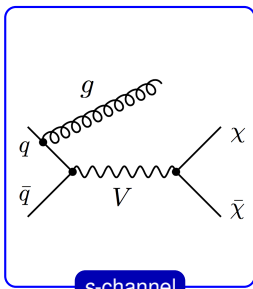


# The Future of Simplified Models

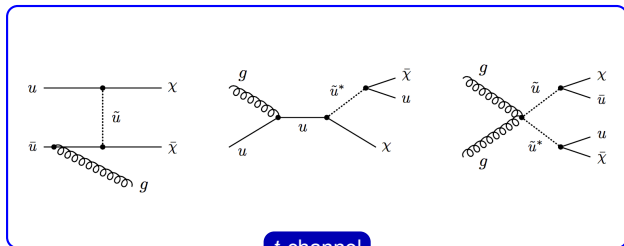
Joachim Kopp



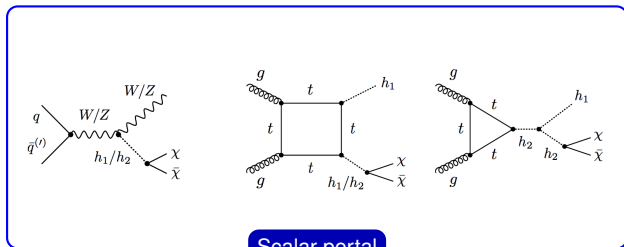
# The current simplified model portfolio



s-channel



t-channel



Scalar portal

see talks by Sarah A. Malik and Oliver Buchmüller

# Outline

- 1 Simplified models off the beaten paths
- 2 Looking for the mediators
- 3 Non-minimal dark sectors
- 4 Data-Driven Phenomenology: Advocacy for Ambulance Chasing
- 5 Summary



# Simplified Models off the Beaten Paths

What may we be missing?

A bottom-up exercise  
in **classifying** simplified models

What may we be missing?

A bottom-up exercise  
in **classifying** simplified models

Allow for **coannihilation**  
→ thermal relics even at weak coupling/heavy mass

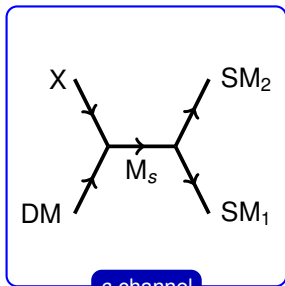
# One step beyond minimal simplified models

## Assumptions:

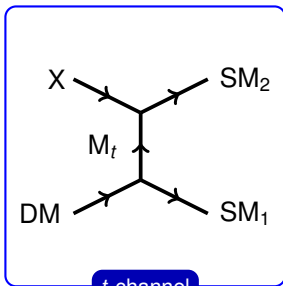
- $\leq 3$  new multiplets:
  - ▶ Dark matter
  - ▶ Coannihilation partner
  - ▶ Mediator
- DM is colorless and electrically neutral
- DM is a thermal relic
- DM annihilation/coannihilation is a 2–2 process
- Tree-level, renormalizable interactions
- New particles are spin 0, 1/2, or 1

Baker et al., *The Coannihilation Codex*, [arXiv:1510.03434](https://arxiv.org/abs/1510.03434)

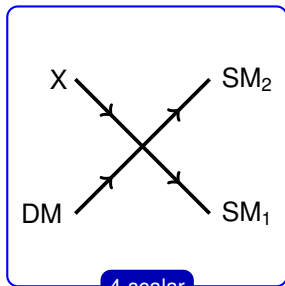
# Four classes of simplified models



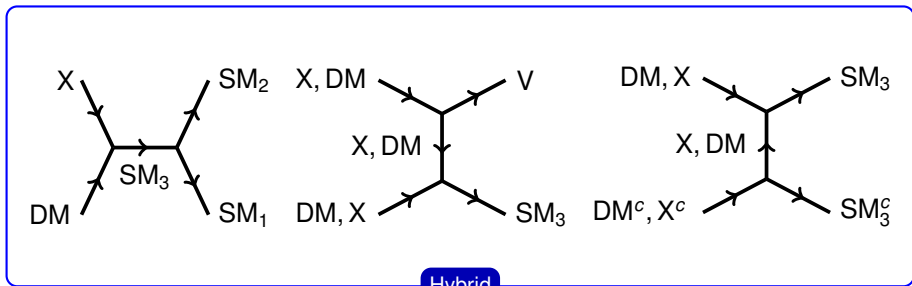
s-channel



t-channel



4-scalar



Hybrid



# Classification procedure

- Choose  $DM$  in  $(1, N, \beta)$  of  $SU(3) \times SU(2) \times U(1)$ .

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- Choose **spin** of **mediator**  $M$ 
  - ▶ Determines the spins of  $DM$  and  $X$

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- Allowed representations for **mediator**  $M$  are now fixed
- Choose **spin** of **mediator**  $M$ 
  - ▶ Determines the spins of **DM** and  $X$
- Allowed couplings to **SM particles** now determined

# Tables, tables, tables

ID	X	$\alpha + \beta$	$M_s$	Spin	(SM <sub>1</sub> SM <sub>2</sub> )	X-DM-SM <sub>3</sub>	$M_s$ -X-X
SU1	(1, N, $\alpha$ )	0	(1, 1, 0)	B	$(u_R \bar{u}_R), (d_R \bar{d}_R), (\ell_R \bar{\ell}_R)$ $(Q_L \bar{Q}_L), (L_L \bar{L}_L), (H H^\dagger)$	H1	✓
SU2				F	$(L_L H)$		
SU3				B	$(Q_L \bar{Q}_L), (L_L \bar{L}_L), (H H^\dagger)$	H1	✓
SU4				F	$(L_L H)$		
SU5		-2	(1, 1, -2)	B	$(d_R \bar{u}_R), (H^\dagger H^\dagger), (L_L L_L)$		✓
SU6				F	$(L_L H^\dagger)$	H2	
SU7				B	$(H^\dagger H^\dagger), (L_L L_L)$		✓ ( $\alpha = \pm 1$ )
SU8				F	$(L_L H^\dagger)$	H2	
SU9		-4	(1, 1, -4)	B	$(\ell_R \ell_R)$		✓ ( $\alpha = \pm 2$ )
SU10	(1, N $\pm 1$ , $\alpha$ )	-1	(1, 2, -1)	B	$(d_R \bar{Q}_L), (\bar{u}_R Q_L), (\bar{L}_L \ell_R)$	H3	
SU11				F	$(\ell_R H)$	H4	
SU12		-3	(1, 2, -3)	B	$(L_L \ell_R)$		
SU13				F	$(\ell_R H^\dagger)$		
SU14	(1, N $\pm 2$ , $\alpha$ )	0	(1, 3, 0)	B	$(Q_L \bar{Q}_L), (L_L \bar{L}_L), (H H^\dagger)$		✓ ( $\alpha = 0$ )
SU15				F	$(L_L H)$		
SU16		-2	(1, 3, -2)	B	$(H^\dagger H^\dagger), (L_L L_L)$		✓ ( $\alpha = \pm 1$ )
SU17				F	$(L_L H^\dagger)$		

DM

X

$M_s$

SM<sub>1</sub>, SM<sub>2</sub>

SM<sub>3</sub>

in (1, N,  $\beta$ ) representation of  $SU(3) \times SU(2) \times U(1)$

coannihilation partner

s-channel mediator

SM particles in coannihilation  $DM + X \rightarrow SM_1 SM_2$

Possible additional vertex  $DM-X-SM_3$

# Tables, tables, tables

ID	X	$\alpha + \beta$	$M_s$	Spin	(SM <sub>1</sub> SM <sub>2</sub> )	X-DM-SM <sub>3</sub>	$M_s$ -X-X
SU1	(1, N, $\alpha$ )	0	(1, 1, 0)	B	$(u_R \bar{u}_R), (d_R \bar{d}_R), (\ell_R \bar{\ell}_R)$ $(Q_L \bar{Q}_L), (L_L \bar{L}_L), (H H^\dagger)$	H1	✓
SU2				F	$(L_L H)$		
SU3				B	$(Q_L \bar{Q}_L), (L_L \bar{L}_L), (H H^\dagger)$	H1	✓
SU4				F	$(L_L H)$		
SU5				B	$(d_R \bar{u}_R), (H^\dagger H^\dagger), (L_L L_L)$		✓

## Tally

In total **161 simplified models** (defined by representations of **DM**, **X** and **M**)  
**49 s-channel**, **105 t-channel**, **7 hybrid**

SU12	(1, N $\pm$ 2, $\alpha$ )	-3	(1, 2, -3)	B	$(\nu_L \bar{\nu}_R)$		
SU13				F	$(\ell_R H^\dagger)$		
SU14				B	$(Q_L \bar{Q}_L), (L_L \bar{L}_L), (H H^\dagger)$		✓ ( $\alpha = 0$ )
SU15				F	$(L_L H)$		
SU16				B	$(H^\dagger H^\dagger), (L_L L_L)$		✓ ( $\alpha = \pm 1$ )
SU17	F	$(L_L H^\dagger)$					

- DM** in  $(1, N, \beta)$  representation of  $SU(3) \times SU(2) \times U(1)$
- X** coannihilation partner
- M<sub>s</sub>** s-channel mediator
- SM<sub>1</sub>, SM<sub>2</sub>** SM particles in coannihilation  $DM + X \rightarrow SM_1 SM_2$
- SM<sub>3</sub>** Possible additional vertex  $DM-X-SM_3$

# LHC pheno classification

Another table ...

	$pp \rightarrow \dots$	Prod. via	Signatures	Search
common	DM + DM + ISR	gauge int. or $SM_1 \in p$ for $t$ -channel	mono-Y + $\cancel{E}_T$	[55,56,62,63,104]
	X ( $\rightarrow SM_1^{\text{soft}} SM_2^{\text{soft}} DM$ ) + X ( $\rightarrow SM_1^{\text{soft}} SM_2^{\text{soft}} DM$ ) + ISR	gauge int. or $SM_2 \in p$ for $t$ -channel	mono-Y + $\cancel{E}_T$ mono-Y + $\cancel{E}_T + \leq 4$ SM	[55,56,62,63,104] Partial coverage [105]
	DM + X ( $\rightarrow SM_1^{\text{soft}} SM_2^{\text{soft}} DM$ ) + ISR	$(SM_1 SM_2) \in p$	mono-Y + $\cancel{E}_T$ mono-Y + $\cancel{E}_T + \leq 2$ SM	[55,56,62,63,104] Partial coverage [105]
s-channel	$M_s (\rightarrow [SM_1 SM_2]^{\text{res}})$ + $M_s (\rightarrow [SM_1 SM_2]^{\text{res}})$	gauge int.	2 resonances	[106–112]
	$M_s (\rightarrow [SM_1 SM_2]^{\text{res}})$ + $M_s (\rightarrow DM + X (\rightarrow SM_1^{\text{soft}} SM_2^{\text{soft}} DM))$		resonance + $\cancel{E}_T$ resonance + $\cancel{E}_T + \leq 2$ SM	No search No search
	$M_s (\rightarrow DM + X (\rightarrow SM_1^{\text{soft}} SM_2^{\text{soft}} DM))$ + $M_s (\rightarrow DM + X (\rightarrow SM_1^{\text{soft}} SM_2^{\text{soft}} DM))$		$\cancel{E}_T + \leq 4$ SM	[113,114,114–124]
	$M_s (\rightarrow [SM_1 SM_2]^{\text{res}})$	$(SM_1 SM_2) \in p$	1 resonance	[125–146]
	$M_s (\rightarrow DM + X (\rightarrow SM_1^{\text{soft}} SM_2^{\text{soft}} DM))$		$\cancel{E}_T + \leq 2$ SM	[120–122,124] [104,147–153]
	$SM_{1,2} + M_s (\rightarrow [SM_1 SM_2]^{\text{res}})$	$SM_{2,1} \in p$	1 resonance + 1 SM	Partial coverage [154,155]
	$SM_{1,2}$ + $M_s (\rightarrow DM + X (\rightarrow SM_1^{\text{soft}} SM_2^{\text{soft}} DM))$		$\cancel{E}_T + 1 \leq 3$ SM	[114,120–124] [147–153,156–158]
$M_t (\rightarrow SM_1 DM)$ + $M_t (\rightarrow SM_1 DM)$		$\cancel{E}_T + \leq 2$ SM	[120–122,124] [104,147–153]	



# LHC pheno classification

Another table ...

	$pp \rightarrow \dots$	Prod. via	Signatures	Search
non	DM + DM + ISR	gauge int. or $SM_1 \in p$ for $t$ -channel	mono- $Y + \cancel{E}_T$	[55,56,62,63,104]
	$X (\rightarrow SM_1^{soft} SM_2^{soft} DM)$	gauge int.	mono- $Y + \cancel{E}_T$	[55,56,62,63,104]
<h2>Frequent features</h2> <ul style="list-style-type: none"> <li>● Cascade decays (1- or 2-step, involves MET)</li> <li>● Soft SM particles (in coannihilation scenarios)</li> <li>● Resonances mediator <math>M_s \rightarrow SM_1 SM_2</math></li> </ul>				
s-ch	$M_s (\rightarrow DM + X (\rightarrow SM_1^{soft} SM_2^{soft} DM))$	$(SM_1 SM_2) \in p$	$\cancel{E}_T + \leq 2 \text{ SM}$	[120–122,124] [104,147–153]
	$SM_{1,2} + M_s (\rightarrow [SM_1 SM_2]^{res})$	$SM_{2,1} \in p$	1 resonance + 1 SM	Partial coverage [154,155]
	$SM_{1,2} + M_s (\rightarrow DM + X (\rightarrow SM_1^{soft} SM_2^{soft} DM))$		$\cancel{E}_T + 1 \leq 3 \text{ SM}$	[114,120–124] [147–153,156–158]
	$M_t (\rightarrow SM_1 DM)$ $+ M_t (\rightarrow SM_1 DM)$		$\cancel{E}_T + \leq 2 \text{ SM}$	[120–122,124] [104,147–153]

## Take-home messages (part 1)

- Look for the **mediators!**
- **Many searches sensitive**  
(though not yet interpreted in terms of DM)
  - ▶ Important role for recasting tools  
(**ATOM**, **CheckMate**, **GAMBIT**, **MadAnalysis 5**,  
**SModelS**, ...)



Looking for the Mediators

# Specific example: a leptoquark model

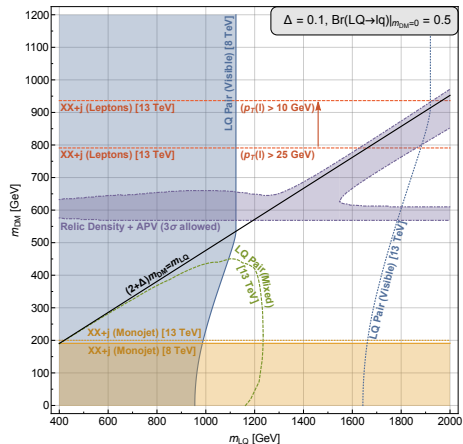
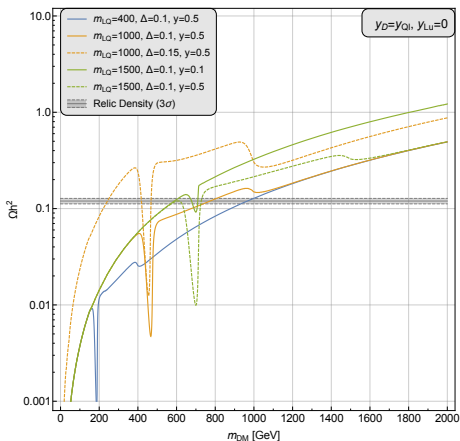
ID	X	$\alpha + \beta$	$M_s$	Spin	(SM <sub>1</sub> SM <sub>2</sub> )	X-DM-SM <sub>3</sub>	$M_s$ -X-X	
ST1	(3, N, $\alpha$ )	$\frac{10}{3}$	$(3, 1, \frac{10}{3})$	B	$(u_R \bar{l}_R)$		$\checkmark(\alpha = -\frac{5}{3})$	
ST2		$\frac{4}{3}$	$(3, 1, \frac{4}{3})$	B	$(d_R \bar{\ell}_R), (Q_L \bar{L}_L), (\bar{d}_R \bar{d}_R)$		$\checkmark(\alpha = -\frac{2}{3})$	
ST3				F	$(Q_L H)$	H5		
ST4			$(3, 3, \frac{4}{3})^{N \geq 2}$	B	$(Q_L \bar{L}_L)$		$\checkmark(\alpha = -\frac{2}{3})$	
ST5				F	$(Q_L H)$	H5		
ST6		$-\frac{2}{3}$	$(3, 1, -\frac{2}{3})$	B	$(\bar{Q}_L \bar{Q}_L), (\bar{u}_R \bar{d}_R), (u_R \ell_R), (Q_L L_L)$		$\checkmark(\alpha = \frac{1}{3})$	
ST7				F	$(Q_L H^\dagger)$	H6		
ST8			$(3, 3, -\frac{2}{3})^{N \geq 2}$	B	$(\bar{Q}_L \bar{Q}_L), (Q_L L_L)$		$\checkmark(\alpha = \frac{1}{3})$	
ST9				F	$(Q_L H^\dagger)$	H6		
ST10		$-\frac{8}{3}$	$(3, 1, -\frac{8}{3})$	B	$(\bar{u}_R \bar{u}_R), (d_R \ell_R)$		$\checkmark(\alpha = \frac{4}{3})$	
ST11	(3, N $\pm$ 1, $\alpha$ )	$\frac{7}{3}$	$(3, 2, \frac{7}{3})$	B	$(Q_L \bar{\ell}_R), (u_R \bar{L}_L)$			
ST12		$\frac{1}{3}$	$(3, 2, \frac{1}{3})$	F	$(u_R H)$			
ST13		$\frac{1}{3}$		B	$(d_R \bar{L}_L), (\bar{Q}_L \bar{d}_R), (u_R L_L)$			
ST14			F	$(u_R H^\dagger), (d_R H)$	H7			
ST15		$-\frac{5}{3}$	$(3, 2, -\frac{5}{3})$	B	$(\bar{Q}_L \bar{u}_R), (Q_L \bar{\ell}_R), (d_R L_L)$			
ST16				F	$(d_R H^\dagger)$			
ST17		(3, N $\pm$ 2, $\alpha$ )	$\frac{4}{3}$	$(3, 3, \frac{4}{3})$	B	$(Q_L \bar{L}_L)$		$\checkmark(\alpha = -\frac{2}{3})$
ST18					F	$(Q_L H)$		
ST19	$-\frac{2}{3}$		$(3, 3, -\frac{2}{3})$	B	$(\bar{Q}_L \bar{Q}_L), (Q_L L_L)$		$\checkmark(\alpha = \frac{1}{3})$	
ST20				F	$(Q_L H^\dagger)$			

# Specific example: a leptoquark model

## Lagrangian

$$\begin{aligned}\mathcal{L} = & \frac{i}{2} \overline{DM} \not{\partial} DM + i \overline{X} \not{\partial} X + |D_\mu M_s|^2 \\ & - \frac{m_{DM}}{2} \overline{DM} DM - m_X \overline{X} X - V(M_s, H) \\ & - \left( y_D \overline{X} M_s DM + y_{Q\ell} \overline{Q}_L M_s \ell_R + y_{Lu} \overline{L}_L M_s^c u_R + \text{h.c.} \right),\end{aligned}$$

# Specific example: a leptoquark model



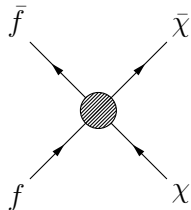
Baker et al., *The Coannihilation Codex*, arXiv:1510.03434



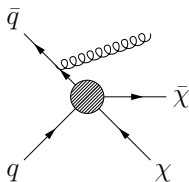
Non-Minimal Dark Sectors

# Dark Matter Production at the LHC

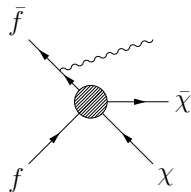
Traditional DM searches: **initial state radiation**



DM pair production  
(invisible @ LHC)

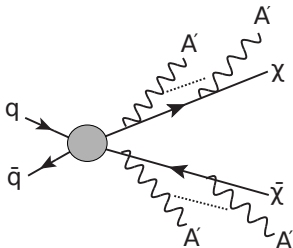


Monojet



Monophoton

How about **final state radiation**?

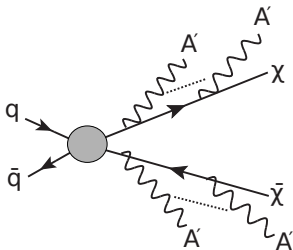




# Model Framework: Self-Interacting DM

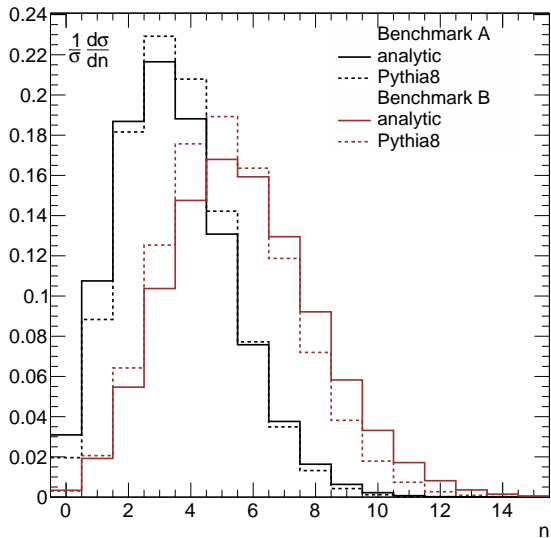
## Dark Sector Lagrangian

$$\mathcal{L}_{\text{dark}} \equiv \bar{\chi}(i\not{\partial} - m_{\chi} + ig_{A'}\not{A}')\chi - \frac{1}{4}F'_{\mu\nu}F'^{\mu\nu} - \frac{1}{2}m_{A'}^2 A'_{\mu}A'^{\mu} - \frac{\epsilon}{2}F'_{\mu\nu}F^{\mu\nu},$$

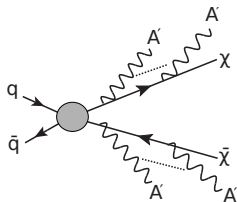


Buschmann JK Liu Wang, [arXiv:1505.07459](https://arxiv.org/abs/1505.07459)

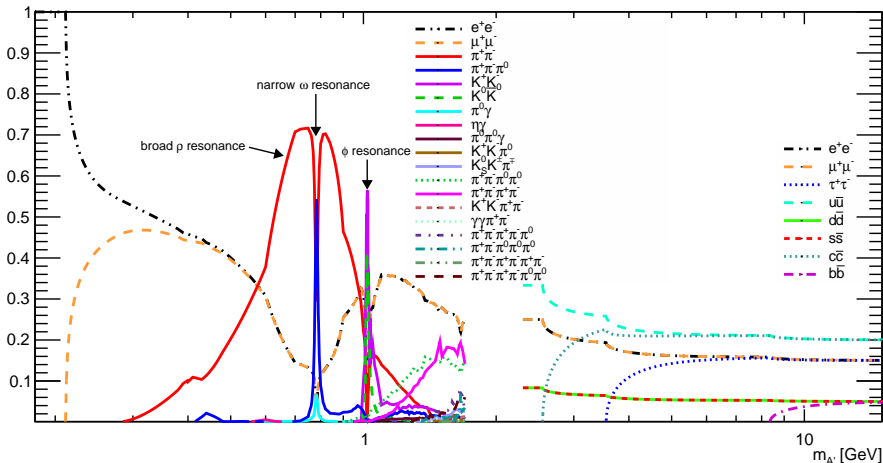
# Dark Radiation — Analytics vs. Numerics



# A' Decay

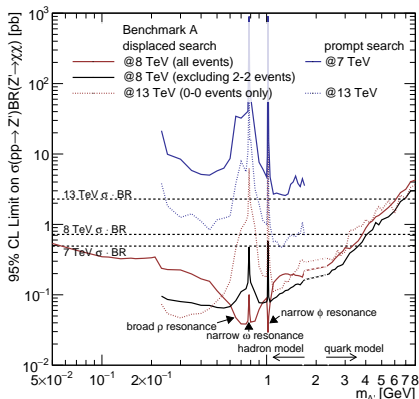
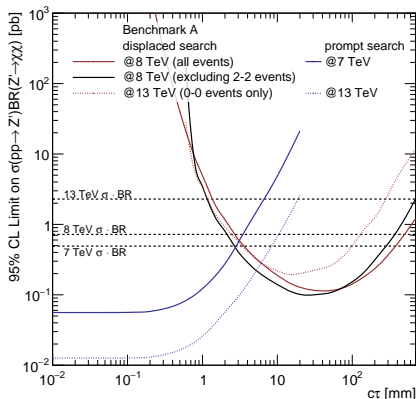


A' branching ratio



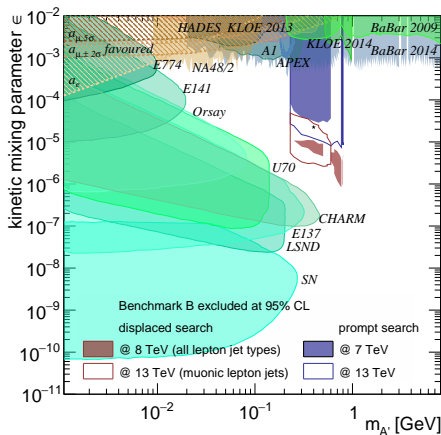
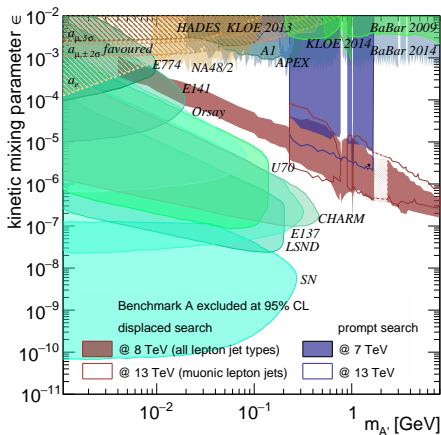
# Phenomenological Results

- Recast ATLAS **prompt** lepton jet search (arXiv:1212.5409)
- Recast ATLAS **displaced** lepton jet search (arXiv:1409.0746)
- Conservative projections for **13 TeV**
  - ▶ Type-0 (muonic lepton jets only) — cannot estimate multijet background



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  - ▶ Type-0 (muonic lepton jets only) — cannot estimate multijet background



## Take-home messages (part 2)

- **Very conventional dark sectors** may lead to **very unconventional signatures**
  - ▶ Lepton jets  
e.g. Buschmann et al., [arXiv:1505.07459](#)
  - ▶ Emerging jets  
e.g. talk by Will Shepherd  
Strassler Zurek, [hep-ph/0604261](#)  
Bai Rajaraman, [arXiv:1109.6009](#)  
Schwaller Stolarski Weiner, [arXiv:1502.05409](#)
  - ▶ Missing energy within jets  
e.g. Cohen Lisanti Lou, [arXiv:1503.00009](#)
  - ▶ ...
- In extended dark sectors, the particle that is **easiest to discover** is typically the one with the **lowest cosmological abundance**  
(large coupling to SM  $\rightarrow$  low relic density)
- Don't take **simplified models** too serious!



# Data-Driven Phenomenology

Why Ambulance Chasing is Awesome

# From Data to Theory

## Traditional Motivation for Simplified Models

Present **experimental results** in a way that allows **theorists** to apply them to **large classes of models**

## Problem

Theorists invent too many **exotic models**  
(which cannot be reduced to the established simplified models)

## My Dream

Exchange **data** on **event-by-event basis**  
(like many **direct/indirect searches**)



# A Pheno-Level Event Record

```
class HEPEventRecord {
    int np; // Number of particles
    HEPParticle *p; // List of particles
    ...
};

class HEPParticle {
    double pT, eta, phi; // 3-momentum
    double E; // Energy
    int type; // Type - jet, photon, e+, e-, etc.
    double em_fraction; // Fraction of E in E-cal
    int n_tracks; // Number of tracks
    bool b_tag;
    int n_spec;
    double *spec; // Analysis-specific info
    ...
};
```

(Heavily inspired by the good old LHC Olympics Format)

- Easy to implement in existing analysis frameworks
- Easy to use
- Customizable

## Criticism

- **Less credit for experimental collaborations**
  - ▶ Share data **after collaboration has exploited it**  
(**ATLAS/CMS** have demonstrated that they can act very fast!)
  - ▶ **Experimentalists' analyses** will always be **more trusted**

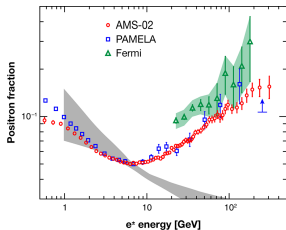
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- **Some will misinterpret/overinterpret data**
  - ▶ Happens anyway
  - ▶ Blame falls on **theorists**
  - ▶ **Not aware of any experiment** whose reputation has suffered from this

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  - ▶ Happens anyway
  - ▶ Blame falls on **theorists**
  - ▶ **Not aware of any experiment** whose reputation has suffered from this
- **Even more ambulance chasing**
  - ▶ IMHO **not a bad thing**

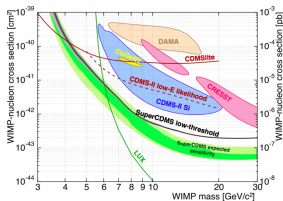
# Recent Ambulances



We've learned about

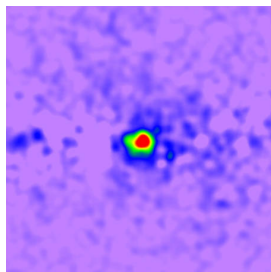
- Pulsars
- Cosmic ray propagation
- Innumerable new DM models
- Sommerfeld enhancement

PAMELA



- Spurred Interest in low-mass DM
- Now main target for many experiments (→ talk by J. Billard)

DAMA/CoGeNT



We've learned about

- High- $E$  astrophysics
- Better CR models
- Analyzing Fermi data (→ talk by M. Lisanti)

Galactic Center Excess

## Take-home messages (part 3)

- Sharing **event-level data**
  - ▶ is **easy**
  - ▶ benefits experimentalists and theorists
  - ▶ **works great** in astrophysics
- **Ambulance chasing**
  - ▶ **moves the field forward**
  - ▶ is educational
  - ▶ is great fun!

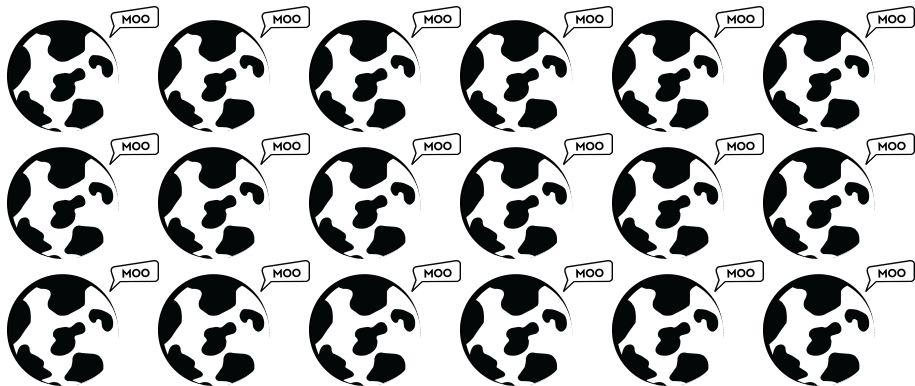
# Summary

- Look for the **mediators!**
- **Many searches sensitive**  
(though not yet interpreted in terms of DM)
- **Very conventional dark sectors** may lead to **very unconventional signatures**
- Don't take **simplified models** too serious!
- Sharing **event-level data** would be great.
- **Ambulance chasing** is a great



**Thank You!**





Bonus Slides

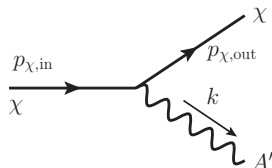
# Dark Radiation Showers — Semi-Analytical Results

Notation, notation, notation, ...

- Incoming (off-shell) DM particle:  $p_{\chi,\text{in}} = (E, 0, 0, p)$
- Outgoing DM particle:  $p_{\chi,\text{out}} = (xE, -k_t, 0, \sqrt{x^2 E^2 - k_t^2 - m_\chi^2})$
- Outgoing dark photon:  $k = ((1-x)E, k_t, 0, \sqrt{(1-x)^2 E^2 - k_t^2 - m_{A'}^2})$
- Virtuality:  $t \equiv (p_{\chi,\text{out}} + k)^2 - m_\chi^2$

Probability for a collinear splitting:

$$\frac{\alpha_{A'}}{2\pi} dx \frac{dt}{t} P_{\chi \rightarrow \chi}(x, t)$$



with the splitting kernel

$$P_{\chi \rightarrow \chi}(x, t) = \frac{1+x^2}{1-x} - \frac{2(m_\chi^2 + m_{A'}^2)}{t}$$

# Dark Radiation Showers — Semi-Analytical Results

Average number radiated dark photons

$$\langle n_{A'} \rangle \simeq \frac{\alpha_{A'}}{2\pi} \int_{x_{\min}}^{x_{\max}} dx \int_{t_{\min}}^{t_{\max}} \frac{dt}{t} P_{\chi \rightarrow \chi}(x).$$

Splitting is a Poisson process.

- Probability for  $m$  splittings

$$p_m = \frac{e^{-\langle n_{A'} \rangle} \langle n_{A'} \rangle^m}{m!}$$

- Probability for **no splitting** (Sudakov factor)

$$\Delta \equiv p_0 = e^{-\langle n_{A'} \rangle}$$

# Dark Radiation — Energy Spectrum of DM Particles

Compute first the **moments** of the  $E$  spectrum  $f_X (X \equiv E_X/E_0)$ :

- Events with one emission

$$\begin{aligned} p_1 \langle X^S \rangle_{1A'} &= e^{-\langle n_{A'} \rangle} \frac{\alpha_{A'}}{2\pi} \int_{x_{\min}}^{x_{\max}} dx x^S \int_{t_{\min}}^{t_{\max}} \frac{dt}{t} P_{X \rightarrow X}(x) \\ &\equiv e^{-\langle n_{A'} \rangle} \langle n_{A'} \rangle \overline{X^S} \end{aligned}$$

- Events with two emissions

$$\begin{aligned} p_2 \langle X^S \rangle_{2A'} &= e^{-\langle n_{A'} \rangle} \left( \frac{\alpha_{A'}}{2\pi} \right)^2 \int_{x_{\min}}^{x_{\max}} dx x^S \int_{t_{\min}}^{t_{\max}} \frac{dt}{t} \int_{x_{\min}}^{x_{\max}} dx' x'^S \int_{t_{\min}}^t \frac{dt'}{t'} P_{X \rightarrow X}(x) P_{X \rightarrow X}(x') \\ &\simeq e^{-\langle n_{A'} \rangle} \frac{1}{2!} \left( \frac{\alpha_{A'}}{2\pi} \right)^2 \int_{x_{\min}}^{x_{\max}} dx x^S \int_{t_{\min}}^{t_{\max}} \frac{dt}{t} \int_{x_{\min}}^{x_{\max}} dx' x'^S \int_{t_{\min}}^{t_{\max}} \frac{dt'}{t'} P_{X \rightarrow X}(x) P_{X \rightarrow X}(x') \\ &= e^{-\langle n_{A'} \rangle} \frac{\langle n_{A'} \rangle^2}{2!} \overline{X^S}^2 \end{aligned}$$

- Events with  $m$  emissions

$$p_m \langle X^S \rangle_{mA'} = e^{-\langle n_{A'} \rangle} \frac{\langle n_{A'} \rangle^m}{m!} \overline{X^S}^m.$$

# Dark Radiation — Energy Spectrum of DM Particles

- Summing over all  $m$

$$\varphi(\mathbf{s} + \mathbf{1}) \equiv \langle X^{\mathbf{s}} \rangle = \sum_{m=0}^{\infty} p_m \langle X^{\mathbf{s}} \rangle_{mA'} = e^{-\langle n_{A'} \rangle (1 - \overline{X^{\mathbf{s}}})}.$$

## Mellin Transform

$$\mathcal{M}[f](\mathbf{s} + \mathbf{1}) \equiv \varphi(\mathbf{s} + \mathbf{1}) \equiv \int_0^{\infty} dX X^{\mathbf{s}} f(X)$$

## Inverse Mellin Transform

$$f(X) = \frac{1}{2\pi i} \int_{c-i\infty}^{c+i\infty} ds X^{-s} \varphi(s)$$

Efficient numerical evaluation using Fast Fourier Transform (FFT)

# Dark Radiation — Energy Spectrum of Dark Photons

With  $Z \equiv E_{A'}/E_0$ :

$$p_m \langle Z^s \rangle_{m A'} = \frac{1}{\langle n_{A'} \rangle} e^{-\langle n_{A'} \rangle} \frac{\langle n_{A'} \rangle^m}{m!} \overline{Z^s} \sum_{k=1}^m \overline{X^s}^{k-1}$$

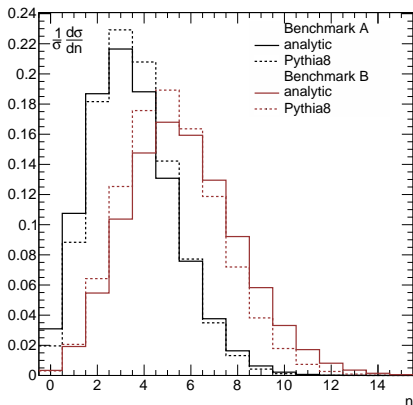
with

$$\overline{Z^s} \equiv \frac{1}{\langle n_{A'} \rangle} \frac{\alpha_{A'}}{2\pi} \int_{x_{\min}}^{x_{\max}} dx (1-x)^s \int_{t_{\min}}^{t_{\max}} \frac{dt}{t} P_{x \rightarrow \chi}(x).$$

Therefore,

$$\varphi(s+1) \equiv \langle Z^s \rangle = \frac{\overline{Z^s}}{\langle n_{A'} \rangle} \frac{1 - e^{-\langle n_{A'} \rangle (1 - \overline{X^s})}}{1 - \overline{X^s}}.$$

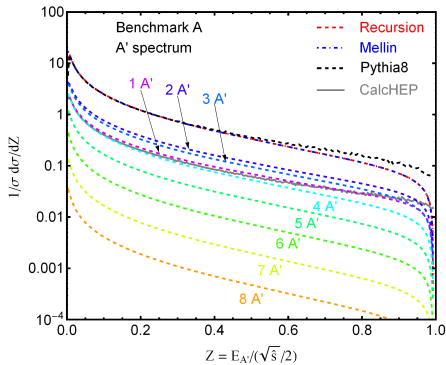
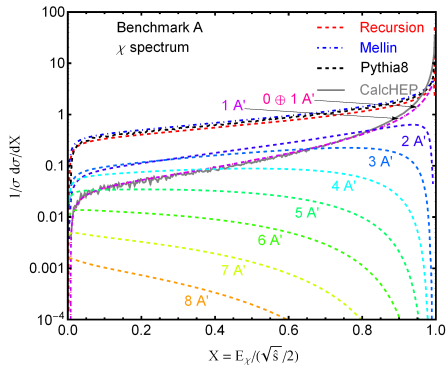
# Dark Radiation — Analytics vs. Numerics



Reasons for minor discrepancies:

- Assumption that integration limits are independent of  $x$ ,  $t$ .
  - ▶ Energy loss in each splitting small
- Neglect of  $t$ -dependence in  $P_{\chi \rightarrow \chi}(x)$

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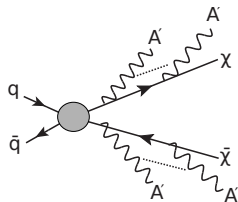


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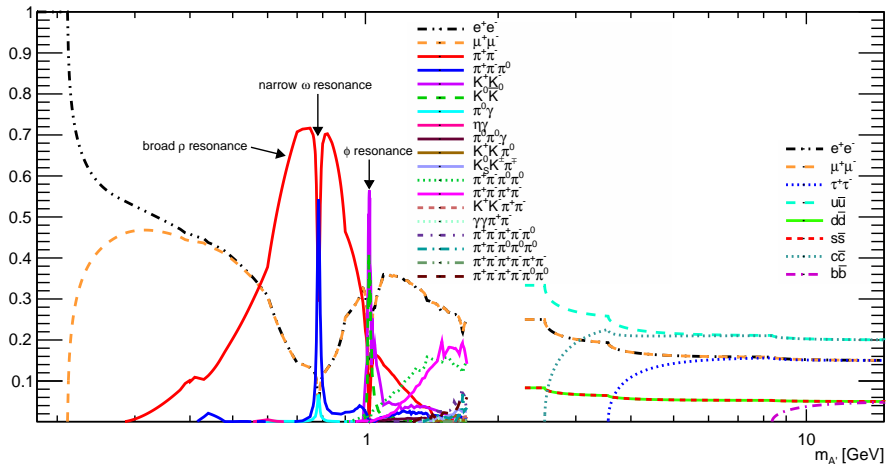
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# A' Decay

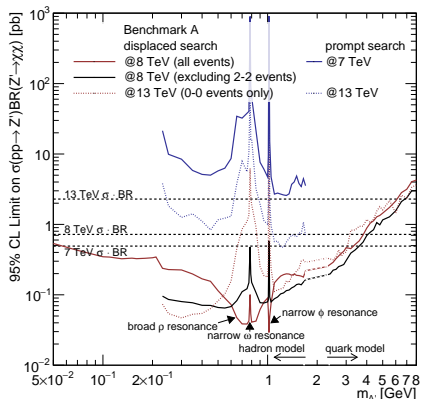
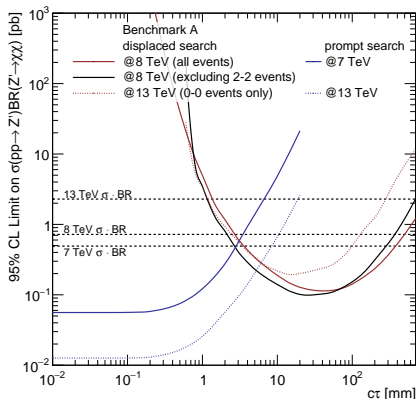


A' branching ratio



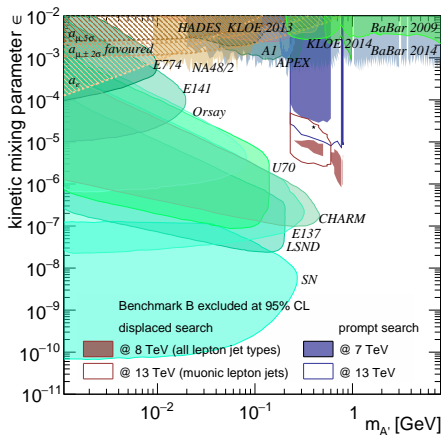
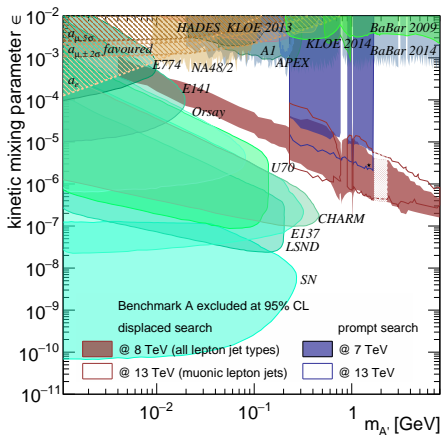
# Phenomenological Results

- Recast ATLAS **prompt** lepton jet search (arXiv:1212.5409)
- Recast ATLAS **displaced** lepton jet search (arXiv:1409.0746)
- Conservative projections for **13 TeV**
  - ▶ Type-0 (muonic lepton jets only) — cannot estimate multijet background



# Phenomenological Results

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# Prompt Lepton Jets

For short  $A'$  lifetime:

- Consider only **muonic lepton jets**
  - ▶ other categories difficult to implement without full detector simulation
- Selection criteria
  - ▶ 1 muon with  $p_T > 18 \text{ GeV}$
  - ▶ or 3 muons with  $p_T > 6 \text{ GeV}$
  - ▶  $|\eta| < 2.5$
  - ▶ Track in the inner detector
  - ▶ Small impact parameter  $|d_0| < 1 \text{ mm}$

# Displaced Lepton Jets

For long  $A'$  lifetime:

- Type 0 (“muonic”) LJ
  - ▶  $\geq 2$  muons (and no calorimeter jets) within  $\Delta R = 0.5$ .
- Type 1 (“mixed”) LJ
  - ▶  $\geq 1$  muon + exactly 1 calorimeter jet
- Type 2 (“calorimeter”) LJ
  - ▶ All other calorimeter jets with **small** EM fraction
  - ▶ Includes  $A' \rightarrow ee$  with large displacement
  - ▶ Includes **hadronic  $A'$  decay modes**

Detector	$A' \rightarrow e^+e^-$	$A' \rightarrow \mu^+\mu^-$	$A' \rightarrow \pi^+\pi^-/K^+K^-$
LJ type	2 (calorimeter)	0 (muonic)	2 (calorimeter)
ID	track	track	track
ECAL	EM fraction	✓	✓
HCAL	✓	✓	✓

Detector	$A' \rightarrow \pi^+\pi^-\pi^0$	$A' \rightarrow K_L^0 K_S^0$
LJ type	2 (calorimeter)	2 (calorimeter)
ID	track	(✓)
ECAL	EM fraction	(✓)
HCAL	✓	✓