

New NLO generator for Bhabha scattering

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Motivation

- Motivation (in brief):
 - No precision generator for polarized Bhabha scattering exists;
 - No NLO Bhabha generator allows to change initial momenta for every event;
 - Precision of BHWIDE and BHLUMI codes is unknown at TeV scale.

Energy spread simulation

Fixed energy mode (simplistic):

```

evt->momentum[0].Set(E_0, Px_0, Py_0, Pz_0);
evt->momentum[1].Set(E_1, Px_1, Py_1, Pz_1);
gen->Initialize(E_cms);
...
for(;;){ gen->GenEvent(evt); }

```

 10^4 evt/s

Variable energy mode (simplistic):

```

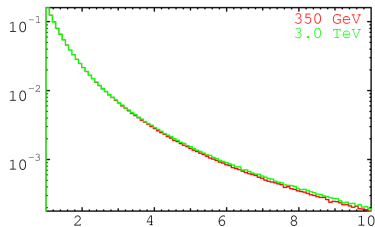
gen->Initialize(E_cms_Min, E_cms_Max);
...
for(;;) {
    evt->momentum[0].Set(E_0, Px_0, Py_0, Pz_0);
    evt->momentum[1].Set(E_1, Px_1, Py_1, Pz_1);
    gen->GenEvent(evt);
}

```

 10^3 evt/s

Energy spread simulation

Variable energy vs. Scaled events



- C.m.s. angular distribution at 350 GeV and 3 TeV;
- Error of event scaling may be up to 1-2%.

- It's reasonable to use fixed-energy generator for analysis.
- The variable-energy generator is useful for final precision run only.

Current status

- (Mostly) completed modules:
 - NLO Bhabha scattering generator code;
 - LO generators for $e^+e^- \rightarrow e^+e^-e^+e^-$ and $e^+e^- \rightarrow e^+e^-\mu^+\mu^-$ processes;
 - The generator seems stable up to $\Theta_e \sim 5mrad$.

- Current focus on:
 - Error budget estimation for polarized scattering;
 - LL YFS exponentiation method implementation for independent cross-check of polarized effects.

Error handling in BHWIDE and BHLUMI

- The physical uncertainties were estimated by comparison to results of independent calculations:
 - LL YFS exponentiation with precise phase space integration of hard photons emission;
 - precise NLO calculation;
 - semi-analytical precise N.NLO calculations (with no MC), ...
- The semi-realistic acceptance fits were used if MC is available.

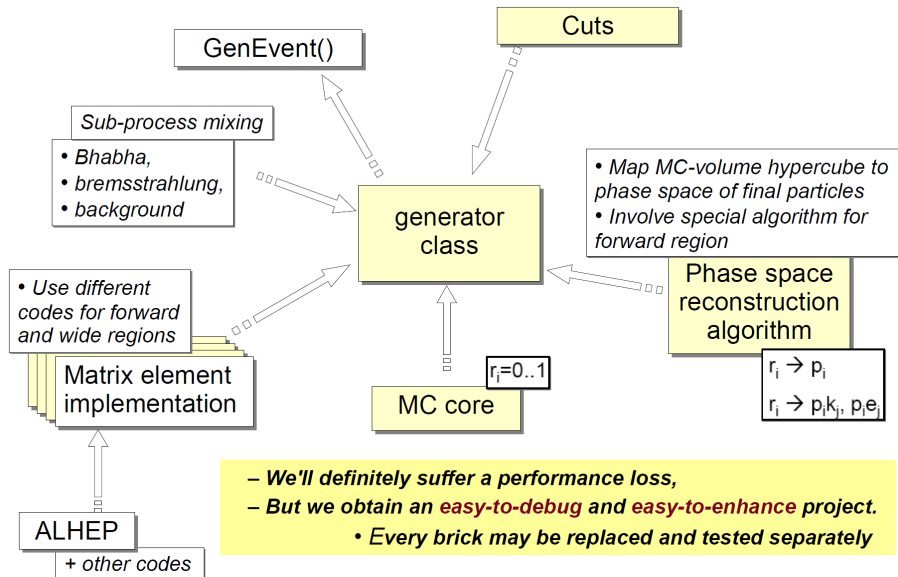
- **I.e. the independent calculation is required to estimate errors properly**

Error handling in new generator

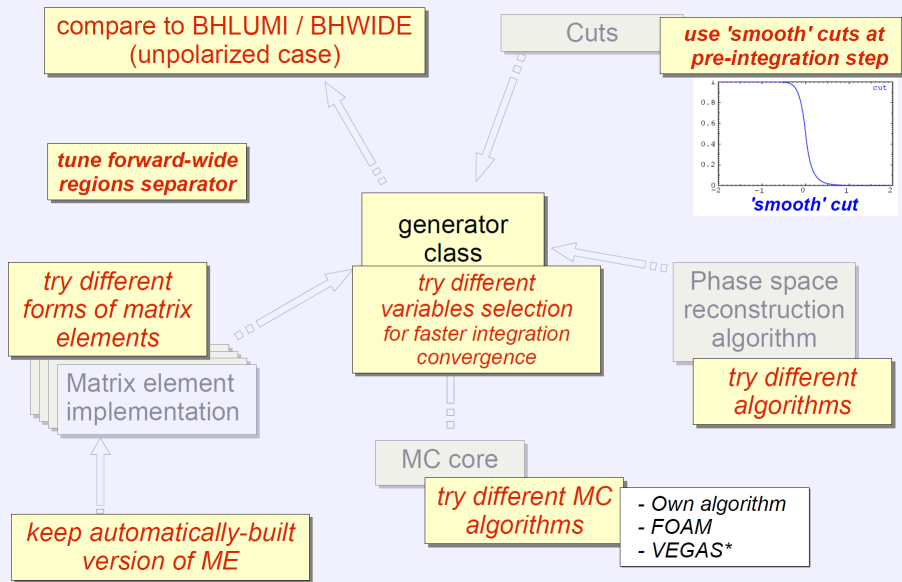
- Unpolarized cross section is compared to BHWIDE (and BHLUMI for small angles):
 - coincidence is within MC integration uncertainty 0.2–0.8% at TeV scale and $\Theta_e > 10\text{mrad}$;
 - better precision requires both higher order RC calculation and improved error handling.

- **There is no independent N.LO code to cross-check polarized effects**
 - We can't estimate the generator precision here!
 - The independent code for YFS-exponentiation will solve the problem.

Brick-based architecture



Cross-checks available

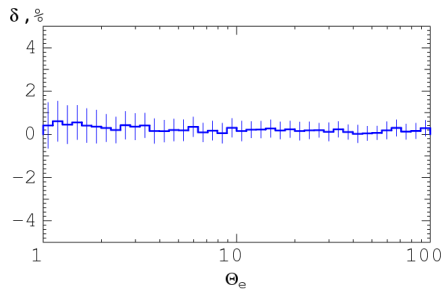
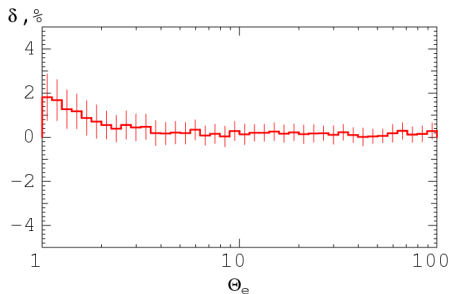


Conclusions

- The difference to BHLUMI of up to 0.8% is sufficient and must be examined.
 - The former requirement of 1%-precision must be improved.
- It seems useless to release the **polarized** NLO generator until its uncertainty will be properly estimated.
 - Current focus on YFS-exponentiation mode for the generator.
 - If no unexpected problem appear, the period of ~ 2 months seems reasonable to finalize the work.
- After the error handling is finalized (and a first version of code is released):
 - need to discuss required user interfaces (LCIO).
 - discuss if it's possible to publish the manual as CLICdp-Note?

Back-up slides

Systematic error example



- left (very early runs): systematic error observed,
- right: difference to BHWIDE within the Monte-Carlo error.

Code performance

	BHWIDE	New generator
initialization, s	~ 10	$\sim 10^3 \dots 10^4$ ~ 10 (using pre-created files)
event rate, s^{-1}	$10^3 \dots 10^4$ unweighted or weighted	10^4 unweighted only
flowed energy event rate, s^{-1}	not available	10^3
memory used	insignificant	up to 0.5 GB

- The huge initialization time is the price for Monte-Carlo algorithm used,
 - both initialization time and required memory will be reduced by the release.

Release announcement (deprecated)

- The generator will be published at <http://cern.ch/~makarenko/bhabha>

The first release will include:

- NLO precision code;
 - user manual with discussion of the methods used;
 - cross-check table to BHWIDE results (for unpolarized scattering);
 - standard output for polarized runs (for further cross-checks).
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- most of the technical problems seems solved;
 - need certain time for:
 - optimization (mostly to reduce the memory used),
 - manual text preparation.