



Current Cosmic Acceleration with Slotheon

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Saha Theory Workshop

Observational evidences of current Cosmic Acceleration

- **Supernovae type Ia**

—————→ **Universe is Accelerating**

standard candles

Their intrinsic luminosity is known

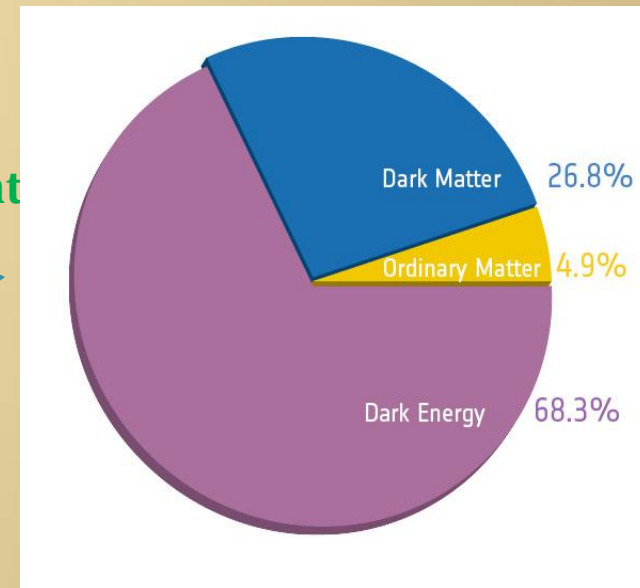
- **Large Scale Structure**

Clustering of matter gives information on cosmological parameters, especially matter content

- **Cosmic Microwave Background**

CMB is a strong pillar of the Big Bang cosmology

It is a powerful tool to use in order to constrain several cosmological parameters



Dynamical dark energy models

$$w \neq -1$$

There are two approaches.

$$G_{\mu\nu} = 8\pi G T_{\mu\nu}$$

(Einstein equation)

(i) Modified gravity

f(R) gravity,
Scalar-tensor theory,
Braneworlds,
Gauss-Bonnet gravity,
.....

(ii) Modified matter

Quintessence,
K-essence,
Tachyon,
Chaplygin gas,
.....

- Should be free from negative energy instability
- Should not conflict with local physics

Slotheon Gravity

$$S_{\text{sloth}} = \frac{1}{2} \int d^4x \sqrt{-g} \left[M_{\text{P}}^2 R - \left(g^{\alpha\beta} - \frac{G^{\alpha\beta}}{M^2} \right) \partial_\alpha \pi \partial_\beta \pi \right]$$

C. Germani *et al*,
PRD 85,103501,2012

Galileon symmetry : $\pi \rightarrow \pi + c + c_\mu x^\mu$

The name comes from the fact that on a given metric

$$H_k \sim \dot{\pi}^2 \left(1 + \frac{G^{tt}}{M^2} \right) \geq \dot{\pi}^2$$

Given a energy density the field is **slower** than its canonical cousin!!

Gravitationally enhanced friction

Slotheon in a Potential

D.Adak, A. Ali, D. Majumdar

PRD 88, 024007, 2013

$$S = \int d^4x \sqrt{-g} \left[\frac{1}{2} \left(M_{\text{pl}}^2 R - \left(g^{\mu\nu} - \frac{G^{\mu\nu}}{M^2} \right) \pi_{;\mu} \pi_{;\nu} \right) - V(\pi) \right] + \mathcal{S}_m \left[\psi_m; e^{2\beta\pi/M_{\text{pl}}} g_{\mu\nu} \right]$$

Crucial coupling between matter and the scalar field

- The action is not galileon invariant
- It is free from Ostrograsky ghost problem
- The slowing of field is solely due to gravitational interaction

Equation of motion in flat FRW background

$$3M_{\text{pl}}^2 H^2 = \rho_m + \rho_r + \frac{\dot{\pi}^2}{2} + \frac{9H^2 \dot{\pi}^2}{2M^2} + V(\pi),$$

$$M_{\text{pl}}^2 (2\dot{H} + 3H^2) = -\frac{\rho_r}{3} - \frac{\dot{\pi}^2}{2} + V(\pi) + \frac{\dot{\pi}^2}{2M^2} (2\dot{H} + 3H^2) + \frac{2H\dot{\pi}\ddot{\pi}}{M^2},$$

$$-\frac{\beta}{M_{\text{pl}}} \rho_m = \ddot{\pi} + 3H\dot{\pi} + \frac{3H^2}{M^2} \left(\ddot{\pi} + 3H\dot{\pi} + \frac{2\dot{H}\dot{\pi}}{H} \right) + V'(\pi).$$

$$\dot{\rho}_m + 3H\rho_m = \frac{\beta}{M_{\text{pl}}} \dot{\pi} \rho_m,$$

$$\dot{\rho}_r + 4H\rho_r = 0.$$

Autonomous System

Equations

$$\begin{aligned}\frac{dx}{dN} &= x \left(\frac{\ddot{\pi}}{H\dot{\pi}} - \frac{\dot{H}}{H^2} \right), \\ \frac{dy}{dN} &= -y \left(\sqrt{\frac{3}{2}} \lambda x + \frac{\dot{H}}{H^2} \right), \\ \frac{d\epsilon}{dN} &= 2\epsilon \frac{\dot{H}}{H^2}, \\ \frac{d\Omega_r}{dN} &= -2\Omega_r \left(2 + \frac{\dot{H}}{H^2} \right), \\ \frac{d\lambda}{dN} &= \sqrt{6} x \lambda^2 (1 - \Gamma),\end{aligned}$$

Variables

$$\begin{aligned}x &= \frac{\dot{\pi}}{\sqrt{6} H M_{\text{pl}}}, & y &= \frac{\sqrt{V(\pi)}}{\sqrt{3} H M_{\text{pl}}}, \\ \epsilon &= \frac{H^2}{2M^2}, & \lambda &= -M_{\text{pl}} \frac{V'(\pi)}{V(\pi)},\end{aligned}$$

$$N \equiv \ln a, \quad \Gamma = \frac{V V_{,\pi\pi}}{V_{,\pi}^2}$$

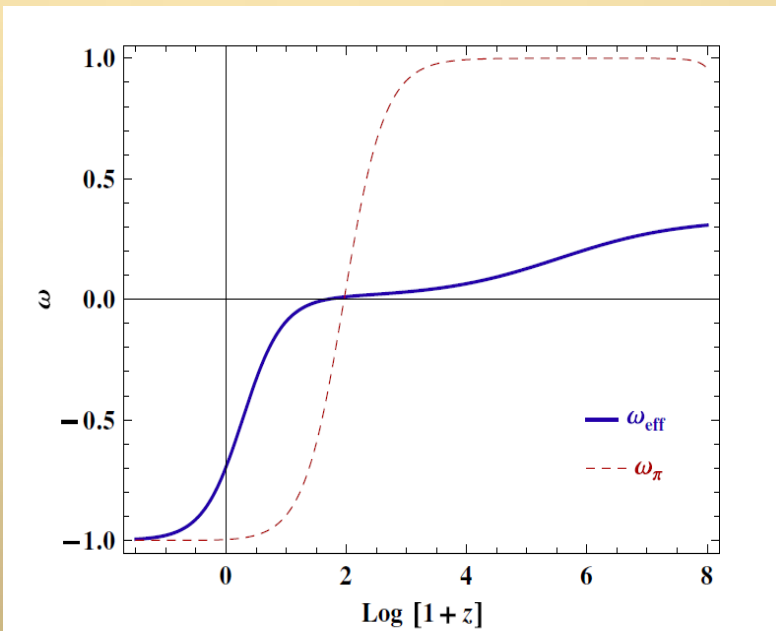
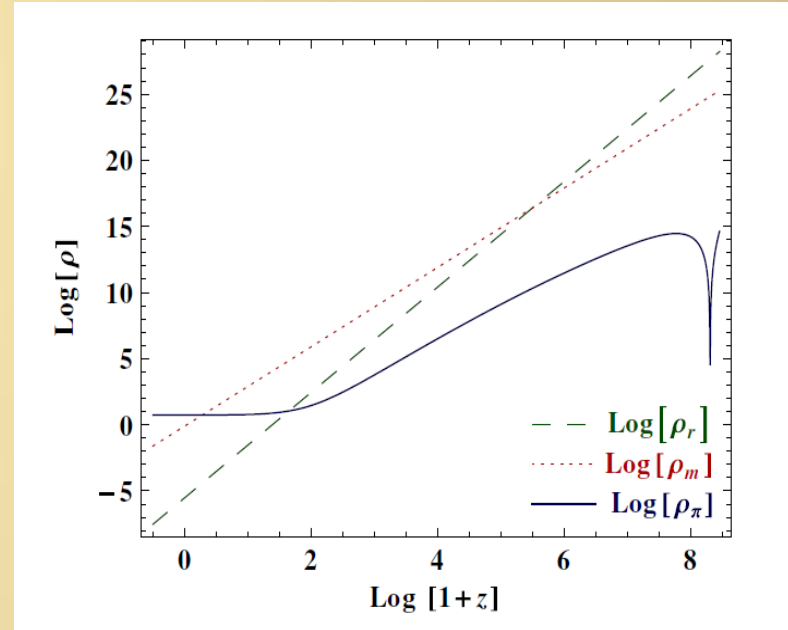
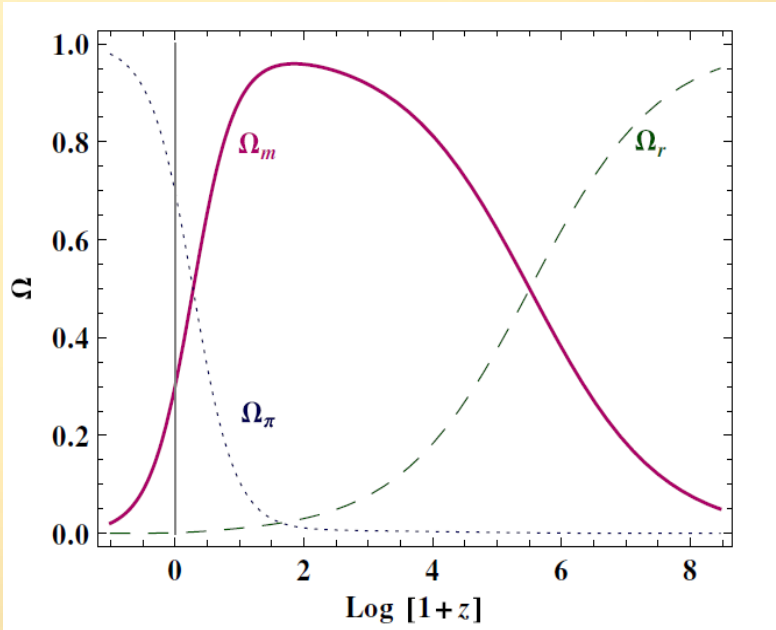
$$\Omega_m = 1 - (x^2(1 + 18\epsilon) + y^2 + \Omega_r)$$

Form of potential

$$V(\pi) = V_0 e^{\frac{-\lambda\pi}{M_{\text{pl}}}}$$

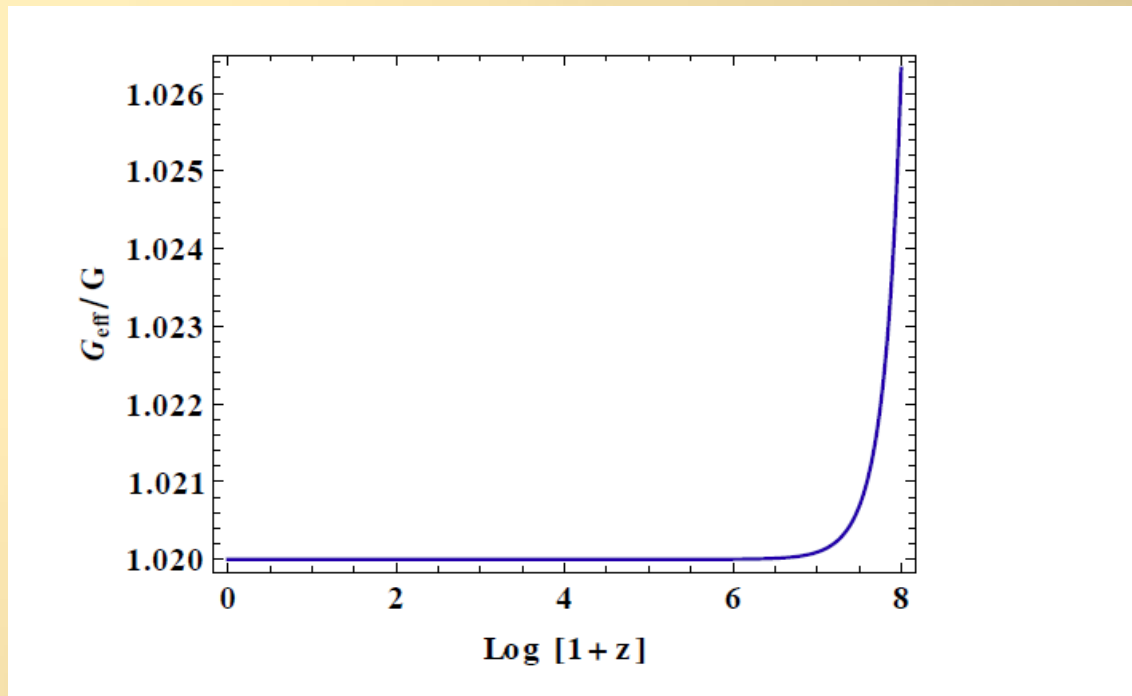
- $\Gamma = 1$ and therefore λ is a constant
- When $\epsilon = 0$ the model is similar to the standard coupled dark energy model

Numerical Analysis



- Tracker behavior of the field is depicted
- Successive sequence of radiation, matter and dark energy epochs are obtained

Cosmological Perturbation



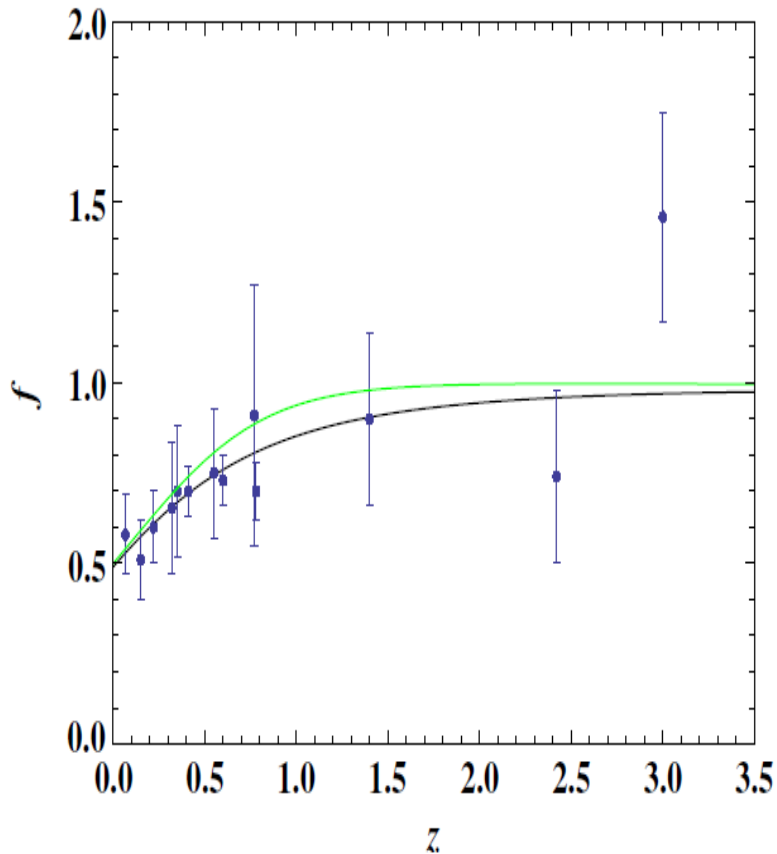
In the de Sitter phase: $G_{\text{eff}} = G(1 + 2\beta^2) > G$

- The model reduces to coupled quintessence scenario at large scale thereby giving strong constraints on β
- G_{eff} is large in the higher redshift which gives different BBN Constraints than in General relativity

Observational Evidences

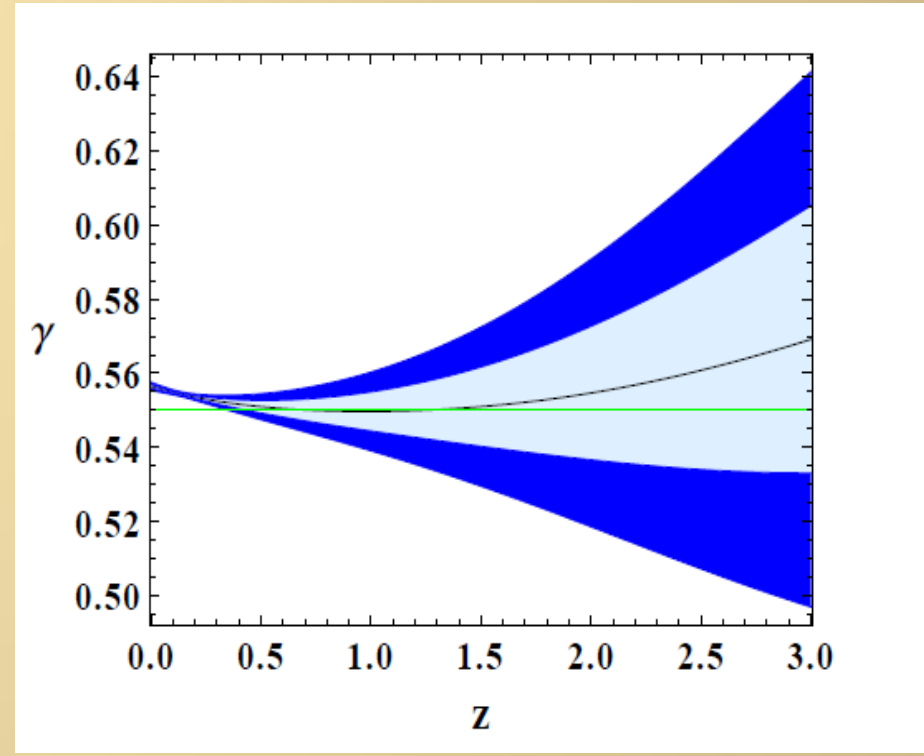
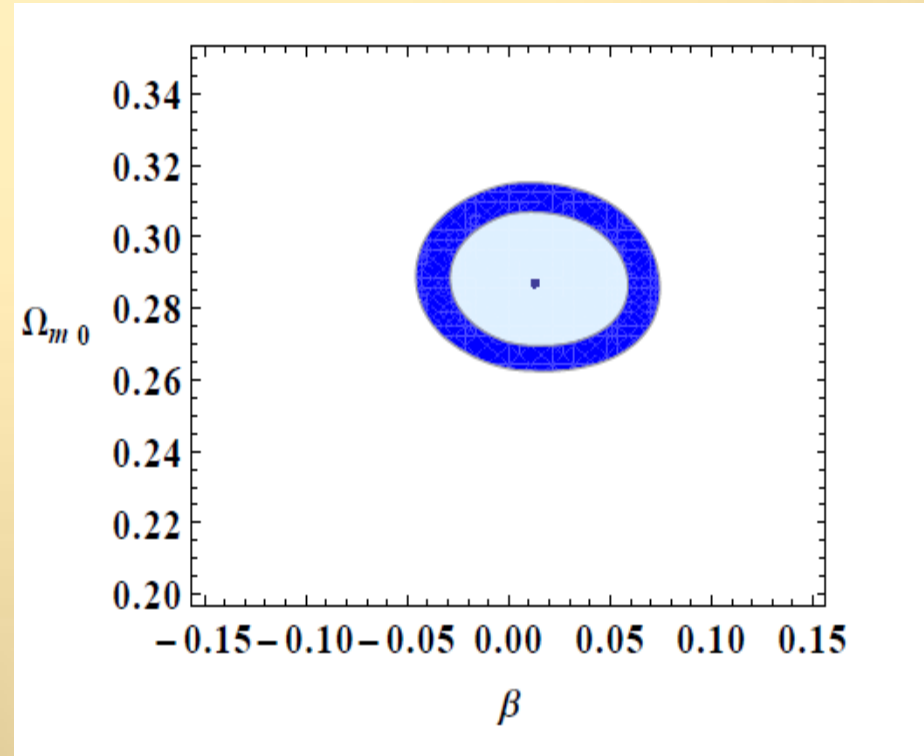
Growth factor:

$$f = \frac{d \ln \delta_m}{d \ln a} .$$



- deviation from Λ CDM model is not significant
- $f < 1$ for all redshift which means growth is in accordance with Einstein- de-Sitter model

$$\chi^2 = \chi_{Growth}^2 + \chi_{SN}^2 + \chi_{BAO}^2 + \chi_{CMB}^2$$



- β is highly constrained
- At present the evolution of γ is consistent with dark energy models
- But the value of γ at all redshift is large compared to that of Λ CDM

Conclusion

- Adding a potential to Slotheon gravity breaks the symmetry but gives a viable cosmology
- With exponential potential the model gives an accelerating universe at late times
- The model is similar to coupled quintessence at late times
- The deviation of growth factor compared to Λ CDM is negligible
- From observational data β is constrained to small range of values and present density of matter is constrained around the concordance values.
- The growth index is large for all redshift compared to Λ CDM

THANK YOU