^{iS(\bar{\psi} + \psi) + i \int dx \partial_x \psi_x - \frac{i}{2} \int dx dy K_{xy} \psi_x \psi_y}$$

$$\text{Tadpole: } \frac{\delta}{\delta j_x} Z \Big|_{j=k=0} = 0$$

$$\text{Gap: } \frac{\delta}{\delta K_{xy}} Z \Big|_{j=k=0} = 0$$

involve  $\langle \psi^n \rangle$ .

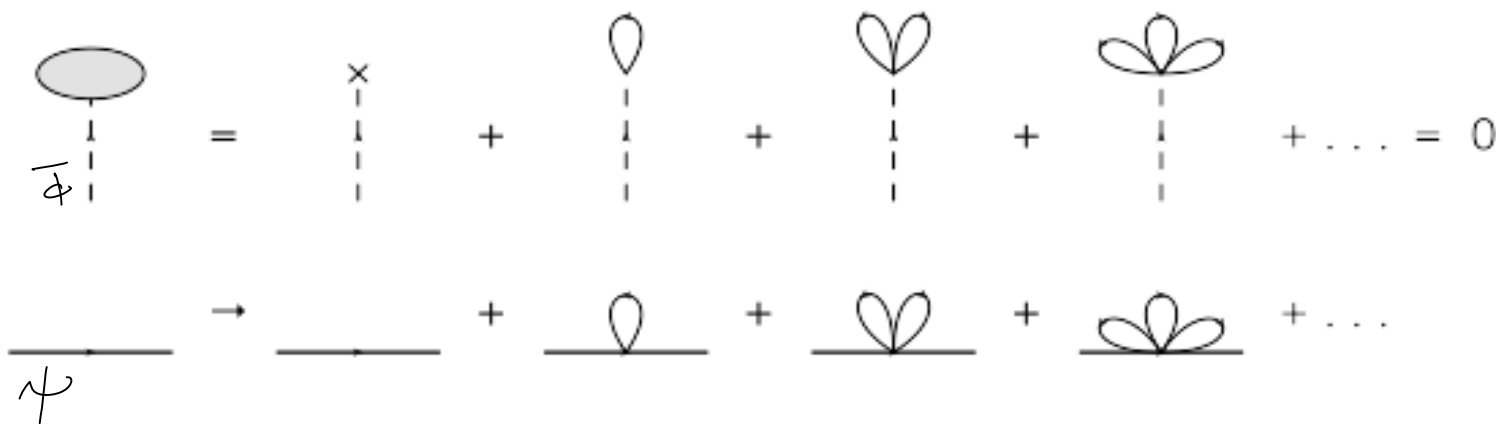
$\cap$  arbitrary:

impossible to solve

# Hartree approximation:

$$\langle \psi^{2n} \rangle = \# \langle \psi^2 \rangle^n, \quad \langle \psi^{2n+1} \rangle = \# \langle \psi^2 \rangle^n \psi$$

truncates system to just  $\bar{\phi}$  and  $\langle \psi^2 \rangle$ .



↻ (& iterate)

## Derivation

$$S(\phi) = \bar{S}_0 + \bar{S}_1 \phi + \bar{S}_2 \phi^2 + S_{int}$$

$$\langle F \rangle_0 = \int \mathcal{D}\psi F e^{i\bar{S}_2 \phi^2}$$

$$\left. \frac{\delta Z}{\delta j_x} \right|_{j=0} = \langle \psi_x \rangle = e^{i\bar{S}_0} \langle \psi_x e^{i\frac{\delta L}{\delta \phi} \phi + i\bar{S}_{int}} \rangle_0$$

$$\approx i e^{i\bar{S}_0} \int d^4 y \left[ \frac{\delta L}{\delta \phi} \langle \psi_x \psi_y \rangle_0 + \langle \psi_x V_{int}(\psi_y) \rangle_0 \right]$$

$$= i e^{i\bar{S}_0} \int d^4 y G_{xy} \left[ \frac{\delta L}{\delta \phi_y} + \langle V_{int}'(\psi_y) \rangle_0 \right] = 0$$



## The diakon

Long wavelength modes important

$$\langle \psi_x \psi_y \rangle_0 \rightarrow \sigma(t)^2 \quad [\text{Cormier, Holman}]$$

$$V^H(\bar{\Phi}, \sigma) \equiv \langle V(\bar{\Phi} + \psi) \rangle_0 = \int_{\mathcal{R}} d\psi \frac{1}{\sqrt{2\pi} \sigma} e^{-\frac{\psi^2}{2\sigma^2}} V(\bar{\Phi} + \psi)$$

$$\left. \begin{array}{l} \text{Tadpole: } (\partial_t^2 + 3H\partial_t) \bar{\Phi} + V_{\bar{\Phi}}^H = 0 \\ \text{Gap: } (\partial_t^2 + 3H\partial_t) \sigma + V_{\sigma}^H = 0 \end{array} \right\} \rightarrow \text{EOMs for } \bar{\Phi} \text{ \& } \sigma$$

Properties of  $V^H(\bar{\phi}, \sigma) = \int_{\mathbb{R}} d\psi \frac{1}{\sqrt{2\pi}\sigma} e^{-\frac{\psi^2}{2\sigma^2}} V(\bar{\phi} + \psi)$

---

$$V^H(\phi, 0) = V(\phi)$$

$$\sigma \partial_{\phi}^2 V^H = \partial_{\sigma} V^H$$

$\Rightarrow M_{\sigma}^2 = M_{\phi}^2 + \sigma^2 \partial_{\phi}^4 V^H$   $\sigma$  also slow rolls!

$\sigma \ll M_p \Rightarrow \sigma$  eqn integrable  $\sigma(N) = \sigma_0 \frac{V_{\phi}(N)}{V_{\phi}(0)}$

## Perturbations

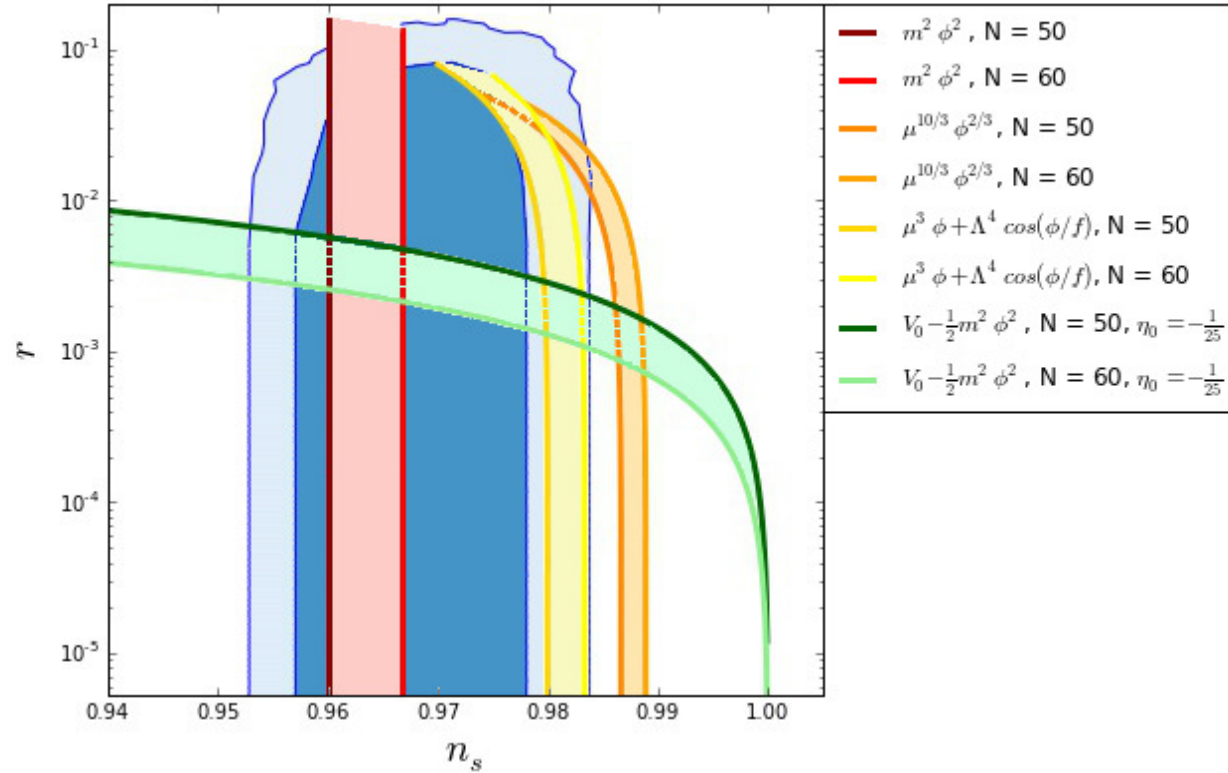
$$P_k = \frac{\langle (V(\phi+\psi) - V(\phi))^2 \rangle_0}{\hbar} \approx \left( \frac{V''}{V_\phi'^2 + V_\sigma'^2} \right)^2 \langle V_\phi'^2 \psi^2 \rangle_0$$

$$\sigma \ll M_p: \quad P_k \approx \frac{V''^2}{V_\phi'^2} \langle \psi^2 \rangle_0 = \frac{V''^2}{V_\phi'^2} \sigma^2$$

$$n_s^{-1} = -\epsilon \left( 4 + \frac{2}{1 + \frac{\sigma^2}{M_p^2}} \right) + 2\gamma \frac{1}{1 + \frac{\sigma^2}{M_p^2}} \rightarrow \begin{cases} -6\epsilon + 2\gamma & \sigma \rightarrow H \\ -4\epsilon & \sigma \gg H \end{cases}$$

$$r = \frac{16\epsilon}{1 + \frac{\sigma^2}{M_p^2}} \rightarrow \begin{cases} 16\epsilon & \sigma \rightarrow H \\ 0 & \sigma \gg H \end{cases}$$

# Specific Models

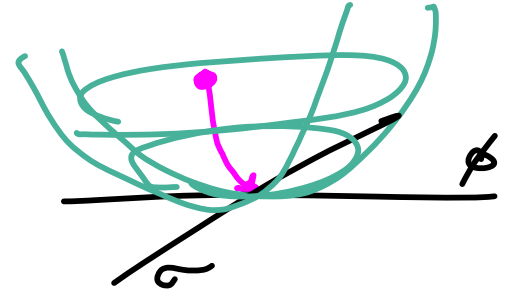


$m^2 \phi^2$ , hilltop, flattened, monodromy

$$V = M^2 \phi^2$$

$$V^H = \frac{1}{2} M^2 (\phi^2 + \sigma^2)$$

$$N_{\text{CMB}} = \frac{\phi^2 + \sigma^2}{4}$$



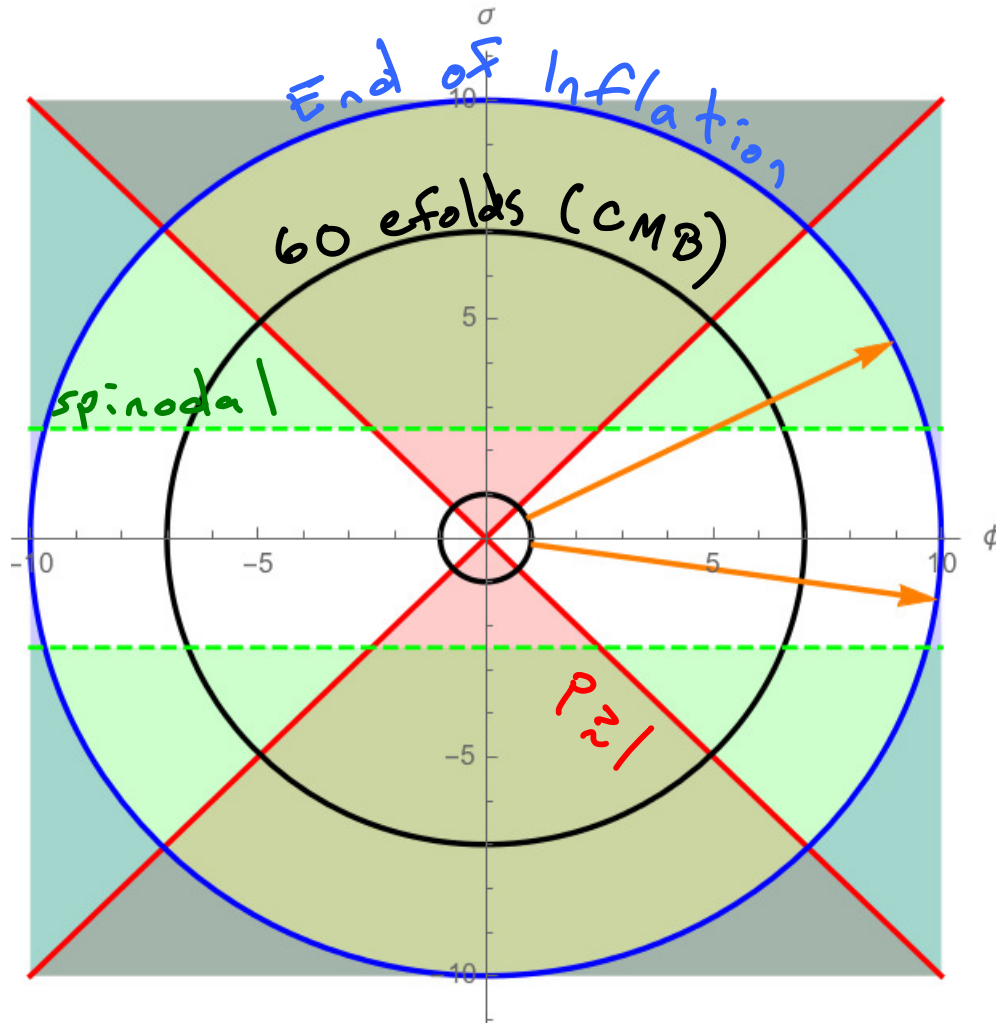
$$P_k = N_{\text{CMB}} \sigma^2 * \left\{ \begin{array}{l} \text{Amplitude set by } \sigma, \\ \text{NOT potential.} \\ P_k \sim 10^{-9} \Rightarrow \sigma_0 \sim 10^{-5} M_p \end{array} \right.$$

$$n_s - 1 = \frac{-2}{N_{\text{CMB}}} * \text{tilt unchanged}$$

$$r \propto \frac{M^2}{\sigma^2} * \text{small } m \Rightarrow \text{small } r$$

Hilltop

$$V^H = V_0 - \frac{1}{2} m^2 (\phi^2 + \sigma^2)$$



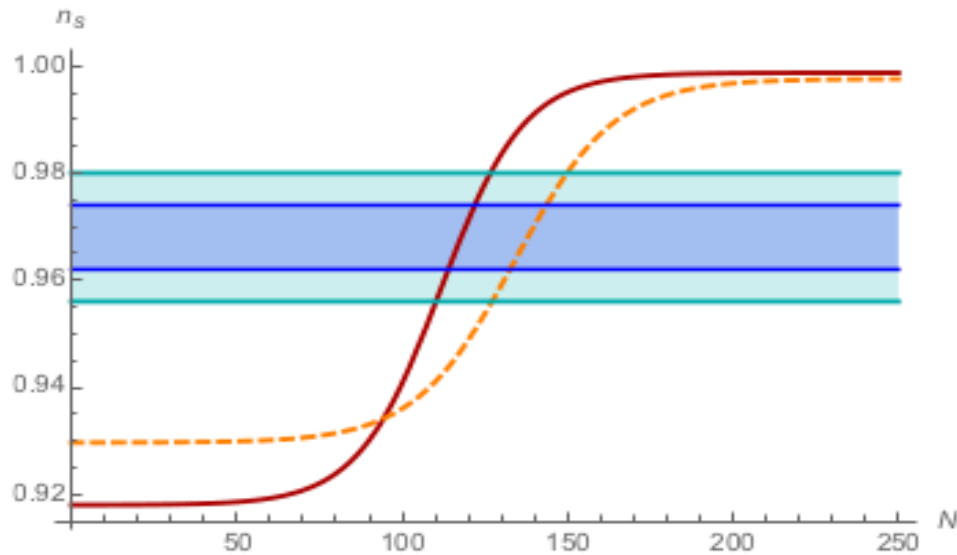
$$P_k \approx \frac{V_0^2}{M_{Pl}^2} \frac{\sigma^2 + H^2}{\phi^2}$$

$$\phi(N) \sim \phi_0 e^{m_0 |N|}$$

$$\sigma(N) \sim \sigma_0 e^{m_0 |N|}$$

$$\phi/\sigma = \text{const}$$

## Hilltop (Continued)



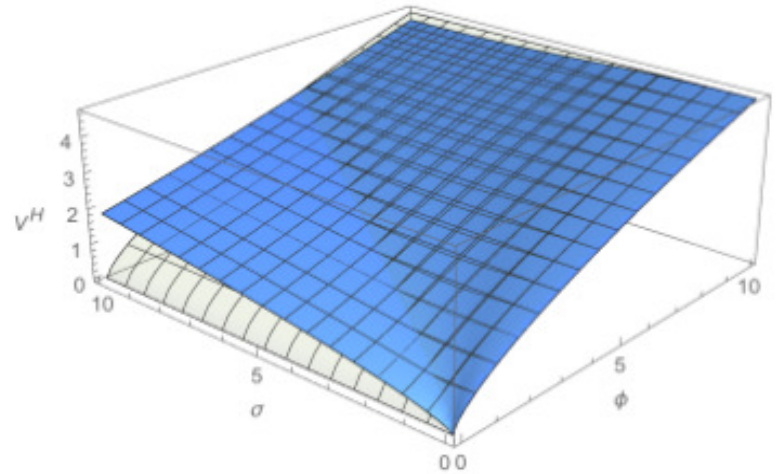
~30 e-folds

- Usual hilltop models are excluded - diarkyon can help!
- Tends towards scale invariance.
- Fairly slow climb

Quartic: 
$$V^H = V_0 - \lambda (\phi^4 + 6\sigma^2 \phi^2 + 3\sigma^4)$$

$$V = \mu^{10/3} \phi^{2/3}$$

$$V_{\text{flat}}^H(\phi, \sigma) = \mu^{10/3} \left[ \frac{\Gamma(5/6) \sigma^{2/3}}{2^{2/3} \pi^{1/2}} {}_1F_1 \left( -\frac{1}{3}; \frac{1}{2}; -\frac{\phi^2}{2\sigma^2} \right) + \frac{\Gamma(4/3) \phi}{2^{1/6} \pi^{1/2} \sigma^{1/3}} {}_1F_1 \left( \frac{1}{6}; \frac{3}{2}; -\frac{\phi^2}{2\sigma^2} \right) \right]$$



IF Starts from the eternal regime:

$$P_k \sim 1 \quad \phi_0 \sim \frac{M_{\text{Pl}}^{9/4}}{\mu^{5/4}} \quad \sigma_0 \sim H_0$$

Then From  $\sigma(N) = \sigma_0 \frac{V_+(N)}{V_+(0)}$

$$\sigma_{\text{CMB}} \sim \left( \frac{M_{\text{Pl}}}{\mu} \right)^{5/6} H_{\text{CMB}} \gg H_{\text{CMB}}$$

$$n_s = .989$$

$$r = 10^{-3}$$

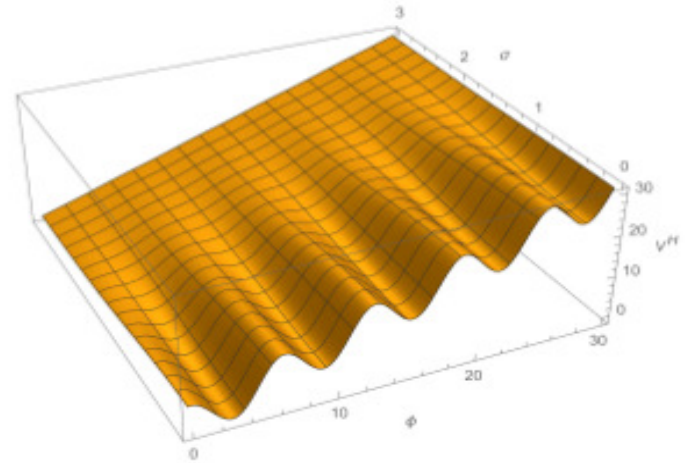
This effect  
HAS to be  
taken  
into account.



## Monodromy

$$V^H(\phi, \sigma) = \mu^3 \phi + \Lambda^4 e^{-\frac{g^2}{2f^2} \sigma} \cos \frac{\phi}{f}$$

Large  $\sigma_0 \Rightarrow$  decreased  $\Lambda_{\text{eff}}^4$



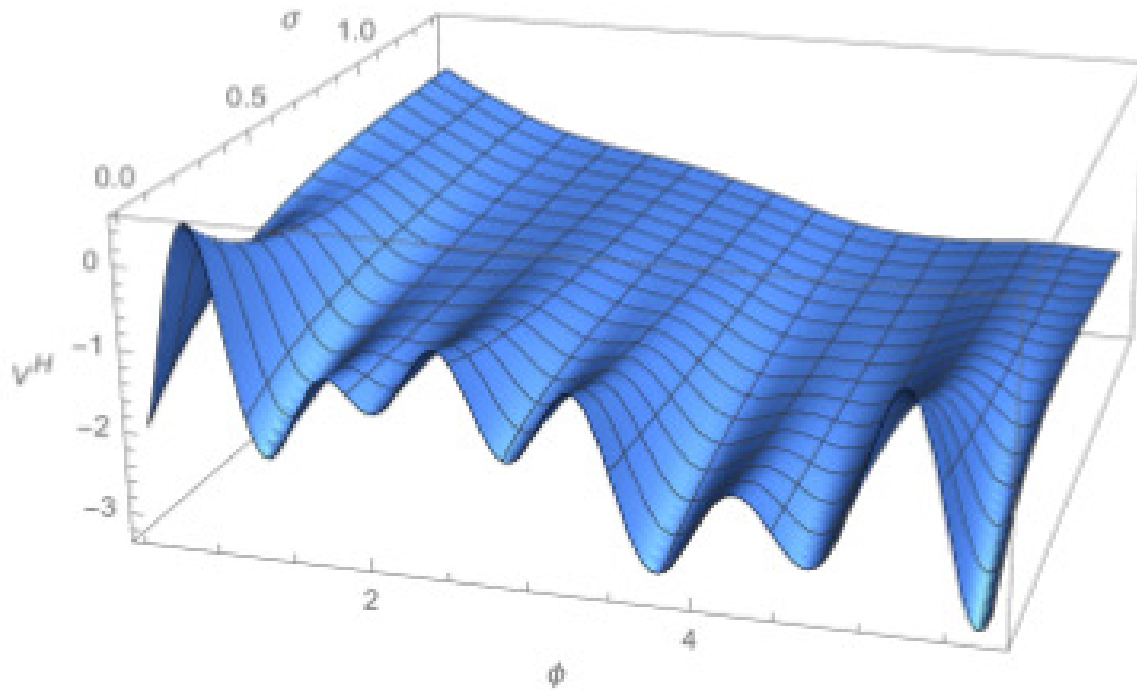
would not expect wiggles in CMB

Disclaimer:  $P_k \sim 10^{-9} \Rightarrow \sigma \sim 10^{-5} M_p$

Screening efficient for  $\sigma \gtrsim 4f$

Iron out wiggles only for  $f < 10^{-6} M_p$ .

## Generic Potential



can "iron out" false vacua and  
produce a slow roll potential

## Conclusions

↳ Spinodal effects have important consequences

Amplitude can depend on initial conditions

LARGE FIELD

decreases  
tensor to scalar ratio

Effect has to be  
taken into account

SMALL FIELD

Spectral tilt tends  
toward scale invariance

Can dampen features  
in the potential

## Future Work

↳ Need to check validity of Hartree approx.

↳ Incorporate into group invariant perturbation theory

↳ Check effects on 3 point function

↳ etc...