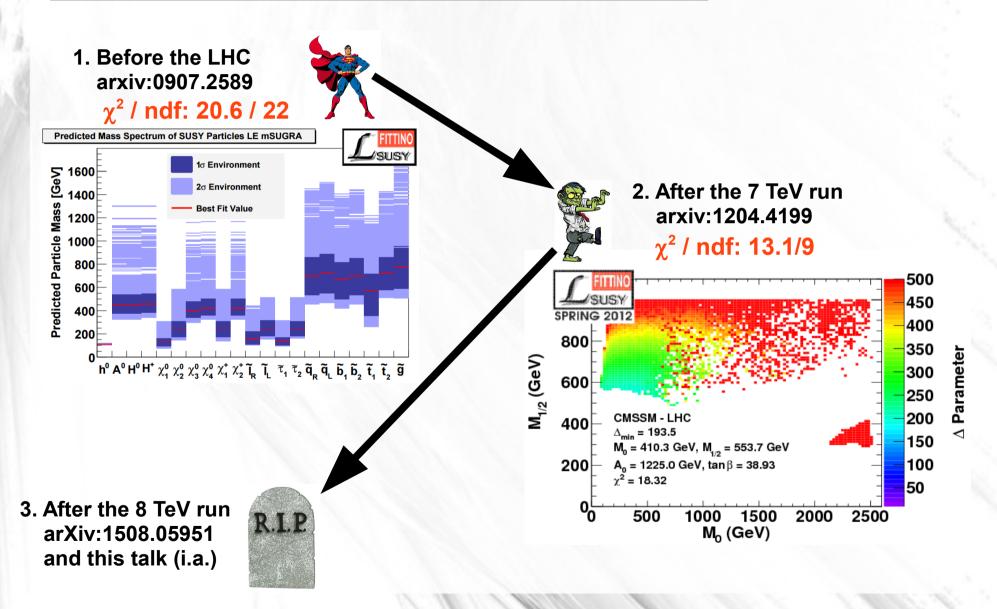
Killing the CMSSM softly



SUSY 2015, 27.08.2015

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Introduction - The Evolution of the CMSSM



Outline

- * The Fittino Framework for Global Fits of SUSY models
 - → Observables & Observable Sets
 - → Scanning and Constraining the Parameter Space
- * The CMSSM after the LHC 8 TeV Run
 - → Best Fit Points & Preferred Parameter Space
 - → The lightest Higgs in the CMSSM
 - → The p-Value of the CMSSM

The Fittino Framework & Outline

- * select sensitive observables
 - ★ low energy observables
 - ★ Higgs boson properties
 - * collider searches for sparticle production
 - * direct/indirect dark matter searches

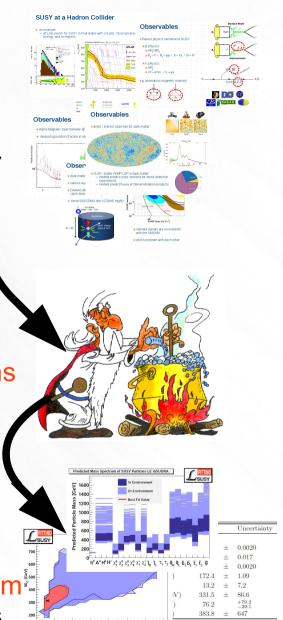
* scan the parameter space

* public codes for calculation of model predictions

 $\star \chi^2$ as a measure for level of agreement

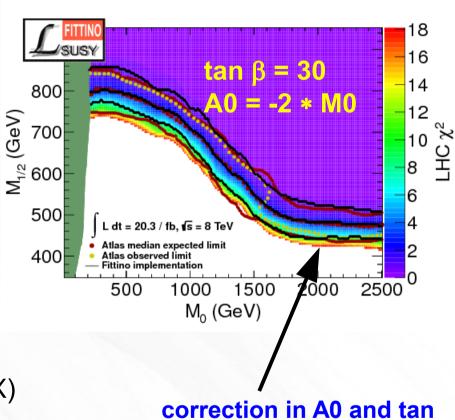
* Markov Chain Monte Carlo for smart sampling

- * statistical analysis
 - ★ frequentist interpretation
 - * preferred parameter regions and mass spectrum
 - ★ calculation of p-value with pseudo experiments



Observables

$a_{\mu} - a_{\mu}^{\mathrm{SM}}$	$(28.7 \pm 8.0) \times 10^{-10}$
$\sin^2 \theta_{ m eff}$	0.23113 ± 0.00021
m_t	$(173.34 \pm 0.27 \pm 0.71) \text{GeV}$
m_W	$(80.385 \pm 0.015) \text{GeV}$
$\Delta m_{\scriptscriptstyle S}$	$(17.719 \pm 0.036 \pm 0.023) \mathrm{ps}^{-1}$
$\mathscr{B}(B_s \to \mu \mu)$	$(2.90 \pm 0.70) \times 10^{-9}$
$\mathscr{B}(b \to s \gamma)$	$(3.43 \pm 0.21 \pm 0.07) \times 10^{-4}$
$\mathscr{B}(B o au u)$	$(1.05 \pm 0.25) \times 10^{-4}$
Ωh^2	0.1187 ± 0.0017



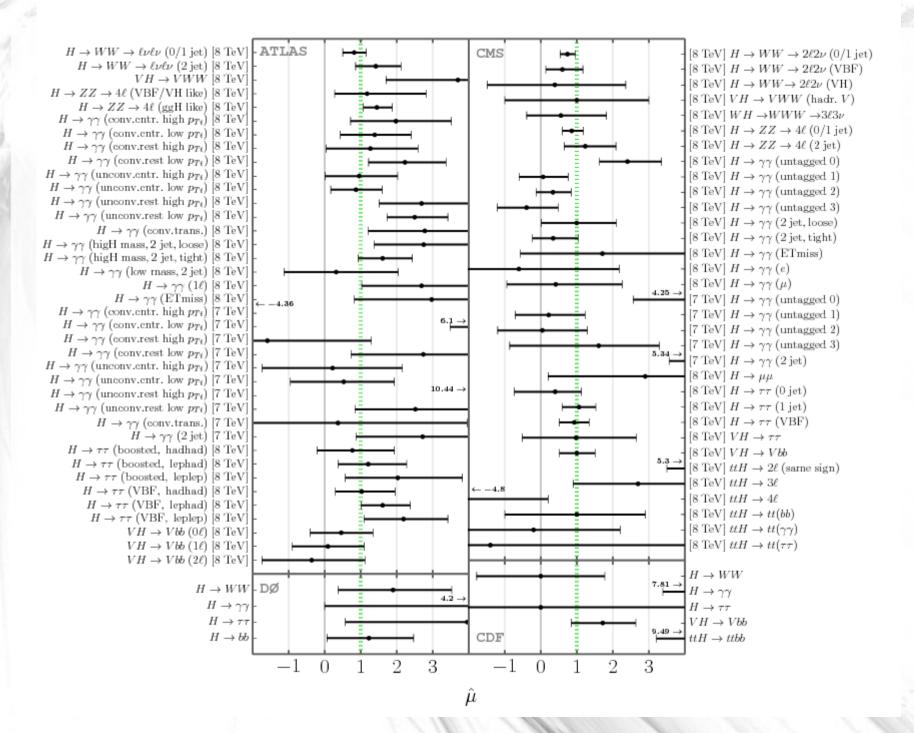
 β via scaling factor

limits on

- → direct detection cross-section (LUX)
- → chargino mass from LEP
- → SUSY production at the LHC

a lot of Higgs measurements

→ implemented via HiggsSignals and HiggsBounds



Higgs Observables Set

- ★ CMSSM can't distinguish between all measurements
- ★ use 3 additional combinations

Experiment, Channel	observed μ	observed m_h
ATLAS, $h \to WW \to \ell \nu \ell \nu$ [80]	$0.99^{+0.31}_{-0.28}$	-
ATLAS, $h \rightarrow ZZ \rightarrow 4\ell$ [80]	$1.43^{+0.40}_{-0.35}$	$(124.3\pm1.1)\text{GeV}$
ATLAS, $h \rightarrow \gamma \gamma$ [80]	$1.55^{+0.33}_{-0.28}$	$(126.8 \pm 0.9) \text{GeV}$
ATLAS, $h ightarrow au au$ [81]	$1.44^{+0.51}_{-0.43}$	-
ATLAS, $Vh \rightarrow V(\overline{bb})$ [82]	$0.17^{+0.67}_{-0.63}$	-
CMS, $h \to WW \to \ell \nu \ell \nu$ [83]	$0.72^{+0.20}_{-0.18}$	-
CMS, $h \rightarrow ZZ \rightarrow 4\ell$ [84]	$0.93^{+0.29}_{-0.25}$	$(125.6 \pm 0.6) \text{GeV}$
CMS, $h \rightarrow \gamma \gamma$ [85]	$0.77^{+0.30}_{-0.27}$	$(125.4 \pm 1.1) GeV$
CMS, $h \rightarrow \tau \tau$ [86]	$0.78^{+0.27}_{-0.27}$	-
CMS, $Vh \rightarrow V(bb)$ [86]	$1.00^{+0.50}_{-0.50}$	-

Medium Obs Set

 \rightarrow Baseline

observed μ	observed m_h
$1.33^{+0.21}_{-0.18}$	$(125.5 \pm 0.8) \mathrm{GeV}$
$1.44^{+0.51}_{-0.43}$	-
$0.17^{+0.67}_{-0.63}$	-
$0.80^{+0.16}_{-0.15}$	$(125.7 \pm 0.6) \mathrm{GeV}$
$0.78^{+0.27}_{-0.27}$	-
$1.00^{+0.50}_{-0.50}$	-
	$1.33^{+0.21}_{-0.18}$ $1.44^{+0.51}_{-0.43}$ $0.17^{+0.67}_{-0.63}$ $0.80^{+0.16}_{-0.15}$ $0.78^{+0.27}_{-0.27}$

Small Obs Set

Experiment, Channel	observed μ	observed m_h
ATLAS+CMS, $h \rightarrow WW, ZZ$ ATLAS+CMS, $h \rightarrow \gamma\gamma$	$0.94^{+0.17}_{-0.16}$ $1.16^{+0.22}_{-0.20}$	$(125.73 \pm 0.45) \mathrm{GeV}$
ATLAS+CMS, $h \rightarrow au au$	$1.11^{+0.24}_{-0.23}$	-
ATLAS+CMS, Vh , $tth \rightarrow bb$	$0.69^{+0.37}_{-0.37}$	-

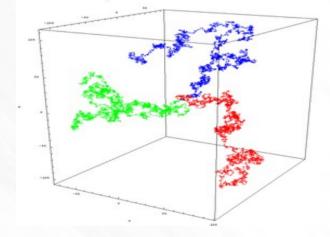
Combined Obs Set

Calculating Model Predictions

- ★ Fittino uses
 - → **SPheno** for the mass Spectrum
 - → SuperIso for the B-meson branching fractions
 - \rightarrow FeynHiggs for Higgs properties, m_w, $\sin\theta_{eff}$, $(g-2)_{\mu}$
 - \rightarrow micrOMEGAs for Ωh^2
 - → DarkSUSY via AstroFit for direct detection cross section
 - → Herwig++/Delphes/Prospino for the emulation of the ATLAS 0-Lepton search

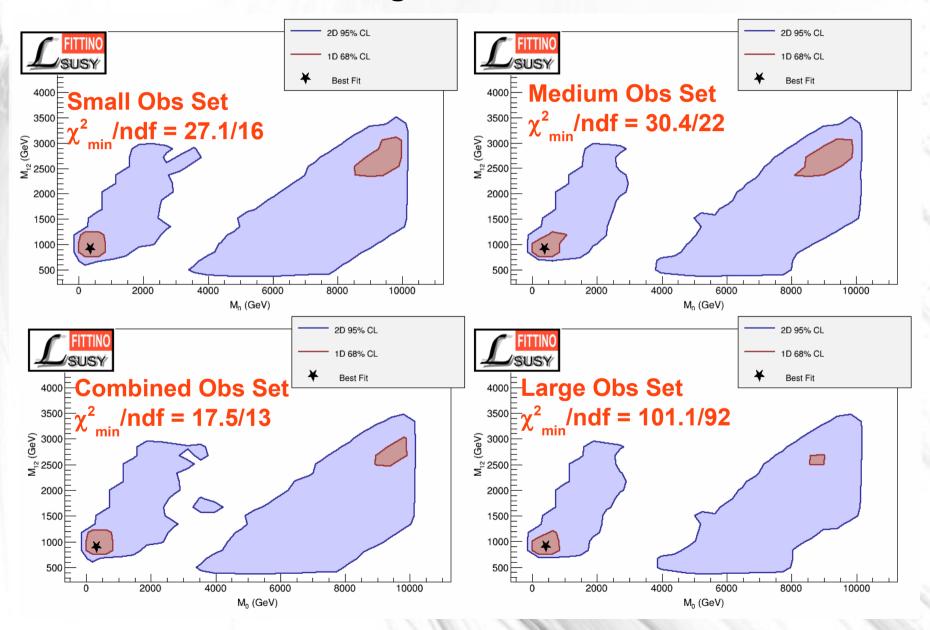
Sampling the Parameter Space

- ★ 3 goals
 - → accurate determination of best fit point
 - → extensive coverage of full parameter space
 - → accurate evaluation of p-value
- * adaptive Markov Chain Monte Carlo
 - → proposal densities adjusted regularly
 - → 20 independent chains
 - → 850 million valid points
 - \rightarrow 100 million points with χ^2 < 100

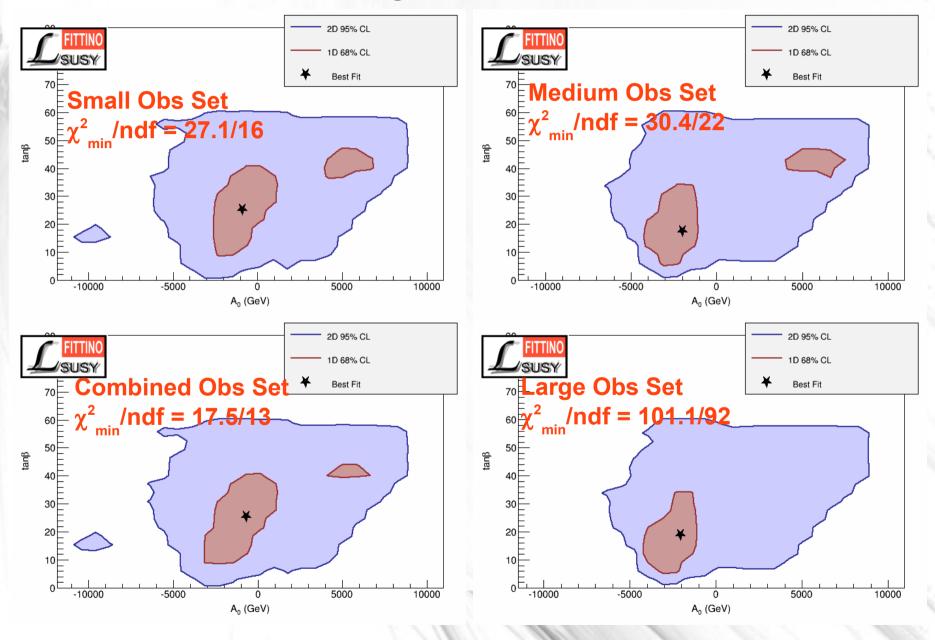


- ★ determination of p-Value
 - → full fit too demanding in terms of CPU time
 - → use original MCMC to find best fit points
 - → conservative estimate of model p-value

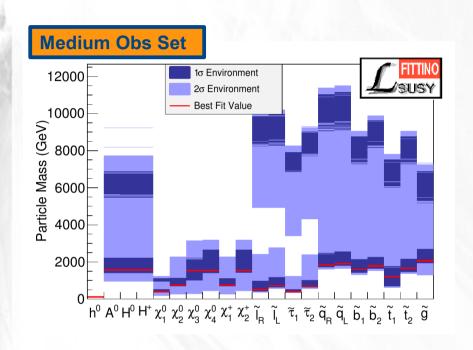
Results I: Parameter Regions & Best Fit Points

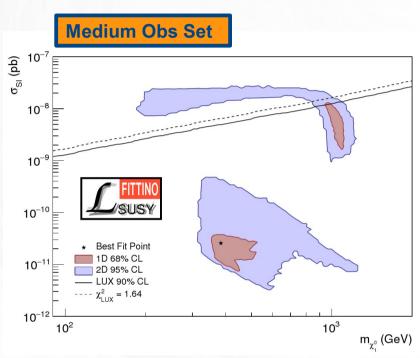


Results I: Parameter Regions & Best Fit Points



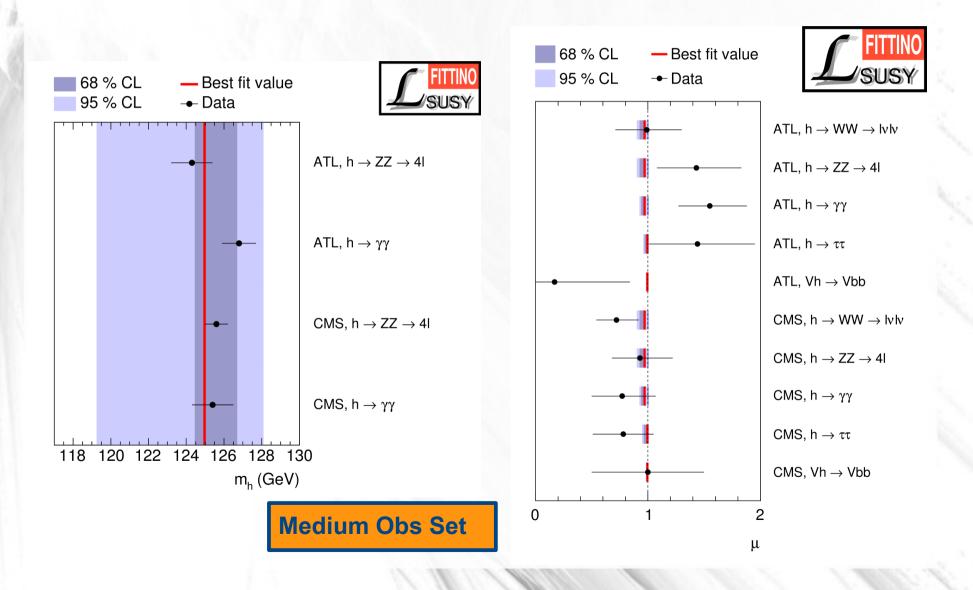
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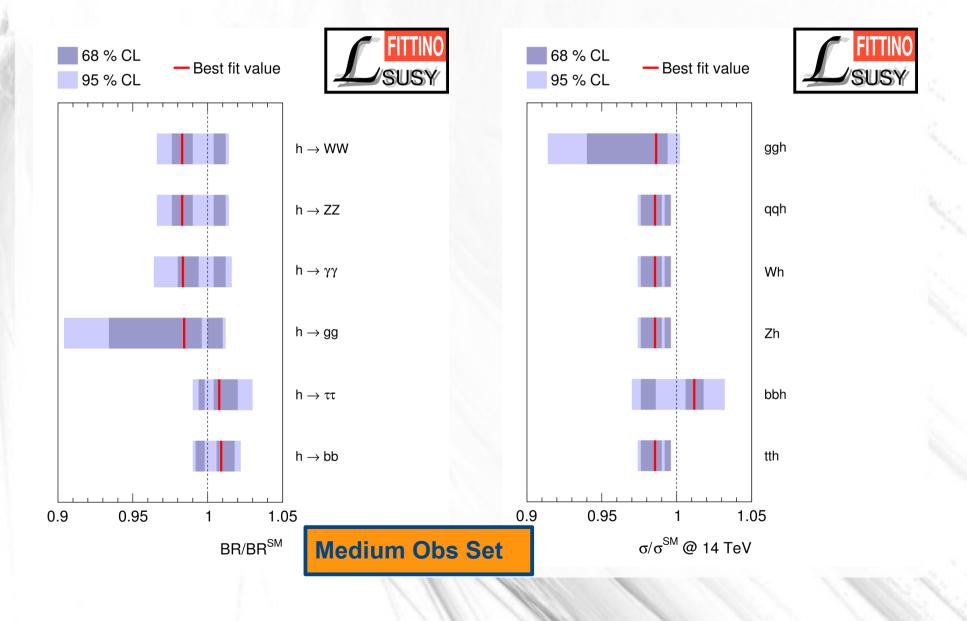


Observable Set	M_0	$M_{1/2}$	A_0	$\tan \beta$
Combined	$327.4~\mathrm{GeV}$	$900.5~\mathrm{GeV}$	-679.6 GeV	25.6
Small	$361.5 \mathrm{GeV}$	$926.3~{ m GeV}$	$-907.9 \; \mathrm{GeV}$	25.3
Medium	$387.4~\mathrm{GeV}$	$918.2~{ m GeV}$	$-2002.8 \; \mathrm{GeV}$	17.7
Large	$418.6~\mathrm{GeV}$	$910.6~{ m GeV}$	$-2041.6 \; { m GeV}$	19.2

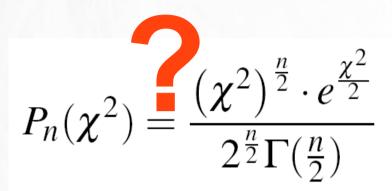
Results II: Lightest Higgs in the CMSSM



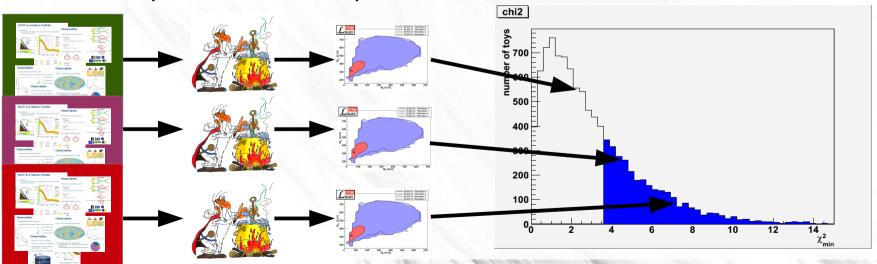
Results II: Lightest Higgs in the CMSSM



- ★ non-gaussian observable set
 - → 1-sided and hard limits
 - → non-gaussian uncertainties
 - → relative uncertainties
 - → highly non-linear model



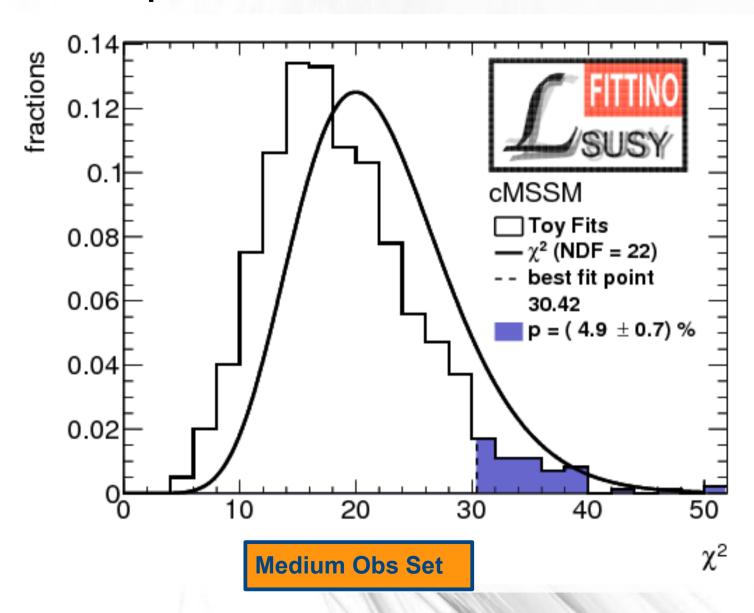
- ★ gaussian χ²-distribution not accurate
 - \rightarrow get true χ^2 -distribution from pseudo measurements
 - → ~1000 pseudo datasets per obs set

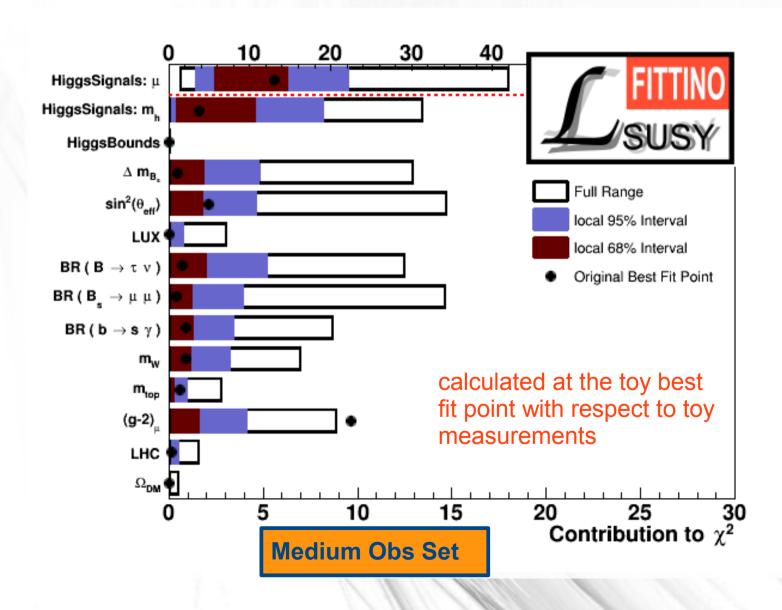


- * Large Obs Set: 84 different measurements in Higgs sector
 - → CMSSM makes the same prediction for several subsets
- * in terms of the p-value, the model can be
 - → punished for bad agreement within the data
 - → rewarded for good agreement within the data
- * p-value should reflect the quality of the model
 - → combine measurements with same prediction
 - → use combination in global fit

medium obs set comes closest to what we need

Experiment, Channel	observed μ	observed m_h
ATLAS, $h \to WW \to \ell \nu \ell \nu$ [80]	$0.99^{+0.31}_{-0.28}$	-
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Observable Set	χ^2/ndf	naive <i>p</i> -value (%)	toy <i>p</i> -value (%)
Small	27.1/16	4.0	1.9 ± 0.4
Medium	30.4/22	10.8	4.9 ± 0.7
Combined	17.5/13	17.7	8.3 ± 0.8
Medium (Focus Point)	30.8/22	10.0	7.8 ± 0.8
Medium without (g-2)	18.1/21	64.1	51 ± 3

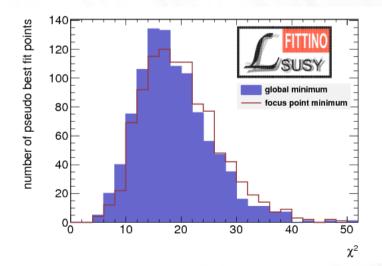
- \rightarrow naive p-value: p-value according to gaussian χ^2 -distribution
- → toy p-value: p-value extracted from pseudo experiments
- → stat.uncertainty: estimated uncertainty on p-value

$$\Delta p = \sqrt{\frac{p \cdot (1-p)}{n_{\text{Toy}}}}$$

Results IV: just one more thing . . .

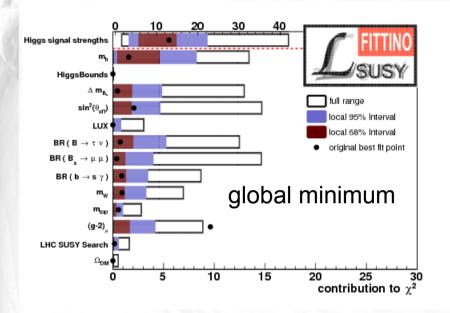


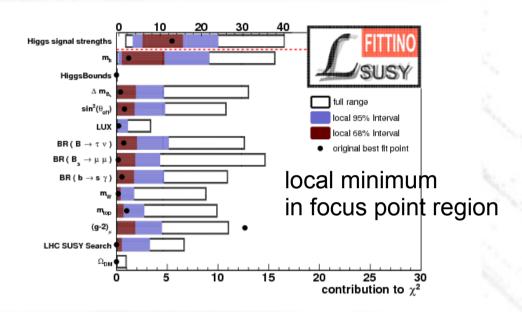
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- → local minimum in focus point region
 - \rightarrow same observable set
 - \rightarrow higher χ^2
 - → higher p-value

Results IV: just one more thing . . .





- → appears to be a real effect
- → cross-checks done
 - → simple toy model with LHC-like limit produces similar effect
 - → effect of reduced sampling density negligible
 - → tested more points in the focus point region

Summary

- ★ global fit of CMSSM with updated observables
 - → low energy measurements
 - → measurements from cosmology/astrophysics
 - → results of direct searches
 - → Higgs observables
- * LHC limits push the CMSSM to a region in which it can no longer accomodate (g-2)_{...}
- ★ accurate determination of the p-value requires pseudo experiments
 → in the ideal case at every single point in the parameter space
- ★ p-value depends strongly on the choice of the observable set
 → combination of measurements with identical predictions crucial
- * we exclude the CMSSM at the 90% CL with the optimal observable set