EW Phase Transition and Baryogenesis

Jing Shu

Institute of Theoretical Physics, Chinese Academy of Science, Beijing (KITPC)

SUSY 2015, Lake Tahoe, California

Outline

The Time to test EWBG (A general overview).
Electroweak Phase Transition in the post-LHC era.
Higgs related CP violation and baryon asymmetry generation in the post-LHC era.
Future prospects of testing EWBG.

Summary and outlook.

Concise Sketch on the Electroweak Baryogenesis after the LHC data

The origin of mass!





Higgs mechanism The origin of electroweak symmetry breaking What big questions can we learn from that?

Higgs ???

The origin of matter

How mass is generated in our universe?



After the electroweak phase transition, the broken phase, all the masses are turning on.

How "positive" matter is generated in our universe? Quite interesting if connected to the mass generation.

The EWBG

Electroweak baryogenesis: generate baryon asymmetry with particle mass generation at EW scale.

Sakharov's condition (EWBG):

Baryon number violation (Sphaleron transitions)

Strongly 1st order PT (SM: crossover) Out of thermal equilibrium

CP violation (SM CPV too small)

Strongly Ist order PT

When the universe is cooling down, if we have strongly 1 st order PT, then we have bubble expanding



Strongly first order phase transition

B number generation



require strongly first order phase transition

$$- m_{\chi}(v)e^{i\xi(v)}$$
$$\mathcal{L}_{X} \sim (\partial_{\mu}\xi)J_{X}^{\mu}$$
$$\partial_{t}\xi = \partial_{v}\xi(\Delta v)v_{w}/L_{w}$$
Behave like a chemical potential term
$$Q_{X} \sim g_{*}(\partial_{t}\xi)T^{2}/6 \sim (\partial_{t}\xi)T^{2}.$$

CPV phase jump generate a net chiral charge close to the bubble wall

Converted to B by sphalerons inside the bubble wall

Lessons from T=0 Higgs physic



EW Phase Transition: For any particle S would contribute to the thermal Higgs effective potential



Charged under SM group: (MSSM stops, 2HDM, etc: LHC Higgs global fits, direct searches on S, etc) A

Higgs mixture (NMSSM, etc, LHC non-Standard Higgs searches, EW precision, etc)

Hidden (EW & Higgs precision, SPPC direct search) C

LHC Higgs data: CPV source



A complex mass term which has vev dependence

suggests that particle χ would contribute to hgg and $h\gamma\gamma$. vertex with CPV

J. S,Y. Zhang, Phys. Rev. Lett. 111 (2013) 091801



if colored



if electric charged

 χ as top quark

Lessons from T=0 Higgs physic

 $m_{\chi}(v)e^{i heta(v)}$

A complex mass term with vev dependence

 Charged under SM group: (MSSM charginos, stops, 2HDM top quark, etc: LHC Higgs global fits, direct CPV searches.)

Singlet Higgs mixture (NMSSM, etc: EWBG driven by singlino. Relatively difficult to test)

Complex Yukawa (vector quarks)?

The EW Phase Transition & Higgs physics

EWPT Classification



Also may from fermion contributions

M. Carena, A. Megevand, M. Quiros, C. Wagner, NPB 716 (2005) 391-351
 H. Davoudiasl, I. Lewis, E. Ponton, PRD 87 (2013) 093001

A Master Formula



Increase the thermal cubic term E term (E is the coefficient of $\phi^3 T$)

• Decrease the T=0 energy difference ΔV (Increase the fine tuning of the Higgs potential)

Large E case (No Higgs mixing)

The 1st case (large E) is indeed the case A & C (The MSSM stops, strongly top Yukawa with N=12 to increase the loop effect)



$$T(\phi, T) \approx \frac{1}{4}\lambda\phi^4 + \frac{1}{2}\left[-\mu^2 + \epsilon_h T^2\right]\phi^2 - T\left[E_{\rm SM}\phi^3 + 2N(r_s)\frac{m_s^3(\phi, T)}{12\pi}\right]$$
$$\frac{1}{2\pi}$$

term $-Tm_s^3(\phi, T)$ has to decrease with phi to compete with positive terms such that there is a 1st PT

Therefore for the single scalar, $\alpha>0$.

Saturday, August 29, 2015

Higgs Global Fits

For the positive EWSB mass square $\alpha > 0$, if it is colored or electric charged.



$$\frac{\sigma(gg \to h)}{\sigma(gg \to h)_{\rm SM}} = \frac{\Gamma(h \to gg)}{\Gamma(h \to gg)_{\rm SM}} = \frac{\hat{c}_{g,\rm SM} + \delta c_g}{\hat{c}_{g,\rm SM}}$$

$$\frac{\Gamma(h \to \gamma \gamma)}{\Gamma(h \to \gamma \gamma)_{\rm SM}} = \frac{\hat{c}_{\gamma,\rm SM} + \delta c_{\gamma}}{\hat{c}_{\gamma,\rm SM}}$$

$$\delta c_g = rac{C(r_s)}{2} rac{lpha v^2}{m_s^2} A_s(au_s)$$

$$\delta c_\gamma = rac{N(r_s)Q_s^2}{24}rac{lpha v^2}{m_s^2}A_s(au_s)$$

Enhanced Higgs production Suppressed Higgs di-photon BR

W. Huang, J. S, Y. Zhang, JHEP 1303 (2013) 164 D. Chung, A. Long, L. T. Wang, PRD 87 (2013) 023509

The MSSM case

Stop (a>0) & Stop like state (a>0) sbottom(a<0)



For MSSM stops, it requires small stop mass ~ 100GeV and large Yukawa couplings to Higgs so it is ruled out by the experiments.

> D. Curtin, P. Jaiswall, P. Meade, JHEP 1208 (2012) 005
> T. Cohen, D. Morrissey, A. Pierce, PRD 86 (2012) 013009

Adding the light sbottom into the spectra can enhance the EWPT strength and cancel the effects from stop on Higgs global fits

W. Huang, J. S,Y. Zhang, JHEP 1303 (2013) 164

Decrease ΔV .

Higgs + singlet (NMSSM)

The 2nd case (decrease the ΔV) is the case B, which involves Higgs singlet mixing (Higgs/2HDM + singlet, the NMSSM, etc)

$$\frac{v_c}{T_c}\simeq \frac{v^4 E}{2\Delta V}$$

Big PT strength with small T=0 potential difference ΔV .

Other aspects NMSSM EWBG:

6 type-1 5 type-3 4 ž з 2 0 0.5 1.5 2 2.5 1 3 3.5 0 ΔV_{num}(10⁸GeV)

W-c. Huang, Z-f. Kang, J. S, P.-w Wu, Jm.Yang, Phys. Rev. D. (2015) 2 025006

J. Kozaczuk, S. Profumo, L.S.Haskins, C.Wainwright, JHEP, 1501 (2015) 144 C. Balazs, A. Mazumdar, E.Pukartas, G.White, JHEP, 1401 (2014) 073

Tuning of the potential

$$\frac{v_c}{T_c}\simeq \frac{v^4 E}{2\Delta V}$$

The natural size of ΔV is $\lambda v^4/2$ $\lambda \sim$ order 0.1 for 125GeV Higgs

The E is a one-loop effect, ~ 0.01 for EW couplings

Higgs +
Singlet
$$\Delta V = \frac{v^4}{2} \left(\tilde{\lambda} - \frac{2\tilde{a}^2 m_s^2}{(m_s^2 + \lambda^2 v^2)^2} \right)$$

M. Carena, N. Shah, C. Wagner, Phys. Rev. D. (2015) 85 036003 The tuning of the Higgs potential is roughly the order E/λ .

Weak coupling strength, more than 10% of tunning

NMSSM PT Patterns

Decoupled Stop, Phase Transition Triggered by the Higgs-singlet sector As our universe cools down HI scenario: 125 Type I PT: it first goes GeV Higgs is the to the S-breaking, lightest EW symmetric phase, then to the 🛑 H2 scenario: 125 EW breaking phase. GeV Higgs is the 2nd lightest

Type III PT: One step PT.

NMSSM

The question is now to understand the parameter dependence of ΔV .

PT parameter

For the 1st time, EWPT in NMSSM is understood semi-analytically!

 $R_{\kappa} \equiv \frac{4\kappa v_s}{A}$ • Type III PT (HI) SFEWPT requires $R_{\kappa} \subset (5, 30)$

• Type III PT (H2): SFEWPT suggests $R_{\kappa} \sim -1$ and $R_{\kappa} \subset (2, 10)$.

Type III PT (H2): SFEWPT suggests $R_{\kappa} \sim -4/3$.

Overall Spectra







Strong Ist PT suggest light Higgs spectra

No PQ limit since the stops are decoupled

Overall Spectra



Challenge in searches, may be ttbar h/A channel? Or other non-Standard Higgs search channels?

Hidden Models (C)

Real singlet with no vevs only couples with the Higgs Direct and indirect searches at the future colliders Future Z pole precision measurements (Two loop effects) . Henning, X-c. Lu, H. Murayama, 1404.1058 ● ILC + CEPC / FCC-ee indirect probe of Higgs-self couplings or A. Katz, M. Perelstein, hZZ, etc. JHEP 1407 (2014), 108 D. Curtin, P. Meade, C-T. Yu SPPC direct production! JHEP 1411 (2014) 127

Higgs self interactions



Higgs self-interaction is believed to be a very good probe of the strongly 1st order EWPT.

A. Nobel, M. Perelstein, Phys. Rev. D. 78 (2008) 063518

Large deviations in general are expected with some exceptional cases with small deviation

A. Katz, M. Perelstein, JHEP 1407 (2014), 108

Indirect search of Higgs self-coupling at the
 future lepton collider
 M. McCullough, Phys. Rev. D 90 (2014), 1, 015001

Di-Higgs production at the HL--LHC or SPPC / FCC-hh (Higgs-self couplings) S. Profumo, M. Ramsey-Musolf, C. Wainwright, P. Winslow Phys. Rev. D 91 (2015), 3, 035018

D. Curtin, P. Meade, C-T. Yu JHEP 1411 (2014) 127

Other aspects!

 $\bigcirc \bigcirc \bigcirc$

The two curves here would results very different EWPT.

Need knowledge of global behavior of Higgs potential in principle



Colliders can only probe the local behavior. Maybe future gravitational detection ?

CPViolation & EDM & Higgs Physics

General connection





Behave like a chemical potential term

$$Q_X \sim g_*(\partial_t \xi) T^2/6 \sim (\partial_t \xi) T^2.$$

Converted to B by sphalerons inside the bubble wall

General connection

Consider a fermion mass X with complex mass term (CPV): $m(v)\overline{X}(1+i\xi(v)\gamma_5)X$.

Expanding around v, CP odd term:

$$im(v)\left\{\xi(v) + \frac{\partial\xi(v)}{\partial \log v}\frac{h}{v}\right\}\bar{X}\gamma_5X$$

An axial rotation of X can remove the 2nd term, results extra terms where F is the gauge field where X is charged.

$$egin{aligned} & ilde{c}_X\left(rac{h}{v}
ight){}_{F ilde{F}} \ & ilde{c}_X=v[\partial\xi(v)/\partial v]. \end{aligned}$$

$$\sim \left(\xi(v) + \frac{\partial \xi(v)}{\partial \log v} \frac{h}{v}\right) F \tilde{F}$$

The CP violating sources from X in EWBG is proportional to the size of the effective operators where X is integrated out.

Affect the Higgs global fits



if colored $\begin{array}{c} H \\ \hline t \\ \hline \\ \gamma \end{array}$

if electric charged

 χ as top quark

 $m_{\chi}(v)e^{i heta(v)}$

A complex mass term which has vev dependence

suggests that particle χ would contribute to hgg and $h\gamma\gamma$. vertex with CPV

There might be more universal results based on Higgs low energy theorem.

J. S, Y. Zhang, Phys. Rev. Lett. 111 (2013) 091801

2HDM example

In order to make a connection with baryogenesis, I must make a model.

$$V = \frac{\lambda_1}{2} (\phi_1^{\dagger} \phi_1)^2 + \frac{\lambda_2}{2} (\phi_2^{\dagger} \phi_2)^2 + \lambda_3 (\phi_1^{\dagger} \phi_1) (\phi_2^{\dagger} \phi_2) + \lambda_4 (\phi_1^{\dagger} \phi_2) (\phi_2^{\dagger} \phi_1) + \frac{1}{2} \left[\lambda_5 (\phi_1^{\dagger} \phi_2)^2 + \text{h.c.} \right]$$

$$\frac{1}{2} \left\{ m_{11}^2(\phi_1^{\dagger}\phi_1) + \left[m_{12}^2(\phi_1^{\dagger}\phi_2) + \text{h.c.} \right] + m_{22}^2(\phi_2^{\dagger}\phi_2) \right\}$$

There are two independent phases from m_{12} and λ_5 .

 $\mathcal{L}_Y = \bar{Q}_L Y_D \phi_1 D_R + \bar{Q}_L Y_U(i\tau_2) \phi_2^* U_R + \bar{L}_L Y_E \phi_1 E_R$

	$\gamma\gamma$	WW^*	ZZ^*
ATLAS	1.17 ± 0.27 20	$0.99^{+0.31}_{-0.28}$ [22]	$1.44^{+0.40}_{-0.33}$ 24
CMS	$1.14^{+0.26}_{-0.23}$ [21]	$0.72^{+0.20}_{-0.18}$ [23]	$0.93^{+0.29}_{-0.25}$ [25]
	bb	au au	
ATLAS	0.52 ± 0.40 26	$1.4^{+0.5}_{-0.4}$ [28]	C
CMS	1.15 ± 0.62 27	0.78 ± 0.27 29	$\mathcal{L}_{h_1}V$

$$\begin{split} c_t &= \frac{\cos\alpha}{\sin\beta}\cos\alpha_b \ , \quad c_b = -\frac{\sin\alpha}{\cos\beta}\cos\alpha_b \\ \tilde{c}_t &= -\cot\beta\sin\alpha_b \ , \quad \tilde{c}_b = -\tan\beta\sin\alpha_b \end{split}$$

 $\mathcal{L}_{h_1VV} = \cos \alpha_b \sin(\beta - \alpha) \mathcal{L}_{hVV}^{SM} \equiv a \mathcal{L}_{hVV}^{SM}$

Bounds from EDM

 When there is a CP odd operator contributes to hgg or $h\gamma\gamma$.

The same operators would contribute to the EDM or CEDM

D. McKeen, M. Pospelov, A. Ritz, PRD, 86, 113004 (2012) $\tilde{c}_{\gamma} \sim \mathcal{O}(10^{-1}) - \mathcal{O}(10^{-2})$

Bounds from neutron EDM and chromo-EDM (CEDM) are much weaker due to small u, d quark charge and Wilson coefficient in RG running.



Second region

Blue points: best fits

sweet spot around $\tan\beta \sim 1$

J. S,Y. Zhang, Phys. Rev. Lett. 111 (2013) 091801

New ACME results

Much Tighter constraints than before:

 $|d_e| < 8.7 \times 10^{-29} ecm$ at 90% C. L.

More than one order improvements

Naively constraints $\tilde{c}_{\gamma} \sim \mathcal{O}(10^{-2}) - \mathcal{O}(10^{-3})$

J. Brod; U. Haisch and J. Zupan, JHEP, 1311, 180 (2013)

But is that really the general case? No need for CPV direct search?

Where are the room for direct CPV searches?



.0 SM 🔸 5

ATLAS only, tan $\beta = 2$



CMS only, $\tan \beta = 0.5$



CMS only, tan $\beta = 0.8$



CMS only, $\tan \beta = 2$



Combined, $\tan \beta = 0.5$



Combined, $\tan \beta = 0.8$



Combined, $\tan \beta = 2$



Ilts

Much Tightly constrained than before:

J. S,Y. Zhang, Phys. Rev. Lett. 111 (2013) 091801

2HDM case

EFT only for illustration



The two pieces can naturally cancel each other

The Higgs is a CP mixture!

2HDM case

Final Results



Future Neutron EDM bound will

Heavy Higgs Searches



C-Y. Chen, S. Dawson, Y. Zhang, JHEP 1508 (2015) 058

There are also heavy Higgs searches bound (strong in the cancellation region)

MSSM case



$$\begin{bmatrix} \frac{de}{e} \end{bmatrix} \approx C \tilde{c}_e^A \sum_{j=1,2} \left(c_{\gamma}^{\tilde{\chi}_j^{\pm}} \ln \frac{1}{z_{\tilde{\chi}_j^{\pm}}^A} + c_{\gamma}^{\tilde{\tau}_j^{\pm}} \ln \frac{1}{z_{\tilde{\tau}_j^{\pm}}^A} \right) \\ - C c_e^H \sum_{j=1,2} \tilde{c}_{\gamma}^{\tilde{\chi}_j^{\pm}} \ln \frac{1}{z_{\tilde{\chi}_j^{\pm}}^H}$$

MSSM with chargino & staus

Negligible CPV in the Higgs sector

EFT only for illustration



Heavy Higgs coupling enhanced by tan beta Non-standard Higgs mediate the cancellation

Allowed CPV for MSSM



Correction in CPSuperH

A sign mistake in the anomalous D of the dipole operator (smaller EDMs at low energy).

No operator mixing effects are considered

Detailed RG running in the Mercury & other EDMs

Update the matrix elements (factor of 10 difference)

• W boson loop included in the Barr-Zee diagram.

The bounds with EDM with colored objects are much weaker

Correction in CPSuperH

$$\gamma_s = \begin{bmatrix} +8C_F & 0 & 0 \\ +8C_F & +16C_F - 4N & 0 \\ 0 & +2N & N + 2n_f + \beta_0 \end{bmatrix}, \quad (36)$$

$$\gamma_f = [-12C_F + 6],$$
 (37)

$$\gamma_f' = \begin{bmatrix} -12C_F & 0\\ 0 & -12C_F \end{bmatrix}, \quad (38)$$

and

$$\gamma_{sf} = \begin{bmatrix} +4 & +4 & 0\\ 0 & 0 & 0 \end{bmatrix}, \quad (39)$$

where N = 3, $C_F = (N^2 - 1)/(2N) = 4/3$, $\beta_0 = (11N - 2n_f)/3$ and n_f is the flavor number.

Now, we explore details of the RG running.

Firstly, we need to use the $n_f = 5$ version of the above RGE for running from Λ (we use M_H in our analysis) to m_b . In which, CP-odd four-fermion operators (33) play a significant role. For our case, we one consider the operators containing the bottom quark for tan β enhancement effects. In addition to coefficients $C_{b(u,d)}$, $C_{(u,d)b}$ that contribute to the light quark CEDM through RGE operator mixing, we also considered the coefficient C_{bb} which mixes with and contributes to the b-quark CEDM. Keeping only the leading logarithmic terms that make additional contributions to the CEDMs of bottom and light quarks at the matching scale $\mu = m_b$, we have bellow m_c scale we use 3 flavors version of RGE. After above processes, we have the neutron EDM

$$d_n = \left(e\zeta_n^u \delta_u + e\zeta_n^d \delta_d\right) + \left(e\tilde{\zeta}_n^u \tilde{\delta}_u + e\tilde{\zeta}_n^d \tilde{\delta}_d\right) + \beta_n^G C_{\tilde{G}}$$
,(44)

with update hadronic matrix elements $\zeta_n^u = 0.82 \times 10^{-8}$, $\zeta_n^d = -3.3 \times 10^{-8}$, $\tilde{\zeta}_n^u = 0.82 \times 10^{-8}$, $\zeta_n^d = 1.63 \times 10^{-8}$ and $\beta_n^G = 2 \times 10^{-20} e \,\mathrm{cm}$ [45].

(2) Mercury EDM

Though the contributions from d_e^E and from the CPodd electron-nucleon interactions

$$\mathcal{L} = C_S \bar{e} i \gamma_5 e \bar{N} N$$

+ $C_P \bar{e} e \bar{N} i \gamma_5 N + C'_P \bar{e} e \bar{N} i \gamma_5 \tau_3 N$, (45)

are also incorporated in the CPsuperH, the mercury EDM is mainly contributed by the nuclear Schiff moment (S). The Schiff moment is generated by long-range, pionexchange mediated P- and T-violating nucleon-nucleon interactions,

$$\mathcal{L}_{\pi NN}^{\text{TVPV}} = \bar{N} \left[\bar{g}_{\pi}^{(0)} \vec{\tau} \cdot \vec{\pi} + \bar{g}_{\pi}^{(1)} \pi^0 + \bar{g}_{\pi}^{(2)} (2\tau_3 \pi^0 - \vec{\tau} \cdot \vec{\pi}) \right] N(46)$$

In a general context, the isoscalar and isovector couplings $\bar{g}_{\pi}^{(0)}$, $\bar{g}_{\pi}^{(1)}$ are dominant over the isotensor coupling $\bar{g}_{\pi}^{(2)}$ [45], so the mercury EDM is approximately given by [45],

$$d_{\text{Hg}} = \kappa_S S \approx \kappa_S \frac{2m_N g_A}{F_{\pi}} \left(a_0 \bar{g}_{\pi}^{(0)} + a_1 \bar{g}_{\pi}^{(1)} \right) ,$$
 (47)

L-g, Bian, T. Liu, J. <mark>S</mark>, Phys. Rev. Lett. 115 (2015) 021801

Codes will be updated in the web soon

Correction in CPSuperH



L-g, Bian, T. Liu, J. S, Phys. Rev. Lett. 115 (2015) 021801

Saturday, August 29, 2015

Something is up in the air?

Category	EDM Limit $(e \cdot cm)$	Experiment	Standard Model value $(e \cdot cm)$
Electron	$8.7 imes10^{-29}$	ThO molecules in a beam 12	10^{-38}
Neutron	$2.9 imes10^{-26}$	Ultracold neutrons in a bottle [11]	10^{-31}
Nucleus	$3.1 imes10^{-29}$	¹⁹⁹ Hg atoms in a vapor cell [13]	10^{-33}

K Kumar, Z-T. Lu, M. R-Musolf 1312.5416

TRIUMF:

The long-term goal, to be reached in 2018 and beyond, is $d_n < 1 \times 10^{-28} e \cdot cm$.

Proton ring: $\sim 1 \times 10^{-29} e \cdot cm$

a projected sensitivity of $10^{-28} - 10^{-29} e \cdot cm$ for 225 Ra, $\sim 10^{-31} - 10^{-32} e \cdot cm$ for 199 Hg.

3 orders of magnitude, close the cancellation region

Think more!

If one of those experiments are before the end of LHC (including HL-LHC)

Do we need direct LHC CPV searches?

or even CPV searches in the future lepton colliders such as CEPC in China?

One possible solution is that we search for CPV in b systems, (LHCb!)

Electroweak Beautygenesis

The key idea is to consider extra inert scalars which dominantly couples to b and s with CPV. H_{int} New source for An extra scalar CPV in Bs mixing triggers 1st h order EWPT. $\rightarrow S$ $\langle H \rangle$ HCPV in b,s in H_{int} the Higgs background

h,

T. Liu, M. Ramsey-Musolf J. S, Phys. Rev. Lett. 108 (2012) 221301
 H. Guo, S. Hong, T. Liu, M. Ramsey-Musolf J. S, in progress

Saturday, August 29, 2015

 $\overline{\mathbf{S}}$

EDMs





No direct Barr-Zee diagram contributions

Weinberg operator contributions are small due to small bottom quark mass

Comments on CPV channels

CPV in h to massive gauge boson coupling: CP odd is dim5, always small (hard to measure).

Measure CPV in gamma gamma or Z gamma requires information in photon polarization.
G.C.E.E.Bishara, X.Grossman, R. Harnik, D. Bobinson, J.S. J. Zupan, J.H.

G.C. F. F. Bishara, Y. Grossman, R. Harnik, D. Robinson, J.S. J. Zupan. JHEP 1404 (2014) 084; Y. Chen, A. Falkowski, I. Low, R. Vega-Morales 1405.6723

Top CPV promising: in ggjj--> h + 2j or ttbar Higgs
 M. Dolan, P. Harris, M. Jankowiak, M. Spannowsky, PRD, 90 (2014), 073008
 Other CPV fermion couplings

R. Harnik, A. Martin, T. Okui, R. Primulando, F. Yu, PRD, 88 (2013) 7, 076009

Various New Ideas

Electroweak Cogenesis (asymmetric DM and generation at the weak scale). C, Cheung, Y. Zhang, JHEP 1309 (2013) 002

- Electroweak baryogenesis from exotic electroweak
 symmetry breaking.
 N. Blinov, J. Kozaczuk, D. Morrissey, C. Wagner, Phys. Rev. D (2015) 035012
- Electroweak baryogenesis from a early time phase J.S, T. Tait, C. Wagner, Phys. Rev. D (2007) 063510

Electroweak baryogenesis from Randall Sundrum
 P. Creminelli, A. Nicolis, R. Rattazzi, JHEP 0203 (2002) 051
 G. Nardino, M. Quiros, A. Wulzer, JHEP 0709 (2007) 077

Some models missing......

Summary & Outlook

A natural connection between EWPT and Higgs physics Strongly 1 st order tells you: strong coupling or shallow potential at T=0.

Future lepton and hadron colliders greatly help this aspect.

- A natural connection between BAU in EWBG and CPV in the Higgs sector (EDMs are more sensitive than colliders).
- After the ACME results, large CPV effects are possible due to cancellation, future EDMs will soon recover those region.

Direct CPV at the LHC or future colliders are worth to look even after future EDMs. Great for both LHC & LHCb!

Prospects: BG versus DM

EWBG & Higgs physics: Collider Searches

> Low energy CPV experiments

DM & WIMPs, Direct Collider Searches

DM Direct detection

Gravitational DM Indirect Waves Searches. Many inputs from the astrophysics and nuclear physics has to be well understood

Fun to explore

Prospects: BG versus DM

t baryogenesis or baryon asymmetry	t dark matter			
find j "Phys.Rev.Lett.,105*" :: more		find j "Phys.Rev.Lett.,1	105** :: more	
Sort by:	Display results:	Sort by:		Display results:
latest first 🛟 desc. 🛟 - or rank by - 🛟	25 results 🛟 sing	latest first 🛟	desc. 🛟 🛛 - or rank by - 🛟	25 results 🛟 single list
HEP 1,208 records found 1	- 25 🕨 🕨 jump to	HEP 1	2,525 records found	1 - 25 🕨 🕨 jump to recor
1. Two-Step Electroweak Baryoge	enesis	1. Limits on o	uark nugget dark m	natter from cosmic ray
Satoru Inoue, Grigory Ovanesyan, Micha	ael J. Ramsey-Muso	Kyle Lawson	(British Columbia U.). 201	5. 5 pp.
ACFI-T15-12	_	Published in E	EPJ Web Conf. 99 (2015)	12005
e-Print: arXiv:1508.05404 [hep-ph] PD	F	DOI: <u>10.1051</u> /	epiconf/20159912005	
References BibTeX LaTeX(US)) <u>LaTeX(EU)</u> <u>Harv</u>	Conference:	C14-08-18.1 Proceedings	
ADS Abstract Service		Refere	nces BibTeX LaTeX(US	<u> 5) LaTeX(EU) Harvmac En</u>
Detailed record		Link to	Fulltext	
		Detailed recor	<u>rd</u>	
2. Baryogenesis via Mesino Oscil	lations			
Akshay Ghalsasi, David McKeen, Ann E	. Nelson. Aug 21, 2	2. Implication	is of the first AMS-0	2 antiproton data for da
e-Print: arXiv:1508.05392 [hep-ph] PD	F	Hong-Bo Jin,	Yue-Liang Wu, Yu-Feng Z	hou. Aug 27, 2015.
References BibTeX LaTeX(US)) <u>LaTeX(EU)</u> <u>Harv</u>	e-Print: arXiv	<u>:1508.06844</u> [hep-ph] <u>Pl</u>	DF
ADS Abstract Service		Refere	<u>nces BibTeX LaTeX(US</u>	<u> 5) LaTeX(EU) Harvmac En</u>
Detailed record		ADS A	bstract Service	
		Detailed recor	<u>rd</u>	
3. Baryogenesis in a CP invariant	theory			
		3. Dark Matte	er and Global Symme	etries