

# The LHC & the universe

G.F. Giudice



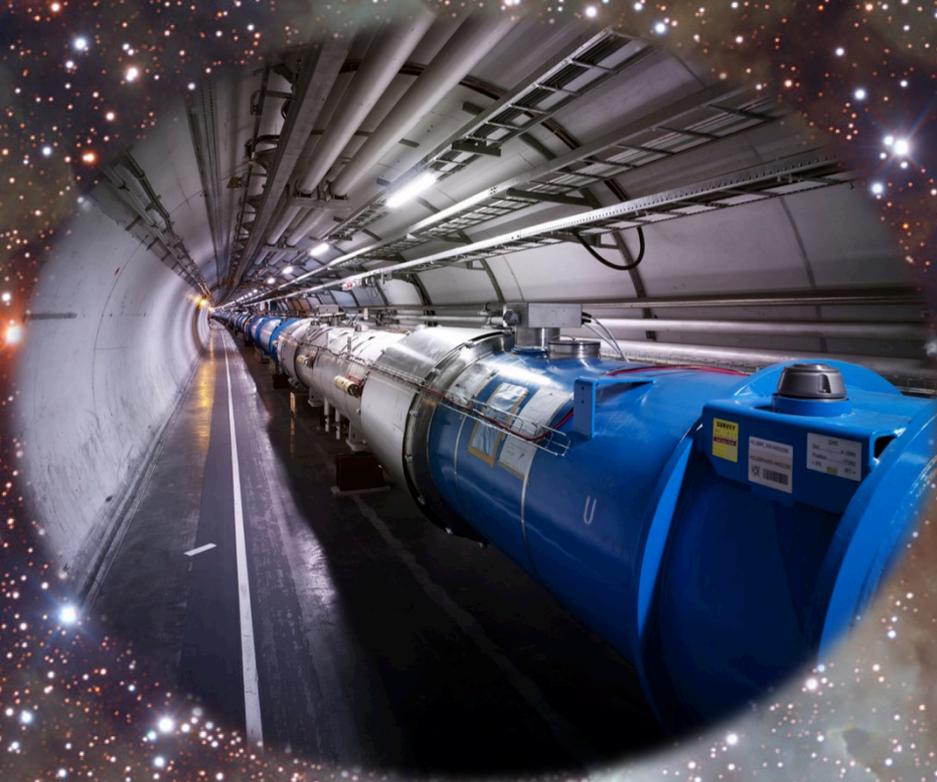
1915 - 2015



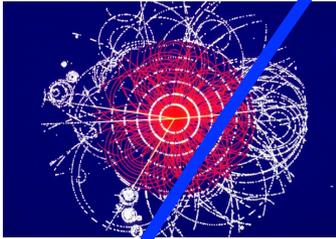
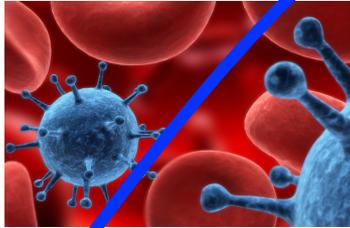
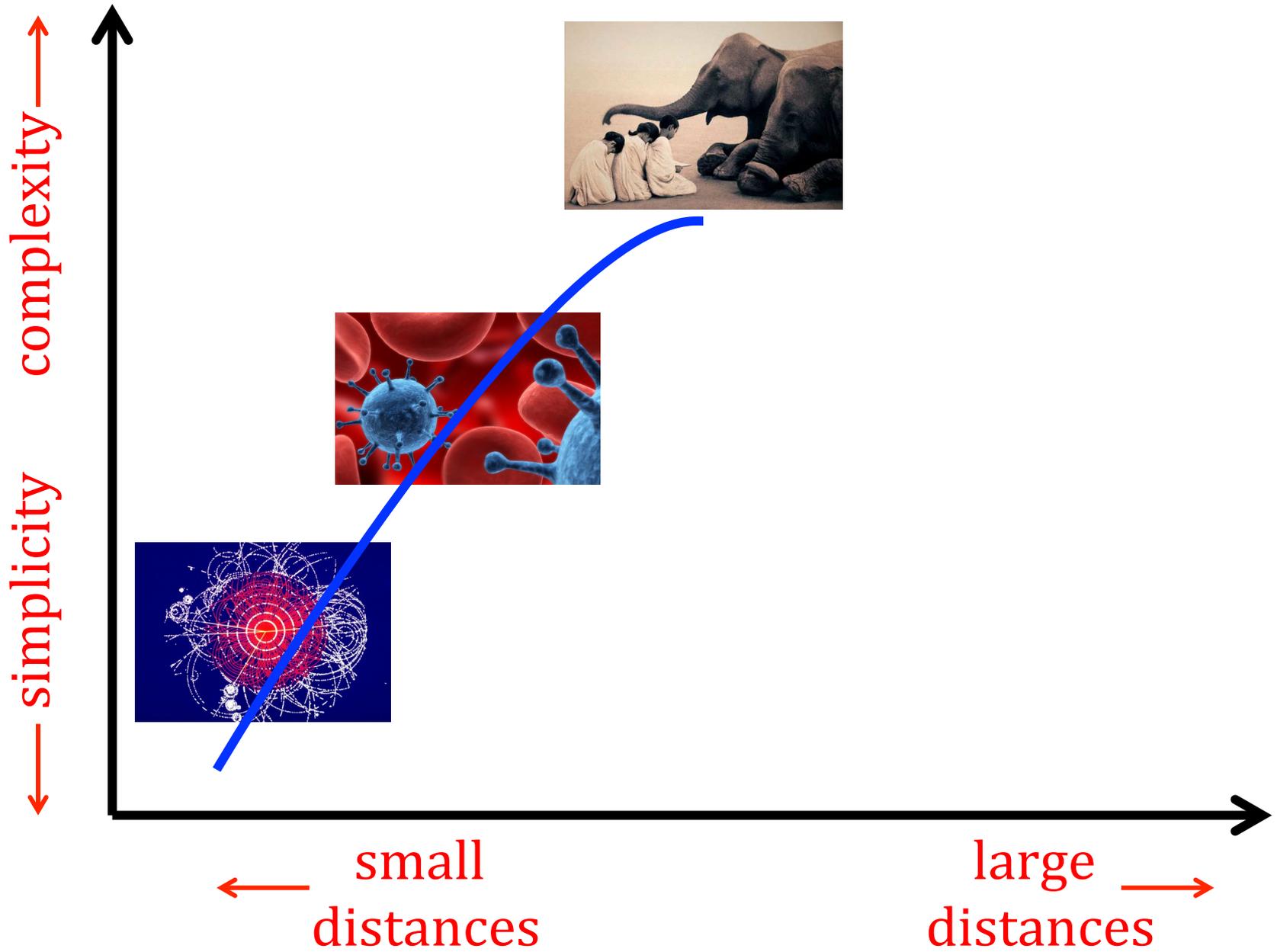
$$G_{\mu\nu} - \Lambda g_{\mu\nu} = \frac{8\pi G}{c^4} T_{\mu\nu}$$

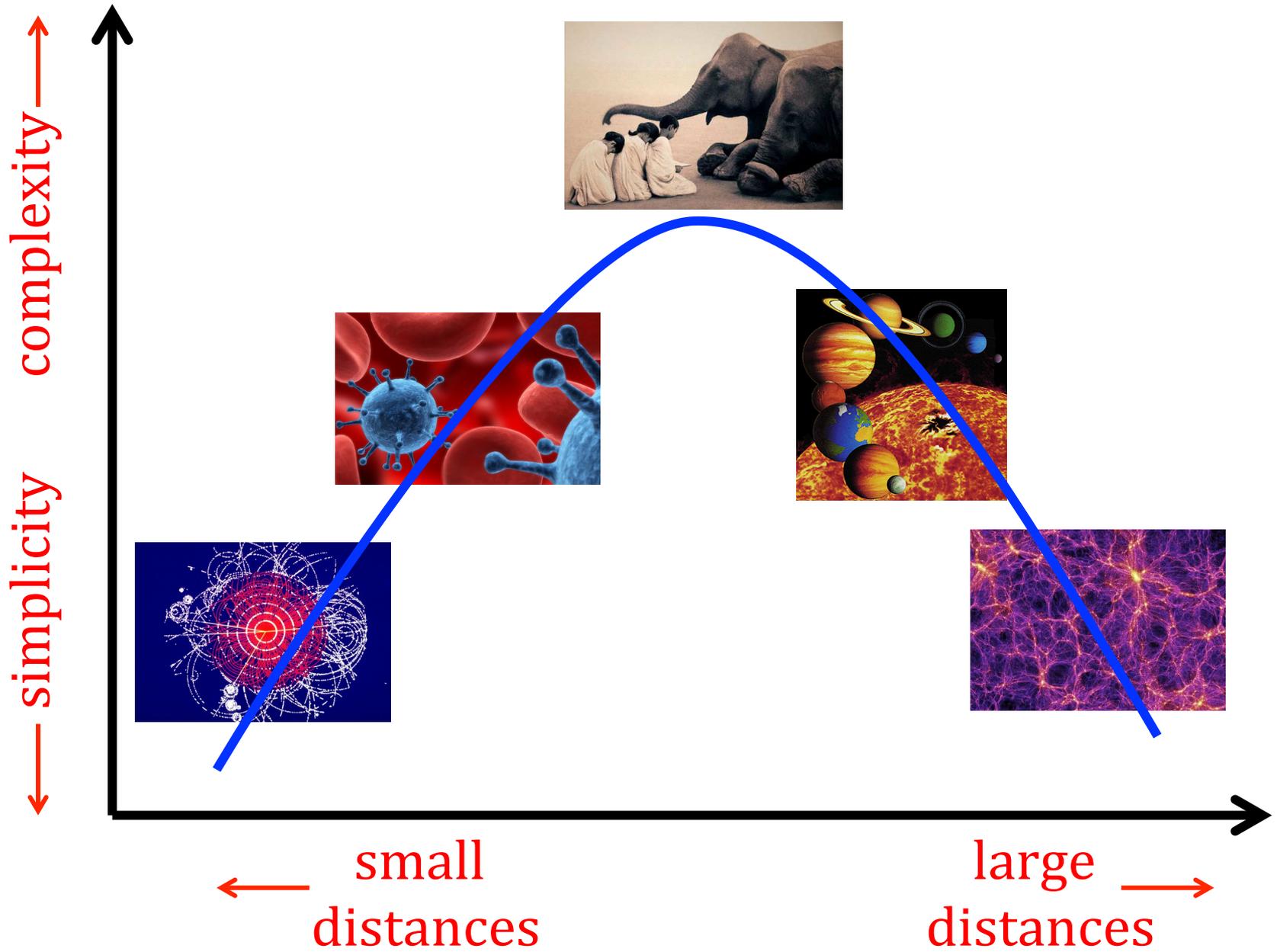
28<sup>th</sup> Texas Symposium on Relativistic Astrophysics  
Geneva, 16 December 2015

The connection between the micro & macro cosm is one of the most astounding achievements of modern physics



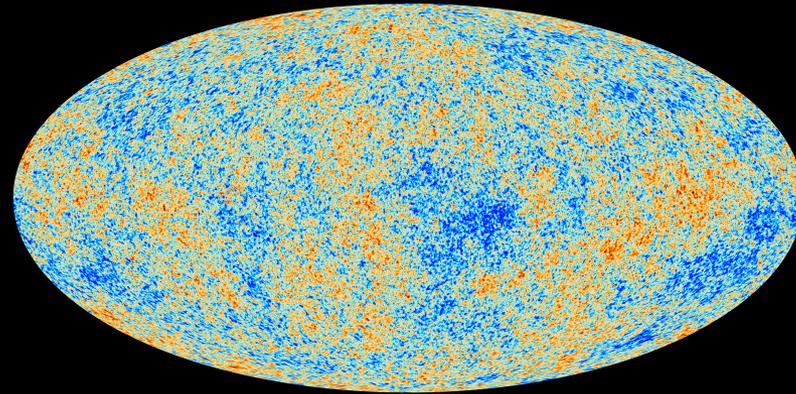
Nature's fundamental laws become apparent only at small distances



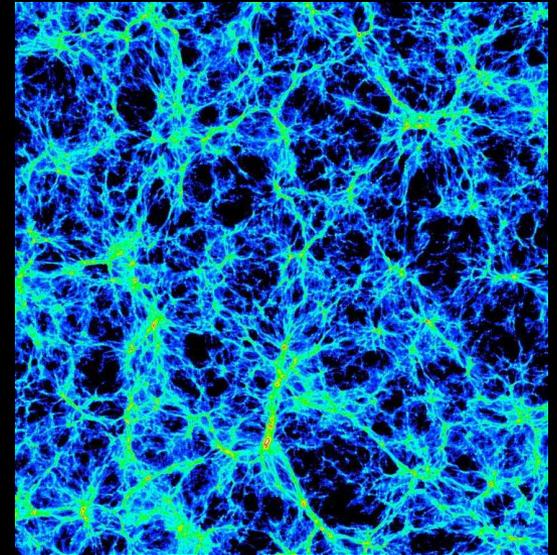


The micro/macro connection is not just a qualitative feature

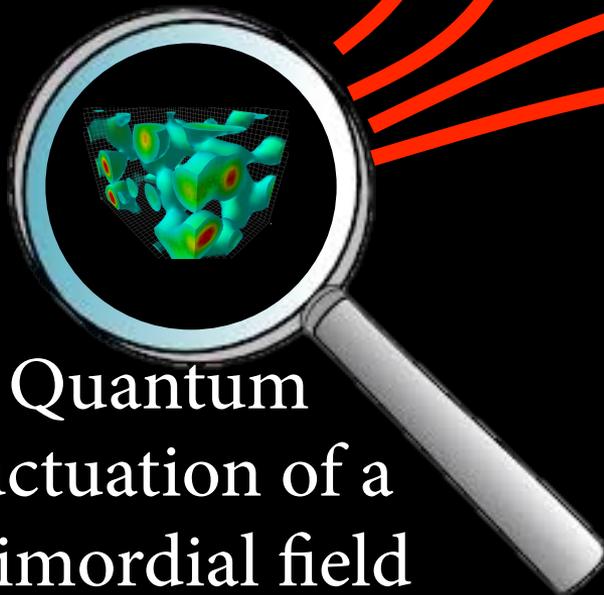
Best example:  
inflation



CMB anisotropy



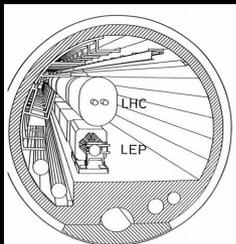
Large scale structures



Quantum fluctuation of a primordial field

# LHC: timeline

21 Mar 1984  
First LHC studies



16 Dec 1994  
LHC approved



10 Sep 2008  
LHC starts



19 Sep 2008  
LHC incident



20 Nov 2009  
LHC restarts



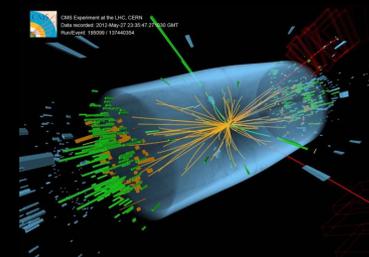
30 Mar 2010  
First collisions  
at 7 TeV



30 Oct 2011  
End of  $pp$  run  
( $5 \text{ fb}^{-1}$ )



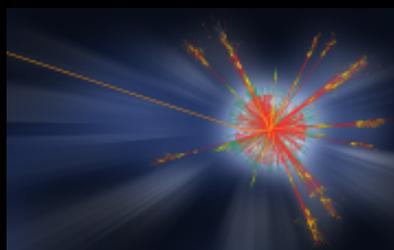
5 Apr 2012  
First collisions  
at 8 TeV



4 Jul 2012  
Higgs discovery



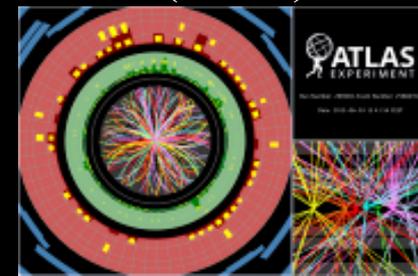
19 Jan 2013  
End of  $pp$  run  
( $20 \text{ fb}^{-1}$ )



03 Jun 2015  
First collisions  
at 13 TeV



04 Nov 2015  
End of  $pp$  run  
( $4 \text{ fb}^{-1}$ )





The last brick of the SM building,  
... but a special brick...

The success of the SM rests on the gauge principle

- elegant
- robust
- predictive

Higgs interactions: first known example  
of non-gauge fundamental forces?

The origin of all SM problems

$$L = (h_{ij} \bar{\psi}_i \psi_j H + \text{h.c.}) - \lambda |H|^4 + \mu^2 |H|^2 - \Lambda_{CC}^4$$

Flavor puzzle

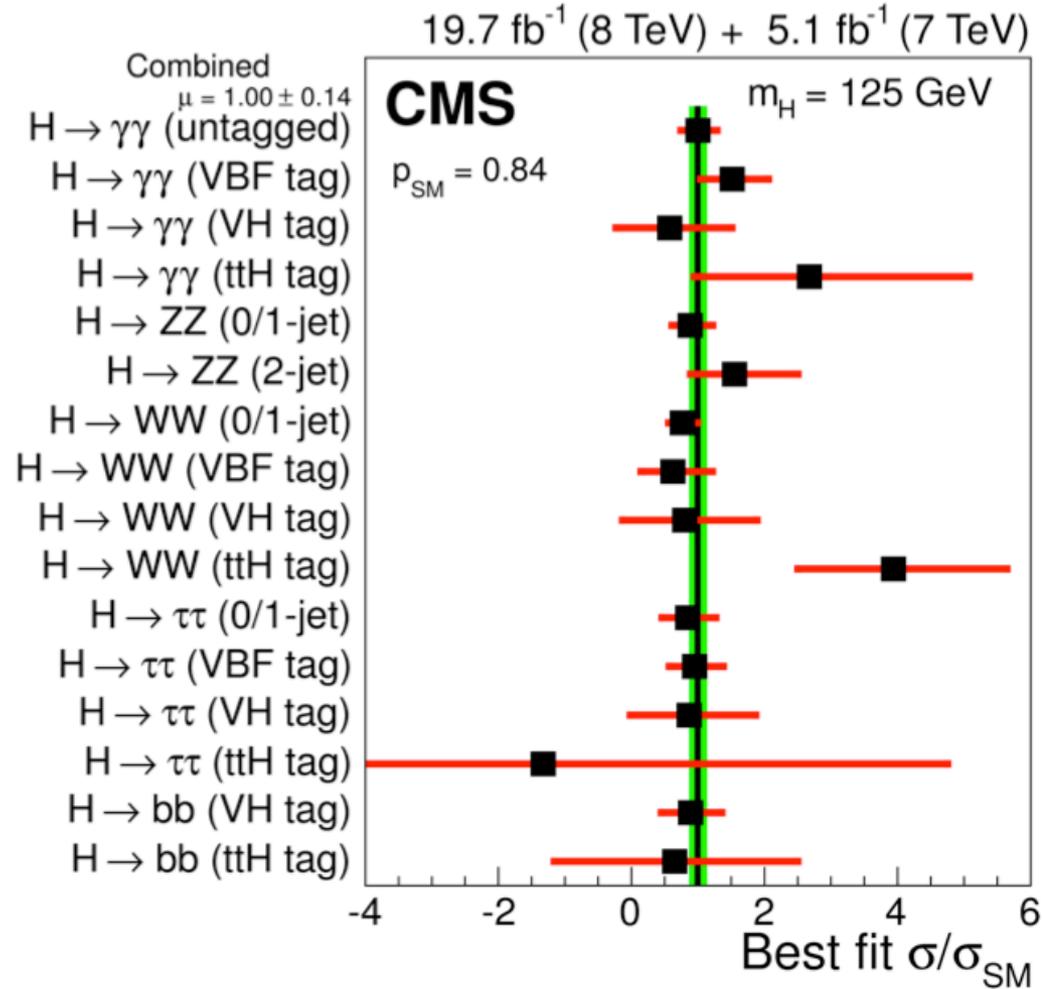
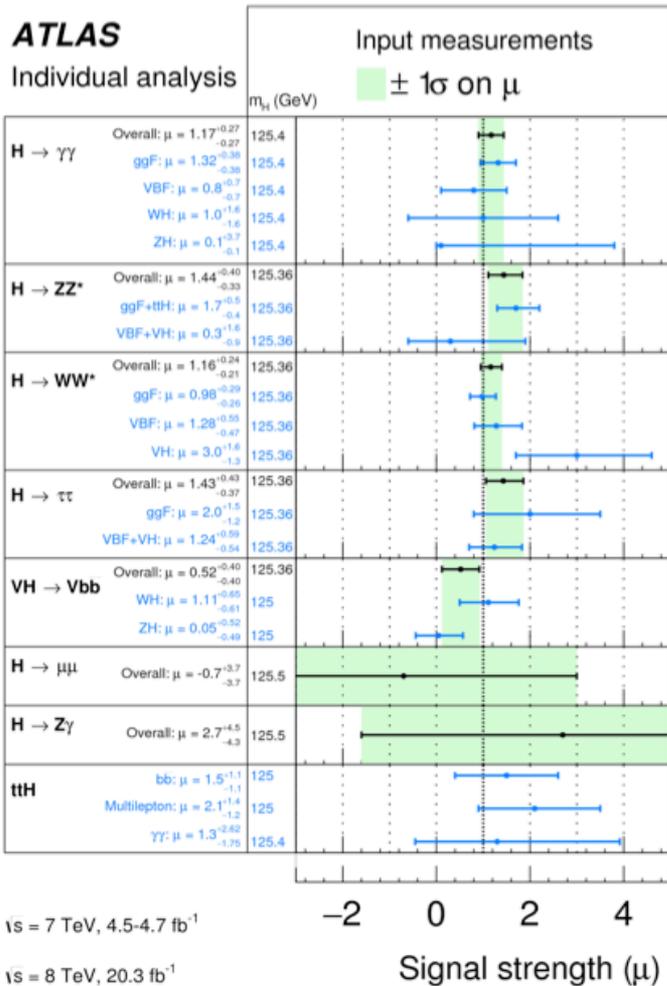
Stability of the potential

Hierarchy problem

Cosmological constant problem



# Is the Higgs a fundamental scalar with non-gauge interactions?



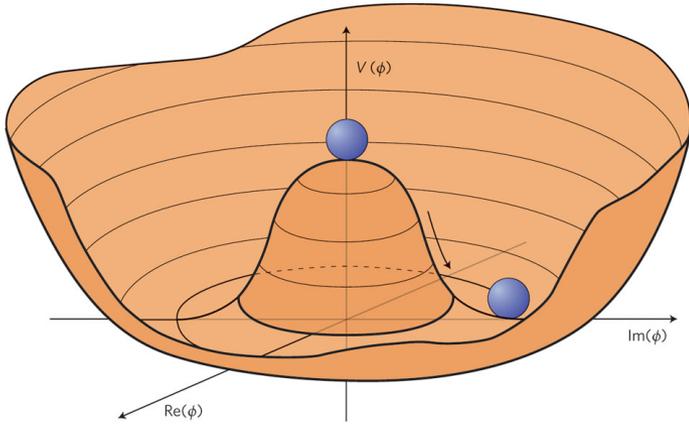
The Higgs is SM-like at the level of 20-40% accuracy

$$\Delta = \frac{v^2}{f^2} \Rightarrow \text{compositeness scale } 4\pi f > \sqrt{\frac{20\%}{\Delta}} \approx 7 \text{ TeV}$$

Does new dynamics reinstate the gauge principle?  
Is the Higgs truly fundamental?  $\Rightarrow$  experiments!

If so, fuel for inflationary model-building  
Scalar fields no longer tools for toy models

Extreme case: the Higgs as inflaton

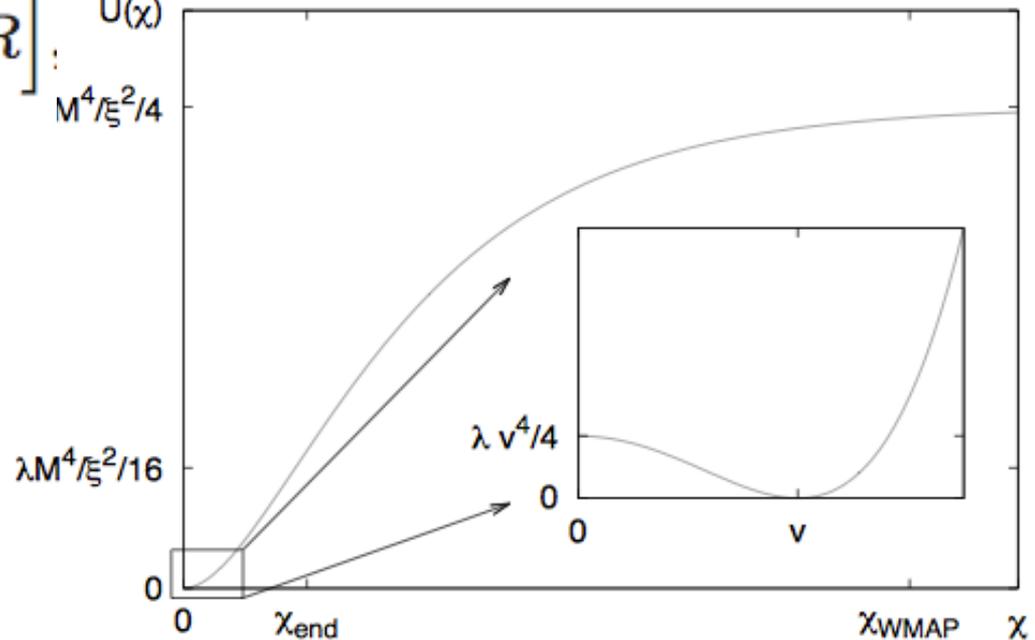


needs large-field modification

Simplest example: Higgs-curvature term

$$\delta S_{\text{NM}} = \int d^4x \sqrt{-g} \left[ -\xi \Phi^\dagger \Phi R \right]; \quad U(\chi) = \frac{M^4}{\xi^2/4}$$

Bezrukov-Shaposhnikov

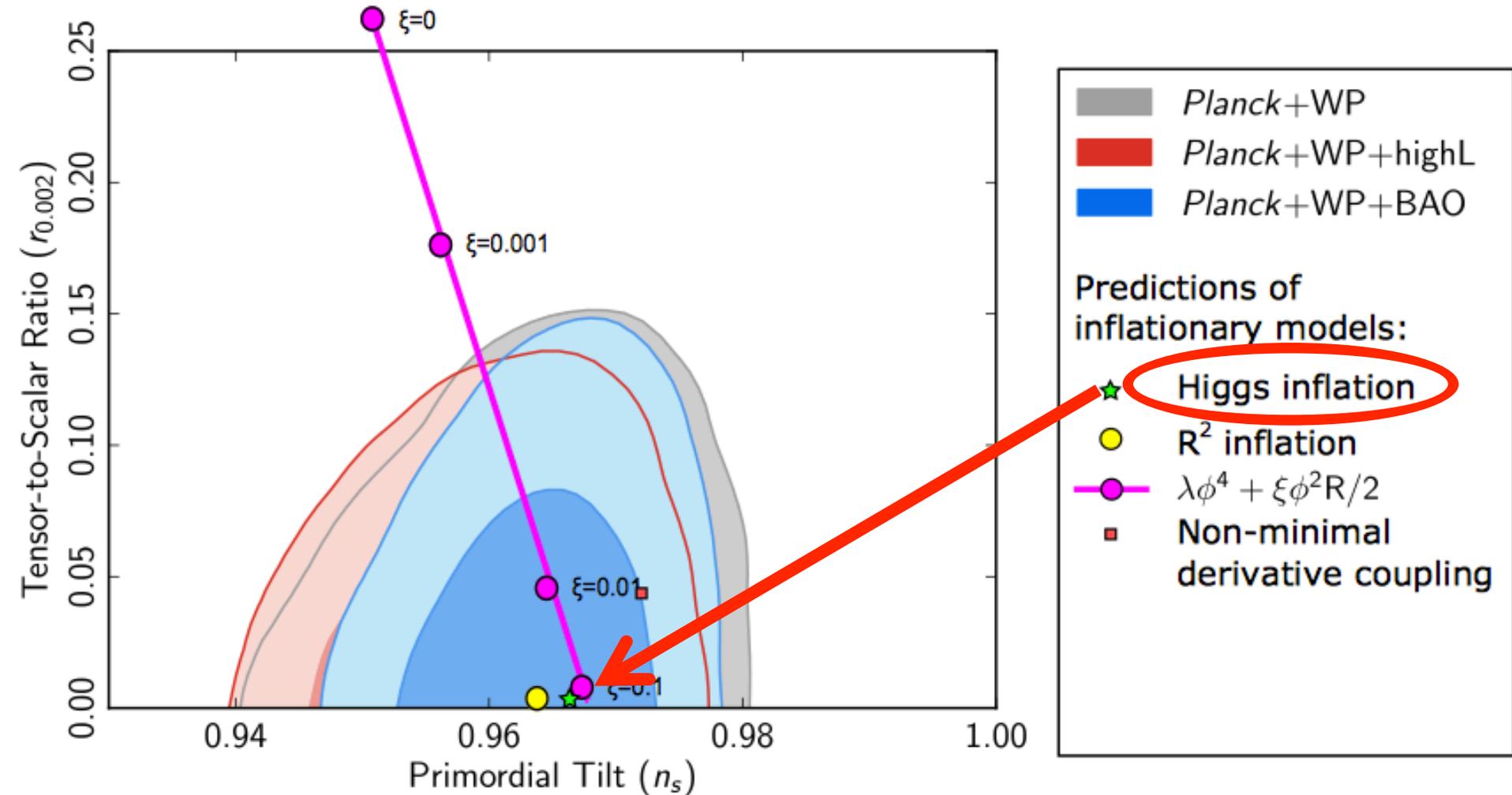


$$\xi \simeq 47\,000\sqrt{\lambda}$$

$$N \simeq 57.7$$

$$n_s \simeq 1 - 8 \frac{4N + 9}{(4N + 3)^2} \simeq 0.967,$$

$$r \simeq \frac{192}{(4N + 3)^2} \simeq 0.0031.$$



Perturbative unitarity violated at  $M_P / \xi$

New physics must appear at that scale

(we cannot trust the potential above the cutoff)

Next-to-simplest example: Higgs + 1 scalar

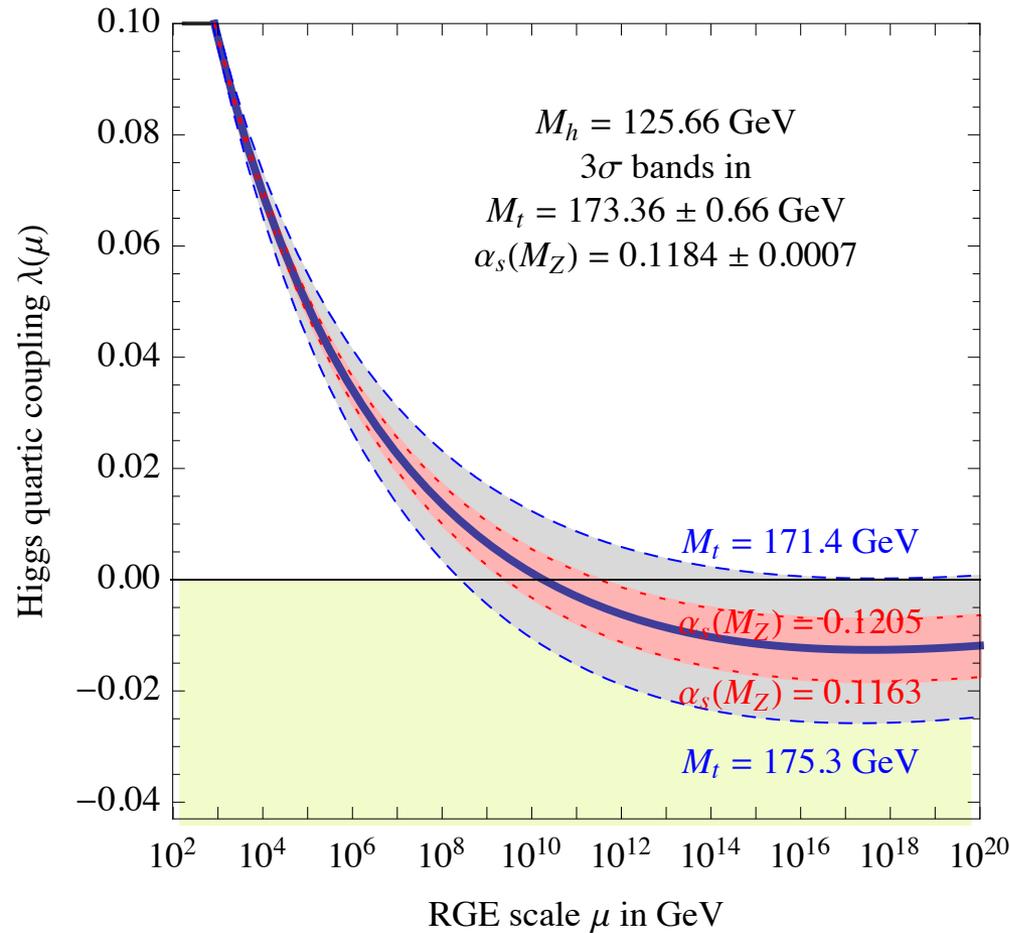
Giudice-Lee

Barbon-Casas-Elias-Espinosa

General lessons from UV-completed models

- Higgs may play a role in inflation, but it is unlikely to be the dominant source
- Inflationary predictions are fairly robust
- Except for special cases, observation of  $r$  can rule out the models

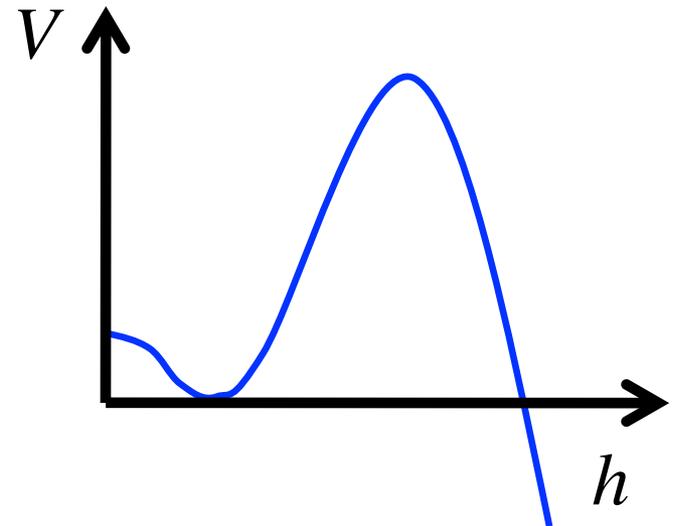
# If SM + Higgs is the full story, new implications for cosmology

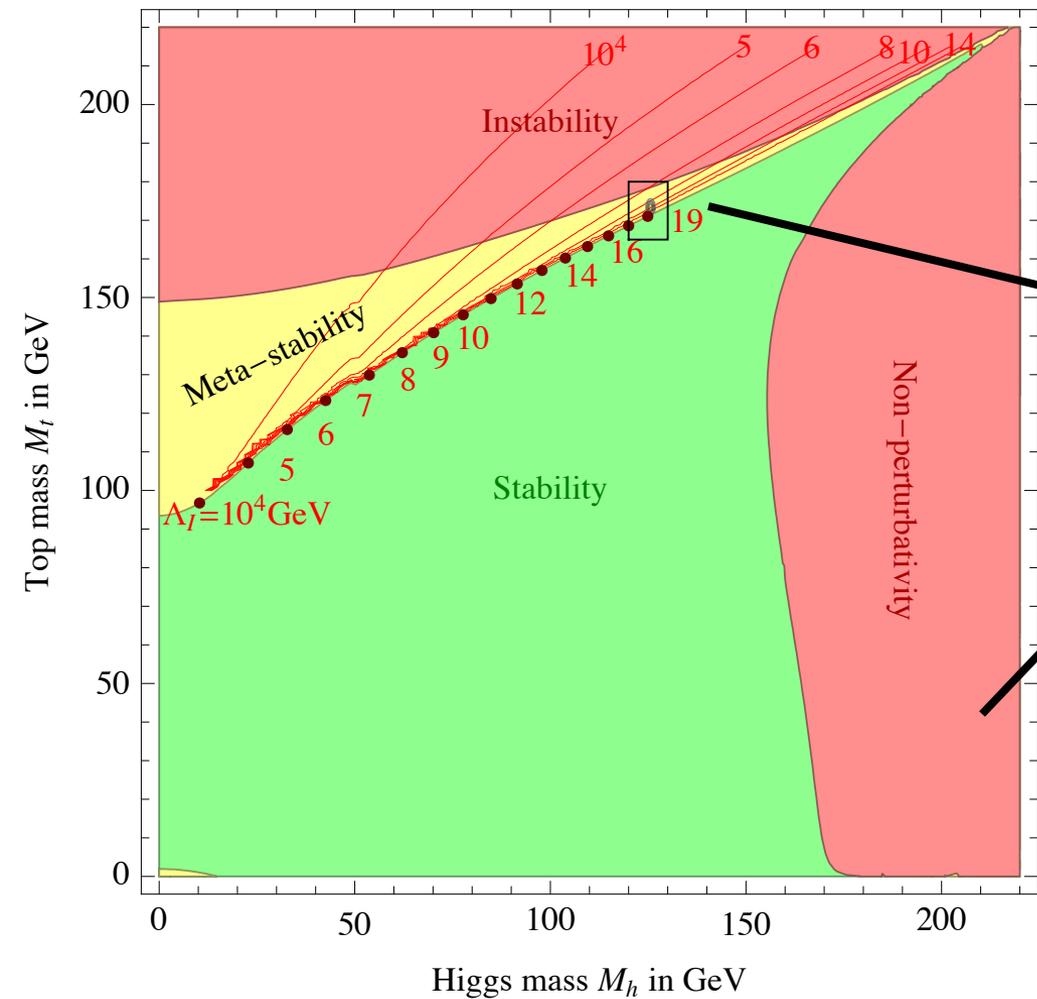


↑ Higgs mass

↓ Top quark mass

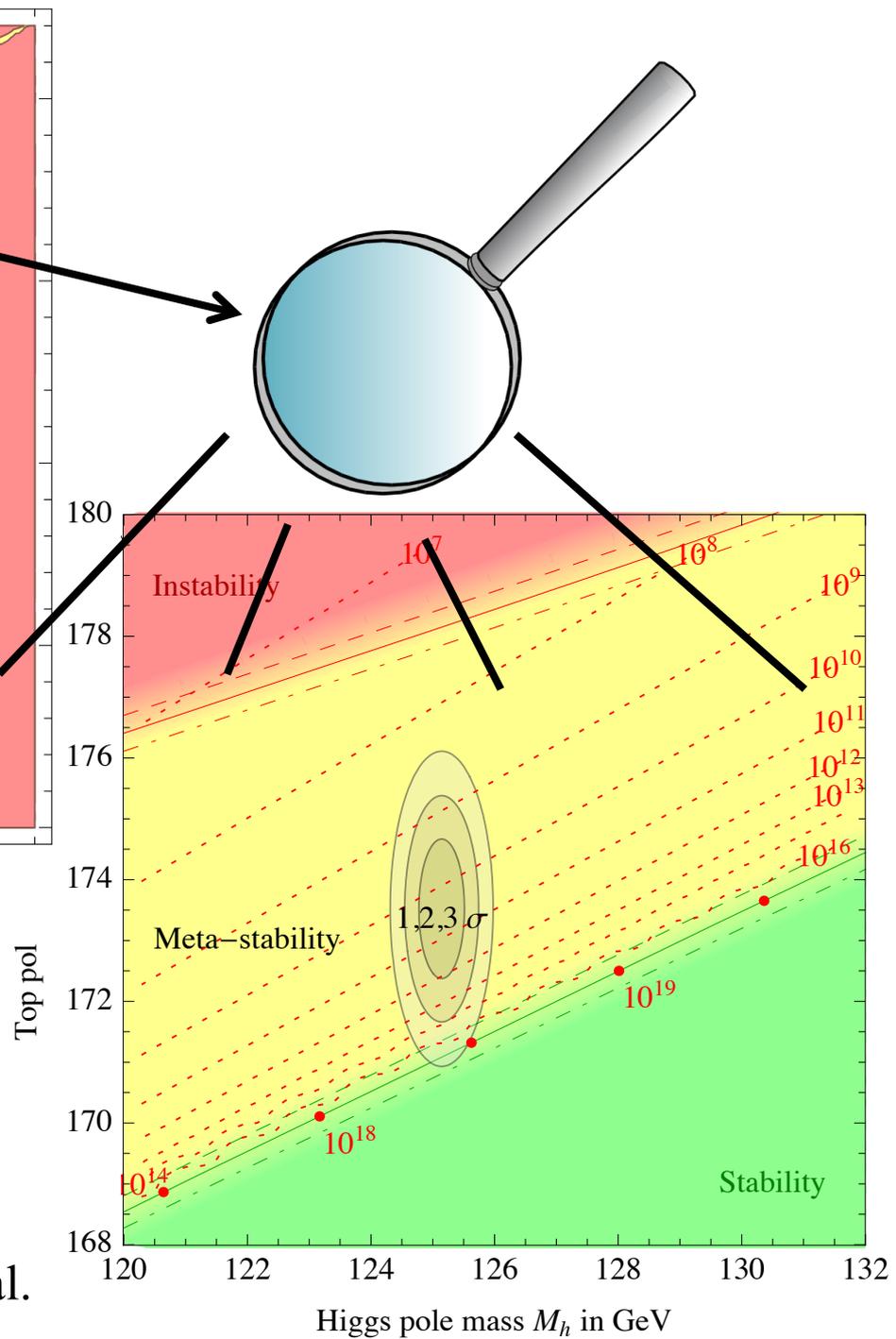
$$\frac{16\pi^2}{3} \mu \frac{d\lambda}{d\mu} = -2y_t^4 + 4y_t^2\lambda + 8\lambda^2 +$$





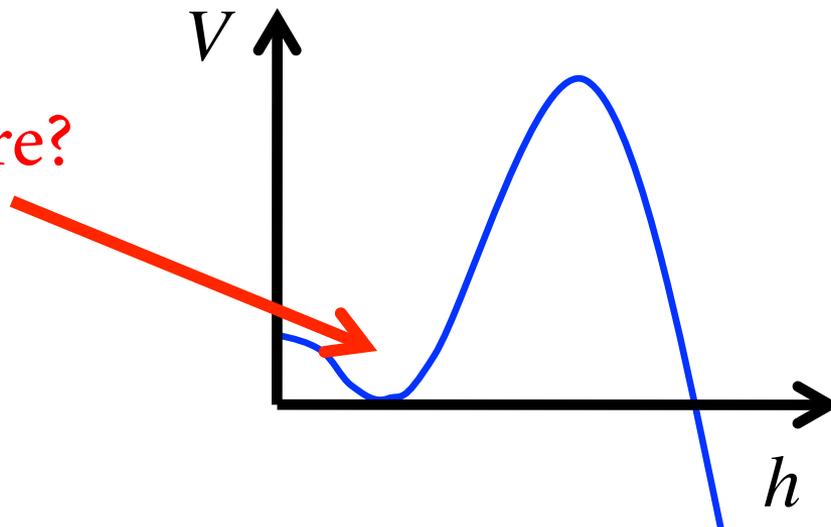
Critical sensitivity  
on SM parameters

Degrassi et al.



# What does metastability imply for the early universe?

1) How did we end up here?

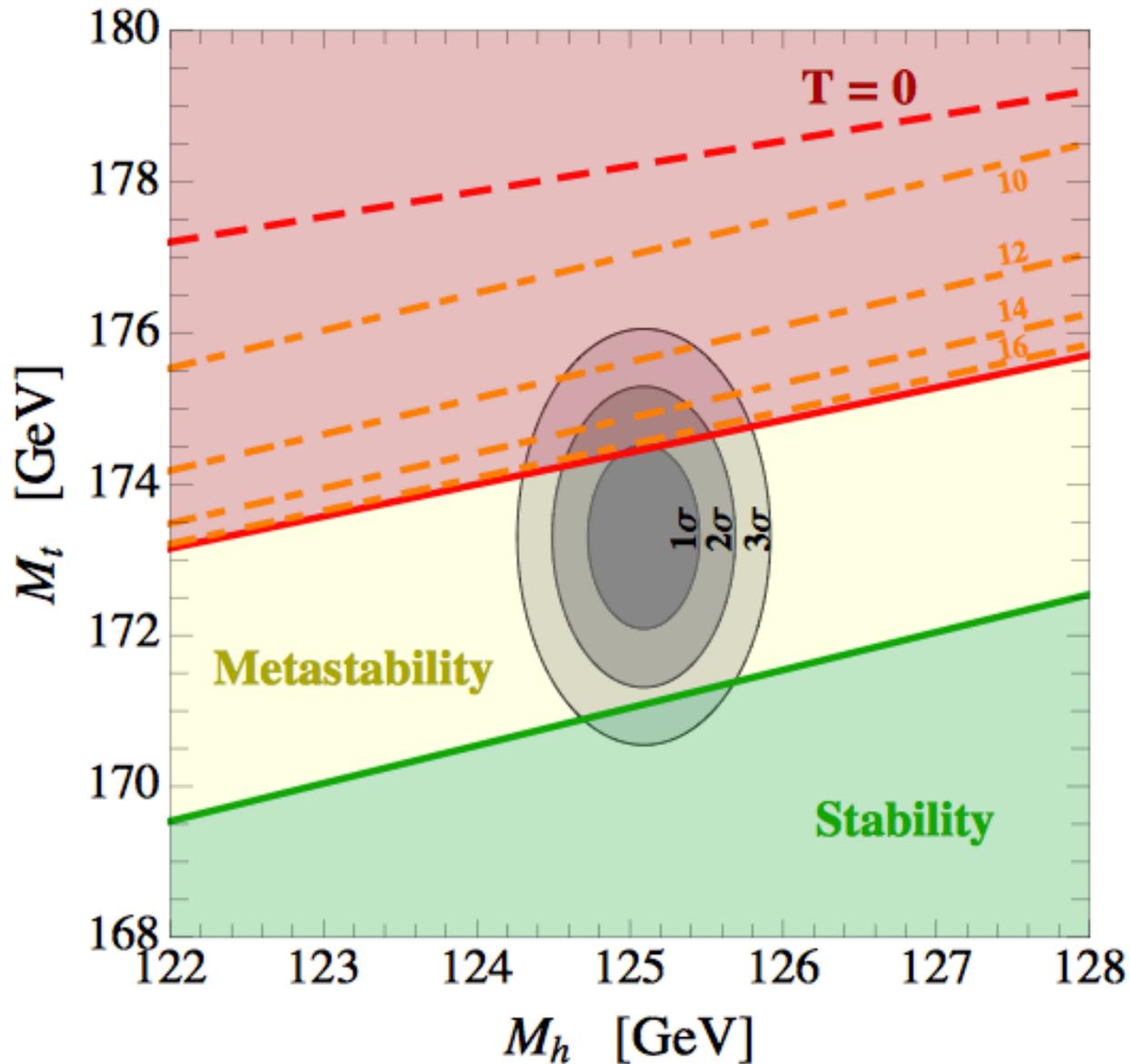


- Initial condition from pre-inflationary thermal phase
- Large  $Rh^2 \approx H^2 h^2$  stabilizing the origin
- Small coupling to inflaton (negligible Higgs contribution to energy density)

$$10^{-10} \leq \lambda_{h\phi} \leq 10^{-6}$$

Lebedev-Westphal

## 2) Why wasn't the vacuum destabilized by thermal effects?

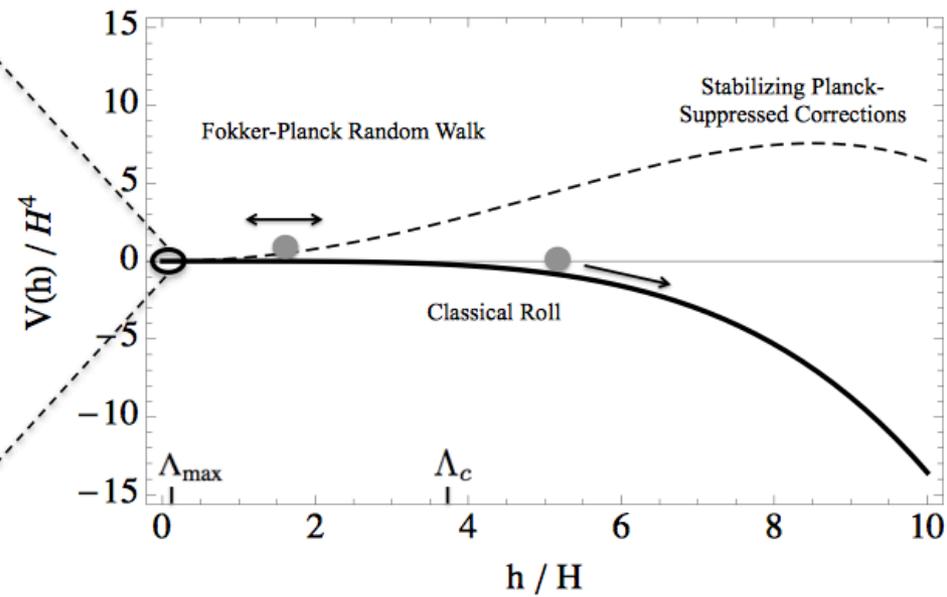
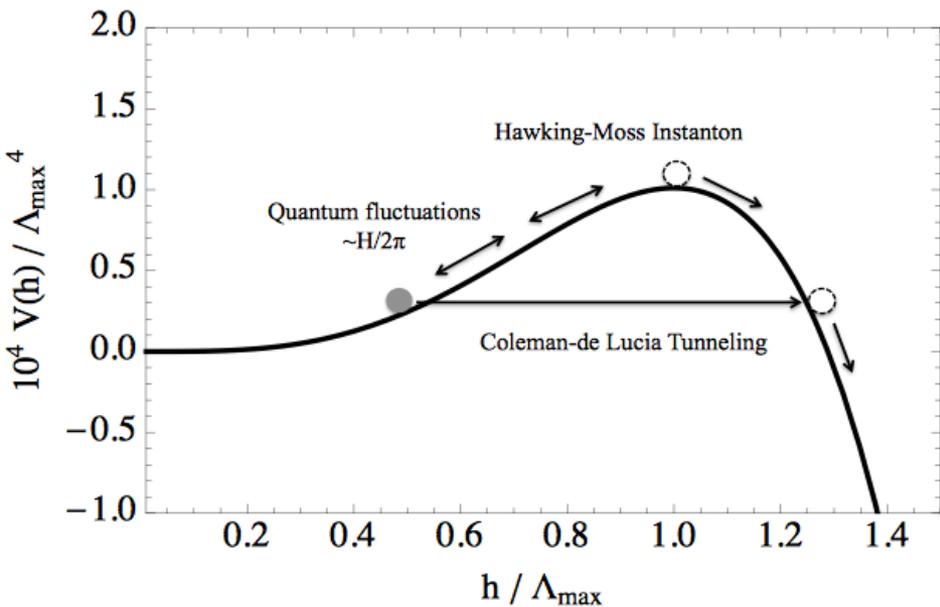


## 2) Why wasn't the vacuum destabilized during inflation?

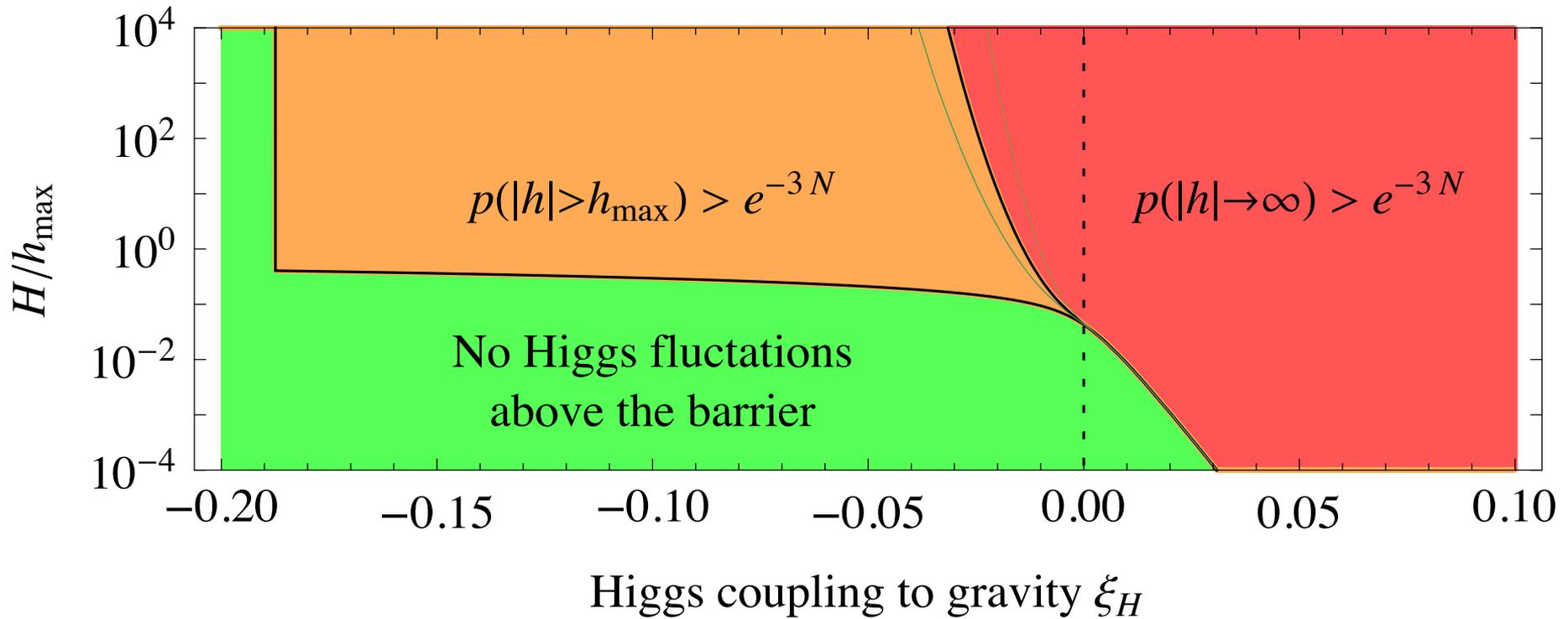
Espinosa-Giudice-Riotto

$$H \ll h_{\max}$$

$$H \gg h_{\max}$$



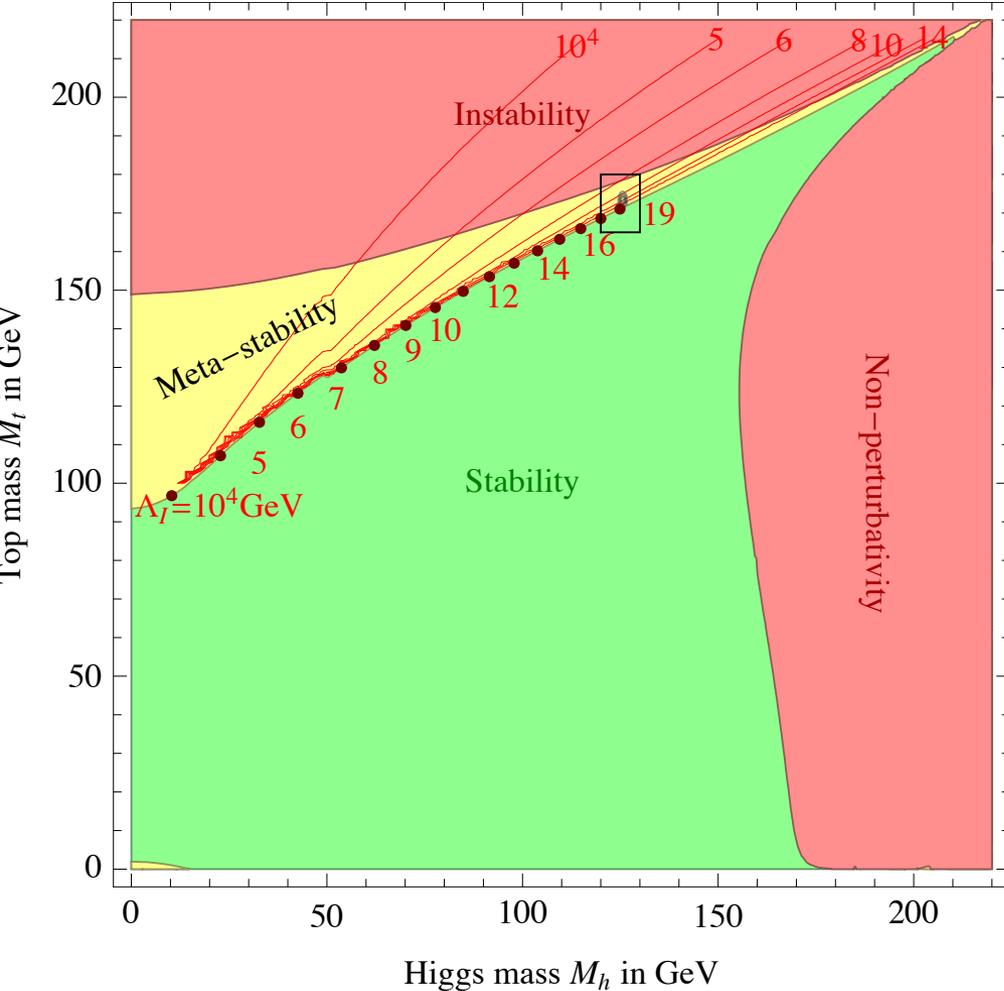
Hook et al.



Espinosa et al.

- Vacuum stability implies a constraint on Hubble during inflation
- Important if tensor modes are detected

$$H \approx 8 \times 10^{13} \text{ GeV } (r/0.1)^{1/2}$$

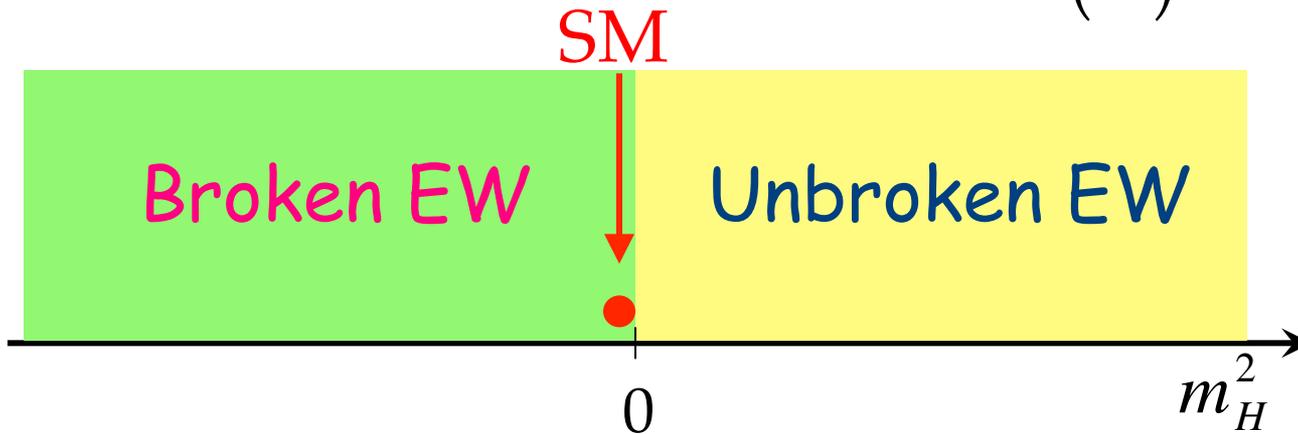


Is there a message in Higgs near-criticality?

- Coincidence?
- Dynamical attractor? Self-organized criticality?
- Statistical explanation in the multiverse?
- Cosmological evolution of fundamental parameters?

# Higgs naturalness as a “criticality” condition

$$V(H) = -m_H^2 |H|^2 + \lambda |H|^4$$

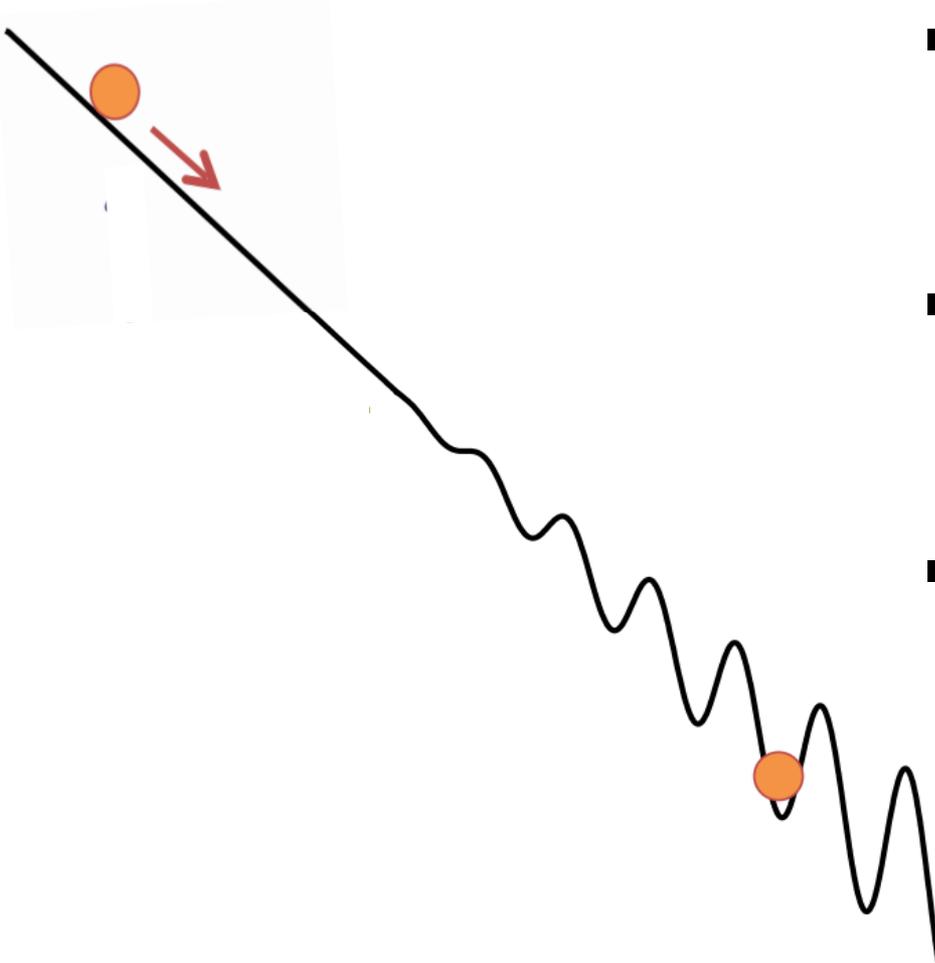


Why is nature so close to the critical line?

- Is  $m_H^2 \approx 0$  special because of symmetry? (supersymmetry, Goldstone symmetry)
- Is  $m_H^2 \approx 0$  special because of SOC?

# An example of Higgs naturalness explained by SOC

Graham-Kaplan-Rajendran

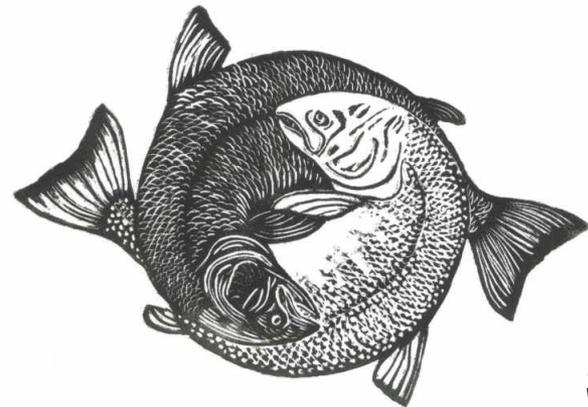
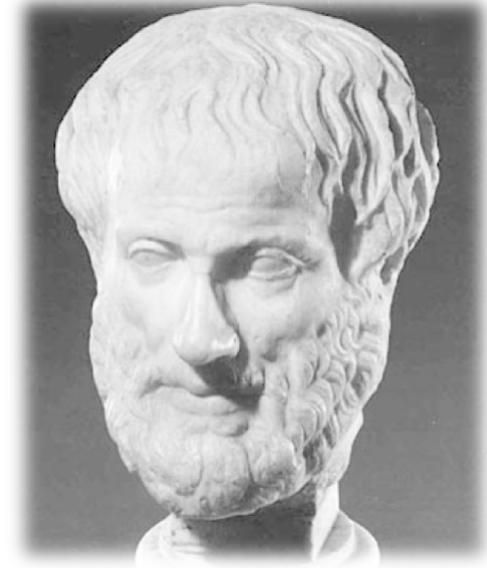


- The relaxion rolls down a shift-breaking potential
- The Higgs mass is scanned during the cosmological evolution
- Back-reaction from EW breaking

# We are facing a fundamental issue that will influence the future strategy of particle physics and early cosmology



The SM is an EFT emerging from a more fundamental theory at the TeV scale



Higgs naturalness and the CC indicate a breakdown of EFT intuition (SOC, multiverse, cosmological selection of parameters, ...)



# LHC: the future

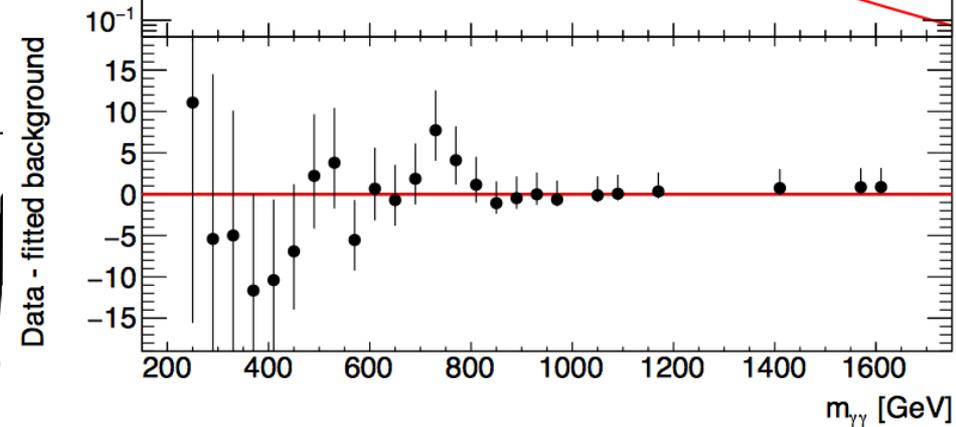
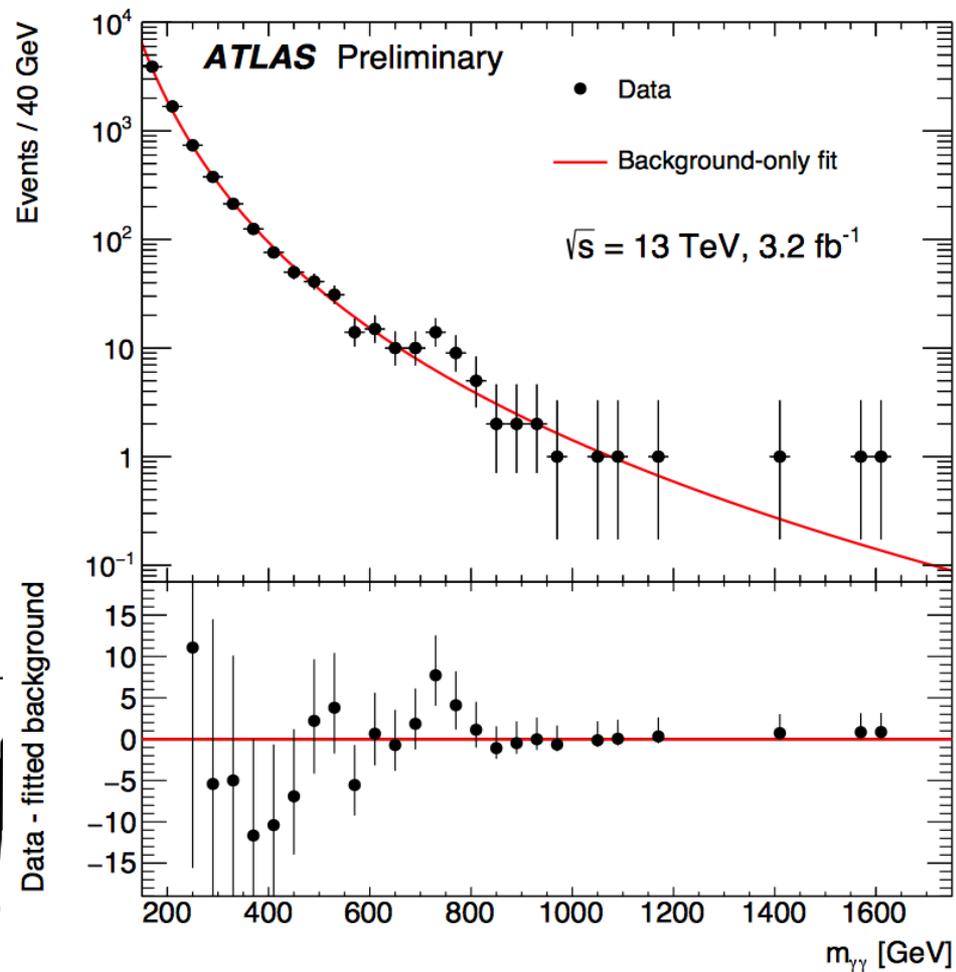
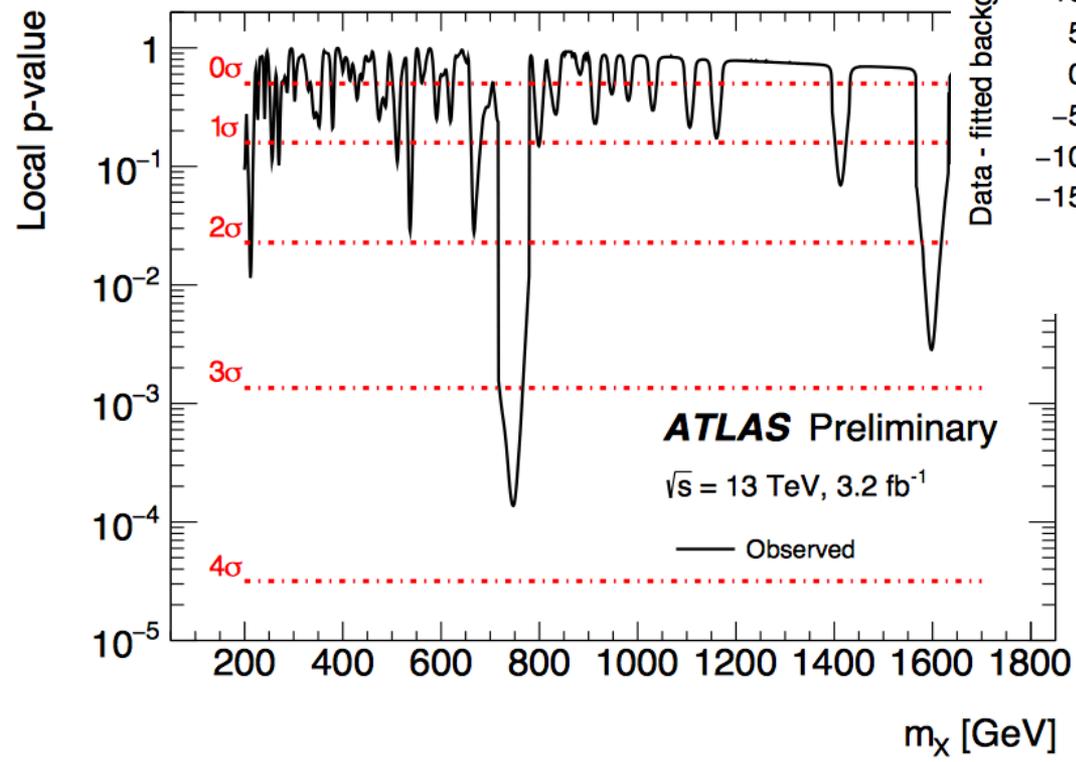
From Apr 2016 to Nov 2018  $\rightarrow 100 \text{ fb}^{-1}$  @ 13 TeV

LS2 until 2021

HL-LHC  $\rightarrow 3000 \text{ fb}^{-1}$

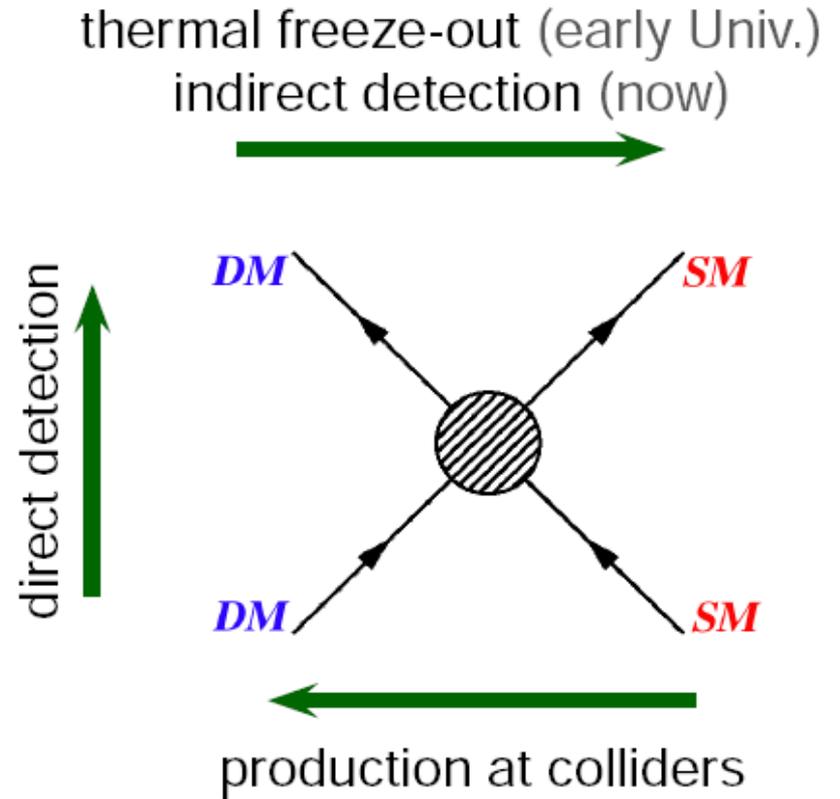
A lot to be learned for cosmology:

- Nature of the Higgs boson
- Higgs naturalness
- EW phase transition
- New physics?
- Dark matter?



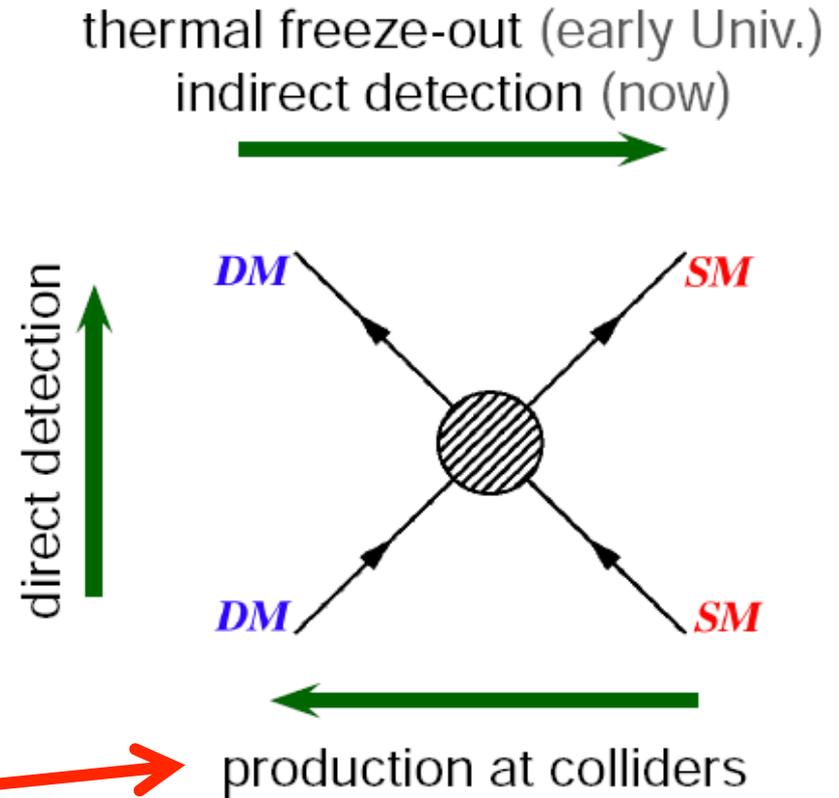
# DM at the LHC

The standard lore...



# DM at the LHC

The standard lore...



... is misleading



- Main signal comes from DM mediators
- EFT have limited use
- Searches are very model dependent

# Conclusions

- Particle physics gives the tools for developing cosmology
- This is why the LHC is a formidable source of information for cosmology
- Lessons from Higgs: near-criticality, EW phase transition
- Naturalness has been challenged by LHC8. The answers from LHC13 will be critical for the strategy to go BSM, hence for cosmology
- The discovery of new physics can revolutionize our understanding of the microworld
- New information on the nature of dark matter