

# Neutrino cosmic viscosity and its effect on the CMB

- von Neumann equation

$$i(\partial_t - H\rho\partial_\rho)\varrho_m = [\mathcal{H}_m, \varrho_m]$$

- with the free neutrino Hamiltonian

$$\mathcal{H}_m = \sqrt{M^2 + p^2}\mathbb{I}$$

- generating an increase in entropy

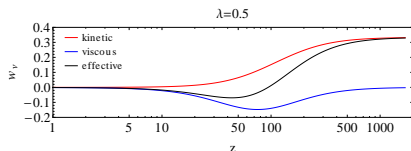
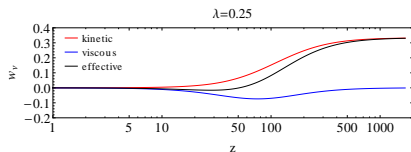
$$\dot{S}_\nu = - \int d\mathbf{q} \text{Tr} \left[ \dot{\varrho}_m \ln \left[ \frac{\varrho_m}{\mathbf{1} - \varrho_m} \right] \right]$$

- and manifesting as a viscous pressure

$$P_\nu = P_k - \frac{\dot{S}_\nu n_a T_a}{9H^2} u_{i;\lambda}$$

- we assume an increase in entropy proportional to the change of the equation of state

$$\dot{S}_\nu = -\lambda \dot{w}_\nu \ln[w_\nu]$$



- which affects the CMB by the ISW effect

