

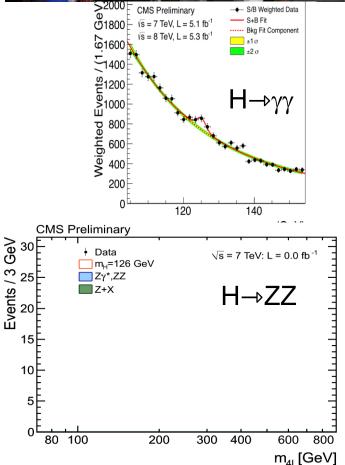
#### CMS:

### Readiness for new energy frontier: first 13 TeV data Planning the long term future

## Outline

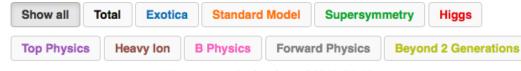
- Following the new boson discovery the success of Standard Model of Nature is paradoxically highlighting its limitations.
- There are fundamental questions which are not yet answered by the Standard Model.
- With the LHC restarting at 13 TeV and opening a new energy domain we expect to be able to address some of these fundamental questions.
- We have prepared this exciting run and we are preparing for the long term future



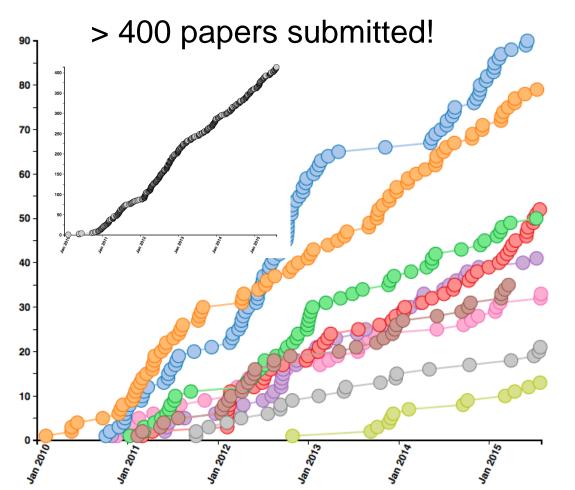




## A rich legacy from Run1







Pub rate steady, ~2.5/week 413 papers submitted: +23 Cosmic ray based +35 ready for CWR or later +17 PubDraft

#### In review process (115):

84 GoingToPreApp or higher 44 in PAG review

## First collisions @ 13 TeV

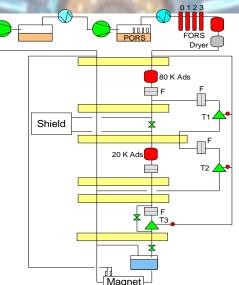
CMS Experiment at the LHC, CERN Data recorded: 2015-May-21 07:59:01.776704 GMT Run / Event / LS: 245194 / 31876157 / 47

## **Readiness for LHC Run2**

- CMS has undergone major changes during the Long Shutdown, amongst other:
  - Added two new endcap planes of muon chambers
  - DAQ: changed the event building fabric and the Timing and control system
  - Replaced the readout devices of the Outer Hadron and the Forward Hadron calorimeter
  - Validated the readout at -20 degrees for the tracker and pixel
  - Implemented the first stage of the Upgraded level 1 calo trigger
- The setup has been re-commissioned using cosmic ray in progressive steps since July 2014: running 24/7 since early February
- The goal for the 2015 campaign is to collect ≤ 10 fb<sup>-1</sup> of luminosity of proton collisions (a period of Heavy ion collisions is foreseen for the last month)

## **Magnet Cryogenics**

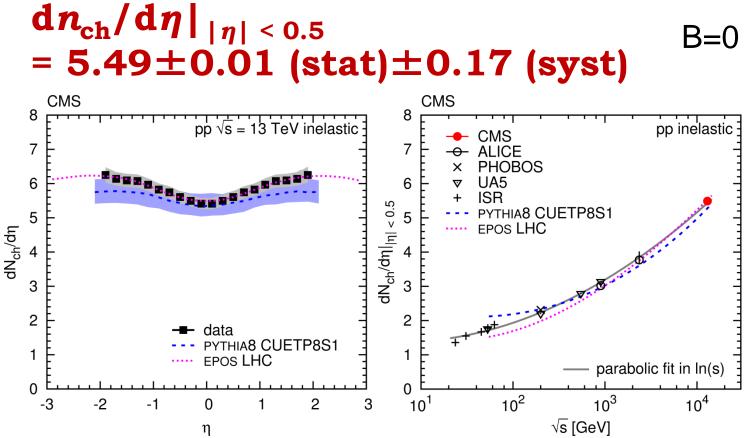
- The restart of the CMS magnet after LS1 was more complicated than anticipated due to problems with the cryogenic system in providing liquid Helium.
- Inefficiencies of the oil separation system of the compressors for the warm Helium required several interventions and delayed the start of routine operation of the cryogenic system.



- The data delivered during the first two weeks of LHC re-commissioning with beams at low luminosity have been collected with B=0
- Currently the magnet can be operated, but the continuous up-time is still limited by the performance of the cryogenic system requiring more frequent maintenance than usual.
- A comprehensive program to re-establish its nominal performance is underway. These recovery activities for the cryogenic system will be synchronized with the accelerator schedule in order to run for adequately long periods.
- A consolidation and repair program is being organized for the next short technical stops and the long TS at the end of the year.

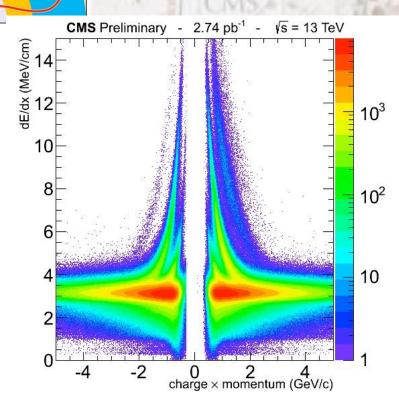


...and we are ready

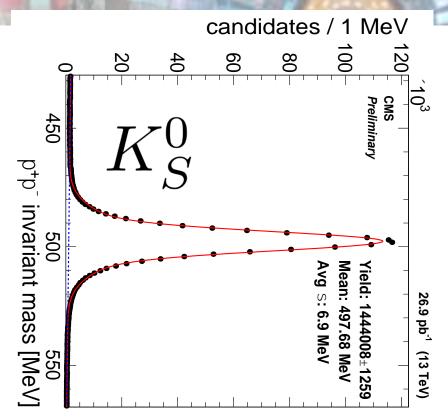


- First measurement of inelastic  $dN_{ch}/d\eta$  at 13 TeV pp collisions.
- First paper submitted by LHC experiment on the 13 TeV data

#### **Tracker Performance**



From right to the left Deuteron, Proton and Kaon bands clearly visible. Particles shown depending on their charge → observe an expected asymmetry in the production of particles/antiparticles for pp collisions.



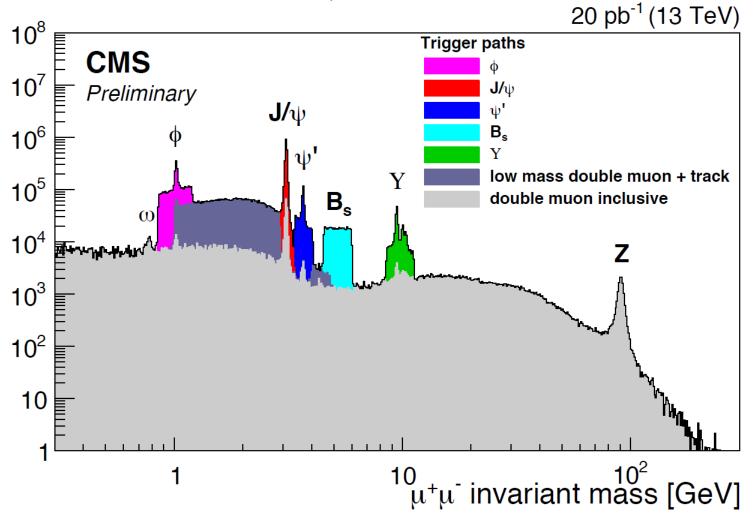
PDG:  $497.614 \pm 0.006$ 

Invarient mass of the pion pairs fitted with a double gaussian and a first order polynomial for the background.

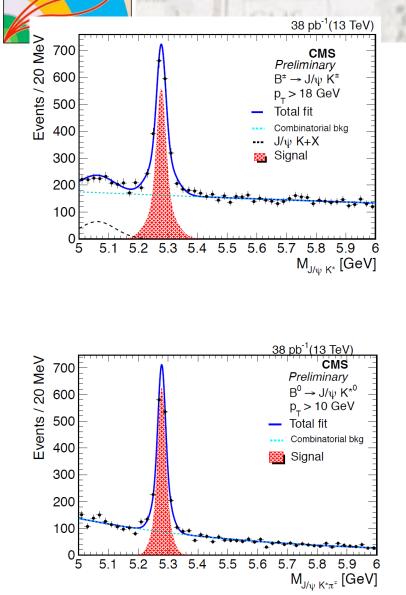
## Trigger

#### **Special Triggers**

Selective mass regions and  $p_{\scriptscriptstyle T}$  cuts

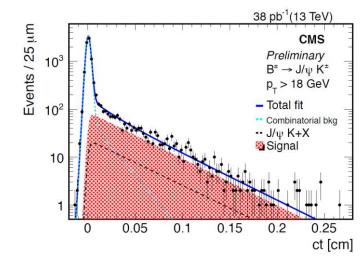


#### Resonances

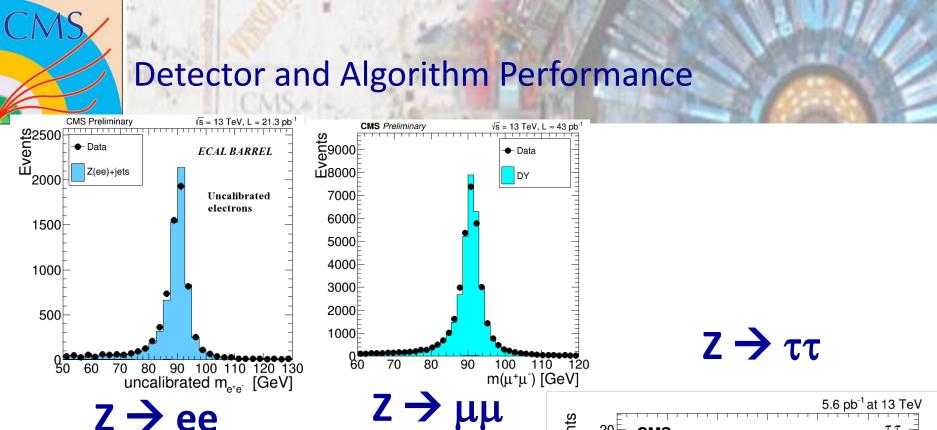


 $B^{\pm}$  → J/ψ K<sup>±</sup> Inclusive J/ψ trigger with p<sub>T</sub> >16 GeV Mass: 5.277 ± 0.001(stat.) GeV

PDG: 5279.26  $\pm$  0.17 MeV

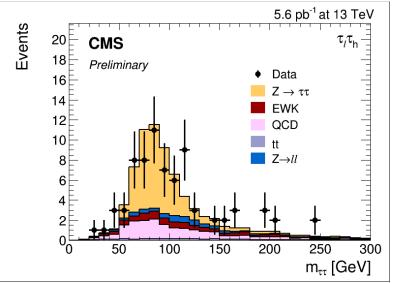


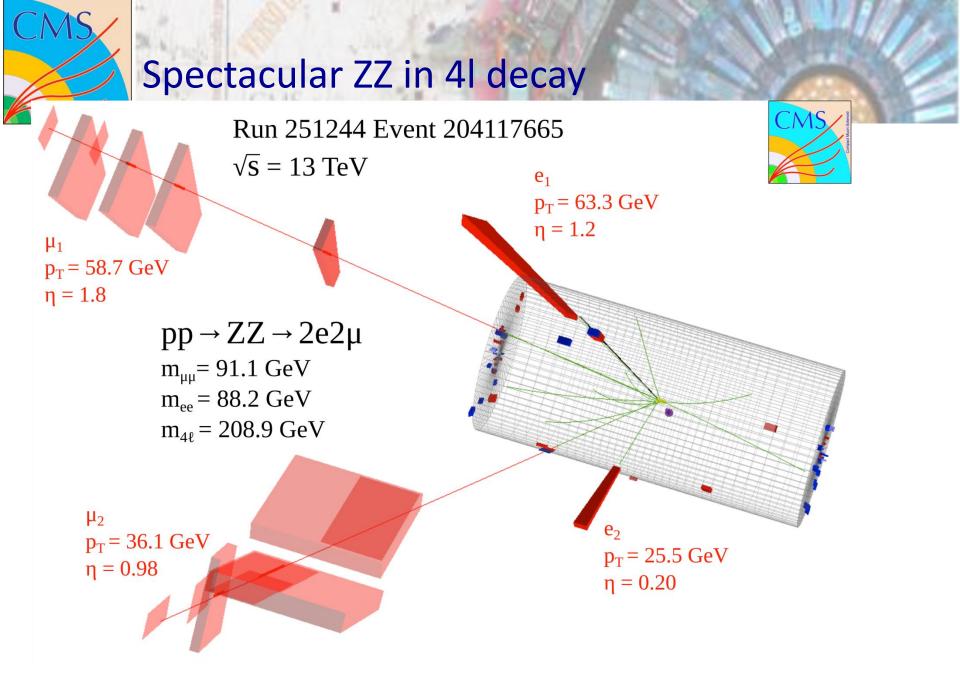
 $B^0_s$  → J/ψ φ with displaced J/ψ and track trigger Mass: 5.369 ± 0.001(stat.) GeV PDG: 5366.7 ± 0.4 MeV

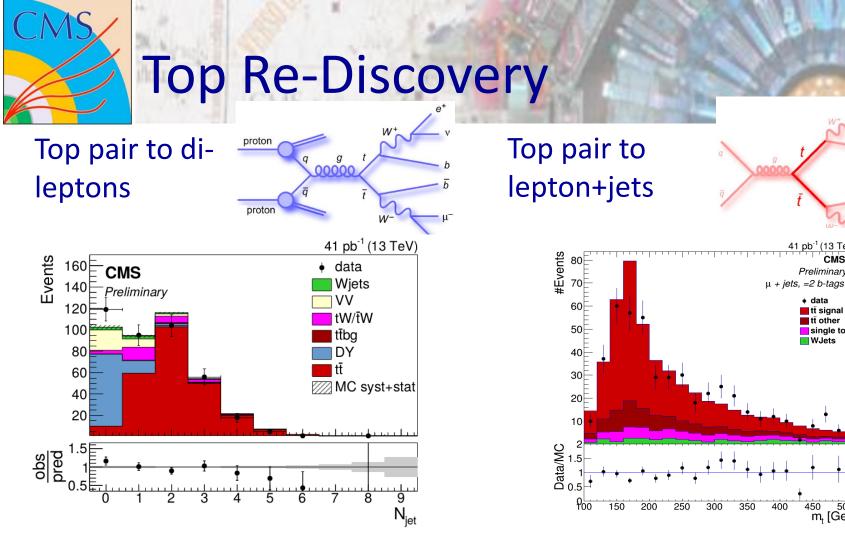


 $Z \rightarrow ee$ 

#### Z – Boson as standard physics candle







Number of hadronic jets for events containing one isolated muon and one isolated electron forming an invariant mass greater than 50 GeV.

Reconstructed hadronic top guark candidate mass for events containing 1 isolated muon ( $p_T$ >30 GeV,  $|\eta$ |<2.4, passing medium identification), four jets (p<sub>T</sub>>30 GeV,  $|\eta| < 2.5$ ), out of which 2 pass the tight b-tagging threshold.

41 pb<sup>-1</sup> (13 TeV)

Preliminary

ltī signal

single top

<sup>450</sup> 500 m<sub>t</sub> [GeV]

tī othe

WJets

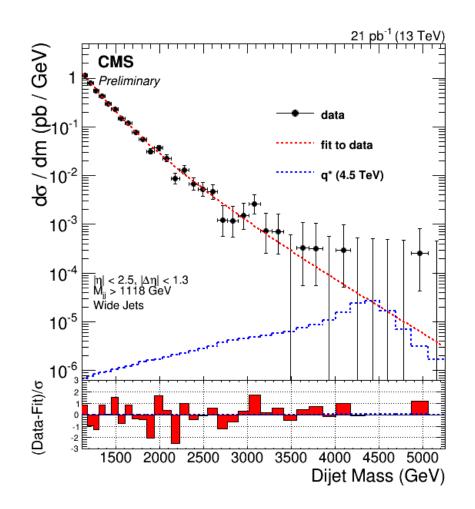
data

CMS



#### **Resonance Search with Di-Jets**

- Data: black points
- Fit to the data: red dashed line
- Blue dashed line: superimposed signal shape of a q\* → qg resonance with M=4.5 TeV
- Performed a maximum likelihood fit
- Rough estimate of the goodness of fit the χ<sup>2</sup> is calculated to χ<sup>2</sup>/ ndof = 31.84 / 31
- Good fit to the data using a background parameterization
- $\rightarrow$  no evidence for a di-jet resonance.



## The Physics objectives: The next years are for frontier exploration

## The past (pre July 4<sup>th</sup>, 2012)

Unified gauge theories of weak and electromagnetic interactions provide an attractive framework for the interpretation of weak-interaction phenomena.<sup>1</sup> Such theories are universal in the prediction that existing data explore only the lowenergy tail of a spectrum of yet-to-be-discovered particles. The most familiar of the hypothetical particles are the massive vector bosons  $W^*$  and  $Z^0$ associated with the observed charged and neutral weak currents. Somewhat more obscure are the massive scalar Higgs bosons which are connected with the spontaneous breakdown of gauge symmetry.

Lee, Quigg, Thacker, PRD, 1 September 1977

## The beacon for the last 40 years

Theoretical considerations<sup>3</sup> suggest that the Higgs-boson mass must exceed about  $4 \text{ GeV}/c^2$ , and we have recently derived a conditional upper bound<sup>4</sup>

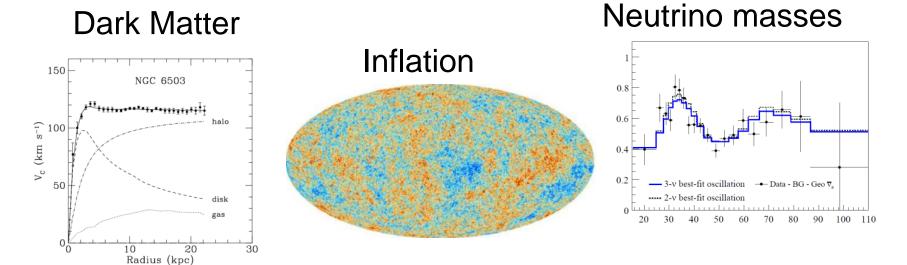
$$M_{H} \leq M_{c} = (8\pi\sqrt{2}/3G_{F})^{1/2} \simeq 1 \text{ TeV}/c^{2},$$

(1.1)

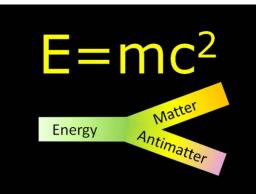
where  $G_F$  is the Fermi constant. The precise meaning of the upper bound is that if  $M_H$  exceeds the critical value  $M_c$ , weak interactions will become strong in the TeV energy regime in the sense that perturbation theory will cease to be a faithful representation of physics.

# Today: answers missing

 Many evidences of physics beyond the Std Model



Matter-Antimatter asymmetry

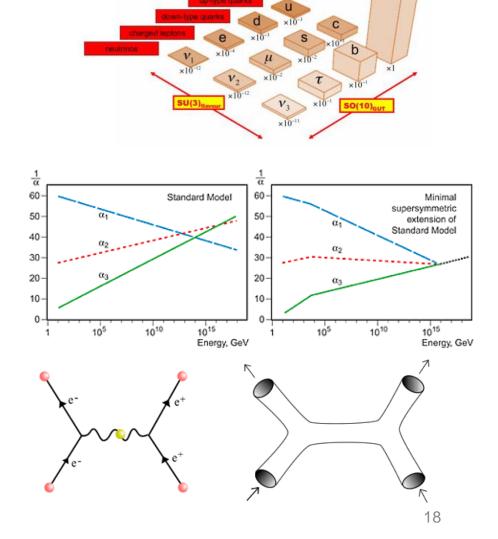


# Today: SM cannot explain

 Flavor families and mass Hierarchy

Unification of forces

• Quantum gravity

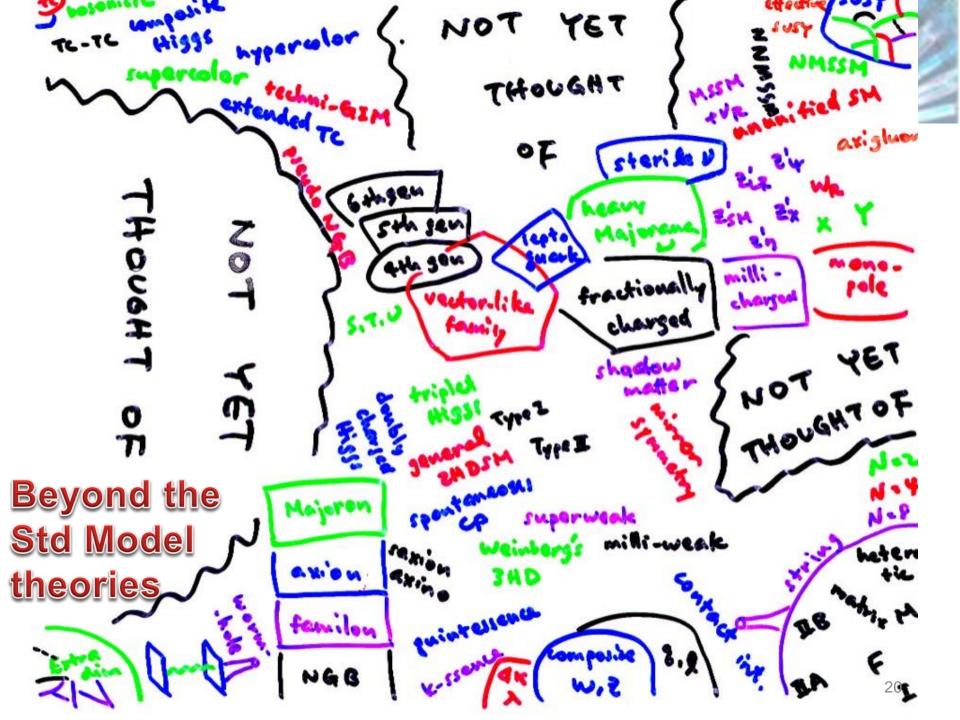




## Today: we lack a beacon

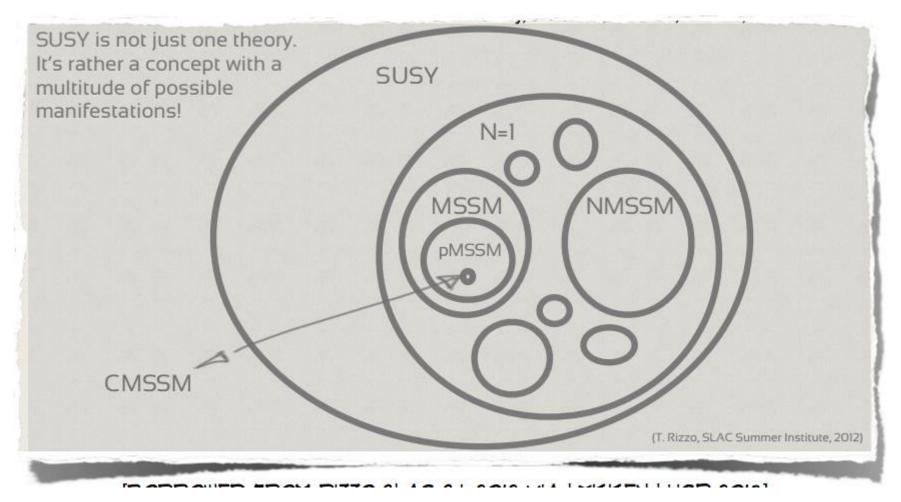
- The heuristic principle of 'Naturalness' DOES NOT WORK: the fact that LHC has not yet found any BSM evidence is EVIDENCE of fine tuning at sub permill level....
- ... so the 'scale' of new physics  $(\Lambda_{\rm NP})$  is not at E-weak scale ... but it is hard to think that the unanswered questions and the internal difficulties of the SM do not imply some new interactions/state of matter

• Lacking a 'guiding light' it is now up to the experiments to trace the route !



## Is SUSY still worth seeking?

SUSY 'space' has been excluded only for the easiest part so far (CMSSM... still a lot to explore..but it might not be easy )

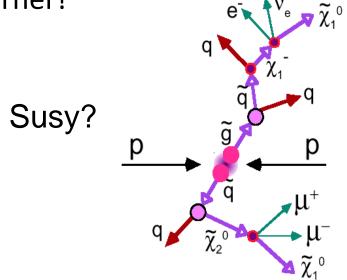


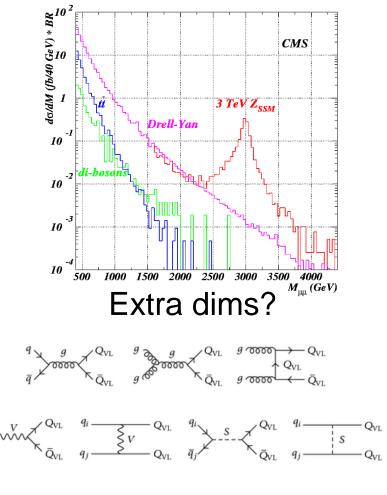


## 2015 and RUN2

A lot at stake!

We know that there must be more than the new scalar boson...and that in order to make things simple the new physics should be around the corner: we have to make sure that we do not miss the corner!





Something else?

#### **Discoveries come early**

#### Courtesy of Gavin Salam

Discovery potential comparison 20fb<sup>-1</sup> @8TeV vs X fb<sup>-1</sup> @13 TeV

30 fb<sup>-1</sup> @ 13 TeV 5 fb<sup>-1</sup> @ 13 TeV system mass [TeV] for 13.00 TeV, 30.00 fb<sup>-1</sup> a a 6 q q Σi qi qi qg system mass [TeV] for 13.00 TeV, 5.00 fb<sup>-1</sup> 6 Σi qi qi q g g g 5 g g 5 4  $\sqrt{\hat{s}}$  @13*TeV* 4  $\sqrt{\hat{s}}$  @13*TeV* з 3 2 2  $\sqrt{\hat{s}} @8TeV$  $\sqrt{\hat{s}}$  @8*TeV* 1 1 0 0 0.5 3.5 0 0.5 1.5 2.5 3.5 1.5 2.5 system mass [TeV] for 8.00 TeV, 20.00 fb<sup>-1</sup> system mass [TeV] for 8.00 TeV, 20.00 fb<sup>-1</sup>



## A primary Energy Frontier objective for the future: Dark matter

## Fritz Zwicky: 1933

Galaxies in the Coma cluster were moving too rapidly.

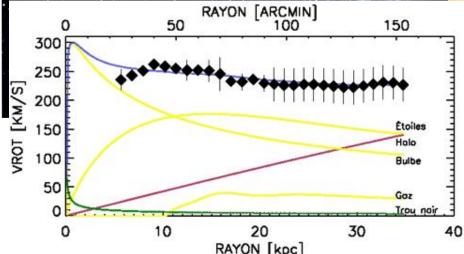
> He proposed "Dunkle Materie" as the explanation.

The beginning of the Dark Matter mistery



## Vera Rubin (1970s) Rotation Curves of Galaxies are Flat







#### The Bullet Cluster:

Two merging clusters: dark matter passes through while atoms get stuck

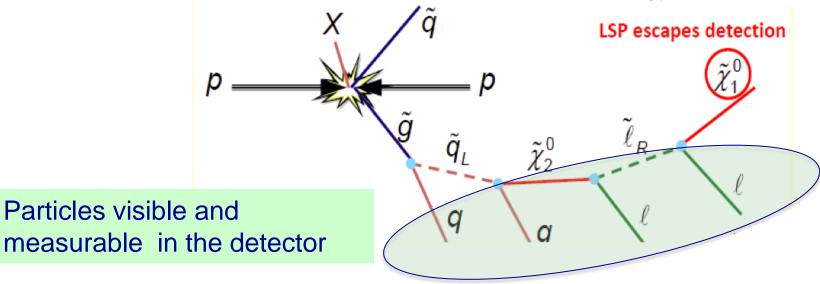
#### Atomic Matter 🔰

#### Dark Matter 🔰



 Direct way is by detecting anomalous Missing energy from final states

Dark Particle which carries 'away' invisible energy



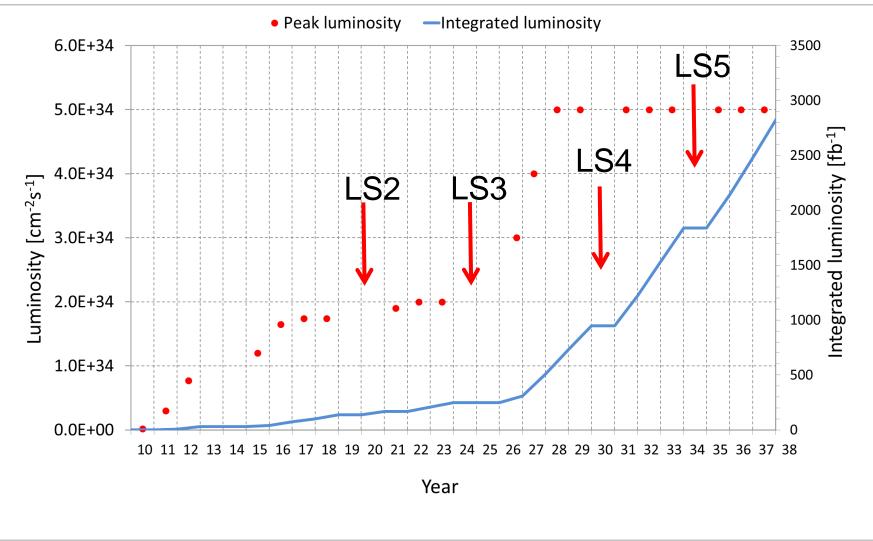
Will not be easy... but we have built our detectors to be able to do it !



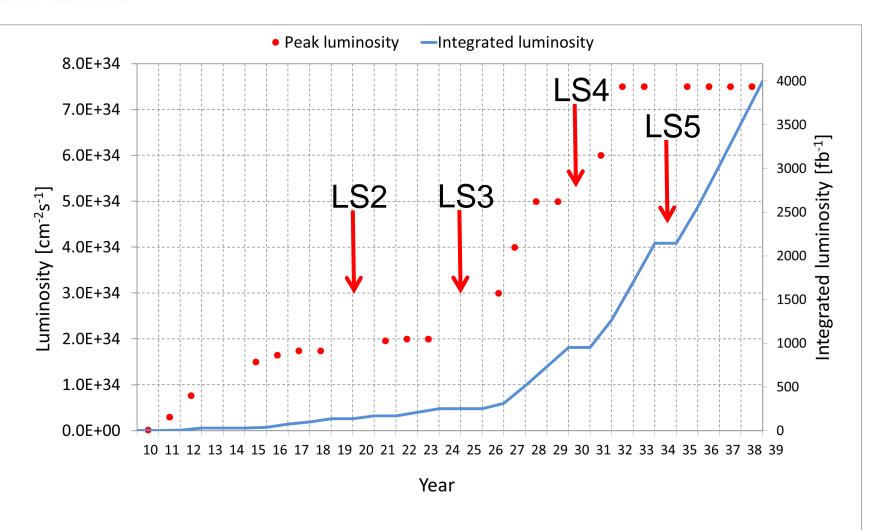
## The long term future

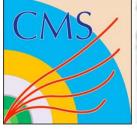
- The CERN council has approved the extension of the LHC program beyond the initial phase (meant to deliver 300 fb<sup>-1</sup> of integrated luminosity)
- The High Luminosity LHC program is now a project : a three years shutdown (LS3) will mark the start of new era with an accelerator able to to reach 10<sup>35</sup> Hz/cm<sup>2</sup> luminosities (levelled to 5 10<sup>34</sup>, possibly 7 10<sup>34</sup> Hz/cm<sup>2</sup>) and aiming to deliver in excess of 3000 fb<sup>-1</sup>
- This will require major upgrades to the LHC detectors where several components will have reached their end of lifetime limits and other will have to be upgraded to exploit the higher instantaneous luminosities and radiation levels
- US P5 :Recommendation 10: Complete the LHC phase-1 upgrades and continue the strong collaboration in the LHC with the phase-2 (HL-LHC) upgrades of the accelerator and both general-purpose experiments (ATLAS and CMS).
   The LHC upgrades constitute our highest-priority near-term large project.
- US and EUROPE are in synch regarding priorities !

## The future: 5x10<sup>34</sup> levelling: a draft plan



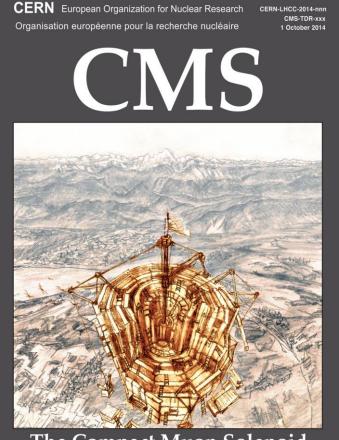
### The future: same beam, leveling at 7.5 10<sup>34</sup> (PU 200 in average)





## CMS Upgrades

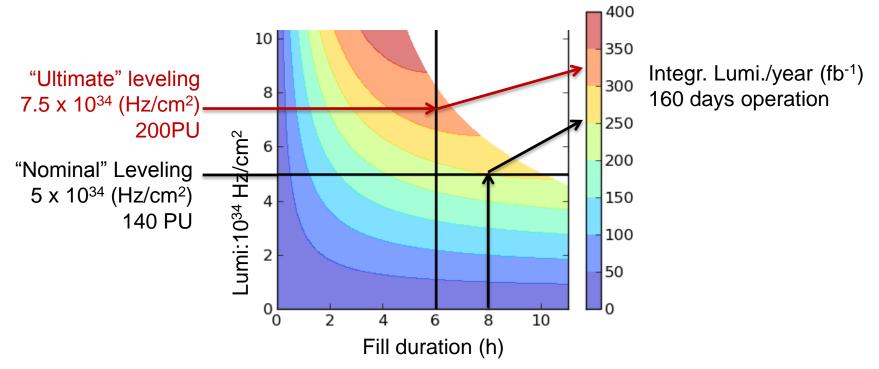
We have submitted in June our Technical Proposal for the High Luminosity LHC upgrade



The Compact Muon Solenoid Phase II Upgrade Technical proposal

# Upgrade scope: the issue is Pileup

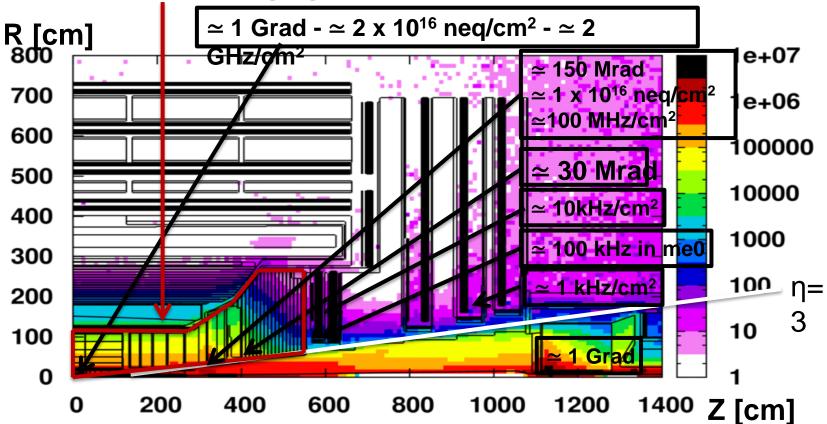
#### Pileup= number of concurrent pp interactions when 2 bunches cross



CMS upgrades enable operation at 200 PileUp, maintain current performance at 140 PU, allowing moderate degradation up to 200 Pileup and radiation tolerance ≥ 3000 fb<sup>-1</sup>

## Radiation dose - neutron fluence particle rates

- 3000 fb<sup>-1</sup> Dose map in [Gy] simulated with MARS and FLUKA
- Numbers in boxes indicate maximum doses neutron equivalent fluence particle rates (for 5 x 10<sup>34</sup> Hz/cm<sup>2</sup>) seen by the various detectors
- These studies show that Tracker & End cap Calorimeters need replacement due to aging

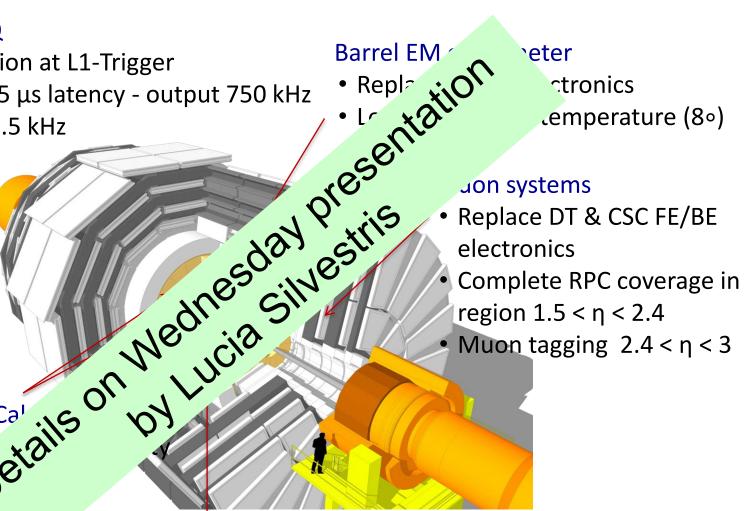


34

#### Summary of the CMS upgrades for Phase-II

#### Trigger/HLT/DAQ

- Track information at L1-Trigger
- L1-Trigger: 12.5 μs latency output 750 kHz
- HLT output  $\simeq$ 7.5 kHz



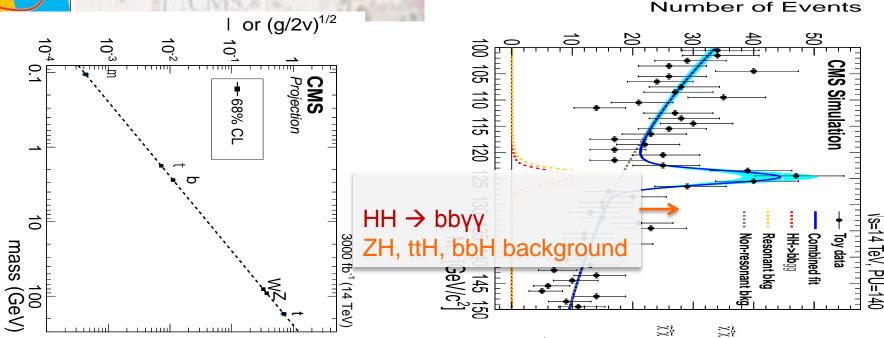
#### Replace Endcap Cal

- Rad. tolerant
- 3D capabili

#### **Replace Tracker**

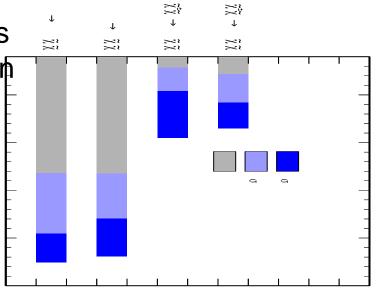
- .ad. tolerant high granularity significantly less material
- 40 MHz selective readout (Pt≥2 GeV) in Outer Tracker for L1-Trigger
- Extend coverage to  $\eta = 3.8$

## **Examples of physics reach**



 $\circ$  2 to 10% precision on BEH couplings

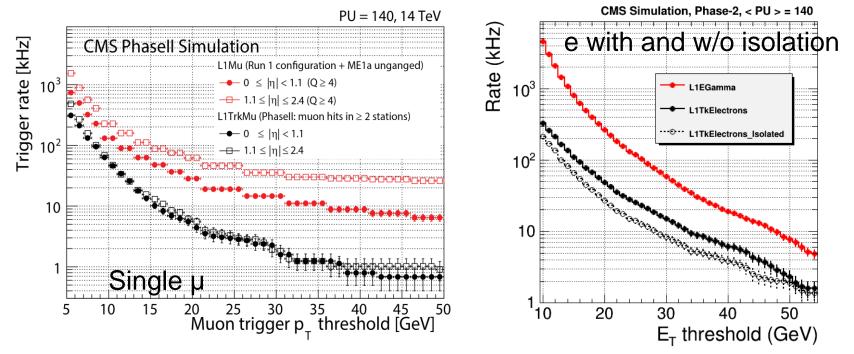
- Evidence of di-BEH boson production
- Access to small cross section SUSY processes
- Several other SM rare processes and BSM physics predictions
- $\,\circ\,$  And increased mass range



#### A major improvement: first level track trigger

#### Powerful scheme to control all inclusive trigger rates at first 40 MHz stage

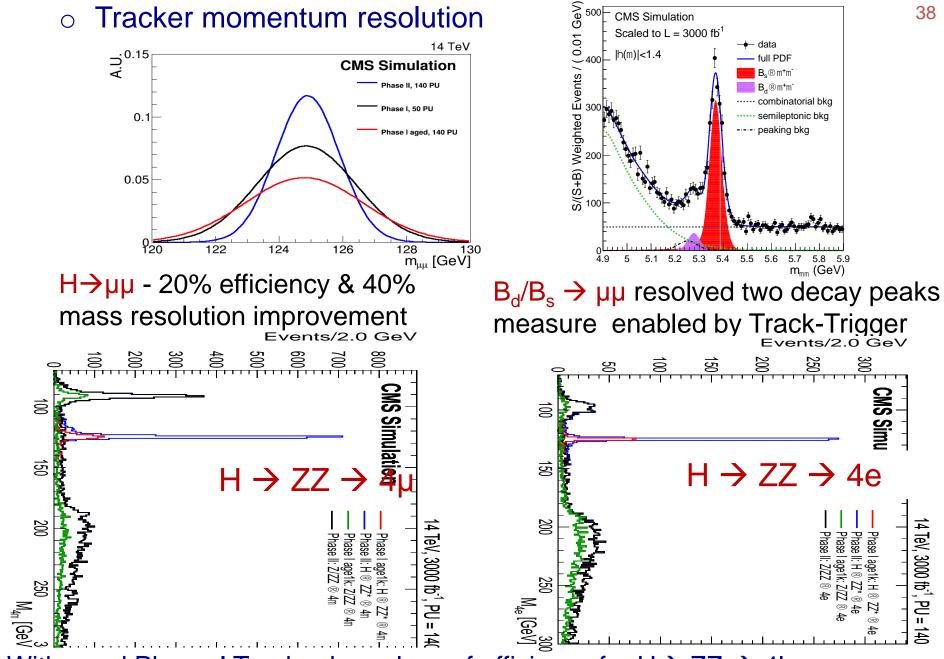
- $\circ$  Single  $\mu$  rate divided by 10
- Single e rate divided by 5(10) w/o (with) isolation
- γγ rate/5 from isolation
- о т efficiency x 2 at same rate
- Vertex  $\simeq$  1 mm resolution → HT & MET rates divided by 10 to 100



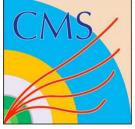
L1-Trigger studies with Phase-I menu thresholds including Track-Trigger:

 $\circ$  Requires  $\simeq$  500/750 kHz rate at 140/200 PU (with 1.5 safety margin)





With aged Phase-I Tracker huge loss of efficiency for  $H \rightarrow ZZ \rightarrow 4I$ Upgrade restores efficiency and extensions increase the acceptance by 20%



#### Discovery can come early ... or late

#### Explored:

- 9 different
  experimental
  signatures.
- 5 different types of SUSY models.

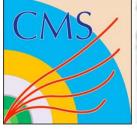
Different types of SUSY models lead to different patterns of discoveries in different final states after different amounts of data

<b>.</b> .	Analysis	Luminosity	Model				
e.		$({\rm fb}^{-1})$	NM1	NM2	NM3	STC	STOC
Exploring experimental signature space	all-hadronic (HT-MHT) search	300					
		3000					
	all-hadronic (MT2) search	300					
		3000					
	all-hadronic $\widetilde{b}_1$ search	300					
		3000					
	1-lepton $\tilde{t}_1$ search	300					
		3000					
	monojet $\widetilde{\mathfrak{t}}_1$ search	300					
		3000					
	$m_{\ell^+\ell^-}$ kinematic edge	300					
		3000					
	multilepton + b-tag search	300					
		3000					
	multilepton search	300					
		3000					
	ewkino WH search	300					
		3000					
ш		$< 3\sigma$ 3 – 5 $\sigma$	$> 5\sigma$				

#### Exploring SUSY model space



- CMS is ready to continue the exciting journey into the mysteries of Nature
- We are all excited by exploring the new energy frontier
- We are preparing our long term future: CMS has been conceived as an Energy frontier experiment, but has demonstrated that with adequate detector a proton collider is also a precision measurement tool
- Stay posted for exciting times ahead !



## Backup

#### 43

## Dark matter : reachable at LHC?

- Weakly Interacting Massive Particles are the best motivated dark matter candidates, Annihilation rate in the early universe determines the density today.
- The annihilation rate comes purely from particle physics and automatically gives the right answer for the relic density!
- $W_{X}h^{2} = \frac{3 \cdot 10^{-27} \, cm^{3} / s}{\langle SN \rangle_{ann}}$
- This is the mass fraction of WIMPs today, and gives the right answer (23%) if the dark matter is weakly interacting and the WIMP mass in the range GeV – 10 TeV

