

Holography for $\mathcal{N} = 2^*$ on S^4

Nikolay Bobev

Instituut voor Theoretische Fysica, K.U. Leuven

CERN

Nov 25 2014

1311.1508 + 14XX.XXXX + in progress

with **Henriette Elvang**, **Daniel Freedman**, **Silviu Pufu**
Uri Kol, **Tim Olson**

- Powerful exact results for supersymmetric field theories on curved manifolds using equivariant localization [Witten], [Nekrasov], [Pestun], ...

Motivation

- Powerful exact results for supersymmetric field theories on curved manifolds using equivariant localization [Witten], [Nekrasov], [Pestun], ...
- The partition function of the theory may be viewed as some “topological invariant” of the manifold.

Motivation

- Powerful exact results for supersymmetric field theories on curved manifolds using equivariant localization [Witten], [Nekrasov], [Pestun], ...
- The partition function of the theory may be viewed as some “topological invariant” of the manifold.
- Supersymmetric localization reduces the path integral of a gauge theory to a finite dimensional matrix integral. Still hard to evaluate explicitly in general!

Motivation

- Powerful exact results for supersymmetric field theories on curved manifolds using equivariant localization [Witten], [Nekrasov], [Pestun], ...
- The partition function of the theory may be viewed as some “topological invariant” of the manifold.
- Supersymmetric localization reduces the path integral of a gauge theory to a finite dimensional matrix integral. Still hard to evaluate explicitly in general!
- Make progress by taking the planar limit for specific 4d $\mathcal{N} = 2$ theories. [Russo], [Russo-Zarembo], [Buchel-Russo-Zarembo]

- Powerful exact results for supersymmetric field theories on curved manifolds using equivariant localization [Witten], [Nekrasov], [Pestun], ...
- The partition function of the theory may be viewed as some “topological invariant” of the manifold.
- Supersymmetric localization reduces the path integral of a gauge theory to a finite dimensional matrix integral. Still hard to evaluate explicitly in general!
- Make progress by taking the planar limit for specific 4d $\mathcal{N} = 2$ theories. [Russo], [Russo-Zarembo], [Buchel-Russo-Zarembo]
- Evaluation of the partition function of planar $SU(N)$, $\mathcal{N} = 2^*$ SYM on S^4 . An infinite number of quantum phase transitions as a function of $\lambda = g_{YM}^2 N$. [Russo-Zarembo]

- $\mathcal{N} = 2^*$ SYM is a theory of a $\mathcal{N} = 2$ vector multiplet and a hyper multiplet in the adjoint. It is a massive deformation of $\mathcal{N} = 4$ SYM. There is a unique supersymmetric Lagrangian on S^4 . [Pestun], [Festuccia-Seiberg]

Motivation

- $\mathcal{N} = 2^*$ SYM is a theory of a $\mathcal{N} = 2$ vector multiplet and a hyper multiplet in the adjoint. It is a massive deformation of $\mathcal{N} = 4$ SYM. There is a unique supersymmetric Lagrangian on S^4 . [Pestun], [Festuccia-Seiberg]
- The result for $N, \lambda \gg 1$ is [Russo-Zarembo]

$$F_{S^4}^{\mathcal{N}=2^*} = -\log \mathcal{Z}_{S^4}^{\mathcal{N}=2^*} = -\frac{N^2}{2} (1 + (mR)^2) \log \frac{\lambda(1 + (mR)^2)e^{2\gamma + \frac{1}{2}}}{16\pi^2},$$

Motivation

- $\mathcal{N} = 2^*$ SYM is a theory of a $\mathcal{N} = 2$ vector multiplet and a hyper multiplet in the adjoint. It is a massive deformation of $\mathcal{N} = 4$ SYM. There is a unique supersymmetric Lagrangian on S^4 . [Pestun], [Festuccia-Seiberg]
- The result for $N, \lambda \gg 1$ is [Russo-Zarembo]

$$F_{S^4}^{\mathcal{N}=2^*} = -\log \mathcal{Z}_{S^4}^{\mathcal{N}=2^*} = -\frac{N^2}{2} (1 + (mR)^2) \log \frac{\lambda(1 + (mR)^2)e^{2\gamma + \frac{1}{2}}}{16\pi^2},$$

- The goal is to calculate $F_{S^4}^{\mathcal{N}=2^*}$ holographically.

Motivation

- $\mathcal{N} = 2^*$ SYM is a theory of a $\mathcal{N} = 2$ vector multiplet and a hyper multiplet in the adjoint. It is a massive deformation of $\mathcal{N} = 4$ SYM. There is a unique supersymmetric Lagrangian on S^4 . [Pestun], [Festuccia-Seiberg]
- The result for $N, \lambda \gg 1$ is [Russo-Zarembo]

$$F_{S^4}^{\mathcal{N}=2^*} = -\log \mathcal{Z}_{S^4}^{\mathcal{N}=2^*} = -\frac{N^2}{2} (1 + (mR)^2) \log \frac{\lambda(1 + (mR)^2)e^{2\gamma + \frac{1}{2}}}{16\pi^2},$$

- The goal is to calculate $F_{S^4}^{\mathcal{N}=2^*}$ holographically.
- Precision test of holography! In AdS_5/CFT_4 one typically compares numbers. Here we have a whole function to match.

Motivation

- $\mathcal{N} = 2^*$ SYM is a theory of a $\mathcal{N} = 2$ vector multiplet and a hyper multiplet in the adjoint. It is a massive deformation of $\mathcal{N} = 4$ SYM. There is a unique supersymmetric Lagrangian on S^4 . [Pestun], [Festuccia-Seiberg]

- The result for $N, \lambda \gg 1$ is [Russo-Zarembo]

$$F_{S^4}^{\mathcal{N}=2^*} = -\log \mathcal{Z}_{S^4}^{\mathcal{N}=2^*} = -\frac{N^2}{2} (1 + (mR)^2) \log \frac{\lambda(1 + (mR)^2)e^{2\gamma + \frac{1}{2}}}{16\pi^2},$$

- The goal is to calculate $F_{S^4}^{\mathcal{N}=2^*}$ holographically.
- Precision test of holography! In AdS_5/CFT_4 one typically compares numbers. Here we have a whole function to match.
- Previous results on holography for $\mathcal{N} = 2^*$ on \mathbb{R}^4 . [Pilch-Warner], [Buchel-Peet-Polchinski], [Evans-Johnson-Petrini] On S^4 the holographic construction is more involved.

- The $\mathcal{N} = 2^*$ SYM theory on S^4
- The gravity dual
- Holographic calculation of $F_{S^4}^{\mathcal{N}=2^*}$
- Comments on $\mathcal{N} = 1^*$ SYM theory on S^4
- Outlook

The $\mathcal{N} = 2^*$ SYM theory on S^4

$\mathcal{N} = 2^*$ SYM on S^4

The field content of $\mathcal{N} = 4$ SYM is

$$A_\mu, \quad X_{1,2,3,4,5,6}, \quad \lambda_{1,2,3,4}.$$

$\mathcal{N} = 2^*$ SYM on S^4

The field content of $\mathcal{N} = 4$ SYM is

$$A_\mu, \quad X_{1,2,3,4,5,6}, \quad \lambda_{1,2,3,4}.$$

Organize this into an $\mathcal{N} = 2$ vector multiplet

$$A_\mu, \quad \psi_1 \equiv \lambda_4, \quad \psi_2 \equiv \lambda_3, \quad Z_3 \equiv \frac{1}{\sqrt{2}}(X_3 + iX_6),$$

and one hypermultiplet

$$\chi_j = \lambda_j, \quad Z_j = \frac{1}{\sqrt{2}}(X_j + iX_{j+3}), \quad j = 1, 2.$$

$\mathcal{N} = 2^*$ SYM on S^4

The field content of $\mathcal{N} = 4$ SYM is

$$A_\mu, \quad X_{1,2,3,4,5,6}, \quad \lambda_{1,2,3,4}.$$

Organize this into an $\mathcal{N} = 2$ vector multiplet

$$A_\mu, \quad \psi_1 \equiv \lambda_4, \quad \psi_2 \equiv \lambda_3, \quad Z_3 \equiv \frac{1}{\sqrt{2}}(X_3 + iX_6),$$

and one hypermultiplet

$$\chi_j = \lambda_j, \quad Z_j = \frac{1}{\sqrt{2}}(X_j + iX_{j+3}), \quad j = 1, 2.$$

Only $SU(2)_V \times SU(2)_H \times U(1)_R$ of the $SO(6)$ R-symmetry is manifest.

$\mathcal{N} = 2^*$ SYM on S^4

The field content of $\mathcal{N} = 4$ SYM is

$$A_\mu, \quad X_{1,2,3,4,5,6}, \quad \lambda_{1,2,3,4}.$$

Organize this into an $\mathcal{N} = 2$ vector multiplet

$$A_\mu, \quad \psi_1 \equiv \lambda_4, \quad \psi_2 \equiv \lambda_3, \quad Z_3 \equiv \frac{1}{\sqrt{2}}(X_3 + iX_6),$$

and one hypermultiplet

$$\chi_j = \lambda_j, \quad Z_j = \frac{1}{\sqrt{2}}(X_j + iX_{j+3}), \quad j = 1, 2.$$

Only $SU(2)_V \times SU(2)_H \times U(1)_R$ of the $SO(6)$ R-symmetry is manifest.

The $\mathcal{N} = 2^*$ theory is obtained by giving a mass term for the hyper multiplet. In flat space this breaks the global symmetry to $SU(2)_V \times U(1)_H$.

$\mathcal{N} = 2^*$ SYM on S^4

The theory is no longer conformal so it is not obvious how to put it on S^4 .

$\mathcal{N} = 2^*$ SYM on S^4

The theory is no longer conformal so it is not obvious how to put it on S^4 .

When there is a will there is a way! [Pestun], [Festuccia-Seiberg], ...

$\mathcal{N} = 2^*$ SYM on S^4

The theory is no longer conformal so it is not obvious how to put it on S^4 .

When there is a will there is a way! [Pestun], [Festuccia-Seiberg], ...

$$\begin{aligned}\mathcal{L}_{\mathcal{N}=2^*}^{S^4} &= \mathcal{L}_{\mathcal{N}=4}^{S^4} + \frac{2}{R^2} \operatorname{tr}(Z_1 \tilde{Z}_1 + Z_2 \tilde{Z}_2 + Z_3 \tilde{Z}_3) \\ &\quad + m^2 \operatorname{tr}(Z_1 \tilde{Z}_1 + Z_2 \tilde{Z}_2) - \frac{m}{2} \operatorname{tr}(\chi_1 \chi_1 + \chi_2 \chi_2 + \tilde{\chi}_1 \tilde{\chi}_1 + \tilde{\chi}_2 \tilde{\chi}_2) \\ &\quad - \sqrt{2} m \operatorname{tr} \left[(\tilde{Z}_1 Z_2 - \tilde{Z}_2 Z_1) Z_3 + (Z_1 \tilde{Z}_2 - Z_2 \tilde{Z}_1) \tilde{Z}_3 \right] \\ &\quad + \frac{im}{2R} \operatorname{tr} (Z_1^2 + Z_2^2 + \tilde{Z}_1^2 + \tilde{Z}_2^2) .\end{aligned}$$

$\mathcal{N} = 2^*$ SYM on S^4

The theory is no longer conformal so it is not obvious how to put it on S^4 .

When there is a will there is a way! [Pestun], [Festuccia-Seiberg], ...

$$\begin{aligned}\mathcal{L}_{\mathcal{N}=2^*}^{S^4} &= \mathcal{L}_{\mathcal{N}=4}^{S^4} + \frac{2}{R^2} \text{tr}(Z_1 \tilde{Z}_1 + Z_2 \tilde{Z}_2 + Z_3 \tilde{Z}_3) \\ &\quad + m^2 \text{tr}(Z_1 \tilde{Z}_1 + Z_2 \tilde{Z}_2) - \frac{m}{2} \text{tr}(\chi_1 \chi_1 + \chi_2 \chi_2 + \tilde{\chi}_1 \tilde{\chi}_1 + \tilde{\chi}_2 \tilde{\chi}_2) \\ &\quad - \sqrt{2} m \text{tr} \left[(\tilde{Z}_1 Z_2 - \tilde{Z}_2 Z_1) Z_3 + (Z_1 \tilde{Z}_2 - Z_2 \tilde{Z}_1) \tilde{Z}_3 \right] \\ &\quad + \frac{im}{2R} \text{tr} \left(Z_1^2 + Z_2^2 + \tilde{Z}_1^2 + \tilde{Z}_2^2 \right) .\end{aligned}$$

Expect to have supergravity scalars dual to

$$\begin{aligned}\mathcal{O}_\phi &= \text{tr}(Z_1 \tilde{Z}_1 + Z_2 \tilde{Z}_2), & \Delta_{\mathcal{O}_\phi} &= 2, \\ \mathcal{O}_\psi &= \text{tr}(\chi_1 \chi_1 + \chi_2 \chi_2 + \tilde{\chi}_1 \tilde{\chi}_1 + \tilde{\chi}_2 \tilde{\chi}_2), & \Delta_{\mathcal{O}_\psi} &= 3, \\ \mathcal{O}_\chi &= \text{tr}(Z_1^2 + Z_2^2 + \tilde{Z}_1^2 + \tilde{Z}_2^2), & \Delta_{\mathcal{O}_\chi} &= 2.\end{aligned}$$

$\mathcal{N} = 2^*$ SYM on S^4

	$SU(2)_V$	$SU(2)_H$	$U(1)_R$
A_μ	0	0	0
Z_3	0	0	+2
$\psi_{1,2}$	1/2	0	+1
$\tilde{\psi}_{1,2}$	1/2	0	-1
$Z_{1,2}$	1/2	1/2	0
$\chi_{1,2}$	0	1/2	-1
$\tilde{\chi}_{1,2}$	0	1/2	+1

In flat space the $\mathcal{N} = 2^*$ theory has $SU(2)_V \times U(1)_H$ global symmetry. The extra coupling \mathcal{O}_χ on S^4 breaks this to $U(1)_V \times U(1)_H$.

$\mathcal{N} = 2^*$ SYM on S^4

	$SU(2)_V$	$SU(2)_H$	$U(1)_R$
A_μ	0	0	0
Z_3	0	0	+2
$\psi_{1,2}$	1/2	0	+1
$\tilde{\psi}_{1,2}$	1/2	0	-1
$Z_{1,2}$	1/2	1/2	0
$\chi_{1,2}$	0	1/2	-1
$\tilde{\chi}_{1,2}$	0	1/2	+1

In flat space the $\mathcal{N} = 2^*$ theory has $SU(2)_V \times U(1)_H$ global symmetry. The extra coupling \mathcal{O}_χ on S^4 breaks this to $U(1)_V \times U(1)_H$.

There is a subtlety. The global symmetry is enhanced to $U(1)_V \times U(1)_H \times U(1)_Y$ in the large N limit. The $U(1)_Y$ is the diagonal subgroup in $U(1)_R \times U(1)_{SL(2,\mathbb{R})}$.

This is proven in flat space. [\[Intriligator\]](#), [\[Intriligator-Skiba\]](#), [\[Pilch-Warner\]](#), [\[Buchel-Peet-Polchinski\]](#)

Expect the same symmetry to emerge on S^4 .

Results from localization

The $SU(N)$ gauge symmetry is generally broken to $U(1)^{N-1}$ by a vev for Z_3 .

$$Z_3 = \text{diag}(a_1, \dots, a_N) .$$

After supersymmetric localization the path integral for the theory on S^4 reduces to a finite dimensional integral over the Coulomb moduli a_i . [Pestun]

$$\mathcal{Z} = \int \prod_{i=1}^N da_i \delta \left(\sum_{i=1}^N a_i \right) \prod_{i < j} (a_i - a_j)^2 \mathcal{Z}_{1\text{-loop}} | \mathcal{Z}_{\text{inst}} |^2 e^{-S_{\text{cl}}} ,$$

where

$$S_{\text{cl}} = \frac{8\pi^2 N}{\lambda} \sum_{i=1}^N a_i^2 , \quad \mathcal{Z}_{1\text{-loop}} = \prod_{i=1}^N \frac{H^2(a_i - a_j)}{H(a_i - a_j + mR) H(a_i - a_j - mR)} ,$$

$$H(x) = \prod_{n=1}^{\infty} \left(1 + \frac{x^2}{n^2} \right)^n e^{-x^2/n} .$$

The function $\mathcal{Z}_{\text{inst}}$ is Nekrasov's partition function with parameters $\varepsilon_1 = \varepsilon_2 = 1/R$.

Results from localization

Russo and Zarembo solved (numerically) this matrix model at large N . They found an infinite number of (quantum) phase transitions as a function of λ .

Results from localization

Russo and Zarembo solved (numerically) this matrix model at large N . They found an infinite number of (quantum) phase transitions as a function of λ .

One caveat. They assume that $\mathcal{Z}_{\text{inst}} = 1$ for large N . This is important and will be checked holographically.

Results from localization

Russo and Zarembo solved (numerically) this matrix model at large N . They found an infinite number of (quantum) phase transitions as a function of λ .

One caveat. They assume that $\mathcal{Z}_{\text{inst}} = 1$ for large N . This is important and will be checked holographically.

The result for $N, \lambda \gg 1$ is

$$F_{S^4} = -\log \mathcal{Z} = -\frac{N^2}{2} (1 + (mR)^2) \log \frac{\lambda(1 + (mR)^2)e^{2\gamma + \frac{1}{2}}}{16\pi^2}.$$

Results from localization

Russo and Zarembo solved (numerically) this matrix model at large N . They found an infinite number of (quantum) phase transitions as a function of λ .

One caveat. They assume that $\mathcal{Z}_{\text{inst}} = 1$ for large N . This is important and will be checked holographically.

The result for $N, \lambda \gg 1$ is

$$F_{S^4} = -\log \mathcal{Z} = -\frac{N^2}{2} (1 + (mR)^2) \log \frac{\lambda(1 + (mR)^2)e^{2\gamma + \frac{1}{2}}}{16\pi^2}.$$

This answer depends on the regularization scheme. The scheme independent quantity is

$$\boxed{\frac{d^3 F_{S^4}}{d(mR)^3} = -2N^2 \frac{mR((mR)^2 + 3)}{((mR)^2 + 1)^2}}$$

This is what one can aim to compute holographically.

For a CFT on S^4 of radius R with cutoff $\epsilon \rightarrow 0$

$$F_{S^4} = \alpha_4 \left(\frac{R}{\epsilon}\right)^4 + \alpha_2 \left(\frac{R}{\epsilon}\right)^2 + \alpha_0 - a_{\text{anom}} \log\left(\frac{R}{\epsilon}\right) + \mathcal{O}(\epsilon/R).$$

Comments

For a CFT on S^4 of radius R with cutoff $\epsilon \rightarrow 0$

$$F_{S^4} = \alpha_4 \left(\frac{R}{\epsilon}\right)^4 + \alpha_2 \left(\frac{R}{\epsilon}\right)^2 + \alpha_0 - a_{\text{anom}} \log \left(\frac{R}{\epsilon}\right) + \mathcal{O}(\epsilon/R).$$

For supersymmetric theories with a supersymmetric regularization scheme $\alpha_4 = 0$.

Comments

For a CFT on S^4 of radius R with cutoff $\epsilon \rightarrow 0$

$$F_{S^4} = \alpha_4 \left(\frac{R}{\epsilon}\right)^4 + \alpha_2 \left(\frac{R}{\epsilon}\right)^2 + \alpha_0 - a_{\text{anom}} \log\left(\frac{R}{\epsilon}\right) + \mathcal{O}(\epsilon/R).$$

For supersymmetric theories with a supersymmetric regularization scheme $\alpha_4 = 0$.

For massive theories there is an extra scale, m , so $\alpha_2 = \alpha_2(m\epsilon)$ and $\alpha_0 = \alpha_0(m\epsilon)$.
Expand this for small $m\epsilon$

$$\alpha_2 = \tilde{\alpha}_2 + m^2 \epsilon^2 \beta_2 + \mathcal{O}(m^4 \epsilon^4),$$

$$\alpha_0 = \tilde{\alpha}_0 + \mathcal{O}(m^2 \epsilon^2).$$

For a CFT on S^4 of radius R with cutoff $\epsilon \rightarrow 0$

$$F_{S^4} = \alpha_4 \left(\frac{R}{\epsilon}\right)^4 + \alpha_2 \left(\frac{R}{\epsilon}\right)^2 + \alpha_0 - a_{\text{anom}} \log\left(\frac{R}{\epsilon}\right) + \mathcal{O}(\epsilon/R).$$

For supersymmetric theories with a supersymmetric regularization scheme $\alpha_4 = 0$.

For massive theories there is an extra scale, m , so $\alpha_2 = \alpha_2(m\epsilon)$ and $\alpha_0 = \alpha_0(m\epsilon)$.
Expand this for small $m\epsilon$

$$\begin{aligned}\alpha_2 &= \tilde{\alpha}_2 + m^2 \epsilon^2 \beta_2 + \mathcal{O}(m^4 \epsilon^4), \\ \alpha_0 &= \tilde{\alpha}_0 + \mathcal{O}(m^2 \epsilon^2).\end{aligned}$$

The nonuniversal contribution to the free energy is

$$\tilde{\alpha}_2 \left(\frac{R}{\epsilon}\right)^2 + \tilde{\alpha}_0 + \beta_2 (mR)^2.$$

Thus 3 derivatives w.r.t. mR eliminate the ambiguity.

For a CFT on S^4 of radius R with cutoff $\epsilon \rightarrow 0$

$$F_{S^4} = \alpha_4 \left(\frac{R}{\epsilon}\right)^4 + \alpha_2 \left(\frac{R}{\epsilon}\right)^2 + \alpha_0 - a_{\text{anom}} \log\left(\frac{R}{\epsilon}\right) + \mathcal{O}(\epsilon/R).$$

For supersymmetric theories with a supersymmetric regularization scheme $\alpha_4 = 0$.

For massive theories there is an extra scale, m , so $\alpha_2 = \alpha_2(m\epsilon)$ and $\alpha_0 = \alpha_0(m\epsilon)$.
Expand this for small $m\epsilon$

$$\begin{aligned}\alpha_2 &= \tilde{\alpha}_2 + m^2 \epsilon^2 \beta_2 + \mathcal{O}(m^4 \epsilon^4), \\ \alpha_0 &= \tilde{\alpha}_0 + \mathcal{O}(m^2 \epsilon^2).\end{aligned}$$

The nonuniversal contribution to the free energy is

$$\tilde{\alpha}_2 \left(\frac{R}{\epsilon}\right)^2 + \tilde{\alpha}_0 + \beta_2 (mR)^2.$$

Thus 3 derivatives w.r.t. mR eliminate the ambiguity.

For nonsupersymmetric theories the 5th derivative of F_{S^4} w.r.t. mR is universal.

The gravity dual

Supergravity setup

Use 5d $\mathcal{N} = 8$ gauged supergravity to construct the holographic dual. Why is this justified?

Supergravity setup

Use 5d $\mathcal{N} = 8$ gauged supergravity to construct the holographic dual. Why is this justified?

- It is a consistent truncation of IIB supergravity on $AdS_5 \times S^5$ which contains fields dual to the lowest dimension operators in $\mathcal{N} = 4$ SYM.
- The gravity dual of $\mathcal{N} = 2^*$ on $\mathbb{R}^{1,3}$ was constructed first in 5d. [Pilch-Warner]

Supergravity setup

Use 5d $\mathcal{N} = 8$ gauged supergravity to construct the holographic dual. Why is this justified?

- It is a consistent truncation of IIB supergravity on $AdS_5 \times S^5$ which contains fields dual to the lowest dimension operators in $\mathcal{N} = 4$ SYM.
- The gravity dual of $\mathcal{N} = 2^*$ on $\mathbb{R}^{1,3}$ was constructed first in 5d. [Pilch-Warner]

From field theory we expect that the solutions we are after possess $U(1)_V \times U(1)_H \times U(1)_Y$ global symmetry. Impose this on the 5d $\mathcal{N} = 8$ supergravity.

Supergravity setup

Use 5d $\mathcal{N} = 8$ gauged supergravity to construct the holographic dual. Why is this justified?

- It is a consistent truncation of IIB supergravity on $AdS_5 \times S^5$ which contains fields dual to the lowest dimension operators in $\mathcal{N} = 4$ SYM.
- The gravity dual of $\mathcal{N} = 2^*$ on $\mathbb{R}^{1,3}$ was constructed first in 5d. [Pilch-Warner]

From field theory we expect that the solutions we are after possess $U(1)_V \times U(1)_H \times U(1)_Y$ global symmetry. Impose this on the 5d $\mathcal{N} = 8$ supergravity.

The result is that only 3 real, 2 Abelian gauge fields and the metric survive in the bosonic sector. The 3 scalars, $\{\phi, \psi, \chi\}$, are dual precisely to the 3 operators

$$\mathcal{O}_\phi, \quad \mathcal{O}_\psi, \quad \mathcal{O}_\chi.$$

It is convenient to use

$$\eta = e^{\phi/\sqrt{6}}, \quad z = \frac{1}{\sqrt{2}}(\chi + i\psi), \quad \tilde{z} = \frac{1}{\sqrt{2}}(\chi - i\psi).$$

In Euclidean signature the fields z and \tilde{z} are independent.

For $\chi = 0$ there is $SU(2)_V \times U(1)_H \times U(1)_Y$ symmetry and we reproduce the Pilch-Warner truncation.

Supergravity setup

The Euclidean Lagrangian is

$$\mathcal{L} = \frac{1}{2\kappa^2} \left[-\mathcal{R} + \frac{12\partial_\mu\eta\partial^\mu\eta}{\eta^2} + \frac{4\partial_\mu z\partial^\mu\tilde{z}}{(1-z\tilde{z})^2} + \mathcal{V} \right],$$
$$\mathcal{V} \equiv -\frac{4}{L^2} \left(\frac{1}{\eta^4} + 2\eta^2 \frac{1+z\tilde{z}}{1-z\tilde{z}} + \frac{\eta^8}{4} \frac{(z-\tilde{z})^2}{(1-z\tilde{z})^2} \right).$$

This is a non-linear sigma model with target $\mathbb{R} \times \mathbb{H}^2$.

Supergravity setup

The Euclidean Lagrangian is

$$\mathcal{L} = \frac{1}{2\kappa^2} \left[-\mathcal{R} + \frac{12\partial_\mu\eta\partial^\mu\eta}{\eta^2} + \frac{4\partial_\mu z\partial^\mu\tilde{z}}{(1-z\tilde{z})^2} + \mathcal{V} \right],$$
$$\mathcal{V} \equiv -\frac{4}{L^2} \left(\frac{1}{\eta^4} + 2\eta^2 \frac{1+z\tilde{z}}{1-z\tilde{z}} + \frac{\eta^8}{4} \frac{(z-\tilde{z})^2}{(1-z\tilde{z})^2} \right).$$

This is a non-linear sigma model with target $\mathbb{R} \times \mathbb{H}^2$.

To preserve the isometries of S^4 take the “domain-wall” Ansatz

$$ds^2 = L^2 e^{2A(r)} ds_{S^4}^2 + dr^2,$$
$$\eta = \eta(r), \quad z = z(r), \quad \tilde{z} = \tilde{z}(r).$$

The masses of the scalars are

$$m_\phi^2 L^2 = m_\chi^2 L^2 = -4, \quad m_\psi^2 L^2 = -3.$$

The BPS equations

Plug the Ansatz in the supersymmetry variations of the 5d $\mathcal{N} = 8$ theory and use the “conformal Killing spinors” on S^4

$$\hat{\nabla}_\mu \zeta = \frac{1}{2} \gamma_5 \gamma_\mu \zeta ,$$

to derive the BPS equations

$$z' = \frac{3\eta'(z\tilde{z} - 1) [2(z + \tilde{z}) + \eta^6(z - \tilde{z})]}{2\eta [\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]} ,$$

$$\tilde{z}' = \frac{3\eta'(z\tilde{z} - 1) [2(z + \tilde{z}) - \eta^6(z - \tilde{z})]}{2\eta [\eta^6(z^2 - 1) + z^2 + 1]} ,$$

$$(\eta')^2 = \frac{[\eta^6(z^2 - 1) + z^2 + 1] [\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]}{9L^2\eta^2(z\tilde{z} - 1)^2} ,$$

$$e^{2A} = \frac{(z\tilde{z} - 1)^2 [\eta^6(z^2 - 1) + z^2 + 1] [\eta^6(\tilde{z}^2 - 1) + \tilde{z}^2 + 1]}{\eta^8(z^2 - \tilde{z}^2)^2} .$$

UV expansion

The (constant curvature) metric on \mathbb{H}^5 is

$$ds_5^2 = dr^2 + L^2 \sinh^2 \left(\frac{r}{L} \right) ds_{S^4}^2 .$$

UV expansion

The (constant curvature) metric on \mathbb{H}^5 is

$$ds_5^2 = dr^2 + L^2 \sinh^2 \left(\frac{r}{L} \right) ds_{S^4}^2 .$$

Solving the BPS equations iteratively, order by order in the asymptotic expansion as $r \rightarrow \infty$, we find

$$e^{2A} = \frac{e^{2r}}{4} + \frac{1}{6}(\mu^2 - 3) + \dots ,$$

$$\eta = 1 + e^{-2r} \left[\frac{2\mu^2}{3} r + \frac{\mu(\mu + v)}{3} \right] + \dots ,$$

$$\frac{1}{2}(z + \tilde{z}) = e^{-2r} [2\mu r + v] + \dots ,$$

$$\frac{1}{2}(z - \tilde{z}) = \mu e^{-r} + e^{-3r} \left[\frac{4}{3} \mu (\mu^2 - 3) r + \frac{1}{3} (2v(\mu^2 - 3) + \mu(4\mu^2 - 3)) \right] + \dots .$$

UV expansion

The (constant curvature) metric on \mathbb{H}^5 is

$$ds_5^2 = dr^2 + L^2 \sinh^2 \left(\frac{r}{L} \right) ds_{S^4}^2 .$$

Solving the BPS equations iteratively, order by order in the asymptotic expansion as $r \rightarrow \infty$, we find

$$e^{2A} = \frac{e^{2r}}{4} + \frac{1}{6}(\mu^2 - 3) + \dots ,$$

$$\eta = 1 + e^{-2r} \left[\frac{2\mu^2}{3} r + \frac{\mu(\mu + v)}{3} \right] + \dots ,$$

$$\frac{1}{2}(z + \tilde{z}) = e^{-2r} [2\mu r + v] + \dots ,$$

$$\frac{1}{2}(z - \tilde{z}) = \mu e^{-r} + e^{-3r} \left[\frac{4}{3}\mu(\mu^2 - 3) r + \frac{1}{3} \left(2v(\mu^2 - 3) + \mu(4\mu^2 - 3) \right) \right] + \dots .$$

Here μ and v are integration constants. Think of them as the “source” and “vev” for the operator \mathcal{O}_χ . Compare to field theory to identify $\mu = imR$.

IR expansion

At $r = r_*$ the S^4 shrinks to zero size. Solve the BPS equations close to $r = r_*$, and require that the solution is smooth to find

IR expansion

At $r = r_*$ the S^4 shrinks to zero size. Solve the BPS equations close to $r = r_*$, and require that the solution is smooth to find

$$e^{2A} = (r - r_*)^2 + \frac{7\eta_0^{12} + 20}{81\eta_0^4} (r - r_*)^4 + \dots,$$

$$\eta = \eta_0 - \left(\frac{\eta_0^{12} - 1}{27\eta_0^3} \right) (r - r_*)^2 \left[1 - \left(\frac{85 + 131\eta_0^{12}}{810\eta_0^4} \right) (r - r_*)^2 + \dots \right],$$

$$\frac{1}{2}(z + \tilde{z}) = \sqrt{\frac{\eta_0^6 - 1}{\eta_0^6 + 1}} \left[\frac{\eta_0^6}{\eta_0^6 + 2} - \frac{2\eta_0^8(4\eta_0^6 + 5)}{15(\eta_0^6 + 2)^2} (r - r_*)^2 + \dots \right],$$

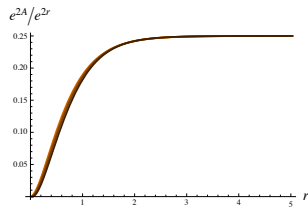
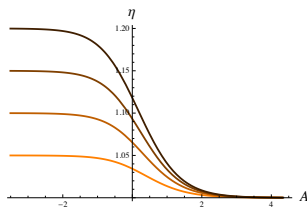
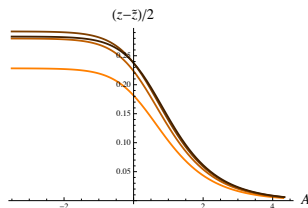
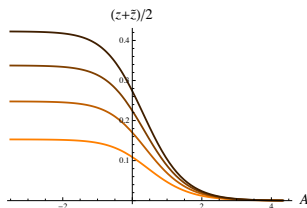
$$\frac{1}{2}(z - \tilde{z}) = \sqrt{\frac{\eta_0^6 - 1}{\eta_0^6 + 1}} \left[\frac{2}{\eta_0^6 + 2} + \frac{\eta_0^2(3\eta_0^{12} - 10\eta_0^6 - 20)}{15(\eta_0^6 + 2)^2} (r - r_*)^2 + \dots \right].$$

Here, η_0 is the only free parameter since one can set $r_* = 0$ by the shift symmetry of the equations.

Numerical solutions

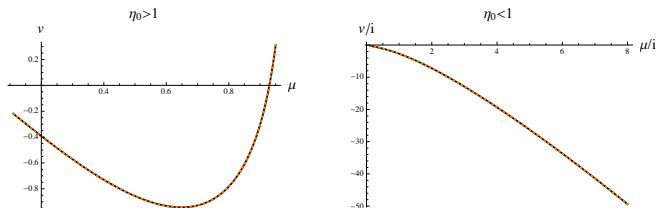
One can find numerical solutions by “shooting” from the IR to the UV. There is a one-parameter family parametrized by η_0 , so

$$v = v(\eta_0), \quad \text{and} \quad \mu = \mu(\eta_0).$$



Numerical solutions

Qualitatively different behavior for $\eta_0 > 1$ and $\eta_0 < 1$.



From the numerical results one can extract the following dependence

$$v(\mu) = -2\mu - \mu \log(1 - \mu^2)$$

Holographic calculation of $F_{S^4}^{\mathcal{N}=2^*}$

Calculating F from supergravity

- By the standard holographic dictionary the partition function of the field theory is mapped to the on-shell action of the supergravity at hand. [Maldacena], [GKP], [Witten]

Calculating F from supergravity

- By the standard holographic dictionary the partition function of the field theory is mapped to the on-shell action of the supergravity at hand. [Maldacena], [GKP], [Witten]
- The on-shell action diverges and one has to regulate it using holographic renormalization. [Skenderis], ...

Calculating F from supergravity

- By the standard holographic dictionary the partition function of the field theory is mapped to the on-shell action of the supergravity at hand. [Maldacena], [GKP], [Witten]
- The on-shell action diverges and one has to regulate it using holographic renormalization. [Skenderis], ...
- There is a subtlety here. If we insist on using a supersymmetric regularization scheme there is a particular finite counterterm that has to be added. Only with it one can successfully compare $\frac{d^3 F}{d\mu^3}$ with the field theory result.

Calculating F from supergravity

- By the standard holographic dictionary the partition function of the field theory is mapped to the on-shell action of the supergravity at hand. [Maldacena], [GKP], [Witten]
- The on-shell action diverges and one has to regulate it using holographic renormalization. [Skenderis], ...
- There is a subtlety here. If we insist on using a supersymmetric regularization scheme there is a particular finite counterterm that has to be added. Only with it one can successfully compare $\frac{d^3 F}{d\mu^3}$ with the field theory result.
- Without knowing this finite counterterm we can only hope to match $\frac{d^5 F}{d\mu^5}$ with field theory.

Calculating F from supergravity

The 5d action + the boundary Gibbons-Hawking term is

$$S_{5D} = \int_M d^5x \sqrt{G} \left\{ -\frac{1}{4} \mathcal{R} + \frac{1}{2} (\partial\phi)^2 + \frac{K}{2} \left((\partial\chi)^2 + (\partial\psi)^2 \right) + \mathcal{V} \right\} - \frac{1}{2} \int_{\partial M} \sqrt{\gamma} \mathcal{K},$$

with $K = (1 - \frac{1}{2}(\chi^2 + \psi^2))^{-2}$.

Calculating F from supergravity

The 5d action + the boundary Gibbons-Hawking term is

$$S_{5D} = \int_M d^5x \sqrt{G} \left\{ -\frac{1}{4} \mathcal{R} + \frac{1}{2} (\partial\phi)^2 + \frac{K}{2} \left((\partial\chi)^2 + (\partial\psi)^2 \right) + \mathcal{V} \right\} \\ - \frac{1}{2} \int_{\partial M} \sqrt{\gamma} \mathcal{K},$$

with $K = (1 - \frac{1}{2}(\chi^2 + \psi^2))^{-2}$.

The (infinite) counterterm action obtained by holographic renormalization is

$$S_{ct} = \int_{\partial M_\epsilon} d^4x \sqrt{\gamma} \left[\frac{3}{2} + \frac{1}{8} \mathcal{R}[\gamma] + \frac{1}{2} \psi^2 + \left(1 + \frac{1}{\log \epsilon} \right) (\phi^2 + \chi^2) \right. \\ \left. - \log \epsilon \left\{ \frac{1}{32} \left[\mathcal{R}[\gamma]^{ij} \mathcal{R}[\gamma]_{ij} - \frac{1}{3} \mathcal{R}[\gamma]^2 \right] + \frac{1}{4} \psi \square_\gamma \psi - \frac{1}{24} \mathcal{R}[\gamma] \psi^2 - \frac{1}{6} \psi^4 \right\} \right],$$

Interestingly S_{ct} is **almost** the same for many 5d models studied in the literature. [FGPW], [GPPZ], [Pilch-Warner],...

Calculating F from supergravity

Use the Bogomolnyi trick to find the finite counterterm. No superpotential for the model with all 3 scalars. In the limits $\chi = 0$ or $\psi = 0$ there is a superpotential.

Calculating F from supergravity

Use the Bogomolnyi trick to find the finite counterterm. No superpotential for the model with all 3 scalars. In the limits $\chi = 0$ or $\psi = 0$ there is a superpotential.

Take $\chi = 0$, then 5d supergravity action action (on \mathbb{R}^4) can be written as

$$S = \int dr d^4x \left\{ e^{4A} \left[-3 \left(A' - \frac{2}{3} \mathcal{W} \right)^2 + \frac{\sqrt{3} e^{-\frac{\phi}{\sqrt{6}}}}{\sqrt{2}} (\phi' + \partial_\phi \mathcal{W})^2 \right. \right. \\ \left. \left. + \frac{2}{(2 - \psi^2)^2} \left(\psi' + \left(1 - \frac{1}{2} \psi^2\right)^2 \partial_\psi \mathcal{W} \right)^2 \right] - \frac{d}{dr} \left(e^{4A} \mathcal{W} \right) \right\}$$

with

$$\mathcal{W} = e^{-2\phi/\sqrt{6}} + \frac{e^{4\phi/\sqrt{6}}}{2} \frac{2 + \psi^2}{2 - \psi^2}, \quad \mathcal{V} = \frac{1}{2} (\partial_\phi \mathcal{W})^2 + \frac{(2 - \psi^2)^2}{8} (\partial_\psi \mathcal{W})^2 - \frac{4}{3} \mathcal{W}^2$$

Calculating F from supergravity

Use the Bogomolnyi trick to find the finite counterterm. No superpotential for the model with all 3 scalars. In the limits $\chi = 0$ or $\psi = 0$ there is a superpotential.

Take $\chi = 0$, then 5d supergravity action action (on \mathbb{R}^4) can be written as

$$S = \int dr d^4x \left\{ e^{4A} \left[-3 \left(A' - \frac{2}{3} \mathcal{W} \right)^2 + \frac{\sqrt{3} e^{-\frac{\phi}{\sqrt{6}}}}{\sqrt{2}} (\phi' + \partial_\phi \mathcal{W})^2 \right. \right. \\ \left. \left. + \frac{2}{(2 - \psi^2)^2} \left(\psi' + (1 - \frac{1}{2} \psi^2)^2 \partial_\psi \mathcal{W} \right)^2 \right] - \frac{d}{dr} (e^{4A} \mathcal{W}) \right\}$$

with

$$\mathcal{W} = e^{-2\phi/\sqrt{6}} + \frac{e^{4\phi/\sqrt{6}}}{2} \frac{2 + \psi^2}{2 - \psi^2}, \quad \mathcal{V} = \frac{1}{2} (\partial_\phi \mathcal{W})^2 + \frac{(2 - \psi^2)^2}{8} (\partial_\psi \mathcal{W})^2 - \frac{4}{3} \mathcal{W}^2$$

The full counterterm action from the Bogomolnyi procedure is

$$S_{\mathcal{W}} = \int d^4x e^{4A} \mathcal{W} = \int d^4x \sqrt{\gamma} \left(\frac{3}{2} + \phi^2 + \frac{1}{2} \psi^2 + \sqrt{\frac{2}{3}} \phi \psi^2 + \frac{1}{4} \psi^4 + \dots \right).$$

The finite counterterm obtained by the Bogomolnyi trick is

$$S_{\text{finite}} = \int d^4x \sqrt{\gamma} \frac{1}{4} \psi^4$$

Calculating F from supergravity

The full renormalized 5d action is

$$S_{\text{ren}} = S_{5\text{D}} + S_{\text{ct}} + S_{\text{finite}} .$$

Calculating F from supergravity

The full renormalized 5d action is

$$S_{\text{ren}} = S_{5\text{D}} + S_{\text{ct}} + S_{\text{finite}} .$$

Differentiate the renormalized action w.r.t. μ to find

$$\frac{dF^{\text{SUGRA}}}{d\mu} = \frac{N^2}{2\pi^2} \text{vol}(S^4) (4\mu - 12v(\mu)) = N^2 \left(\frac{1}{3}\mu - v(\mu) \right) .$$

Calculating F from supergravity

The full renormalized 5d action is

$$S_{\text{ren}} = S_{5\text{D}} + S_{\text{ct}} + S_{\text{finite}} .$$

Differentiate the renormalized action w.r.t. μ to find

$$\frac{dF^{\text{SUGRA}}}{d\mu} = \frac{N^2}{2\pi^2} \text{vol}(S^4) (4\mu - 12v(\mu)) = N^2 \left(\frac{1}{3}\mu - v(\mu) \right) .$$

Finally we arrive at the supergravity result

$$\frac{d^3 F^{\text{SUGRA}}}{d\mu^3} = -N^2 v''(\mu) = -2N^2 \frac{\mu(3 - \mu^2)}{(1 - \mu^2)^2} .$$

Set $\mu = imR$ and compare this to field theory

$$\frac{d^3 F_{S^4}}{(mR)^3} = -2N^2 \frac{mR((mR)^2 + 3)}{((mR)^2 + 1)^2} .$$

Calculating F from supergravity

The full renormalized 5d action is

$$S_{\text{ren}} = S_{5\text{D}} + S_{\text{ct}} + S_{\text{finite}} .$$

Differentiate the renormalized action w.r.t. μ to find

$$\frac{dF^{\text{SUGRA}}}{d\mu} = \frac{N^2}{2\pi^2} \text{vol}(S^4) (4\mu - 12v(\mu)) = N^2 \left(\frac{1}{3}\mu - v(\mu) \right) .$$

Finally we arrive at the supergravity result

$$\frac{d^3 F^{\text{SUGRA}}}{d\mu^3} = -N^2 v''(\mu) = -2N^2 \frac{\mu(3 - \mu^2)}{(1 - \mu^2)^2} .$$

Set $\mu = imR$ and compare this to field theory

$$\frac{d^3 F_{S^4}}{d(mR)^3} = -2N^2 \frac{mR((mR)^2 + 3)}{((mR)^2 + 1)^2} .$$

Lo and behold

$$\boxed{\frac{d^3 F_{S^4}}{d(mR)^3} = \frac{d^3 F^{\text{SUGRA}}}{d\mu^3}}$$

Calculating F from supergravity

The full renormalized 5d action is

$$S_{\text{ren}} = S_{5\text{D}} + S_{\text{ct}} + S_{\text{finite}} .$$

Differentiate the renormalized action w.r.t. μ to find

$$\frac{dF^{\text{SUGRA}}}{d\mu} = \frac{N^2}{2\pi^2} \text{vol}(S^4) (4\mu - 12v(\mu)) = N^2 \left(\frac{1}{3}\mu - v(\mu) \right) .$$

Finally we arrive at the supergravity result

$$\frac{d^3 F^{\text{SUGRA}}}{d\mu^3} = -N^2 v''(\mu) = -2N^2 \frac{\mu(3 - \mu^2)}{(1 - \mu^2)^2} .$$

Set $\mu = imR$ and compare this to field theory

$$\frac{d^3 F_{S^4}}{d(mR)^3} = -2N^2 \frac{mR((mR)^2 + 3)}{((mR)^2 + 1)^2} .$$

Lo and behold

$$\boxed{\frac{d^3 F_{S^4}}{d(mR)^3} = \frac{d^3 F^{\text{SUGRA}}}{d\mu^3}}$$

Without the finite counterterm we can only match $\frac{d^5 F^{\text{SUGRA}}}{d\mu^5}$ with field theory.

Comments on $\mathcal{N} = 1^*$

$$\mathcal{N} = 1^*$$

- Perform a similar analysis for mass deformations of $\mathcal{N} = 4$ SYM with only $\mathcal{N} = 1$ supersymmetry.
- Supersymmetric localization does not seem to work for $\mathcal{N} = 1$ theories on S^4 .
[Festuccia, Komargodski, ...]
- Supergravity may offer some insight about the dual field theory!

$$\mathcal{N} = 1^*$$

- Perform a similar analysis for mass deformations of $\mathcal{N} = 4$ SYM with only $\mathcal{N} = 1$ supersymmetry.
- Supersymmetric localization does not seem to work for $\mathcal{N} = 1$ theories on S^4 .
[Festuccia, Komargodski, ...]
- Supergravity may offer some insight about the dual field theory!

The Lagrangian for $\mathcal{N} = 1^*$ on S^4 of radius R is

$$\begin{aligned}\mathcal{L}_{\mathcal{N}=1^*}^{S^4} &= \mathcal{L}_{\mathcal{N}=4}^{S^4} + \frac{(2 + m_i^2 R^2)}{R^2} \text{tr}(Z_i \tilde{Z}_i) - \frac{m_i}{2} \text{tr}(\chi_i \chi_i + \tilde{\chi}_i \tilde{\chi}_i) \\ &\quad - \frac{\sqrt{2} m_i}{2} \varepsilon^{ijkl} \text{tr}(\tilde{Z}_i Z_j Z_k + Z_i \tilde{Z}_j \tilde{Z}_k) \\ &\quad + \frac{i m_i}{2R} \text{tr}(Z_i^2 + \tilde{Z}_i^2) .\end{aligned}$$

In supergravity this will mean at least 8 real scalars. Possible in principle but too cumbersome...

$$\mathcal{N} = 1^*$$

There are special cases which allow for an explicit analysis.

- Take $m_1 = m_2 = m_3$. In flat space this is the GPPZ flow. It has $SO(3)$ global symmetry. On S^4 we need 4 supergravity scalars.
- Take $m_2 = m_3 = 0$. In flat space this is the Leigh-Strassler flow. On S^4 we need 3 supergravity scalars.

$$\mathcal{N} = 1^*$$

There are special cases which allow for an explicit analysis.

- Take $m_1 = m_2 = m_3$. In flat space this is the GPPZ flow. It has $SO(3)$ global symmetry. On S^4 we need 4 supergravity scalars.
- Take $m_2 = m_3 = 0$. In flat space this is the Leigh-Strassler flow. On S^4 we need 3 supergravity scalars.

For both examples we have explicit supergravity truncations and BPS solutions on S^4 .

The hard part is to extract the third derivative of the free energy from the numerical solutions.

No results from localization to guide us.

A glimpse of the field theory answer from conformal perturbation theory in the planar limit and in the small m expansion.

Summary

- We found a 5d supergravity dual of $\mathcal{N} = 2^*$ SYM on S^4 .
- After careful holographic renormalization we computed the universal part of the free energy of this theory.
- The result is in exact agreement with the supersymmetric localization calculation in field theory.
- This is a precision test of holography in a non-conformal Euclidean setting.
- The results extend immediately to $\mathcal{N} = 2^*$ mass deformations of quiver gauge theories obtained by \mathbb{Z}_k orbifolds of $\mathcal{N} = 4$ SYM. [Azeyanagi-Hanada-Honda-Matsuo-Shiba]
- Extension of these results to $\mathcal{N} = 1^*$ SYM.

- Uplift of the $\mathcal{N} = 2^*$ solution to IIB supergravity. Will allow for a holographic calculation of the Wilson line vev. In addition one can study probe D3-branes.
- Holography for $\mathcal{N} = 2^*$ on other 4-manifolds.
- Extensions to other $\mathcal{N} = 2$ theories in 4d with holographic duals, e.g. pure $\mathcal{N} = 2$ SYM?
- Extensions to other dimensions.

- Uplift of the $\mathcal{N} = 2^*$ solution to IIB supergravity. Will allow for a holographic calculation of the Wilson line vev. In addition one can study probe D3-branes.
- Holography for $\mathcal{N} = 2^*$ on other 4-manifolds.
- Extensions to other $\mathcal{N} = 2$ theories in 4d with holographic duals, e.g. pure $\mathcal{N} = 2$ SYM?
- Extensions to other dimensions.
- Can we see some of the large N phase transitions argued to exist by Russo-Zarembo in IIB string theory?
- Revisit supersymmetric localization for $\mathcal{N} = 1$ theories on S^4 . Can one find the exact partition function (modulo ambiguities)? [Gerchkovitz-Gomis-Komargodski]
- Massive deformations of class \mathcal{S} SCFTs from holography?
- Broader lessons for holography from localization?

THANK YOU!