### Some considerations on the sensitivity calculation

- A coherent matrix formalism is adopted to treat correlation between different detectors and different neutrino flavors
- Inclusion of different systematic uncertainties contributions is straightforward (at the moment only the flux systematic is properly treated)
- Some results on the sensitivity for the ve appearance from anomalous νμ oscillations

#### General assumptions

- Up to now only the ve appearance channel has been considered
- Up to now only two detectors have been considered in the computations:
  - Lar1-ND detector @100 m
  - ICARUS (T600 / T1200) detector @600 m
- Only the measured ve have been considered in the sensitivity to retain only the genuine appearance signal (i.e. no direct measurement of the  $\nu\mu$  component enters in the sensitivity computation)
- Three years and 6.6e20 pot have been assumed for both detectors

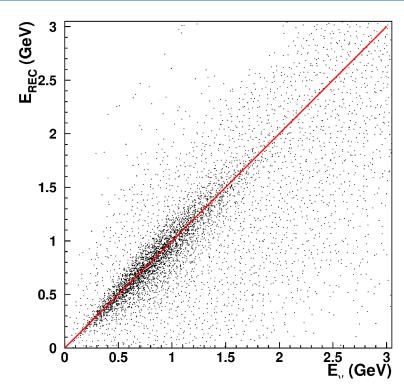
#### Global flow of the computation

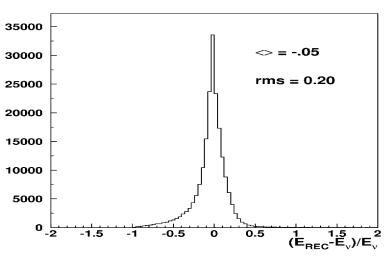
- Unoscillated spectra at the different detector locations (true v energy E<sub>v</sub>)
- Beam systematic uncertainty on the E<sub>v</sub> spectra (including correlations between different flavors/detector positions)
- Cross section systematic
- Simulation of the experimental reconstruction of the signal (reconstructed v energy E<sub>REC</sub>)
- Contributions from misidentified v interactions
- Detector/reconstruction dependent systematic uncertainties
- Estimated sensitivity
- Coherent matrix formalism to treat all systematic uncertainties and the detector response

### The ve signal: event energy reconstruction 1

The energy reconstruction has been obtained by MC simulation of the intrinsic ve events inside the detector, separately correcting the visible hadronic and leptonic energy to account for the average undetected and or non contained energy as a fuction of the vertex position (thanks to Paola).

The average reconstructed energy is close to the true neutrino energy with a smearing of ≈ 20%

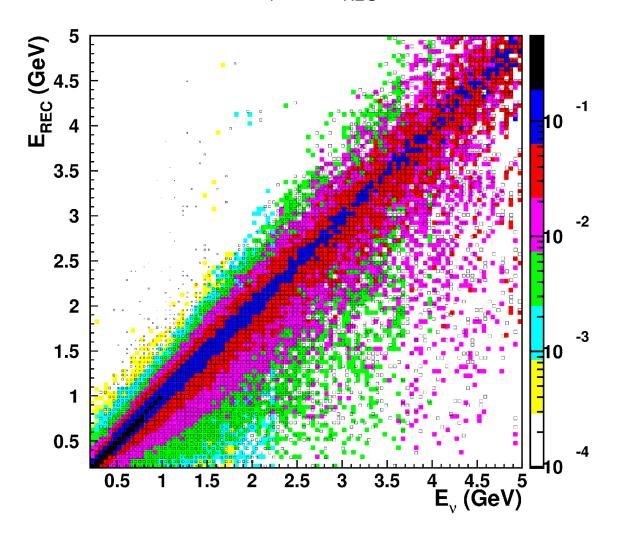




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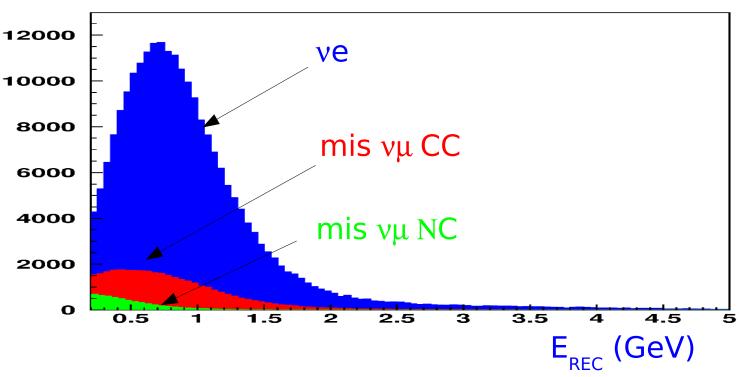
### The ve signal: event energy reconstruction 2

• The detector response is simulated with a smearing matrix  $S_{ee}(j_{v},i_{REC})$  transforming  $\mathbf{E}_{v}$  to  $\mathbf{E}_{REC}$ 



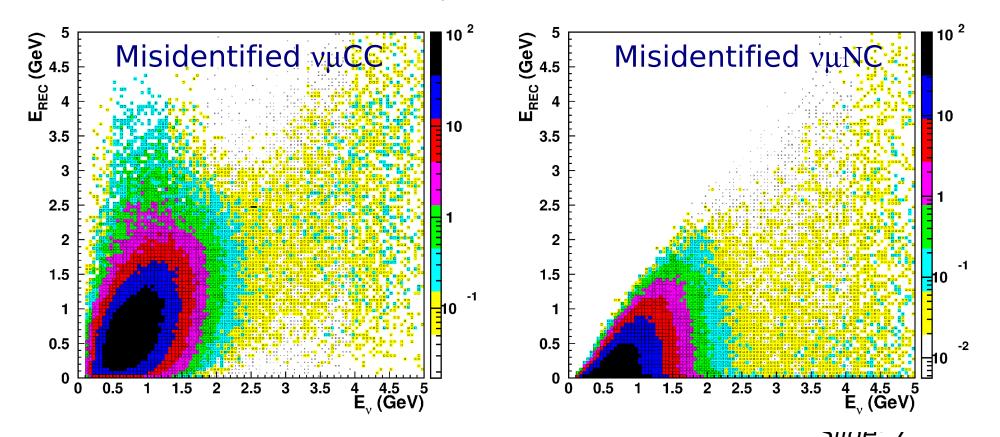
#### Background from misidentified NC and CC

- A rough estimate of the misidentified NC and CC is adopted
- The reconstructed energy for these components is computed with a scheme similar to the one adopted for the true ve events



The reconstructed energy for the  $v\mu$  misidentified as ve is computed with matrices  $B_{\mu e}^{NC,CC}(j_{\nu\mu},i_{REC})$ transforming  $E_{\nu}$  to  $E_{REC}$  and properly accounting for the relative normalization

$$n_{i_{REC}}^{NC,CC} = \sum_{j_{v\mu}} B_{\mu e}^{NC,CC}(j_{v\mu},i_{REC}) n_{j_{v\mu}}$$



#### The expected ve spectra

 At each detector the reconstructed unoscillated ve spectrum is factorized as

$$n_{e0} = S_{ee} \quad v_e CC_0 + (B_{\mu e}^{NC} + B_{\mu e}^{CC}) \quad v\mu CC_0$$
No oscillation

Using the previously defined reconstruction matrices  $S_{ee}$  for ve and  $B_{\mu e} = (B_{\mu e}^{NC} + B_{\mu e}^{CC})$  for the misidentified  $\nu\mu$ CC and  $\nu\mu$ NC and the unoscillated neutrino spectra  $\nu_e$   $CC_0$  and  $\nu\mu$   $CC_0$ 

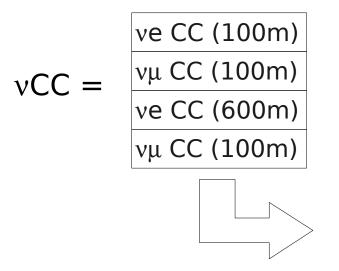
 in case of oscillation the reconstructed ve spectrum at each detector

$$n_{e\ Osc} = S_{ee} \ \nu_{e} CC_{Osc} + (B_{\mu e}^{NC} + B_{\mu e}^{CC}) \ \nu \mu CC_{Osc}$$
 With oscillation

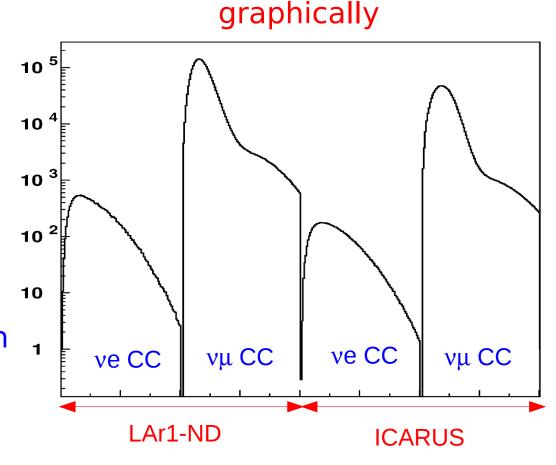
is expected to be generated by the oscillated neutrino spectra  $v_e CC_{Osc}$  and  $v\mu CC_{Osc}$ 

### The full neutrino spectra

The ve and vμ at the different sites (Lar1-ND @100m +ICARUS @ 600m) are packed together into a 4-fold vector vCC



 With such a choice both the correlations between different flavors and between different detectors can be fully exploited in the analysis

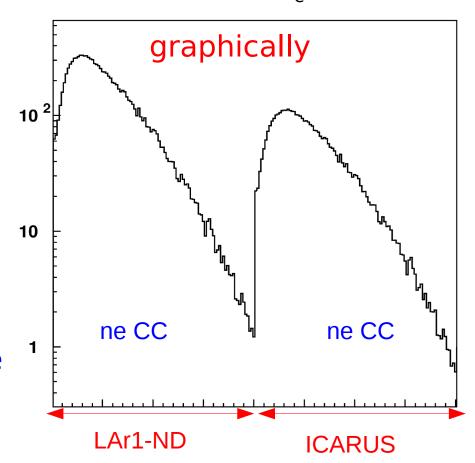


### The reconstructed ve spectra

 The reconstructed electron neutrino events at the two sites are packed together into a two-fold vector n<sub>e</sub>

$$n_e = \frac{\text{ne (100m)}}{\text{ne (600m)}}$$

 The correlations between the two detectors can be fully exploited in the analysis

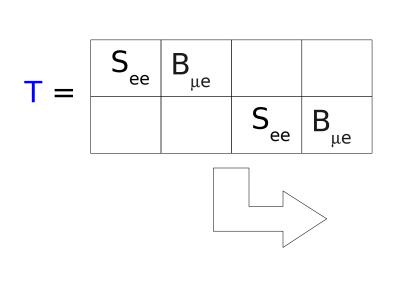


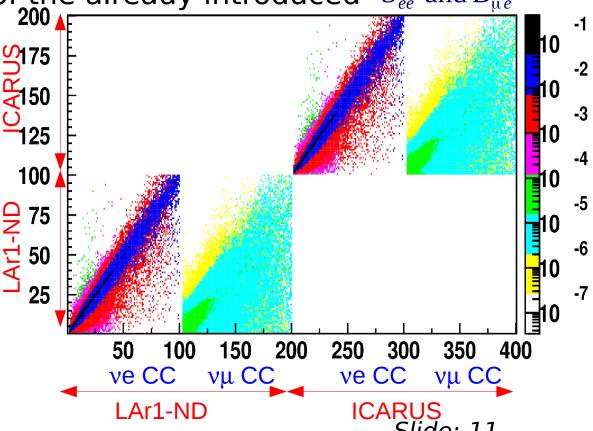
# From true vCC to reconstructed electron spectra n (1)

The measurement of the neutrino spectra by the detectors corresponds to a linear transformation mapping the 4-fold vector vCC to the two-fold vector n with a 2x4 -fold matrix T:

$$n_e = T \ \nu CC$$

T is a proper combination of the already introduced  $S_{ee}$  and  $B_{\mu e}$ 





# From true vCC to reconstructed electron spectra n<sub>e</sub>(2)

The linear dependence of n on vCC

$$n_e = T \nu CC$$

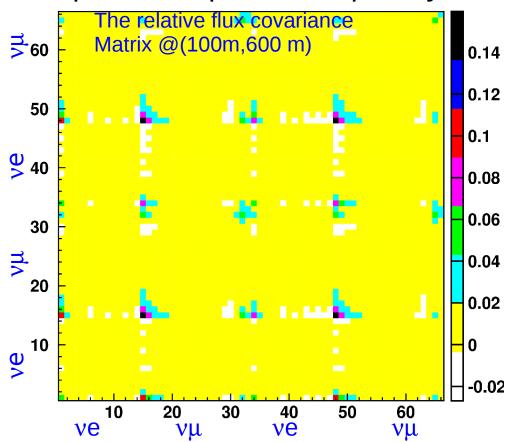
 Permits a straigthforward transformation from the vCC error matrix V(vCC) to the measurement error matrix U(n<sub>a</sub>)

$$U(n_e) = T V(vCC) T^T$$

• In this scheme the inclusion of additional contributions to the error matrix is also quite natural and simple both for the terms affecting the vCC spectrum (like e.g. those generated by the neutrino cross section systematic error) and for those directly associated to the measured quantities (the most trivial example is the statistical fluctuations affecting  $n_a$ )

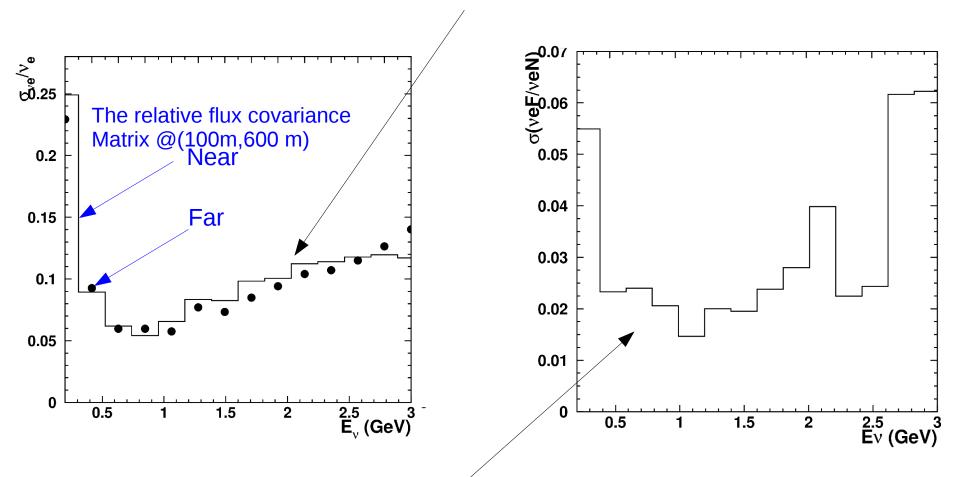
## Flux systematic contribution to V(vCC) error matrix

- In the V(vCC) error matrix is contained the BNB flux systematic error, provided by Corey (SBN Doc 47-v1) as a 4fold error matrix correlating both ve and vμ flavors and the 100m and 600m sites.
- These matrixes are quite complex with plenty of information



## Flux systematic (2)

 Typical uncertainty on the diagonal for ve is ≈6-7% (representative of the absolute error in predicting the flux at a detector



flux errors cancel out to ≈2% in the comparison of Near and Far detector

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## Flux systematic (3)

- It is useful to look at the correlation between the integrated fluxes of the ve and  $v\mu$  and at different location
- The absolute prediction for the integral flux is 6% ( $v\mu$ ) 7% (ve) in each detector.
- Each flavor permits to predict the other one with similar (≈6%) precision.
- The prediction of the same flavor Far from Near is very precise:  $4\%(v\mu)$  9‰ (ve)

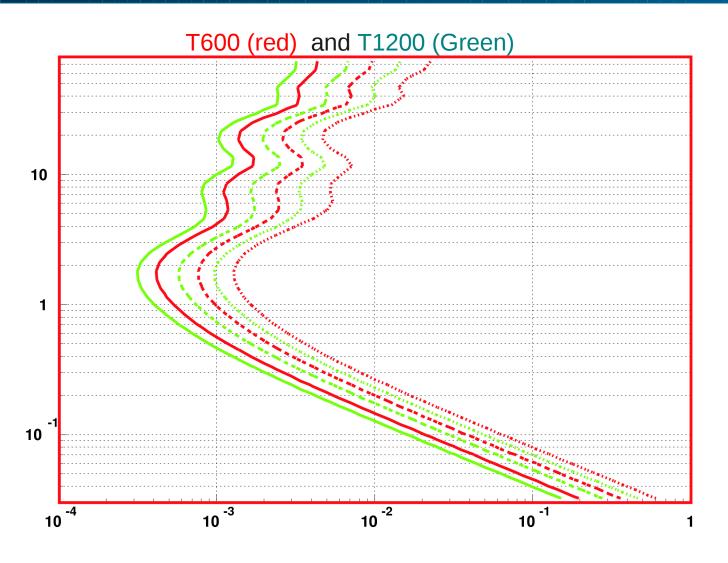
		Near		Far		
		νe	νμ	ve	νμ	
Near	νe	7	6	0,9	6	
	νμ	6	6	6	0,4	
Far	νe	0,9	6	7	6	
	νμ	6	0,4	6	6	
	N.B. Numbers are in %					

 The same conclusions are true also for the reconstructed electron neutrino events n<sub>e</sub>

## Sensitivity computation

- At the moment the oscillation excess is searched for exploiting the ratio  $R_{eNF}(E_{REC})=n_{e\ Far}/n_{e\ Near}$ . In this way we make use of the excellent precision predicted by the BNB MC for extrapolating the Far beam once it has been measured the Near site.
- R<sub>enf</sub>(E<sub>rec</sub>) measures the *change* in the oscillation probability with the distance i.e. the *relative* variations of spectrum at Far Vs Near
- R<sub>eNF</sub>(E<sub>REC</sub>), a part from second order effects, is robust against any common Near/Far systematic uncertainty
- This approach is robust even against common spectral unmodeled distortions in the Near and Far sites which could generate a fake signal in the case of a shape only analysis
- the only residual systematic effects still affecting R<sub>enf</sub>(E<sub>REC</sub>) are due to different behavior/response/experimental in the Near w.r.t. Far detector: all this kind of effects have to be carefully considered and kept under control

# Example of computation



### Still progressing

- Finalize the study of required efficiency/exposure to achieve the necessary sensitivity
- Parameterize the effect of different kind of systematics, common and relative in view of establishing their maximum acceptable level, to provide feed back on the requirements which have to be fulfilled by the experiment
- More robust estimate of the misidentification v bck and bracketing its impact on the final sensitivity
- Addition of possible cosmic bck and study its possible impact on the experimental sensitivity
- Address the νμ disappearance
- ...