

# SUPERSYMMETRIC DARK MATTER AFTER LHC RUN I

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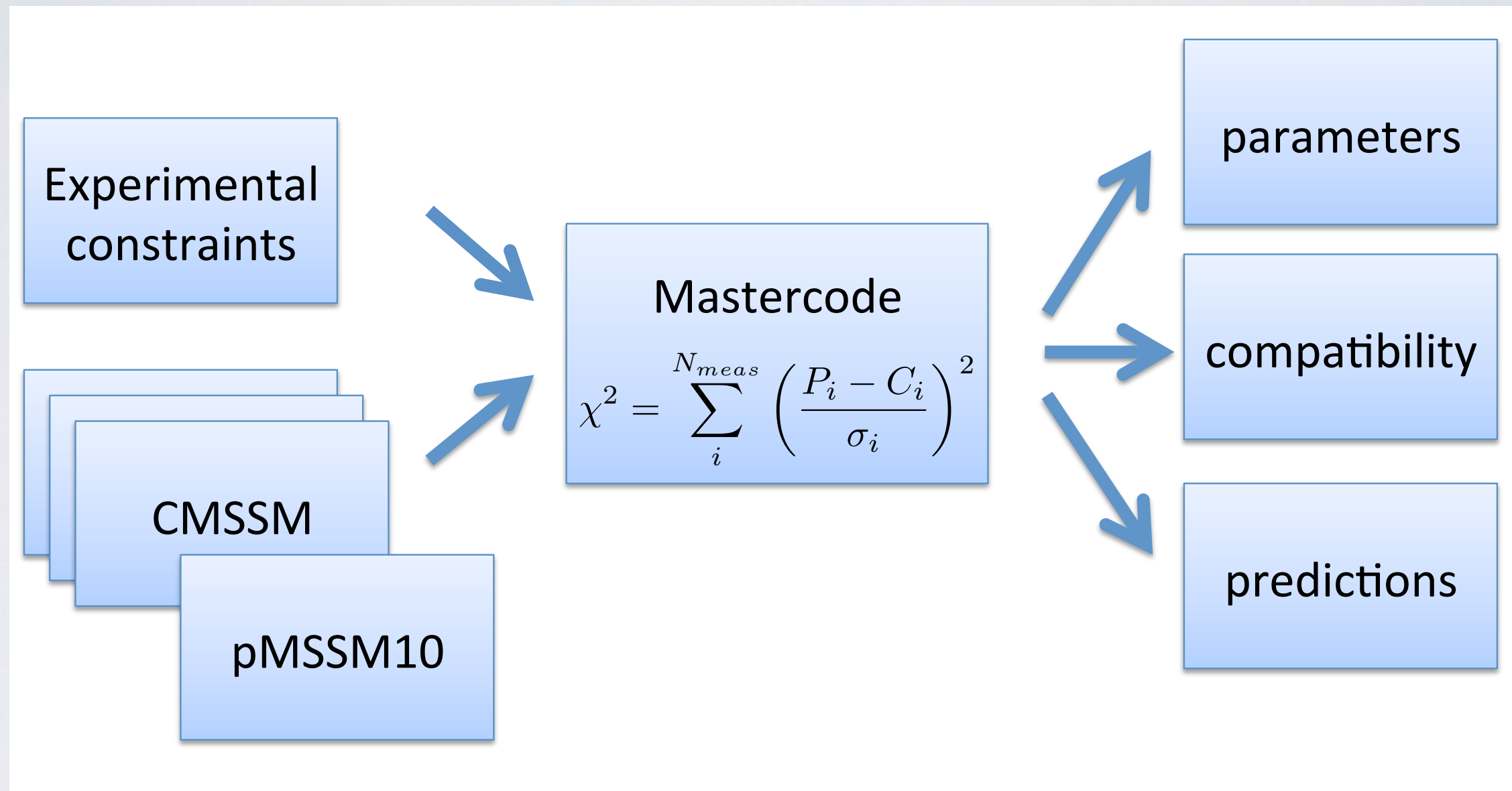




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- Joint theory and experimental collaboration.
- Experimental: CMS, LHCb
- Theory: SUSY, DM, Flavour, Precision Electroweak/Higgs

# THE GLOBAL FIT GAME



# EXPERIMENTAL CONSTRAINTS

We use a suite of constraints from

- Higgs Physics
- Precision Electroweak
- Direct Detection and Cosmology
- Flavour Physics
- LHC SUSY Searches

SoftSusy, FEWZ, FeynHiggs, SuFla, SuperIso, Micromegas, SSARD,  
HiggsSignals, HiggsBounds, ATOM

See also yesterday's talk.



# DARK MATTER PHENOMENOLOGY

Fits provide a rich dataset.

How is relic density set in the pMSSM?

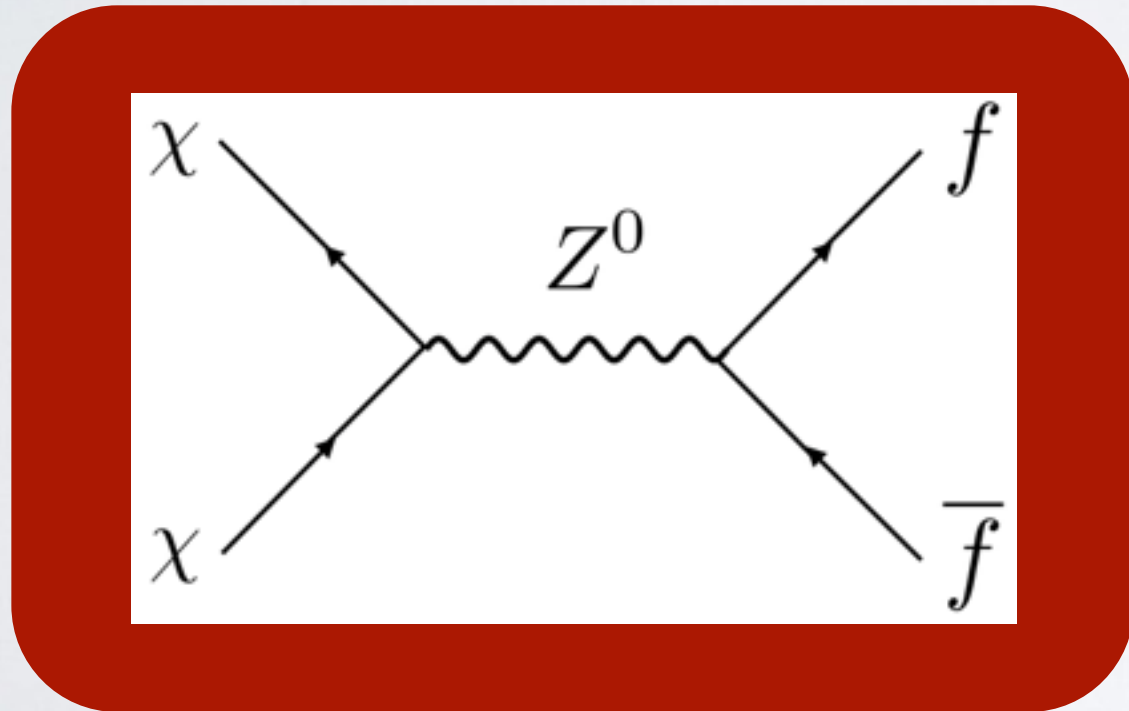
How does LHC probe the pMSSM by mechanism?

Direct detection prospects?

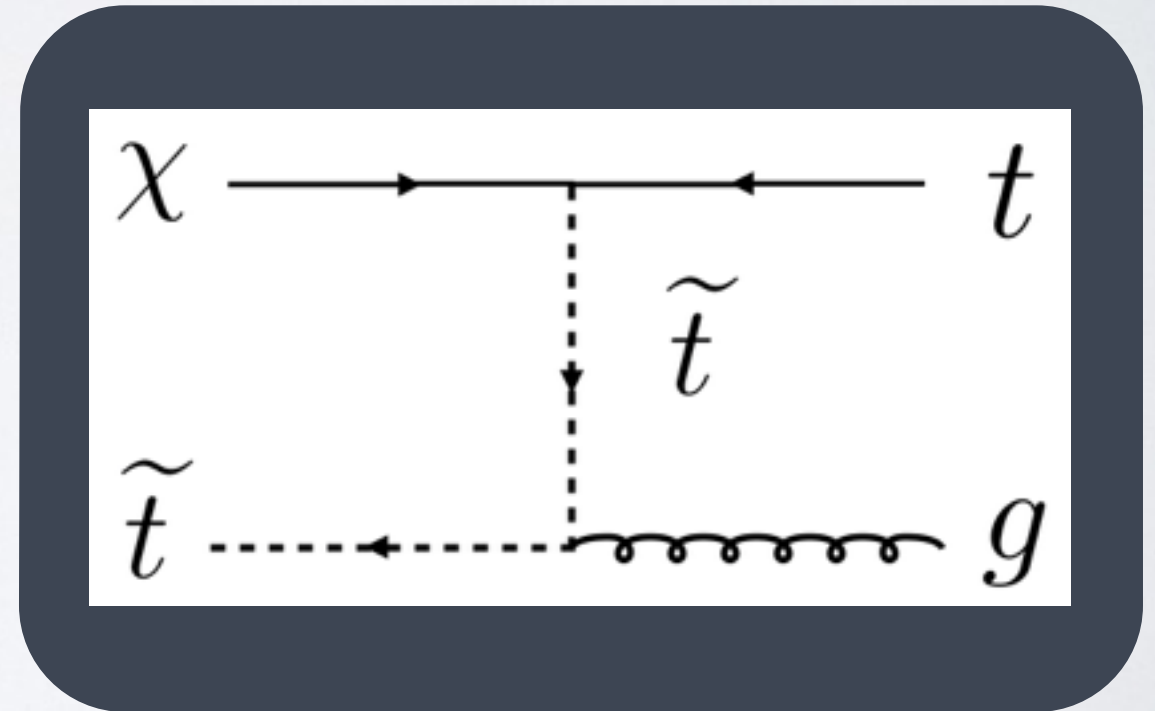
# RELIC DENSITY MECHANISMS

Relic density depletion requires relations between sparticle masses.

In the MSSM this happens through resonant DM annihilation ('funnel') or co-annihilation



Resonant/funnel/s-channel



Co-annihilation/t-channel

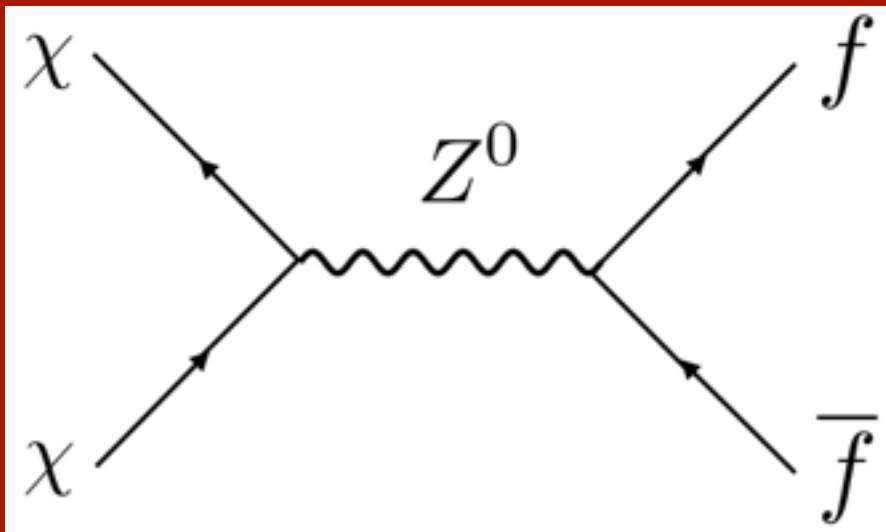
# RELIC DENSITY MECHANISMS

Also for:

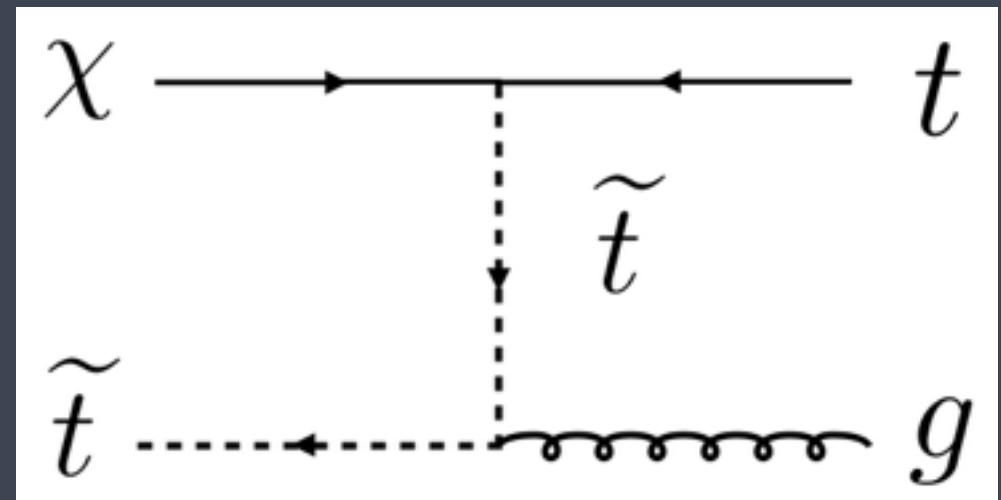
- Light Higgs  $h$
- Heavy Higgs  $A/H$

Also for:

- stau co-annihilation
- chargino co-annihilation



Resonant/funnel/s-channel



Co-annihilation/t-channel

# RELIC DENSITY MECHANISMS

How to quantify this?

$$\tilde{\tau}_1 \text{ coann. (pink)} : \left( \frac{m_{\tilde{\tau}_1}}{m_{\tilde{\chi}_1^0}} - 1 \right) < 0.15,$$

$$\tilde{\chi}_1^\pm \text{ coann. (green)} : \left( \frac{m_{\tilde{\chi}_1^\pm}}{m_{\tilde{\chi}_1^0}} - 1 \right) < 0.1,$$

$$\tilde{t}_1 \text{ coann. (grey)} : \left( \frac{m_{\tilde{t}_1}}{m_{\tilde{\chi}_1^0}} - 1 \right) < 0.2,$$

$$A/H \text{ funnel (blue)} : \left| \frac{M_A}{m_{\tilde{\chi}_1^0}} - 2 \right| < 0.4,$$

$$\text{focus point (cyan)} : \left( \frac{\mu}{m_{\tilde{\chi}_1^0}} - 1 \right) < 0.3.$$

$$h \text{ funnel (magenta)} : \left| \frac{M_h}{m_{\tilde{\chi}_1^0}} - 2 \right| < 0.4,$$

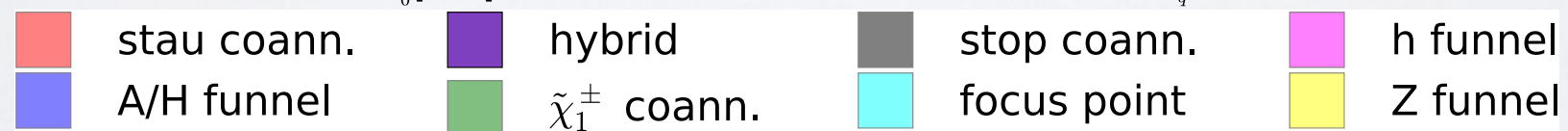
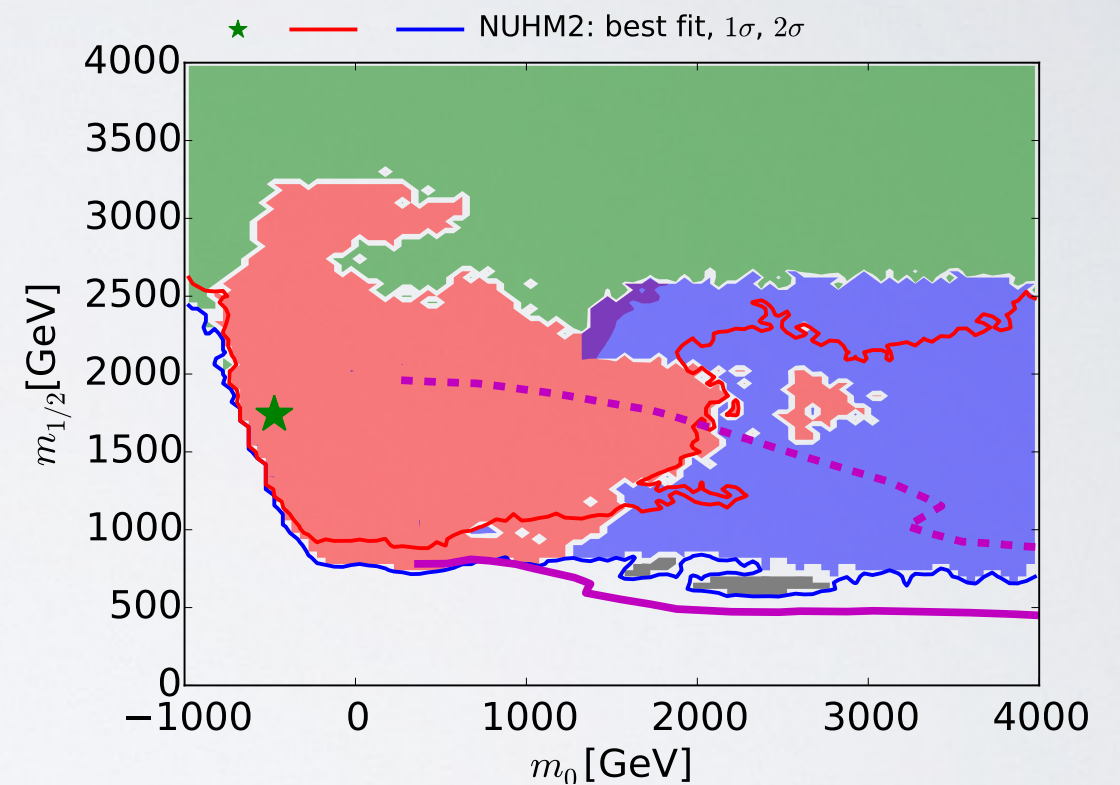
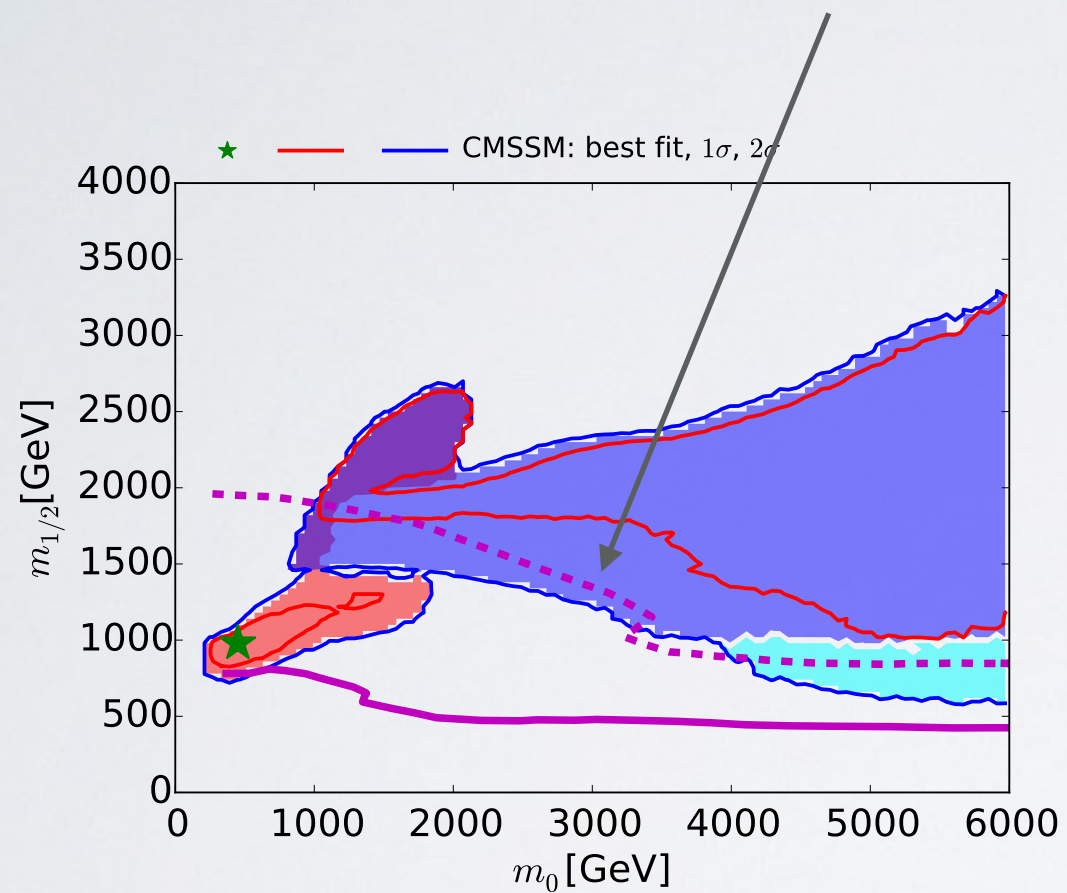
$$Z \text{ funnel (orange)} : \left| \frac{M_Z}{m_{\tilde{\chi}_1^0}} - 2 \right| < 0.4.$$

Conditions cross-checked from Micromegas output



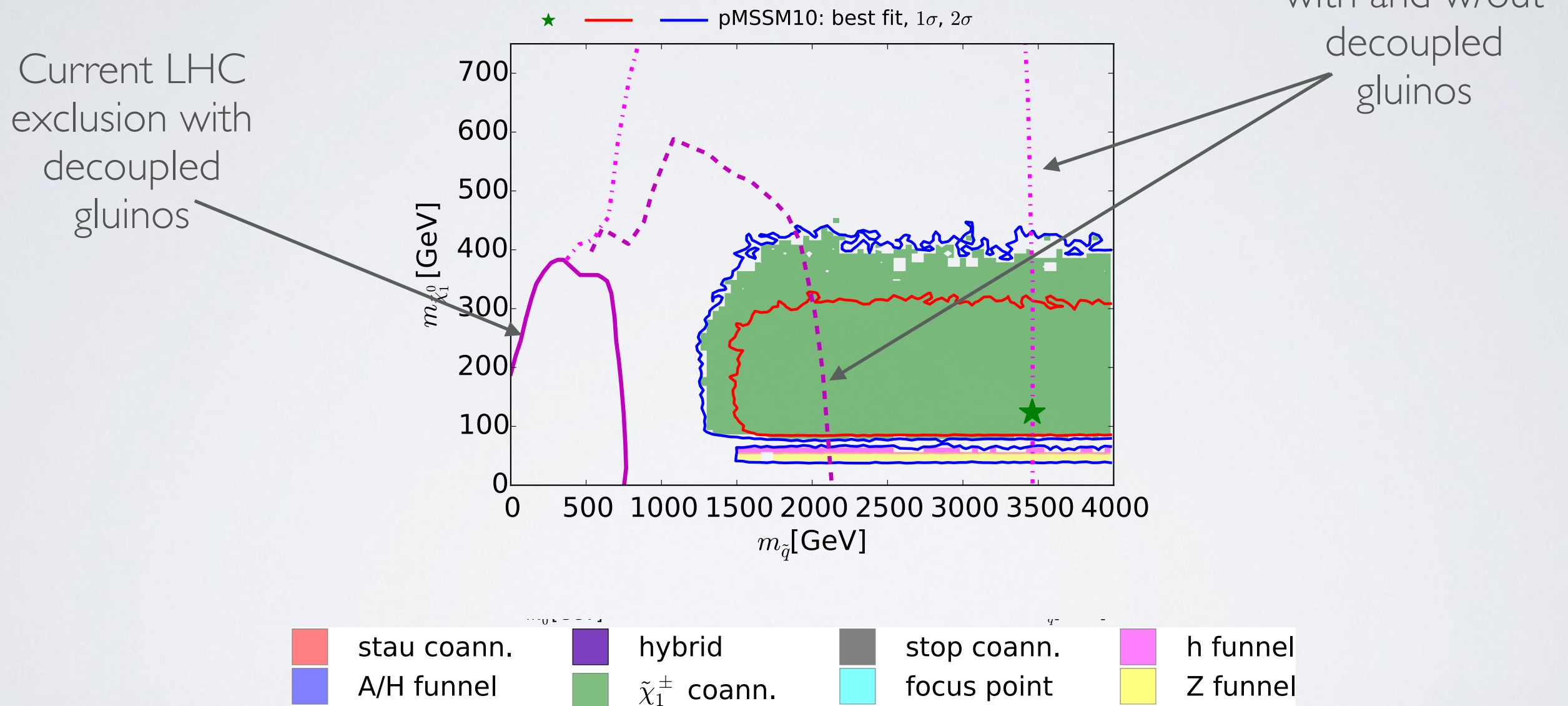
# CONSTRAINED MODELS

300/fb 95% exclusion estimate



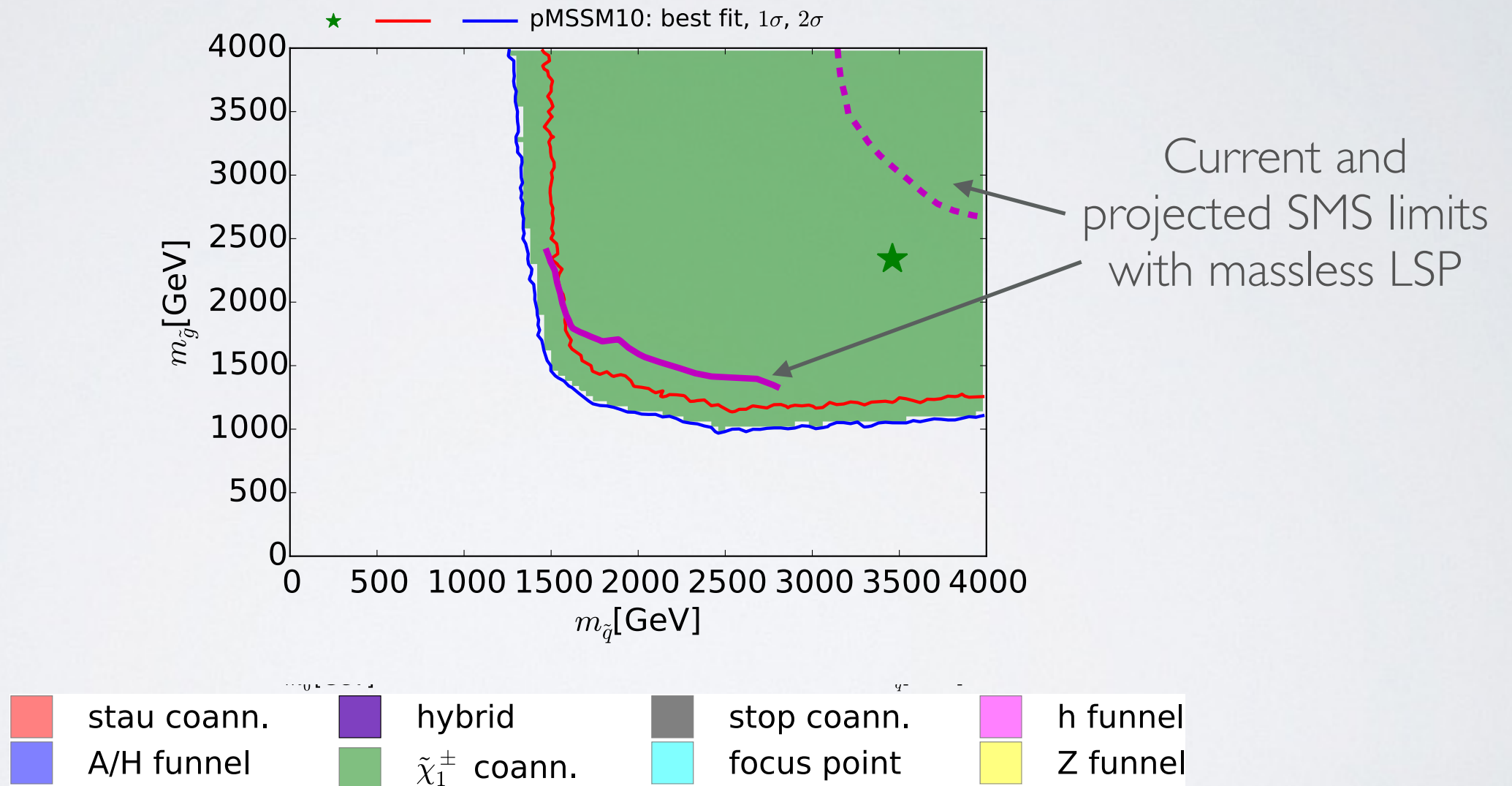
# PMSSM RESULTS

## Squark Neutralino mass plane



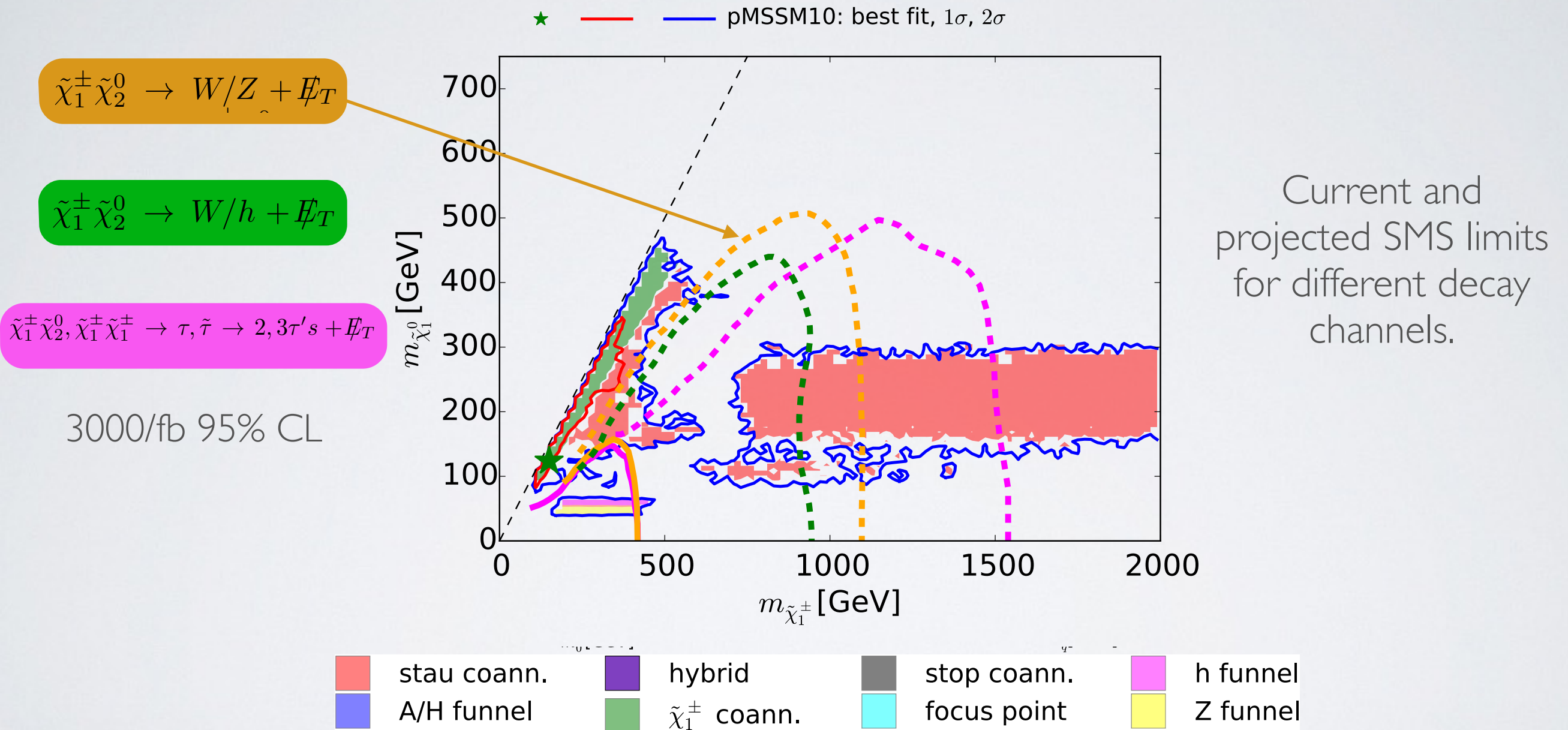
# PMSSM RESULTS

## Squark Gluino mass plane



# PMSSM RESULTS

## Lightest chargino-neutralino mass plane

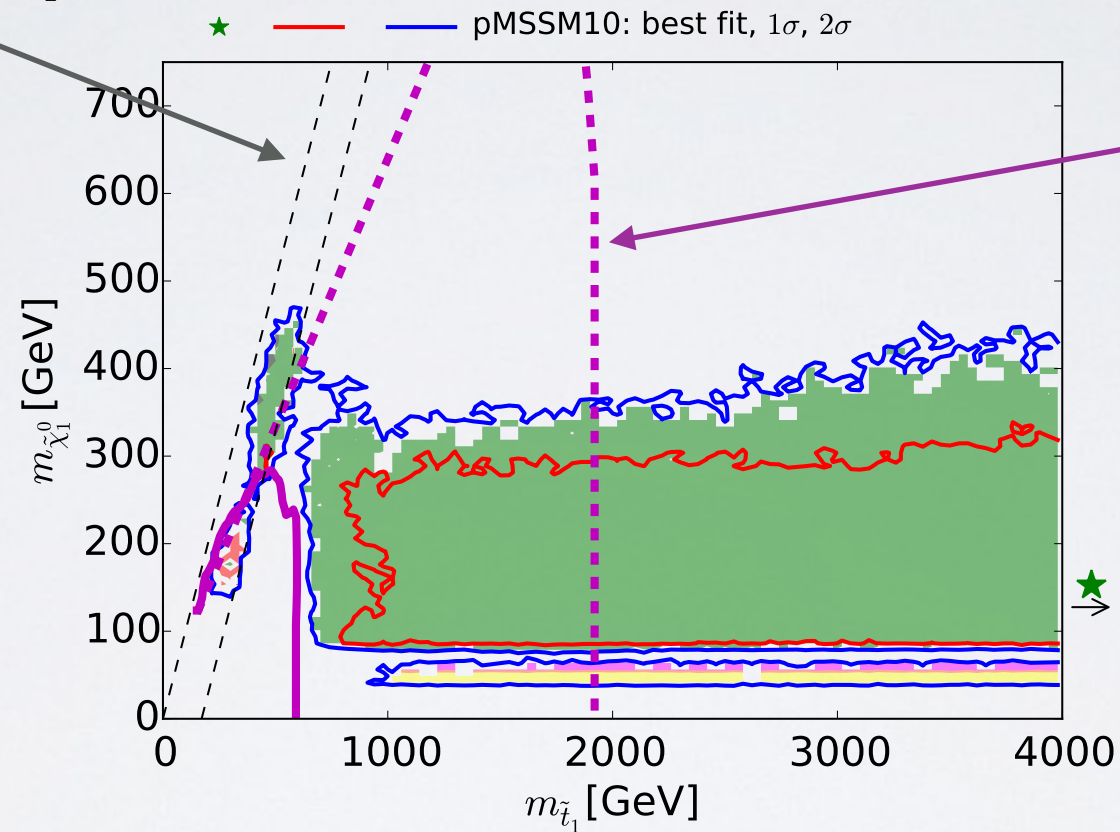




# PMSSM RESULTS

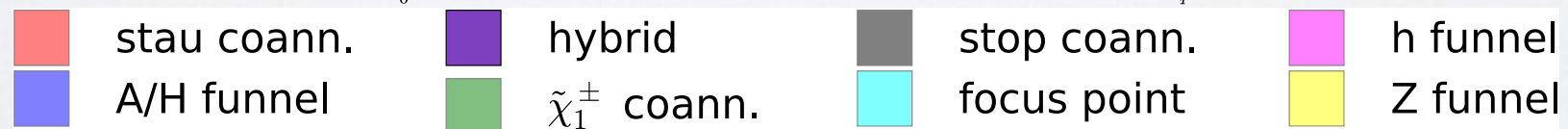
## Stop neutralino mass plane

$$m_{\tilde{t}_1} = m_{\tilde{\chi}_1^0} \text{ and } m_{\tilde{t}_1} = m_t + m_{\tilde{\chi}_1^0}.$$



$$\tilde{t}_1 \rightarrow b \tilde{\chi}_1^\pm$$

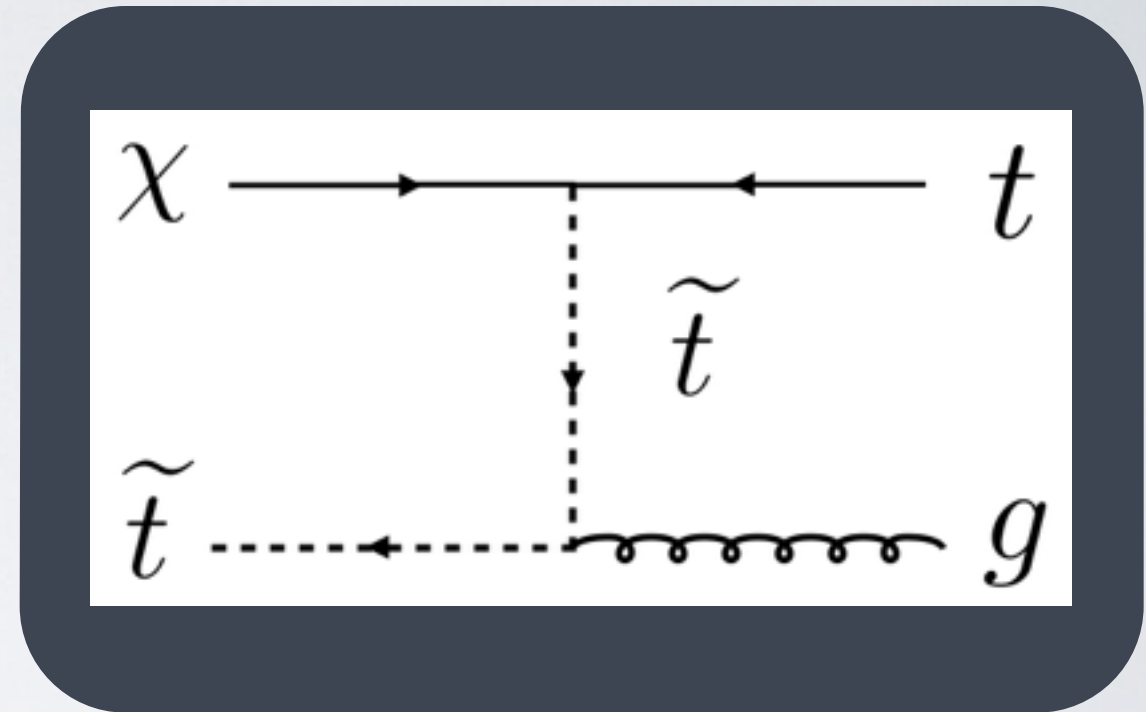
3000/fb 95% CL  
100% BR



# PMSSM RESULTS

Co-annihilation requires  
LSP and other sparticle to  
be near degenerate.

Possibility of long-lived  
particles?



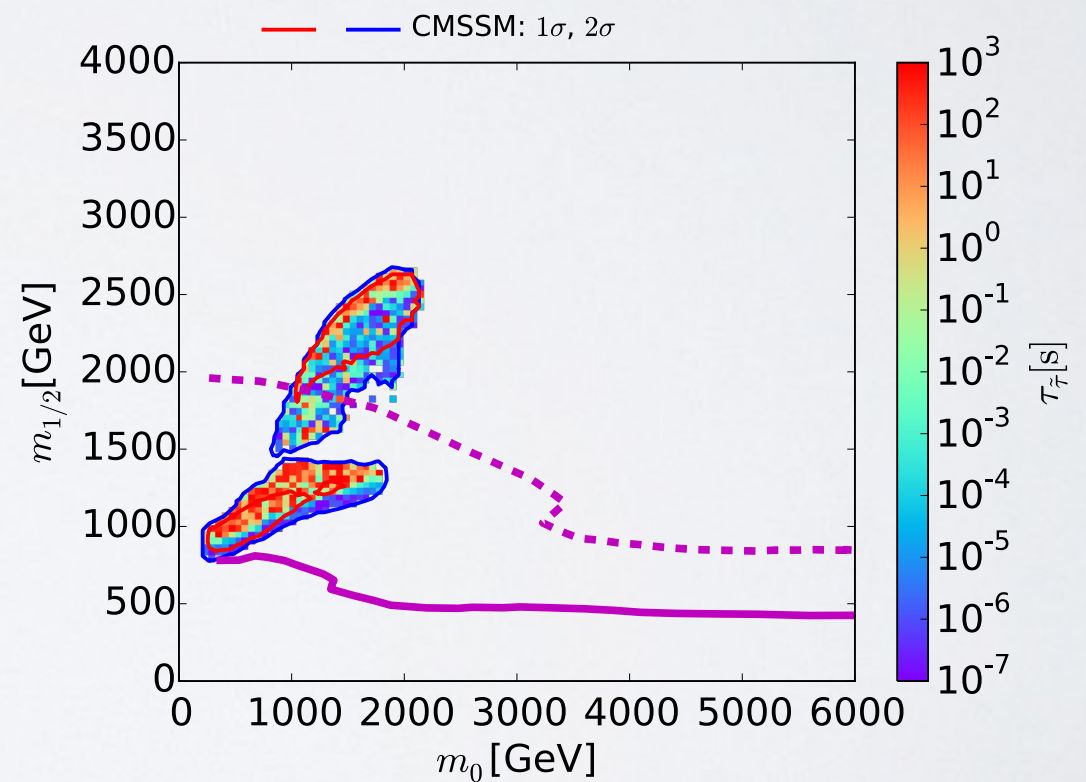
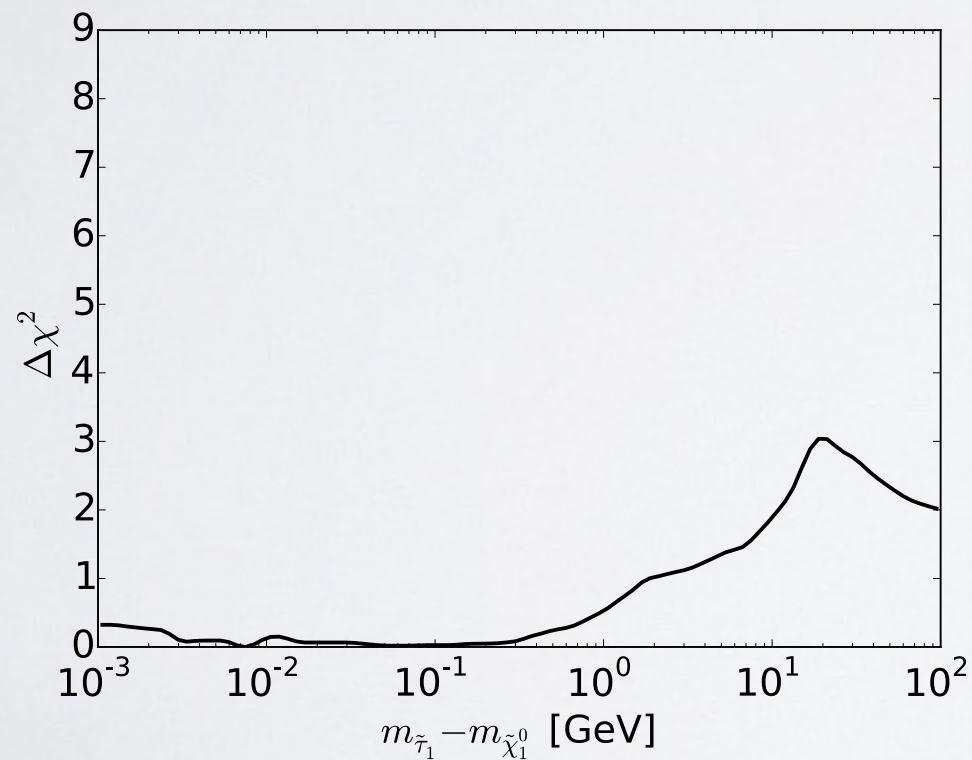
This possibility is not realized in the pMSSM

# LONG LIVED SPARTICLES

In constrained models, squark/gluino limits also push up the LSP mass

Heavier LSP implies greater NLSP degeneracy for correct annihilation cross-section

CMSSM

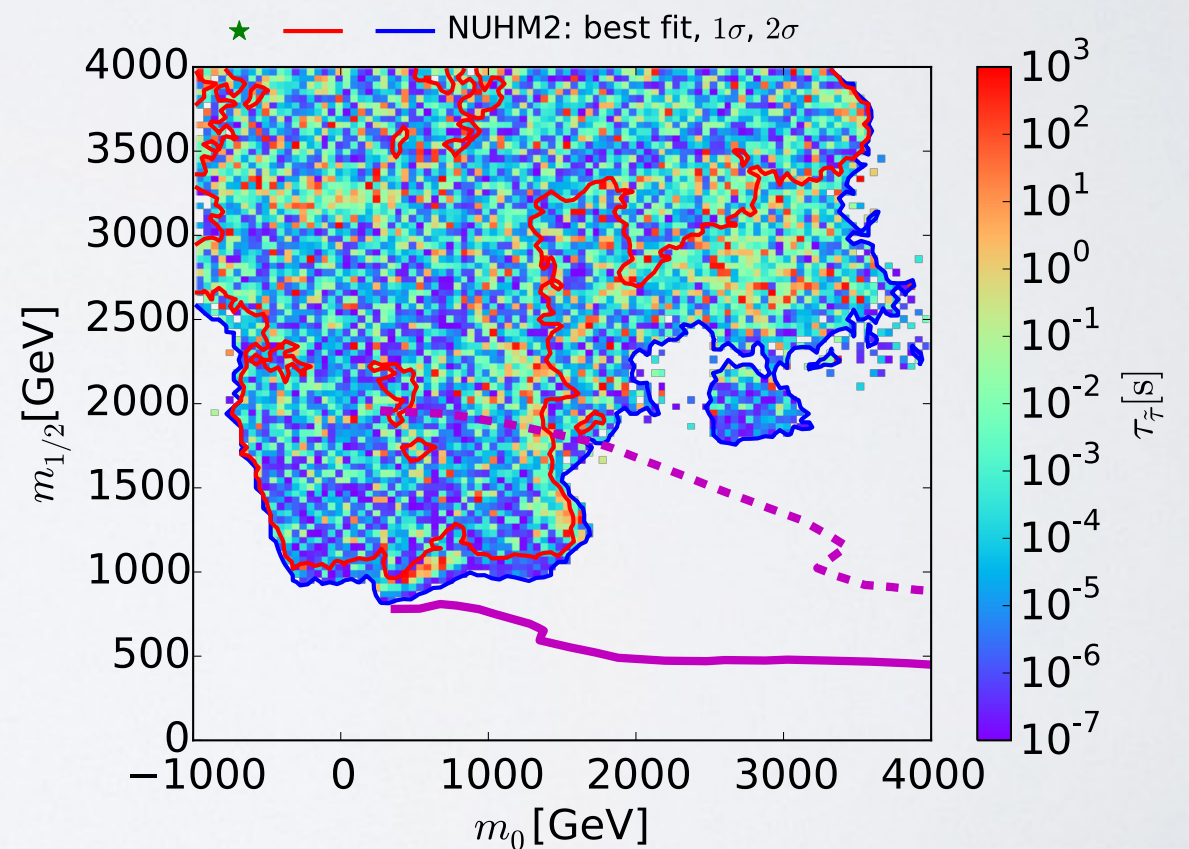
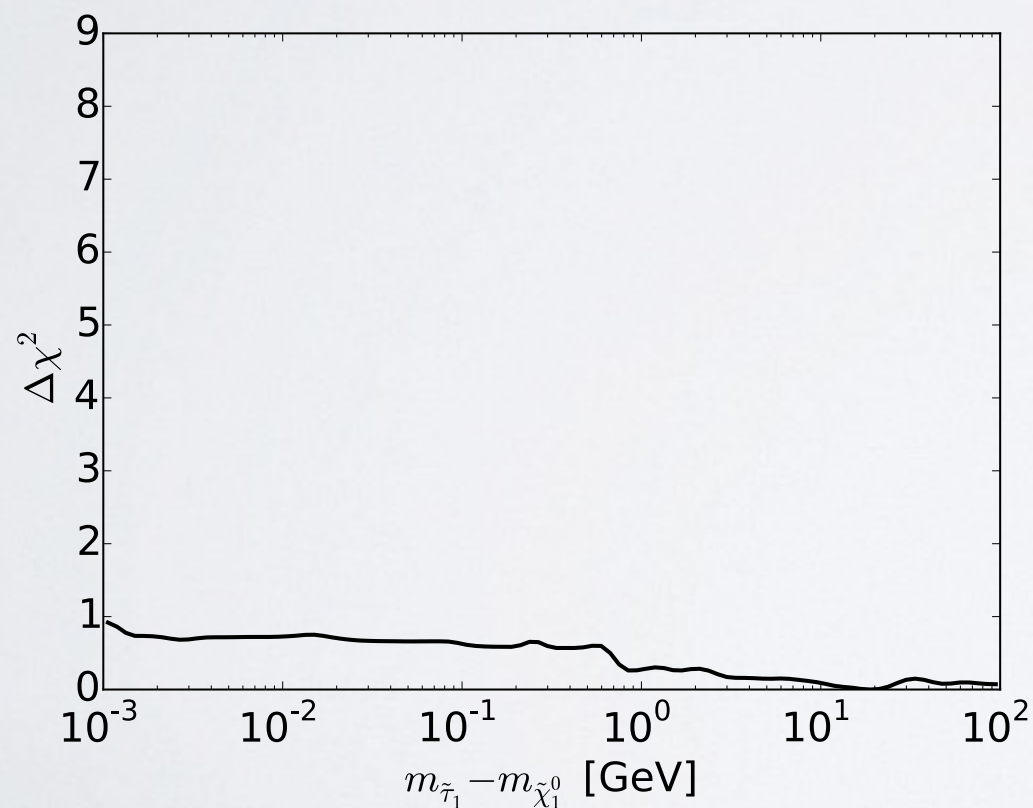


# LONG LIVED SPARTICLES

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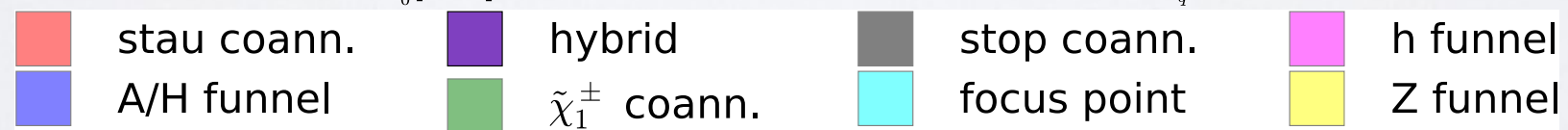
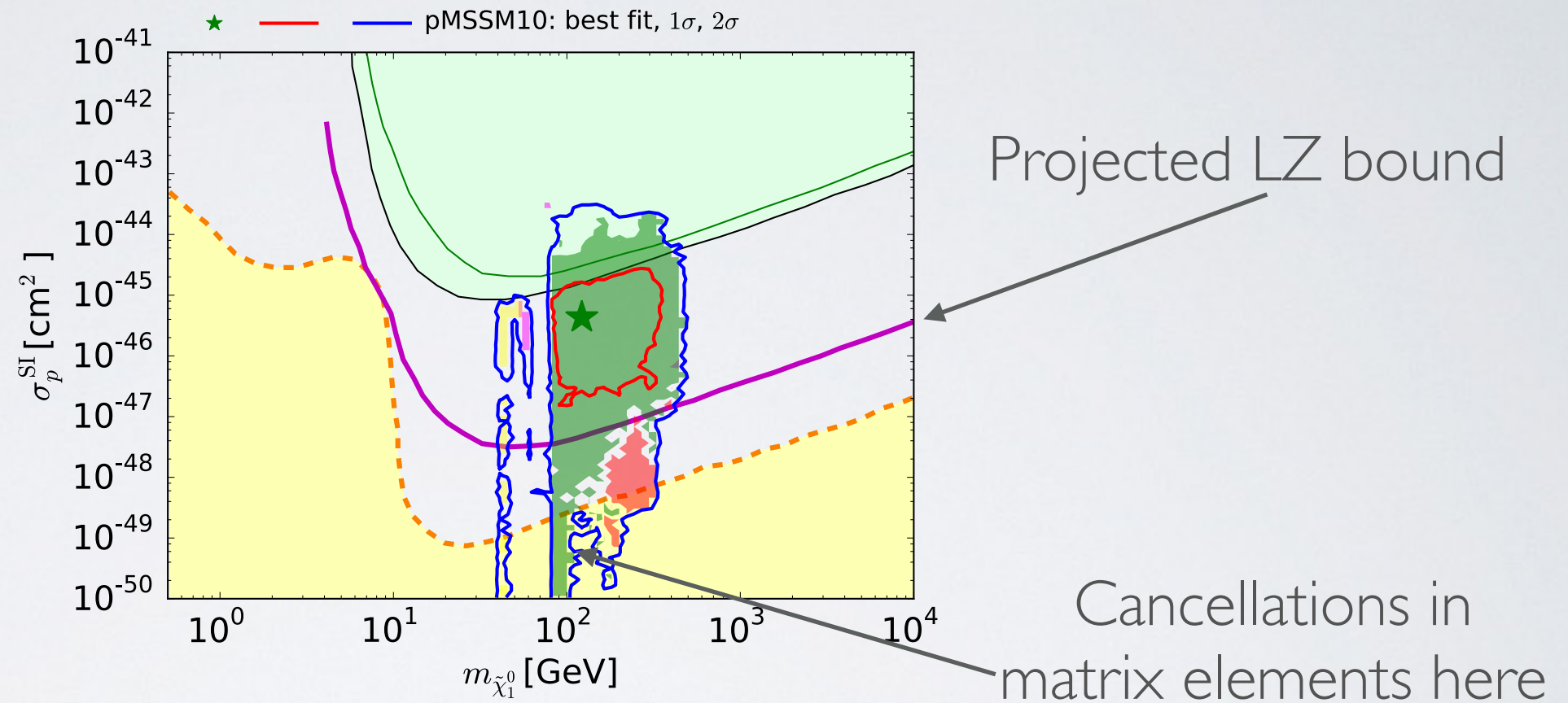
Heavier LSP implies greater NLSP degeneracy for correct annihilation cross-section

NUHM2





# DIRECT DETECTION PHENOMENOLOGY



# SUMMARY OF DETECTABILITY

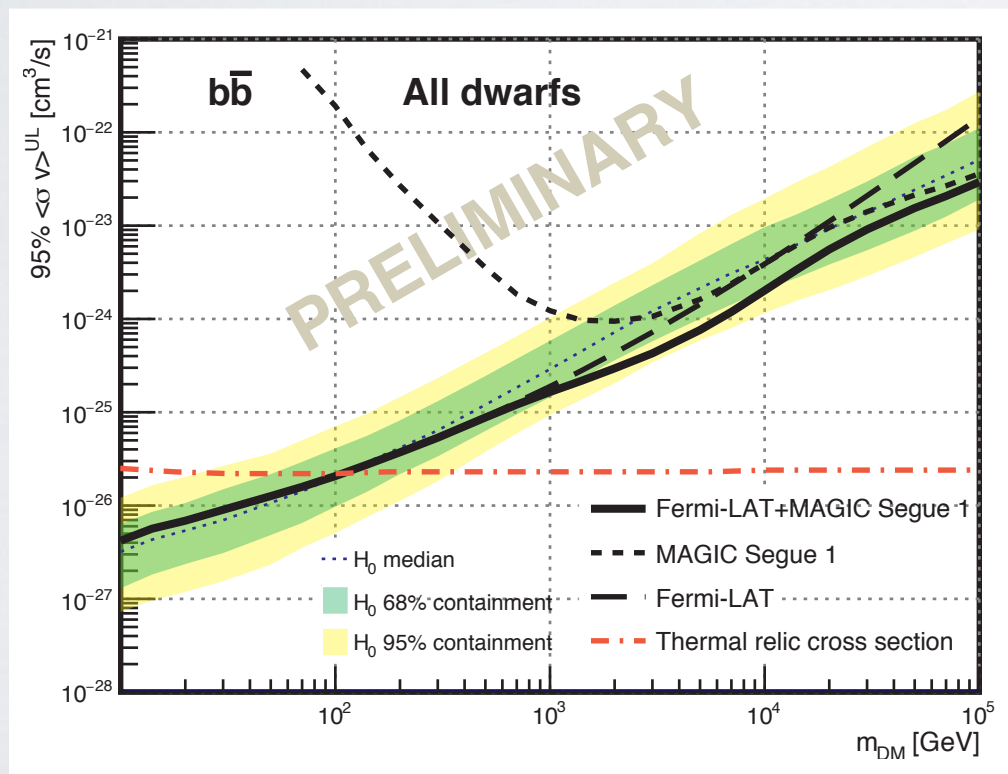
DM mechanism	Exp't	Models			
		CMSSM	NUHM1	NUHM2	pMSSM10
$\tilde{\tau}_1$ coann.	LHC DM	$\checkmark \cancel{E}_T, \checkmark \text{ LL}$ ( $\checkmark$ )	( $\checkmark \cancel{E}_T, \checkmark \text{ LL}$ ) ( $\checkmark$ )	( $\checkmark \cancel{E}_T, \checkmark \text{ LL}$ ) $\times$	( $\checkmark \cancel{E}_T$ ), $\times \text{ LL}$ $\times$
$\tilde{\chi}_1^\pm$ coann.	LHC DM	— —	$\times$ $\checkmark$	$\times$ $\checkmark$	( $\checkmark \cancel{E}_T$ ) ( $\checkmark$ )
$\tilde{t}_1$ coann.	LHC DM	— —	— —	$\checkmark \cancel{E}_T$ $\times$	— —
$A/H$ funnel	LHC DM	$\checkmark A/H$ $\checkmark$	( $\checkmark A/H$ ) $\checkmark$	( $\checkmark A/H$ ) ( $\checkmark$ )	— —
Focus point	LHC DM	( $\checkmark \cancel{E}_T$ ) $\checkmark$	— —	— —	— —
$h, Z$ funnels	LHC DM	— —	— —	— —	( $\checkmark \cancel{E}_T$ ) ( $\checkmark$ )

# INDIRECT DETECTION

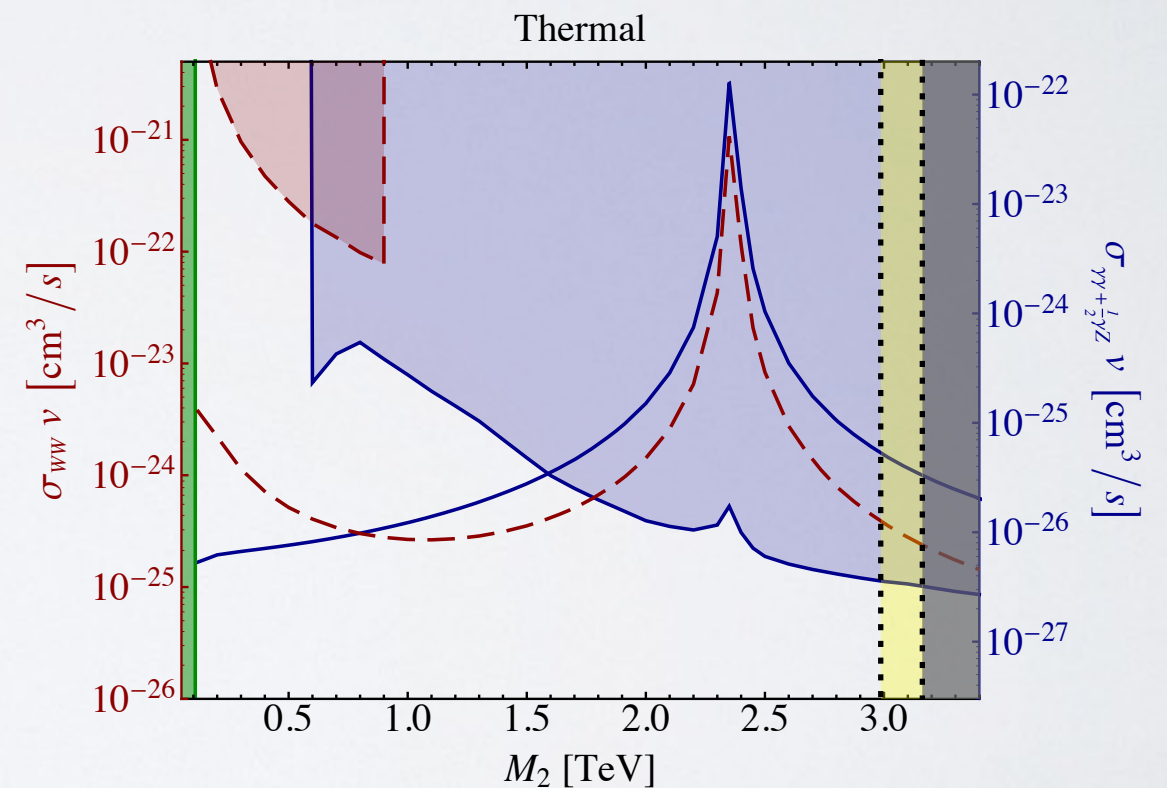
Also Weniger's talk

Will be interesting to incorporate recent Fermi-LAT dwarf satellite limits

Indirect detection constraints from Fermi/HESS constrain heavier (wino) states



Rico, Wood, Drlica-Wagner, Aleksic 2015



Cohen, Lisanti, Pierce, Slatyer 2013

# SUMMARY

DD and Collider searches will probe variety of DM mechanisms in near future

Charged track searches in constrained models

Run II + CTA/SKA/HESS/Fermi-LAT = Interesting times ahead!