



Lecture 7 Lattice design with MAD-X

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Lecture 8 – MADX introduction and examples

- This lecture is based on one by Ted Wilson (CERN) which was in turn based on a lecture by V. Ziemann (Uppsala University)
 - Installing/running MADX (on windows)
 - Input of elements and beamlines
 - Beta functions, tunes, dispersion
 - Matching
 - Examples

All credit to V. Ziemann for example input files More examples are on the MAD-X website!

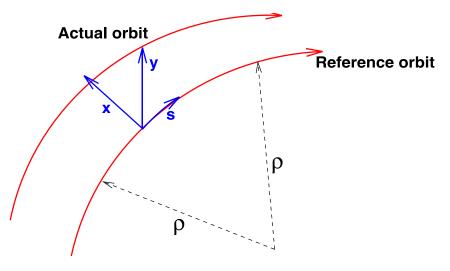


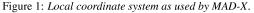


What is MAD-X?

"A program for accelerator design and simulation with a long history" Developed from previous versions (MAD, MAD-8, then finally MAD-X in 2002). User guide: http://mad.web.cern.ch/mad/uguide.html

- Uses a sequence of elements placed sequentially along a *reference orbit*
- *Reference orbit* is path of a charged particle having the central design momentum of the accelerator through idealised magnets (no fringe fields)
- The reference orbit consists of a series of straight line segments and circular arcs
- local curvilinear right handed coordinate system (x, y, s)









What can it do?

- You can input magnets (& electrostatic elements) according to the manual
 - User guide: http://madx.web.cern.ch/madx/
- Calculate beta functions, tune, dispersion, chromaticity, momentum compaction numerically.
- Generates tables and plots (.ps)
- (tip: you might need to install ghostscript to view plots)





Why choose MAD-X?

- There are any number of tracking and beam optics codes, but MAD-X is *widely used* (especially at CERN), *well maintained* and *well documented*.
 - The more lattice design you do the more you will appreciate this about MAD-X!
- What it can't do:
 - Acceleration and tracking simultaneously
 - Not so accurate at large excursions from closed orbit (as in an FFAG)
 - Complicated magnet geometries
 - Field maps





Installing MADX

Go to 'Updates'

Production version 5.02.00

Choose your system & download:

NOTE: You may need to make it executable

-rw-r--r--@ 1 Suzie staff 9.7M 4 Nov 11:41 madx-macosx64 suzie-imac:2013Test Suzie\$ chmod 777 madx-macosx64 suzie-imac:2013Test Suzie\$ ls -lh total 19968 -rwxrwxrwx@ 1 Suzie staff 9.7M 4 Nov 11:41 madx-macosx64 suzie-imac:2013Test Suzie\$

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European Laboratory	for Particle Physics			
MAD - N	ABP Accelerator Beam Physics (
ntroduction	Index of /madx	/releases/las	st-p	ro
Archives				
Bugs & requests Communication	Name	Last modified	Size	Description
Contributors	3			
Development	Parent Directory		-	
ocumentation	The second secon		7.8M	
Project	The second secon	15-Mar-2013 11:58	9.2M	Linux 32 bit library
Reference	-	15-Mar-2013 11:58	11M	Linux 64 bit library
Releases Resources	-	15-Mar-2013 11:58	7.9M	Linux 64 bit library
Roadmap	1ibmadx-macosx32.a	15-Mar-2013 11:57	9.4M	MacOS X 32 bit library
Jpdates		15-Mar-2013 11:57	8.4M	MacOS X 32 bit library
shortcuts	libnadx-macosx64.a	15-Mar-2013 11:57	11M	MacOS X 64 bit library
subject index	1ibmadx-macosx64.dylib	15-Mar-2013 11:57	9.8M	MacOS X 64 bit library
user's guide	1ibptc-linux32.a	15-Mar-2013 11:58	4.1M	Linux 32 bit library
examples	1ibptc-linux32.so	15-Mar-2013 11:58	4.4M	Linux 32 bit library
updates events	1ibptc-linux64.a	15-Mar-2013 11:58	5.7M	Linux 64 bit library
	1ibptc-linux64.so	15-Mar=2013 11:58	4.7M	Linux 64 bit library
	1ibptc-macosx32.a	15-Mar-2013 11:57	5.4M	MacOS X 32 bit library
	1ibptc-macosx32.dylib	15-Mar-2013 11:57	4.8M	MacOS X 32 bit library
	1ibptc-macosx64.a	15-Mar-2013 11:57	6.1M	MacOS X 64 bit library
	1ibptc-macosx64.dylib	15-Mar-2013 11:57	5.7M	MacOS X 64 bit library
	madx-info.txt	15-Mar-2013 11:57	73	SVN release information
	2 madx-linux32	15-Mar-2013 11:58	11M	Linux 32 bit binary
	2 madx-linux64	15-Mar-2013 11:58	13M	Linux 64 bit binary
	2 madx-macosx32	15-Mar-2013 11:57	8.4M	MacOS X 32 bit binary
	2 madx-macogx64	15-Mar-2013 11:57	9.7M	MacOS X 64 bit binary
	adx-src.tgz	15-Mar=2013 11:57	14M	Source code tarball
	madx-win32.exe	15-Mar-2013 11:57	6.9M	Windows 32 bit binary
	madx-win64.exe	15-Mar-2013 11:57	8.6M	Windows 64 bit binary
	nundiff-linux32	15-Mar-2013 11:58	164K	Linux 32 bit binary
	nundiff-linux64		184K	
	nundiff-macosx32	15-Mar-2013 11158	187K	
	-			NacOS X 32 bit binary
	10	15-Mar-2013 11:57	205K	MacOS X 64 bit binary
	10-	15-Mar-2013 11:57	244K	Windows 32 bit binary
	nundiff-win64.exe	15-Mar-2013 11:57	282K	Windows 64 bit binary
updated 2013.10.21				
visits: 102875				
mad support copyright				
W3C HTML				





How to run MADX

- In command prompt:
- Go to directory (with madx.exe and input files)
- madx.exe < inputfile > outputfile
- Or you can add the madx.exe location to your path (if you know how...)





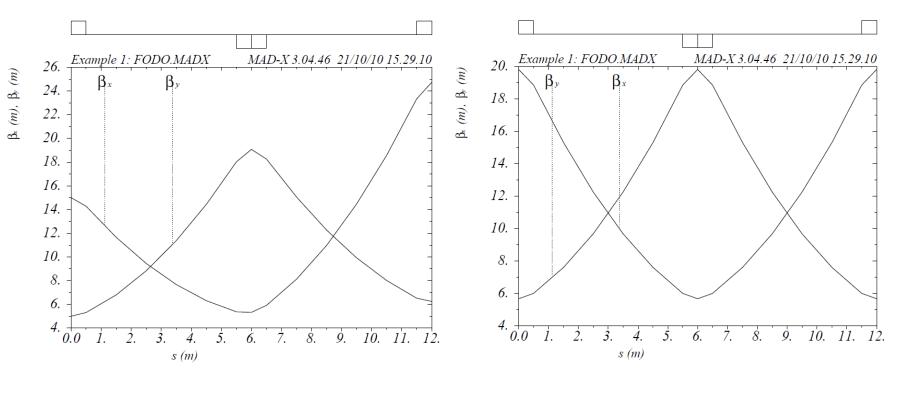
An input file – Basic FODO

	// comments out a line
<pre>// MADX Example 1: FODO cell // Author: V. Ziemann, Uppsala University // Date: 060910 // Upperson Currents Currents of (11 (2012, EOD, MADY))</pre>	TITLE at top of output
<pre>// UPDATED SUZIE SHEEHY 04/11/2013 FOR MADX ` TITLE,'Example 1: FOD0.MADX'; BEAM, PARTICLE=ELECTRON,PC=3.0;</pre>	Define particle type and momentum (pc) GeV/c
D: DRIFT,L=1.0; QF: QUADRUPOLE,L=0.5,K1=0.2; QD: QUADRUPOLE,L=0.5,K1=-0.2;	—— Elements are given in the manual including definitions of L, K1 etc
FODO: LINE=(QF,5*(D),QD,QD,5*(D),QF); <	Define a 'line' –can be a cell or a whole beamline that you will USE.
USE, PERIOD=FODO; TWISS,SAVE,BETX=15.0,BETY=5.0; PLOT,HAXIS=S, VAXIS=BETX, BETY, NOVERSION=TRU	 Calculate beta functions from starting values UE, TITLE='unmatched beta functions';
//Here MATCH is used as a single command this USE, PERIOD=FODO; MATCH, SEQUENCE=FODO; TWISS, SAVE; PLOT,HAXIS=S, VAXIS=BETX, BETY, NOVERSION=TRU	TWISS (internal table)
<pre>Value, TABLE(SUMM,Q1); Value, TABLE(SUMM,Q2); WRITE,TABLE=SUMM,FILE=print.dat;</pre>	——— Match the periodic solution (+Plot that)
5/11/14	Output to tables





Result of a MADX run



With βx starting at 15m (βy at 5m)

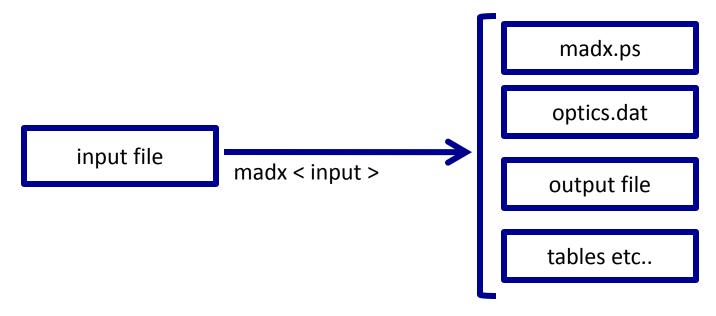
If the line is matched (periodic) $\beta_{\text{start}} = \beta_{\text{end}}$





Output files

- Optics.dat
- Your specified output file madx< input > output
- Can specify tables







Add bending magnets

- Can introduce your own parameters (watch the :=)
- Can use alternative 'sequence' format
- Let's add dipoles
- Look at dispersion

TITLE, 'Example 2: FODO2.MADX'; BEAM, PARTICLE=ELECTRON, PC=3.0; DEGREE:=PI/180.0; // for readability

QF: QUADRUPOLE,L=0.5,K1=0.2; // still half-length QD: QUADRUPOLE,L=1.0,K1=-0.2; // changed to full length B: SBEND,L=1.0,ANGLE=15.0*DEGREE; // added dipole

FODO: SEQUENCE, REFER=ENTRY, L=12.0; QF1: QF, AT=0.0; B1: B, AT=2.5; QD1: QD, AT=5.5; B2: B, AT=8.5; QF2: QF, AT=11.5; ENDSEQUENCE;

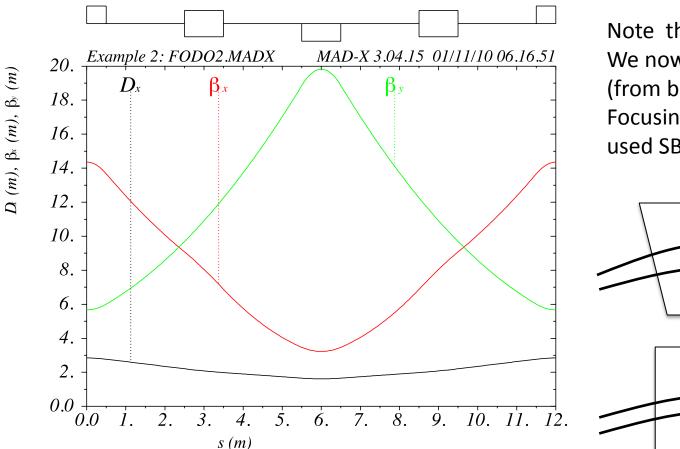
USE, PERIOD=FODO; //MATCH, SEQUENCE=FODO; //Uncomment to match SELECT,FLAG=SECTORMAP,clear; SELECT,FLAG=TWISS,column=name,s,betx,bety; TWISS, file=optics.dat,sectormap;

PLOT, HAXIS=S, COLOUR=100, VAXIS=DX, BETX, BETY, INTERPOLATE=TRUE;

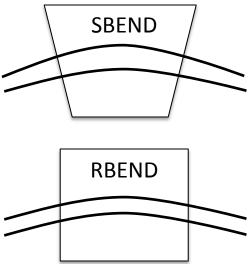
Value, TABLE(SUMM,Q1); Value, TABLE(SUMM,Q2);



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Note the colours! We now have dispersion (from bending magnet) Focusing changed as we used SBEND







	000			optics.dat	📄 optics.dat		
	NAME	%05s "TWISS"					
	a TYPE	%05s "TWISS"					
	@ SEQUENCE	%04s "FODO"					
	@ PARTICLE	%08s "ELECTRON"					
	@ MASS	%le 0.000510	0998902				
	@ CHARGE	%le	-1				
	@ ENERGY		0000044				
	@ PC	%le	3				
	@ GAMMA		.854187				
	<pre>@ KBUNCH @ BCURRENT</pre>	%le %le	1 0				
	@ SIGE	%le	0 A				
	@ SIGE @ SIGT	%le	0				
		%le	8				
SELECT, FLAG=SECTOF	@ EX	%le	1				
	@ EY	%le	1				
SELECT, FLAG=TWISS,	@ ET	%le	1				
	@ LENGTH	%le	12				
TWISS, file=optics	@ ALFA	%le 0.0910	0331259				
-	@ ORBIT5	%le	-0				
NZ •11	@ GAMMATR		4364527				
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	@ Q2		3459932				
	@ DQ1 @ DQ2		2485772 6020063				
	@ DXMAX		4272104				
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	@ XCOMAX	%le	õ				
	@ YCOMAX	%le	õ				
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 and an out 	@ BETYMAX	%le 18.85	5637615				
	@ XCORMS	%le	0				
	@ YCORMS	%le	0				
	@ DXRMS		8159047				
	@ DYRMS	%le	0				
	@ DELTAP	%le	0 0				
	@ SYNCH_1 @ SYNCH_2	%le %le	0 A				
You can customise	a SYNCH 3	%le	0				
Tou can customise	@ SYNCH_4	%le	ñ				
	A SYNCH 5	%le	õ				
<pre>select,flag=my sec</pre>	@ TITLE	%21s "Example 2:	FODO2.MADX"				
—	@ URIGIN	%20s "MAD-X 3.04.	.15 Darwin"				
Or even select by comp	@ DATE	%08s "01/11/10"					
		%08s "08.52.11"					
coloct flog-my coc	* NAME		S BETX				
<pre>select,flag=my_sec</pre>	3 %S "EODO#CTID	T.I.	%le %le				
	"FODO\$STAR "OF1"	1	0 14.36014477 0.5 13.67114589	5.674619595 6.007906055			
	"DRIFT_0"		2.5 9.081076899				
	"B1"		3.5 6.880079877	12.25970476			
	"DRIFT_0"		5.5 3.471640314	18.85637615			
	"QD1"		6.5 3.471640314	18.85637615		-	
	"DRIFT_0"		8.5 6.880079877	12.25970476		Ψ.	
5/11/14	"B2"	Latti	iĉe desigh የምም	1AD-X ^{9.690010744}		11.	
-, ,		_0.00.					





Matching

- Matching lets MAD-X do the tedious work for you!
- Before MATCH select at least one sequence (USE)
- Initiated by the MATCH command
- Initiating: MATCH, SEQUENCE='name1', 'name2',..,'nameX';
- Can define constraints & variables (magnets) to achieve aim MATCH, SEQUENCE = FODO; CONSTRAINT,SEQUENCE=FODO, RANGE=#E, MUX=0.16666666, MUY=0.25; VARY, NAME=QF->K1, STEP=1E-6; VARY, NAME=QD->K1, STEP=1E-6; LMDIF,CALLS=500,TOLERANCE=1E-20; ENDMATCH;





Matching input file

TITLE, 'Example 3: MATCH1.MADX'; BEAM, PARTICLE=ELECTRON, PC=3.0;

D: DRIFT,L=1.0;

- QF: QUADRUPOLE, L=0.5, K1:=0.2;
- QD: QUADRUPOLE, L=0.5, K1:=-0.2;

FODO: LINE=(QF,5*(D),QD,QD,5*(D),QF); USE, PERIOD=FODO;

//...match phase advance at end of cell to 60 and 90 degrees
MATCH, SEQUENCE=FODO;
CONSTRAINT,SEQUENCE=FODO,RANGE=#E,MUX=0.1666666666,MUY=0.25;
VARY,NAME=QF->K1,STEP=1E-6;
VARY,NAME=QD->K1,STEP=1E-6;
LMDIF,CALLS=500,TOLERANCE=1E-20;
ENDMATCH;

SELECT,FLAG=SECTORMAP,clear; SELECT,FLAG=TWISS,column=name,s,betx,alfx,bety,alfy; TWISS, file=optics.dat,sectormap;

PLOT,HAXIS=S, VAXIS=BETX, BETY; Value, TABLE(SUMM,Q1); // verify result Value, TABLE(SUMM,Q2);

Print out final values of matching

Matching commands





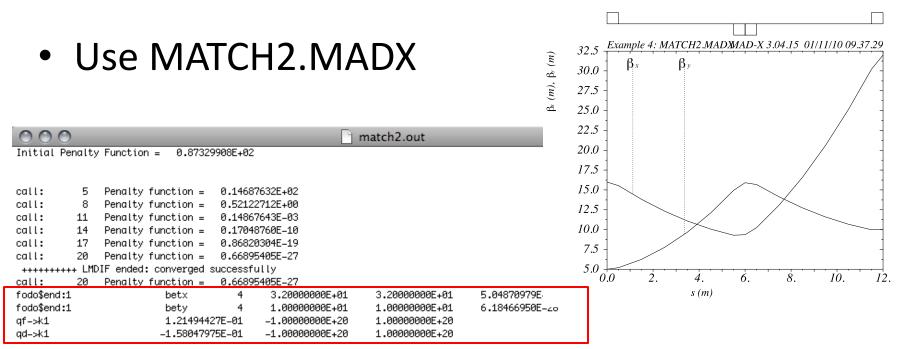
Matching example

• Demonstration MATCH1.MADX





Fitting beta functions



GXPLOT-X11 1.50 initialized

plot number = 1

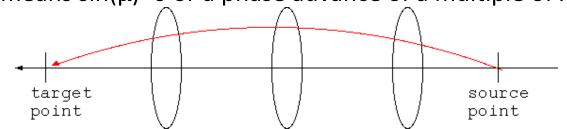
// MADX Example 1: FODO matching final beta function





Transfer matrix matching

- Sometimes want to constrain transfer matrix elements to some value.
- For example $R_{16}=0$ and $R_{26}=0$ will make the horizontal position and angle independent of the momentum after a beamline.
- This is called an 'Achromat'.
- Other versions are imaginable
- point-to-point imaging \rightarrow R12 = 0.
 - This means $sin(\mu)=0$ or a phase advance of a multiple of π .







Examples in MAD-X

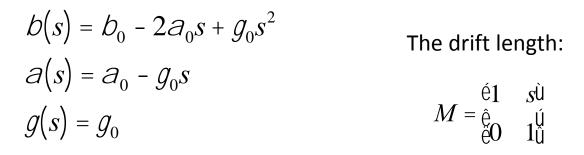
- FODO arcs
- Dispersion suppressor
- 'Telescopes' for low-β
- Synchrotron radiation lattices + achromats



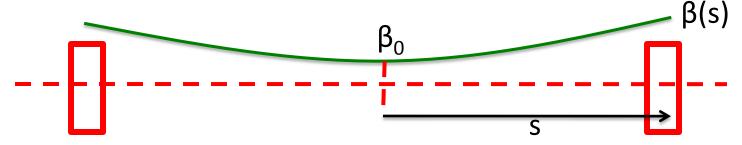


Is that it?

- 'the not-so-ideal world'
- What happens to α, β, γ if we stop focusing for a distance?



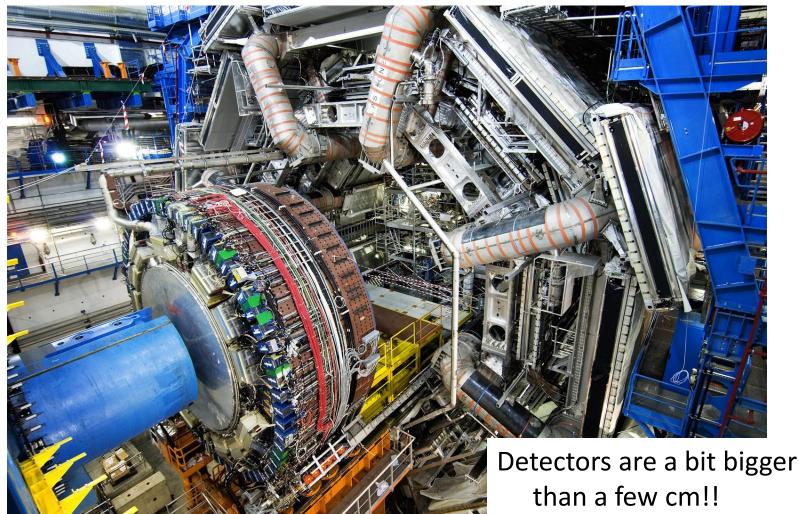
- If we take the center of a drift ($\alpha_0=0$), we find $b(s) = b_0 + \frac{s^2}{b_0}$
- It doesn't matter what you do $-\beta$ will grow!







Seems fine, until...







FODO Arcs

- Usually in colliders take beam between interaction regions
- Simple and tunable (β_x large in QF β_v at QD)
- Moderate quad strengths
- The beam is not round
- In arcs dipoles generate dispersion

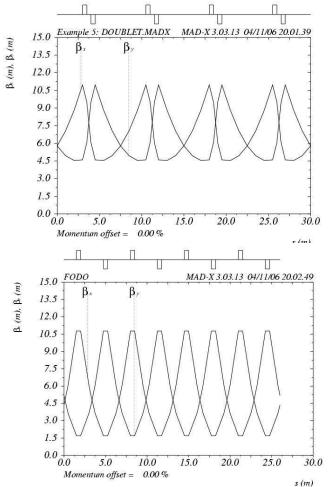
In QD





FD Doublet Lattice

- More space between quads
- Stronger quad strengths
- Round beams
- Used in CTF3 linac

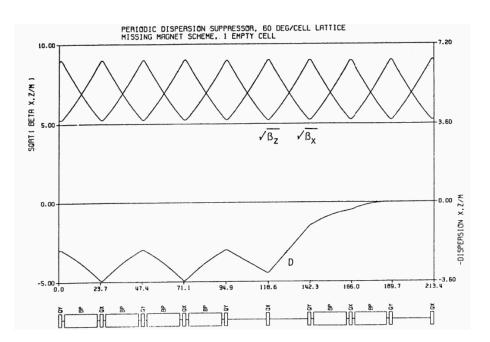






Dispersion suppressor

- Want small spot size at interaction point
- Spot size: $\sqrt{\epsilon\beta + (D\Delta p/p)^2}$
- Missing magnet dispersion suppression scheme
- Works with proper phase advance between elements

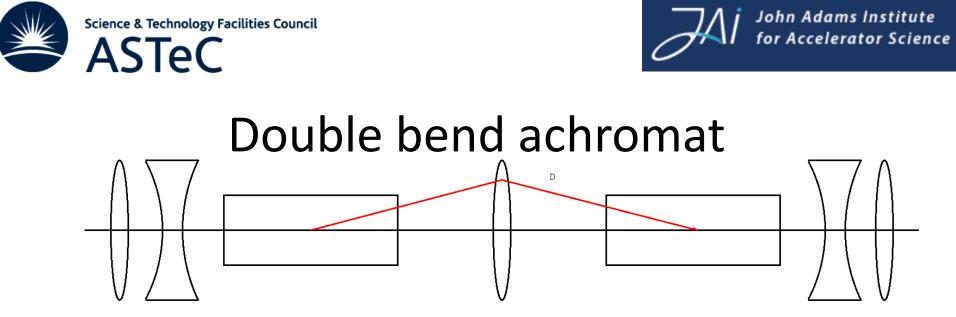






Telescope and low β Source f2 plane f1 point-to-point IΡ parallel-to-parallel 12=f2 f2 11=f1 f1 $\begin{array}{ccc} \mathfrak{X} & \mathbf{0} & f_1 \ddot{\mathbf{0}} \\ \mathbf{C} \\ \dot{\mathbf{C}} - 1/f_1 & \mathbf{0} \dot{\mathbf{g}} \end{array}$ For one module with $I_1 = f_1$ $R = \frac{\mathfrak{E} - f_2 / f_1}{\mathfrak{C}} \qquad 0 \qquad \ddot{\mathcal{C}}$ For both modules:

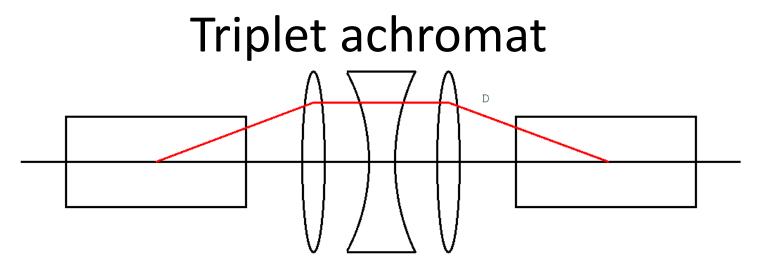
- Used in colliders to achieve small beam size at IP
- Doublet or Triplet
- Point-to-point R₁₂=0
- Parallel to parallel R₂₁=0
- R₁₁=demagnification
- Ratio of focal lengths
- Needs to work in both planes with doublets/triplets



- One dipole generates dispersion and the next, which is 180 degrees apart will take it out again
- Remember: the dispersion is the orbit of a particle with slightly too high momentum w.r.t the reference particle
- Quadrupoles are used to make βx in dipoles small





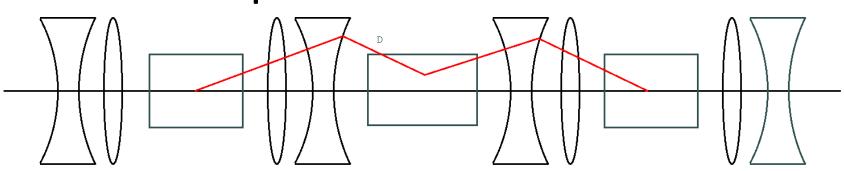


- Do the 180 degrees in the horizontal plane and the beta matching by quads between dipoles
- very compact, few magnets, but not flexible





Triple bend achromat



- Small emittance.
- Very flexible due to large number of quadrupoles.
- Adjacent drift space can be made long to accommodate undulators/wigglers.





Resources

- Many examples available at the MAD-X website
 - A helpful 'primer' by W. Herr:

http://madx.web.cern.ch/madx/doc/madx primer.pdf

You can always ask me or another lecturer – though we can't promise to know the answer!