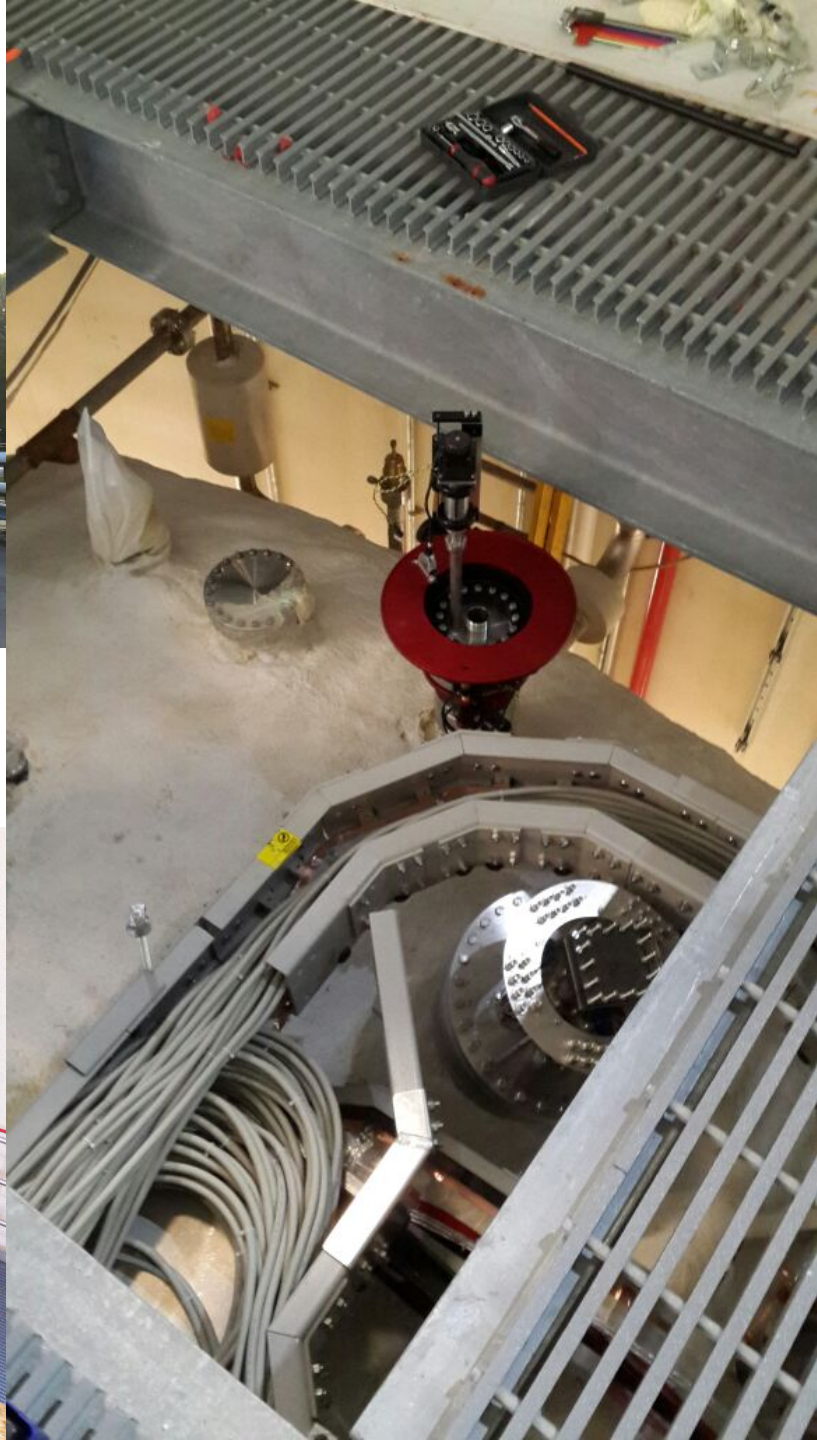
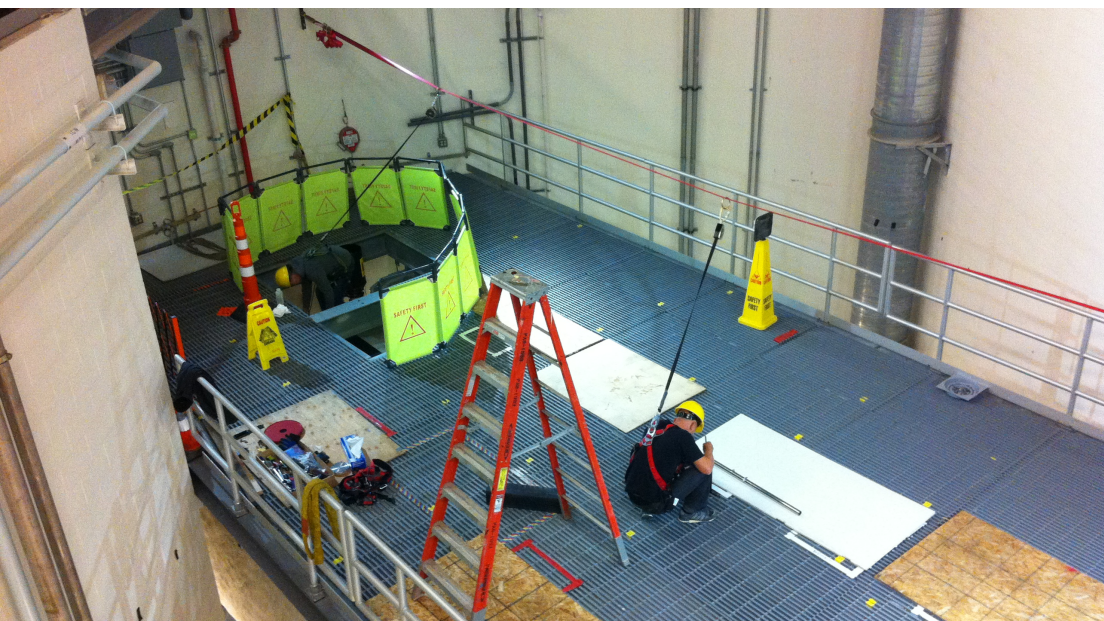


Cosmogenics

MicroBooNE



We have a building
We have a detector



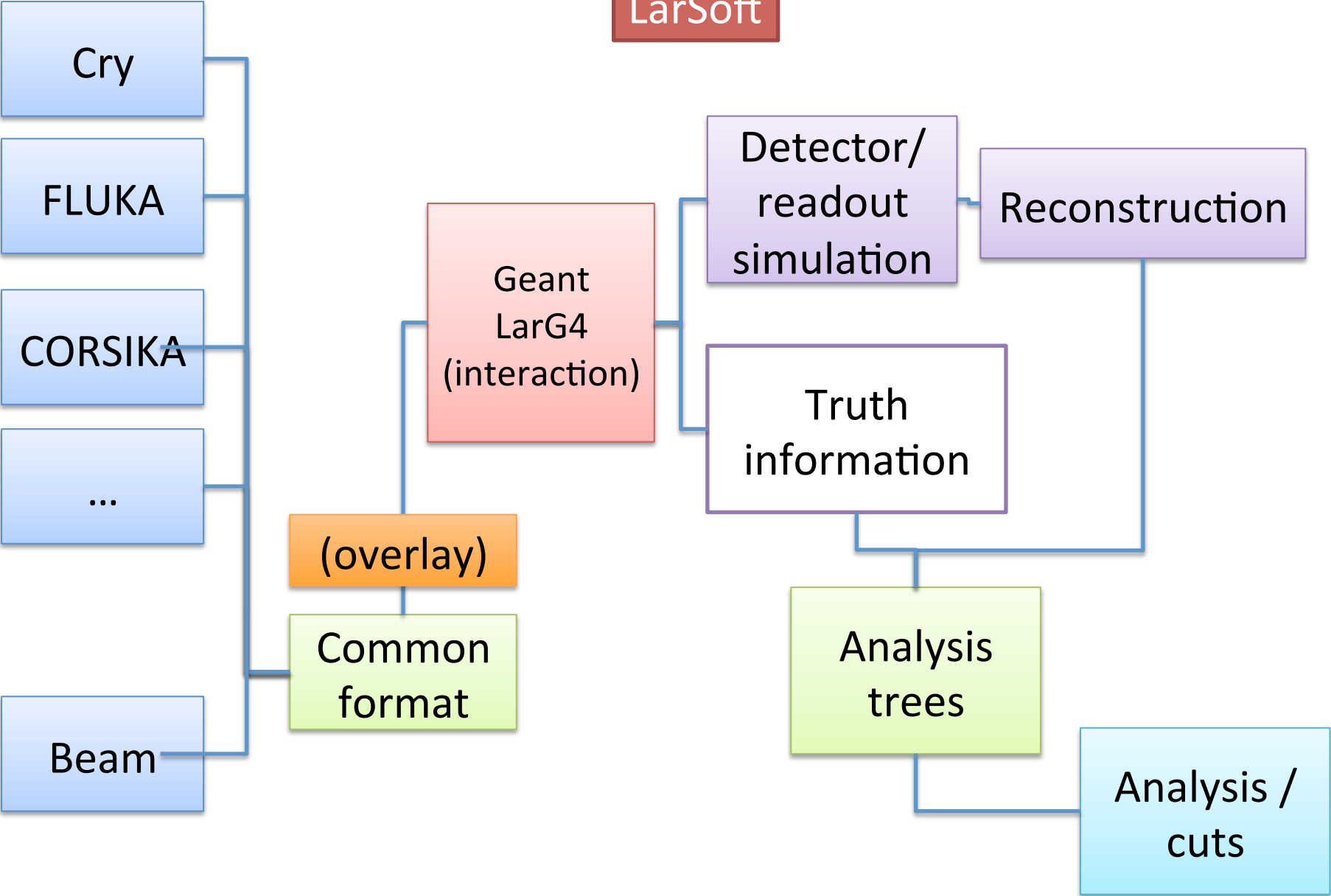
MicroBooNE work on Cosmogenics

- First measurements with cosmic rays
- Cosmic ray removal tools
- Task Force to address the over-burden decision (foreseen option for uB)
- Use LarSoft
 - considerable effort went/going to set up the infrastructure, good investment in the future

Work WG1 (from end of July talk)

- Simulation
 - Detailed FLUKA studies for ICARUS available (aim at higher statistics)
 - Profit from LBNE support and expertise
 - Better interaction simulation for μB / LAr1ND (re-interactions in dirt and detector)
- Cuts
 - Electron / photon separation (Compton + Pair)
 - Study a list of cuts with efficiencies (in close collaboration with WG2 for the signal). See slide.
- In spill / in drift time
 - Additional background in data sample
 - Overlap, reconstruction issues (energy , topology, ...)
 - Possibility of “veto” (external/internal)
 - Light system performance

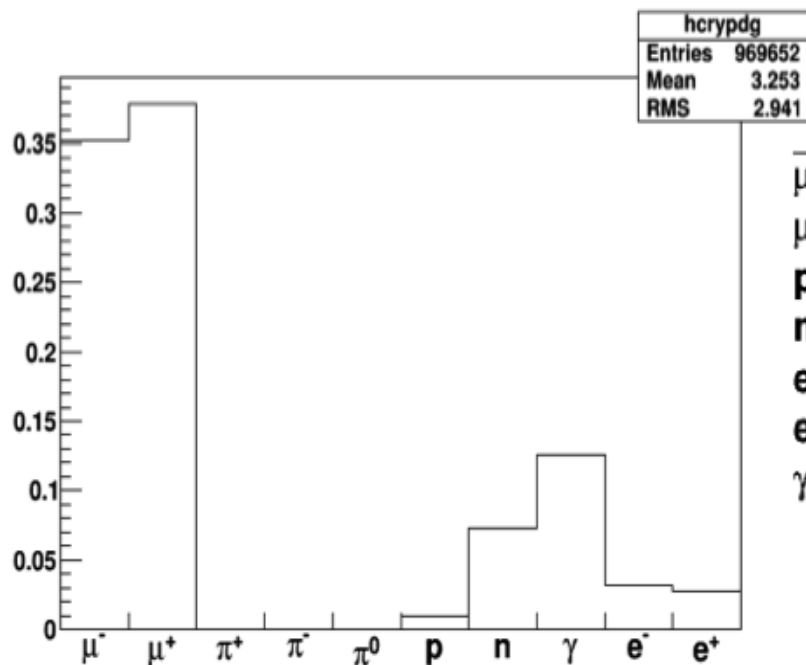
LarSoft



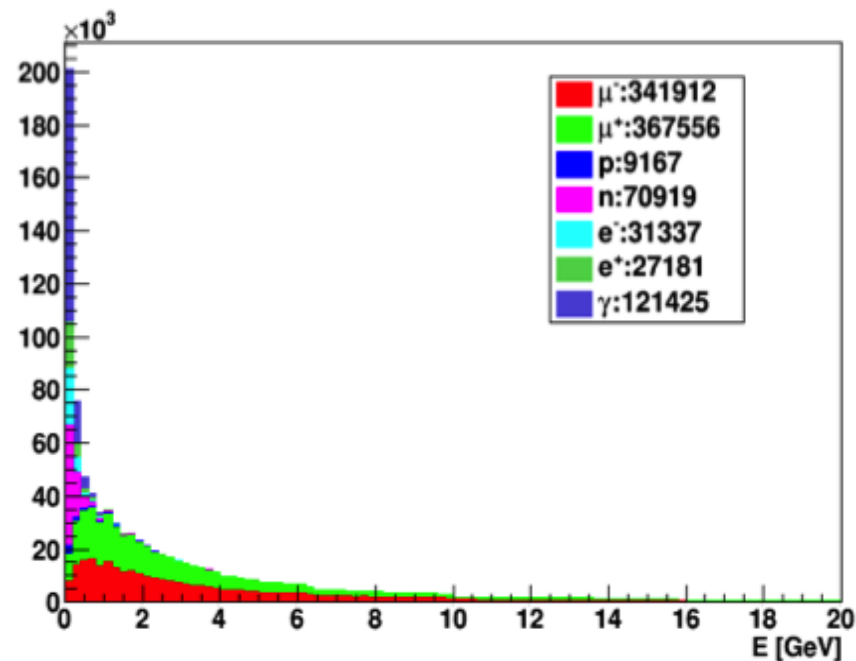
Sorted CRY Particles spectra

(rates for Cryostat+TPC)

Initial sample size = 19900 events



μ^- :341912
 μ^+ :367556
 p:9167
 n:70919
 e^- :31337
 e^+ :27181
 γ :121425



Known CRY undercounting issue:

No corrections for Neutrons/Protons/Photon applied!

Geant physics processes

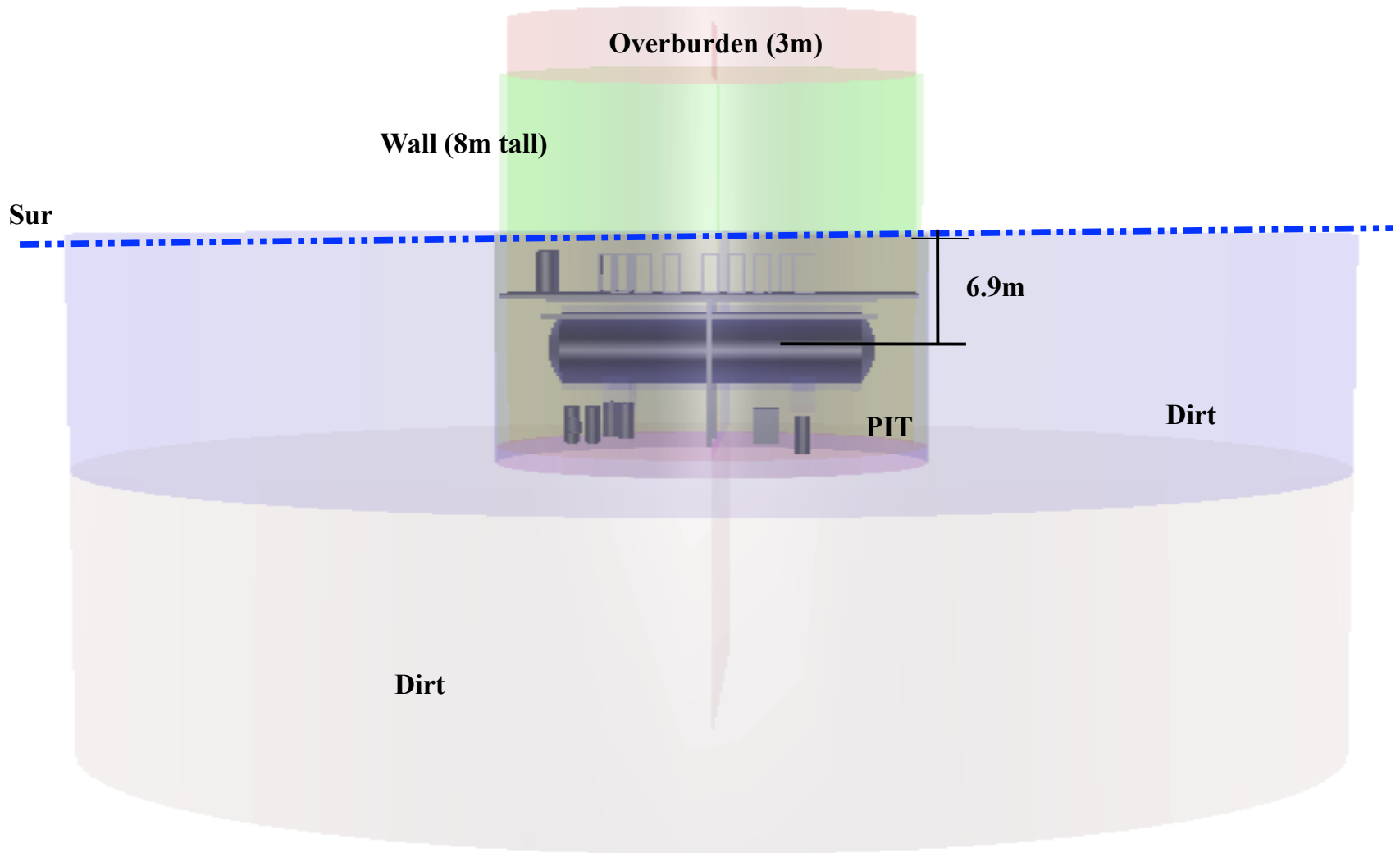
B. Eberly

Fraction of daughter particles, by species, created by each physics process

Interaction	μ^-	μ^+	e^-	e^+	γ	n	p	π^-	π^+	π^0
Ionization			69.7							
Bremsstrahlung					89.6					
Compton			22.0							
Photoelectric			7.8							
Annihilation					9.2					
Pair Production			0.4	95.5						
μ^- capture			0.02		0.6	27.2	7.6			
Hadron Elastic							6.2			
Decays	100	100	0.002	4.5	0.005					0.06
Inelastic					0.5	72.4	85.7	100	100	94.3
n Capture					0.2					
ElectroNuclear					0.001	0.08	0.3			
CHIPS Nuclear Capture						0.3	0.3			

All GEANT particles (no TPC fiducial check!)

Overburden Geometry



Effect of 3m overburden

S. Gollapinni

Initial sample size = 19300 events

All rates shown below are for particles that pass the TPC fiducial volume

	W/o overburden	W/ 3m overburden
All	308955	259585
All Primary	207122	193267
All Secondary	101833	66318

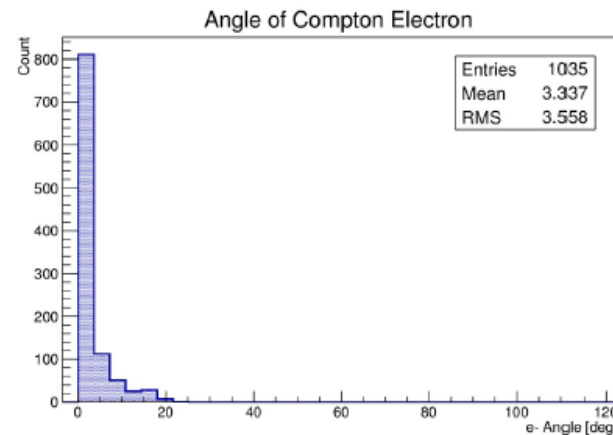
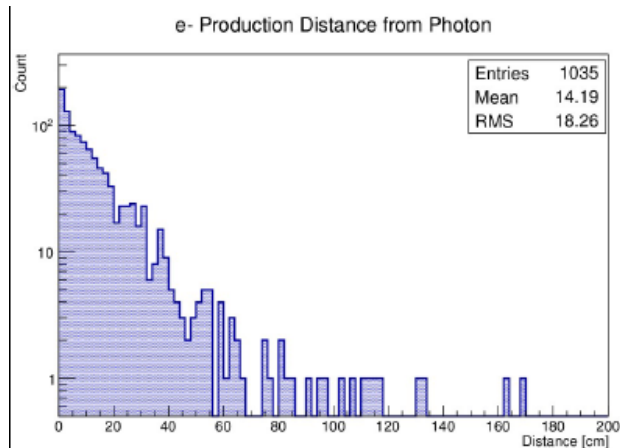
	Primary noOB	Primary 3mOB	% less	Secondary noOB	Secondary 3mOB	% less
mu-	101215	95346	-6%	16	8	-50%
mu+	104383	97687	-7%	27	12	-55%
N	446	108	-76%	19332	7701	-60%
P	112	9	-92%	3922	1382	-65%
γ	869	113	-87%	29842	18161	-39%
e-	54	2	-96%	1094	1030	-6%
e+	41	2	-95%	1948	2183	12%

NUMBERS AS ILLUSTRATION ONLY!

More analyses using the framework

- E.g. Muons with Brem -> Compton

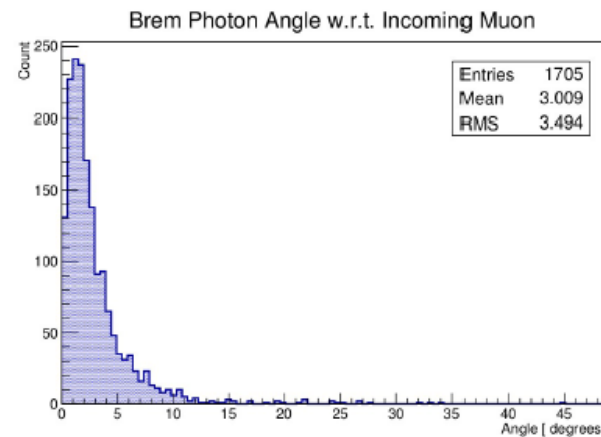
D. Caratelli



Top left: Distance from 200 MeV photon production and compton e-production

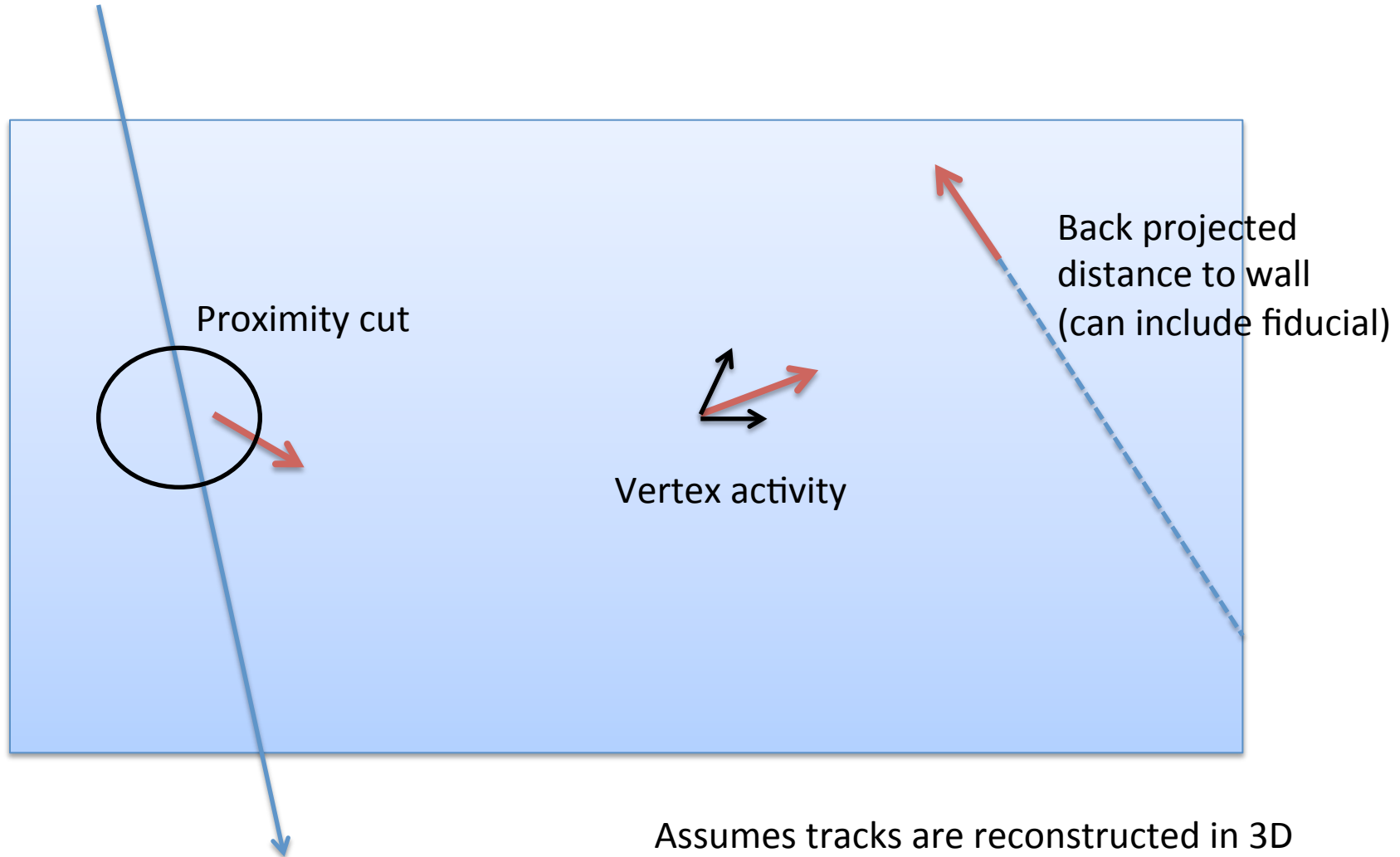
Top right: Angle between compton e-direction and photon's original direction

Bottom right: Angle between mu- and bremsstrahlung photon's direction



BACKUP

Cuts, some ideas

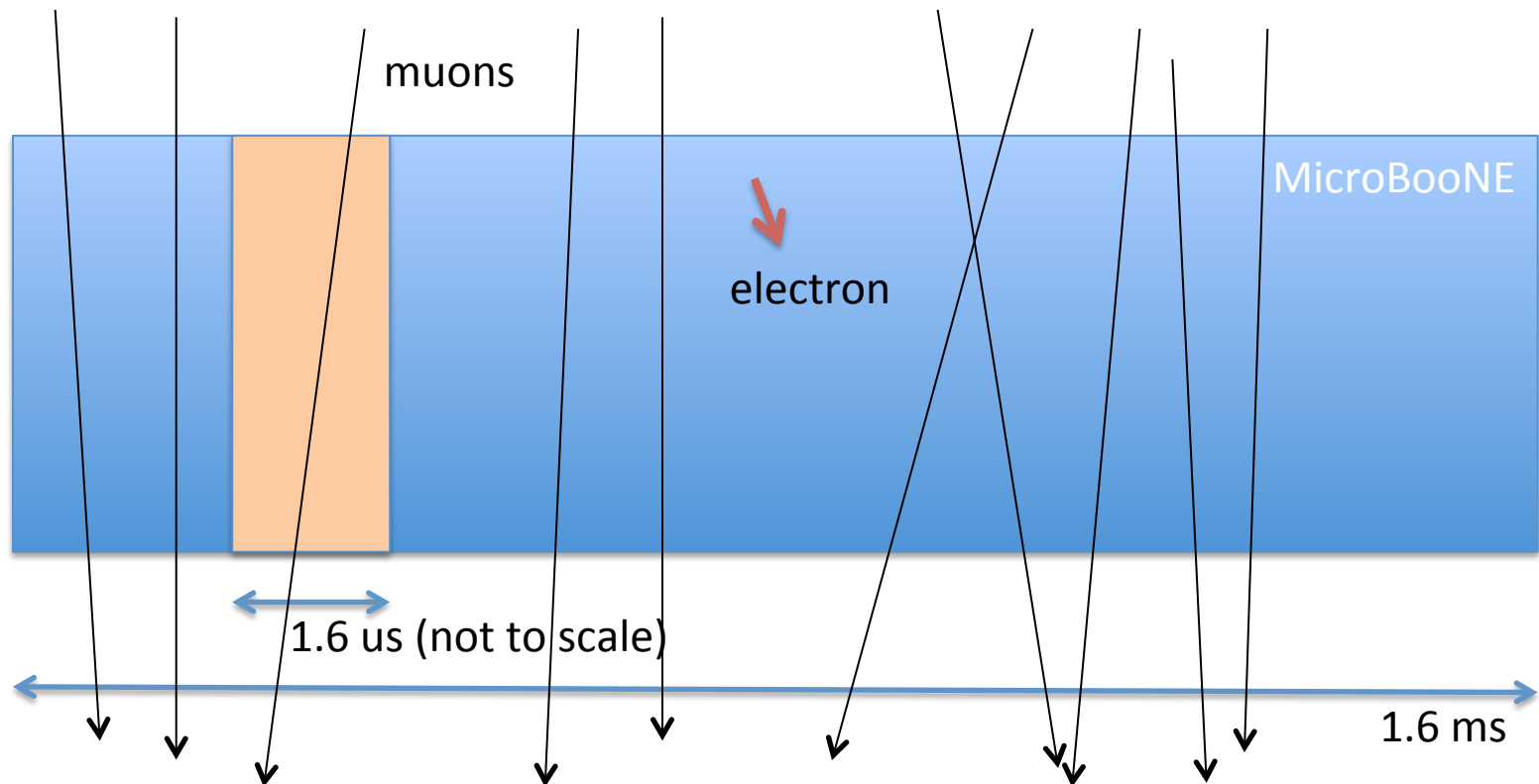


Proximity cut

Vertex activity

Back projected
distance to wall
(can include fiducial)

Assumes tracks are reconstructed in 3D
(2D plus unique time association)



- Electron (light pulse) outside the spill time (alone) will not be selected for analysis
- Without any attempt to cut them, the electrons in the spill time are of the same order as the intrinsic nue (~ 600 for $6.6E20$ pot)
- Additional contribution from other activity in the event that gets in the spill time window (overlap and ambiguity because of multiple light pulses)
 - Probability in uB to have a muon in the beam spill is ~ 10 muons in 1.6 ms $\rightarrow 1/100$ muons in 1.6 us
 \rightarrow At most a factor of $1.6\text{ms}/1.6\mu\text{s}/100=10$ times more electrons entering the analysis, which needs to be resolved by the optical system timing and position (or an external muon veto)
- For other beam related tracks (in the beam time: numuCC) we can remove the events if it overlaps with a cosmic electron. It removes $1/200$ beam events or accept e.g. as numuCC. More careful studies for dirt interactions needed.