

Nucleon correlations in neutrino oscillation experiments

1. Introduction
2. 2-body current in neutrino interactions
3. Neutrino oscillation experiments
4. Neutrino interaction measurements
5. Conclusion



Teppei Katori

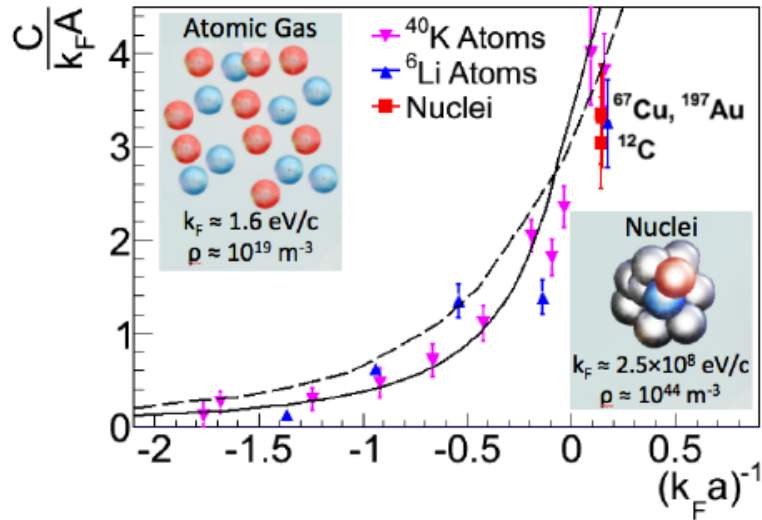
Queen Mary University of London

CLAS data mining group meeting, MIT, MA, USA, Aug. 9, 2014

1. Unexpected connection...

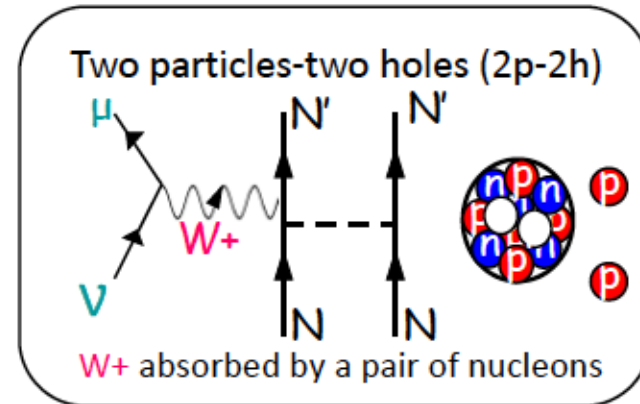
Atomic gas and nuclear gas

- They share same property in terms of "contact".



Short range correlation and neutrino oscillation experiments

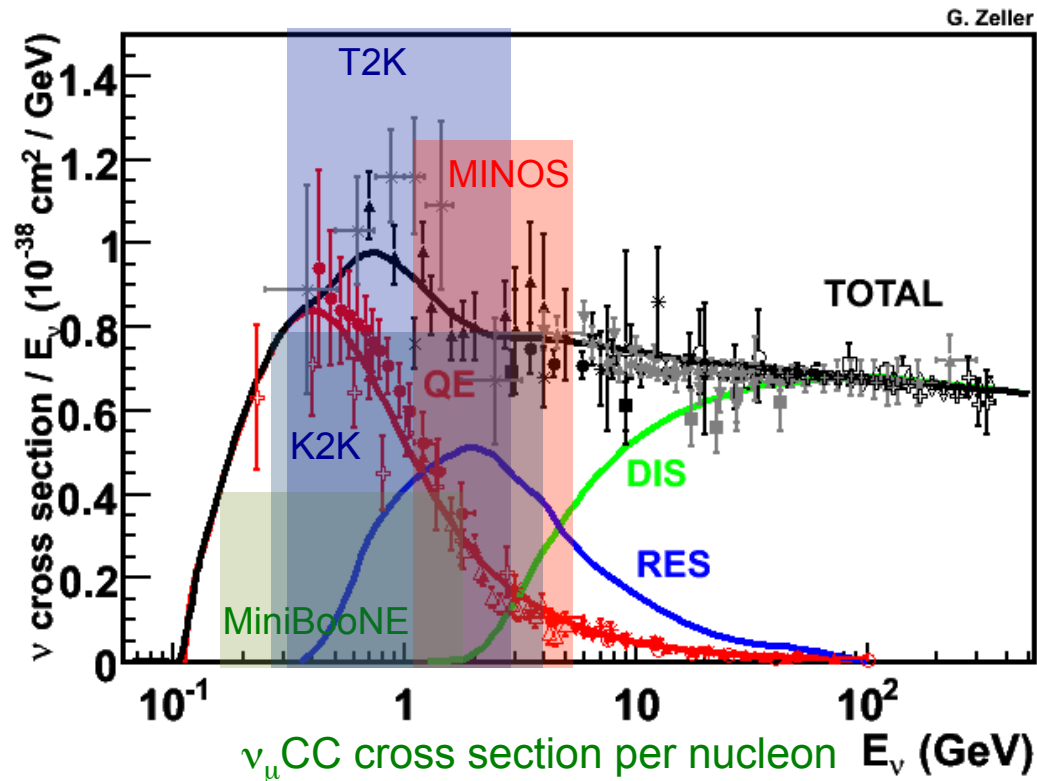
- Significant fraction of interaction(?)
- A correct model must be in our simulation



1. Next generation neutrino oscillation experiments

Neutrino oscillation experiments

- Past to Present: K2K, MiniBooNE, MINOS, T2K
- Present to Future: T2K, NOvA, PINGU, JUNO, HyperK, LBNF



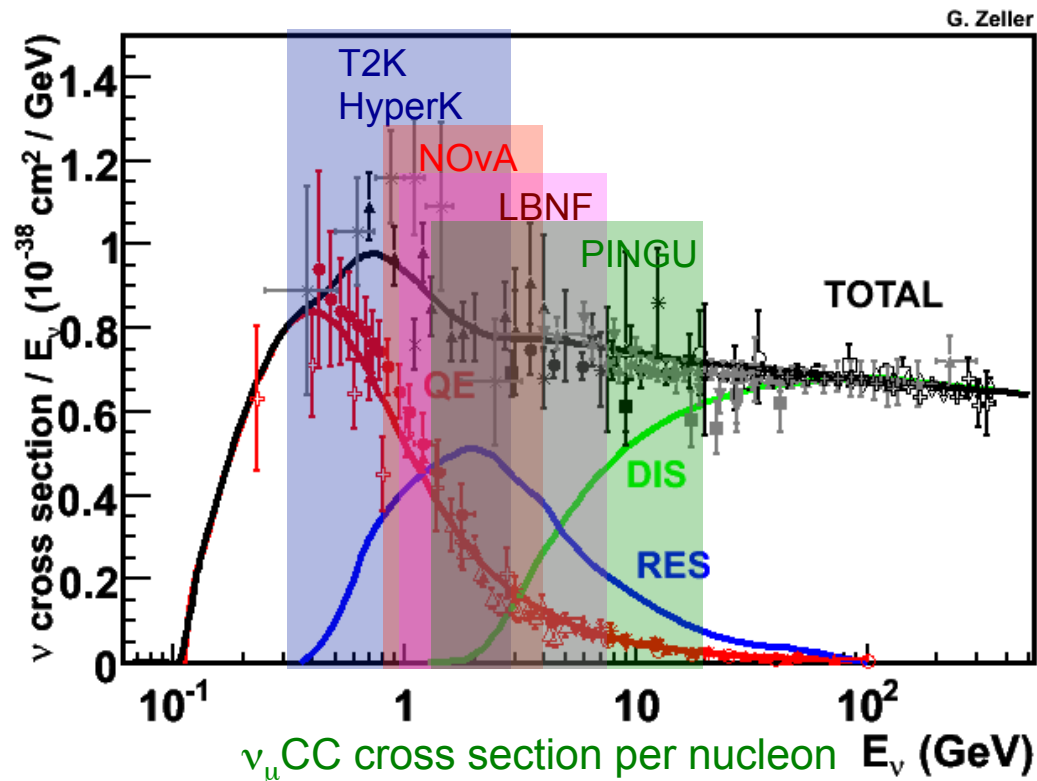
$$P_{\mu \rightarrow e}(L/E) = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 (eV^2) \frac{L(km)}{E(GeV)} \right)$$

1. Next generation neutrino oscillation experiments

Neutrino oscillation experiments

- Past to Present: K2K, MiniBooNE, MINOS, T2K
- Present to Future: T2K, NOvA, PINGU, JUNO, HyperK, LBNF

Typical oscillation experiment (L~100-1000km) always choose 1-10 GeV energy region (only exception is reactor neutrino experiment)



$$P_{\mu \rightarrow e}(L/E) = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 (eV^2) \frac{L(km)}{E(GeV)} \right)$$

1. Neutrino interaction model building

Adjust all vector response from available (e,e') data

e.g.) QE scattering (Fermi gas model)

$$(W_{\mu\nu})_{ab} = \int_{E_{lo}}^{E_{hi}} f(\vec{k}, \vec{q}, w) T_{\mu\nu} dE : \text{hadronic tensor}$$

$f(\vec{k}, \vec{q}, w)$: nucleon phase space distribution

$T_{\mu\nu} = T_{\mu\nu}(F_1, F_2, F_P, F_A)$: nucleon form factors

$F_A(Q^2) = g_A / (1 + Q^2/M_A^2)^2$: Axial vector form factor

1. Neutrino interaction model building

Adjust all vector response from available (e,e') data

known from (e,e') known from (e,e')

tiny ($\sim m^2/M^2$) less known from (e,e')

e.g.) QE scattering (Fermi gas model)

$$(W_{\mu\nu})_{ab} = \int_{E_{lo}}^{E_{hi}} f(\vec{k}, \vec{q}, w) T_{\mu\nu} dE : \text{hadronic tensor}$$

$f(\vec{k}, \vec{q}, w)$: nucleon phase space distribution

$T_{\mu\nu} = T_{\mu\nu}(F_1, F_2, F_P, F_A)$: nucleon form factors

$F_A(Q^2) = g_A / (1 + Q^2/M_A^2)^2$: Axial vector form factor

Often used a free parameter (we do know $M_A \sim 1$ from electro- π production data...)

QE: axial form factor is the only unknown part

Resonance: C_5^A form factor is the only unknown part, etc

1. Neutrino interaction model building

Adjust all vector response from available (e,e') data

→ all uncertainties are usually in axial part

1. Quasielastic

- all vector form factors are from (e,e') data (BBA form factors)
- axial form factor is the only unknown part
- assuming dipole form, axial mass is the only one parameter with large error

2. Delta resonance

- all vector form factors are from (e,e') data (MAID, for example)
- axial form factors are reduced to one form factor (Adler's theorem)
- C_5^A form factor is the only unknown part
- assuming dipole form, axial mass and $C_5^A(Q^2=0)$ are the two parameters with large errors

How to choose these parameters is a part of “open questions of neutrino interaction physics”

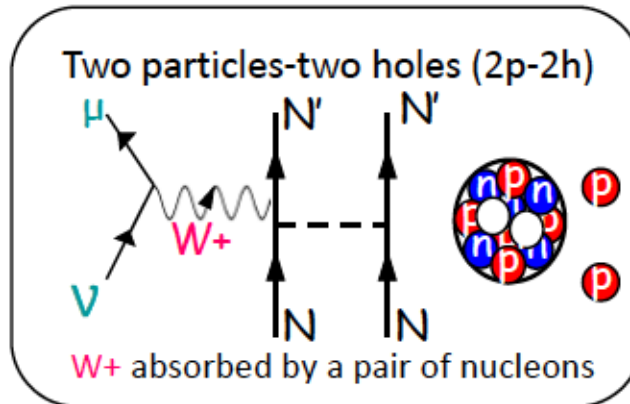
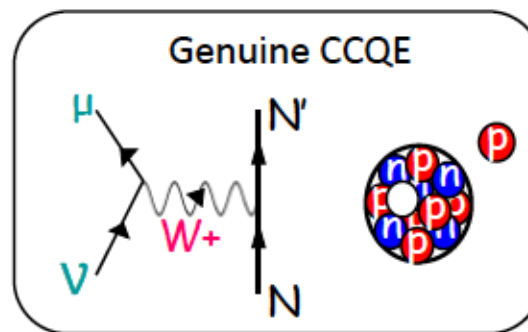
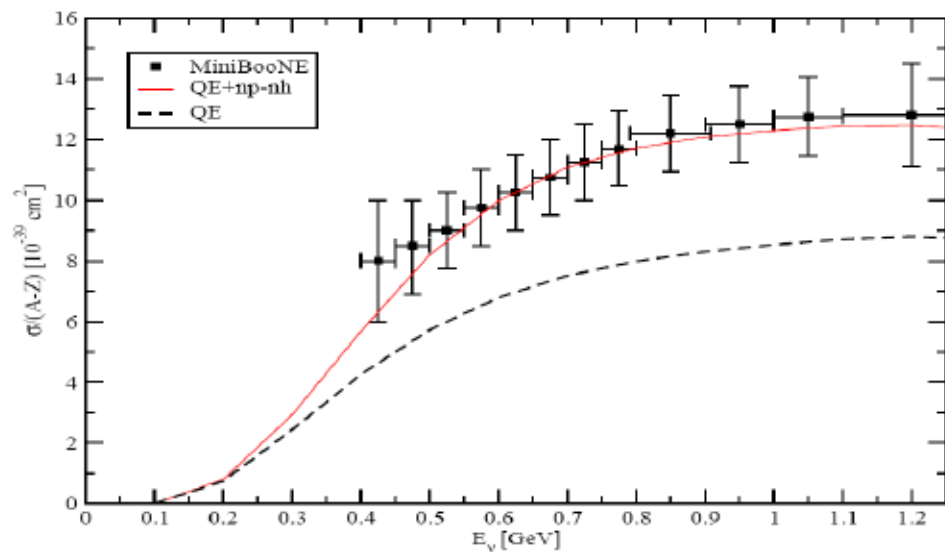
2. Presence of 2p-2h effect in neutrino interactions

Neutrino oscillation experiments

- Many modern neutrino experiments (K2K, MiniBooNE, etc) show data is higher than prediction.
- Martini et al showed np-nh effect can add up 30-40% more cross section!

An explanation of this puzzle

Inclusion of the multinucleon emission channel (np-nh)



M. Martini, M. Ericson, G. Chanfray, J. Marteau *Phys. Rev. C* 80 065501 (2009)

Agreement with MiniBooNE without increasing M_A

2. Presence of 2p-2h effect in neutrino interactions

Neutrino oscillation experiments

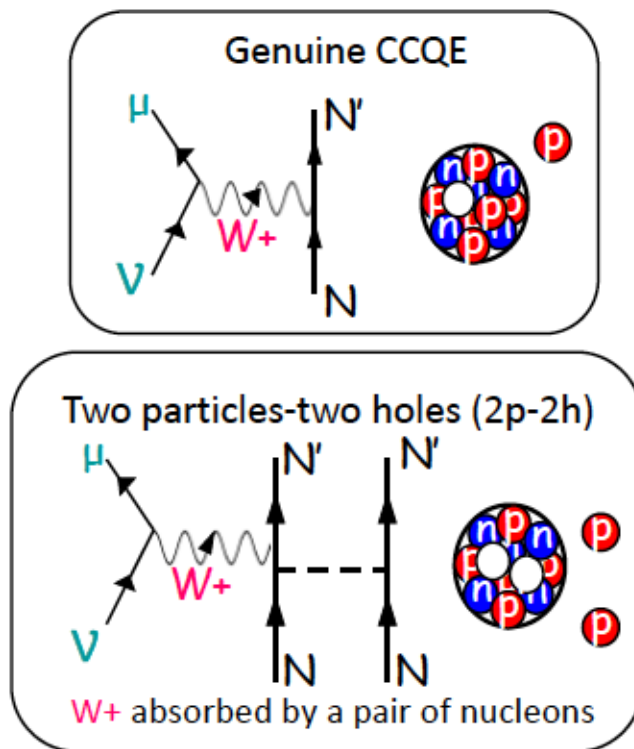
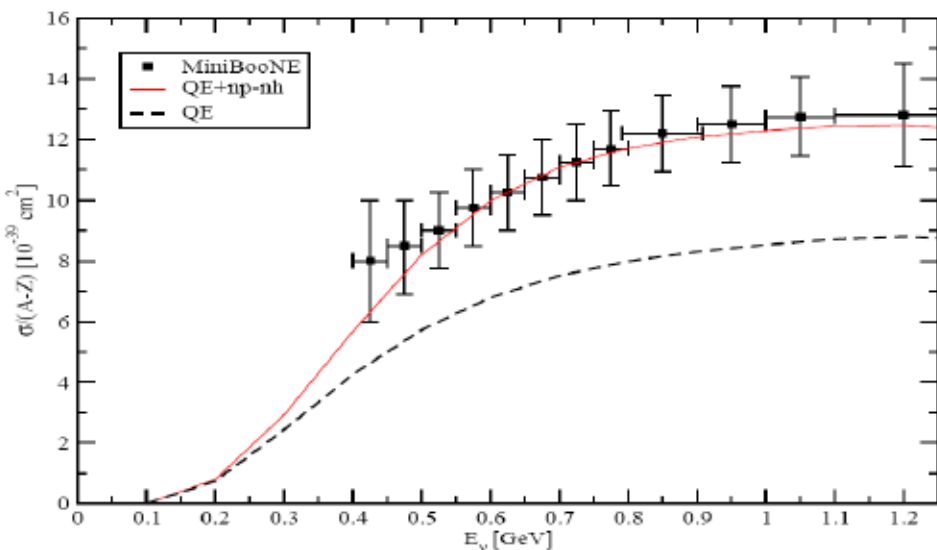
- Many modern neutrino experiments
- Martini et al showed np-nh effect

An explanation

$$E_\nu^{QE} = \frac{ME_\mu - 0.5m_\mu^2}{M - E_\mu + p_\mu \cos\theta_\mu} \quad \text{"QE assumption"}$$

$$Q_{QE}^2 = -m_\mu^2 + 2E_\nu^{QE}(E_\mu - p_\mu \cos\theta_\mu)$$

Inclusion of the multinucleon emission channel (np-nh)



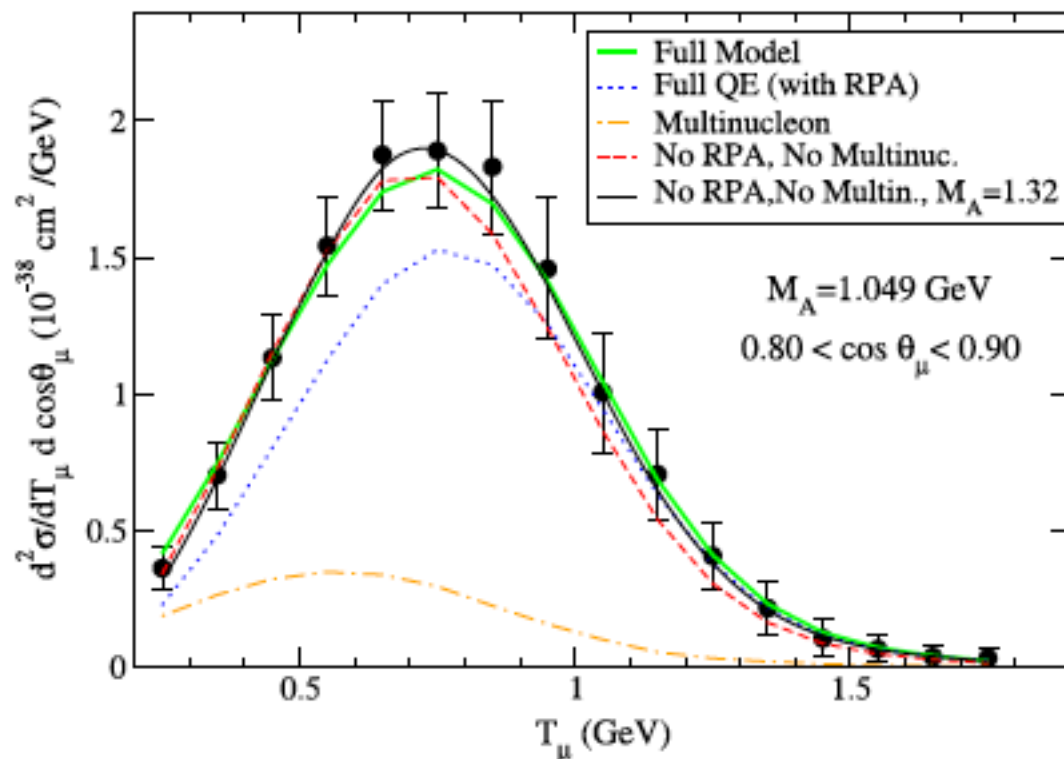
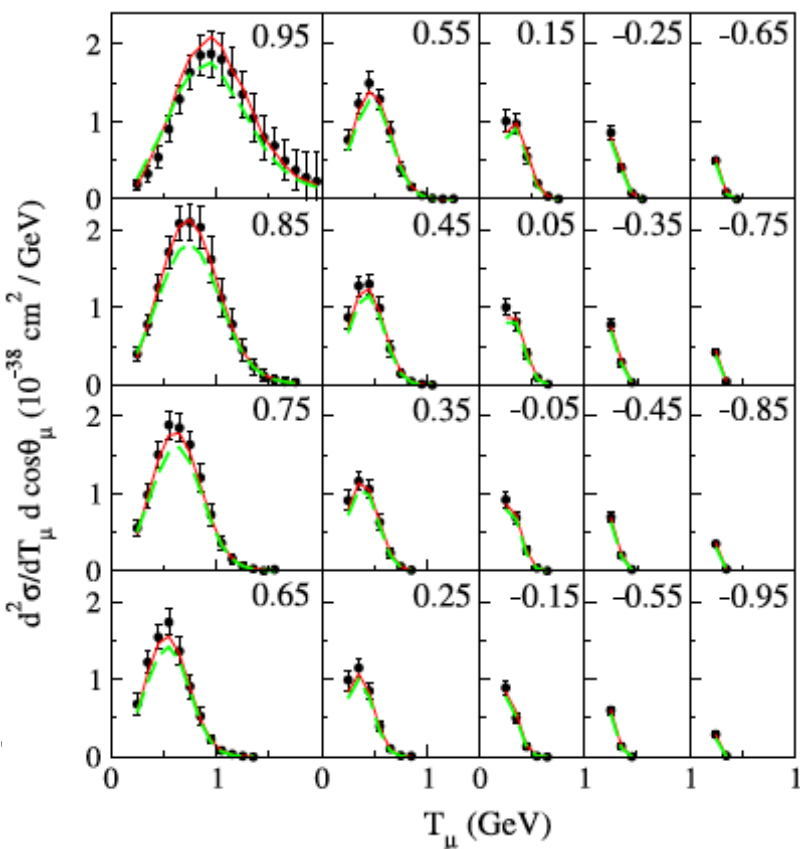
M. Martini, M. Ericson, G. Chanfray, J. Marteau *Phys. Rev. C* 80 065501 (2009)

Agreement with MiniBooNE without increasing M_A

2. Presence of 2p-2h effect in neutrino interactions

Neutrino oscillation experiments

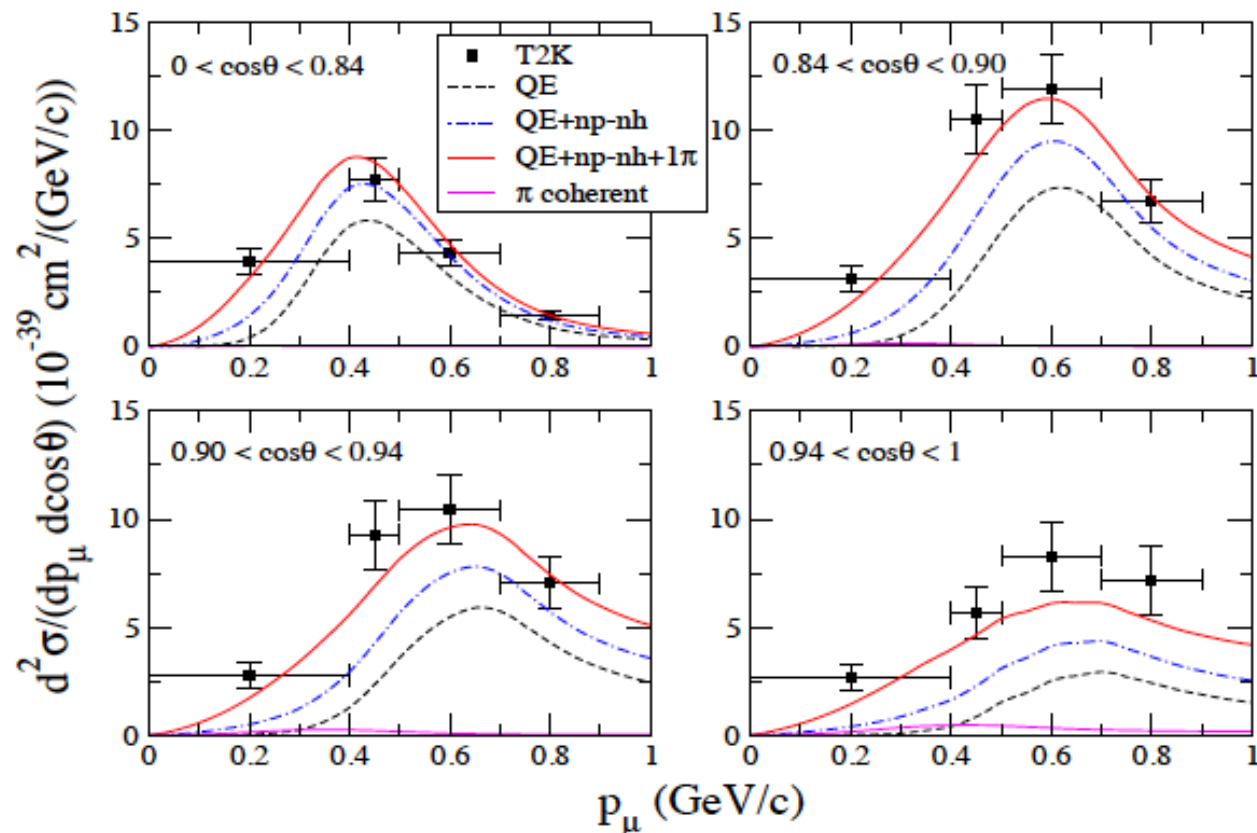
- Many modern neutrino experiments (K2K, MiniBooNE, etc) show data is higher than prediction.
- Martini et al showed np-nh effect can add up 30-40% more cross section!
- Later groups got similar results.
- Especially, MiniBooNE CCQE double differential cross section is explained by np-nh effect (and RPA)!



2. Presence of 2p-2h effect in neutrino interactions

Neutrino oscillation experiments

- Many modern neutrino experiments (K2K, MiniBooNE, etc) show data is higher than prediction.
- Martini et al showed np-nh effect can add up 30-40% more cross section!
- Later groups got similar results.
- Especially, MiniBooNE CCQE double differential cross section is explained by np-nh effect (and RPA)!
- Same model also describes T2K CC inclusive data (MiniBooNE flux prediction is right)



2. Presence of 2p-2h effect in neutrino interactions

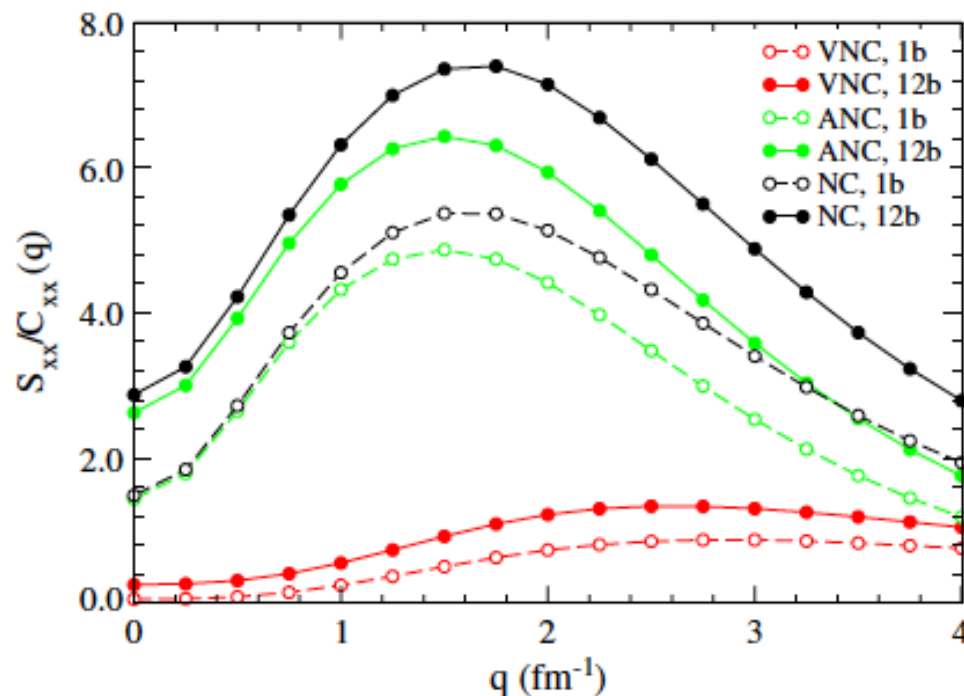
Neutrino oscillation experiments

- Many modern neutrino experiments (K2K, MiniBooNE, etc) show data is higher than prediction.
- Martini et al showed np-nh effect can add up 30-40% more cross section!
- Later groups got similar results.
- Especially, MiniBooNE CCQE double differential cross section is explained by np-nh effect (and RPA)!
- Same model also describes T2K CC inclusive data (MiniBooNE flux prediction is right)
- Consistent result is obtained by standard nuclear physics approach (SNPA, ab initio calculation)

It is difficult to translate to experimental observables, but they observed a similar large tensor correlation in axial vector part.

This enhancement is dominated by T=0 n-p pair.

If all these are true, this effect must be included in the simulation for better prediction of neutrino interactions.



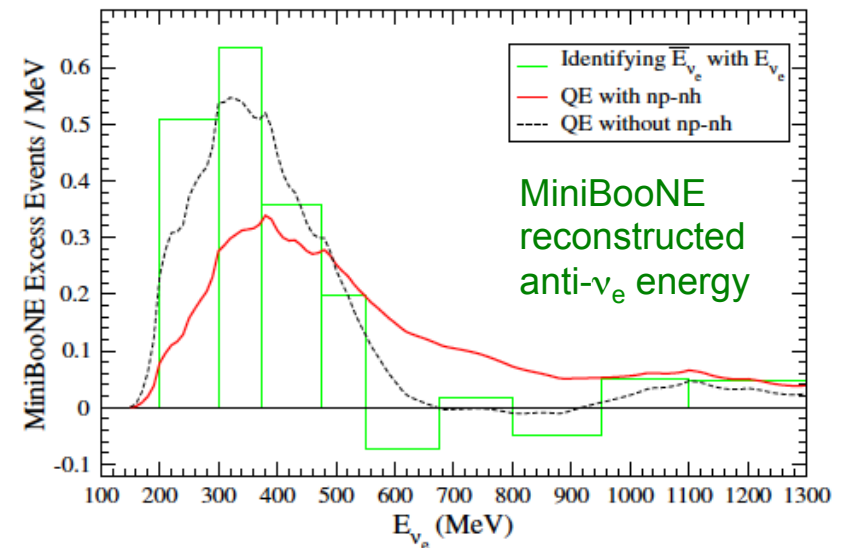
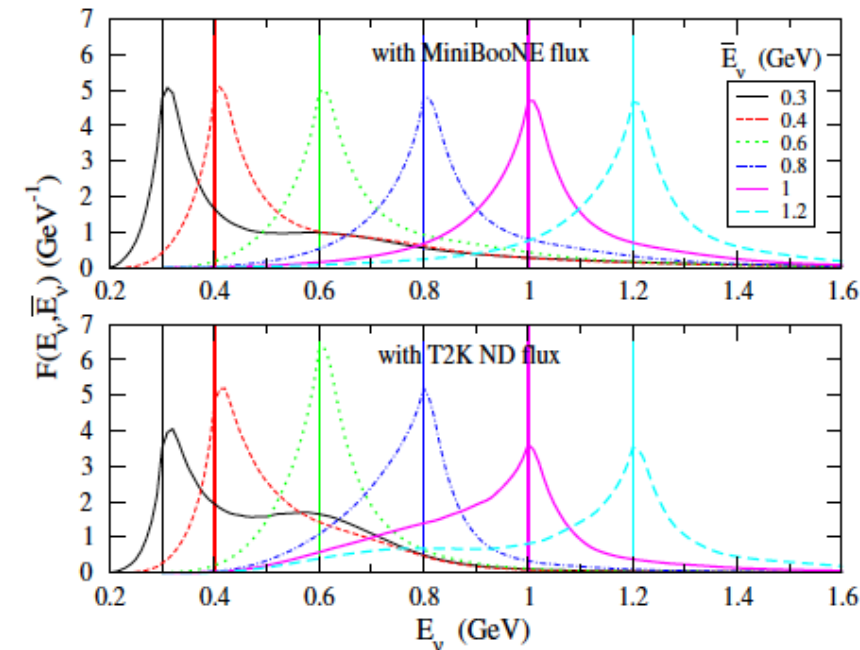
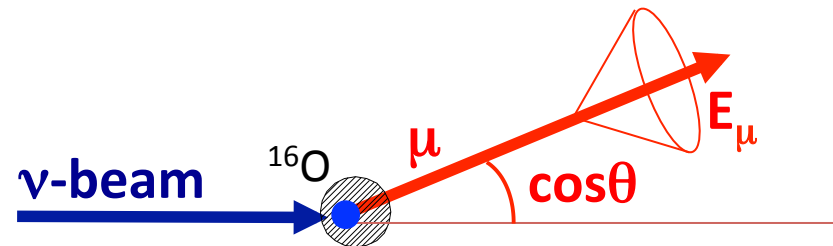
3. Neutrino oscillation experiment

Super-Kamiokande and Hyper-Kamiokande

- Neutrino energy reconstruction is based on lepton kinematics, assuming CCQE interaction.
- 2p-2h contribution mimics CCQE, and cause mis-calculation of neutrino energy.

$$E_{\nu}^{QE} = \frac{ME_{\mu} - 0.5m_{\mu}^2}{M - E_{\mu} + p_{\mu} \cos \theta_{\mu}}$$

$$Q_{QE}^2 = -m_{\mu}^2 + 2E_{\nu}^{QE}(E_{\mu} - p_{\mu} \cos \theta_{\mu})$$

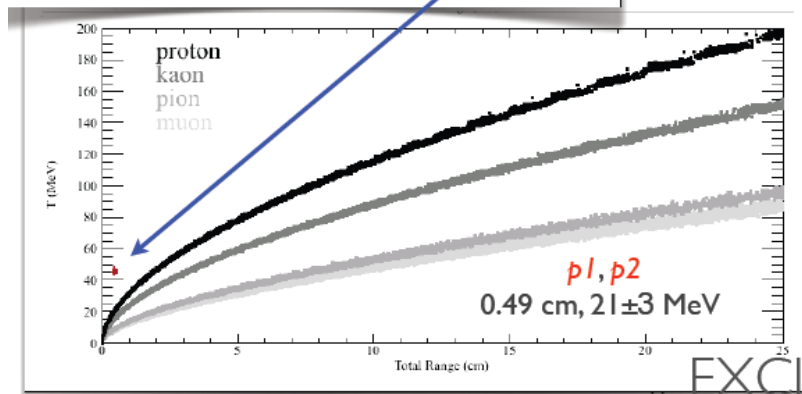
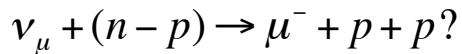
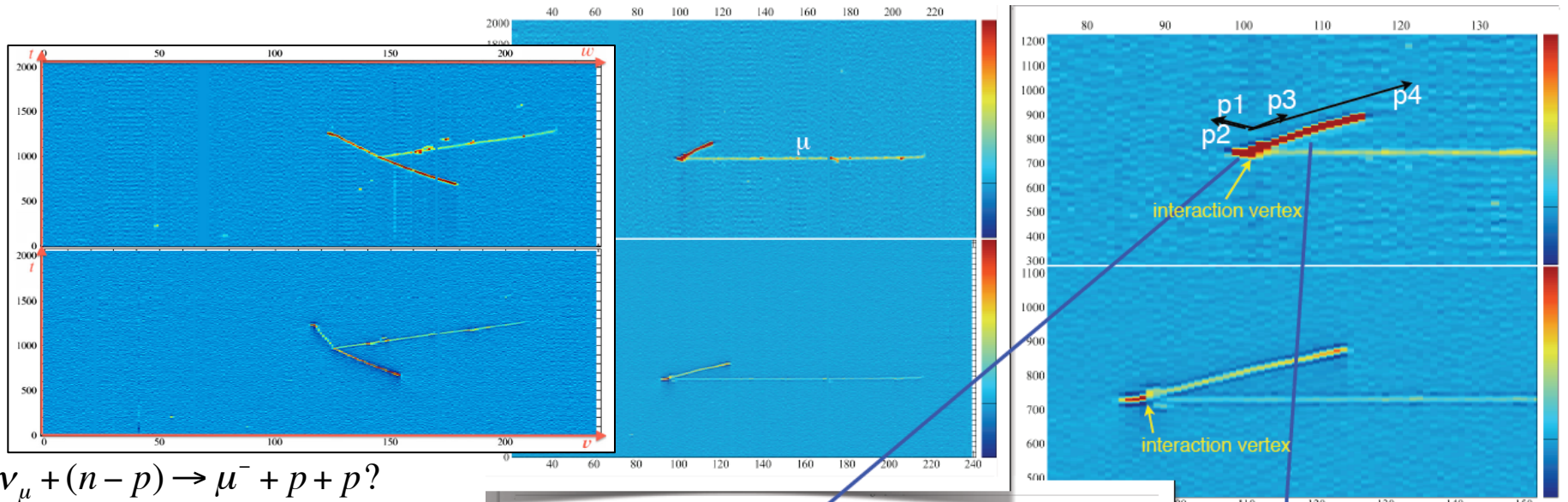


3. Neutrino oscillation experiment

LBNE (Liquid Argon TPC)

- They try to measure all outgoing hadrons, on top of the leading lepton
- Neutrons are worry (we are 4 GeV, energy carried away by neutrons are significant)
- Study is on-going to understand outgoing nucleons (exclusive topological cross section)

$$E_\nu = E_{lepton} + \sum E_{hadron}$$



p4: 12 cm, 128±7 MeV
 p3: 0.6 cm, 24±3 MeV

compatible with
 1 μ 4p

DATA:

EXCLUSIVE TOPOLOGIES

4. Neutrino-induced 2 nucleon emission

Charged-current NNSRC $\nu_{\mu} + (n - p) \rightarrow \mu^{-} + p + p$

- Naively we expect 2 outgoing protons (easier than (e,e') experiment?!)
- General structure is known from (e,e') experiments

Theory overview on neutrino-nucleon (-nucleus) scattering

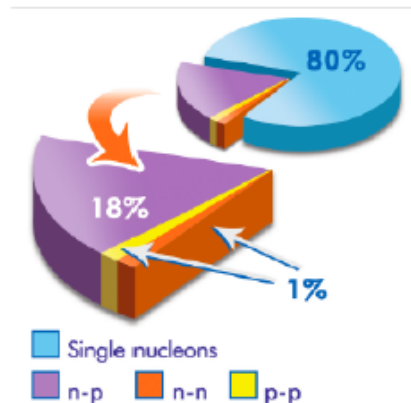
└ Theoretical models

└ Correlations

Nucleon correlations

^{12}C From (e,e'), (e,e'p), and (e,e'pN) Results

- 80 +/- 5% single particles moving in an average potential
 - 60 - 70% independent single particle in a shell model potential
 - 10 - 20% shell model long range correlations
- 20 +/- 5% two-nucleon short-range correlations
 - 18% np pairs (quasi-deuteron)
 - 1% pp pairs
 - 1% nn pairs (from isospin symmetry)
- Less than 1% multi-nucleon correlations



INT Workshop 4 December 2013

Jefferson Lab

from Higinbotham



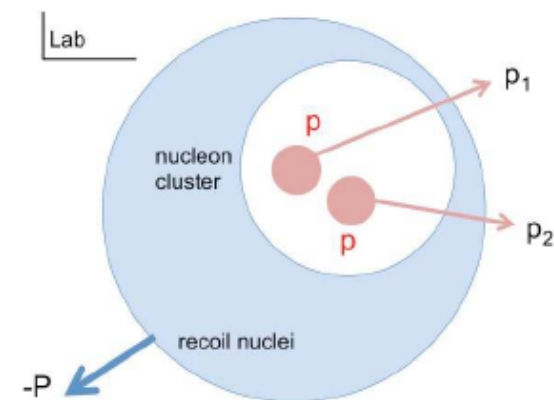
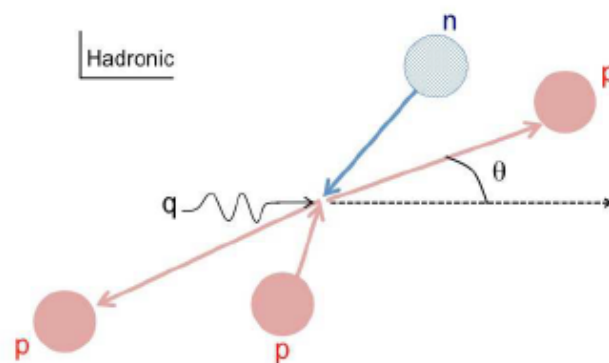
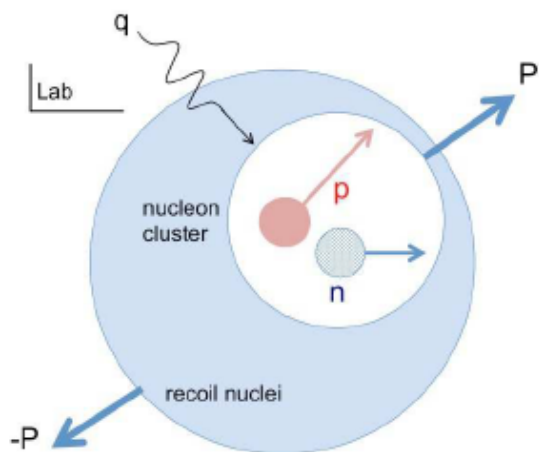
4. Neutrino-induced 2 nucleon emission

Charged-current NNSRC $\nu_{\mu} + (n - p) \rightarrow \mu^{-} + p + p$

- Naively we expect 2 outgoing protons (easier than (e,e') experiment?!)
- General structure is known from (e,e') experiments

Nucleon cluster model

- Based on these information, we simulate 2-nucleon emission in our MC
- energy-momentum vector and 2 nucleons make center mass system (hadronic system)
- isotropic decay is boosted back, to simulate 2 outgoing nucleons
- here, (n-p) pair is maybe 80% or so, but higher than (n-n) or (p-p) pairs.



4. Neutrino-induced 2 nucleon emission

Charged-current NNSRC $\nu_{\mu} + (n - p) \rightarrow \mu^{-} + p + p$

- Naively we expect 2 outgoing protons (easier than (e,e') experiment?!)
- General structure is known from (e,e') experiments

Nucleon cluster model

- Based on these information, we simulate 2-nucleon emission in our MC
- energy-momentum vector and 2 nucleons make center mass system (hadronic system)
- isotropic decay is boosted back, to simulate 2 outgoing nucleons
- here, (n-p) pair is maybe 80% or so, but higher than (n-n) or (p-p) pairs.

What neutrino interaction community want to know

→ Any knowledge to improve this picture

- 1. Introduction
- 2. 2 body current
- 3. Neutrino oscillation
- 4. Neutrino interaction
- 5. Conclusion

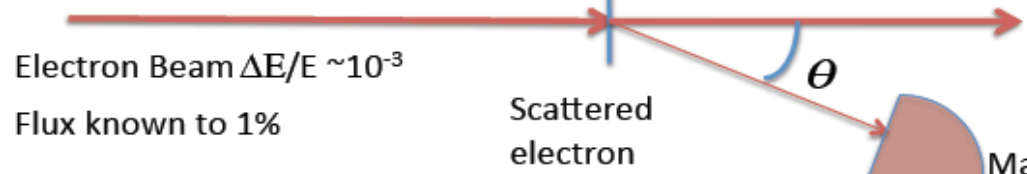
4. Neutrino interaction measurement

In neutrino physics, we don't know the energy of incoming neutrino

- We don't have ω and $|q|$ (and initial nucleon momentum, light-cone fraction, etc)
- We need a model works fine in all kinematic space

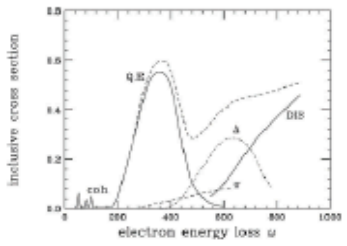
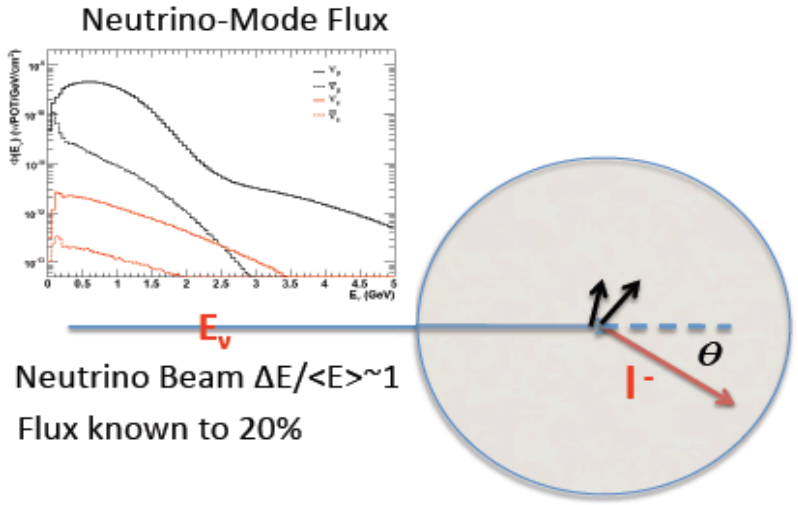
Contrast of e-N with ν -N Experiments

Electron



"description of neutrino data will require a new paradigm, suitable for application to processes in which the lepton kinematics is not fully determined" - Benhar

Neutrino



Don't know E_ν !!!
What's ω ???
What's \vec{q} ????
QE peak ????

Very Different Situation from inclusive electron scattering!!

4. Neutrino interaction measurement

In neutrino physics, we don't know the energy of incoming neutrino

- We don't have ω and $|q|$ (and initial nucleon momentum, light-cone fraction, etc)
- We need a model works fine in all kinematic space

This fact tricked many successful nuclear models (all impulse approximation based models)

Spectral function

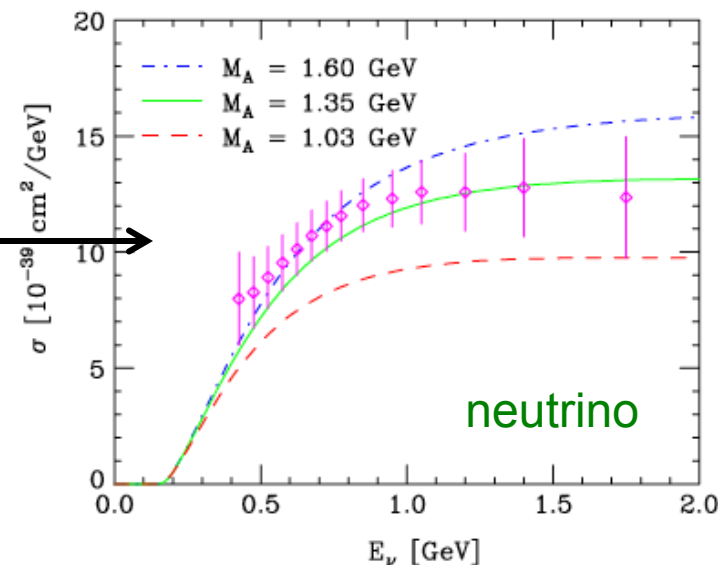
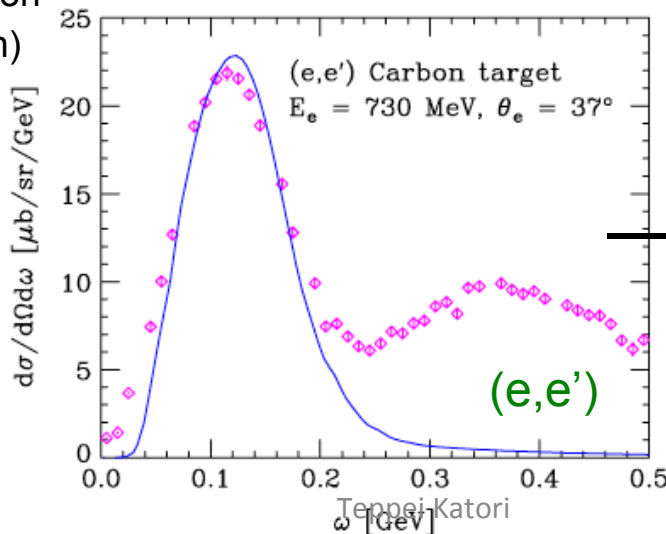
- It cannot describe low q region and dip region.
- (new model takes into account low q region)

Superscaling

- It cannot describe low q region.
 - $f_L = f_T$ is a bad assumption
- (new model include 2p2h)

DWIA, etc

Yo! M_A (axial mass) is 1.3 (GeV/c), da?



5. Conclusion, what we want to know...

Isospin

- Is 20% (n-p) pair reasonable? Is there any energy dependence?
(ω -|q| dependence is not measurable, but we can put in simulation, if we know that is right)

Momentum sharing

- Is it isotropic? can sometimes one nucleon takes more energy-momentum? if so, what kind of distribution is that?
- Are initial momentums back-to-back? can sometimes deviate? if so, what kind of distribution is that?
- Is there any energy dependence in this picture?

Separation energy

- Is there any “cost” to liberate (n-p) pair? are both nucleons on-shell?

Nuclear explosion

- In fact, LArTPC can measure de-excitation gamma rays (<1 MeV) from the nuclear remnant. This is a great advantage of vertex detector (=neutrino experiment) than arm spectrometer (electron experiment). Any there any predictions for that? 2 nucleon emissions should leave nuclei more unstable...

5. Conclusion, nuclear model we want

Interaction type

- It should work on all interactions, QE, resonance, transition, to DIS
- T2K, MicroBooNE=QE dominant, NOvA, LBNE=resonance dominant, PINGU=DIS dominant

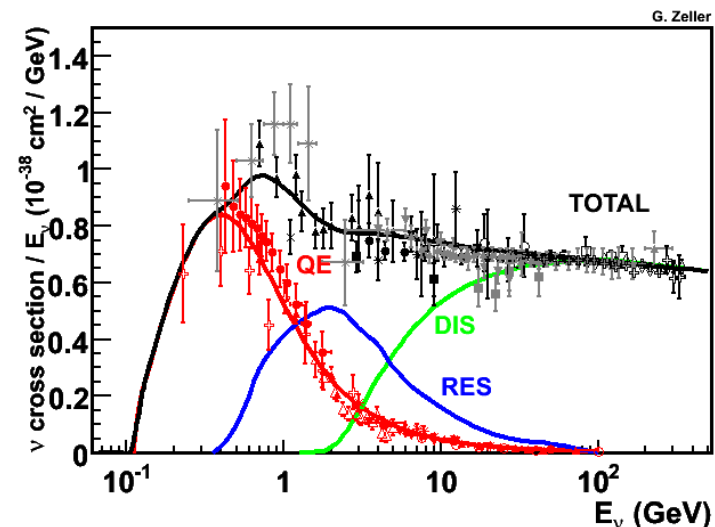
Kinematic region

- It should work in all kinematic space, low $\omega-|q|$ to high $\omega-|q|$.
- Too low ω (most of models break down) and too high $|q|$ (many RPA approaches fails) can be ignored?

Target material

- It should work on heavy element, Fiducial volume is made by carbon, water, argon, etc, but the surrounding material is made of heavier elements (and they produce more interactions)

Thank you for your attention!



5. NuSTEC collaboration

NuSTEC School at Fermilab, Oct. 20-30, 2014

- NuSTEC is CTEQ-like collaboration to improve neutrino interaction models
- School is designed for an introduction to both nuclear and particle physics of neutrino-nucleus scattering
- Ideal for experimental and theoretical advanced graduate students and young postdocs

<http://nustec2014.phys.vt.edu/>

Lecturers

Luis Alvarez-Ruso (IFIC), Rocco Schiavilla (ODU), Bill Donnelly (MIT), Juan Nieves (IFIC), Omar Benhar (Roma), Toru Sato (Osaka), Pawel Danielewicz (MSU), Jeff Owens (FSU), Pilar Coloma (FERmilab), Tom Dealty (Oxford), Mitch Soderberg (Syracuse), Chris Mauger (Los Alamos)

Thank you for your attention!



Neutrino Cross-Section Newsletter

- Teppei Katori's one-person journal club
- Discuss the latest papers (both theory and experiment)
- news around the neutrino cross-section community
- 1 or 2 per month, depending on how many new papers

<https://pprc.qmul.ac.uk/~katori/nu-xsec.html>