# The Hunt for Heavy Neutrinos at the FCC

Alain Blondel University of Geneva with many thanks to S. Ganjour, M. Mitra, S. Pascoli, E. Graverini, P. Mermod, N. Serra, M. Shaposhnikov

30/2015

## Where is Everybody?











## At higher masses -- or at smaller couplings?





THE STANDARD MODEL IS COMPLETE .....

## But at least 3 pieces are still missing



neutrinos have mass... and this very probably implies new degrees of freedom → Right-Handed, Almost «Sterile» (very small couplings) Neutrinos completely unknown masses (meV to ZeV), nearly impossile to find. .... but could perhaps explain all: DM, BAU,v-masses



### some REFERENCES

	PHYSICAL REVIEW D	VOLUME 29, N	UMBER 11	1 JUNE 1984		
		Extending limits on ne	utral heavy leptons	PUBLISHED FOR SISSA BY 2 S RECEIVE: Receive: Rec	PRINGER 23, 2013 25, 2013	
		Michael Gronau <sup>*</sup> Department of Physics, Syracuse University, Syracuse, New York 132.		PUBLISHED: January 29, 2014 First look at the physics case of TLEP		
	FLAVOUR(267104)-ERC-23	TUM-HEP 850/12 SISSA 25/2012	/EP CFTP/12-013	arxiv:1308.61	76	
	arxiv:1208.3654		The TLEP Design Study Working Group M. Bicer, <sup>a</sup> H. Duran Yildiz, <sup>b</sup> I. Yildiz, <sup>c</sup> G. Coignet, <sup>d</sup> M. Delmastro, <sup>d</sup> T. Alexopoulos, <sup>c</sup> C. Grojean, <sup>f</sup> S. Antusch, <sup>b</sup> T. Sen, <sup>h</sup> HJ. He, <sup>f</sup> K. Potamianos, <sup>f</sup> S. Haug, <sup>b</sup> A. Moreno, <sup>f</sup> A. Heiser, <sup>m</sup> V. Sanz, <sup>a</sup> G. Gomez-Ceballos, <sup>o</sup> M. Klute, <sup>a</sup> M. Zanetti, <sup>a</sup> LT. Wang, <sup>p</sup> M. Dam, <sup>g</sup> C. Boehm, <sup>r</sup> N. Glover, <sup>r</sup> F. Krauss, <sup>r</sup> A. Lenz, <sup>r</sup> M. Syphers, <sup>*</sup>			
	Higgs Decays in the Low Scale Type I See-Saw Model C. Garcia Cely <sup>a</sup> ), A. Ibar					
				-ee CERN-F	PE/96–195 December 1996	
	theories of the	A. Blondel (presenter)*, E. Graverini <sup>b</sup> , N. Serra <sup>b</sup> , M. Shaposhnikov			ons	
	and mixings with <sup>k</sup> DPNC, University of Geneva, Qual Ansermet 24, CH-1205 Geneva, Switzgrlan <sup>k</sup> Physik Institu, University of Zurich, CH-8057 Zurich, Switzgrland <sup>c</sup> UTDP EVEL CH-105 Learners Switzgrland			Neutral Heavy Dept	0115	
	The Role of Sterile Newson and the second states and the second st			uced in Z Decays		
L	Astrophysics			DELPHI Collaboration		
	Alexey Boyarsky*†, Oleg F	tuchayskiy <sup>‡</sup> and Mikhail Shaposhni	CC design study a	and FCC-ee <u>http://cern.ch/</u>	/ <mark>fcc-ee</mark>	
The $\nu$ MSM, Dark Matter and Neutrino Masses http://indico.cern.ch/category/5684/						
Takehiko Asaka, Steve Blanchet, and Mikhail Shaposhnikov						
Preprint typeset in JHEP style - HYPER VERSION FERMILAB-PUB-08-086-T, NSF-KITP-08-54, MADPH-06-1466, DCPT/07/198, IPPP/07/99					7/99	
ł	Phys.Lett.B631:151-156,20	005 CH-1015 Lausanne, Switzerland				
2	arXiv:hep-ph/0503065	2005)	The Search for He	avy Maiorana Neutrinos		
					50000	
	30/01/2015 Alain	Blondel Right handed ne	Anupama Atre <sup>1,2</sup> , Tao Han <sup>2,3</sup>	<sup>8,4</sup> , Silvia Pascoli <sup>5</sup> , Bin Zhang <sup>4</sup> *		



#### **Electroweak eigenstates**





Adding masses to the Standard model neutrino 'simply' by adding a Dirac mass term (Yukawa coupling)

$$m_D v_L v_R$$
,  $m_D \overline{v_L} v_R$ 

implies adding a right-handed neutrino (new particle)

**<u>No SM symmetry</u>** prevents adding then a term like

 $m_M \overline{v_R}^c v_R$ 



 $\nabla_{\mathbf{R}}$ 

and this simply means that a neutrino turns into a antineutrino (the charge conjugate of a right handed antineutrino is a left handed neutrino!)

It is perfectly conceivable ('natural'?) that both terms are present -> 'see-saw'

B. Kayser, the physics of massive neutrinos (1989)





#### **Mass eigenstates**

See-saw in a general way :  $\mathcal{L} = \frac{1}{2} (\bar{\nu}_L, \, \bar{N}_R^c) \begin{pmatrix} 0 & m_D \\ m_D^T & M_R \end{pmatrix} \begin{pmatrix} \nu_L^c \\ N_R \end{pmatrix}$  $M_R \neq 0$  $m_{D} \neq 0$ **Dirac + M**ajorana mass terms  $\tan 2\theta = \frac{2m_D}{M_D - 0}$  $\ll 1$  $m_{\nu} = \frac{1}{2} \begin{bmatrix} (0 + M_R) - \sqrt{(0 - M_R)^2 + 4 m_D^2} \end{bmatrix} \simeq -m_D^2 / M_R$  $M = \frac{1}{2} \begin{bmatrix} (0 + M_R) + \sqrt{(0 - M_R)^2 + 4 m_D^2} \end{bmatrix} \simeq M_R$ general formula if  $m_D \ll M_R$  $M_R \neq 0$ see-saw  $M_R = 0$  $M_{R} \neq 0$  $m_{\rm D} \neq 0$  $m_D \neq 0$  $m_D = 0$ <u>Dirac + Majorana</u> m Dirac only, (like e- vs e+): <u>Majorana only</u>  $\begin{array}{cccc} v_{\mathbf{L}} & v_{\mathbf{R}} & \overline{v}_{\mathbf{R}} & \overline{v}_{\mathbf{L}} \\ 0 & \frac{1}{2} & 0 \end{array}$ m m  $v_{L}$  $\bar{\nu}_{R}$  $\begin{array}{cccc} \nu_L & N_R & \bar{\nu}_R & N_L \\ I_{weak} = \begin{array}{ccc} \frac{1}{2} & 0 & \frac{1}{2} & 0 \end{array}$ 1/2 I<sub>weak</sub>= 4 states of equal masses 2 states of equal masses 4 states, 2 mass levels Some have I=1/2 (active) All have I=1/2 (active)  $m_1$  have ~I=1/2 (~active) Some have I=0 (sterile) m<sub>2</sub> have ~I=0 (~sterile)

ENSIS

There even exists a scenario that claims to explain everything: the vMSM

can generate Baryon Asymmetry of Universe

if m<sub>N2.N3</sub> > 140 MeV

Shaposhnikov et al

TeV

GeV

MeV

keV



 $N_{2}, N_{3}$ 







## **Manifestations of right handed neutrinos**

one family see-saw :  

$$\theta \approx (m_D/M)$$
  
 $m_v \approx \frac{m_D^2}{M}$   
 $m_N \approx M$   
 $|\mathbf{U}|^2 \propto \theta^2 \approx m_v / m_N$ 

 $v = vL\cos\theta - N^c_R\sin\theta$  $N = N_R\cos\theta + v_L^c\sin\theta$ 

what is produced in W, Z decays is:  $v_L = v \cos\theta + N \sin\theta$  v = light mass eigenstate N = heavy mass eigenstate  $\neq v_L$ , active neutrino which couples to weak inter. and  $\neq N_R$ , which does'nt.

can be larger with 3 families

- -- mixing with active neutrinos leads to various observable consequences
  - -- if very light (eV) , possible effect on neutrino oscillations (short baseline)
  - -- if in keV region (dark matter), monochromatic photons from galaxies with  $E=m_N/2$
- -- possibly measurable effects at High Energy
  - If N is heavy it will decay in the detector (not invisible)
  - PMNS matrix unitarity violation and deficit in Z «invisible» width
  - $\rightarrow$  Higgs, W, Z exotic decays  $H \rightarrow v_i \overline{N}_i$  and  $Z \rightarrow v_i \overline{N}_i$ ,  $W \rightarrow I_i \overline{N}_i$
  - $\rightarrow$  also in charm and b decays via W<sup>\*</sup>-> I<sub>i</sub>  $\overline{N}_i$
  - $\clubsuit$  violation of unitarity and lepton universality in Z, W or  $\tau\,$  decays
  - -- etc... etc...

-- Couplings are small  $(m_v / m_N)$  (but who knows?) and generally out of reach of hadron colliders (but this deserves to be revisited for detached vertices @LHC, HL-LHC, FCC-hh)



**Indirect effects** 

- -- neutrino Majorana mass term can lead to lepton number violating processes by virtual neutrino exchange and to flavour violation
  - -- neutrinoless double beta decay (the most powerful one)
  - -- FCNC ( $\mu \rightarrow e\gamma$ ) etc...
  - -- at a Z factory :  $Z \rightarrow \tau \mu$   $Z \rightarrow \tau e$  Z->  $\tau \tau$ ,  $\tau \rightarrow \mu \gamma$   $\tau \rightarrow e \gamma etc...$





#### Indirect constraints from lepton flavour violation are weak

arxiv:1208.3654

#### Higgs Decays in the Low Scale Type I See-Saw Model

C. Garcia Cely<sup>a)</sup>, A. Ibarra<sup>a)</sup>, E. Molinaro<sup>b)</sup> and S. T.  $Petcov^{c,d)}$ <sup>1</sup>





## Goal performance of e+ e- colliders



(possibly even 10<sup>13</sup> with crab-waist) (few years)



complementarity

NB: ideas for lumi upgrades:

ILC arxiv:1308.3726 (not in TDR. Upgrade at 250GeV by reconfiguration after 500 GeV running; under discussion) FCC-ee (crab waist)





- -- invisible widths
- -- FCNC
- -- direct search of heavy neutrino decays





Phys.Rept.427:257-454,2006

 $N_v = 2.984 \pm 0.008$ 

This is determined from the Z line shape scan and dominated by the measurement of the hadronic cross-section at the Z peak maximum →

The dominant systematic error is the theoretical uncertainty on the Bhabha cross-section (0.06%) which represents an error of  $\pm 0.0046$  on N<sub>v</sub>



Improving on  $N_{\nu}$  by more than a factor 2 would require a large effort to improve on the Bhabha cross-section calculation!



#### NEUTRINO COUNTING AT THE Z-PEAK AND RIGHT-HANDED NEUTRINOS

#### C. JARLSKOG

CERN, CH-1211 Geneva 23, Switzerland and Department of Physics, University of Stockholm, S-113 46 Stockholm, Sweden

Received 20 February 1990

We consider the implications of extending the minimal standard model, with *n* families of quarks and leptons, by introducing an arbitrary number of right-handed neutrinos, for neutrino-counting via the "invisible width" of the Z. It is shown that the effective number of neutrinos,  $\langle n \rangle$ , satisfies, the inequality  $\langle n \rangle \leq n$ , where  $\langle n \rangle$  is defined by  $\Gamma(Z \rightarrow neutrinos) \equiv \langle n \rangle \Gamma_0$  and  $\Gamma_0$ is the standard width for one massless neutrino. Thus, in the case of three families, the neutrino-counting can give a result which is less than three, if there are right-handed neutrinos.

### Theorem.

In the standard model, with *n* left-handed lepton doublets and N-n right-handed neutrinos, the effective number of neutrinos,  $\langle n \rangle$ , defined by

 $\Gamma(\mathbf{Z} \rightarrow \text{neutrinos}) \equiv \langle n \rangle \Gamma_0$ ,

where  $\Gamma_0$  is the standard width for one massless neutrino, satisfies the inequality

Alain Blond  $\langle n \rangle \leq n$ .



(15)

FCC Neutrino counting at TLEP



given the very high luminosity, the following measurement can be performed

$$N_{v} = \frac{\frac{\gamma Z(inv)}{\gamma Z \to ee, \mu\mu}}{\frac{\Gamma_{v}}{\Gamma e, \mu} (SM)}$$

The common  $\gamma$  tag allows cancellation of systematics due to photon selection, luminosity etc. The others are extremely well known due to the availability of O(10<sup>12</sup>) Z decays.

The full sensitivity to the number of neutrinos is restored, and the theory uncertainty on  $\frac{\Gamma_{\nu}}{\Gamma e}$  (*SM*) is very very small.

A good measurement can be made from the data accumulated at the WW threshold where  $\sigma$  ( $\gamma$  Z(inv)) ~4 pb for  $|\cos\theta_{\gamma}| < 0.95$ 

A better point may be 105 GeV (20pb and higher luminosity) may allow  $\Delta N_{\nu}$ =0.0004? Alain Blondel Right handed neutrinos at the FCC Aspen 2015





### Lepton flavour violating Z decays with 10<sup>13</sup> Z decays

A. Abada et al, <u>arXiv:1412.6322</u> Indirect searches for sterile neutrinos at a highluminosity Z-factory <u>A. Abada</u>, <u>V. De Romeri</u>, <u>S. Monteil</u>, <u>J. Orloff</u>, <u>A. M. Teixeira</u>



Figure 4: The "3+1 model": on the left  $(\sin^2 \theta_{14}, m_4)$  parameter space of the sterile state, displaying the regimes for BR $(Z \to e\mu)$  for a NH light neutrino spectrum. Line and colour code as in Fig. 3 (dark green points are associated with  $10^{-13} \leq \text{BR}(Z \to e\mu) \leq 10^{-9}$ , while light green ones correspond to BR $(Z \to e\mu) \geq 10^{-9}$ ). On the right,  $(\sin^2 \theta_{34}, m_4)$  displaying with the same colour code the corresponding regimes for BR $(Z \to \mu\tau)$ .





## **RHASnu's production in Z decays**

**Production:** 

$$BR \ (\mathbf{Z}^{0} \to \nu_{m} \overline{\nu}) = BR \ (\mathbf{Z}^{0} \to \nu \overline{\nu}) \ |U|^{2} \ \left(1 - \frac{m_{\nu_{m}}^{2}}{m_{\mathbf{Z}^{0}}^{2}}\right)^{2} \left(1 + \frac{1}{2} \frac{m_{\nu_{m}}^{2}}{m_{\mathbf{Z}^{0}}^{2}}\right)$$

multiply by 2 for anti neutrino and add contributions of 3 neutrino species (with different  $|U|^2$ )





21







Search for heavy neutral leptons

search e<sup>+</sup> e<sup>-</sup>  $\rightarrow$  v N N $\rightarrow$  v( $\gamma/Z$ )<sup>\*</sup>  $\rightarrow$  monojet

> Find: one event in 4x10<sup>6</sup>Z:











A.B, Elena Graverini, Nicola Serra, Misha Shaposhnikov







TLEP expected sensitivity to HNL (NH)











## SHIP



NB very large detector caverns for FCC-hh may allow very large FCC-ee detector (R=15m?) leading to improved reach at lower masses.





# -- also search for same sign muons or electrons at the LHC (e.g. CMS: )





ATLAS arXiv:1203.5420. CMS arXiv:1207.6079.

limits at  $|U|^2 \sim 10^{-2-5}$  level

7 GeV run papers on 8 GeV run in preparation





**Preliminary projection for LHC (P. Mermod, very preliminary)** 







We have seen that the Z factory offers a clean method for detection of Heavy Right-Handed neutrinos Ws are less abundant at the lepton colliders

At the 100 TeV hadron machine the W is the dominant particle.



There is a lot of /pile-up/backgrounds/lifetime/trigger issues which need to be investigated.

- BUT.... in the regime of long lived HNLs the simultaneous presence of
- -- the initial lepton from W decays
- -- the detached vertex with kinematically constrained decay allows for a significant background reduction.

But it allows also a characterization **both in flavour and charge** of the produced neutrino, thus information of the flavour sensitive mixing angles and a test of the fermion violating nature of the intermediate (Majorana) particle.

**VERY** interesting...





## Conclusions

- 1. H, Z factory will investigate invisible widths with exp precision of 0.15% and 0.0004  $\Gamma_{\rm v}$
- 2. The quest for the

«Right-Handed-Almost-Sterile-See-saw-partners neutrinos» (dextrinos? RHASnus? Heavy Neutral Leptons? Shaposhninos? heavinos?)

### is not desperate at all

for Majorana mass at weak scale. In particular it may lead to spectacular 'detached vertex' signatures in a beam dump experiment (SHIP) or in Z-> neutrino decays at a Tera-Z factory like FCC-ee,

**Observation in W decays at the FCC-hh would be very interesting...** 





**SPARES** 





 $h \rightarrow \nu N$  allowed if N is light enough Lepton universaility constraints must be met  $heta_{
u_e-N} < 0.05, \quad heta_{
u_\mu-N} < 0.05, \quad heta_{
u_\tau-N} < 0.08$ 



Chen, He, Tandean, Tsai (2011) Alain Blondel Right handed neutrinos at the FCC Aspen 2015





or





determine the number of neutrinos from the radiative returns

e+e- $\rightarrow \gamma Z (\rightarrow v \bar{v})$ 



```
CERN-TH.5528/89
```

NEUTRINO COUNTING

```
G. Barbiellini<sup>1</sup>, X. Berdugo<sup>2</sup>, G. Bonvicini<sup>3</sup>, P. Colas<sup>4</sup>, L. Mirabito<sup>4</sup>,
C. Dionisi<sup>5</sup>, D. Karlen<sup>6</sup>, F. Linde<sup>7</sup>, C. Luci<sup>8</sup>, C. Mana<sup>8</sup>, C. Matteuzzi<sup>9</sup>,
O. Nicrosini<sup>10</sup>, R. Ragazzon<sup>1</sup>, D. Schaile<sup>11</sup>, F. Scuri<sup>1</sup> and L. Trentadue*),<sup>12</sup>
```

in its original form (Karlen) the method only counts the 'single photon' events and is actually less sensitive than claimed. It has poorer statistics and requires running ~10 GeV above the Z pole. Systematics on photon selection are not small.

present result:  $N_v = 2.92 \pm 0.05$ 



#### NEUTRINO COUNTING AT THE Z-PEAK AND RIGHT-HANDED NEUTRINOS

#### C. JARLSKOG

CERN, CH-1211 Geneva 23, Switzerland and Department of Physics, University of Stockholm, S-113 46 Stockholm, Sweden

Received 20 February 1990

We consider the implications of extending the minimal standard model, with *n* families of quarks and leptons, by introducing an arbitrary number of right-handed neutrinos, for neutrino-counting via the "invisible width" of the Z. It is shown that the effective number of neutrinos,  $\langle n \rangle$ , satisfies, the inequality  $\langle n \rangle \leq n$ , where  $\langle n \rangle$  is defined by  $\Gamma(Z \rightarrow neutrinos) \equiv \langle n \rangle \Gamma_0$  and  $\Gamma_0$ is the standard width for one massless neutrino. Thus, in the case of three families, the neutrino-counting can give a result which is less than three, if there are right-handed neutrinos.

### Theorem.

In the standard model, with *n* left-handed lepton doublets and N-n right-handed neutrinos, the effective number of neutrinos,  $\langle n \rangle$ , defined by

 $\Gamma(\mathbf{Z} \rightarrow \text{neutrinos}) \equiv \langle n \rangle \Gamma_0$ ,

where  $\Gamma_0$  is the standard width for one massless neutrino, satisfies the inequality

Alain Blond  $\langle n \rangle \leq n$ .



(15)