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Variability of the γ -ray emission from the binary PSR B1259-63

We examine temporal evolution of the observed γ -ray intensity from the direction of the binary PSR B1259-63 using only numbers of detected very high energetic photons. Simple and straightforward method is applied to the published data obtained with the Imaging Atmospheric Cherenkov Technique during campaigns around three periastron passages of the binary system. Changes attributable to the variations of the intrinsic source activity were recognized at high levels of significance.

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