

# Search for Hidden Particles



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# Particle physics: now what's the matter?



- Higgs @ 125 GeV in agreement with the **Standard Model**
- Couplings as predicted by **SM**
- **No** direct observation of **BSM** particles
- No significant deviations from **SM** in Flavour Physics

But...

- NP @ TeV scale needed to recover **naturalness**
- $m_H, m_t$  close to **stability** bound
- Hierarchy
- **Baryogenesis** (matter-antimatter asym.) not explained by CKM
- **Neutrino masses** and oscillations not explained by SM
- What is **dark matter**?



i.e. How to fix most SM problems w/o introducing new physic principles [2]

Leptons and quarks get their mass through the Yukawa interaction:

$$\mathcal{L} = m \psi_L^\dagger h \psi_R + c.c.$$

Three Generations of Matter (Fermions) spin 1/2

	I	II	III	
mass - charge - name	2.4 MeV 1/3 <b>u</b> Left up Right	1.27 GeV 2/3 <b>c</b> Left charm Right	173.2 GeV 2/3 <b>t</b> Left top Right	0 0 <b>g</b> gluon
Quarks	4.8 MeV -1/3 <b>d</b> Left down Right	104 MeV -1/3 <b>s</b> Left strange Right	4.2 GeV -1/3 <b>b</b> Left bottom Right	0 0 $\gamma$ photon
	0 0 $\nu_e$ electron neutrino	0 0 $\nu_\mu$ muon neutrino	0 0 $\nu_\tau$ tau neutrino	91.2 GeV 0 <b>Z<sup>0</sup></b> weak force
Leptons	0.511 MeV -1 <b>e</b> Left electron Right	105.7 MeV -1 $\mu$ Left muon Right	1.777 GeV -1 $\tau$ Left tau Right	126 GeV 0 <b>H</b> Higgs boson spin 0
				80.4 GeV -1 <b>W<sup>±</sup></b> weak force

Bosons (Forces) spin 1

Neutrinos need a right-handed partner to get their masses like the other SM fermions!



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Bosons (Forces) spin 1

Neutrinos need a right-handed partner to get their masses like the other SM fermions!

Three extra  $N$  field could exist (*sterile neutrinos*). They would be  $SU(2)$  singlets, as they are still unobserved.

**Sterile neutrinos** would mix to  $\nu_{e,\mu,\tau}$  with very small couplings  $U_{e,\mu,\tau}^2$ .

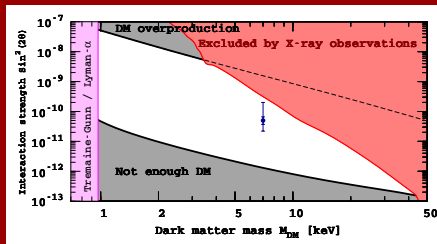


i.e. How to fix most SM problems w/o introducing new physic principles [2]

Suitable values of  $m_N$  and  $U_f^2$  [2] allow to simultaneously explain:

- $\nu$  oscillation: two massive states  $N_2, N_3$  give mass to active  $\nu$ s through the Seesaw mechanism
- dark matter:  $N_1$ , likely to have mass in the keV region, could have a lifetime greater than the age of the Universe
- matter-antimatter asymmetry: baryogenesis induced by leptogenesis if  $N$  has a Majorana mass term [10].

Two recent observations provide hints of the existence of a 7 keV dark matter ( $N_1$ ?) candidate [8, 9].



# A new dedicated experiment

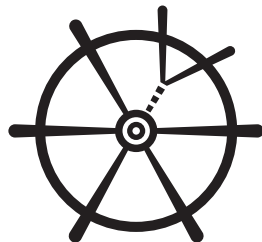
## Search for Hidden Particles



Small hints of a 7 keV DM candidate are not enough to confirm or rule out the  $\nu$ MSM: need to look for  $N_2, N_3$ !  $\rightarrow$  SHiP project [6, 7].

The proposed facility enables the investigation of a wide range of BSM models.

- collaboration rapidly growing
- working on the technical proposal



# SHiP

### My contribution

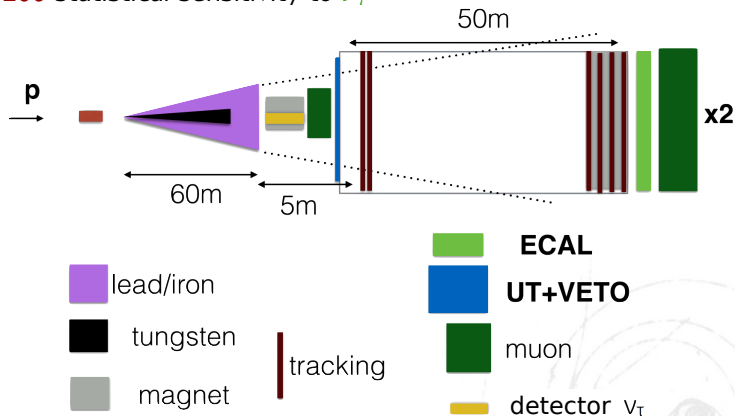
- investigate SHiP's physics reach in the parameter space of  $\nu$ MSM sterile neutrinos  $N_2, N_3$
- determine SHiP's discovery potential in the "Vector Portal"



## Search for Hidden Particles

A proposed fixed-target experiment at the SPS (400 GeV protons) aimed at the study of **long-lived weakly interacting particles** [6, 7].

- $\times 10\,000$  statistical sensitivity to **Heavy Neutral Leptons** & co.
- $\times 200$  statistical sensitivity to  $\nu_\tau$

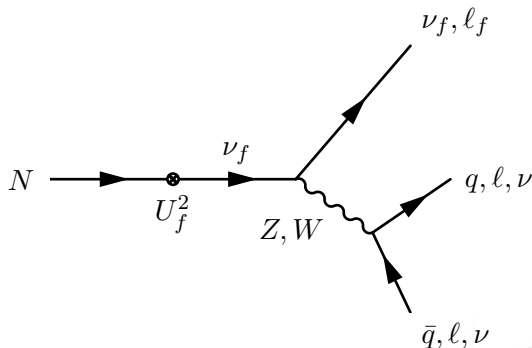


# HNL mixing to Standard Model



i.e. How to detect sterile neutrinos

HNLs can be produced in decays where a  $\nu$  is replaced by a  $N$  (kinetic mixing, **low  $\mathcal{BR}$** ). Main neutrino sources in SHiP:  $c$  and  $b$  mesons.



They can then decay again to SM particles through **mixing** ( $U^2$ ) with a SM neutrino. This (now **massive**) neutrino can decay to a large amount of final states through emission of a  $Z^0$  or  $W^\pm$  boson [11].



# Estimating SHiP's physics reach

HNL production  $\times$  Experimental acceptance



- Number of detected HNL events:

$$\Phi(p.o.t) \times \sigma(pp \rightarrow NX) \times \mathcal{P}_{vtx} \times \mathcal{BR}(N \rightarrow visible) \times \mathcal{A}$$

with

$$\sigma(pp \rightarrow NX) \propto \chi_{cc}, \chi_{bb}, U_f^2$$

$$\mathcal{BR}(N \rightarrow visible) \propto U_f^2$$

- HNL production:
  - $\chi_{cc}, \chi_{bb}$  from simulations
  - $\mathcal{BR}(m_N, U_f^2)$  parametrised according to theory [11]
- Now we only need to compute the **daughters acceptance  $\mathcal{A}$** :
  - HNLs kinematics obtained from simulation
  - every decay channel with **detectable** daughters is simulated

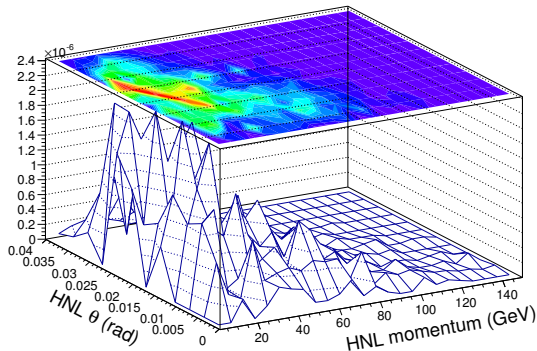
# Estimating SHiP's physics reach



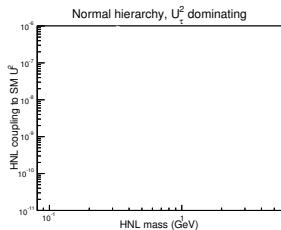
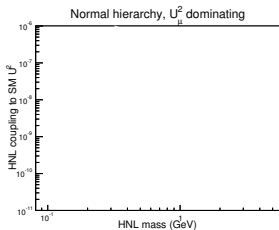
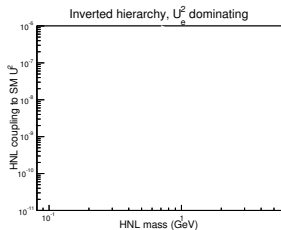
$$\Phi(p.o.t) \times \mathcal{BR}(pp \rightarrow NX) \times \mathcal{P}_{vtx} \times \mathcal{BR}(N \rightarrow visible) \times \mathcal{A}$$

- **Pythia8** used to retrieve the spectrum of  $c$  and  $b$  mesons in SHiP. Heavy neutrinos produced in kinematically-allowed decay chains.
- HNL spectra are re-weighted by the probability that HNLs decay with a vertex inside SHiP's tracking volume:  $\mathcal{P}_{vtx} \sim \int_{SHiP} e^{-l/c\gamma\tau} dl$

Weighted PDF for model 2,  $m_N=1.8$  GeV,  $U_\mu=10^{-9}$



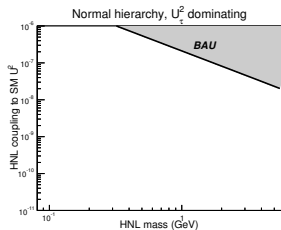
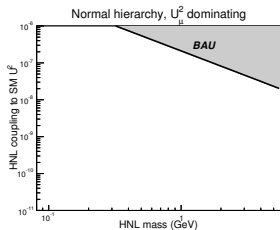
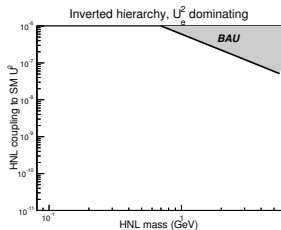
# The parameter space of the $\nu$ MSM



Results must be interpreted according to the **hierarchy of  $\nu$  masses** and to the relative strength of the  $U_e^2$ ,  $U_\mu^2$  and  $U_\tau^2$  couplings.

Three “**extreme**” benchmark models [11] will be shown.

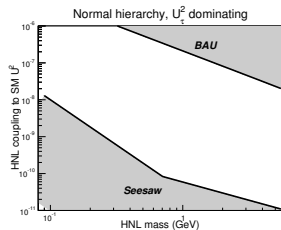
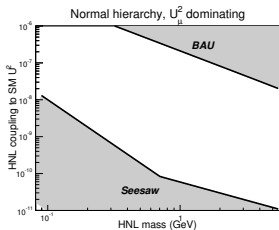
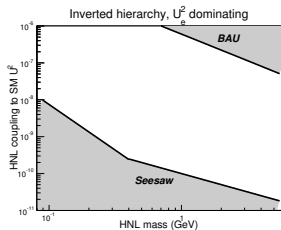
# The parameter space of the $\nu$ MSM



The **cosmologically interesting** parameter space is limited by physical constraints [11]:

1. by the requirement that sterile neutrinos explain the **matter-antimatter asymmetry** (baryogenesis induced by leptogenesis)

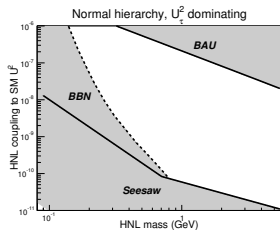
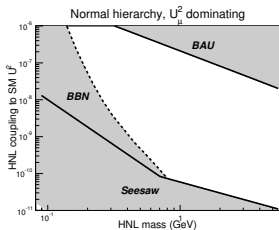
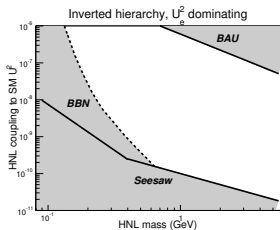
# The parameter space of the $\nu$ MSM



The **cosmologically interesting** parameter space is limited by physical constraints [11]:

2. by the **Seesaw** mechanism: adding an extra singlet neutrino field allows to extend the Lagrangian with a Majorana mass term. The active and sterile neutrino masses  $\nu$  are then inversely proportional

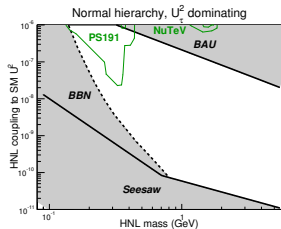
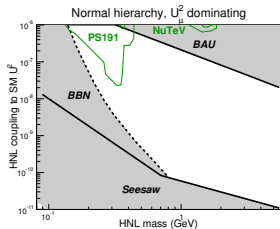
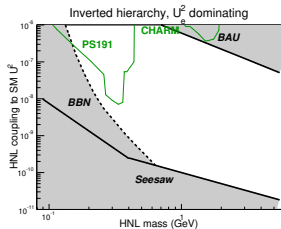
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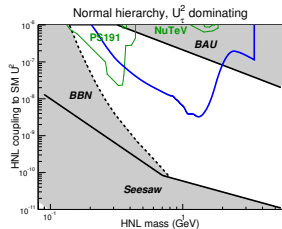
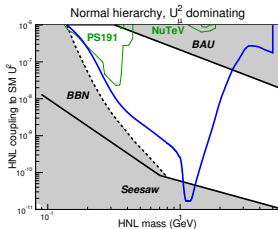
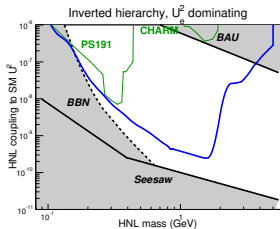
- at low  $M_N$ , by **Big Bang Nucleosynthesis** constraints (observations of the relative abundance of primordial nuclei)

# The parameter space of the $\nu$ MSM



Searches already performed by e.g. PS191, CHARM, NuTeV, but **most** of the parameter space is **still unexplored!**

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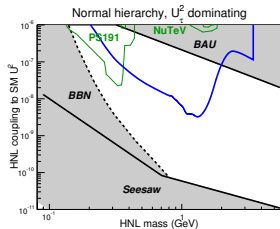
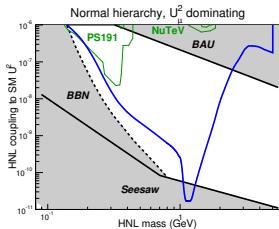
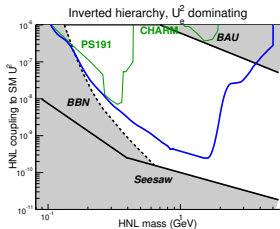


Searches already performed by e.g. PS191, CHARM, NuTeV, but **most** of the parameter space is **still unexplored!**

**Blue lines:** area of the  $\nu$ MSM parameter space that can be ruled out if no event is observed in SHiP (90% C.L.).



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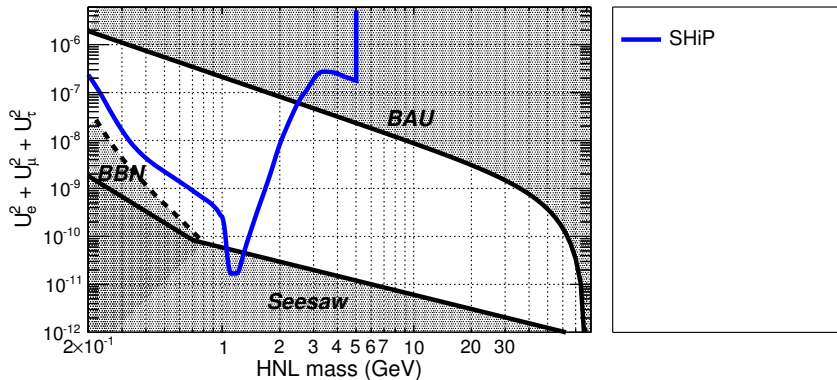
- the remaining area below  $m_N \simeq 350$  keV can be covered by NA62.
- what about  $m_N \gg 1$  GeV?

# Yet another reason to build a $Z^0$ factory

i.e. How to close the HNL parameter space



- Possible  $N$  sources for  $m_N \gg 1$  GeV:  $W^\pm$ ,  $Z^0$

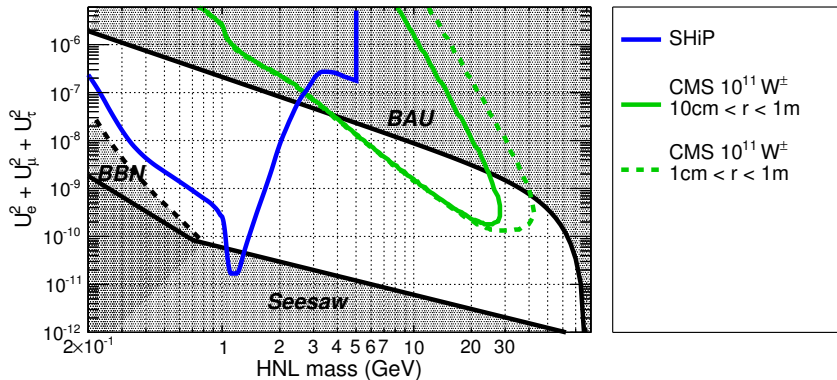


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- $W \rightarrow \ell N$  at LHC: extremely large BG, difficult triggering/analysis.

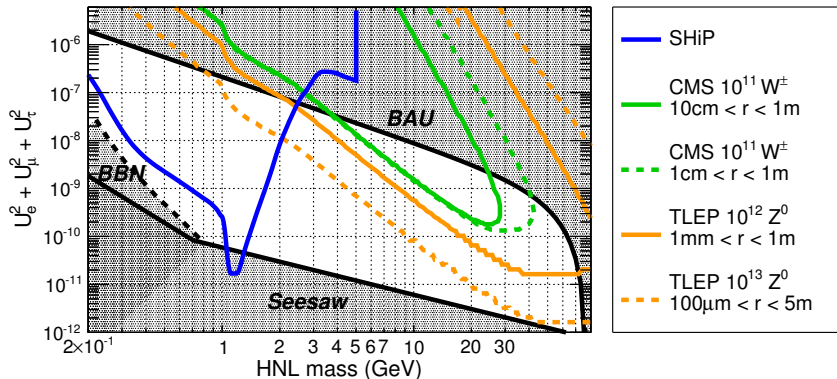


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- $W \rightarrow \ell N$  at LHC: extremely large BG, difficult triggering/analysis.
- $Z \rightarrow \nu N$  at  $e^+e^-$  collider [3]: clean signature, low BG.

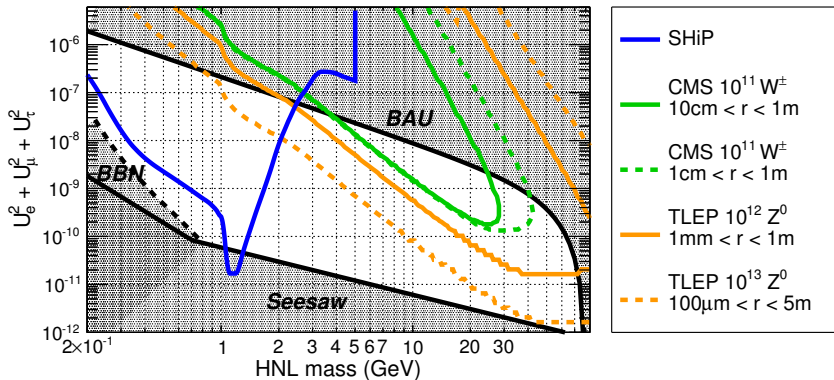
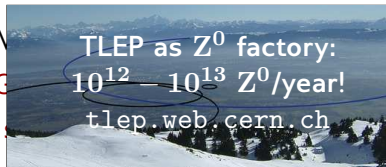


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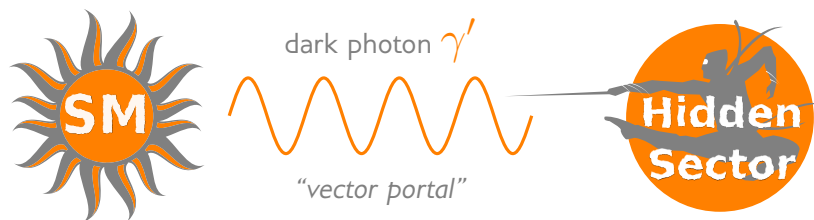
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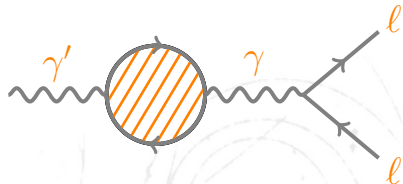
# The vector portal to the hidden sector



Most BSM models predict a set of SM-neutral unobserved particles that do not interact with SM except through a “messenger” particle.

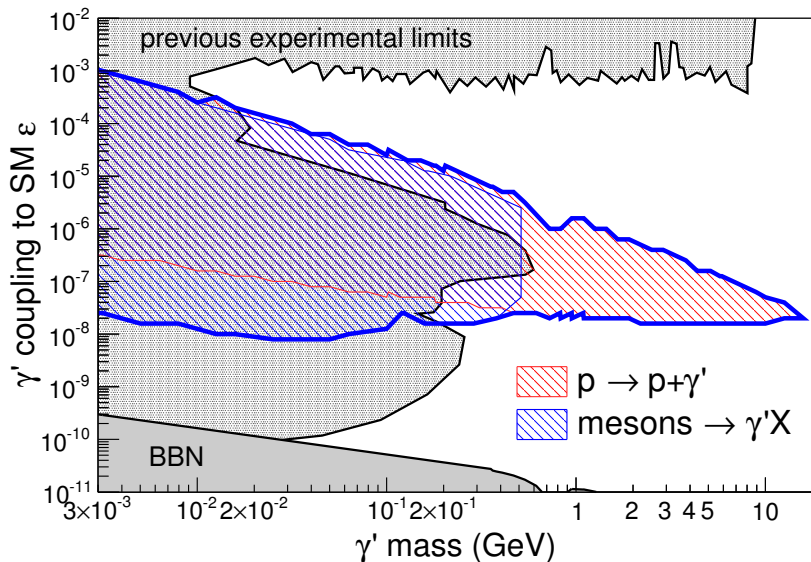


Sensitivity to  $\gamma'$  investigated as for HNLs:  $\gamma'$  mixes to  $\gamma$  through loops of particles charged under both the SM and HS. Extra production mode: **proton bremsstrahlung** [1, 5].



# SHiP's contribution to the vector portal

Sensitivity to dark photons





- Three **fundamental questions** left open by the Standard Model can be answered by the  $\nu$ MSM, with no need for new physics principles
  - **searches for  $N_2, N_3$**  are necessary to **confirm or rule out** the  $\nu$ MSM
  - design of the **dedicated SHiP experiment**
- SHiP's **discovery potential** was investigated
  - **most** of the parameter space can be covered by **SHiP** ( $m_N < 2 - 3$  GeV) and by **future  $Z$  factories** ( $m_N > 3$  GeV)
- SHiP can be used to look for **signatures of other BSM physics**
  - SHiP's physics reach into the **vector portal** determined





**Thanks!**



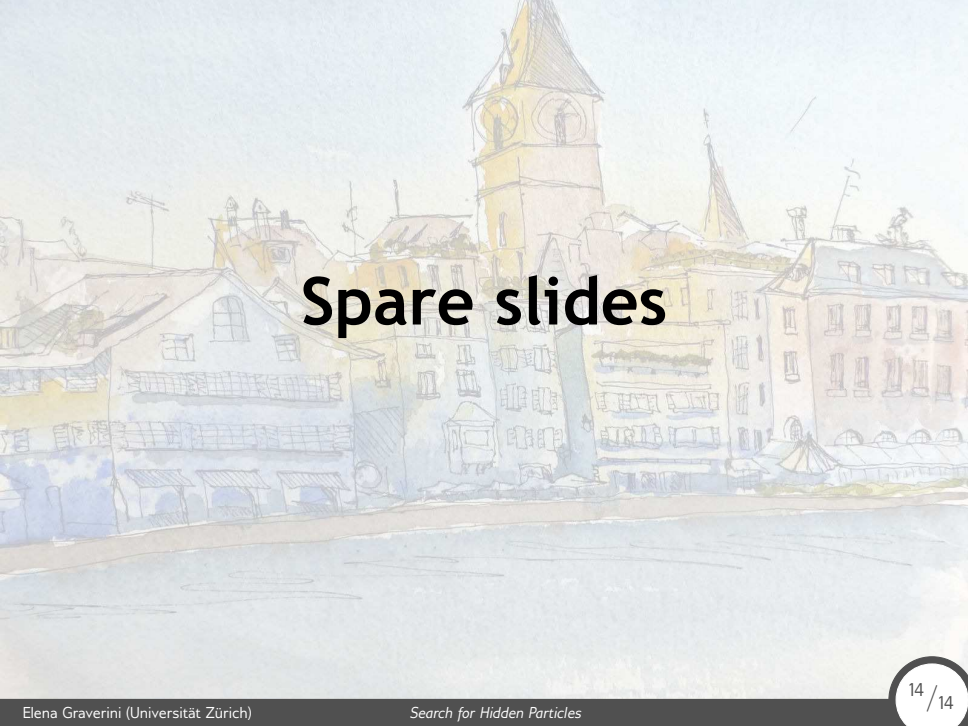
- [1] S. Andreas. “Light Weakly Interacting Particles: Constraints and Connection to Dark Matter”. PhD thesis. Universität Hamburg, 2013.
- [2] T. Asaka, S. Blanchet, and M. Shaposhnikov. “The nuMSM, dark matter and neutrino masses”. In: *Phys.Lett.* B631 (2005), pp. 151–156.
- [3] M. Bicer et al. “First Look at the Physics Case of TLEP”. In: *JHEP* 1401 (2014), p. 164. DOI: 10.1007/JHEP01(2014)164. arXiv: 1308.6176 [hep-ex].
- [4] M. Bicer et al. “First Look at the Physics Case of TLEP”. In: *JHEP* 1401 (2014), p. 164. DOI: 10.1007/JHEP01(2014)164. arXiv: 1308.6176 [hep-ex].



- [5] J. Blümlein and J. Brunner. “New Exclusion Limits on Dark Gauge Forces from Proton Bremsstrahlung in Beam-Dump Data”. In: *Phys.Lett. B*731 (2014), pp. 320–326. DOI: 10.1016/j.physletb.2014.02.029. arXiv: 1311.3870 [hep-ph].
- [6] W. Bonivento et al. “Answers to the questions from the SPSC concerning the SPSC-EOI-010”. CERN-SPSC-2013-024 / SPSC-EOI-010. 2014.
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- [8] A. Boyarsky et al. “An unidentified line in X-ray spectra of the Andromeda galaxy and Perseus galaxy cluster”. In: (2014). arXiv: 1402.4119 [astro-ph.CO].



- [9] E. Bulbul et al. “Detection of An Unidentified Emission Line in the Stacked X-ray spectrum of Galaxy Clusters”. In: *Astrophys.J.* 789 (2014), p. 13. DOI: 10.1088/0004-637X/789/1/13. arXiv: 1402.2301 [astro-ph.CO].
- [10] L. Canetti, M. Drewes, and M. Shaposhnikov. “Matter and Antimatter in the Universe”. In: *New J.Phys.* 14 (2012), p. 095012. DOI: 10.1088/1367-2630/14/9/095012. arXiv: 1204.4186 [hep-ph].
- [11] D. Gorbunov and M. Shaposhnikov. “How to find neutral leptons of the  $\nu$ MSM?” In: *JHEP* 0710 (2007), p. 015. DOI: 10.1007/JHEP11(2013)101, 10.1088/1126-6708/2007/10/015. arXiv: 0705.1729 [hep-ph].



# Spare slides

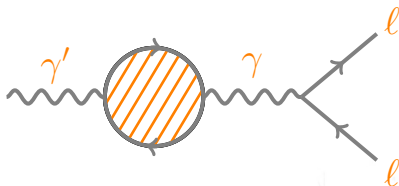


- **SHiP:** it is designed to be a zero-BG experiment: evacuated decay volume, veto chambers, event topology... Active or passive muon shielding.
- **CMS:** every event with either three leptons or one lepton and two jets in the final state!
- **TLEP:**  $W^*W^*$ ,  $Z^*Z^*$  and  $Z^*\gamma^*$  backgrounds suppressed by displacement of the secondary vertex.



## The vector portal [5]

- Extra  $U(1)'$  symmetry with gauge boson  $\gamma'$  (dark photon)
- $U(1)'$  broken by Higgs-like mechanism  $\rightarrow$  non-zero  $m_{\gamma'}$
- Mix to  $\gamma$  through kinetic mixing to particles charged under both  $U(1)$  and  $U(1)'$ :  $\mathcal{L}_{eff} = \mathcal{L}_{SM} + \mathcal{L}_{hidden} + \frac{\epsilon}{2} A'_{\mu\nu} F^{\mu\nu}$



# Estimating SHiP's physics reach

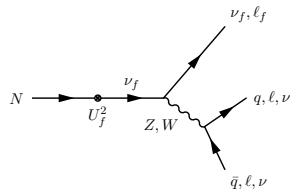
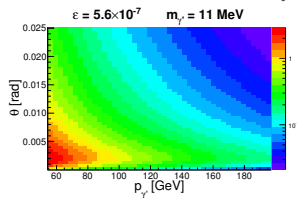


1. Production:  $10^{20}$  p.o.t.  
HNL:  $D_s \rightarrow \ell N$ ,  $D \rightarrow K \ell N$   
 $\gamma'$ :  $p \rightarrow p \gamma'$ , meson decays  $\rightarrow \gamma' X$
2.  $(p, \theta)$ -PDF weighted with the probability that the particle decays inside the SHiP volume ( $\mathcal{P}_{vtx}$ )
3. Vertex acceptance:  $\int_{SHiP} e^{-l/c\gamma\tau} dl$
4. Simulate decays and compute daughters acceptance ( $\mathcal{A}$ )
5. Count the number of events:

$$N = \Phi(\text{p.o.t}) \times BR_{\text{prod}} \times U_f^2 \times BR_{\text{visible}} \times \mathcal{P}_{vtx} \times \mathcal{A}$$

where  $U_f^2$  is the mixing angle to SM particles.

**HNL reach at TLEP:**  $Z^0 \rightarrow \nu N$ , similar procedure.





# HNL lifetime

## and branching ratios



For a given  $N$  mass, its lifetime was computed on the basis of the widths of its kinematically allowed decay channels, according to the formulas in [11]. I included all the main  $N$  decay channels:

- $N \rightarrow H^0 \nu$ , with  $H^0 = \pi^0, \rho^0, \eta, \eta'$
- $N \rightarrow H^\pm \ell^\mp$ , with  $H = \pi, \rho$
- $N \rightarrow 3\nu$
- $N \rightarrow \ell_i^\pm \ell_j^\mp \nu_j$
- $N \rightarrow \nu_i \ell_j^\pm \ell_j^\mp$

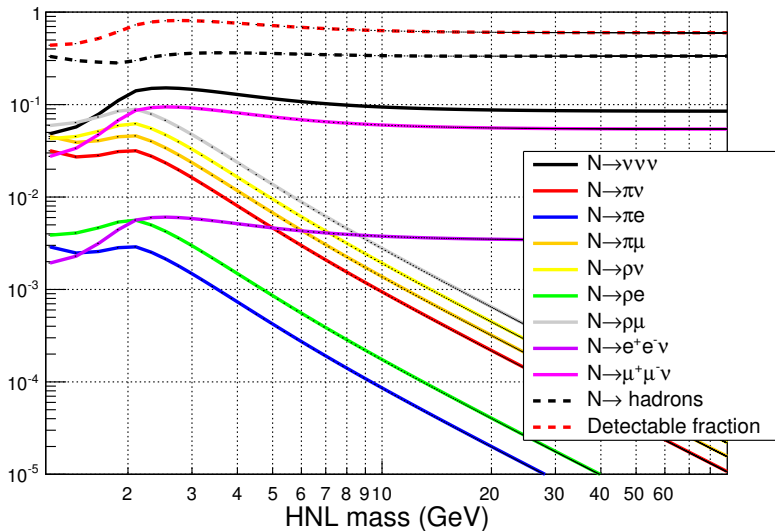
The formulas in [11] are valid **up to**  $m_N \simeq 1$  GeV. For  $m_N \gg \lambda_{QCD}$ , the two quarks do no longer hadronise together and should be considered free. In this regime ( $m_N \geq 2$  GeV) I extrapolated the three-body leptonic formulas to the quark case to compute the total hadronic decay width.

# HNL lifetime

and branching ratios



Branching ratios for HNL (model:  $U^2 = [6.25e-10, 1e-08, 2.5e-09]$ )





All decay channels into  $\geq 2$  charged particles were taken to be **visible**.

**Decay chains** such as  $N \rightarrow \rho^0 \nu$  followed by  $\rho^0 \rightarrow \pi^+ \pi^-$  were taken into account, with the pion pair as final state particles.

Final-state  $\pi^0$ s were taken not to be detectable, even if the photon pair could be reconstructed with a high-granularity calorimeter. For decay chains like  $N \rightarrow \rho^\pm \ell$  with  $\rho^\pm \rightarrow \pi^\pm \pi^0$ , the requirement that the two photons fall into the acceptance of a toy detector of radius 2.5 m decreases the final-state detection efficiency to  $\sim 27\%$ , while it is  $\sim 37\%$  for a  $(\pi^\pm, \ell)$  final state.

# How to estimate the number of expected events

For SHiP



1. Charm mesons are the main source of HNLs in SHiP
2. Simulate protons on target  $\implies$  store an **ntuple of charm mesons** with their secondary vertices
3. Fix the model (fix the  $U_f^2$ ) and the HNL mass  $m_N$
4. According to the dominating  $U_f^2$ , simulate the  $D_s \rightarrow \ell N$  and  $D^{\pm,0} \rightarrow K^{0,\mp} \ell^{\pm} N$  decays or the  $D_s \rightarrow \tau \nu_\tau, \tau \rightarrow \mu \nu N$  decay chain
5. Momentum and angle of the outgoing  $N$  are stored in a binned bi-dimensional PDF
6. Each bin of the PDF is weighted with the **probability that a  $N$  with that kinematics decays into SHiP's fiducial volume**, and the total probability  $\mathcal{P}_{vtx}$  is computed as the **integral of the weighted PDF**

# How to estimate the number of expected events

For SHiP



7. For each of the kinematically allowed decays into detectable daughters, decays inside SHiP's fiducial volume are simulated to compute SHiP's acceptance to the daughters ( $\mathcal{A}$ ). Decaying  $N$  are sampled from the weighted PDF. The total branching ratio  $BR(visible)$  into detectable particles is computed.
8. For each of the two fiducial volumes, the number of expected events is computed as:

$$N_{events} = N_{p.o.t.} \times 2\chi_{cc} \times BR(production, f) \times U_f^2 \\ \times \mathcal{P}_{vtx} \times BR(visible) \times \mathcal{A}$$

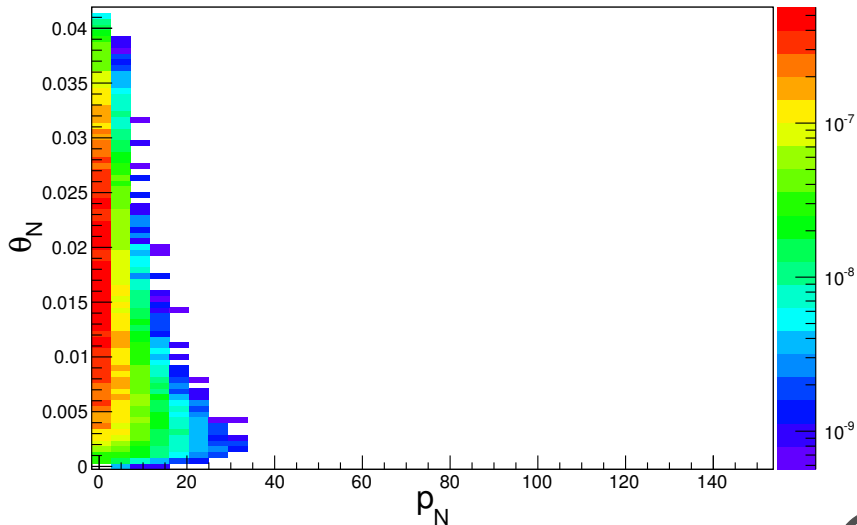
It depends on the three mixing angles  $U_f^2$  and on  $m_N$ .

# How to estimate the number of expected events

For SHiP



weightedPDF\_vol1\_m0.51\_couplings1.75e-09\_2.8e-08\_6.65e-09



# How to estimate the number of expected events



For TLEP, assuming running as  $Z^0$  factory

The expected statistic is  $10^{12} - 10^{13}$   $Z^0$ /year. The following considerations is applicable to any  $e^+e^-$  collider capable of producing a large quantity of  $Z^0$  bosons.

The only production channel considered is  $Z^0 \rightarrow \nu\bar{\nu}$ , with one neutrino mixing to  $N$ . An onion-like detector is assumed.

1. Fix the model (fix the  $U_f^2$ ) and the HNL mass  $m_N$
2. The  $Z^0$  decays in place to  $\nu N$ . The boost to the  $N$  is then:

$$\gamma = \frac{m_Z}{2m_N} + \frac{m_N}{2m_Z}$$

# How to estimate the number of expected events



For TLEP, assuming running as  $Z^0$  factory

3. Compute the  $N$  lifetime  $\tau$ . Compute the probability  $\mathcal{P}_{vtx}$  that it decays inside the tracking volume:

$$\mathcal{P}_{vtx} = \frac{1}{c\gamma\tau} \int_{R_{min}}^{R_{max}} e^{-\frac{r}{c\gamma\tau}} dr$$

where  $R_{min}$  corresponds to the minimum  $SV$  displacement ( $\sim$  inner tracker resolution)

4. Compute the total branching ratio  $BR(visible)$  into lepton pairs ( $ee$ ,  $\mu\mu$  and  $e\mu$ ). The detector efficiency is assumed to be 100%
5. The number of expected events is computed as:

$$N_{events} = N_{Z^0} \times BR(Z^0 \rightarrow 2\nu) \times 2 \times \sum_f U_f^2 \times \mathcal{P}_{vtx} \times BR(visible)$$

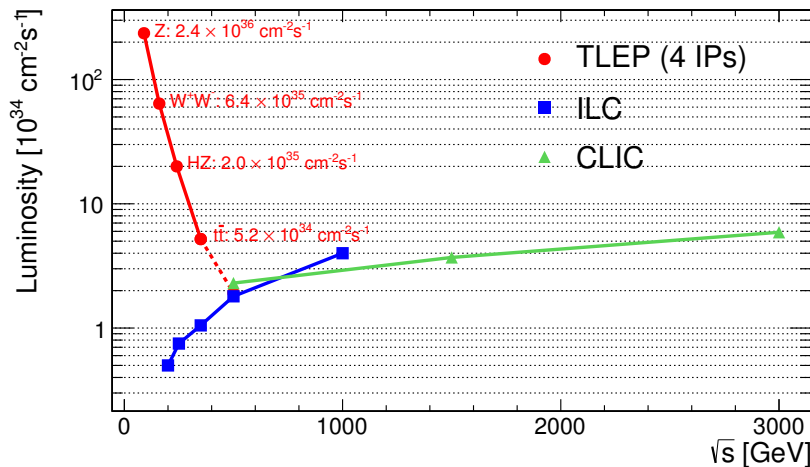


# How to estimate the number of expected events

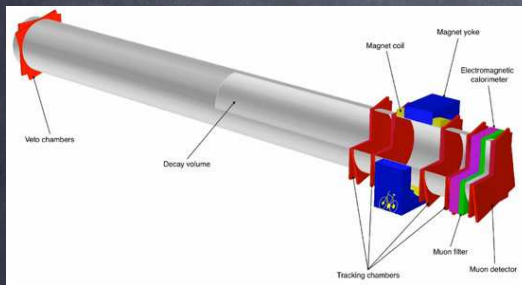


For TLEP, assuming running as  $Z^0$  factory

Expected luminosity is large at the  $Z^0$  mass [4]:



# Decay Volume



- ④ Vacuum tank (similar to NA62) with  $1e-2$  mbar (instead of  $1e-5$  mbar)
- ④ NA62-like straw chambers, 120 $\mu$ m resolution and 0.5% X0/X
- ④ LHCb-like magnet 0.5Tm over 5m
- ④ LHCb-like shashlik calorimeter
- ④ Veto chambers at the entrance of the vacuum tank to veto muons and strangeness from surroundings

# Geometry

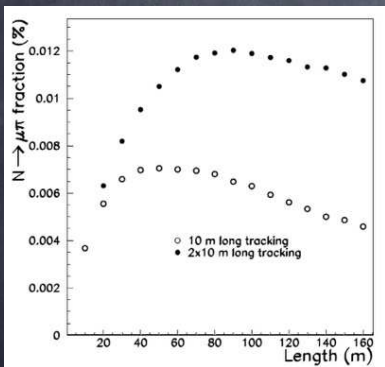


Figure 8: Fraction of HNL in the detector acceptance as a function of the length of the fiducial volume. Open circles: a single spectrometer following a fiducial volume of a given length. Full circles: two spectrometers in series, each following a fiducial volume of half the given length. The spectrometer length is fixed to 10 m.

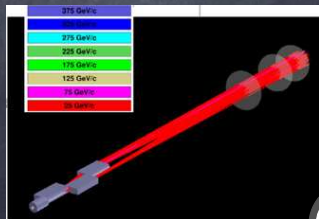
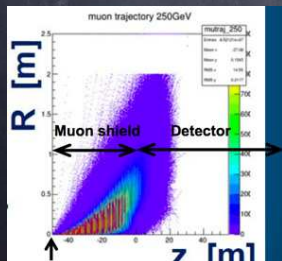
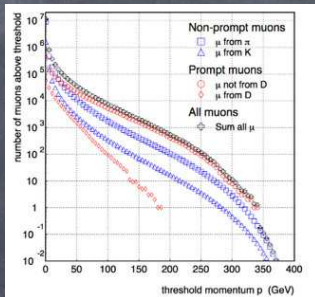
# Muon Shield

N. Serra, Flavour Physics Conference, Vietnam, 01.08.2014

Muon rate  $5 \times 10^9$  muons/spill

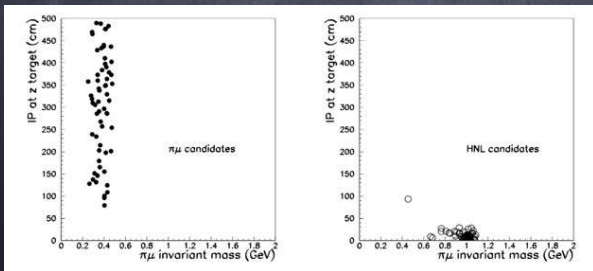
- Acceptable rate  $< 10^5$  muons per spill
- Main source of muons from  $\eta$ ,  $\eta'$ ,  $\omega$ , etc...
- Studying solution with passive or active filter

## Schematics



# Background rejection

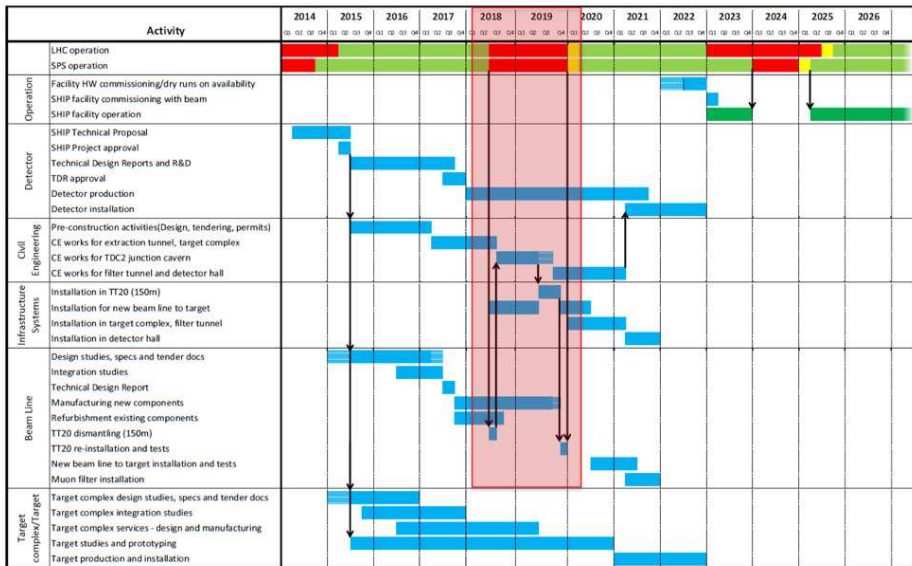
- $2 \times 10^4$  neutrino interactions per  $2 \times 10^{20}$ POT in the decay volume at atmospheric pressure, negligible at  $10^{-2}$ mbar
- $K_L$  production from  $\nu + A \rightarrow K_L(\rightarrow \mu\nu\pi)X$
- 10%  $\nu$  interactions produce  $\Lambda$  and  $K^0$  in acceptance
- Majority of the decays in the first 5m of the decay volume
- Muon filter to reduce background from muon DIS to a negligible level



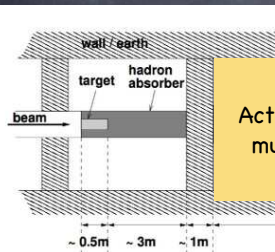
Fighting hard to design  
a zero background  
experiment

# N. Serra, Flavour Physics Conference, Vietnam, 01.08.2014





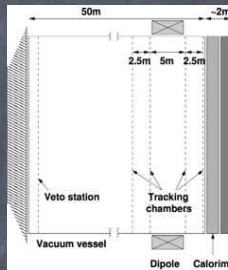
# Tau neutrinos



Active/passive  
muon shield

neutrino tau  
detector

5 m



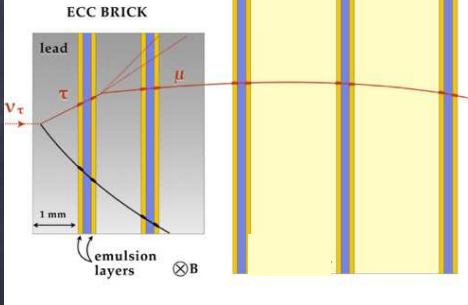
- The same optimization for sterile neutrinos in the GeV region also maximises the flux of  $\nu_\tau$
- Source of  $\nu_\tau$  and  $\bar{\nu}_\tau$  is  $D_s \rightarrow \tau \nu_\tau$
- Also high rate of  $\nu_e$  from charm



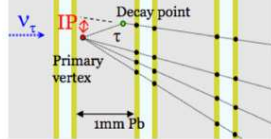
# Tau neutrino detector

N. Serra, Flavour Physics Conference, Vietnam, 01.08.2014

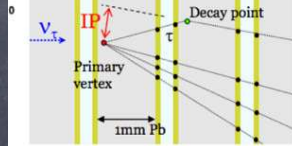
## Muon spectrometer



### Short flight decay



### Long flight decay



- $\nu_\tau$  target: Opera-like bricks, laminated lead and nuclear emulsions (for micrometric resolution)
- 750 Opera-like bricks, to be replaced 10 times
- Muon spectrometer to measure charge and momentum and give time stamp

## Working hypotheses

- Detector located  $\sim 60\text{m}$  from the proton target
- Charm production cross-sections in p-W affected by large uncertainties
- Compare with DONUT to extrapolate the expected numbers
- Energy dependence of  $\sigma_{cc}$  and  $\nu_\tau$  cross-section, acceptance: production  $\sim 0.36$ , detector acceptance  $\sim 0.2$ , energy dependence of the  $\nu_\tau$  cross-section  $\sim 0.52 \rightarrow \text{DONUT/SHIP} \sim 26$
- $2 \times 10^{20}$  pot for SHIP compared to  $3.6 \times 10^{17}$  DONUT  $\rightarrow \sim 550$  in favour of SHIP
- Overall rate SHIP/DONUT  $\sim 20$
- DONUT observed 9 events with a background of 1.5  $\rightarrow 7.5 \pm 3$  (40% error)
- 150 events expected with the same mass (260 kg)
- Measurement of  $\nu_\tau$  and anti- $\nu_\tau$  cross-section, including the study of structure functions sets the scale for the mass:  $\sim 6$  tons for  $\sim 3400 \nu_\tau$  interactions
- Assume OPERA-like bricks (8.3 kg) and wall target structure:  $\sim 750$  bricks