Inflation and Electroweak vacuum stability after LHC, PLANCK and BICEP2 (JCAP12(2014)001)

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Discoveries

@ Large Hadron Collider :

Higgs Particle Aass 122-127 Gev

Precision measurements of scalar spectral index and primordial non-Gaussinity — > Hinting towards single-field inflationary scenarios

@ BICEP2 :

Claimed (dust?) to have measured (primordial) large tensor-to-scalar ratio — Puts inflationary scale close to GUT scale

Open issues

- Observed Higgs and top masses may lead to <u>unstable electroweak vacuum</u> at higher energy scales (around 10¹⁰ GeV)
- We still do not know the particle physics
 origin of inflaton
- Large tensor-to-scalar ratio observation by BICEP2 might claim that <u>Higgs is not</u> <u>inflaton</u>

Aims

- To include inflaton in a particle
 physics model
- To stablilize the Electroweak
 vacuum all the way upto the Planck
 scale

Model

@ Symmetry Group :

 $SU(3)_C \otimes SU(2)_L \otimes U(1)_Y \otimes U(1)_{B-L}$

 3 extra right handed neutrinos to cancel all the gauge as well as gravitational anomalies, 1 extra gauge boson and 1 extra heavy scalar field

Electroweak symmetry breaks at 256
 GeV and B-L symmetry breaks at GUT scale

- Scalar Lagrangian :
- $\mathcal{L} = (\mathcal{D}_{\mu}S)^{\dagger}(\mathcal{D}^{\mu}S) + (\mathcal{D}_{\mu}\Phi)^{\dagger}(\mathcal{D}^{\mu}\Phi) + m_{s}^{2}(S^{\dagger}S) + m_{\phi}^{2}(\Phi^{\dagger}\Phi)$ $-\lambda_{1}(S^{\dagger}S)^{2} - \lambda_{2}(\Phi^{\dagger}\Phi)^{2} - \lambda_{3}(S^{\dagger}S)(\Phi^{\dagger}\Phi)$
 - The full symmetry group and the low energy effective SM are well separated in energy scales
 - Decoupling theorem presence of new heavy particles of the extended theory would not affect the quartic coupling of the SM Higgs at low scales No Vacuum Stability ??

Threshold effect

- Threshold effect : modifies the evolution of Higgs quartic coupling at lower scale due to presence of a heavy scalar below the instability scale
- Scalar potential :

 $V(S,\Phi) = \lambda_1 (S^{\dagger}S)^2 - m_s^2 (S^{\dagger}S) + \lambda_2 \left(\Phi^{\dagger}\Phi - \frac{1}{2}v_{\phi}^2\right)^2 + \lambda_3 (S^{\dagger}S) \left(\Phi^{\dagger}\Phi - \frac{1}{2}v_{\phi}^2\right)^2$

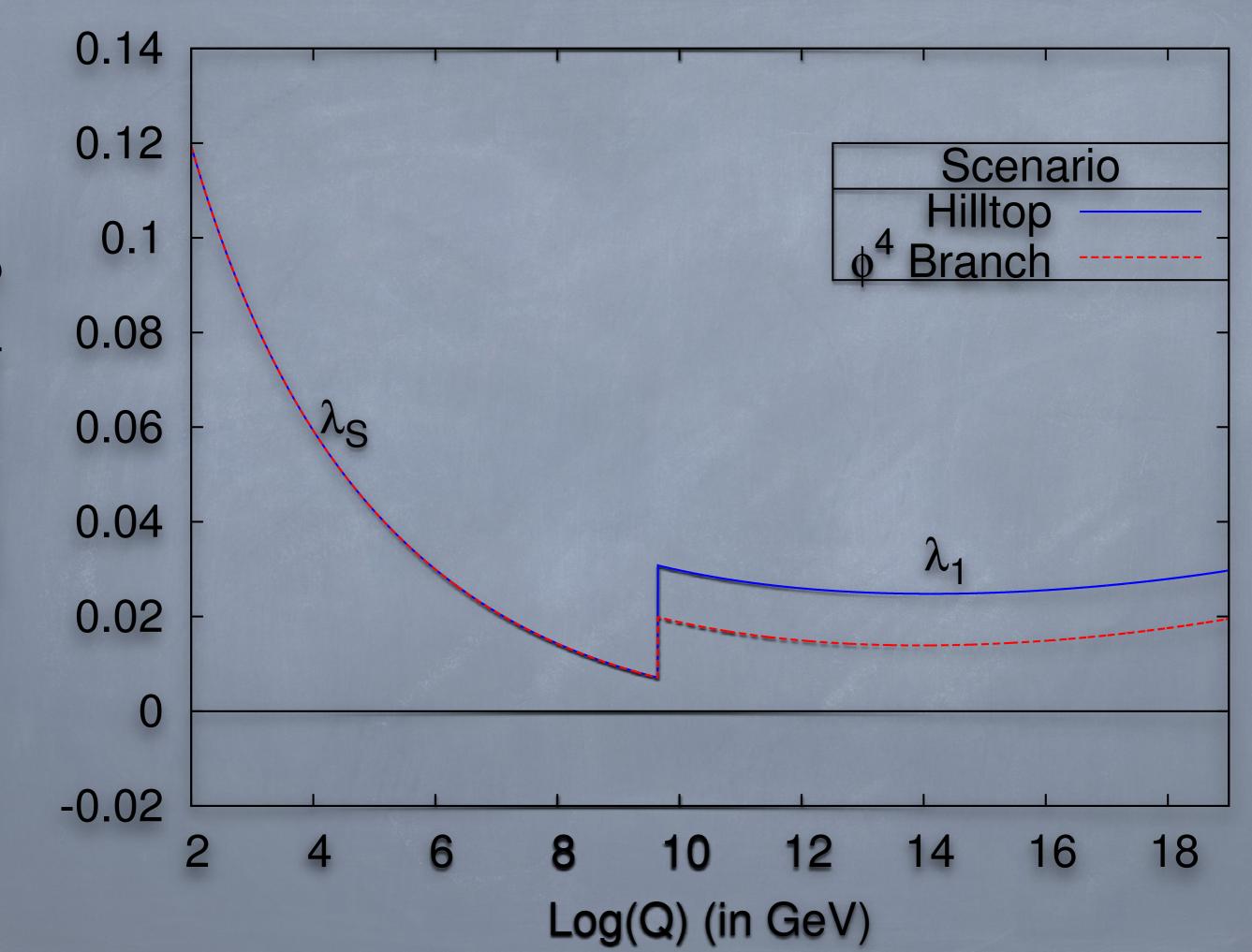
The EOM of the heavy scalar at its minima

$$\Phi^{\dagger}\Phi = \frac{1}{2}v_{\phi}^2 - \frac{\lambda_3}{2\lambda_2}S^{\dagger}S$$

Effective scalar potential below the heavy scalar's mass-scale :

 $V(S)|_{\text{eff}} = \left(\lambda_1 - \frac{\lambda_3^2}{4\lambda_2}\right) (S^{\dagger}S)^2 - m_s^2(S^{\dagger}S)$ $\equiv \lambda_S (S^{\dagger}S)^2 - m_s^2(S^{\dagger}S)$

The Electroweak vacuum is stable all the way upto the Planck scale



Inflation

@ Inflation is driven by the radiatively corrected quartic inflaton potential $V(\phi_0) = \frac{1}{4}\lambda_2(\phi_0^2 - v_{\phi}^2)^2 + a\lambda_2 \log\left(\frac{\phi_0}{v_{\phi}}\right)\phi_0^4$ @ Where the radiatively corrected term has the coefficient $a \equiv \frac{1}{16\pi^2\lambda_2} \left(20\lambda_2^2 + 2\lambda_3^2 + 2\lambda_2 \left(\sum_i (Y_i^{N_R})^2 - 24g_{B-L}^2 \right) + 96g_{B-L}^4 - \sum_i (Y_i^{N_R})^4 \right)$ Define : $u = (1 + a + 4a \ln \phi_0 / v_\phi) / a$ The observables are

$$r = \frac{128M_{\rm Pl}^2}{\phi_0^2} \frac{u^2}{(u-1)^2}$$

$$n_s = 1 - \frac{8M_{\rm Pl}^2}{\phi_0^2} \frac{3u^2 - u + 4}{(u-1)^2}$$

$$\frac{dn_s}{d\ln k} = -\frac{64M_{\rm Pl}^4}{\phi_0^4} \left[\frac{u(3u^3 - 4u^2 + 15u + 10)}{(u-1)^4} \right]$$

$$\mathcal{P}_{\mathcal{R}} = \frac{\lambda_2}{768\pi^2} \left(\frac{\phi_0}{M_{\rm Pl}}\right)^6 \frac{a(u-1)^3}{u^2}$$

Hillop inflation

 $u_* = -0.333$ $r_* = 0.015$ $\lambda_2 \sim 1.89 \times 10^{-13}$ $m_{\phi} \sim 4.3 \times 10^9 \,\mathrm{GeV}$ $\frac{dn_s}{d\ln k}\Big|_{k_*} \sim 1.07 \times 10^{-4}$

In accordance with PLANCK observations

ϕ^4 -branch Inflation $u_* = -11.001$ $r_* = 0.203$

 $\begin{array}{ll} \lambda_2 & \sim & 3.6 \times 10^{-13} \\ m_{\phi} & \sim & 6.0 \times 10^9 \, \mathrm{GeV} \\ \frac{dn_s}{d \ln k} \bigg|_{k_*} & \sim & -5.6 \times 10^{-4} \end{array}$

@ In accordance with PLANCK & BICEP2

Summary

- Aims : to stabilize the electroweak vacuum all the way upto Planck scale and accommodate inflaton in the particle picture
- o We extend the SM by $U(1)_{B-L}$ symmetry
- Due to presence of an extra heavy scalar Higgs quartic coupling receives a threshold correction which helps keep the vacuum stable till the Planck scale
- The real part of the new heavy scalar is the inflaton
- Inflation can occur in two different branches which are in accordance with observation.

Thank You !!