# Off-axis beam dynamics and optimisation strategies 

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## Overview

- Off-axis beam dynamics
- Combined function magnets
- Off-axis beams through linear optics
- Nonlinear extension
- Optimisations strategies
- Local vs. global optimisation
- Examples for complex system


## Combined function magnets

Normal dipole magnet


Combined
function dipole


- Parallel faces of dipole yoke causes uniform magnetic field
- Only dipole component


## Combined function magnets



## Combined function magnets

Normal dipole magnet


- Parallel faces of dipole yoke causes uniform magnetic field
- Only dipole component

Combined function dipole


- Non-parallel faces changes the magnetic field
- Dipole component
- AND quadrupole component


## Hill's Equation

$x^{\prime \prime}=K x$
$K=\frac{1}{\rho^{2}}$ (Dipole)
$K=k_{1}$ (Quadrupole)
$K=\frac{1}{\rho^{2}}+k_{1}$ (Combined function magnet)
Transfer matrix:

$$
\begin{aligned}
& \left(\begin{array}{cccccc}
\cos \sqrt{K} s & \frac{1}{\sqrt{K}} \sin \sqrt{K} s & 0 & 0 & 0 & \frac{2}{\rho K} \sin ^{2} \frac{\sqrt{K} s}{2} \\
-\sqrt{K} \sin \sqrt{K} s & \cos \sqrt{K} s & 0 & 0 & 0 & \frac{1}{\rho \sqrt{K}} \sin \sqrt{K} s \\
0 & 0 & \cosh \sqrt{k_{1}} s & \frac{1}{\sqrt{k_{1}}} \sinh \sqrt{k_{1}} s & 0 & 0 \\
0 & 0 & \sqrt{k_{1}} \sinh \sqrt{k_{1}} s & \cosh \sqrt{k_{1}} s & 0 & 0 \\
0 & 0 & 0 & 1 & \alpha \\
-\frac{1}{\rho \sqrt{K}} \sin \sqrt{K} s & -\frac{2}{\rho K} \sin ^{2} \sqrt{K} s & 0 & 0 & 0 & 1
\end{array}\right) \\
& 0 \\
& 0
\end{aligned}
$$

## Simplified transfer matrix

- Only consider horizontal plane
- Vertical plane acts like a normal quadrupole
- Ignore longitudinal dynamics for this lecture!

$$
\begin{gathered}
\binom{x_{1}}{x_{1}^{\prime}}=\left(\begin{array}{cc}
\cos \sqrt{K} s & \frac{1}{\sqrt{K}} \sin \sqrt{K} s \\
-\sqrt{K} \sin \sqrt{K} s & \cos \sqrt{K} s
\end{array}\right)\binom{x_{0}}{x_{0}^{\prime}} \\
\left(\begin{array}{c}
D_{x, 1} \\
D_{x, 1}^{\prime} \\
1
\end{array}\right)=\left(\begin{array}{ccc}
\cos \sqrt{K} s & \frac{1}{\sqrt{K}} \sin \sqrt{K} s & \frac{2}{\rho K} \sin ^{2} \sqrt{K} s \\
\sqrt{K} \sin \sqrt{K} s & \cos \sqrt{K} s & \frac{1}{\rho \sqrt{K}} \sin \sqrt{K} s \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
D_{x, 0} \\
D_{x, 0}^{\prime} \\
1
\end{array}\right)
\end{gathered}
$$

The $3^{\text {rd }}$ row of the matrix is useful for matrix calculations

## Off-axis beams through linear optics

- A combined function magnet is NOT equivalent to travelling off-axis through a quadrupole!


## Trajectory through quadrupole



## Trajectory through quadrupole

Beam trajectory through a quadrupole is well known solution from Hill's equation:

$$
\begin{aligned}
&\binom{x_{1}}{x_{1}^{\prime}}=\left(\begin{array}{cc}
\cos \sqrt{k_{1}} s & \frac{1}{\sqrt{k_{1}}} \sin \sqrt{k_{1}} s \\
-\sqrt{k_{1}} \sin \sqrt{k_{1}} s & \cos \sqrt{k_{1}} s
\end{array}\right)\binom{x_{0}}{x_{0}^{\prime}} \\
& \operatorname{Or}\binom{x_{1}}{x_{1}^{\prime}}=\left(\begin{array}{cc}
\cosh \sqrt{k_{1}} s & \frac{1}{\sqrt{k_{1}}} \sinh \sqrt{k_{1}} s \\
\sqrt{k_{1}} \sinh \sqrt{k_{1}} s & \cosh \sqrt{k_{1}} s
\end{array}\right)\binom{x_{0}}{x_{0}^{\prime}}
\end{aligned}
$$

(focusing quad)
(defocusing quad)

But what about the dispersion?
As $\rho$ varies, we cannot use the same equations as for a combined function magnet...
For simplicity, we will define the dispersion function as:

$$
\left(\begin{array}{c}
D_{x, 1} \\
D_{x, 1}^{\prime} \\
1
\end{array}\right)=\left(\begin{array}{ccc}
M_{11} & M_{12} & D_{q} \\
M_{21} & M_{22} & D_{q}^{\prime} \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
D_{x, 0} \\
D_{x, 0}^{\prime} \\
1
\end{array}\right)
$$

Where $M_{i j}$ are the quadrupole transfer matrix elements
$D_{q}$ and $D^{\prime}{ }_{q}$ are the dispersive contributions of the quadrupole, which we shall determine...

## Off-axis dispersion

Recall that dispersion can be defined as:

$$
\begin{aligned}
& D_{x}=M_{12}(l) \int_{0}^{l} \frac{\widetilde{M}_{11}(s)}{\rho} d s-M_{11}(l) \int_{0}^{l} \frac{\widetilde{M}_{12}(s)}{\rho} d s \\
& D^{\prime}{ }_{x}=M_{22}(l) \int_{0}^{l} \frac{\widetilde{M}_{11}(s)}{\rho} d s-M_{21}(l) \int_{0}^{l} \frac{\widetilde{M}_{12}(s)}{\rho} d s
\end{aligned}
$$

But we know that $\rho$ is not constant, so we need to find an expression for this...

$$
\rho(s)=\frac{d L}{d \theta}=\frac{\left(1+x^{\prime 2}\right)^{\frac{3}{2}}}{x^{\prime \prime}}
$$

But the Hill's equation, $x^{\prime \prime}=-k_{1} x$, can be used to simplify this equation:

$$
\rho(s)=-\frac{\left(1+x^{\prime 2}\right)^{\frac{3}{2}}}{k_{1} x}
$$

If $k_{1}>0$ we obtain the solution for a focusing quadrupole If $k_{1}<0$ we obtain the solution for a defocusing quadrupole

## Off-axis dispersion

Since we know that:

$$
\begin{aligned}
x(s) & =\widetilde{M}_{11}(s) x_{0}+\widetilde{M}_{12}(s) x_{0}^{\prime} \\
x^{\prime}(s) & =\widetilde{M}_{21}(s) x_{0}+\widetilde{M}_{22}(s) x_{0}^{\prime}
\end{aligned}
$$

Then:

$$
\begin{gathered}
D_{q}=k_{1} \int_{0}^{l_{q}} \frac{\left(M_{12}(l) \widetilde{M}_{11}(s)-M_{11}(l) \widetilde{M}_{12}(s)\right) x}{\left(1+x^{\prime 2}\right)^{\frac{3}{2}}} d s \\
=k_{1} \int_{0}^{l_{q}} \frac{\left(M_{12}(l) \widetilde{M}_{11}(s)-M_{11}(l) \widetilde{M}_{12}(s)\right)\left(x_{0} \widetilde{M}_{11}+x^{\prime}{ }_{0} \widetilde{M}_{12}\right)}{\left(1+\left(x_{0} \widetilde{M}_{21}+x^{\prime}{ }_{0} \widetilde{M}_{22}\right)^{2}\right)^{\frac{3}{2}}} d s \\
=k_{1} \int_{0}^{l_{q}} \frac{\left(M_{22}(l) \widetilde{M}_{11}(s)-M_{21}(l) \widetilde{M}_{12}(s)\right)\left(x_{0} \widetilde{M}_{11}+x^{\prime}{ }_{0} \widetilde{M}_{12}\right)}{\left(1+\left(x_{0} \widetilde{M}_{21}+x^{\prime}{ }_{0} \widetilde{M}_{22}\right)^{2}\right)^{\frac{3}{2}}} d s
\end{gathered}
$$

So we now have expressions for the dispersive contribution

## Example case: 1-quad local orbit bump



Dipole:
$\left(\begin{array}{ccc}\cos \theta & \rho \sin \theta & \rho(1-\cos \theta) \\ -\frac{\sin \theta}{\rho} & \cos \theta & \sin \theta \\ 0 & 0 & 1\end{array}\right)$

Drift:

$$
\left(\begin{array}{lll}
1 & L & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)
$$

## Close the orbit bump

Determine x and $\mathrm{x}^{\prime}$ just after the $1^{\text {st }}$ dipole:

$$
\binom{x_{d 1}}{x_{d 1}^{\prime}}=\binom{\rho(1-\cos \theta)}{\tan \theta}
$$

Determine $x$ and $x^{\prime}$ just before the quadrupole:

$$
\binom{x_{q 1}}{x_{q 1}^{\prime}}=\left(\begin{array}{ll}
1 & L \\
0 & 1
\end{array}\right)\binom{\rho(1-\cos \theta)}{\tan \theta}=\binom{\rho-\rho \cos \theta+L \tan \theta}{\tan \theta}
$$

Required x and $\mathrm{x}^{\prime}$ at the end of the quadrupole to close the orbit:

$$
\binom{x_{q 2}}{x_{q 2}^{\prime}}=\binom{M_{11} x_{q 1}+M_{12} x_{q 1}^{\prime}}{M_{21} x_{q 1}+M_{22} x_{q 1}^{\prime}}=\binom{x_{q 1}}{-x_{q 1}^{\prime}}
$$

Solving these simultaneous equations, we get:

$$
\frac{x_{q 1}}{x_{q 1}^{\prime}}=\frac{M_{12}}{1-M_{11}}=-\frac{1+M_{21}}{M_{22}}
$$

Now we have closed the orbit bump, but we can use this to help with the dispersion integrals

## Dispersion through the cell

To keep the maths easier, let's just look at the dispersion through the quadrupole:

$$
\left(\begin{array}{c}
D_{q 2} \\
D_{q 2}^{\prime} \\
1
\end{array}\right)=\left(\begin{array}{ccc}
M_{11} & M_{12} & D_{q} \\
M_{21} & M_{22} & D_{q}^{\prime} \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
D_{q 1} \\
D_{q 1}^{\prime} \\
1
\end{array}\right)
$$

If we first assume that the beam travels on-axis through the quadrupole ( $D_{q}=D_{q}^{\prime}=0$ ) then if $\frac{D_{q 1}}{D_{q 1}^{\prime}}=\frac{x_{q 1}}{x_{q 1}^{\prime}}$ we know that this cell would be dispersion free at the ends...

Dispersion after dipole:

$$
\binom{D_{x 1}}{D_{x 1}^{\prime}}=\binom{\rho(1-\cos \theta)}{\sin \theta}
$$

Dispersion at the start of the quad:

$$
\left(\begin{array}{c}
D_{q 1} \\
D_{q 1}^{\prime} \\
1
\end{array}\right)=\left(\begin{array}{lll}
1 & L & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
\rho(1-\cos \theta) \\
\sin \theta \\
1
\end{array}\right)=\left(\begin{array}{c}
\rho(1-\cos \theta)+L \sin \theta \\
\sin \theta \\
1
\end{array}\right)
$$

## Dispersion through the cell

So:

$$
\begin{gathered}
\frac{D_{q 1}}{D_{q 1}^{\prime}}=\frac{\rho(1-\cos \theta)}{\sin \theta}+L \\
\frac{x_{q 1}}{x_{q 1}^{\prime}}=\frac{\rho(1-\cos \theta)}{\sin \theta} \cos \theta+L
\end{gathered}
$$

If we assume that $\theta \ll 1$ (small angle approximation) then $\cos \theta \approx 1$ and the ratios are equal.

So in this case, the dispersion for an on-axis beam will be zero at the end of the cell. What about an off-axis beam?

Assuming the small angle approximation, the transfer matrix for a dipole is:

$$
\left(\begin{array}{ccc}
\cos \theta & \rho \sin \theta & \rho(1-\cos \theta) \\
-\frac{\sin \theta}{\rho} & \cos \theta & \sin \theta \\
0 & 0 & 1
\end{array}\right) \approx\left(\begin{array}{ccc}
1 & \rho \theta & \frac{\rho \theta^{2}}{2} \\
-\frac{\theta}{\rho} & 1 & \theta \\
0 & 0 & 1
\end{array}\right)
$$

## Dispersion through the cell

Since we know that all terms that don't depend on $D_{q}$ and $D_{q}^{\prime}$ cancel through the cell, we can simply consider the propagation of $D_{q}$ and $D_{q}^{\prime}$ through the second part of the cell.
After the $2^{\text {nd }}$ drift:

$$
\left(\begin{array}{c}
D_{x 2} \\
D_{x 2}^{\prime} \\
1
\end{array}\right)=\left(\begin{array}{ccc}
1 & L & 0 \\
0 & 1 & 0 \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
D_{q}+\cdots \\
D_{q}^{\prime}+\cdots \\
1
\end{array}\right)=\left(\begin{array}{c}
D_{q}+L D_{q}^{\prime}+\cdots \\
D_{q}^{\prime}+\cdots \\
1
\end{array}\right)
$$

At the end of the cell:

$$
\left(\begin{array}{ccc}
1 & \rho \theta & \frac{\rho \theta^{2}}{2} \\
-\frac{\theta}{\rho} & 1 & \theta \\
0 & 0 & 1
\end{array}\right)\left(\begin{array}{c}
D_{q}+L D_{q}^{\prime}+\cdots \\
D_{q}^{\prime}+\cdots \\
1
\end{array}\right)=\left(\begin{array}{c}
D_{q}+(L+\rho \theta) D_{q}^{\prime} \\
-\frac{\theta}{\rho} D_{q}+\left(1-\frac{\theta L}{\rho}\right) D_{q}^{\prime} \\
1
\end{array}\right)
$$

## Dispersion through the cell

This is often known as RESIDUAL DISPERSION
In low emittance machines, this can be a limitation on performance:

- Difficult to locate sources of residual dispersion, so difficult to correct it
- Beam based alignment, such as DISPERSION FREE STEERING, will reduce this
- Depends on beam position jitter, so often varies pulse to pulse
- Can also depend on charge jitter if wakefields are a problem.

It is beyond the scope of this lecture, but it can be shown that it is not possible to have zero residual dispersion through any optical system when the beam travels off-axis.

## Nonlinear extension

Consider a beam travelling through a higher order multipole, like a sextupole:
On-axis:
$B_{y}=\frac{p}{c} k_{n} x^{n}$
For example, sextupole field: $B_{y}=\frac{p}{c} k_{2} x^{2}$


If the beam travels off-axis by a distance $\delta x$, then the field becomes:
$B_{y}=\frac{p}{c} k_{n}(x+\delta x)^{n}=\sum \frac{p}{c} k_{n} \frac{n!}{k!(n-k)!} x^{n-k} \delta x^{k}$
For sextupole: $B_{y}=\frac{p}{c} k_{2} x^{2}+2 \frac{p}{c} k_{2} x \delta x+\frac{p}{c} k_{2} \delta x^{2}$



Dipole term

Sextupole term

Travelling off-axis through a multipole introduces lower order terms

## Nonlinear extension

We will only consider the dipole and quadrupole terms:

$$
\begin{gathered}
B_{y}=\frac{p}{c} k_{n} \delta x^{n}+n \frac{p}{c} k_{n} \delta x^{n-1} x+o\left(x^{2}\right) \\
K x=\left(n k_{n} \delta x^{n-1}\right) \frac{\delta x}{n}+\left(n k_{n} \delta x^{n-1}\right) x
\end{gathered}
$$

So the Hill's equation is:

$$
x^{\prime \prime}+\left(\left(n k_{n} \delta x^{n-1}\left(x, x^{\prime}, s\right)\right) \frac{\delta x\left(x, x^{\prime}, s\right)}{n x}+\left(n k_{n} \delta x^{n-1}\left(x, x^{\prime}, s\right)\right)\right) x=0
$$

But for $n>1$ (i.e. sextupole or higher), this is nonlinear and cannot be solved analytically. If we consider splitting the multipole into a large number of thin slices, then $\delta x$ can be considered constant and each slice can be treated as a quadrupole and dipole term.

## Nonlinear extension

If we consider the Hill's equation for an off-axis quadrupole:

$$
K x=k_{1} \delta x+k_{1} x
$$

Comparing this to the Hill's equation for an off-axis multipole:

$$
K x=\left(n k_{n} \delta x^{n-1}\right) \frac{\delta x}{n}+\left(n k_{n} \delta x^{n-1}\right) x
$$

Then we can compare terms:

$$
\tilde{k}_{1}=n k_{n} \delta x^{n-1}
$$

And the off-axis multipole term becomes:

$$
K x=\tilde{k}_{1} \frac{\delta x}{n}+\tilde{k}_{1} x
$$

Therefore travelling off-axis through a multipole a distance $\delta x$ is equivalent to travelling off-axis through a quadrupole by a distance $\frac{\delta x}{n}$.

- This means that the residual dispersion can be reduced, but still not removed.
- The strong dependence of magnetic field to radial position means that the sensitivity to beam jitter increases by a factor $n$; so nonlinear optics can cause more problems.


## Optimisation strategies

Consider a short beamline consisting of 3 sections:

1) A quadrupole matching section
2) An achromatic arc section
3) Another quadrupole matching section


## Local optimisation

Local optimisation:

- Match initial beam parameters into arc cell
- Match optics and dispersion in arc cell
- Match arc cell beam parameters into final optics

Advantages:

- Easy to implement in simulations such as MADX, PLACET, ELEGANT...
- Modular: can modify each section independently
- Good for linear transverse optics
- Matching Twiss parameters
- Dispersion

Disadvantages

- Not good for complex problems:
- Chromatic corrections
- Nonlinear optics
- Longitudinal optics


## Global optimisation

## MATCH EVERYTHING TOGETHER!

Advantage

- Very good at optimising complex problems

Disadvantages

- More difficult to implement
- Large number of parameters to optimise: optimal solution can be difficult to find
- Often good to use local optimisation before global optimisation.


## Example: CLIC drive beam TAL



## Locally optimised solution



Emittance growth:
Horizontal: $\quad 138 \mu \mathrm{mRad}$
Vertical:
Longitudinal $\quad 1.8 \mu \mathrm{mGeV}$

## Globally optimised solution



Emittance growth:

Horizontal:
Vertical:
Longitudinal
$3.0 \mu \mathrm{mRad}$
$2.0 \mu \mathrm{mRad}$
-0.0012 $\mu \mathrm{mGeV}$

## Dispersion energy dependence




R56 energy dependence



