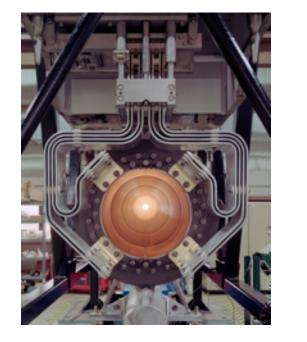
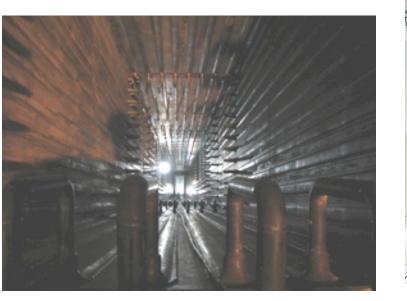
Results from Long-Baseline Neutrino Experiments









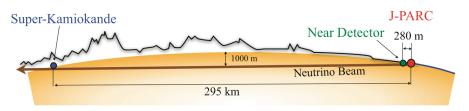
Alysia Marino, University of Colorado at Boulder PIC, Bloomington, Sept. 17, 2014

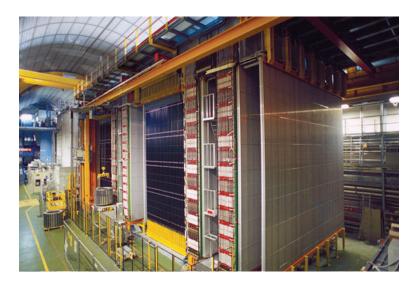
Outline

- Neutrino Mixing
- Recent Neutrino Oscillation Results
 - MINOS
 - ▶ T2K
 - ▶ OPERA
- Results expected in near future
 MINOS+
 - ► NOvA







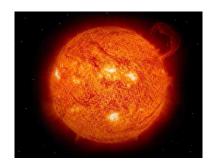


Neutrino Oscillations



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- Very **compelling evidence** that neutrinos change flavor
 - Fewer atmospheric muon neutrinos vanish while crossing Earth
 - Confirmed with accelerator neutrino beams.





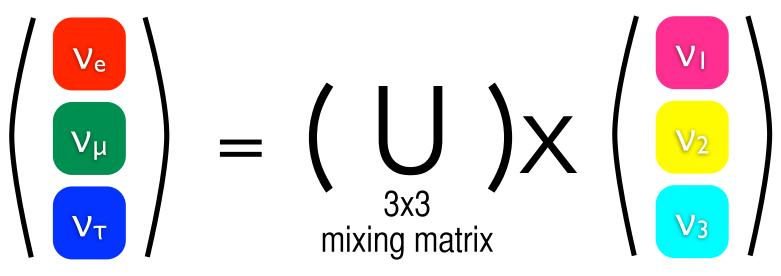
- Too few electron neutrinos from the solar core are observed. But the total flux of all 3 flavors matches the expected number.
- Confirmed with reactor neutrinos over 100km distances.
- Missing reactor neutrinos at shorter distances
- Electron neutrino appearance in a muon neutrino beam

Neutrino Mixing

• Neutrino flavor states are a mixture of neutrino mass states.

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 Interference between the mass and flavor eigenstates causes the observed flavor to oscillate with time.



• In general neutrino transition probability to another flavor is given by: $P(\nu_{\alpha} \rightarrow \nu_{\beta}) = \delta_{\alpha\beta} - 4 \sum_{i>j} \Re(U^*_{\alpha i} U_{\beta i} U_{\alpha j} U^*_{\beta j}) \sin^2 [\Delta m^2_{ij} L/4E]$ $-2 \sum_{i>j} \Im(U^*_{\alpha i} U_{\beta i} U_{\alpha j} U^*_{\beta j}) \sin [\Delta m^2_{ij} L/2E]$ Matrix Elements $\Delta m^2_{ij} = m^2_{j} - m^2_{j}$

Mixing Matrix

- Unitary matrix relates mixing between flavor and mass states
 - θ_{23} and Δm^2_{32} Atmospheric/ Accelerator neutrinos $\theta_{23} \sim 45^{\circ}$

δ, $θ_{13}$ and $Δm^2_{31}$ reactor anti-neutrinos and accelerator neutrinos $θ_{13} \sim 9^{\circ}$ θ_{12} and Δm_{21}^2 Solar neutrinos/ reactor anti-neutrinos $\theta_{12} \sim 34^\circ$

$$U = \begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \times \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{bmatrix} \times \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$

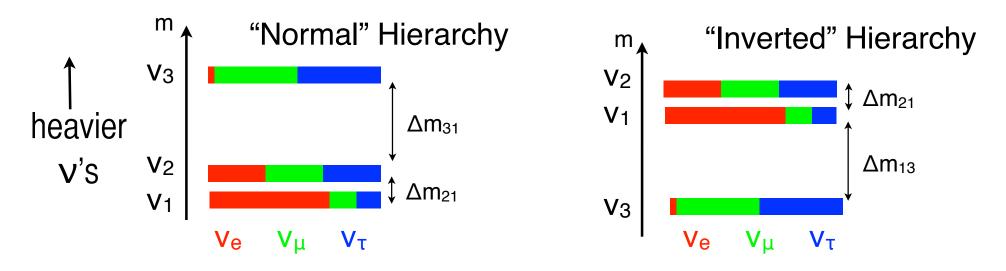
where $s_{ij} = sin(\theta_{ij})$ and $c_{ij} = cos(\theta_{ij})$

• Anti- ν depend on U^{*}

Neutrino Masses

Two different mass difference scales

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• Sign of Δm_{21} is known from solar ν from mass effects, but sign of Δm_{31} isn't, so two possible orderings of masses

Vacuum Oscillation Probability

- L/E scale relevant for recent accelerator beams oscillation effects are dominated by m₃↔m₂ and m₃↔m₁ mixing
- v_{μ} Disappearance in a v_{μ} Beam (no matter, no CP)

 $P(\nu_{\mu} \rightarrow \nu_{\mu}) \simeq 1 - \sin^2 2\theta_{23} \cdot \sin^2 \left(\Delta m_{32}^2 L/4E\right)$

• v_e Appearance in a v_{μ} Beam (no matter, no CP) $P(\nu_{\mu} \rightarrow \nu_e) \simeq \sin^2 2\theta_{13} \cdot \sin^2 \theta_{23} \cdot \sin^2 (\Delta m_{31}^2 L/4E)$

• Can also look for appearance of τ neutrinos

Increasing need for simultaneous 3 flavor fits

Matter & CP Effects

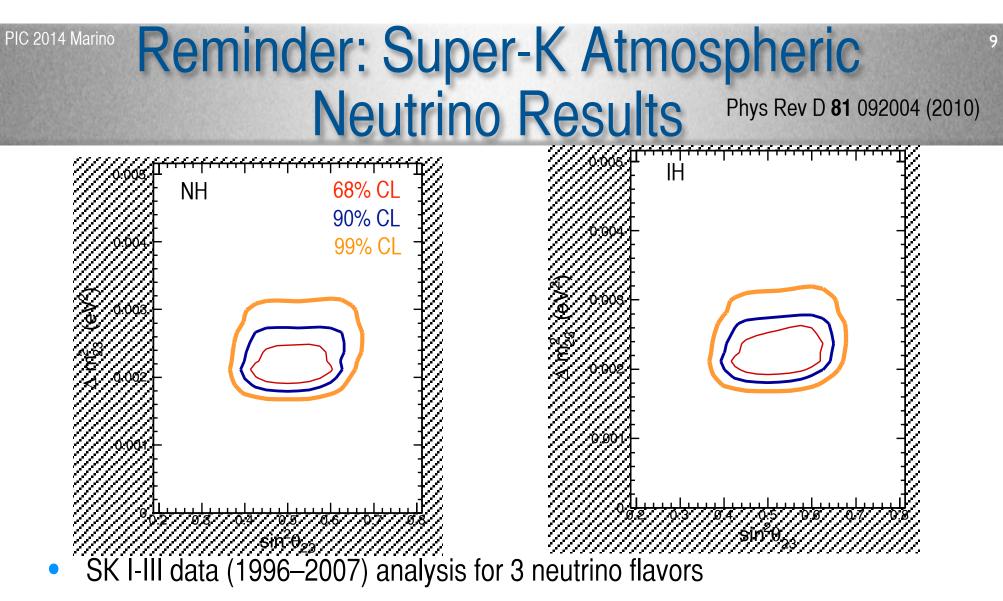
• In the presence of **matter**, the appearance probability changes to

$$P(\nu_{\mu} \to \nu_{e}) \simeq \sin^{2} 2\theta_{13} \cdot \frac{\sin^{2} \theta_{23}}{(A-1)^{2}} \cdot \sin^{2} \left((A-1)\Delta m_{31}^{2}L/4E\right)$$

where $A = \sqrt{2}G_{F}N_{\Delta m_{31}^{2}}$
Depends on hierarchy

- A has opposite sign for anti- ν . So even with $\delta=0$, $P(\nu_{\mu} \rightarrow \nu_{e})$ increases while $P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e})$ decreases for NH, so an apparent CP violation.
- For $\delta \neq 0$, there are additional terms including $-\frac{|\Delta m_{21}^2|}{|\Delta m_{31}^2|} \sin \delta \sin 2\theta_{12} \sin 2\theta_{13} \sin 2\theta_{23} \cos \theta_{13}}{A(1-A)} \sin \Delta \sin (A\Delta) \sin ((1-A)\Delta)$ where $\Delta = \Delta m_{31}^2 L/4E$

• Changes sign for anti- ν . So positive δ will decrease $P(\nu_{\mu} \rightarrow \nu_{e})$ and increase $P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_{e})$.



- No hierarchy preference; Results consistent with both $\theta_{13}=0$ and non-zero values
- Preliminary results at Neutrino 2014 for SK I-IV (1996-present, 4581 live days) which have a slight preference for NH and weak preference for $\theta_{13\neq}$ 0.

MINOS: Main Injector Neutrino Oscillation Search

Soudan

MN

Duluth "

IA

MO

Lake

uperio

WI

Fermilab

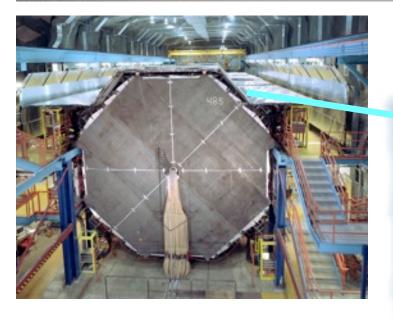
IL

Madison

Lake Michigan

IN

MI



Far Detector:

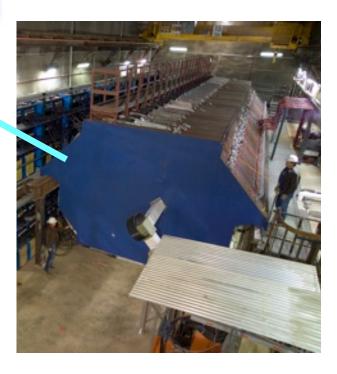
- 735 km from target
- 705 m underground in Soudan, MN
- 5.4 kton mass
- 8 m tall

Magnetized iron/ scintillator sampling calorimeters

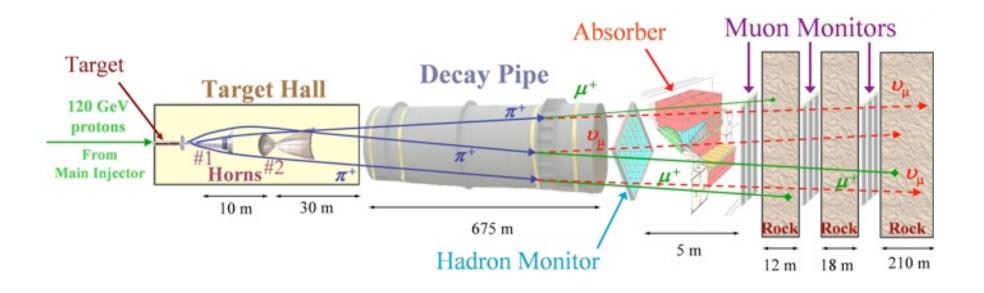
Near Detector:

 $\mathbf{0}$

- 103 m underground at FNAL
- 1 km from target
- 1 kton mass



NuMI Beam

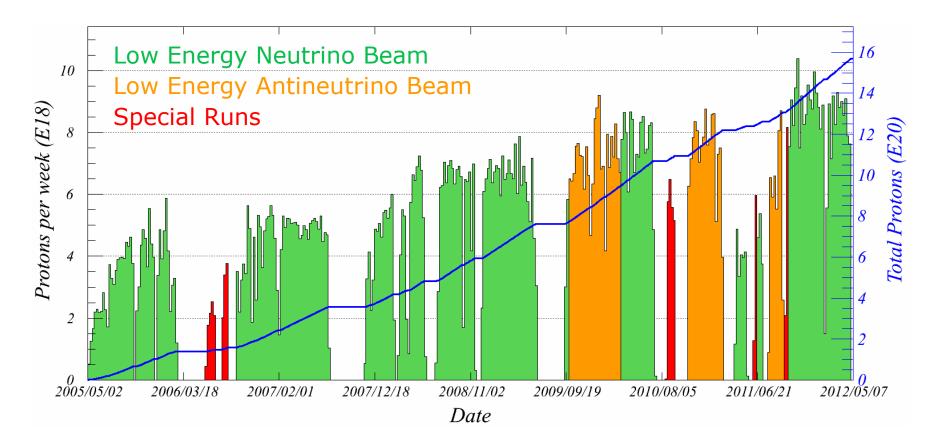


- 120 GeV protons from the Main injector strike a graphite target
- 2 horns focus the pions and kaons, which decay in a 675 m long decay pipe
- 93% ν_μ, 6% ν
 _μ, 1% ν_e
- 400 kW design power

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MINOS Running



- Neutrino Mode: 10.71x10²⁰ protons on target
- Anti-neutrino Mode: 3.61x10²⁰ protons on target
- 37.88 kton years of atmospheric neutrinos

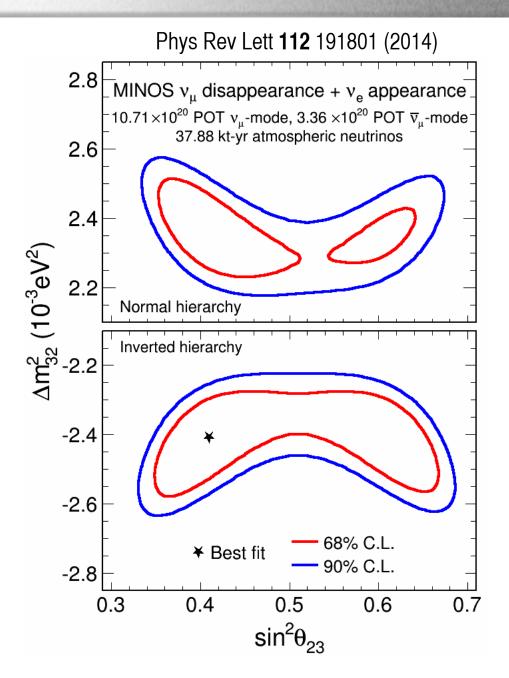
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Beam + Atmospheric Neutrinos

• Combined fit for v_{μ} disappearance and v_{μ} -> v_{e} appearances using NuMI beam data and atmospheric neutrinos

- Includes anti-neutrino data
- 3 flavor fit where Δm²₃₂, θ₂₃,
 θ₁₃, δ_{CP} are varied
- θ₁₃ constrained by reactor data

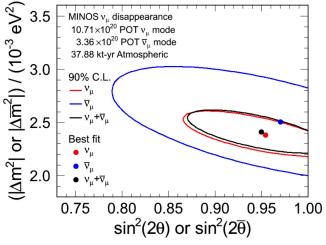


Anti-Neutrino Resulto

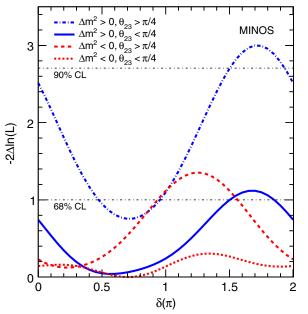
- v_{μ} disappearance for v and anti-v beam + atmospheric neutrinos
 - ▶ anti-v and v have similar best fit points and strongly overlapping 90% CL regions

- v_e appearance for v and anti-v beam
 - Data prefer inverted hierarchy, especially for $\theta_{23} > \pi/4$
 - At 68% CL disfavors 31% of the δ/octant/ hierarchy phase space, especially normal hierarchy with θ₂₃ >π/4 and large δ

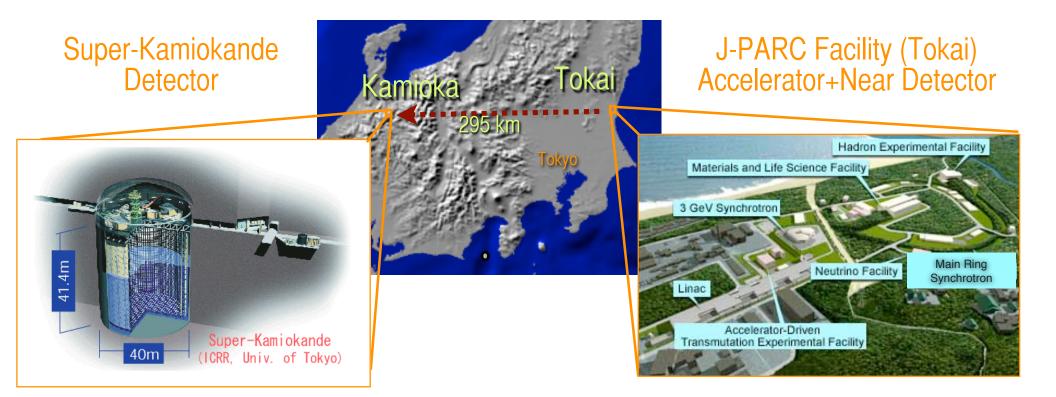
Rev Lett 110 251801 (2013)



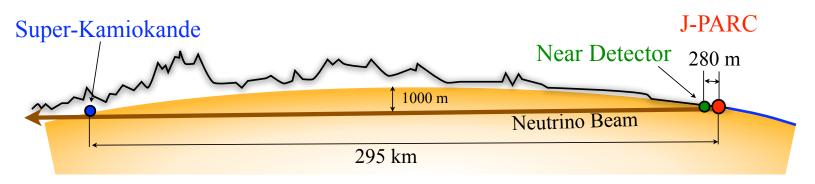




T2K: Tokai-to-Kamioka



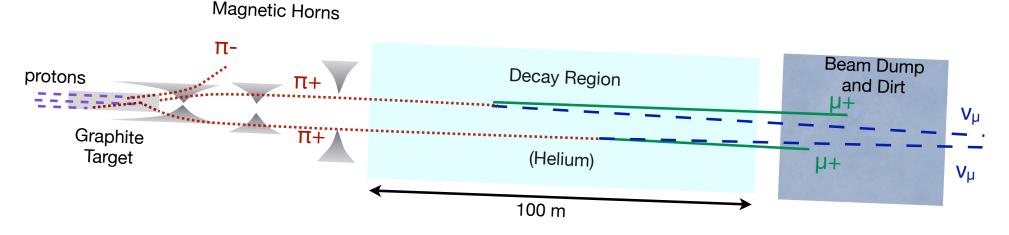
Long-baseline neutrino experiment in Japan with 295 km baseline



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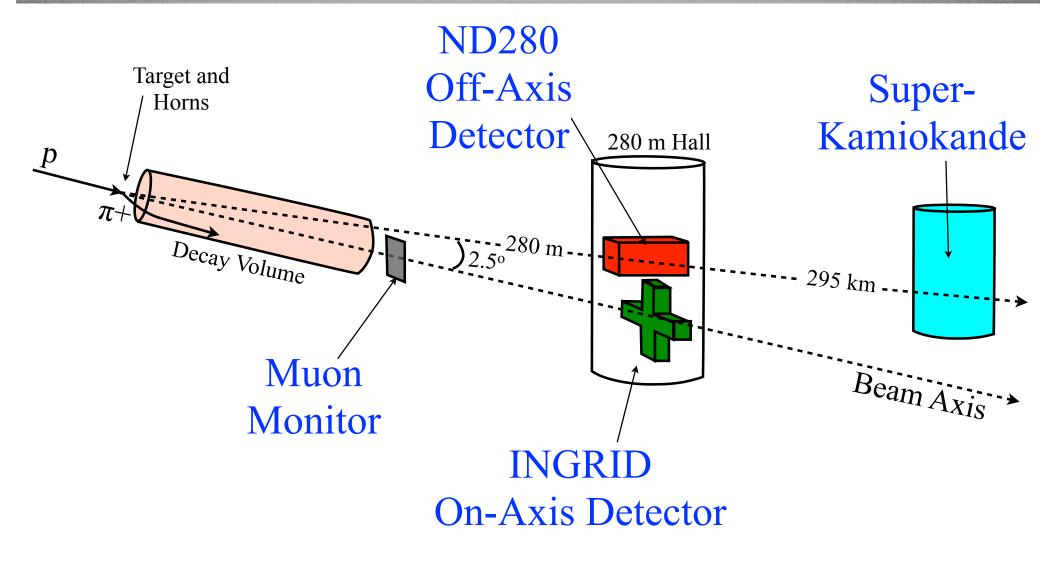


• 30 GeV protons strike graphite target, producing πs and Ks

- 3 magnetic horns to focus π and K into the desired direction
- Beam is ~95% $\nu_{\mu},~4\%~\overline{\nu_{\mu}},~1\%~\nu_{e}$
- Designed for 750 kW



T2K Detectors

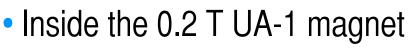


Not to scale!

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ND280: Off-Axis Near Detector

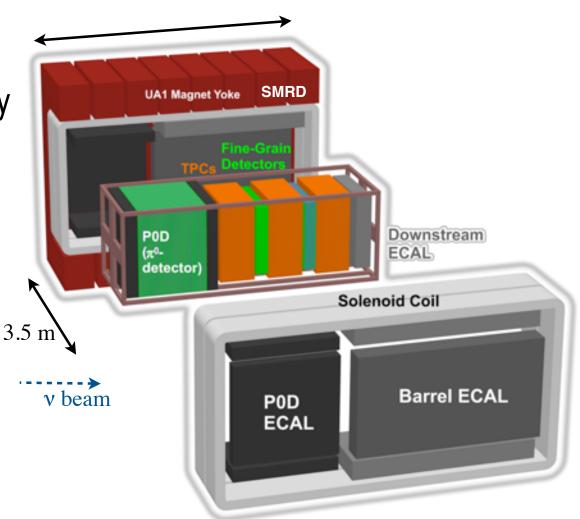
7m



- ~ 10,000 ν interactions per day
- Tracker
 - Fine-grained scintillator tracker + 3 TPCs (and water targets)
- Pi-Zero Detector
 - Brass/Lead/plastic scintillator calorimeter (and water targets)
- ECAL
 - Lead/plastic scintillator calorimeter

Side Muon Range Detector

 Plastic scintillator paddles interspersed between magnet yoke layers



PIC 2014 Marino Super-Kamiokande: Far Detector

- Large volume water Cherenkov detector
- 50 kton of pure H₂O (22.5 kton fiducial volume)

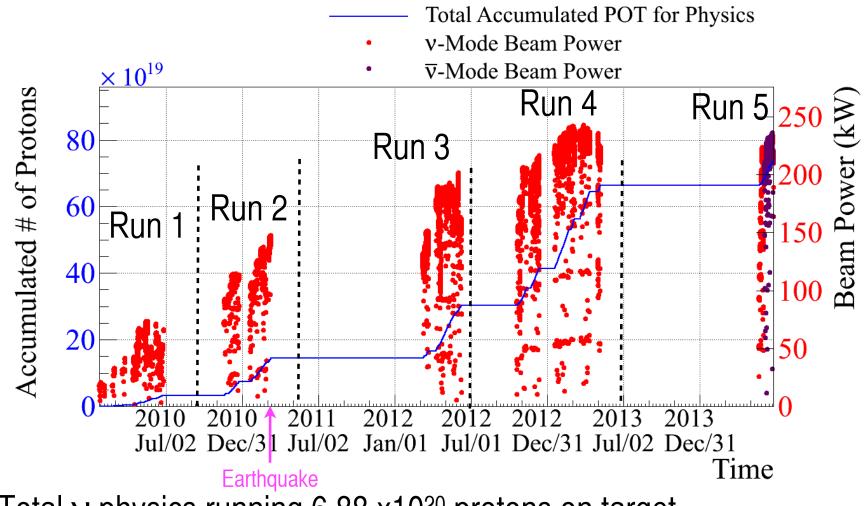
40m

11,000 phototubes

41.4m

- Outer layer with 1885 phototubes to reject external events
- Expect a handful of events per day at full beam power.

T2K Beam Running



- Total ν physics running 6.88 x10²⁰ protons on target
- Total anti- ν physics running 0.51 x10²⁰ protons on target
- Only 10% of expected protons

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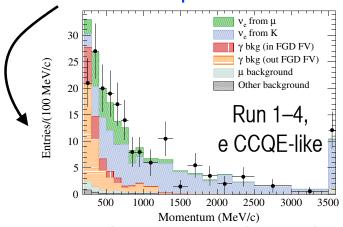
Measurements Using Only Near Detectors

Papers from On-Axis Near Detector

• Inclusive v_{μ} CC cross section on Fe and C, arXiv:1407.4256, accepted by Phys Rev D

Papers from Off-Axis Near Detector Data

- Inclusive ν_{μ} CC cross section on C, Phys Rev D 87 092003 (2013)
- Inclusive v_e CC cross section on C, arXiv: 407.7389
- ► Measurement of ve component of the Beam, Phys Rev D 89 092003 (2014)

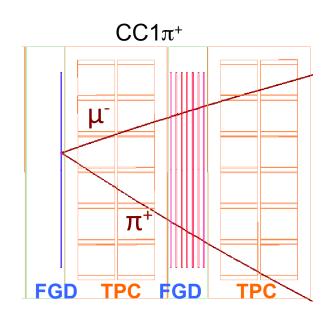


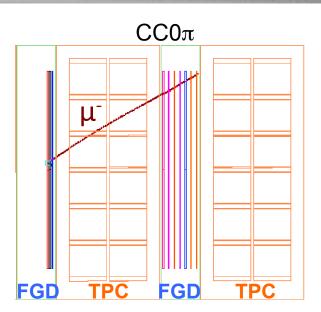
Measured/Predicted $= 1.01\pm0.1$

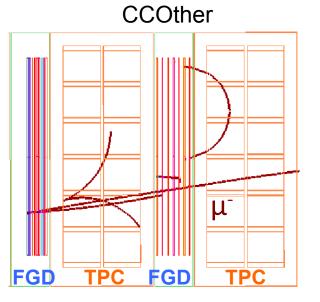
• Preliminary cross sections results have been also presented on a number of topics including v_{μ} CC quasi-elastic cross section on C, NC elastic scattering cross section, and CC coherent pion production

Near Detector Samples

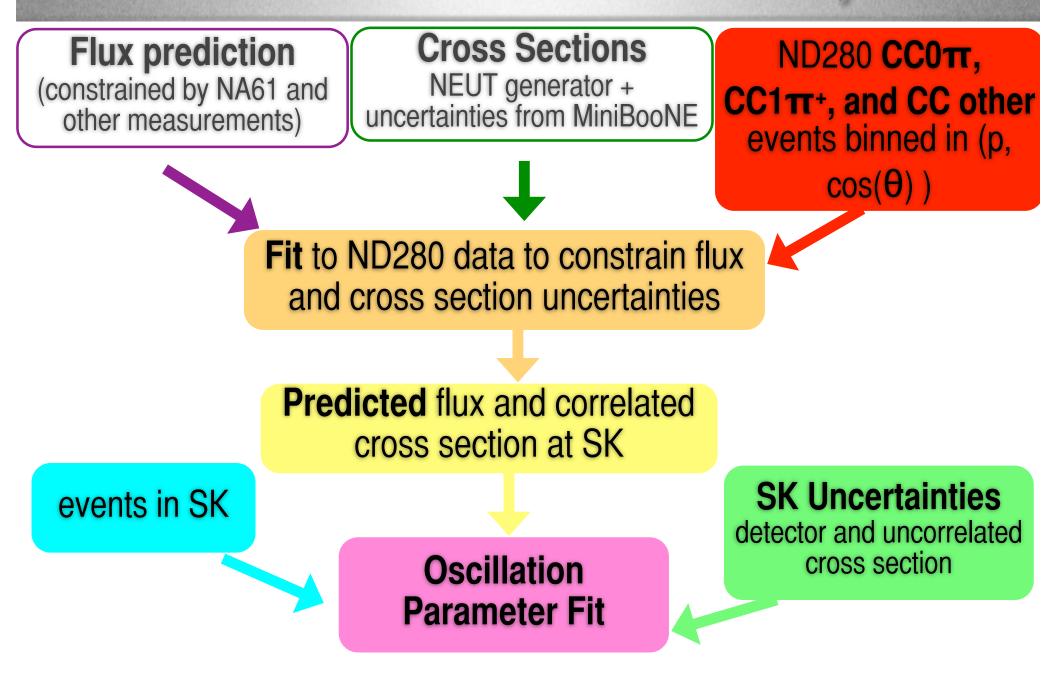
- 3 ND samples used as constraints for oscillation analyses
 - CC0π
 - CC1π+
 - CC Other



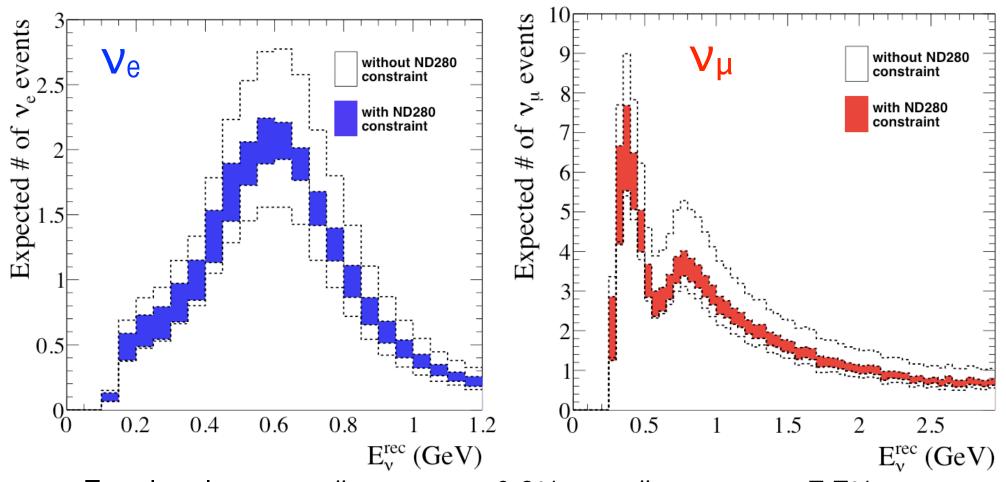




Constraints for Oscillation Analysis



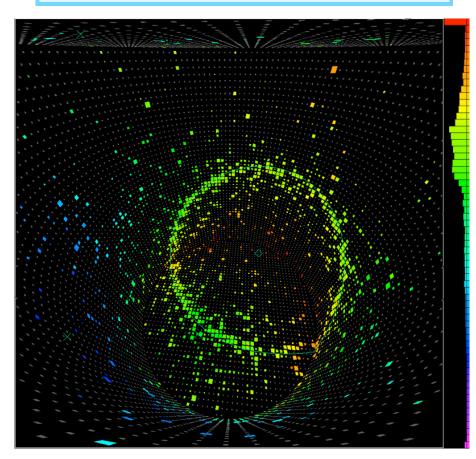
Constraints and Uncertainties



- Fractional error on # v_e events: 6.8%; on # v_μ events: 7.7%
- After constraint, largest sys uncertainty due to difference in nuclear targets

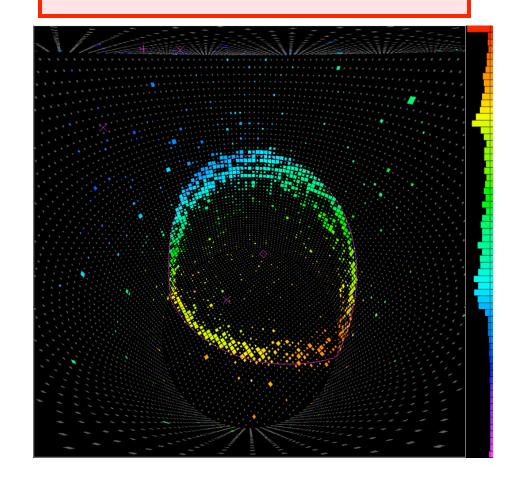
Event Selection in Super-K

Ve sample Fully-contained event with single electron-like ring, E<1250 MeV



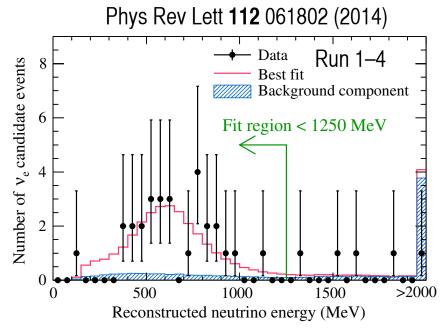
V_{μ} sample

Fully-contained event with single muon-like ring



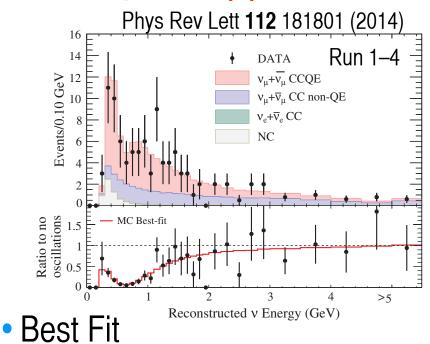
ν_e Appearance and ν_μ Disappearance

ν_e Appearance



- •Best Fits at $\delta_{cp}=0$, $\Delta m^2_{32}=2.51 \times 10^{-3}$ eV², sin² $\theta_{23}=0.5$
 - ► NH: sin²2θ₁₃=0.140±^{0.038}0.032
 - ► IH: sin²2θ₁₃=0.170±^{0.045}0.037
- •7.3 σ significance compared to bkg

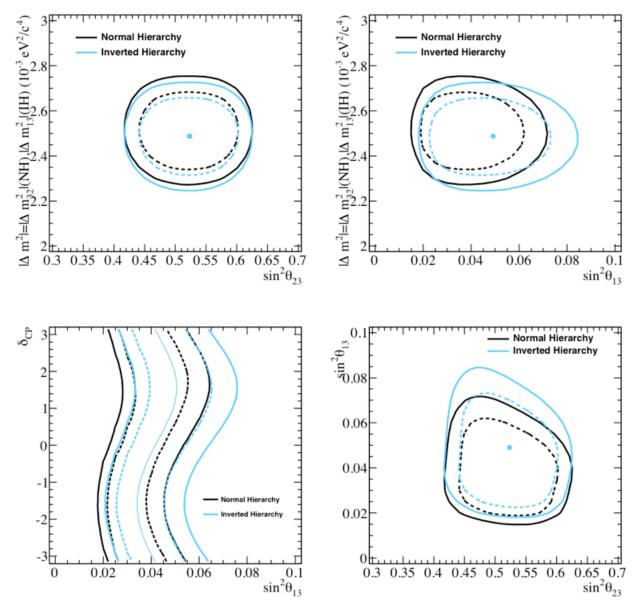
ν_{μ} Disappearance



- ► NH: $\Delta m_{32}^2 = (2.51 \pm 0.10) \times 10^{-3} \text{ eV}^2$, sin² $\theta_{23} = 0.514 \pm 0.055_{0.056}$
- IH: Δm²₃₂=(2.48±0.10)x10⁻³ ev², sin²θ₂₃=0.511±0.055
- Most precise measurement of θ_{23}

T2K Combined 3-flavor Fit

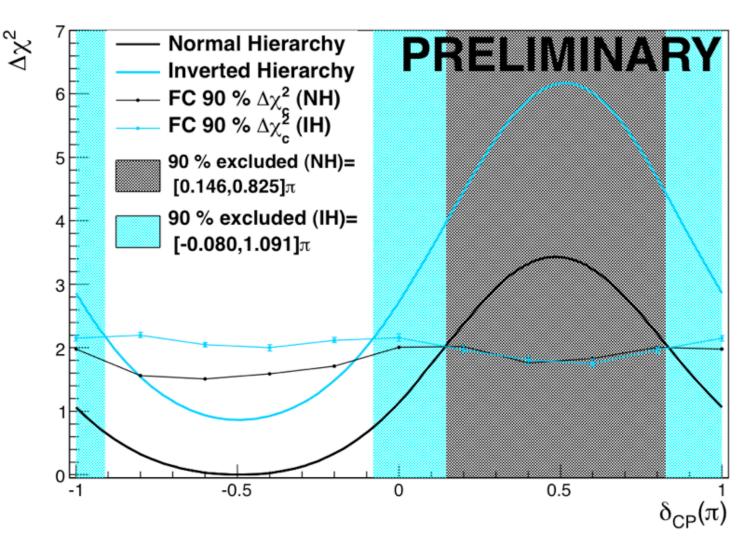
T2K Preliminary



• Combined 3-flavor fit for T2K ν_{μ} disappearance and ν_{e} appearance data

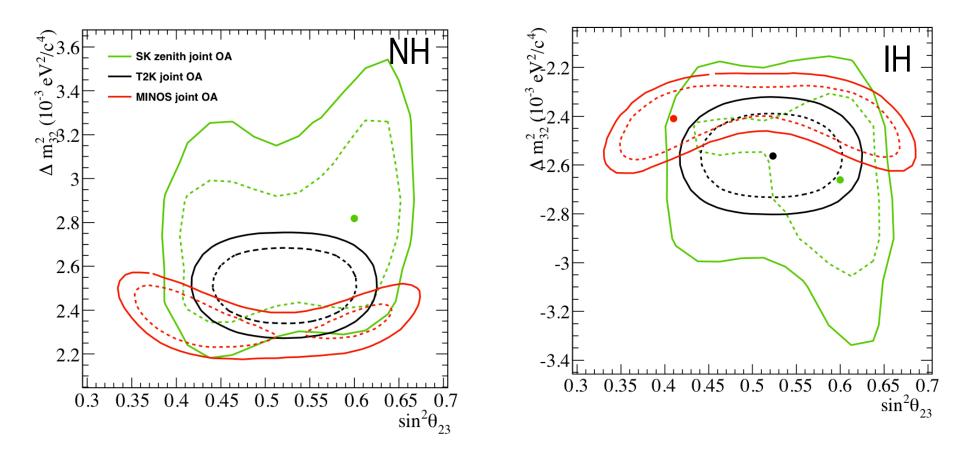
Hierarchy and δ_{CP}

- Includes reactor constraint on θ_{13}
- Preference for negative δ

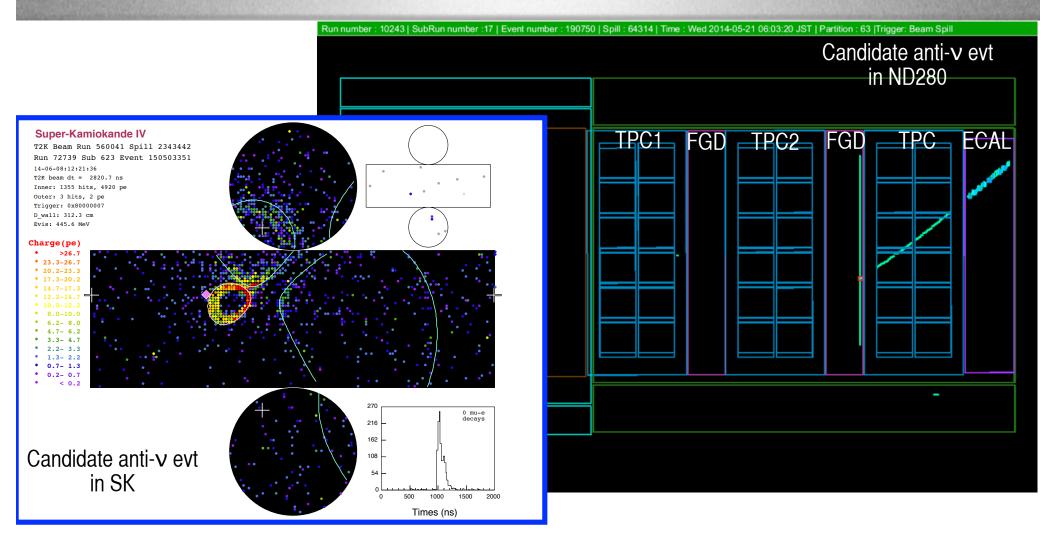


Comparison of Masses and Angles

- Simultaneous appearance/disappearance 3-flavor fits for each experiment
- Allowed regions for mixing angles and mass splitting overlap very well



Anti-Neutrinos in T2K

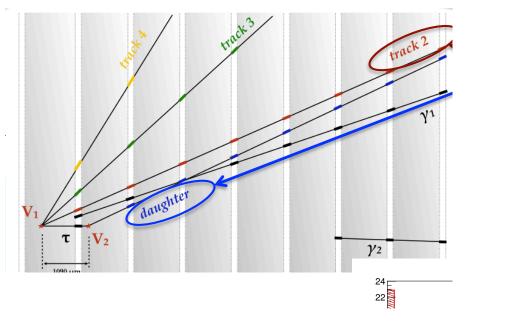


- In June 2014 began data taking with anti-ν beam
- In future will take similar amounts of anti- ν and ν data

Tau Appearance

- OPERA 2008–2012 data in Phys Rev D 89 051102 (2014)
- A 4th τ appearance event discovered in more recent data, no oscillation now excluded at the 4.2 σ CL

4th candidate, S. Dusini, Neutrino 2014



• Super-K in Phys Rev Lett **110** 181802 (2013) also reported 3.8 σ (evidence for tau appearance in events resulting from hadronic tau decays. $D = \frac{L}{R_{lead}(p)} \frac{\rho_{average}}{\rho_{lead}} = 0.40^{+0.04}_{-0.05}$

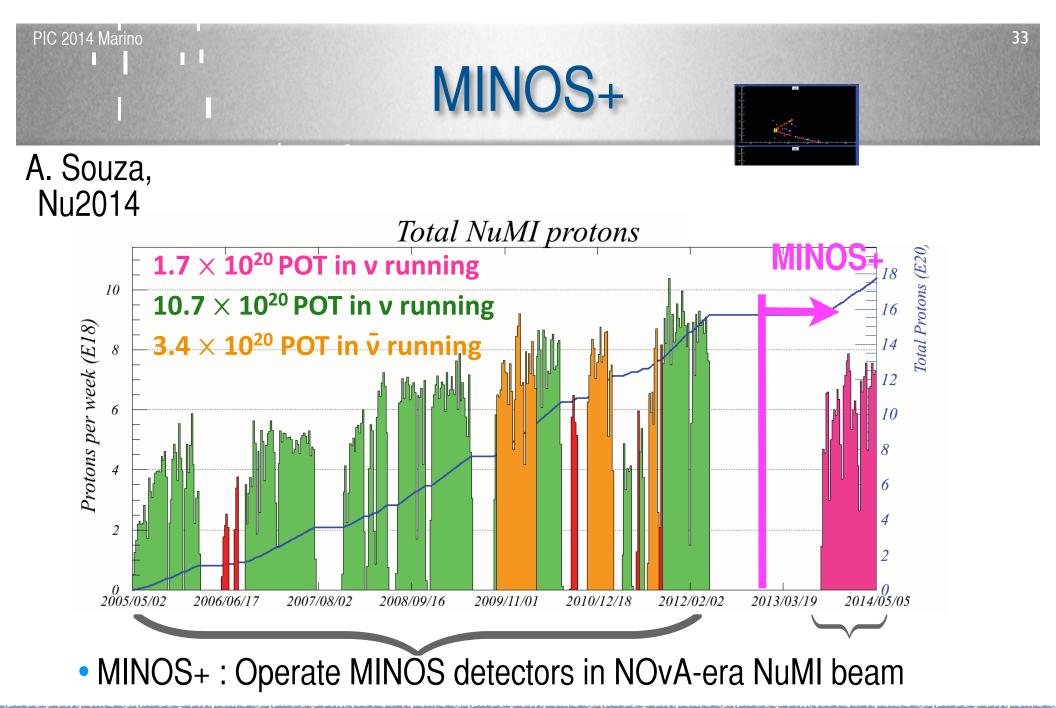
 $\operatorname{Pt}_{Y}\left(\operatorname{GeV}(c)\right)$

-0.2

-0.4

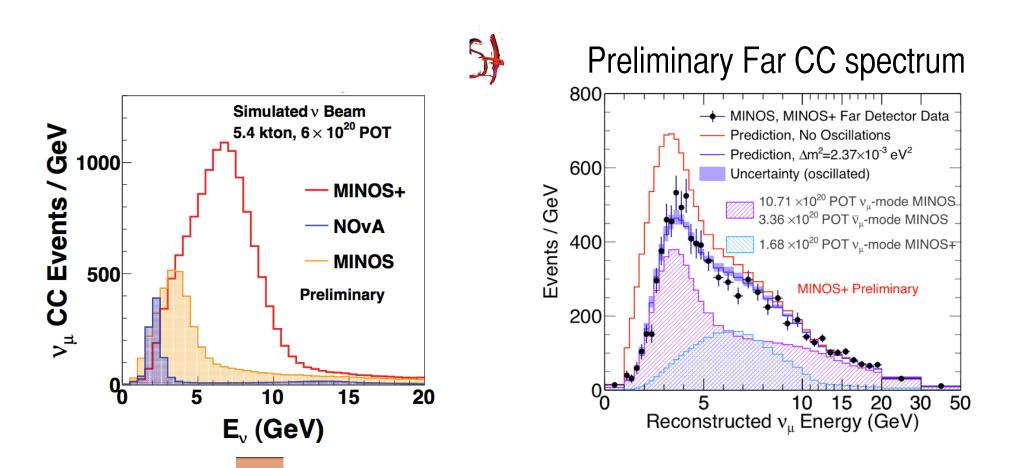
-0.4

Near Future Accelerator-Based Neutrino Experiments



MINOS+ Physics

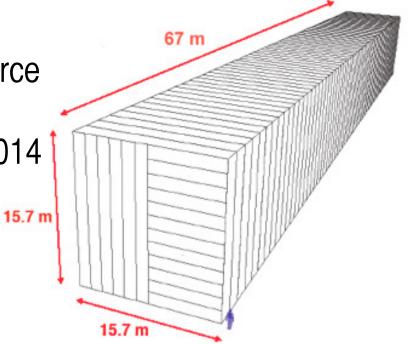
- Beam is higher in energy and has a greater flux per POT
- Goals include precision 3-flavor mixing tests, sterile neutrinos, non-standard interactions, neutrino cross section measurements



ΝΟνΑ

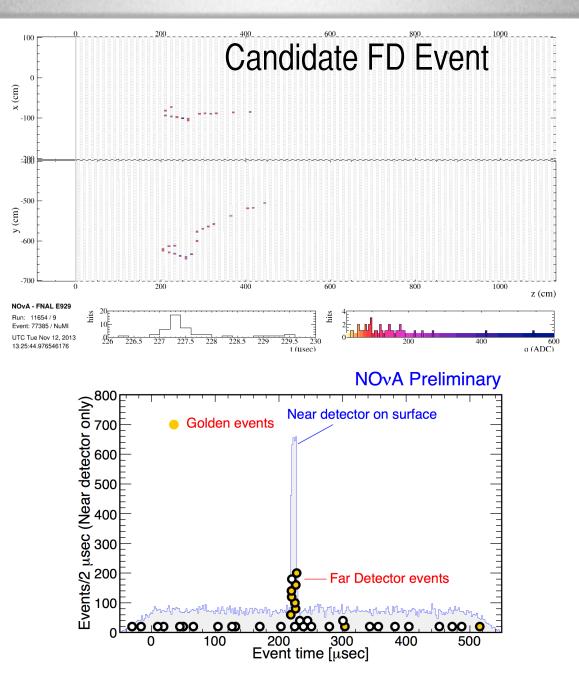
- active scintillator detectors, 14 kton far det
- Off-axis from NuMI beam, 810 km from source
- First beam April 2013
- Near and Far Detectors complete in April 2014







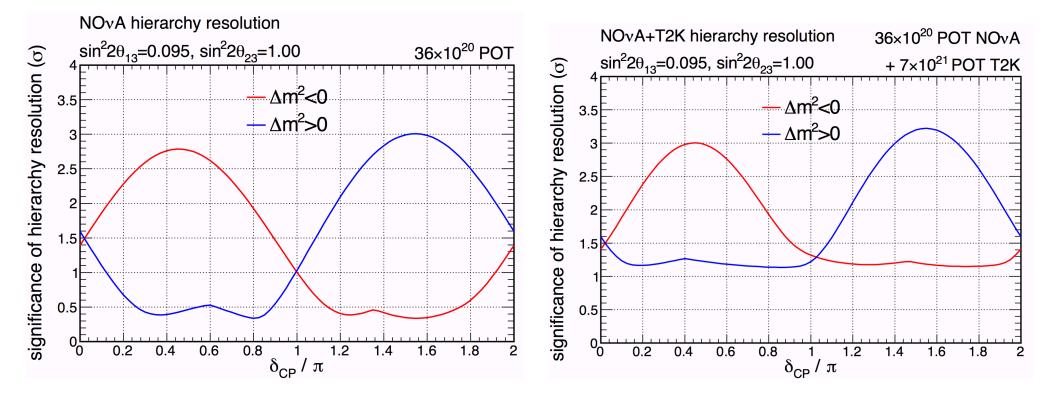
NOvA Candidate Events



Hierarchy Sensitivity

NOvA

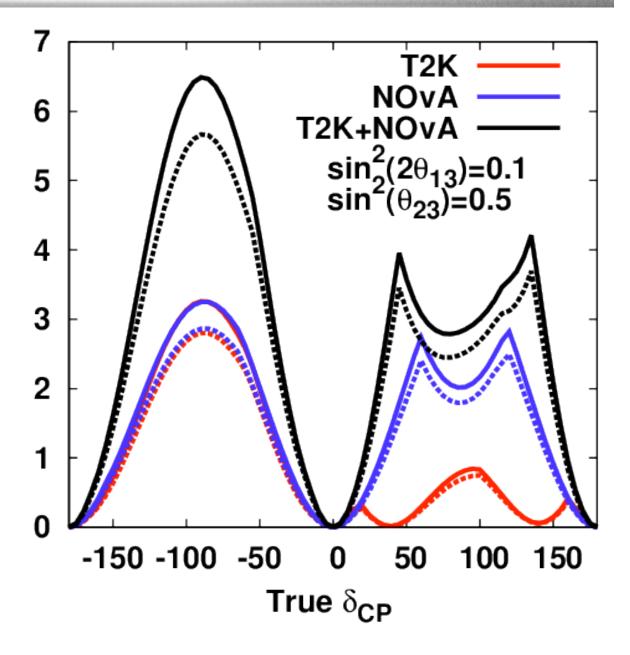
NOvA+T2K



- NOvA has 2σ resolution of hierarchy over 1/3rd of δ space
- NOvA+T2K have better than 1σ resolution over all space

δ_{CP} Sensitivity

- Assumes hierarchy is unknown
- Dashed lines include systematics
- T2K+NOvA have more than 1σ sensitivity over 75% of the δ_{CP} range



Summary

- We are moving into an era of precision measurements of neutrino oscillation parameters in the 1–3 and 2–3 sector with acceleratorbased neutrino experiments.
- MINOS and T2K have observed the disappearance of ν_{μ} and the appearance of ν_{e}
- MINOS operations have concluded, joint 3-flavor analysis performed of beam (v+anti-v) + cosmics
- T2K began taking anti- ν data in 2014
- MINOS+ and NOvA have started data collection and will present physics result in the near future