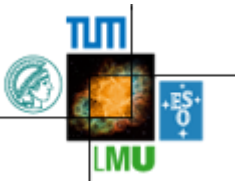


Waiting for New Physics

Andrzej J. Buras
(Technical University Munich, TUM-IAS)

Portoroz, April 2015



A Journey to the Very Short Distance Scales:

1676 - 2046

Microuniverse

10^{-6}m

**Bacteriology
Microbiology**

Nanouniverse

10^{-9}m

Nanoscience

Femtouniverse

10^{-15}m

**Nuclear Physics
Low Energy Elementary
Particle Physics**

Attouniverse

10^{-18}m

**High Energy Particle
Physics (present)**

**High Energy Proton-Proton
Collisions at the LHC**

$5 \cdot 10^{-20}\text{m}$

**Frontiers of Elementary
Particle Physics in 2010's**

**High Precision Measurements
of Rare Processes (Europe,
Japan, USA)**

10^{-21}m

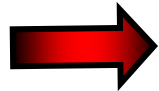
Zeptouniverse



**New Physics beyond the SM
must exist !!!**



**It is our duty to find it.
If not at the LHC then through
high precision experiments.**



**Quark Flavour Physics
Lepton Flavour Violation
EDMs + $(g-2)_{\mu,e}$**

2015-2025 : Expedition
Attouniverse → Zeptouniverse
 $10^{-18}\text{m} \rightarrow 10^{-21}\text{m}$

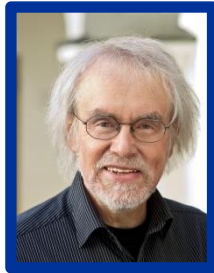
Advanced ERC Grant at the TUM Institute for Advanced Study Zeptouniverse Base Camp



**HEADQUARTERS
ERC-Flavour**



Present ERC Flavour Team



AJB



J.Girrbach-Noe



G.Isidori



S.Pokorski



F. De Fazio



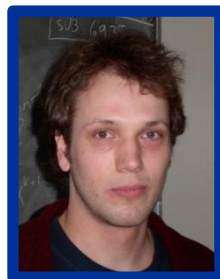
D.Buttazzo



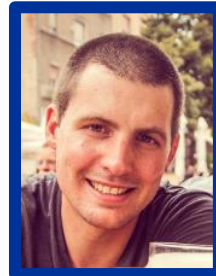
G.Buchalla



A.Ibarra



C.Bobeth



R.Kneijens

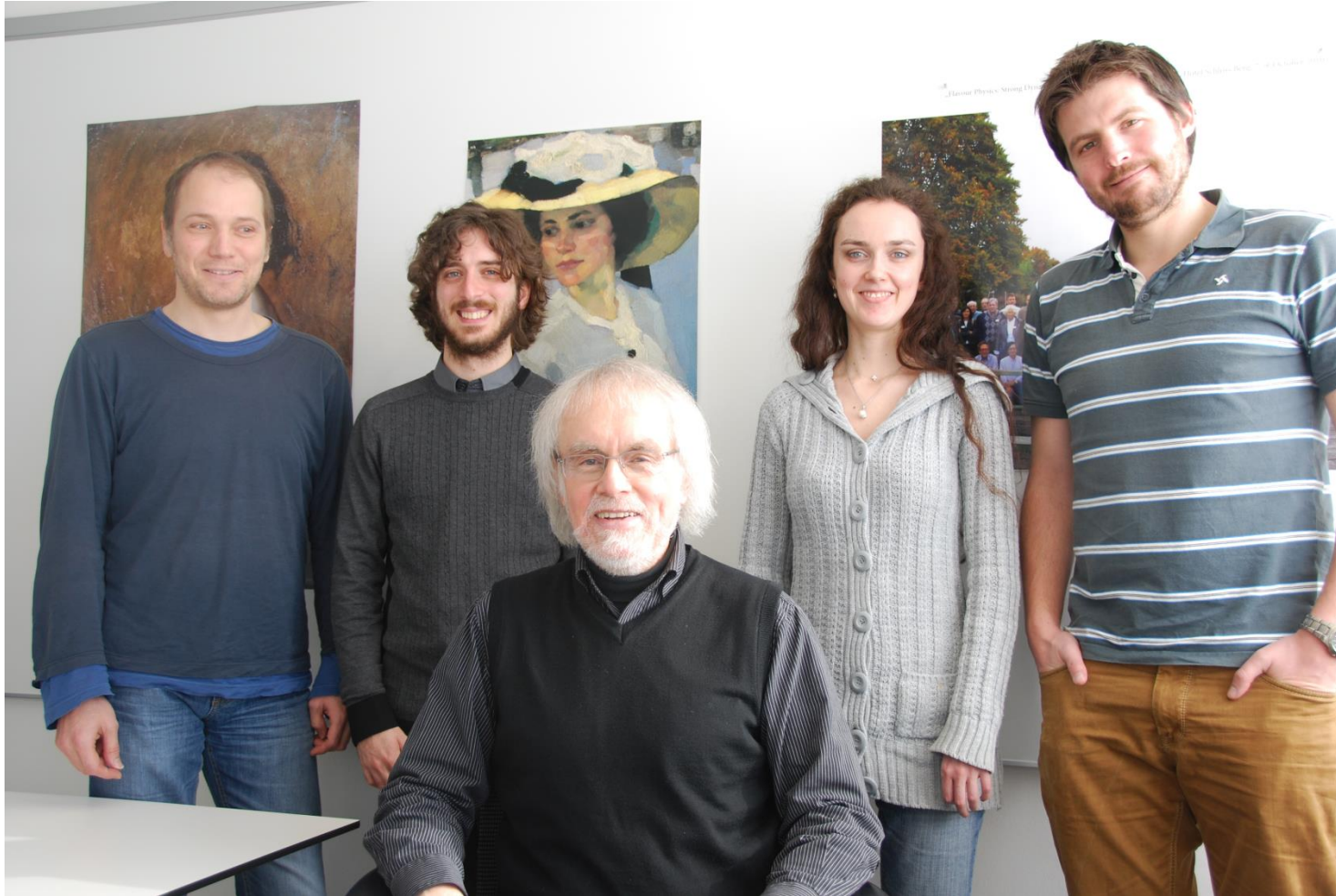


M.Ratz



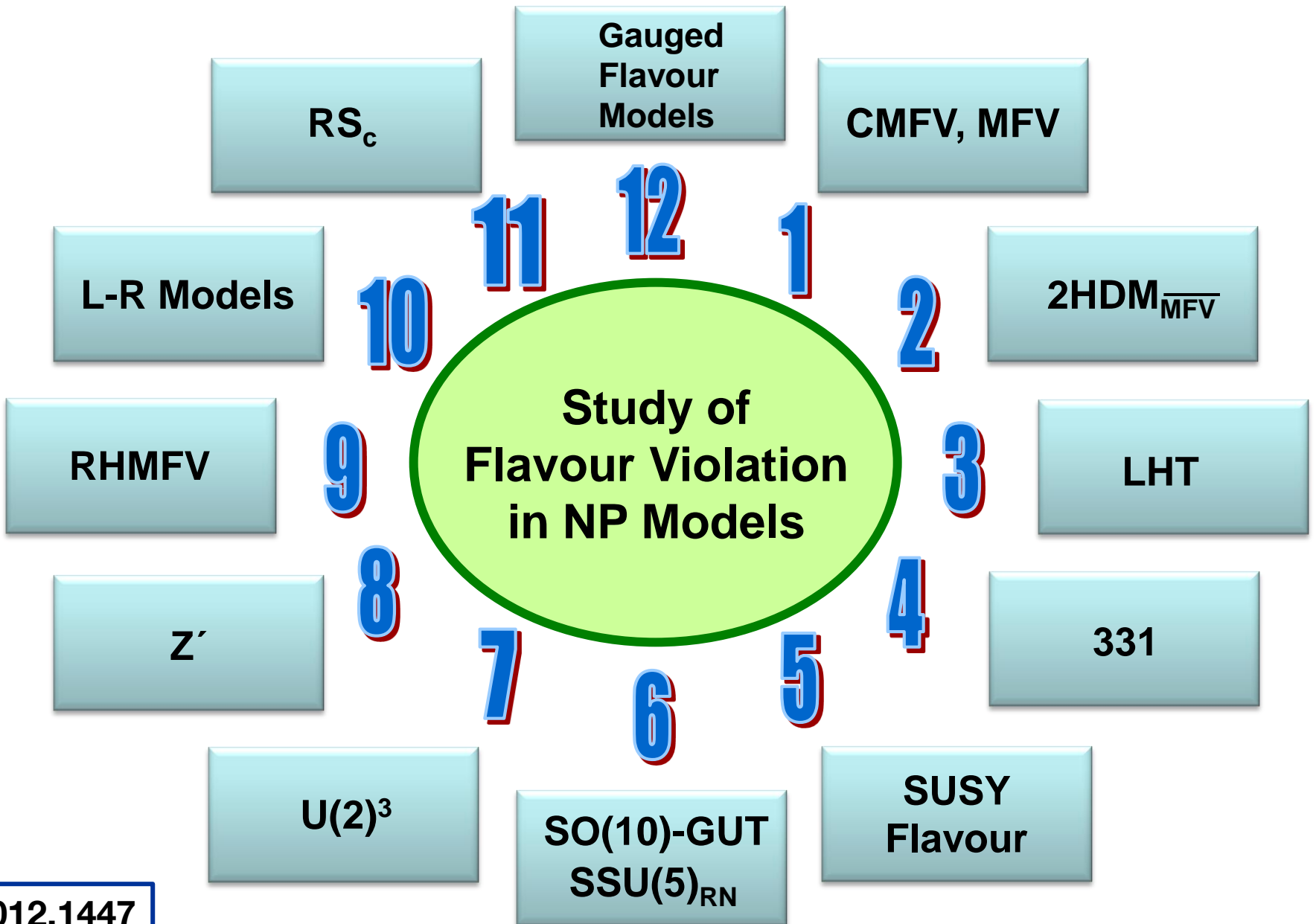
O.Cata

IAS Local Team



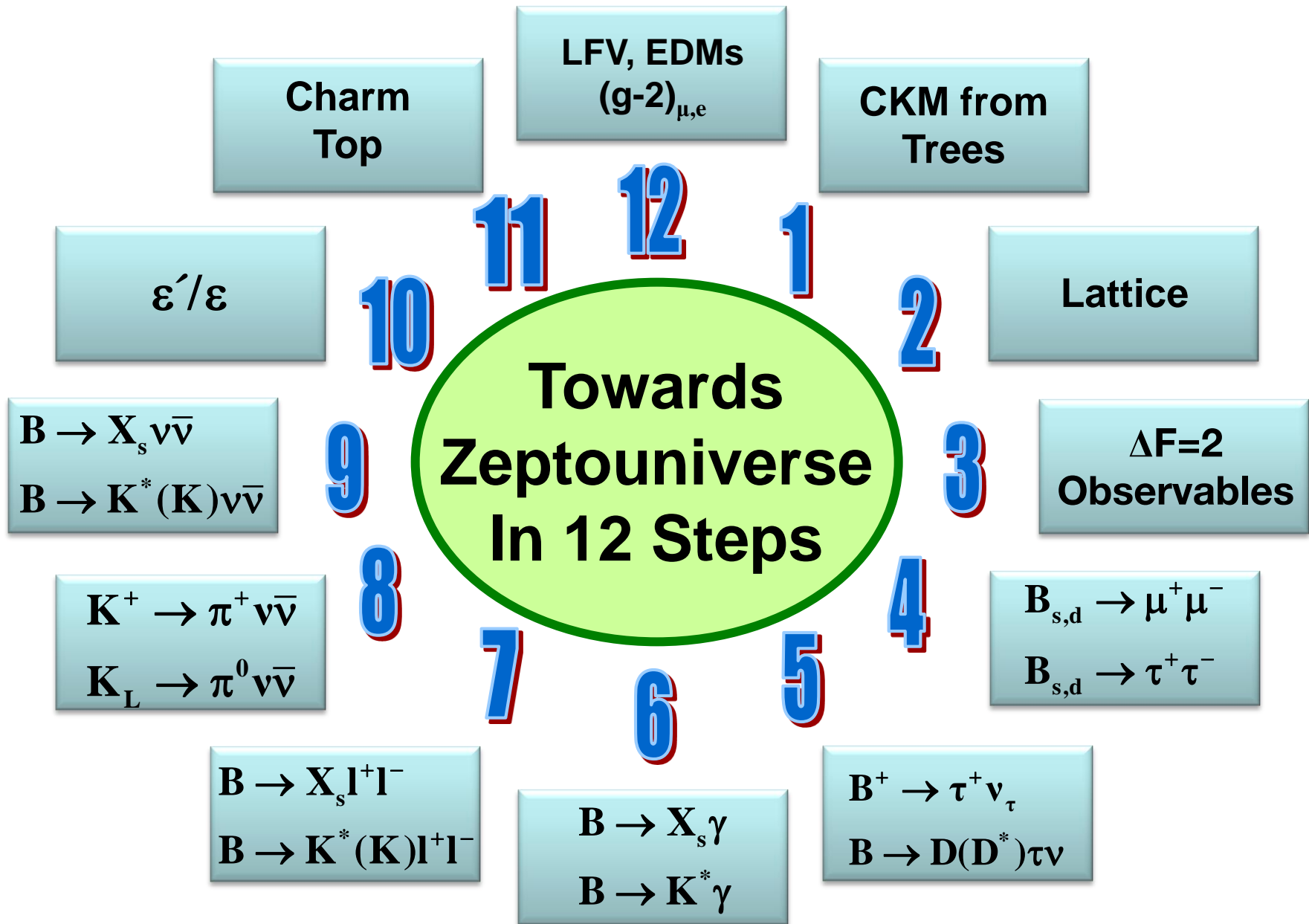
IAS Local Team





1012.1447
1204.5065

1306.3755



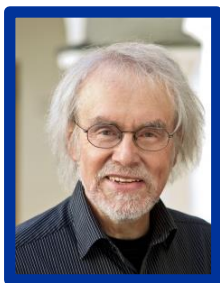
Topics for Next 23 min

- 1.** News on $B \rightarrow K^*(K)\nu\bar{\nu}$
- 2.** News on $K^+ \rightarrow \pi^+\nu\bar{\nu}$, $K_L \rightarrow \pi^0\nu\bar{\nu}$, ε'/ε

1.

News on $B \rightarrow K^* (K) \nu \bar{\nu}$

1409.4557



AJB



J.Girrbach-Noe



Christoph Niehoff



D.Straub

The Power of $B \rightarrow K^*(K)\nu\bar{\nu}$

1. Theoretically cleaner than $b \rightarrow s\mu^+\mu^-$ transitions (factorization exact, only formfactor uncertainties)
2. Sensitive to right-handed currents

Motivation for New *) Analysis

3. Improvement on formfactors (Lattice, LCSR) Bharucha
Straub
Zwicky
4. NLO Electroweak Corrections (Brod, Gorbahn, Stamou, 1009.0947)
5. New Data on $B \rightarrow K^*(K)l^+l^-$ put stronger constraints.

*) Altmannshofer, AJB, Straub, Wick (0902.0160)

Basic Structure

$$\mathbf{H}_{\text{eff}}^{\text{SM}} = -\frac{4G_F}{\sqrt{2}} \mathbf{V}_{tb} \mathbf{V}_{ts}^* \mathbf{C}_L^{\text{SM}} \mathbf{O}_L + \text{h.c}$$

$$\mathbf{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} \mathbf{V}_{tb} \mathbf{V}_{ts}^* (\mathbf{C}_L \mathbf{O}_L + \mathbf{C}_R \mathbf{O}_R) + \text{h.c}$$

$$\mathbf{O}_L \propto (\bar{s} \gamma_\mu \mathbf{P}_L b) (\bar{\nu} \gamma^\mu \mathbf{P}_L \nu) \quad \mathbf{O}_R \propto (\bar{s} \gamma_\mu \mathbf{P}_R b) (\bar{\nu} \gamma^\mu \mathbf{P}_L \nu)$$

$$\varepsilon = \frac{\sqrt{|\mathbf{C}_L|^2 + |\mathbf{C}_R|^2}}{|\mathbf{C}_L^{\text{SM}}|}$$

$$\eta = \frac{-\text{Re}(\mathbf{C}_L \mathbf{C}_R^*)}{|\mathbf{C}_L|^2 + |\mathbf{C}_R|^2} \neq 0$$

only for
RH currents

$$\mathbf{R}_K \neq \mathbf{R}_{K^*}$$

$$\mathbf{R}_K \equiv \frac{\text{Br}(\mathbf{B} \rightarrow \mathbf{K} \nu \bar{\nu})}{\text{Br}^{\text{SM}}(\mathbf{B} \rightarrow \mathbf{K} \nu \bar{\nu})} = (1 - 2\eta) \varepsilon^2$$

$$\mathbf{R}_{K^*} \equiv \frac{\text{Br}(\mathbf{B} \rightarrow \mathbf{K}^* \nu \bar{\nu})}{\text{Br}^{\text{SM}}(\mathbf{B} \rightarrow \mathbf{K}^* \nu \bar{\nu})} = (1 + 1.34\eta) \varepsilon^2$$

Predictions for

$$\mathbf{F}_L(\mathbf{K}^*) \quad \text{(Longitudinal polarization fraction)}$$

$$\text{Br}(\mathbf{B} \rightarrow \mathbf{X}_s \nu \bar{\nu})$$



(ϵ, η) : Parameters for $b \rightarrow sv\bar{\nu}$ Transitions

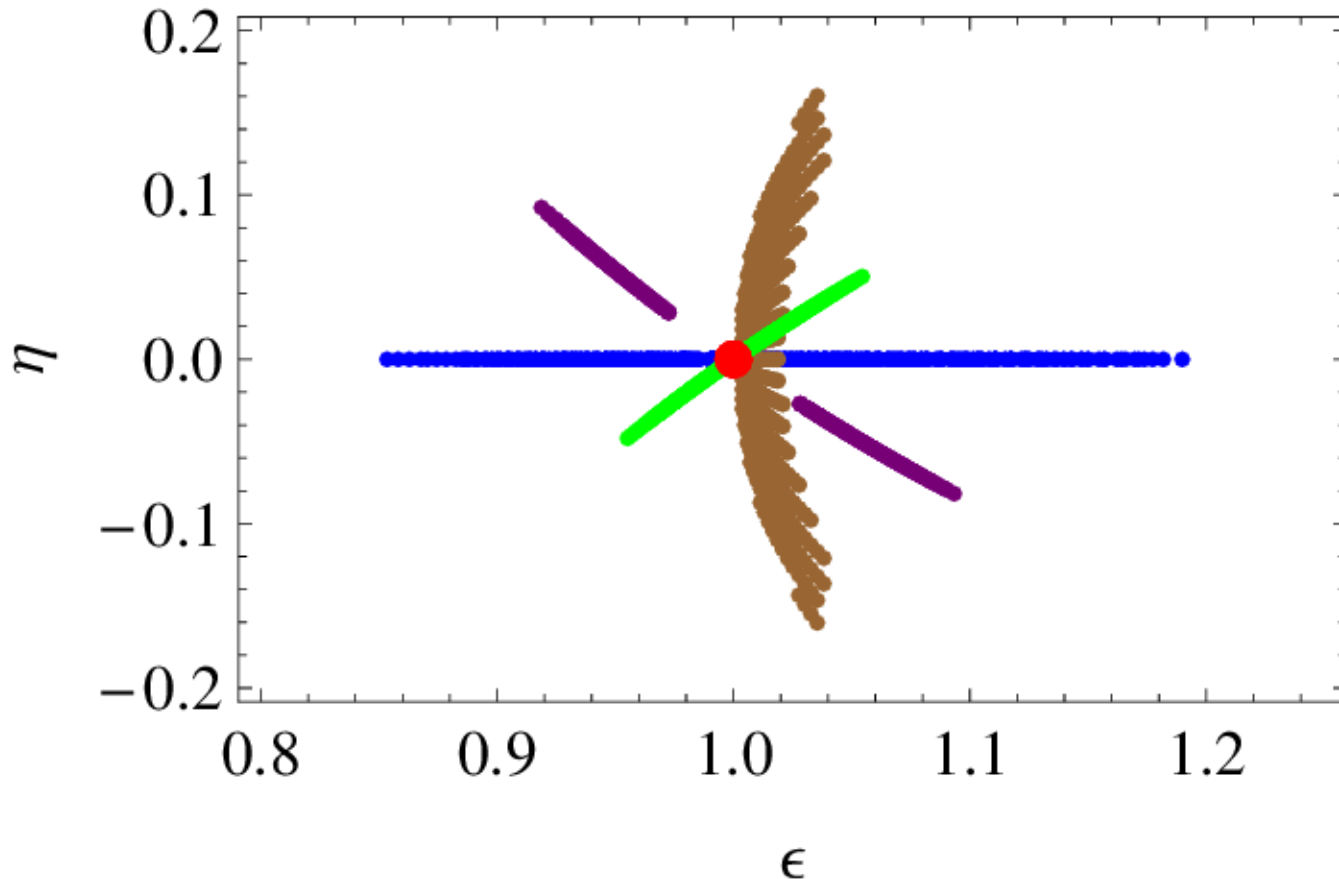
$B \rightarrow K^* v\bar{\nu}$

$B \rightarrow K v\bar{\nu}$

$B \rightarrow X_s v\bar{\nu}$

1211.1896

AJB, de Fazio, Girrbach-Noe



Z'

- LHS
- RHS
- LRS
- ALRS

Powerful tests of right-handed currents

Altmannshofer, AJB, Straub, Wick
0902.0160

Updated SM Prediction

Exp

$$\text{Br}(B^+ \rightarrow K^+ \nu \bar{\nu})_{\text{SM}} = (3.98 \pm 0.43 \pm 0.19) \cdot 10^{-6}$$

$$< 1.7 \cdot 10^{-5} \text{ (BaBar)}$$

$$\text{Br}(B^0 \rightarrow K^{*0} \nu \bar{\nu})_{\text{SM}} = (9.19 \pm 0.86 \pm 0.50) \cdot 10^{-6}$$

$$< 5.5 \cdot 10^{-5} \text{ (Belle)}$$

(formfactor) (CKM)*

Larger by 40%
than previous
estimates

Better:

$$\text{Br}(B^+ \rightarrow K^+ \nu \bar{\nu}) = \left[\frac{|V_{cb}|}{0.0409} \right]^2 (3.98 \pm 0.43) \cdot 10^{-6}$$

$$\text{Br}(B^0 \rightarrow K^{*0} \nu \bar{\nu}) = \left| \frac{|V_{cb}|}{0.0409} \right|^2 (9.19 \pm 0.86) \cdot 10^{-6}$$

$$F_L = 0.47 \pm 0.03$$

* $|V_{cb}| = 0.0409 \text{ (10)}$

Correlation with $b \rightarrow sl^+l^-$

Neutrinos and charged leptons related by $SU(2)_L$ symmetry.

$$\mathbf{O}_9^{(l)} = (\bar{s} \gamma_\mu \mathbf{P}_{L(R)} \mathbf{b}) (\bar{l} \gamma^\mu l) \quad \mathbf{O}_{10}^{(l)} = (\bar{s} \gamma_\mu \mathbf{P}_{L(R)} \mathbf{b}) (\bar{l} \gamma^\mu \gamma_5 l)$$

SM-EFT = OPE with dim = 6 invariant under SM gauge symmetry.

(1008.4884)

Grzadkowski et al.

Examples:

Z'

$$\mathbf{C}_L^{\text{NP}} = \frac{\mathbf{C}_9^{\text{NP}} - \mathbf{C}_{10}^{\text{NP}}}{2} \quad \mathbf{C}_R = \frac{\mathbf{C}'_9 - \mathbf{C}'_{10}}{2}$$

Z

$$\mathbf{C}_L^{\text{NP}} = \mathbf{C}_{10}^{\text{NP}} \quad \mathbf{C}_R = \mathbf{C}'_{10}$$

Z, Z'
with
 $Z-Z'$
mixing

$$\mathbf{C}_L^{\text{NP}} = \frac{\mathbf{C}_9^{\text{NP}} - \mathbf{C}_{10}^{\text{NP}}}{2} + 3 \frac{\tilde{\mathbf{C}}_Z}{2} \quad \mathbf{C}_R = \frac{\mathbf{C}'_9 - \mathbf{C}'_{10}}{2} + 3 \frac{\tilde{\mathbf{C}}'_Z}{2}$$

$\tilde{\mathbf{C}}_Z, \tilde{\mathbf{C}}'_Z$ model dependent ($Z-Z'$ mixing)

(Example 331)

AJB + Fulvia + Jennifer

(1311.6729)

Anomalies in $B \rightarrow K(K^*)\mu^+\mu^-$, $B_s \rightarrow \mu^+\mu^-$

Matias et al
 Altmannshofer+Straub,
 Jäger et al
 Bobeth et al
 Hiller+Schmaltz
 Hurth et al

Reproduced with

$$C_9^{\text{NP}} \approx -C_{10}^{\text{NP}} < 0$$

$$b \rightarrow s\nu\bar{\nu}$$

$$: C_L = C_L^{\text{SM}} + C_L^{\text{NP}}$$

$$C_L^{\text{SM}} < 0$$

BGNS
 (1409.4557)

$$\begin{aligned} Z' &: C_L^{\text{NP}} = \frac{C_9^{\text{NP}} - C_{10}^{\text{NP}}}{2} \approx C_9^{\text{NP}} \rightarrow \text{enhancement of } B \rightarrow K(K^*)\nu\bar{\nu} \\ Z &: C_L^{\text{NP}} = C_{10}^{\text{NP}} \approx -C_9^{\text{NP}} \rightarrow \text{suppression of } B \rightarrow K(K^*)\nu\bar{\nu} \end{aligned}$$

If

$$C_9^{\text{NP}} < 0 \quad C_{10}^{\text{NP}} = 0$$



enhancement by Z'
 no effect from Z

See also leptoquark models: Hiller + Schmaltz, Fajfer et al.
 Nardecchia et al.

3 Correlated Anomalies

(LHCb)

Matias et al
 Altmannshofer+Straub,
 Jäger et al
 Bobeth et al
 Hiller+Schmaltz
 Hurth et al
 Fajfer et al
 Crivellin et al
 Nardecchia et al

$$R_{K\mu\mu} = \frac{\text{Br}(B^+ \rightarrow K^+ \mu^+ \mu^-)^{[15,22]}}{\text{Br}(B^+ \rightarrow K^+ \mu^+ \mu^-)_{SM}^{[15,22]}} < 1$$

$$R_{K^*\mu\mu} = \frac{\text{Br}(B^0 \rightarrow K^{*0} \mu^+ \mu^-)^{[15,19]}}{\text{Br}(B^0 \rightarrow K^{*0} \mu^+ \mu^-)_{SM}^{[15,19]}} < 1$$

$$R_{\mu\mu} = \frac{\text{Br}(B_s \rightarrow \mu^+ \mu^-)}{\text{Br}(B_s \rightarrow \mu^+ \mu^-)_{SM}} < 1$$

Can be reproduced partly by Z but fully by Z' with left-handed quark FCNC couplings.

$$C_9^{NP} \approx -C_{10}^{NP}$$

(V) (A) $\mu^+ \mu^-$

$$R_{K^*(K)} \equiv R_{K^*(K)\nu\bar{\nu}}$$

can distinguish between Z and Z' solution

$$R_{K^*(K)}(Z') > 1$$

$$R_{K^*(K)}(Z) < 1$$

Z' wins over Z

**Z' can reproduce all LHCb anomalies
including $\text{Br}(\text{B}^+ \rightarrow \text{K}^+ \mu^+ \mu^-) < \text{Br}(\text{B}^+ \rightarrow \text{K}^+ e^+ e^-)$**

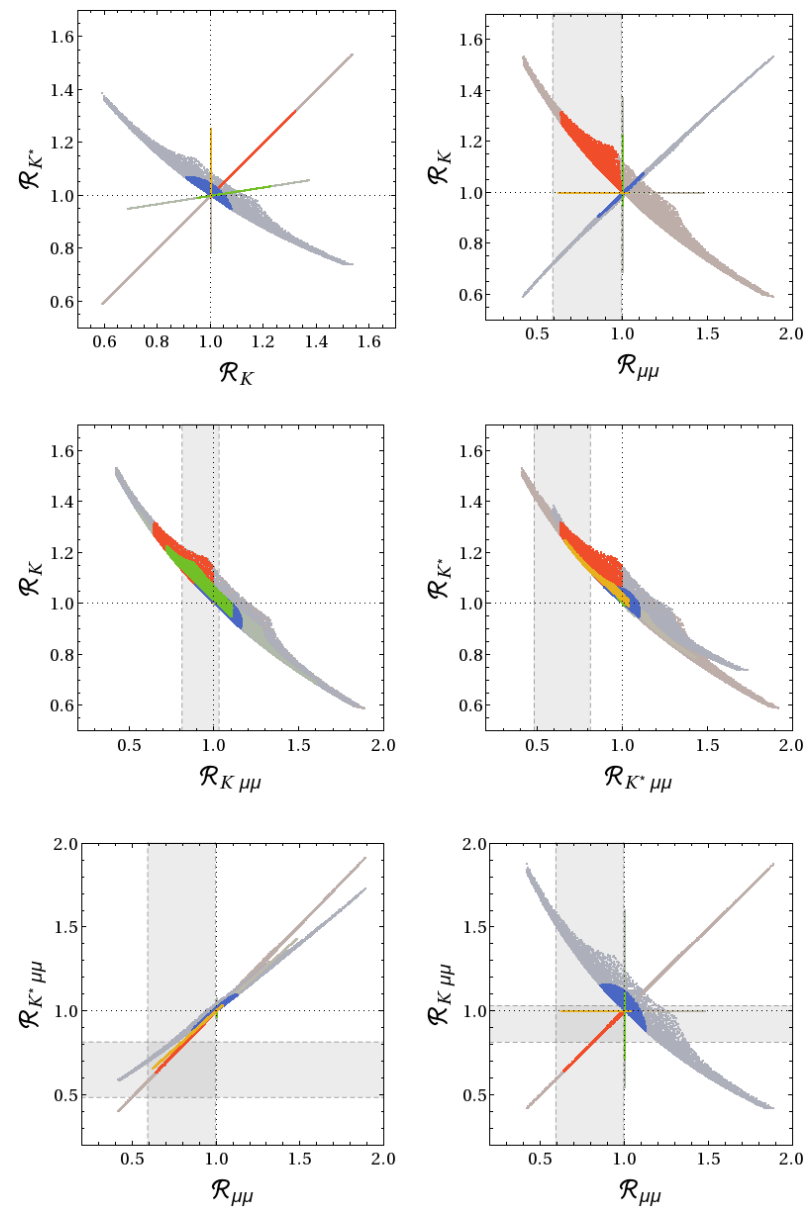
$$\mathbf{Z} : \left\{ g_V^{\mu\mu}(\mathbf{Z}) \text{ small} \right\} \rightarrow \mathbf{C}_9^{\text{NP}} \text{ small}$$
$$\text{Br}(\text{B}^+ \rightarrow \text{K}^+ \mu^+ \mu^-) \approx \text{Br}(\text{B}^+ \rightarrow \text{K}^+ e^+ e^-)$$

$B \rightarrow K(K^*)\nu\bar{\nu}, B \rightarrow \mu^+\mu^-, B \rightarrow K(K^*)l^+l^-$

AJB
Girrbach-Noe
Niehoff
Straub

1409.4557

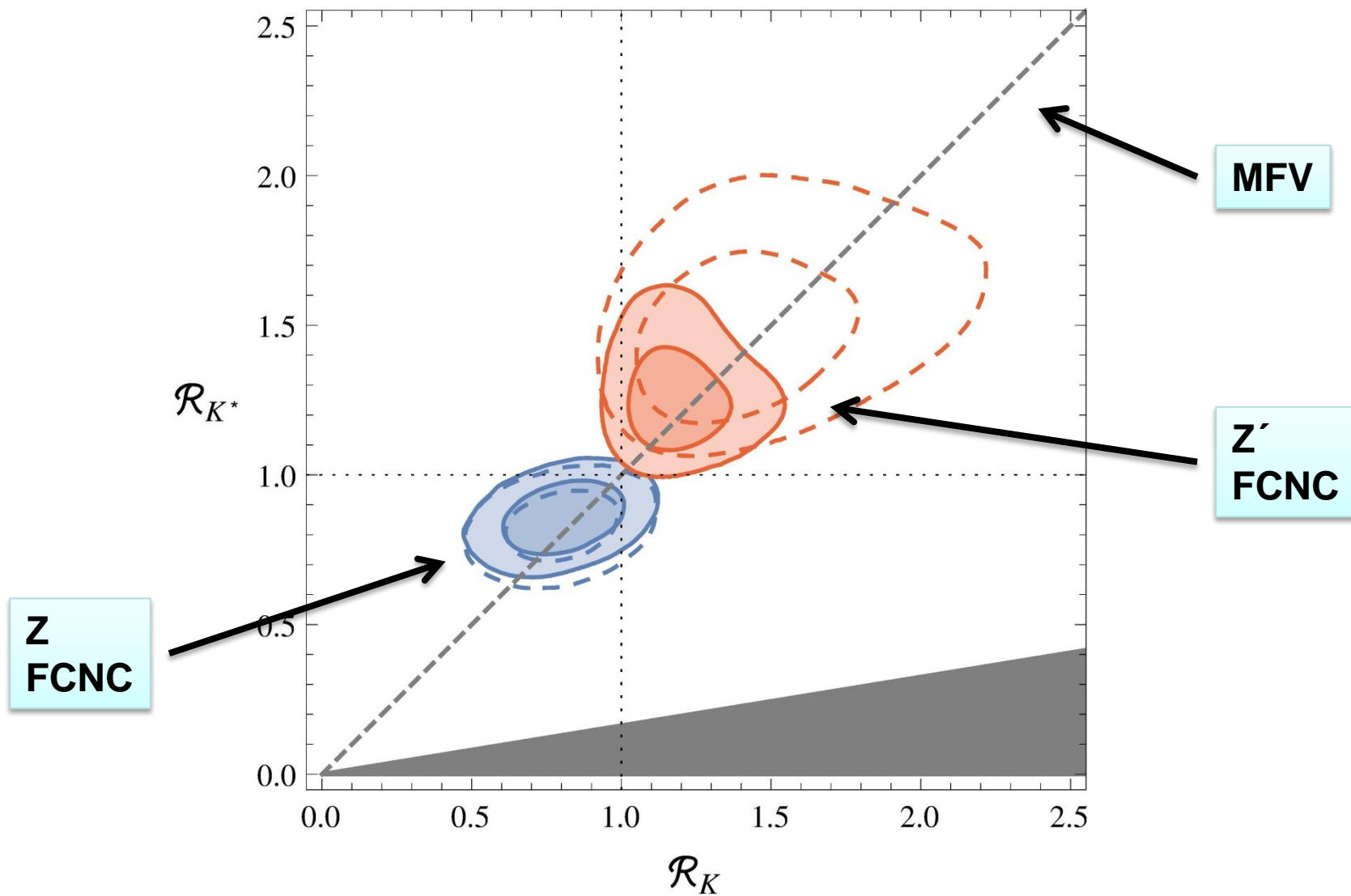
Z'



- LHS
- RHS
- LRS
- ALRS

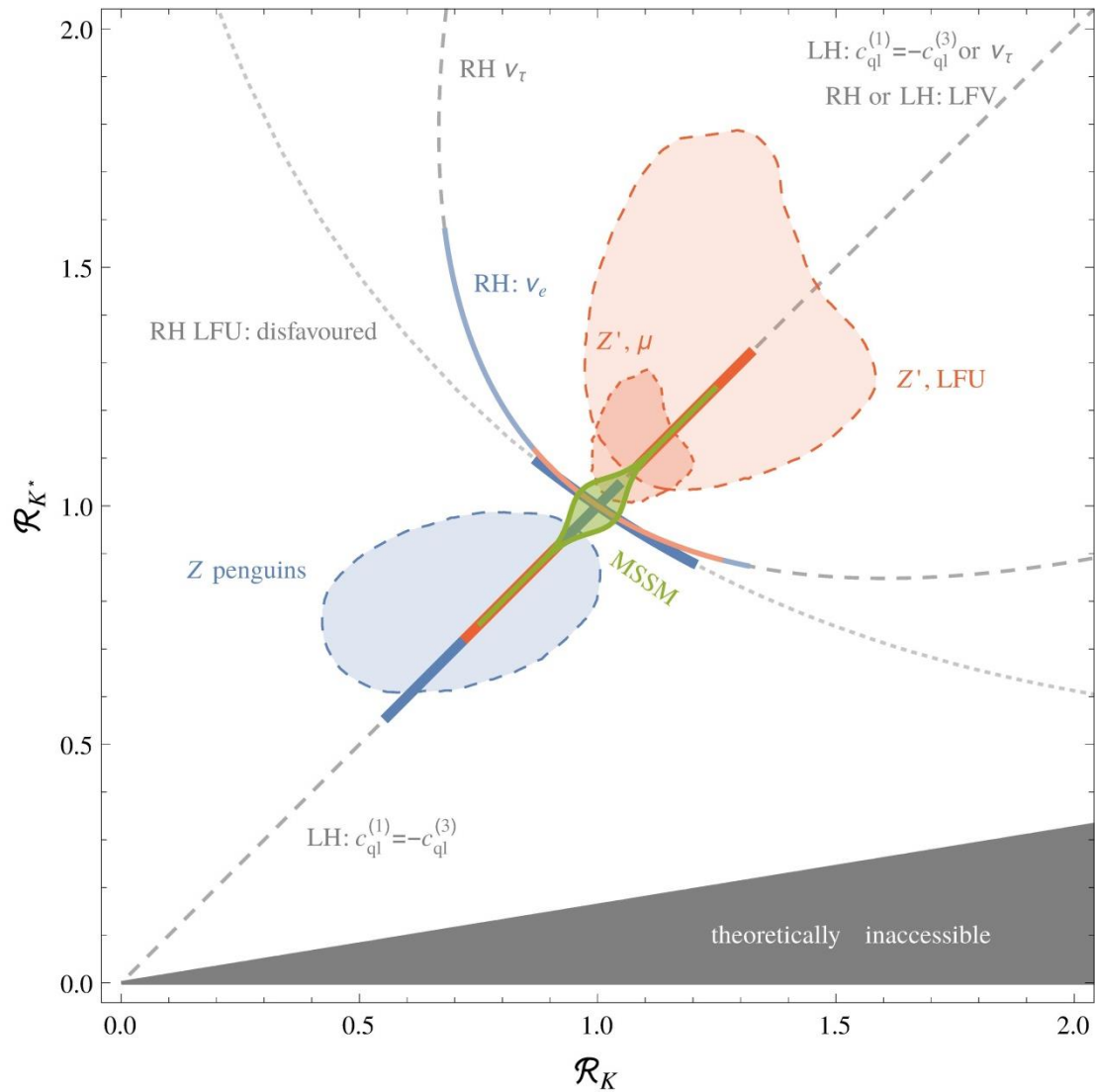
$$B \rightarrow K(K^*)\nu\bar{\nu}$$

BGNS 1409.4557



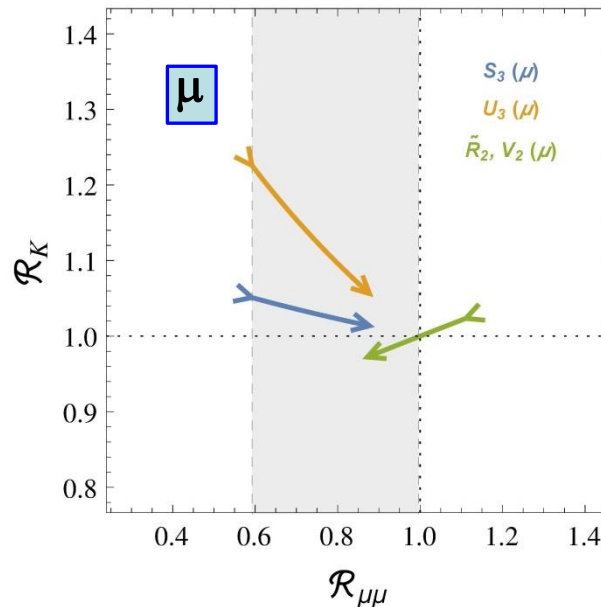
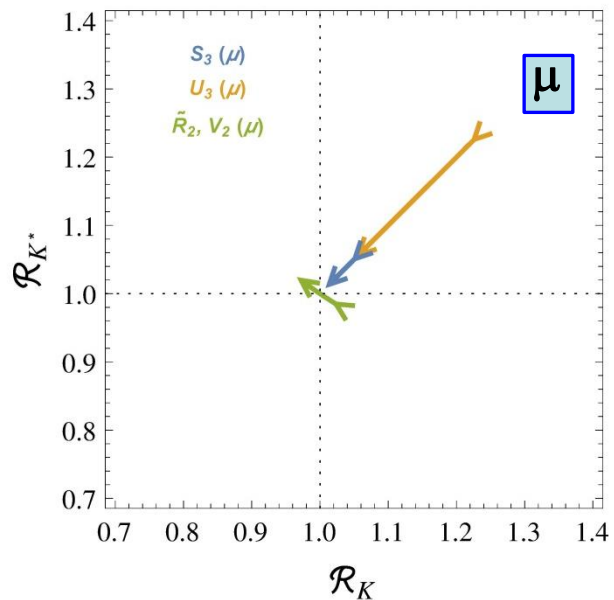
$B \rightarrow K(K^*)\nu\bar{\nu}$

BGNS 1409.4557

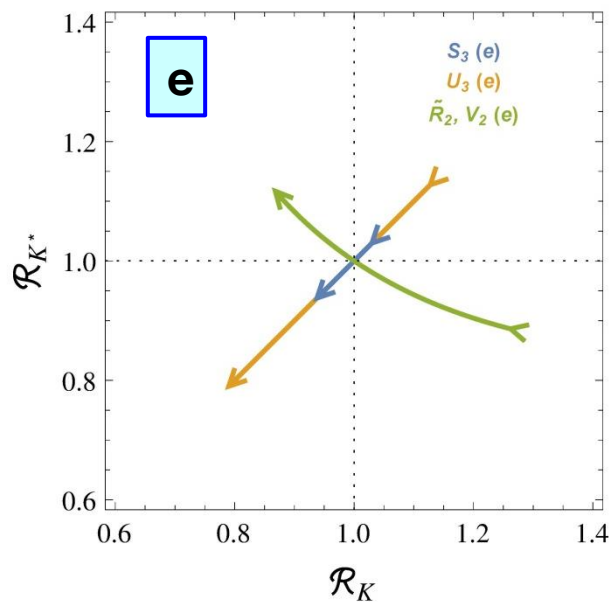


Leptoquarks at Work

BGNS, 1409.4557



Large effects
for τ and S_1



$SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$

Spin

S_3	$\bar{3}$	3	$1/3$	0
\bar{R}_2	3	2	$1/6$	0
U_3	3	3	$2/3$	1
V_2	$\bar{3}$	2	$5/6$	1
S_1	$\bar{3}$	1	$1/3$	0

Summary on $B \rightarrow K(K^*)\nu\bar{\nu}$

$R_K \neq R_{K^*}$ will be problematic for :

MFV, Z' with LH couplings, certain PC scenarios
MSSM, $SU(2)_L$ singlet or triplets LQ, 331

NP
at most
 $\pm 30\%$

$R_K \neq R_{K^*}$ can distinguish between Z and Z'
explanation of $b \rightarrow sl^+l^-$ anomalies

$b \rightarrow s\mu^+\mu^-$ Data



max $\pm 60\%$ NP for LFU

$\pm 20\%$ if NP only in muons and ν_μ

In the presence of LF non-universality NP in $b \rightarrow s\nu\bar{\nu}$
could be LARGE ! Examples: NP only in (τ, ν_τ) Leptoquarks

Main message:

Finding small NP effects in $b \rightarrow s\mu^+\mu^-$ would not
imply necessarily small NP effects in $b \rightarrow s\nu\bar{\nu}$

Intermezzo

Status of $B_{s,d} \rightarrow \mu^+ \mu^-$

The first
NLO QCD
Calculation
of $B_{s,d} \rightarrow \mu^+ \mu^-$

Buchalla + AJB (Nucl. Phys. B400 (1993) 225)

- Reduction of μ_t dependence in $m_t(\mu_t)$
- Finding missing factor of two in branching ratios.



Values of $\text{Br}(B_s \rightarrow \mu^+ \mu^-)_{\text{SM}} \sim 3 - 4 \cdot 10^{-9}$ were
 $\text{Br}(B_d \rightarrow \mu^+ \mu^-)_{\text{SM}} \sim 1 \cdot 10^{-10}$ with us
for last
15 years

Theoretical Improvements
over years

: Buchalla, AJB; Misiak, Urban (~1998)

September
2013

Recently: full NLO Electroweak, NNLO QCD

Bobeth, Gorbahn, Hermann, Misiak, Stamou, Steinhauser

Data (LHCb+CMS)

$$\bar{\text{Br}}(B_s \rightarrow \mu^+ \mu^-)_{\text{SM}} = (3.65 \pm 0.23) \cdot 10^{-9}$$

$$\text{Br}(B_d \rightarrow \mu^+ \mu^-)_{\text{SM}} = (1.06 \pm 0.09) \cdot 10^{-10}$$

$$(2.8 \pm 0.7) \cdot 10^{-9}$$

$$(3.6^{+1.6}_{-1.4}) \cdot 10^{-10}$$

Warning: $|V_{cb}|$ ($|V_{ts}|$) Dependence

BGHMSS

use

$$|V_{cb}|_{\text{incl}} \approx 42 \cdot 10^{-3}$$

$$\Rightarrow \bar{\text{Br}}(B_s \rightarrow \mu\bar{\mu})_{\text{SM}} = (3.65 \pm 0.23) \cdot 10^{-9} \\ (2.8 \pm 0.7) \cdot 10^{-9} \\ \text{(LHCb+CMS)}$$

But
for

$$|V_{cb}|_{\text{excl}} \approx 39 \cdot 10^{-3} \Rightarrow \bar{\text{Br}}(B_s \rightarrow \mu\bar{\mu})_{\text{SM}} \approx (3.1 \pm 0.2) \cdot 10^{-9}$$

Different
Route

(AJB 2003)
(Knegjens 2014)

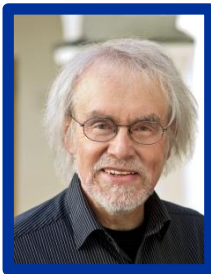
$$\bar{\text{Br}}(B_s \rightarrow \mu\bar{\mu})_{\text{SM}} = (3.5 \pm 0.2) \cdot 10^{-9} \left[\frac{(\Delta M_s)^{\text{SM}}}{(\Delta M_s)^{\text{Data}}} \right] \left[\frac{1.33}{\hat{B}_s} \right]$$

(No V_{cb} , F_{B_s} dependence)

↑
Lattice

2.

News on $K^+ \rightarrow \pi^+ \nu \bar{\nu}$, $K_L \rightarrow \pi^0 \nu \bar{\nu}$, ε'/ε



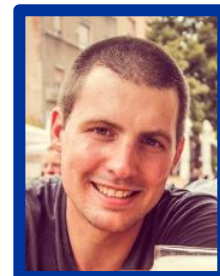
AJB



D. Buttazzo



J. Girrbach-Noe



R. Knegjens

$K^+ \rightarrow \pi^+ \nu \bar{\nu}$ and $K_L \rightarrow \pi^0 \nu \bar{\nu}$ in the SM

QCD Corrections:

NLO Buchalla, AJB; Misiak, Urban (93, 98)
 NNLO AJB, Gorbahn, Haisch, Nierste (2005)

NLO EW Corrections:

Large m_t : Buchalla, AJB (1997)
 Exact NLO (m_t): Brod, Gorbahn, Stamou (2010)
 " " (m_c): Brod, Gorbahn (2008)

LD Effects:

Isidori, Mescia, Smith (2005)
 Mescia, Smith (2007)

+ Isospin breaking corrections



TH uncertainties at the level of 2% in BR

Unique in Flavour Physics !!

But significant parametric uncertainties

due to $|V_{ub}|, |V_{cb}|, \gamma$

Data

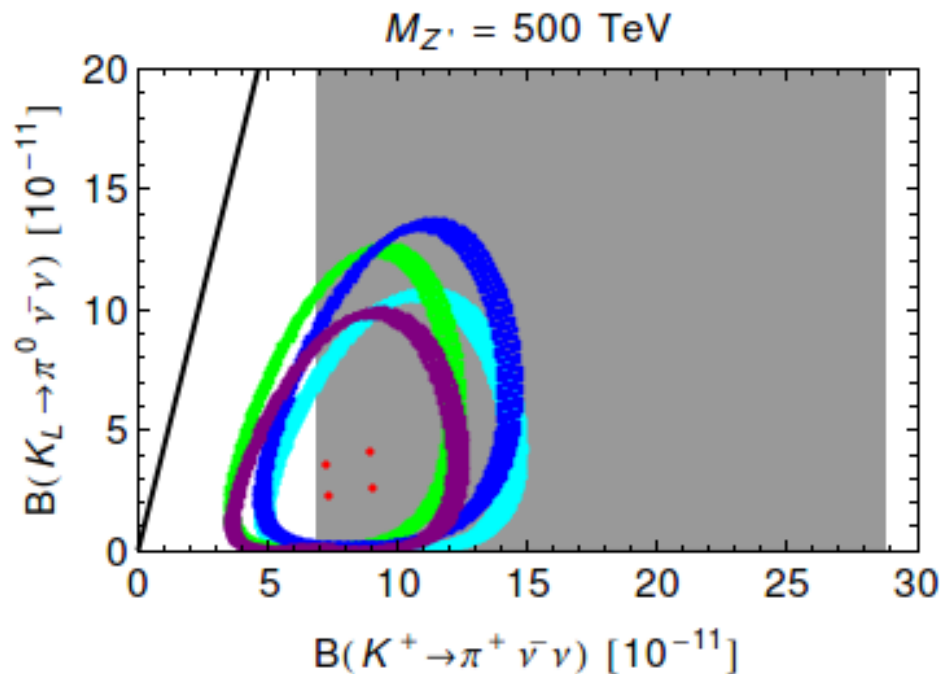
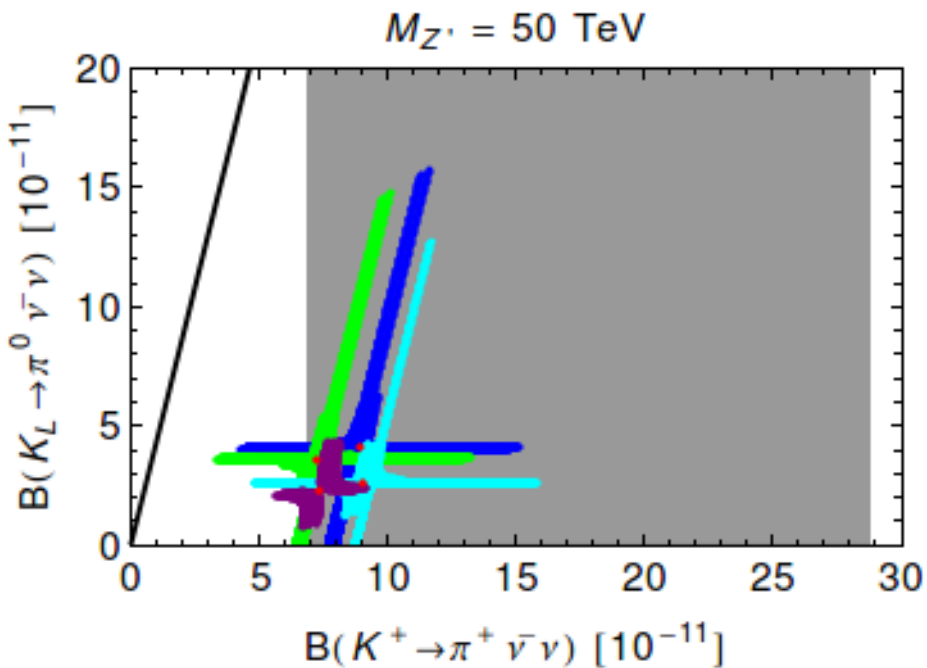
$$\begin{aligned} \text{Br}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) &= (17.3 \pm 11) \cdot 10^{-11} \\ \text{Br}(K_L \rightarrow \pi^0 \nu \bar{\nu}) &\leq 2.6 \cdot 10^{-8} \end{aligned}$$

General Properties

- 1.** $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ **CP-conserving**
- 2.** $K_L \rightarrow \pi^0 \nu \bar{\nu}$ **CP-violating**
- 3.** **Both sensitive to New Physics (NP)**
 $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ bounded by $K_L \rightarrow \mu^+ \mu^-$
 $K_L \rightarrow \pi^0 \nu \bar{\nu}$ bounded by ε'/ε
- 4.** **The correlation between $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ and $K_L \rightarrow \pi^0 \nu \bar{\nu}$ depends on the ε_K constraint (Blanke 0904.2528)**
- 5.** **Can probe scales far above LHC.**

Heavy Z' at Work

AJB, Buttazzo, Girrbach-Noe, Kneijens, 1407.0728



ϵ_K constraint

General discussion:
Blanke 0904.2528

No ϵ_K constraint

Motivations for New Analysis

1. NA62 in progress: 10% measurement of $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ in 2018.

2. Stress CKM uncertainties in $\text{Br}(K^+ \rightarrow \pi^+ \nu \bar{\nu})$, $\text{Br}(K_L \rightarrow \pi^0 \nu \bar{\nu})$

3. Point out correlation between

$K^+ \rightarrow \pi^+ \nu \bar{\nu}$, $B_s \rightarrow \mu^+ \mu^-$ and γ
(NA62) (LHCb+CMS) (LHCb)

Basically
no CKM
uncertainties

4. Update correlation between

$K^+ \rightarrow \pi^+ \nu \bar{\nu}$, $K_L \rightarrow \pi^0 \nu \bar{\nu}$ and β

(Buchalla, AJB, 94)
(AJB, Fleischer, 00)

5. Use most recent lattice input for CKM

6. Provide the present best value in SM

Strategy A: Use Tree Level Determination of CKM

$$|V_{ub}|_{\text{excl}} = (3.72 \pm 0.14) \cdot 10^{-3}$$

$$|V_{cb}|_{\text{excl}} = (39.36 \pm 0.75) \cdot 10^{-3}$$

$$|V_{ub}|_{\text{incl}} = (4.40 \pm 0.25) \cdot 10^{-3}$$

$$|V_{cb}|_{\text{incl}} = (42.21 \pm 0.78) \cdot 10^{-3}$$



$$|V_{ub}|_{\text{avg}} = (3.88 \pm 0.29) \cdot 10^{-3}$$

$$|V_{cb}|_{\text{avg}} = (40.7 \pm 1.4) \cdot 10^{-3}$$

$$\gamma = \left(73.2 \begin{matrix} +6.3 \\ -7.0 \end{matrix} \right)^\circ$$

$$\overline{\text{Br}}(B_s \rightarrow \mu^+ \mu^-) = (3.4 \pm 0.3) \cdot 10^{-9}$$
$$\overline{\text{Br}}(B_s \rightarrow \mu^+ \mu^-)_{\text{exp}} = (2.8 \pm 0.7) \cdot 10^{-9}$$

$$\text{Br}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) = (8.4 \pm 1.0) \cdot 10^{-11}$$
$$\text{Br}(K_L \rightarrow \pi^0 \nu \bar{\nu}) = (3.4 \pm 0.6) \cdot 10^{-11}$$



CKM Uncertainties

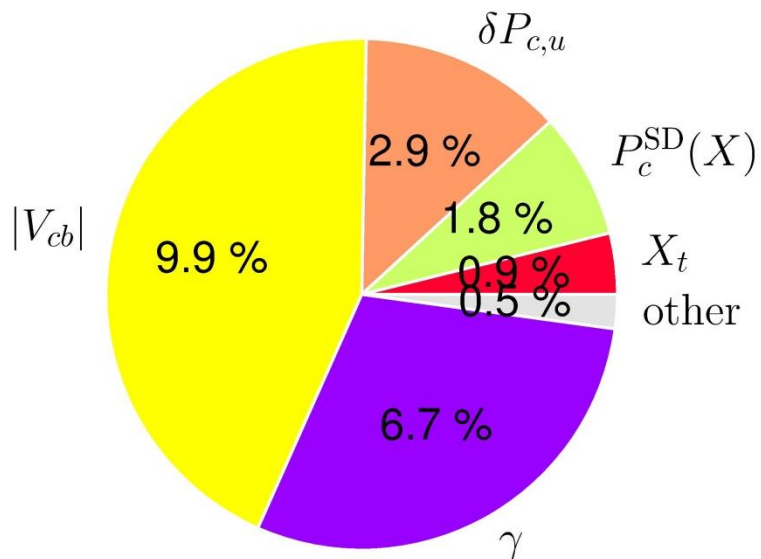
$$\text{Br}(\text{K}^+ \rightarrow \pi^+ \nu \bar{\nu}) = (8.39 \pm 0.30) \cdot 10^{-11} \left[\frac{|\mathbf{V}_{cb}|}{0.0407} \right]^{2.8} \left[\frac{\gamma}{73.2^\circ} \right]^{0.71}$$

$$\text{Br}(\text{K}_L \rightarrow \pi^0 \nu \bar{\nu}) = (3.36 \pm 0.09) \cdot 10^{-11} \left[\frac{|\mathbf{V}_{ub}|}{3.88 \cdot 10^{-3}} \right]^2 \left[\frac{|\mathbf{V}_{cb}|}{0.0407} \right]^2 \left[\frac{\sin \gamma}{\sin(73.2)} \right]^2$$

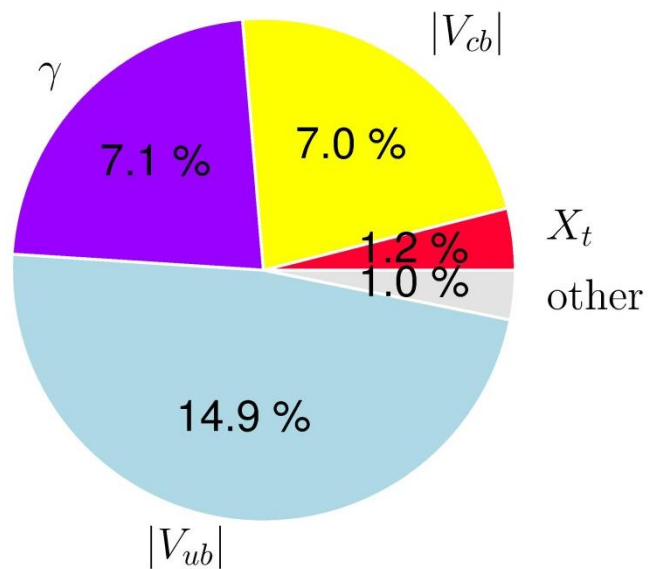
$$\text{Br}(\text{K}^+ \rightarrow \pi^+ \nu \bar{\nu}) = (65.3 \pm 3.1) \left[\bar{\text{Br}}(\text{B}_s \rightarrow \mu^+ \mu^-) \right]^{1.4} \left[\frac{\gamma}{70^\circ} \right]^{0.71} \left[\frac{227 \text{ MeV}}{F_{\text{B}_s}} \right]^{2.8}$$

Error Budgets

$$\mathcal{B}(K^+ \rightarrow \pi^+ \nu \bar{\nu})$$



$$\mathcal{B}(K_L \rightarrow \pi^0 \nu \bar{\nu})$$



Update: 1503.02693

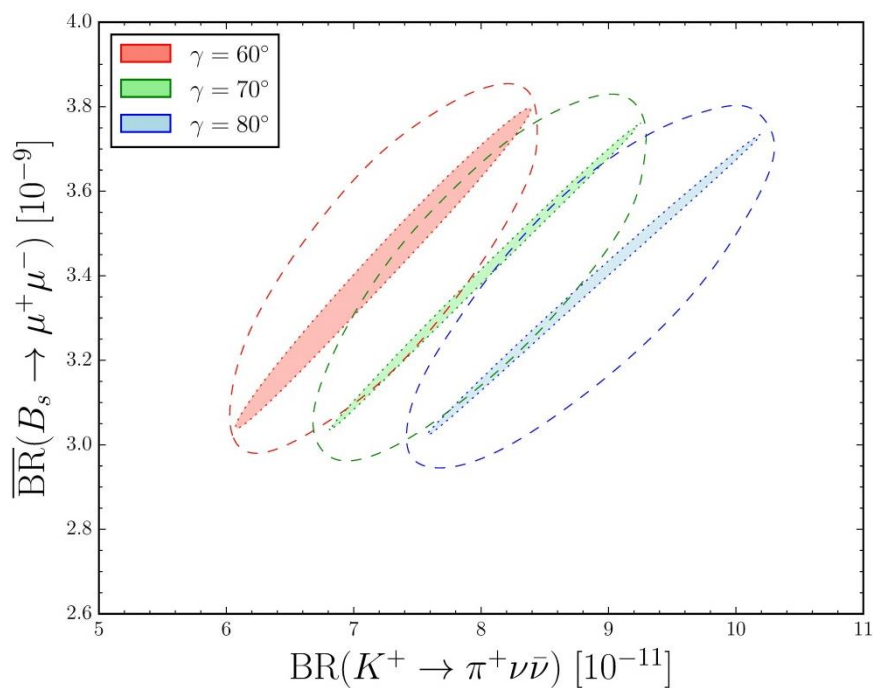
$$P_c = 0.404 \pm 0.024$$

$$X_t = 1.481 \pm 0.005_{\text{th}} \pm 0.008_{\text{exp}}$$

Correlations

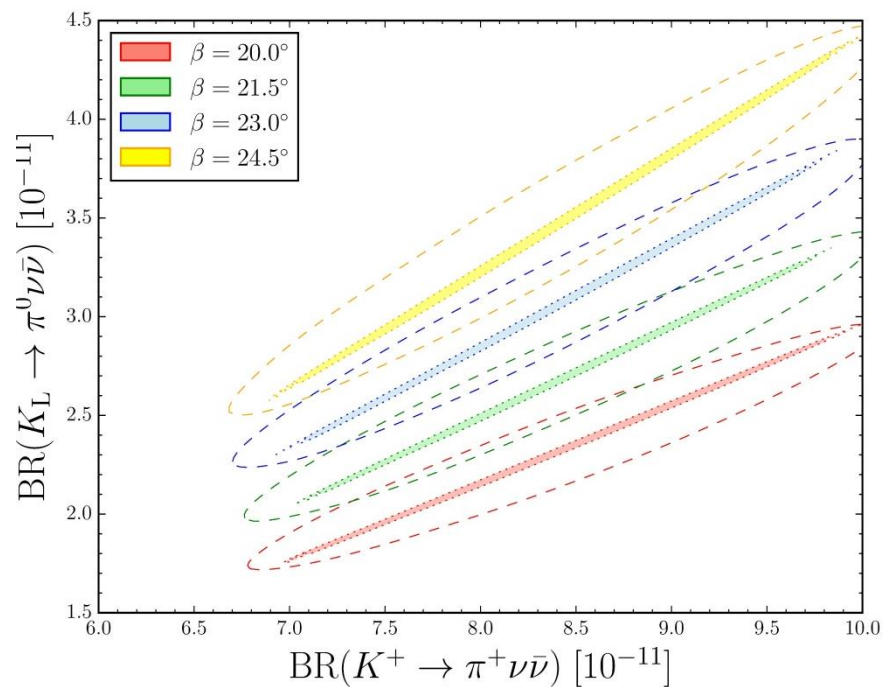
$$B_s \rightarrow \mu^+ \mu^-, K^+ \rightarrow \pi^+ \nu \bar{\nu}, \gamma$$

BBGK (2015)



$$K^+ \rightarrow \pi^+ \nu \bar{\nu}, K_L \rightarrow \pi^0 \nu \bar{\nu}, \beta$$

Buchalla, AJB (94)



Strategy B: use ε_K , ΔM_s , ΔM_d , $S_{\psi K_s}$

$$|V_{cb}| = (42.4 \pm 1.0) \cdot 10^{-3}$$

$$|V_{ub}| = (3.61 \pm 0.13) \cdot 10^{-3}$$

$$\gamma = (69.5 \pm 5.0)^\circ$$

$$\text{Br}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) = (9.1 \pm 0.7) \cdot 10^{-11}$$
$$\text{Br}(K_L \rightarrow \pi^0 \nu \bar{\nu}) = (3.0 \pm 0.3) \cdot 10^{-11}$$

$$\text{UTfit} : |V_{cb}| = (41.7 \pm 0.6) \cdot 10^{-3}$$

$$|V_{ub}| = (3.63 \pm 0.12) \cdot 10^{-3}$$

$$\text{CKMfitter} : |V_{cb}| = (41.2 \pm 1.0) \cdot 10^{-3}$$

$$|V_{ub}| = (3.55 \pm 0.16) \cdot 10^{-3}$$

New Analysis of ε'/ε in the SM

BBGK

QCD Penguins:

$$B_6^{(1/2)} = 1.0 \pm 0.2$$

Large N

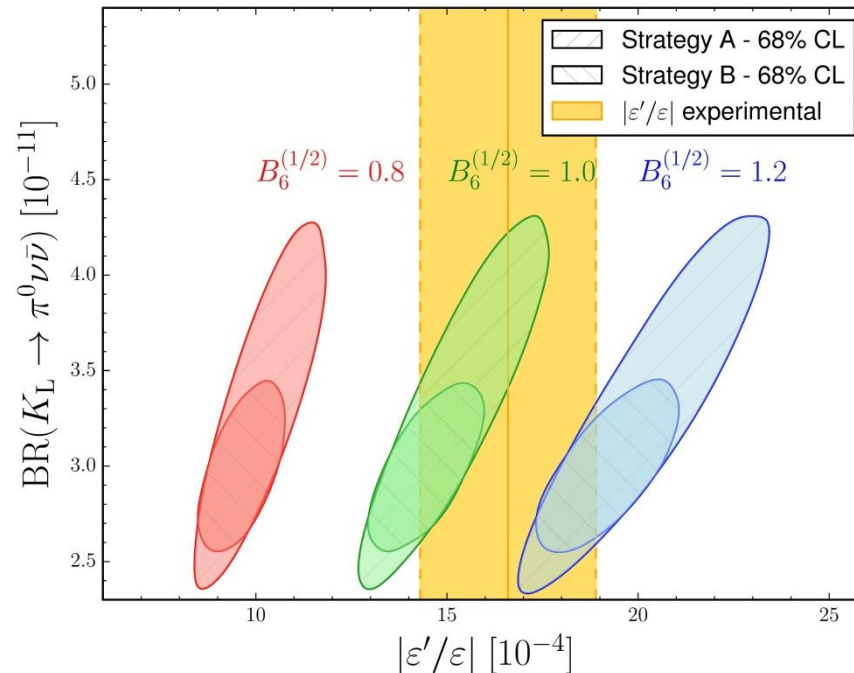
EW Penguins:

$$B_8^{(2/3)} = 0.76 \pm 0.05$$

RBC-UKQCD

$$\text{Re}(\varepsilon'/\varepsilon) = (16.5 \pm 2.6) \cdot 10^{-4}$$

NA48, KTeV



$K^+ \rightarrow \pi^+ \nu \bar{\nu}$ and $K_L \rightarrow \pi^0 \nu \bar{\nu}$ beyond SM

Review Mod. Phys.: AJB, Schwab, Uhlig (2008) (0405132)
AJB, Buttazzo, Knegjens: hep-ph-1504.xxxx

MFV, $U(2)^3$:

20-30% effects, strong correlation between K^+ and K_L

No MFV :

Correlation depends on the presence or absence of ε_K constraint, size on ε'/ε , $K_L \rightarrow \mu^+ \mu^-$

FCNCs Z :

Enhancements by factors 2-3 over SM still possible (ε'/ε constraint important)

FCNCs Z' :

Still larger enhancements possible as ε'/ε constraint can be eliminated in a model independent analysis but not in specific models with known flavour diagonal quark couplings.

**More info
in BBK to
appear soon**

see Rob Knegjens (Moriond)

Finale: Vivace !

Main Message

Rare K, B_s, B_d Decays will play crucial role in identifying New Physics hopefully present on the route

Attouniverse → Zeptouniverse

Coming Years

: Flavour Precision Era

**LHC
Upgrade
E = 14 TeV
(CERN)**

**Precision
B_{d,s} – Meson
Decays
LHC
KEK (Japan)**

**$K^+ \rightarrow \pi^+ \nu \bar{\nu}$ ($\sim 10^{-10}$) (CERN)
 $K_L \rightarrow \pi^0 \nu \tilde{\nu}$ ($\sim 3 \cdot 10^{-11}$) J-PARC
(Japan)**

**Lepton Flavour
Violation**

$\mu \rightarrow e \gamma$
 $\mu \rightarrow e e e$

**Electric
Dipole
Moments**

$(g-2)_\mu$

**Improved
Lattice
Gauge Theory
Calculations**

Neutrinos

**Exciting Times are just
ahead of us !!!**

Superstars and Stars of Quark Flavour Physics

Superstars

$\varepsilon_K, \Delta M_s, \Delta M_d, S_{\psi K_s}$ (TH)
 $B_s \rightarrow \mu^+ \mu^-, B_d \rightarrow \mu^+ \mu^-, S_{\psi\phi}(\phi_s)$ (LHCb, CMS, ATLAS)
 $B \rightarrow K \nu \bar{\nu}, B \rightarrow K^* \nu \bar{\nu}, B \rightarrow X_s \nu \bar{\nu}$ (Belle II)
 $K^+ \rightarrow \pi^+ \nu \bar{\nu}, K_L \rightarrow \pi^0 \nu \bar{\nu}$ (NA62, J-Parc)

Stars

$B \rightarrow K^* \mu^+ \mu^- \quad B \rightarrow K \mu^+ \mu^-$
 $B \rightarrow D^* \tau \nu_\tau \quad B \rightarrow D \tau \nu_\tau$

Old Superstar

ε'/ε will strike back
provided B_6 (QCD Penguins)
will be precisely known.

B_8 (EW Penguins)
 $\approx 0.76 \pm 0.05$
(UK-QCD)

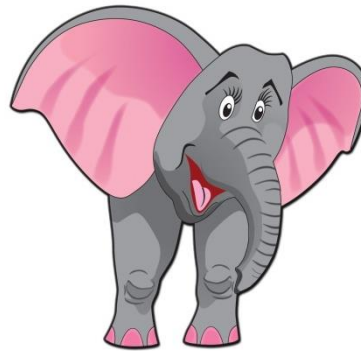
A Zeptouniverse Vision



Seen only in

$$K \rightarrow \pi \nu \bar{\nu}$$

$$B_{d,s} \rightarrow \mu^+ \mu^-$$

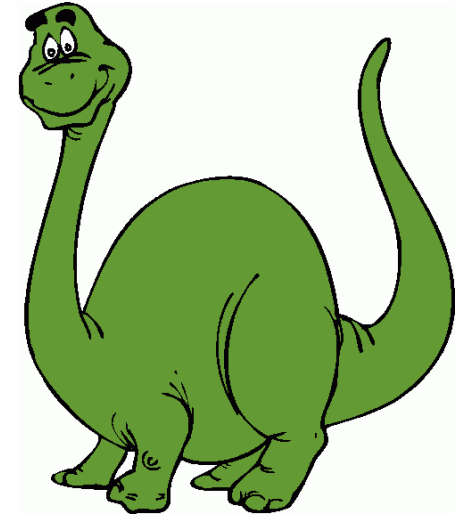


Seen in

$$\text{Rare } B_d$$

$$K \rightarrow \pi \nu \bar{\nu}$$

$$B_{d,s} \rightarrow \mu^+ \mu^-$$



Seen in

$$\text{Rare } B_s$$

$$\text{Rare } B_d$$

$$K \rightarrow \pi \nu \bar{\nu}$$

$$B_{d,s} \rightarrow \mu^+ \mu^-$$

Final Message

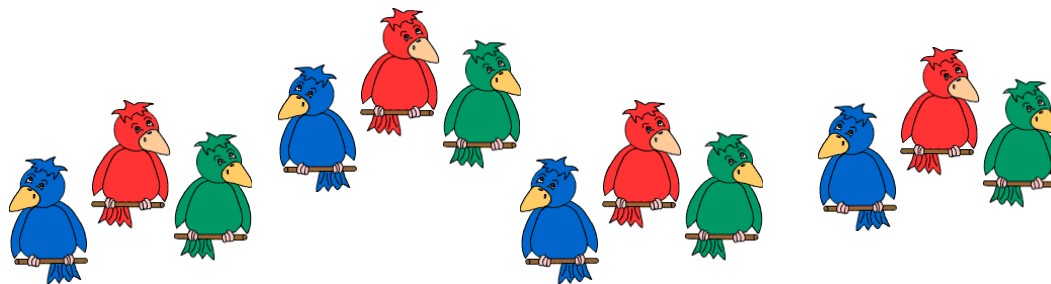
**Great hopes to
see many oases
on the way**

Attouniverse → Zeptouniverse

Final Message

**Great hopes to
see many oases
on the way**

**Attouniverse → Zeptouniverse
and**



at the LHC

Backup

Warning: $|V_{cb}|$ ($|V_{ts}|$) Dependence

BGHMSS

use

$$|V_{cb}|_{\text{incl}} \approx 42 \cdot 10^{-3}$$

$$\Rightarrow \bar{\text{Br}}(B_s \rightarrow \mu\bar{\mu})_{\text{SM}} = (3.65 \pm 0.23) \cdot 10^{-9} \\ (2.9 \pm 0.7) \cdot 10^{-9} \\ \text{(LHCb+CMS)}$$

But
for

$$|V_{cb}|_{\text{excl}} \approx 39 \cdot 10^{-3} \Rightarrow \bar{\text{Br}}(B_s \rightarrow \mu\bar{\mu})_{\text{SM}} \approx (3.1 \pm 0.2) \cdot 10^{-9}$$

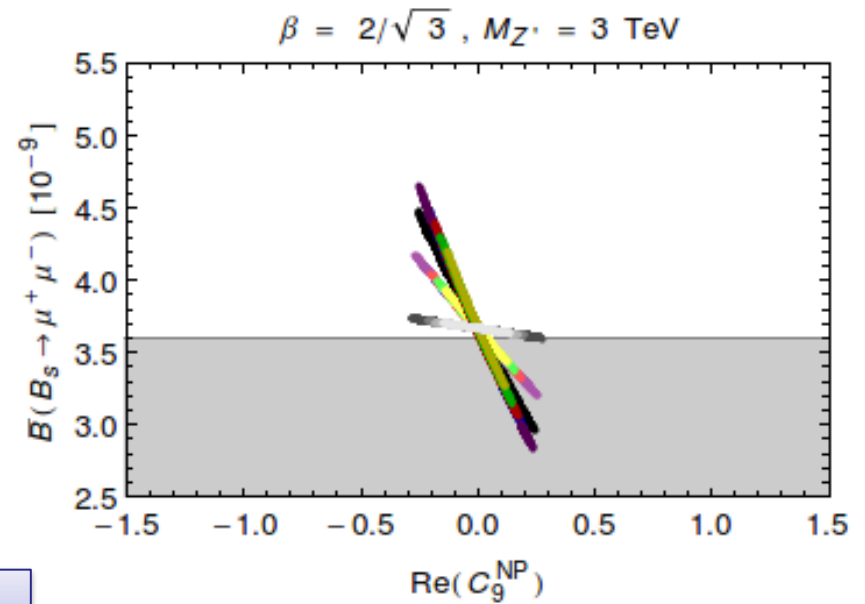
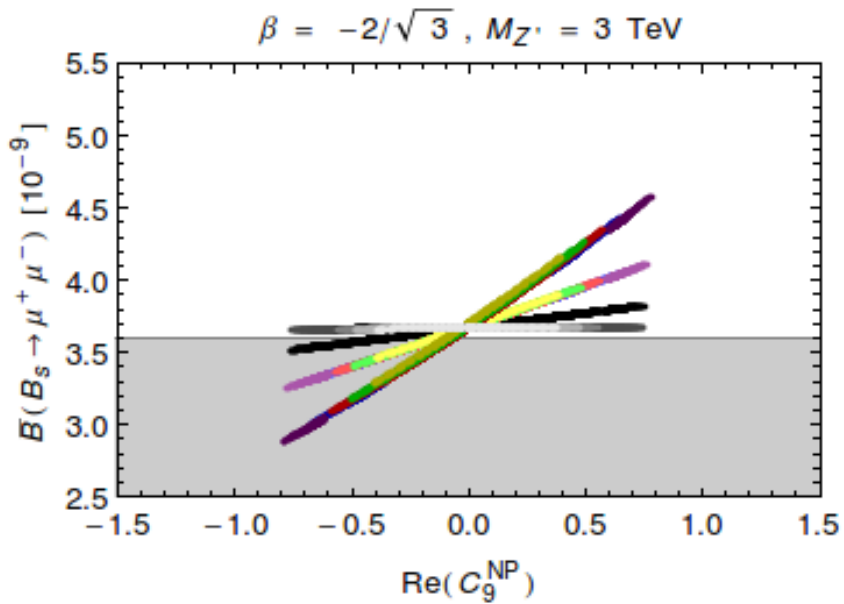
Different
Route

(AJB 2003)
(Knegjens 2014)

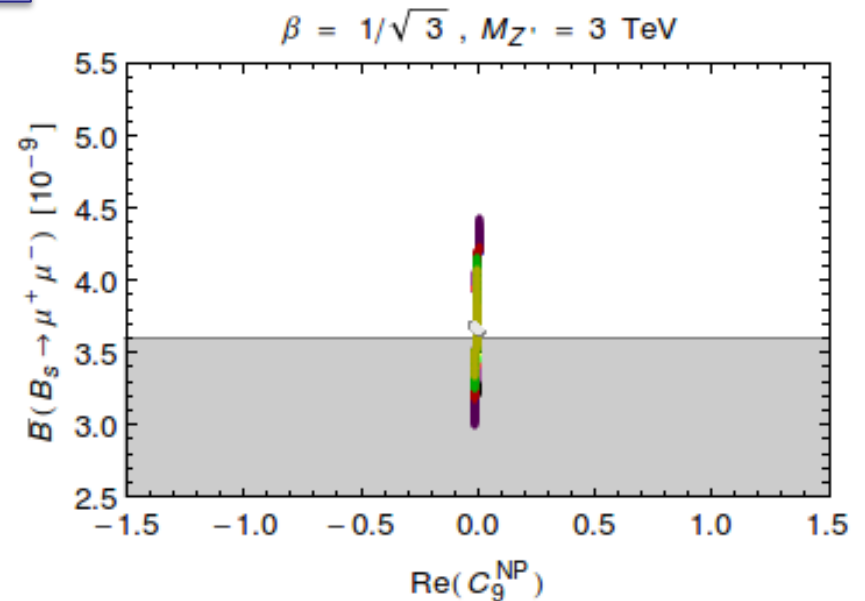
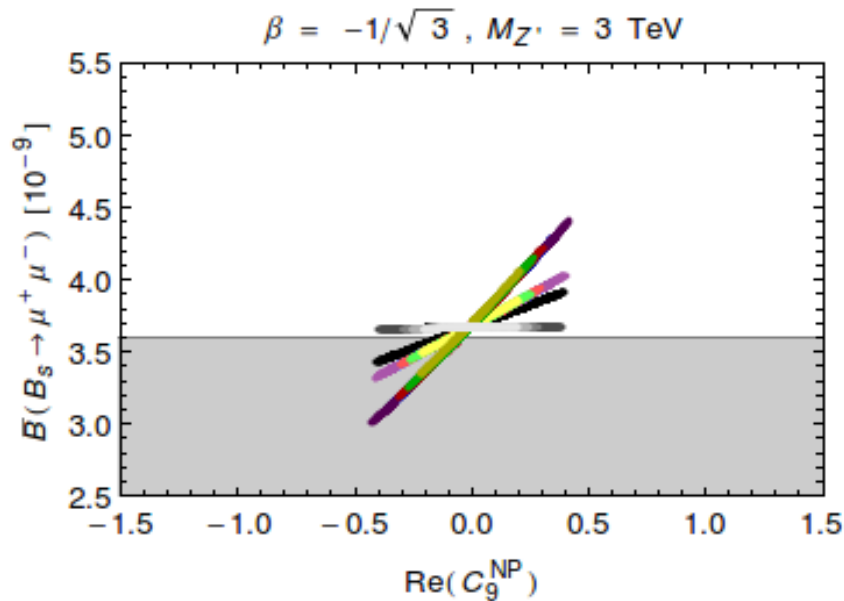
$$\bar{\text{Br}}(B_s \rightarrow \mu\bar{\mu})_{\text{SM}} = (3.5 \pm 0.2) \cdot 10^{-9} \left[\frac{(\Delta M_s)^{\text{SM}}}{(\Delta M_s)^{\text{Data}}} \right] \left[\frac{1.33}{\hat{B}_s} \right]$$

(No V_{cb} , F_{B_s} dependence)

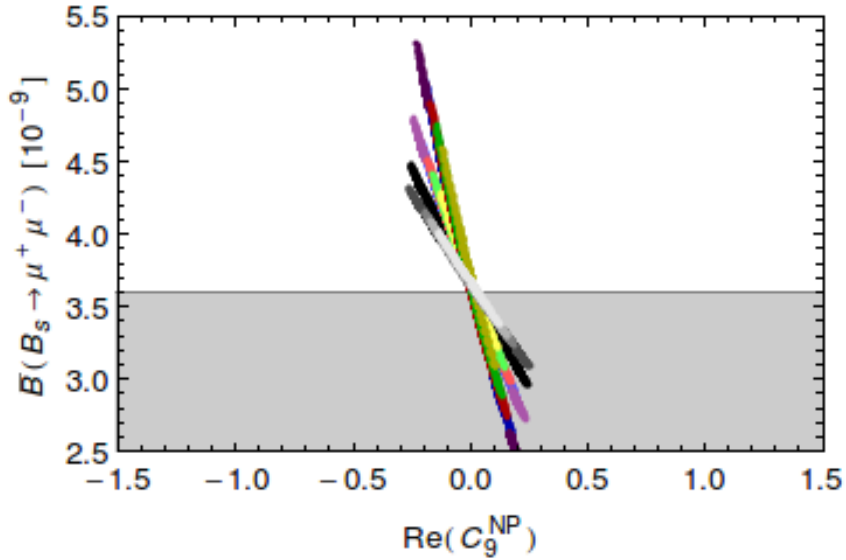
↑
Lattice



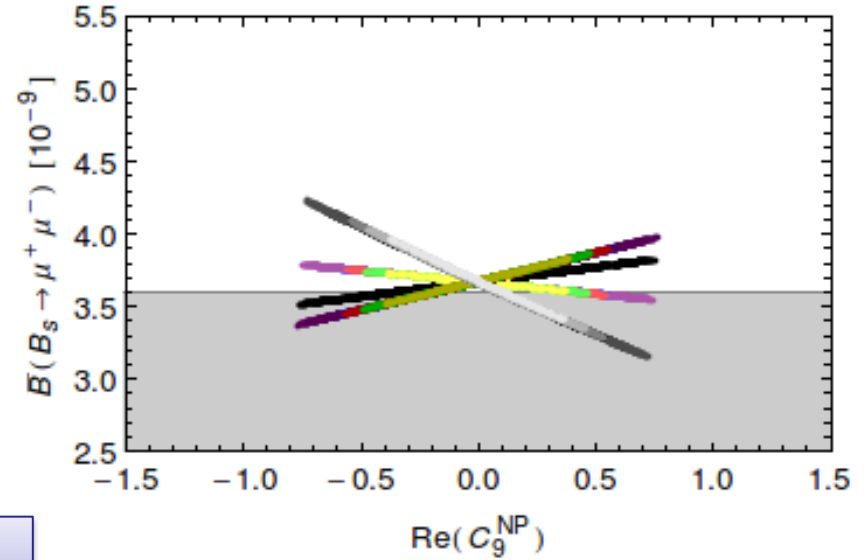
F₁



$$\beta = -2/\sqrt{3}, M_{Z'} = 3 \text{ TeV}$$

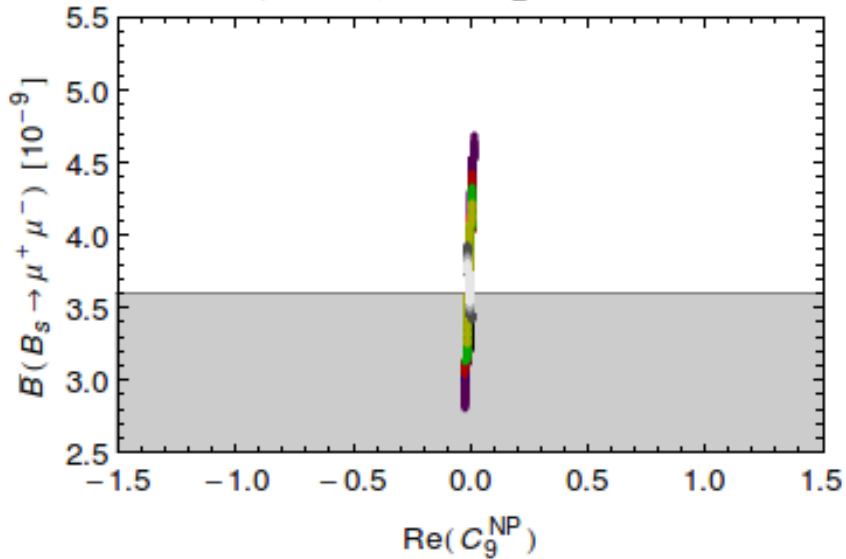


$$\beta = 2/\sqrt{3}, M_{Z'} = 3 \text{ TeV}$$

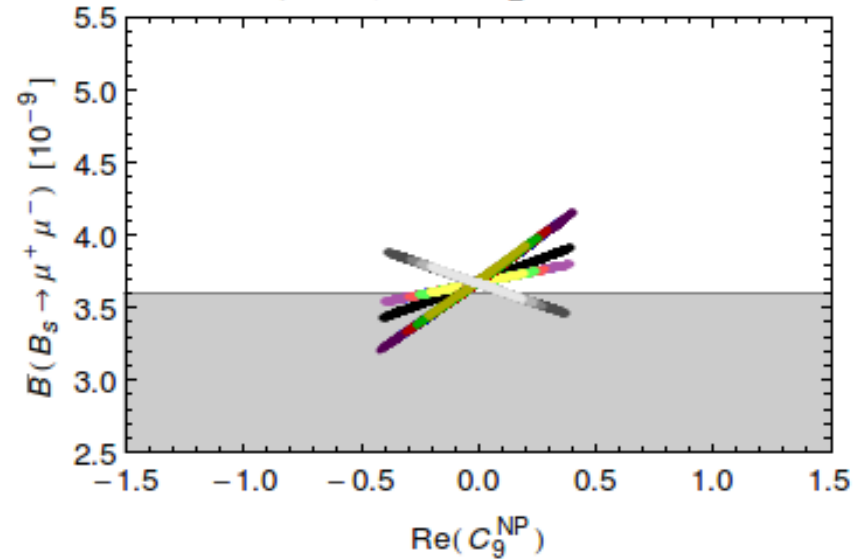


F₂

$$\beta = -1/\sqrt{3}, M_{Z'} = 3 \text{ TeV}$$



$$\beta = 1/\sqrt{3}, M_{Z'} = 3 \text{ TeV}$$



Effective Theory Approach

($\Delta F=2$)

$$H_{\text{eff}}(\Delta F = 2) = \underbrace{H_{\text{eff}}^{\text{SM}}(\Delta F = 2)}_{\text{Must be precisely known to identify NP}} + H_{\text{eff}}^{\text{NP}}(\Delta F = 2)$$

$$H_{\text{eff}}^{\text{NP}}(\Delta F = 2) = \sum_{ij} \frac{c_{ij}}{\Lambda_{\text{NP}}^2} \underbrace{Q_{ij}(\Delta F = 2)}_{\text{4-quark operators}}$$

Utfitters
Isidori, Nir, Perez

For $c_{ij} = 0(1)$ sensitivity to physics $\Lambda_{\text{NP}} > 1000 \text{ TeV}$ (LR operators)
($\varepsilon_K, \Delta M_K$)

But with the help of $\Delta F=2$ only it is not possible to learn with ET about the nature of the dynamics at Λ_{NP}

We need

$\Delta F=1$ transitions : Rare K, $B_{s,d}$, D decays



Effective Theory Approach

($\Delta F=1$)

$$H_{\text{eff}}(\Delta F = 1) = H_{\text{eff}}^{\text{SM}}(\Delta F = 1) + H_{\text{eff}}^{\text{NP}}(\Delta F = 1)$$

$$H_{\text{eff}}^{\text{NP}}(\Delta F = 1) = \sum_{ij} \frac{d_{ij}}{\Lambda_{\text{NP}}^2} Q_{ij}(\Delta F = 1)$$

Limitations
of ET :

- a) ET does not provide concrete relations between the c_{ij} ($\Delta F=2$) and d_{ij} ($\Delta F=1$) present in concrete models.

Impossible
to incorporate

Impact of $\Delta F=2$ transitions on
rare $K, B_{s,d}$ decays

Beyond
ET :

- b) ET does not provide relations between different coefficients in concrete models:

Example

331 Models

In models with Z' and Z FCNCs their contribution can interfere constructively and destructively implying very different results

AJB
De Fazio
Girrbach-Noe
1405.3850

R measurement

SM expectations: (S.Fajfer, J.Kamenik, I.Nisandzic, PRD 85, 094025 (2012))

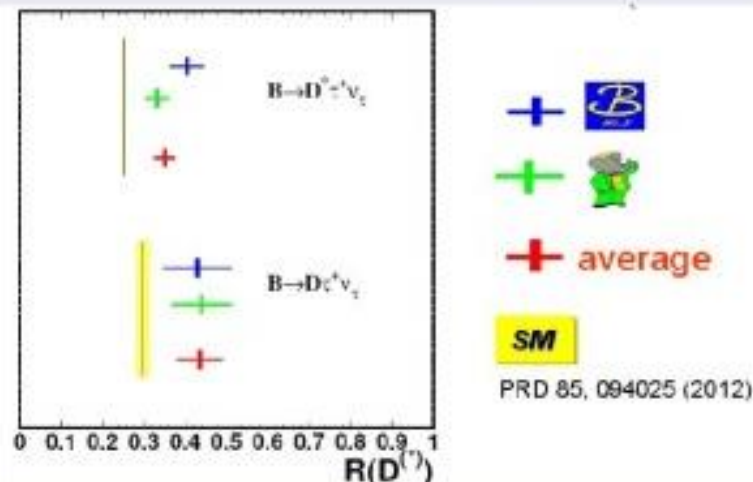
$$R(D) = 0.297 \pm 0.017, R(D^*) = 0.252 \pm 0.003$$

BABAR SM deviations

- $R(\bar{D}^*)$ 2.7σ
- $R(\bar{D})$ 2.0σ
- $R(\bar{D}^{(*)})$ 3.4σ

Belle average SM deviations

- $R(\bar{D}^*)$ 3.0σ
- $R(\bar{D})$ 1.4σ
- $R(\bar{D}^{(*)})$ 3.3σ



Belle and BABAR average deviation from SM

- $R(\bar{D}^*)$ 3.8σ
- $R(\bar{D})$ 2.4σ
- $R(\bar{D}^{(*)})$ 4.8σ

Observed deviations between observable and SM expectations for $R_{D^{(*)}}$ are not only due to improvement of experimental results but also reduction theoretical uncertainties.

LQCD expectations : A. Bailey, et al., Phys. Rev. Lett. 109, 071802, (2012), arXiv:1206.4992 [hep-ph].

$$R(D) = 0.316 \pm 0.012 \pm 0.007$$

Simple Tests in the Coming Years



Sign of $S_{\psi\phi}$



$$\frac{\text{Br}(\mathbf{B}_d \rightarrow \mu^+ \mu^-)}{\text{Br}(\mathbf{B}_s \rightarrow \mu^+ \mu^-)} = \frac{\tau(\mathbf{B}_d) m_{\mathbf{B}_d} F_{\mathbf{B}_d}^2}{\tau(\mathbf{B}_s) m_{\mathbf{B}_s} F_{\mathbf{B}_s}^2} \left| \frac{\mathbf{V}_{td}}{\mathbf{V}_{ts}} \right|^2$$



$$\frac{\text{Br}(\mathbf{B}_s \rightarrow \mu^+ \mu^-)}{\text{Br}(\mathbf{B}_d \rightarrow \mu^+ \mu^-)} = \frac{\hat{\mathbf{B}}_d \tau(\mathbf{B}_s) \Delta M_s}{\hat{\mathbf{B}}_s \tau(\mathbf{B}_d) \Delta M_d}$$



$$\text{Br}(\mathbf{K}^+ \rightarrow \pi^+ \nu \bar{\nu}); \quad \text{Br}(\mathbf{K}_L \rightarrow \pi^0 \nu \bar{\nu})$$



Lepton Flavour Violation

$$\mu \rightarrow e\gamma, \quad \mu \rightarrow 3e, \quad \tau \rightarrow 3\mu$$

$$\tau \rightarrow e\gamma, \quad \tau \rightarrow 3e$$

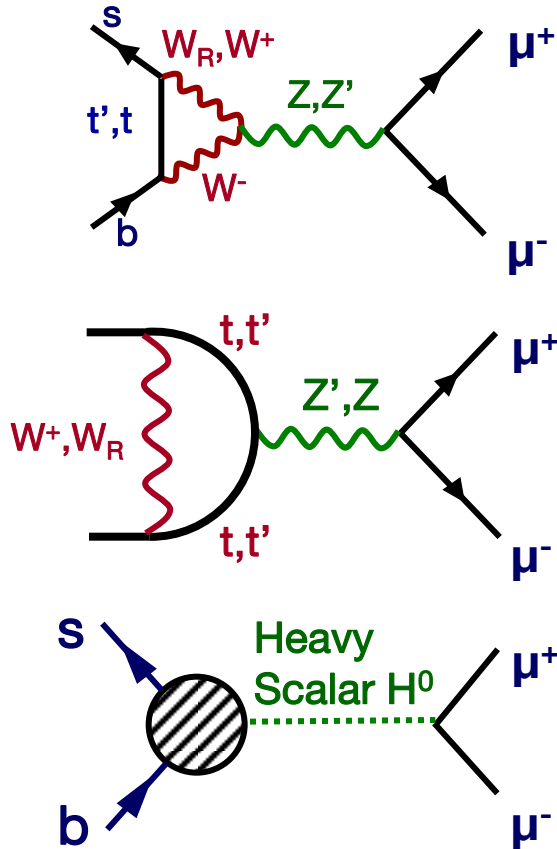
$$\tau \rightarrow \mu\gamma$$



ε'/ε provided QCD Penguin hadronic matrix under control

Standard
Candles
of
Flavour
Physics

$B_s \rightarrow \mu^+ \mu^-$ Beyond the Standard Model



Other Z-Penguins
and Boxes

$$\text{SM: } (3.2 \pm 0.2) \cdot 10^{-9}$$

**Model Independent
Limit (95% C.L.)**

$$\text{Br}(B_s \rightarrow \mu^+ \mu^-) < 5.6 \cdot 10^{-9}$$

**Altmannshofer, Paradisi,
Straub 1111.1257**

$$\frac{(\tan \beta)^6}{M_H^4}$$

in SUSY

$$\text{Br}(B_s \rightarrow \mu^+ \mu^-) < 11 \cdot 10^{-9}$$

In the case of

$$\text{Br}(B_s \rightarrow \mu^+ \mu^-) > 6 \cdot 10^{-9}$$

**distinction between Z, Z' and H^0
possible**

Minimal Effective Model with Right-Handed Currents

AJB, Gemmler, Isidori (1007.1993)

- Explains the difference $|V_{ub}|_{\text{excl}} \neq |V_{ub}|_{\text{incl}}$
- Softens $B^+ \rightarrow \tau^+ \nu_\tau$ problem (large V_{ub})

But with large $S_{\psi\phi}$ predicted: (2010)

Large $\text{Br}(B_s \rightarrow \mu^+ \mu^-)$, SM-like $\text{Br}(B_d \rightarrow \mu^+ \mu^-)$, too large $S_{\psi K_s}$

Impact of small $S_{\psi\phi}$ from LHCb (2012) (Relief !!)

SM-like $\text{Br}(B_s \rightarrow \mu^+ \mu^-)$, $\text{Br}(B_d \rightarrow \mu^+ \mu^-)$, $S_{\psi K_s}$ ok
can be large

LHT after LHCb Data

**Our 2006
Predictions**
(Blanke et al.)

$\text{Br}(B_{s,d} \rightarrow \mu^+ \mu^-)$ within 40% from SM

$$|S_{\psi\phi}| \leq 0.25$$

$\{S_{\psi\phi} > 0.20\} \Rightarrow \left\{ \begin{array}{l} \text{No New Physics Effects} \\ \text{in } K^+ \rightarrow \pi^+ \nu \bar{\nu}, K_L \rightarrow \pi^0 \nu \bar{\nu} \end{array} \right\}$

**Concerning
B-Physics**

LHCb Data = Relief for LHT model *)

**Concerning
K-Physics**

**LHCb opened the road to large NP effects
in rare K-decays within LHT model** *)

*)

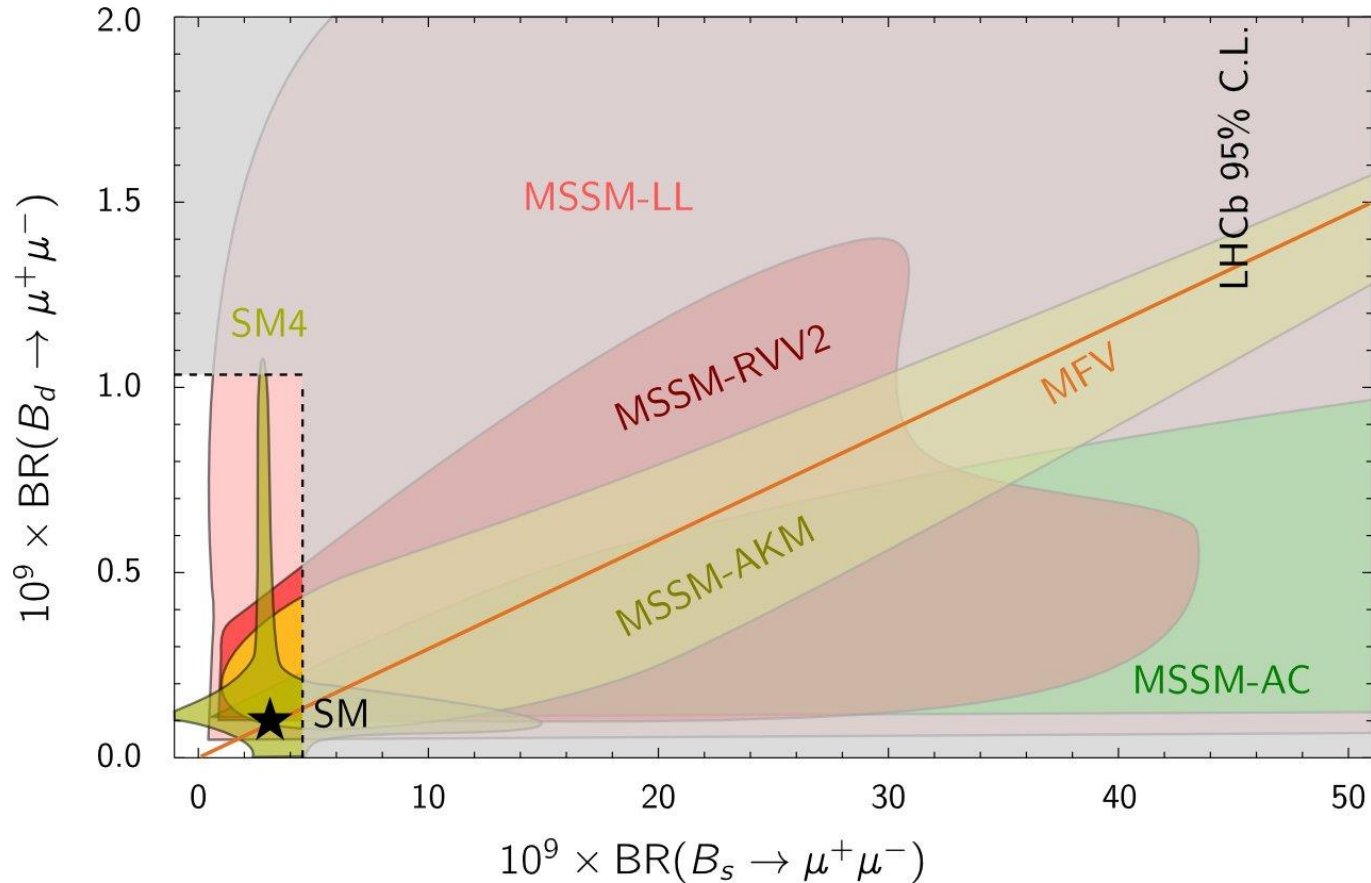
**The same impact of LHCb on Rare B
and K decays within RS_c model**

**Effects in
 $B_{s,d} \rightarrow \mu^+ \mu^-$
even smaller**

Supersymmetric Models Facing LHCb Data

ABGPS

Straub 1012.3893



Models with new left-handed currents favoured

Can $|V_{ub}|_{\text{excl}} \neq |V_{ub}|_{\text{incl}}$ be explained through right-handed currents?

Crivellin; Chen + Nam; Feger, Mannel et al.; AJB, Gemmler, Isidori

RHMFV

Works better with small $S_{\psi\phi}$

$$|V_{ub}|_{\text{excl}} = 3.12 (26) \cdot 10^{-3}$$

$$|V_{ub}|_{\text{inc}} = 4.27 (38) \cdot 10^{-3}$$

$$\varepsilon \approx \frac{v_L}{v_R}$$

$$|V_{ub}|_{\text{excl}} = |V_{ub}^L + a\varepsilon^2 V_{ub}^R|$$

$$|V_{ub}|_{\text{inc}} \approx |V_{ub}^L|$$

Generally: in principle yes

But a very detailed analysis of $SU(2)_L \otimes SU(2)_R \otimes U(1)_{B-L}$ with $g_L \neq g_R$; $V_L \neq V_R$ (mixing) including FCNC constraints + EWP constraints shows that in this concrete model the effect of RH currents too small !!

Blanke
AJB
Gemmler
Heidsieck
(1111.5014)

Comparison of Simplest Models

	$\Delta \varepsilon_K $	ΔM_d	ΔM_s	$\Delta S_{\psi K_s}$	$\Delta S_{\psi\phi}$	Favoured $ V_{ub} $
CMFV	+	+	+	0	0	exclusive
2HDM_{MFV}	0	\pm	\pm	-	+	inclusive
U(2)³	+	\pm	\pm	- 0 +	+ 0 -	inclusive exclusive

$$\left(\frac{\Delta M_s}{\Delta M_d}\right)_{\text{CMFV}} = \left(\frac{\Delta M_s}{\Delta M_d}\right)_{\text{MU(2)^3}} = \left(\frac{\Delta M_s}{\Delta M_d}\right)_{\text{SM}}$$

$$S_{\psi K_s} = \sin(2\beta + 2\varphi_{\text{new}})$$

$$S_{\psi\phi} = \sin(2|\beta_s| - 2\varphi_{\text{new}})$$

(the same relation for $B_{s,d} \rightarrow \mu^+ \mu^-$)

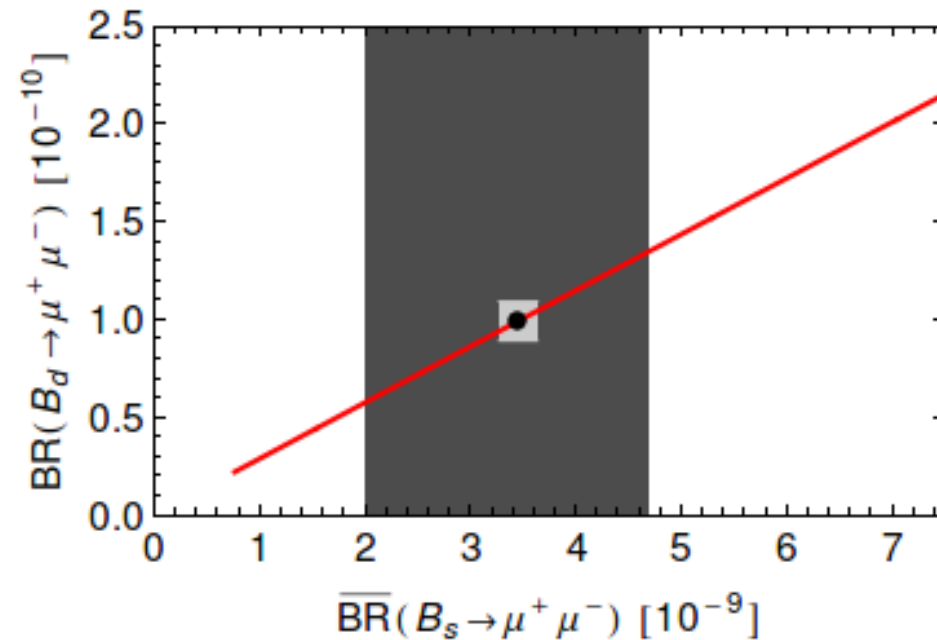
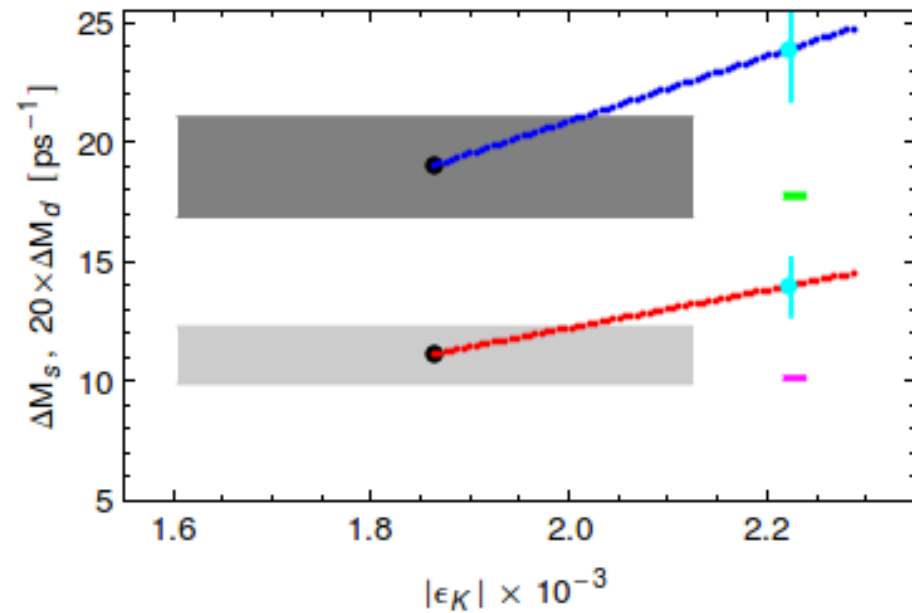
$$\beta = F(|V_{ub}|, \gamma)$$

(weak)

Constrained Minimal Flavour Violation

AJB + J. Girrbach (2012)

0007085
0310208
0604057



Tension within CMFV

EXP

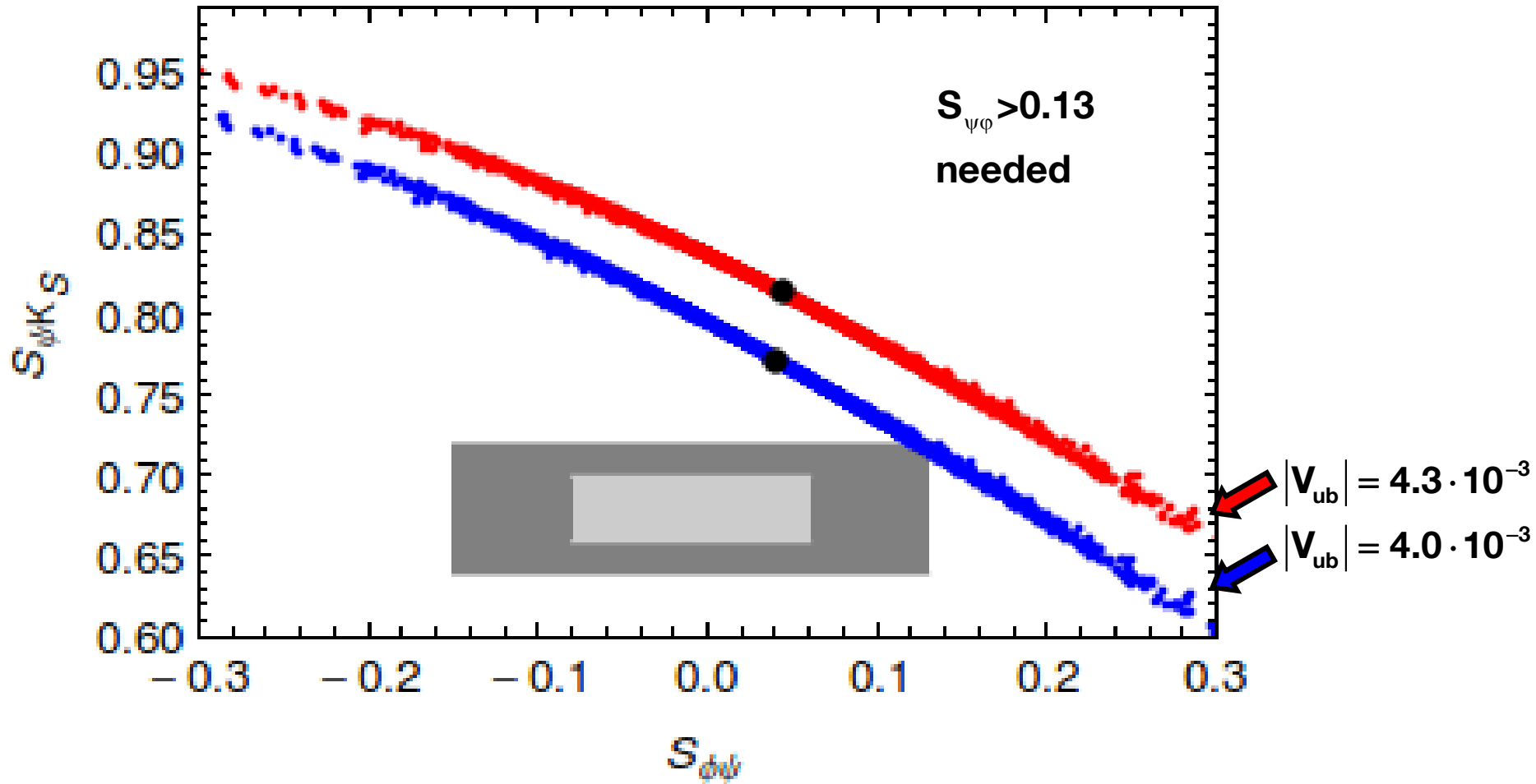
**Similar tension in
Gauged Flavour Models:
AJB, Merlo, Stamou (2011)**

EXP

$$\overline{Br}(B_s \rightarrow \mu^+ \mu^-) = \left(3.2^{+1.5}_{-1.2} \right) \cdot 10^{-9}$$

2HDM_{MFV} Facing LHCb Data

AJB, Girschbach, Nagai (2013)

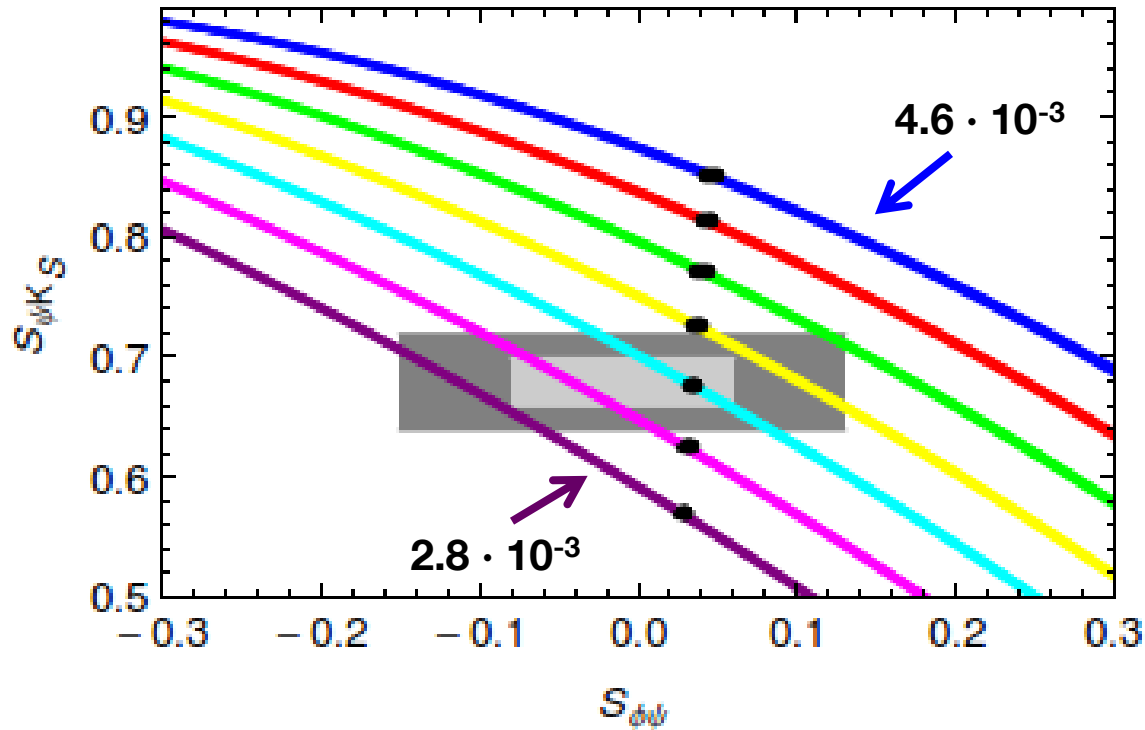


AJB, Carlucci, Gori, Isidori; 1005.5310

AJB, Isidori, Paradisi; 1007.5291

$S_{\psi K_S} - S_{\psi\phi} - |V_{ub}|$ Correlation in $U(2)^3$

Important test of $U(2)^3$ Models

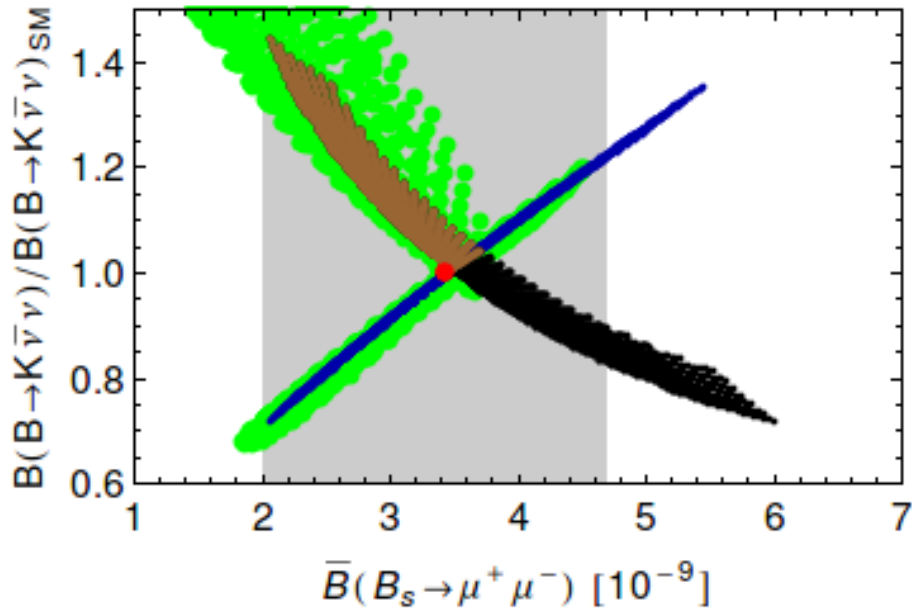


$\gamma = 68^\circ$

In the $U(2)^3$ Symmetric World we could determine $|V_{ub}|$ without significant hadronic uncertainties (QCD penguins)

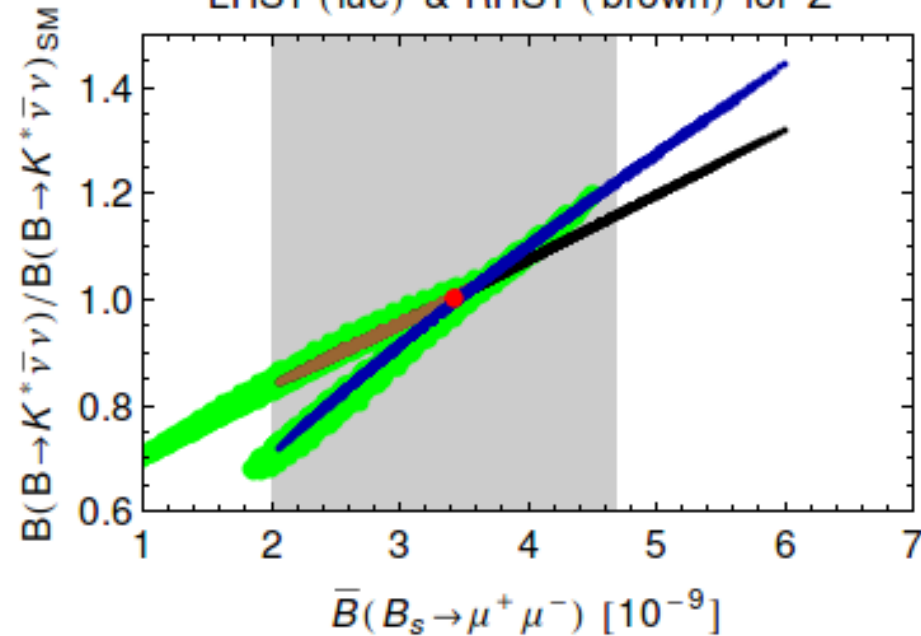
Distinguishing Left-Handed Currents from Right-Handed Currents

LHS1 (blue) & RHS1 (brown) for Z'



AJB, De Fazio, Girrbach
1211.1896

LHS1 (blue) & RHS1 (brown) for Z'



Altmannshofer et al.
0902.0160

■ : forbidden by
 $b \rightarrow sll$

■ : allowed by
 $b \rightarrow sll$

Z'-Couplings

$\bar{\Delta}_L^{sb}(\mathbf{Z}')$
complex

$\bar{\Delta}_A^{\mu\bar{\mu}}(\mathbf{Z}')$ $\bar{\Delta}_V^{\mu\bar{\mu}}(\mathbf{Z}')$
real

$\bar{\Delta}_L^{db}(\mathbf{Z}')$
complex

$\Delta M_s, S_{\psi\phi}$

×

$B_d \rightarrow K^* \mu^+ \mu^-$

×

×

×

$B_s \rightarrow \mu^+ \mu^-$

×

×

$B_d \rightarrow \mu^+ \mu^-$

×

×

$\Delta M_d, S_{\psi\phi}$

×

coupling

$$\bar{\Delta}_i = \frac{\Delta_i}{M_{Z'}}$$

reduced coupling

Z-Couplings $\Delta_{A,V}^{\mu\bar{\mu}}$ fixed

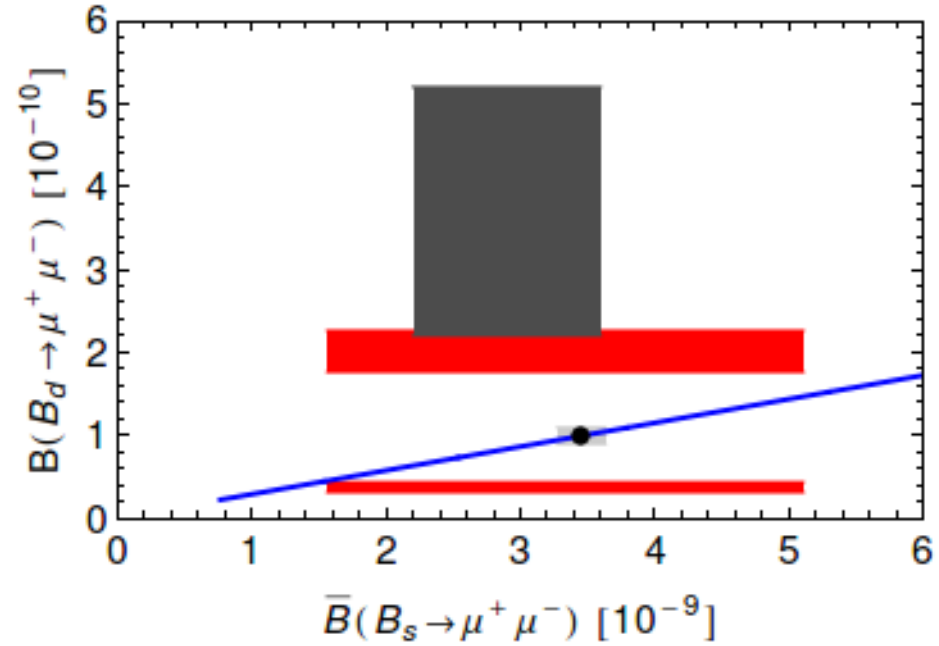
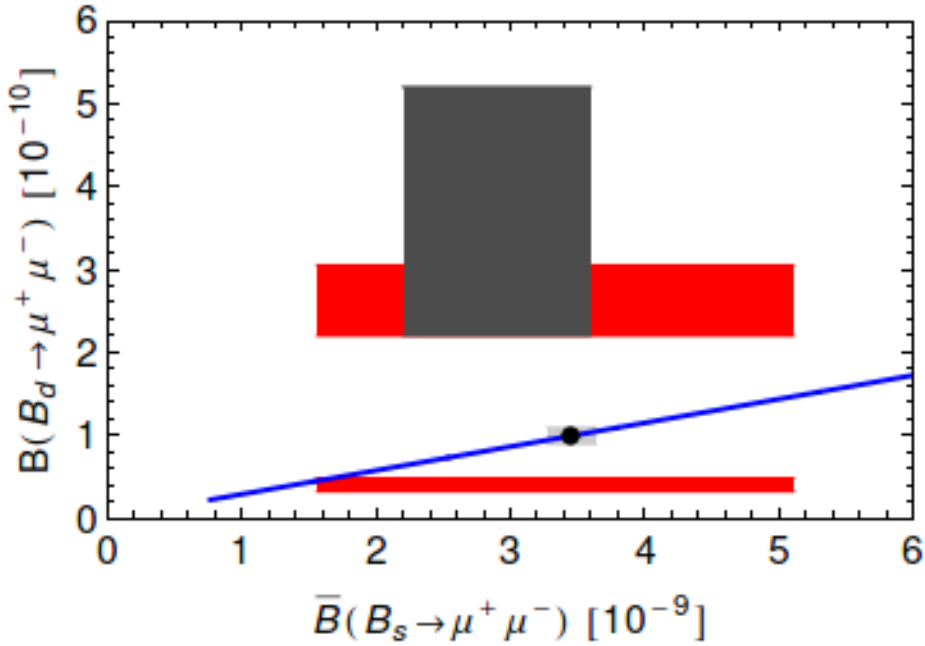
$$SU(2)_L : \Delta_L^{vv}(\mathbf{Z}') = \frac{\Delta_V^{\mu\bar{\mu}}(\mathbf{Z}') - \Delta_A^{\mu\bar{\mu}}(\mathbf{Z}')}{2}$$

$b \rightarrow sv\bar{\nu}$

Violation of CMFV (Z)

High V_{ub}

Low V_{ub}



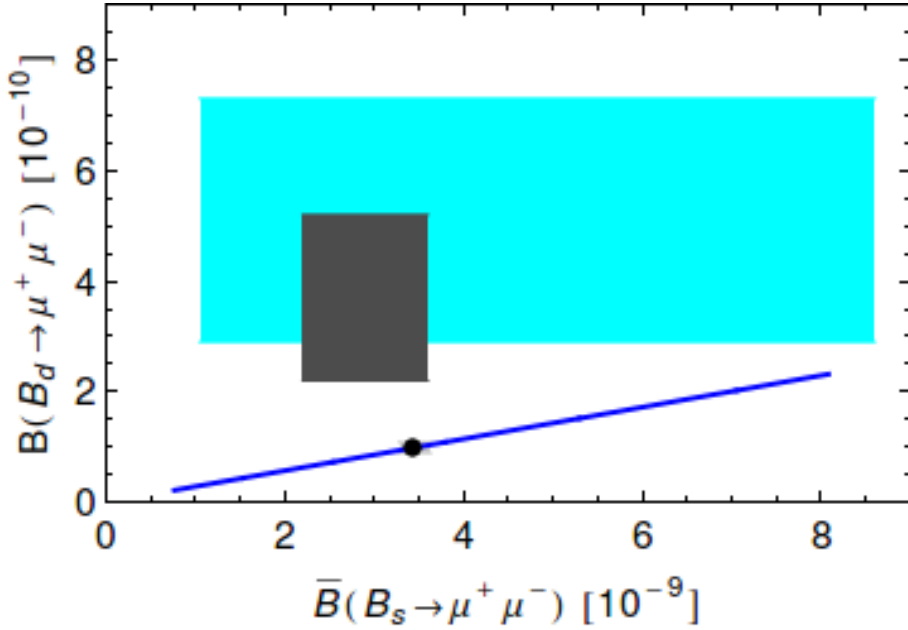
CMS + LHCb



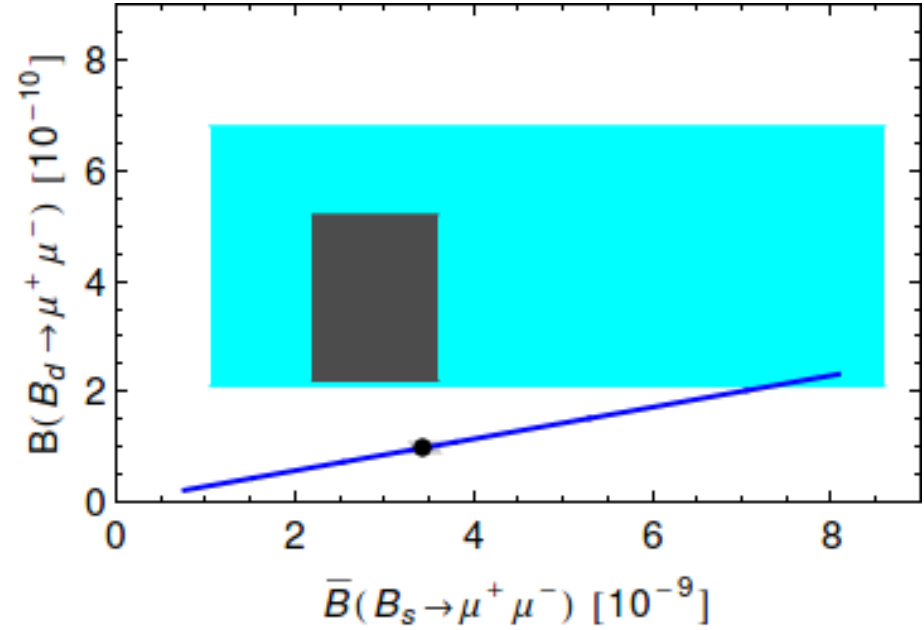
SM

Violation of CMFV (Z)

High V_{ub}



Low V_{ub}



CMS + LHCb



SM

Anomalies in $B_d \rightarrow K^* \mu^+ \mu^-$

(24 angular observables. Good agreement with SM but three deviations)

LHCb

SM

(Altmannshofer + Straub)

$$\langle \mathbf{F}_L \rangle_{[1.6]} = 0.66 \pm 0.07$$

$$0.77 \pm 0.04$$

$$\langle \mathbf{S}_5 \rangle_{[1.6]} = 0.10 \pm 0.10$$

$$-0.14 \pm 0.02$$

$$\langle \mathbf{S}_4 \rangle_{[14,16]} = -0.07 \pm 0.11$$

$$0.29 \pm 0.07$$

(Not understood in any model)

Extensive Analyses:

Descotes-Genon, Matias, Virto (1307.5683)

Altmannshofer + Straub (1308.1501)



DMV
AS

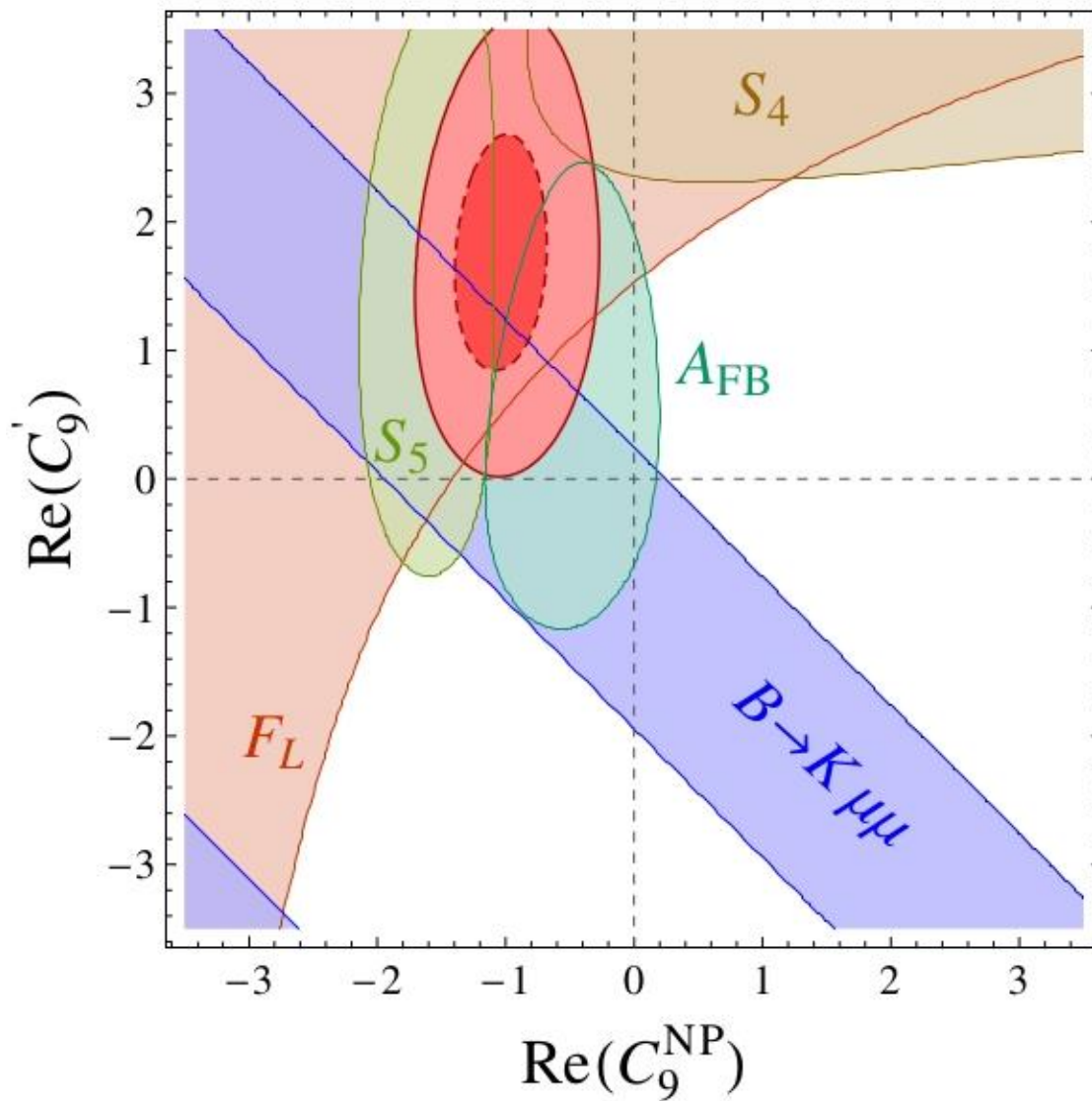
$$C_{7\gamma}^{\text{NP}} < 0, C_9^{\text{NP}} < 0$$

or

AS

$$C_9^{\text{NP}} < 0, C_9' \approx -C_9^{\text{NP}}$$

right-handed



**Altmannshofer
 Straub
 (1308.1501)**

Left-handed Z' and Z FCNC Couplings Facing $B_d \rightarrow K^* \mu^+ \mu^-$ Anomalies

(AJB + Girrbach, 1309.2466)

Z

fails because of small vector coupling to muons when $\Delta M_{s,d}$ constraints taken into account.

Z'

Suggested by Descotes-Genon, Matias, Virto (1307.5683)

Softens $\langle F_L \rangle$, $\langle S_5 \rangle$ anomalies

provided $C_9^{NP} \approx -1.5$ in a correlated manner

See also Altmannshofer Straub (1308.1501)

Note: In Z' models $C_{7\gamma}^{NP} = 0$ (1211.1896)



Optimal solutions to $\langle F_L \rangle$, $\langle S_5 \rangle$ (1309.2466)

Fails for $\langle S_4 \rangle$ must be SM-like

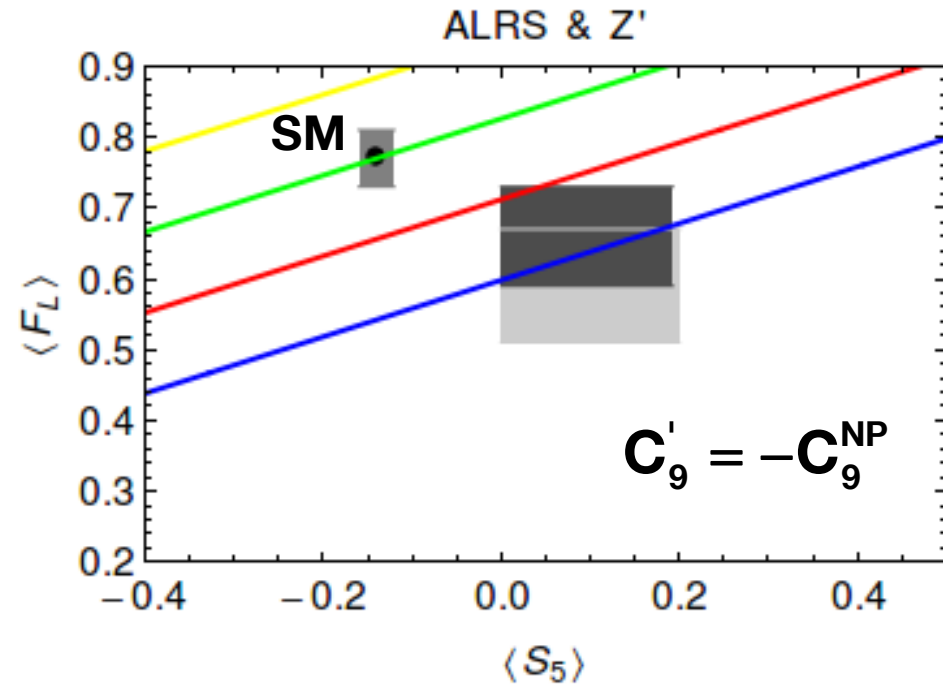
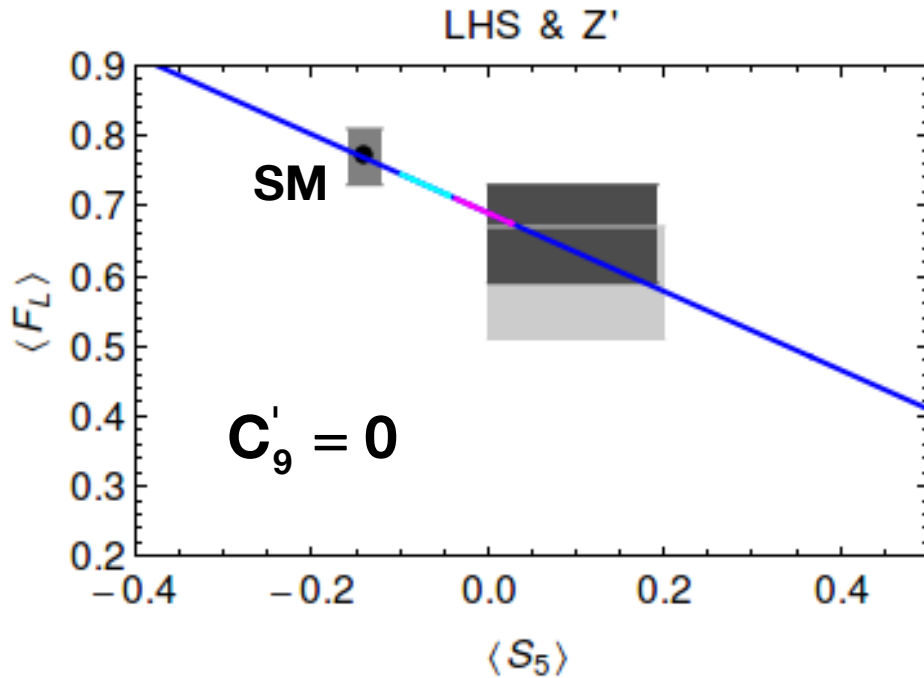
$$C_9^{NP} \neq 0, C_9' = 0 \quad (\text{LHS})$$

$$C_9^{NP} \neq 0, C_9' \approx -C_9^{NP} \quad (\text{ALRS})$$

New Correlations

(General Z')

(AJB + Girrbach, 1309.2466)



— $C_9^{NP} = -(0.8 \pm 0.3)$

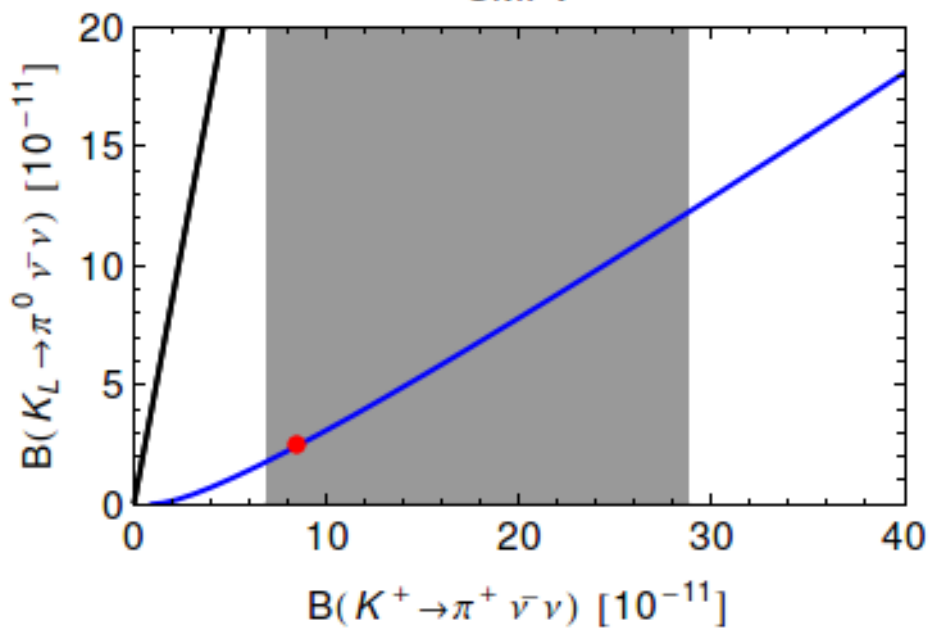
— $C_9^{NP} = -(1.6 \pm 0.3)$

— $C_9^{NP} = -1.0$

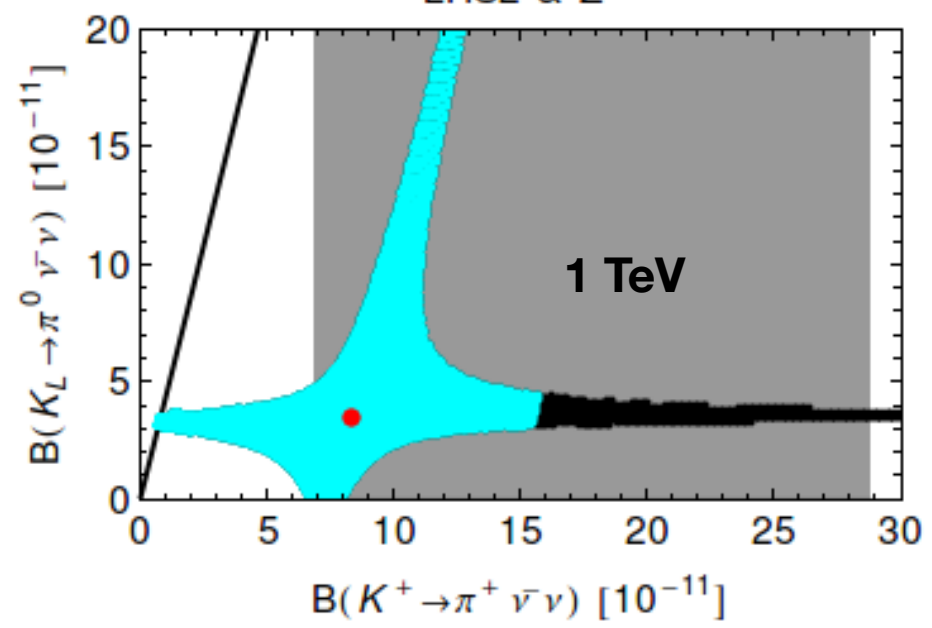
— $C_9^{NP} = -2.0$

— $C_9^{NP} = 0$

CMFV



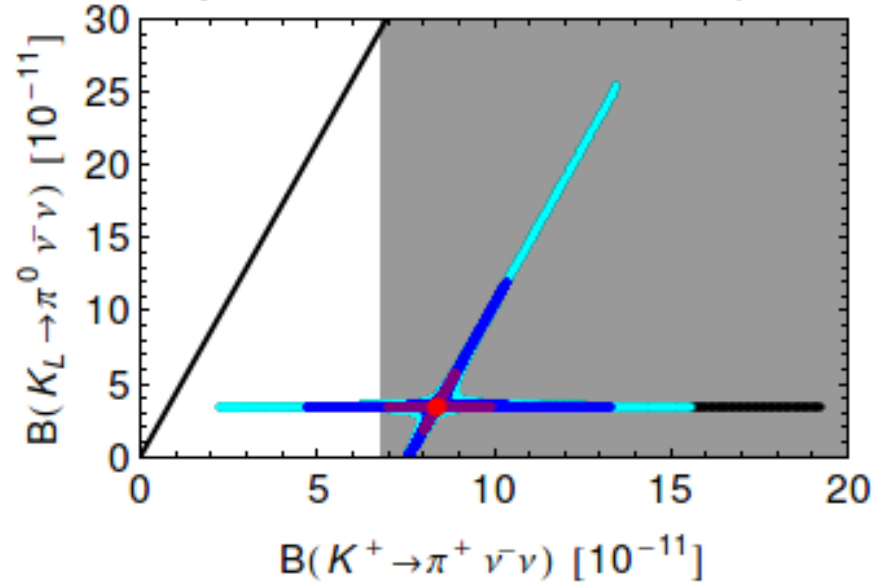
LHS2 & Z'



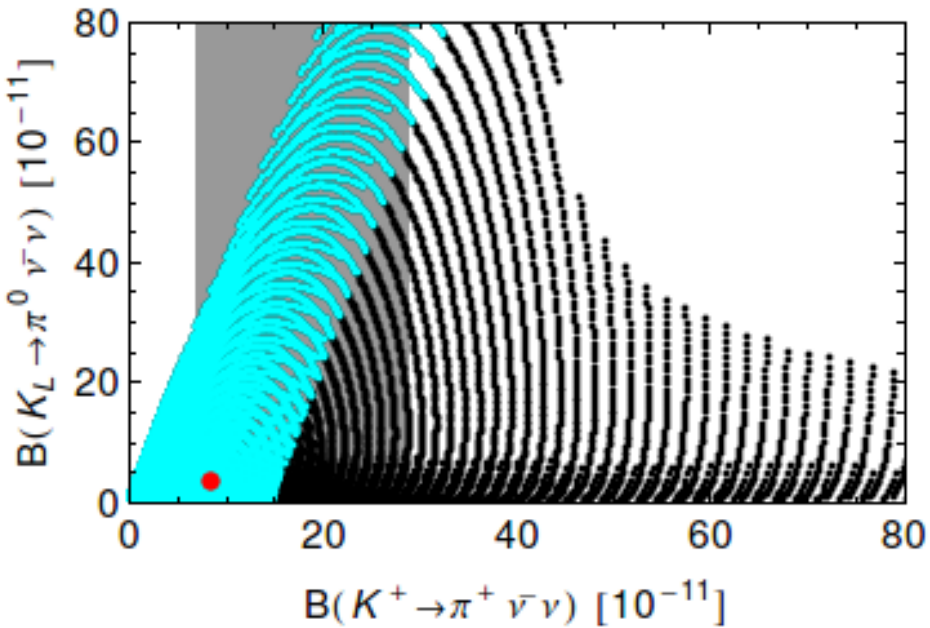
Sensitivity to $M_{Z'}$ beyond the LHC

■ : forbidden by $K_L \rightarrow \mu^+ \mu^-$

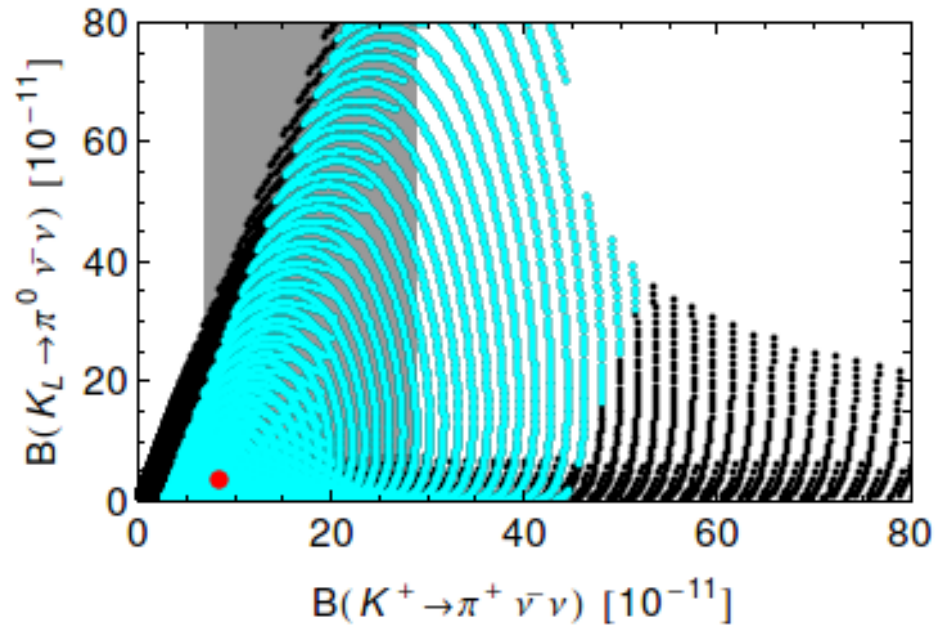
LHS2, Cyan: 5TeV, Blue: 10TeV, Purple: 30TeV



LHS2 & Z



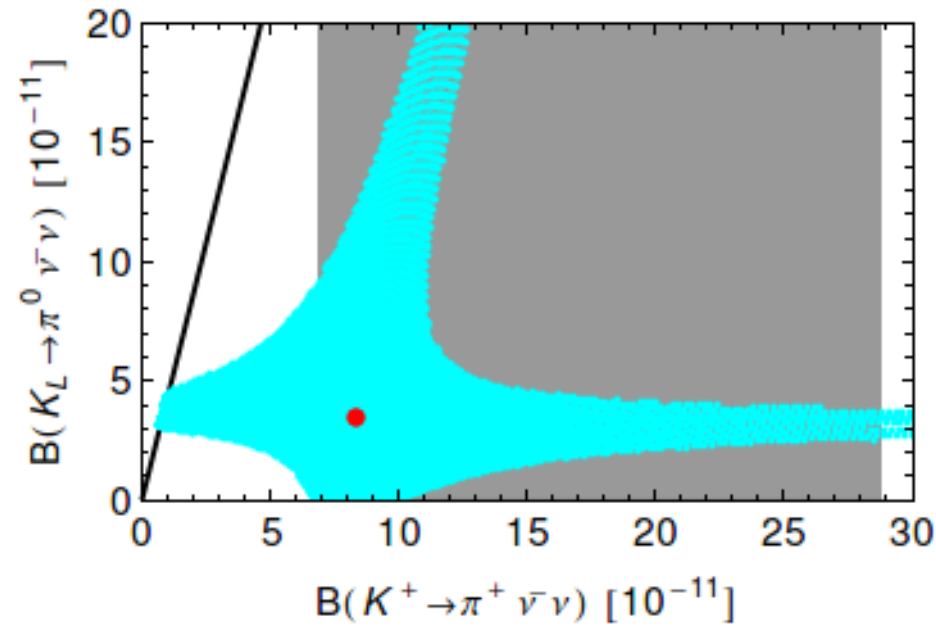
RHS2 & Z



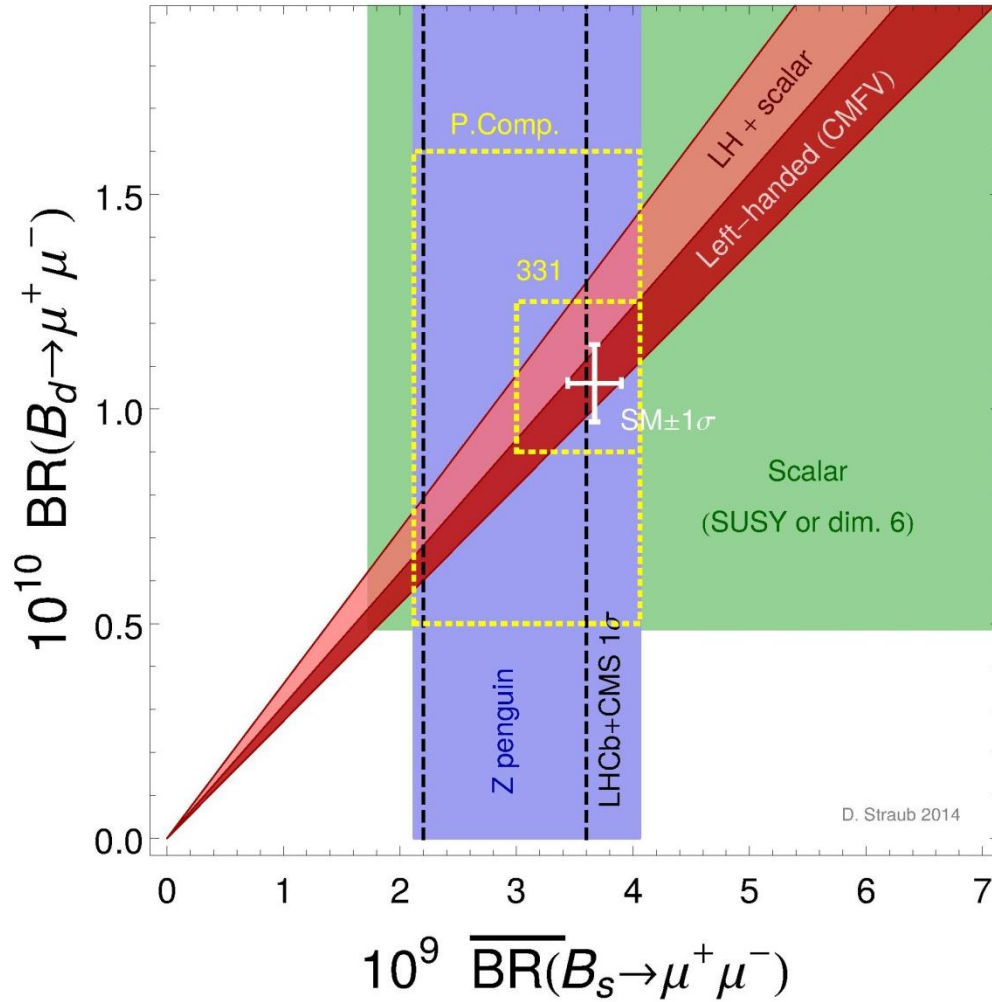
■ : forbidden by
 $K_L \rightarrow \mu^+ \mu^-$

LHS, RHS
 LRHS

LRS2 & Z



Straub's Plot 2014



Two Versions of Effective Theories

1.

$$L_{\text{eff}} = L_{\text{SM}} + \sum_i \frac{c_i}{\Lambda_i^2} Q_i$$

Wilson Coefficients

Local operators

which results from integrating out heavy fields (c_i, Λ_i depend on parameters of a given theory)
Very powerful framework in flavour physics, RG, etc.

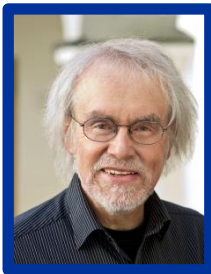
2.

The coefficients c_i, Λ_i are free parameters.
Completion unknown. Very limited framework in flavour physics except for cases when flavour symmetries and their breakdown are assumed:
MFV $(U(3))^3, U(2)^3, \dots$

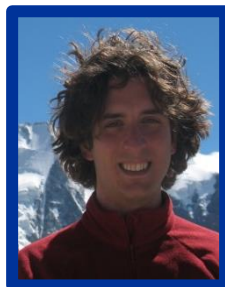
ET \equiv

1.

**Can we reach Zeptouniverse
 10^{-21}m ($\sim 200\text{ TeV}$)
by means of Quark Flavour Physics?**



AJB



D. Buttazzo



J. Girrbach-Noe



R. Kneijens

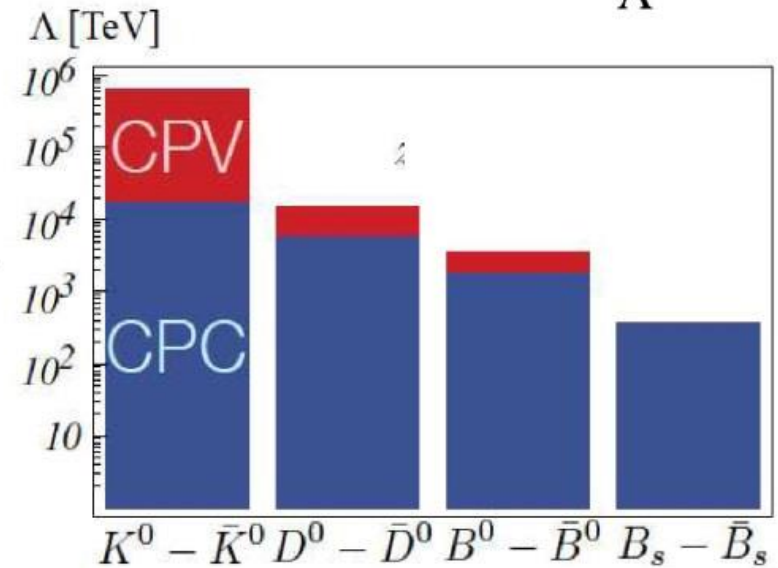
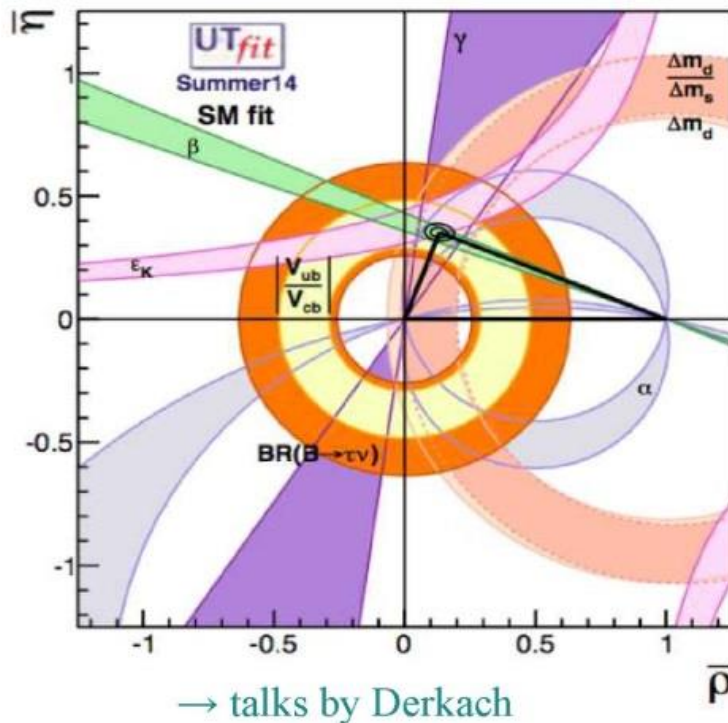
See also Charles et al.
(1309.2293)

- Are there other sources of flavor symmetry breaking (beside the SM Yukawa couplings)?

- What determines the observed pattern of quark & lepton mass matrices?

That's the question addressed by precision measurements (& searches) of flavor-changing processes of quarks & charged-leptons → So far everything seems to fit well with the SM → Strong limits on NP

$$\mathcal{L}_{\text{eff}} = \mathcal{L}_{\text{SM}} + \frac{1}{\Lambda^2} \mathcal{O}_{\Delta F=2}$$



Kamenik '13, G.I. '13, ...
[wide literature]

The sensitivity of $\Delta F=2$ processes to scales $\Lambda_{\text{NP}} > 1000 \text{ TeV}$ is impressive !!!

Yet

Three points to be made in this talk

1

New Physics at these scales cannot be measured in K, B_s , B_d , D rare decays (NP effects negligible)



2

We cannot learn much about the nature of this physics through $\Delta F=2$ processes and Effective Theory approach except when flavour symmetries $U(3)^3$ (MFV), $U(2)^3$ are involved.

3

We need badly rare decays to learn about physics beyond the LHC.

?

What are the maximal scales at which NP can be seen in rare K, B_s , B_d , D decays?

L and R Quark Couplings in Tree Level FCNCs

$\Delta F=2$

Cannot distinguish between L and R

(square)

$$\varepsilon_K, \Delta M_{s,d} \sim ag_L^2 + ag_R^2 + cg_L g_R$$

$|c| \gg |a|$
Hadronic matrix elements + RG

K: $c \sim 150 a$ **$B_{s,d}$:** $c \sim 7 a$

$\Delta F=1$

Can distinguish between L and R

A

Decays governed by V-quark couplings (γ_μ)

: $K^+ \rightarrow \pi^+ \nu \bar{\nu}, K_L \rightarrow \pi^0 \nu \bar{\nu}, B \rightarrow K \nu \bar{\nu}$

L \rightarrow R

No sign flip in NP contribution

B

Decays governed by A-quark couplings ($\gamma_\mu \gamma_5$)

: $K_L \rightarrow \mu^+ \mu^-, B_{s,d} \rightarrow \mu^+ \mu^-, B \rightarrow K^* \nu \bar{\nu}$

Sign flip in NP contribution

L \rightarrow R



Correlations A \leftrightarrow B
change to Anticorrelations A \leftrightarrow B

DNA - Charts

1306.3755

AJB + Girrbach



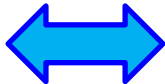
- **SM-like**



- **suppression relative to SM**



- **enhancement relative to SM**



correlation



anti-correlation

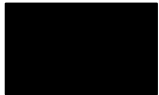
DNA - Charts

1306.3755

AJB + Girrbach



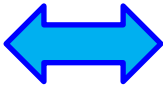
- SM-like



- suppression relative to SM



- enhancement relative to SM



correlation



anti-correlation



Searching for New Physics on the Way to Zeptouniverse

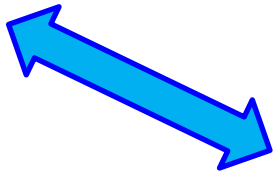
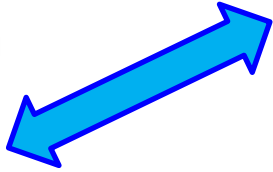


CMFV

ϵ_K

ΔM_s

ΔM_d



$S_{\psi\phi}$

$S_{\psi K_S}$

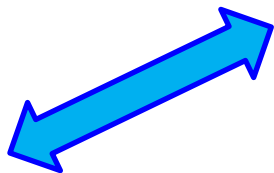
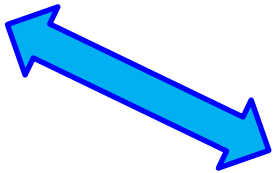
$B_s \rightarrow \mu\bar{\mu}$

$B_d \rightarrow \mu\bar{\mu}$



$K^+ \rightarrow \pi^+ \nu\bar{\nu}$

$K_L \rightarrow \pi^0 \nu\bar{\nu}$



$B \rightarrow K^{(*)} \nu\bar{\nu}$

$U(2)^3$

ϵ_K

ΔM_s



ΔM_d

$S_{\psi\phi}$



$S_{\psi K_s}$

$B_s \rightarrow \mu\bar{\mu}$



$B_d \rightarrow \mu\bar{\mu}$

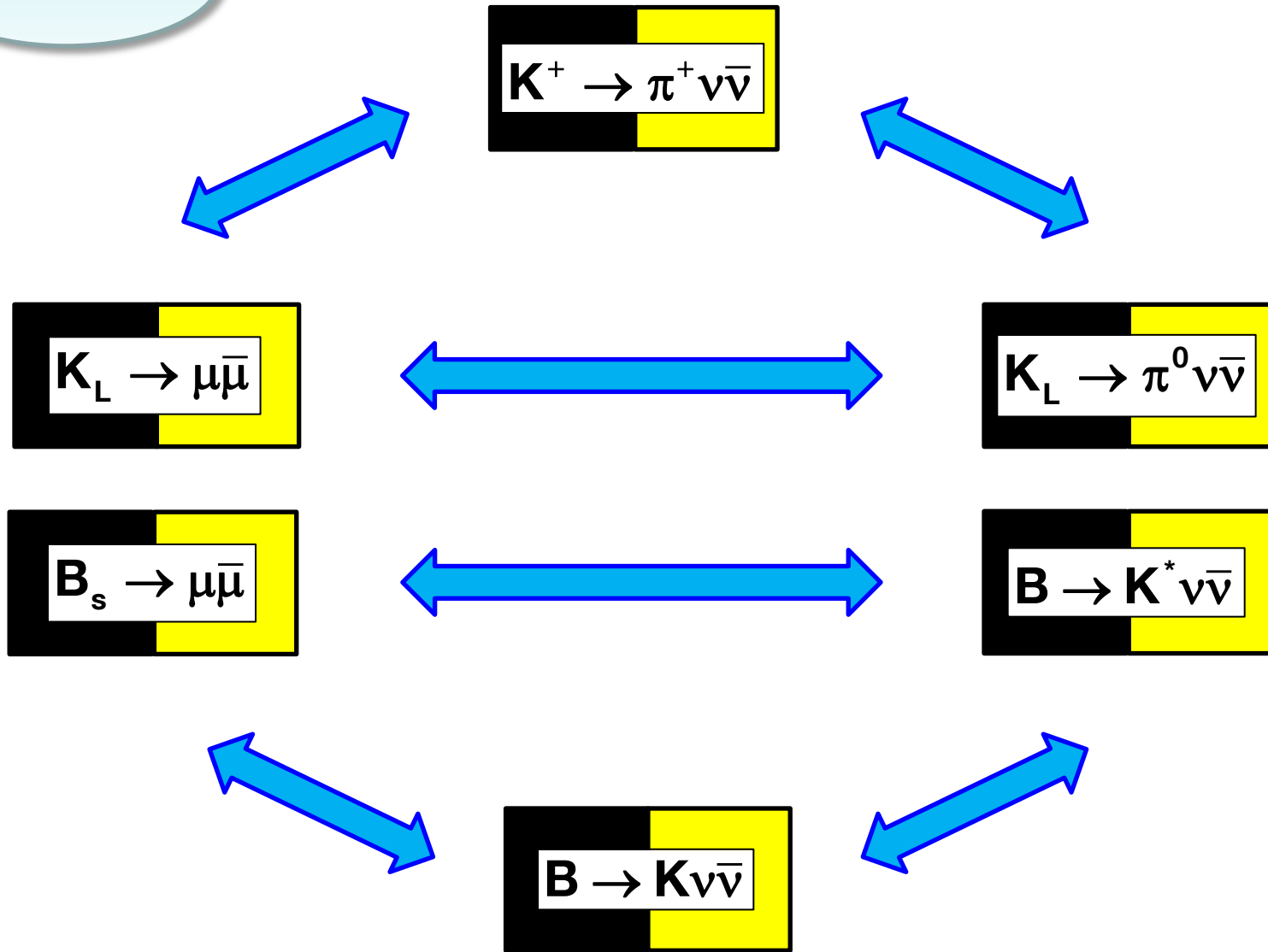
$K^+ \rightarrow \pi^+ \nu\bar{\nu}$



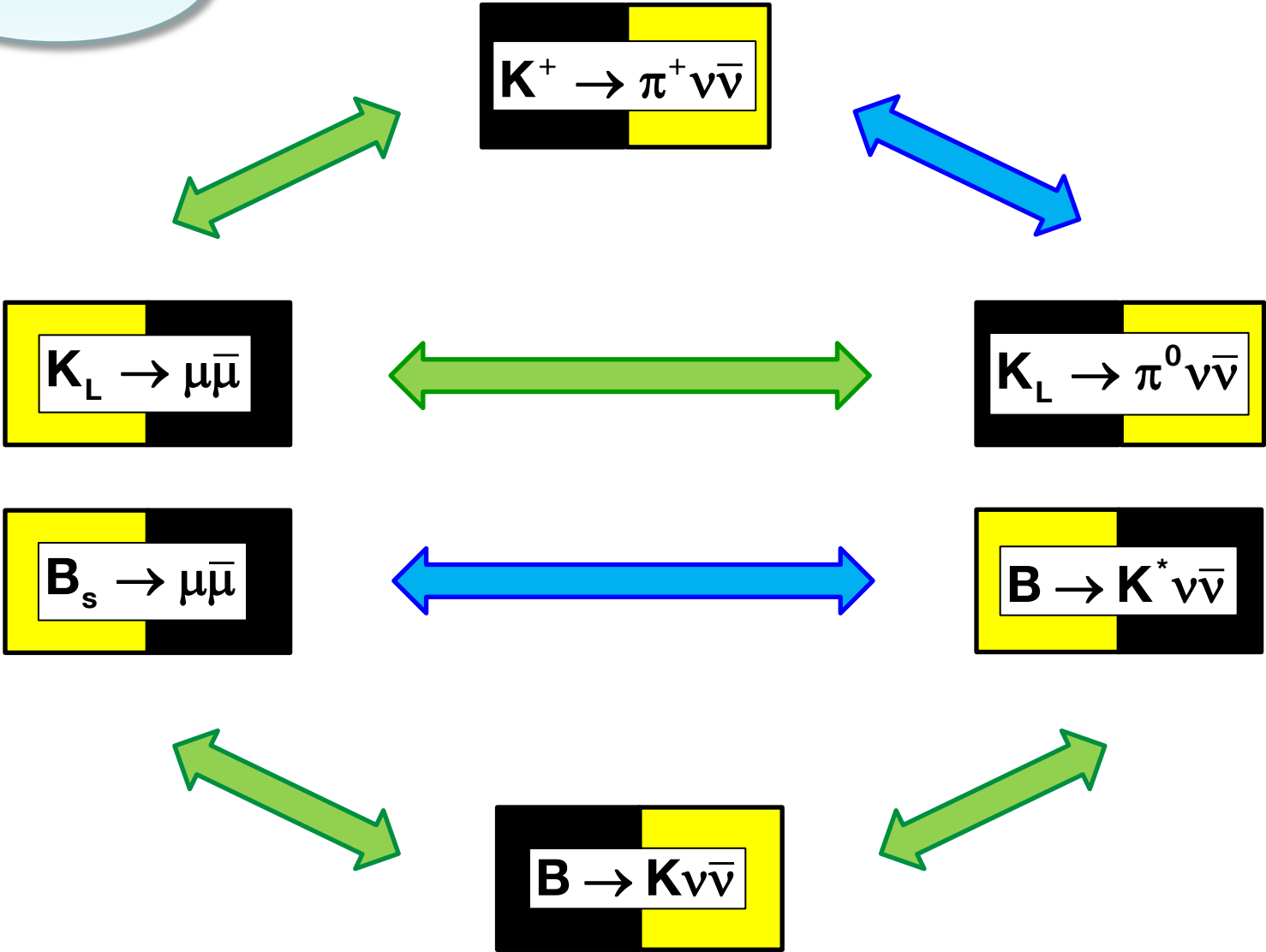
$K_L \rightarrow \pi^0 \nu\bar{\nu}$

$B \rightarrow K^{(*)} \nu\bar{\nu}$

Z'/Z LHS



Z'/Z RHS



Can we reach Zeptouniverse through Quark Flavour Physics ?

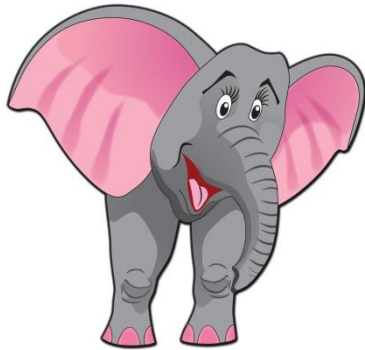
AJB, Buttazzo, Girrbach-Noe, Kneijens, 1407.0728

Due to limits of theory and experiment the answer depends on whether Zeptouniverse is “populated” by

Can we reach Zeptouniverse through Quark Flavour Physics ?

AJB, Buttazzo, Girrbach-Noe, Kneijens, 1407.0728

Due to limits of theory and experiment the answer depends on whether Zeptouniverse is “populated” by



In QFT :

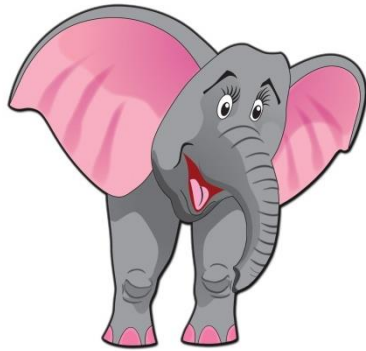
**Large
couplings**

**(still consistent
with perturbativity)**

Can we reach Zeptouniverse through Quark Flavour Physics ?

AJB, Buttazzo, Girrbach-Noe, Kneijens, 1407.0728

Due to limits of theory and experiment the answer depends on whether Zeptouniverse is “populated” by



In QFT :

**Large
couplings**

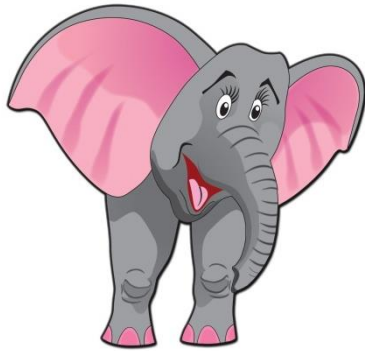
**Moderate
couplings**

**(still consistent
with perturbativity)**

Can we reach Zeptouniverse through Quark Flavour Physics ?

AJB, Buttazzo, Girrbach-Noe, Kneijens, 1407.0728

Due to limits of theory and experiment the answer depends on whether Zeptouniverse is “populated” by



In QFT :

**Large
couplings**

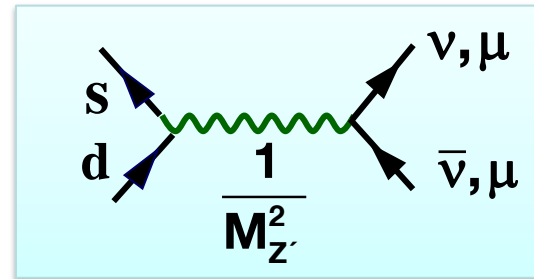
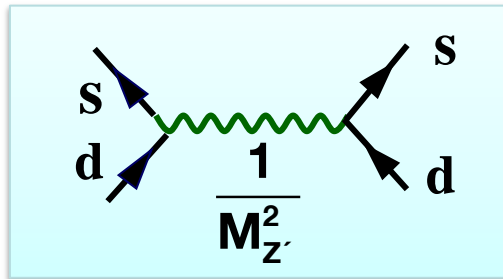
**Moderate
couplings**

**Small
couplings**

**(still consistent
with perturbativity)**

Answer within Z' -Models

(Stringent correlations between $\Delta F=2$ and $\Delta F=1$)



Similar
for B_s, B_d, D, \dots

For fixed lepton couplings, after $\Delta F=2$ constraints, NP effects in rare decays decrease as $1/M_{Z'}$.



Strategy:

Assume largest g_{ij} and $g_{\nu\nu}, g_{\mu\mu}$ couplings subject to $\Delta F=2$ constraints on g_{sd}, g_{sb}, g_{db}

$g_{ij} \approx 3$ still allowed by perturbativity
but often not by $\Delta F=2$ constraints.

NP effects should still be sufficiently large
to be able to see correlations.

Main Messages from this Study

(Maximal Resolution of Short Distance Scales)

1

If only g_L or g_R flavour changing Z' couplings to quarks present and $\Delta F=2$ constraints taken into account:

$K \rightarrow \pi \nu \bar{\nu} \sim 200 \text{ TeV}$

B_d physics: $\sim 15 \text{ TeV}$

B_s physics: $\sim 15 \text{ TeV}$

Maximal scales that can be explored

2

If $g_L = \pm g_R$ the scales are lower:
LR operator in $\Delta F=2$ enhanced through
RG + chiral enhancement in $\Delta M_K, \epsilon_K$



Smaller couplings



Lower scales at which NP dynamics can be tested

3

In order to probe scales above 50 TeV even with B_s, B_d physics we need either left-handed or right-handed elephants:

but

$g_L \neq g_R$ $g_R \neq 0$ $g_L \neq 0$
 $g_L \gg g_R$ or $g_R \gg g_L$

Cannot be distinguished through $\Delta F=2$ observables

Allows us to obtain significant NP effects in rare K, $B_{s,d}$ decays while satisfying $\Delta F=2$ constraints



(Help from LR operators with some tuning)

Important:

But:

Can be distinguished through correlations in rare K and B decays

(See DNA Charts)

Can we reach Zeptouniverse through Quark Flavour Physics ?

(Z)

AJB, Buttazzo, Girrbach-Noe, Kneegjens, 1407.0728

If only left-handed
or only right-handed
couplings present in NP

:

Only with K rare Decays
 $B_s \sim 15 \text{ TeV}$, $B_d \sim 15 \text{ TeV}$

If both LH and RH
present but
 $g_L^{ij} \ll g_R^{ij}$ or $g_L^{ij} \gg g_R^{ij}$

:

$K \rightarrow \pi \nu \bar{\nu}$: $\Lambda_{\text{NP}}^{\text{max}} \simeq 2000 \text{ TeV}$
 B_d : $\Lambda_{\text{NP}}^{\text{max}} \simeq 160 \text{ TeV}$
 B_s : $\Lambda_{\text{NP}}^{\text{max}} \simeq 160 \text{ TeV}$

Can we reach Zeptouniverse through Quark Flavour Physics ?

(Z)

AJB, Buttazzo, Girrbach-Noe, Kneijens, 1407.0728

If only left-handed
or only right-handed
couplings present in NP

:

Only with K rare Decays
 $B_s \sim 15 \text{ TeV}$, $B_d \sim 15 \text{ TeV}$

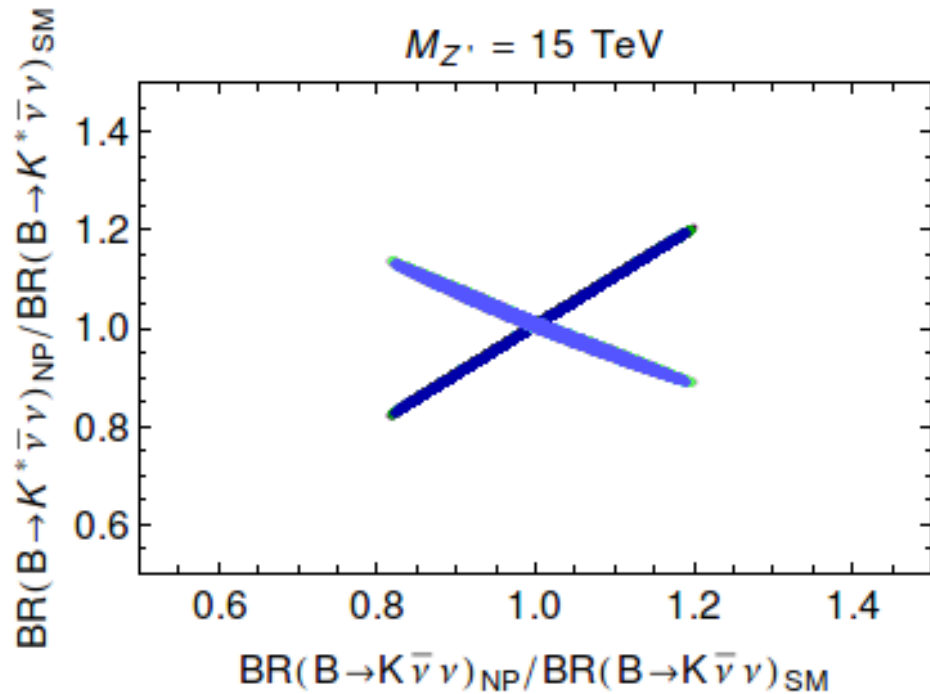
If both LH and RH
present but
 $g_L^{ij} \ll g_R^{ij}$ or $g_L^{ij} \gg g_R^{ij}$

:

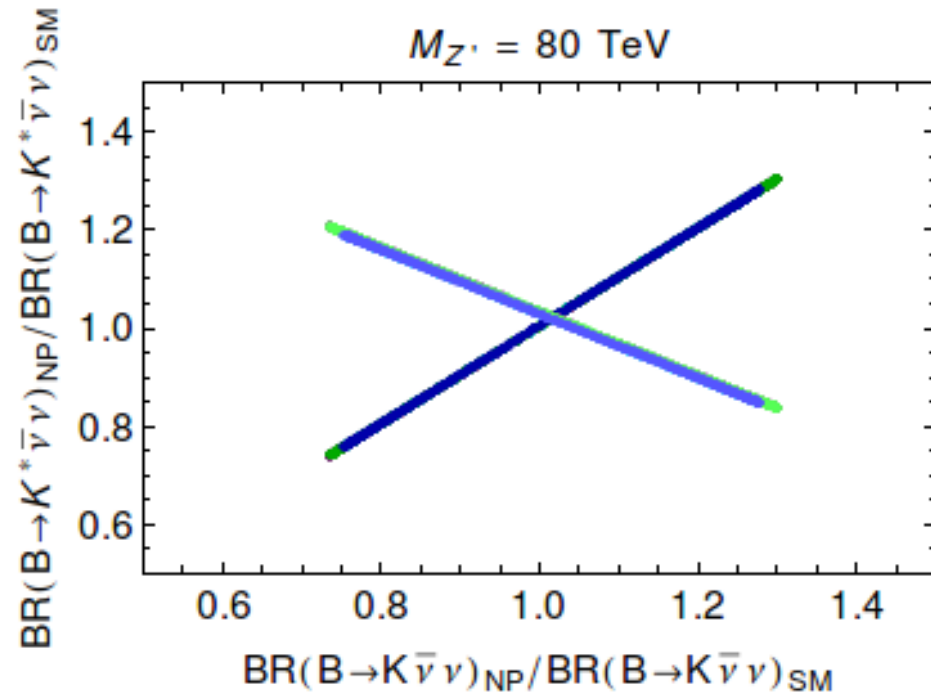
$K \rightarrow \pi \nu \bar{\nu}$: $\Lambda_{\text{NP}}^{\text{max}} \simeq 2000 \text{ TeV}$
 B_d : $\Lambda_{\text{NP}}^{\text{max}} \simeq 160 \text{ TeV}$
 B_s : $\Lambda_{\text{NP}}^{\text{max}} \simeq 160 \text{ TeV}$

Yes we can !!

Heavy Z' at Work

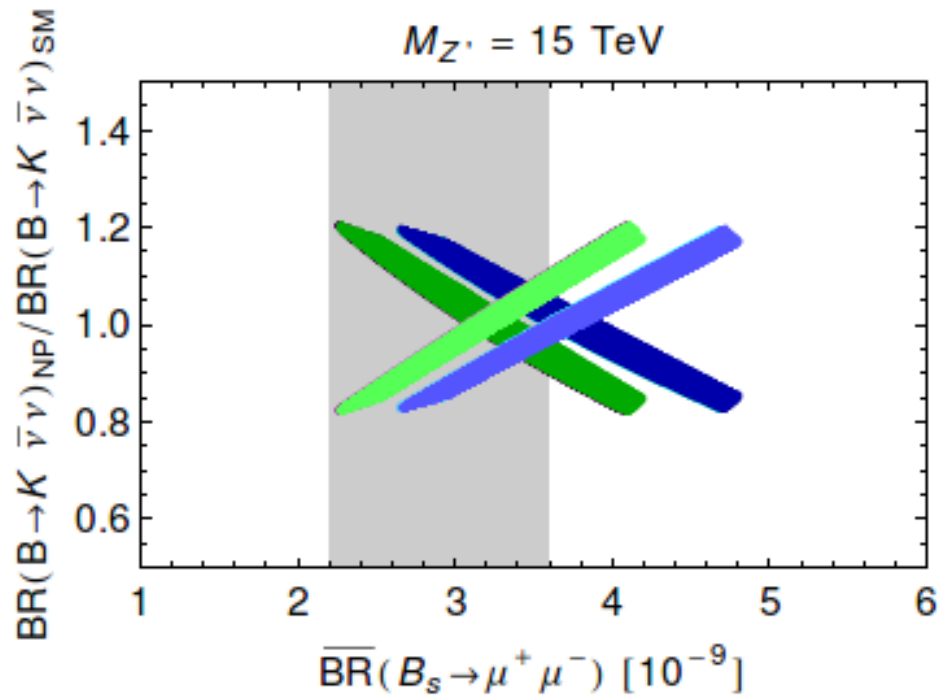


$\Delta F=2$ constraint

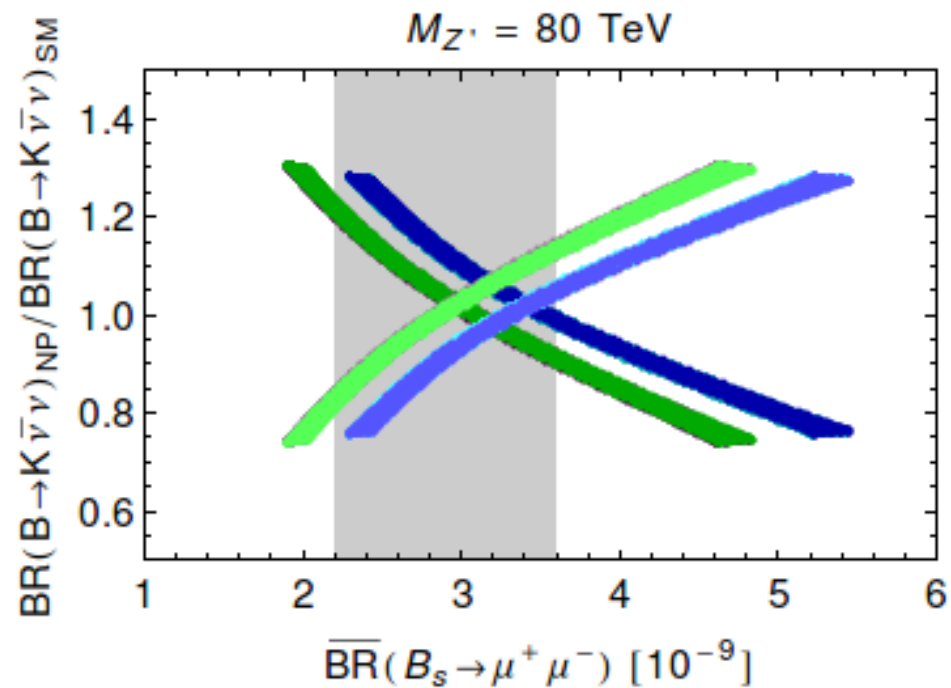


No $\Delta F=2$ constraint

Heavy Z' at Work



$\Delta F=2$ constraint



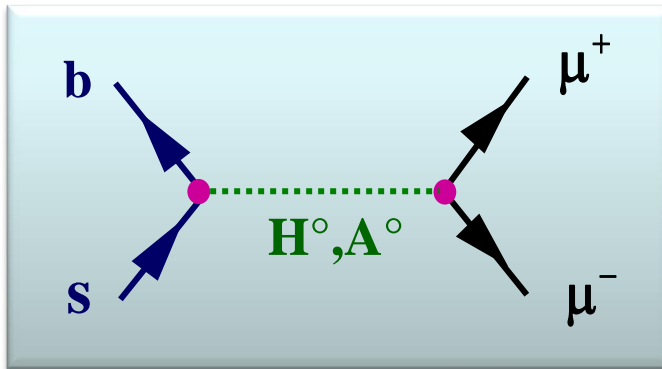
No $\Delta F=2$ constraint

Can we reach Zeptouniverse through S and P

AJB, Buttazzo, Girrbach-Noe, Kneijens, 1407.0728

Yes :

$$B_{s,d} \rightarrow \mu^+ \mu^-$$



S : ≈ 350 TeV

P : ≈ 700 TeV

Pseudoscalars more powerful than scalars because of the interference with SM contribution

Similar to $K \rightarrow \pi \nu \bar{\nu}$ (Z): No tuning necessary to reach Zeptouniverse

S= H^0

P= A^0