

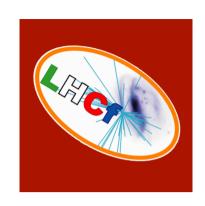
LHCf experiment: forward physics at LHC for cosmic rays study

Marina Del Prete (University and INFN of Florence) on behalf of the LHCf collaboration

International Conference on New Frontiers in Physics 23 August 2015 – 30 August 2015 Kolymbari (Crete)

OUTLINE

- Physics motivations of LHCf
- Experimental setup
- Latest analysis:
 - neutron energy spectra in 7 TeV p-p collisions
 - $\pi^0 P_T$ spectra in 5.02 TeV p-Pb collisions
 - Feynman Scaling
 - LPM Effect
- Run 13 TeV 2015
- Conclusion and Future plan



High Energy Cosmic Rays

High energies -> low flux and indirect measurements

Extensive air shower observation:

- Longitudinal distribution
- Lateral distribution
- Arrival direction



Astrophysical parameters:

- Spectrum
- Composition
- Source distribution

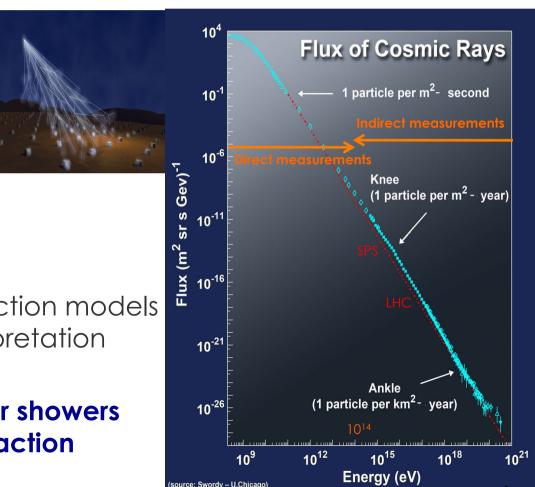


Uncertainty of hadron interaction models -> uncertainty of results interpretation



Monte Carlo simulations of air showers with accurate hadronic interaction models are needed





The role of the accelerators experiments

particles

Disk of particles

Muon detectors

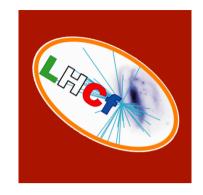


Forward energy spectrum

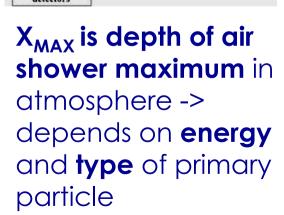
- if softer shallow development
- If harder deep penetrating

Inelasticity $k = 1 - E_{lead.} / E_{beam}$

- If large k (π⁰s take more energy)
 -> rapid development
- If small k (baryons take more energy) -> deep penetrating



by LHCf studing neutrons photons



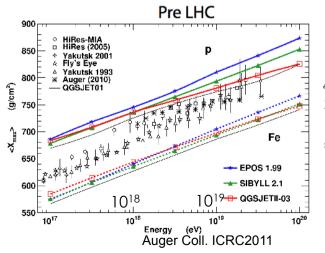
interaction

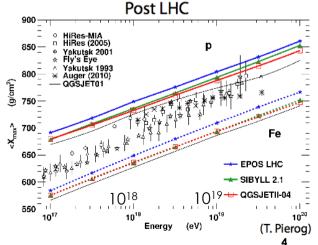
Shower core

Cherenkov light detectors

Charged particle

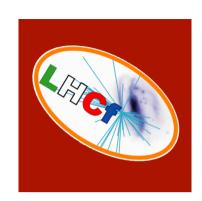
Nuclear effects by LHCf: study of p-Pb collisions

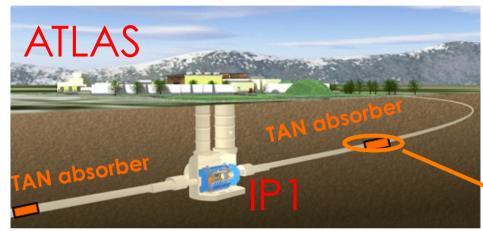




LHCf: experimental setup

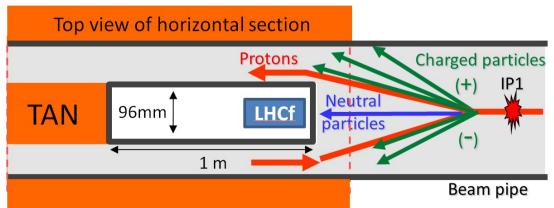
LHCf has the aim to understand the hadronic interactions in the collisions between high-energy CR and earth atmosphere





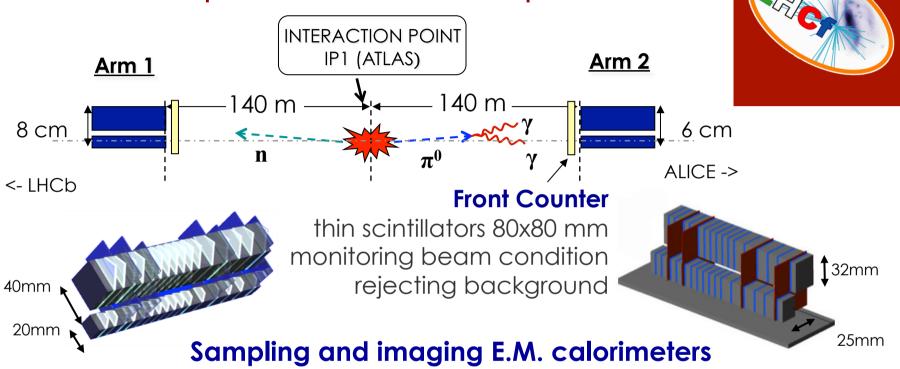
LHCf detectors are located inside the TAN absorbers for neutral particles

Charged particles are deviated by D1 dipole magnets



The LHCf detectors measure energy and impact point of γ and n

LHCf: experimental setup



Absorber: 16 W layers (44 r.l., $1.55\lambda_1$ in total)

Energy measurement: plastic scintillator tiles ->

Resolution (>100 GeV) < 5% for γ and ~ (35 ÷ 40)% for n

4 tracking layers: XY-SciFi(Arm1) -> Position resolution < 200μm

XY-Silicon μ -strip(Arm2) -> Position resolution ~ 40 μ m

Detectors has two independent calorimeter towers \rightarrow reconstruction of the mass of $\gamma\gamma$ system

LHCf collaboration

- O. Adriania, E. Bertia, L. Bonechib, M. Bongia, G. Castellinic,
- R. D'Alessandroa, M. Del Pretea, M. Grandib,
- M. Haguenauere, Y. Itowf,g, K. Kasaharah, K. Kawadeg,
- Y. Makino⁹, K. Masuda⁹, Y. Matsubara⁹, E. Matsubayashi⁹,
- H. Menjoi, G. Mitsukag, Y. Murakig, P. Papinib, A.-L. Perroti,
- S. Ricciarini^{c,b}, T. Sako^g, N. Sakurai^g, Y. Shimitsu^h, Y. Sugiura^g,
- T. Suzuki^h, T. Tamura^k, A. Tiberio^{a,b}, S. Torii^h, A. Tricomi^{l,m}, W. C. Turnerⁿ, K. Yoshida^o, Q. Zho
 - a) University of Florence, Italy
 - b) INFN Section of Florence, Italy
 - c) IFAC-CNR, Florence, Italy
 - d) Ecole Polytechnique, Palaiseau, France
 - e) Kobayashi Maskawa Institute for the Origin of Particles and the Universe, Nagoya University, Nagoya, Japan
 - f) Solar-Terrestrial Environment Laboratory, Nagoya University, Japan
 - g) RISE, Waseda University, Japan
 - h) Graduate School of Science, Nagoya University, Japan
 - i) CERN, Switzerland
 - j) Kanagawa University, Japan
 - k) University of Catania, Italy
 - I) INFN Section of Catania, Italy
 - m) LBNL, Berkeley, California, USA
 - n) Shibaura Institute of Technology, Japan





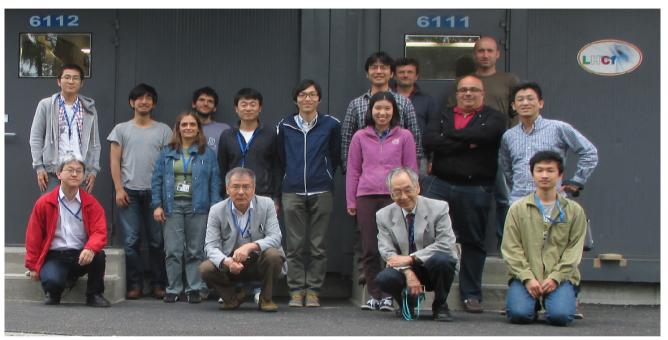








LHCf collaboration at CERN





 In front of LHCf control room at IP1 in June 2015

 In LHCf control room at IP1 in June 2015 ... meeting and shift for special Run p-p 13 TeV

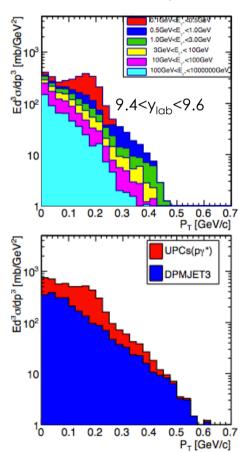




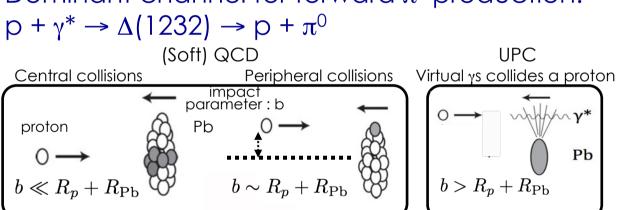
π^0 analysis in p-Pb collisions at 5.02 TeV (1.3x10¹⁶ eV proton hitting the atmosphere)



- Only Arm2 installed (better position resolution than Arm1)
- Data both at p-side (20Jan-1Feb) and Pb-side (1fill, 4Feb), thanks to the swap of the beams



Dominant channel for forward π^0 production:

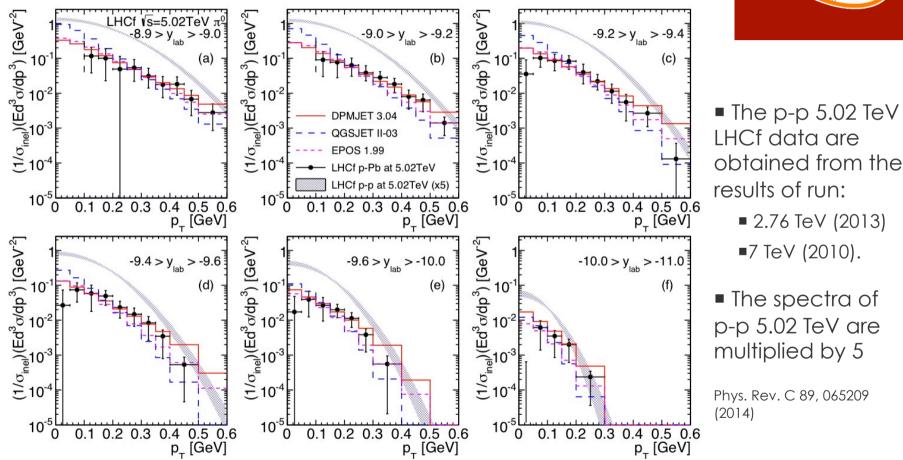


UPC contribution to P_T spectra is estimated from MC simulations (using Weizsacker Williams approximation for γ^* spectrum and SOPHIA model for p- γ^* collisions)

About half of the observed π^0 s originate from UPC Need to subtract UPC component

π^0 analysis in p-Pb collisions at 5.02 TeV and p-p collisions at 2.76 and 7 TeV

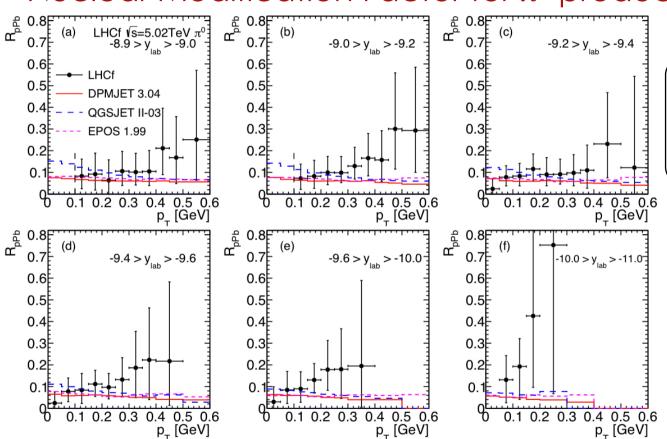
(4.1x10¹⁵ and 2.6x10¹⁶ eV proton hitting the atmosphere)

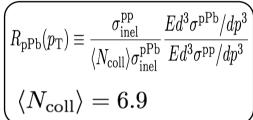


- p-Pb Data (filled circles) are in good agreement with DPMJET and EPOS
- p-Pb spectra are harder than p-p spectra.

π^{0} analysis in p-Pb collisions at 5.02 TeV

Nuclear Modification Factor for π^0 production





Large y_{lab} -> large experimental uncertainty.

Better results with ATLAS combined analysis: ATLAS does not detect the diffractive events

Data and MC shows strong suppression

0.1 0.2 0.3 0.4 0.5 0.6

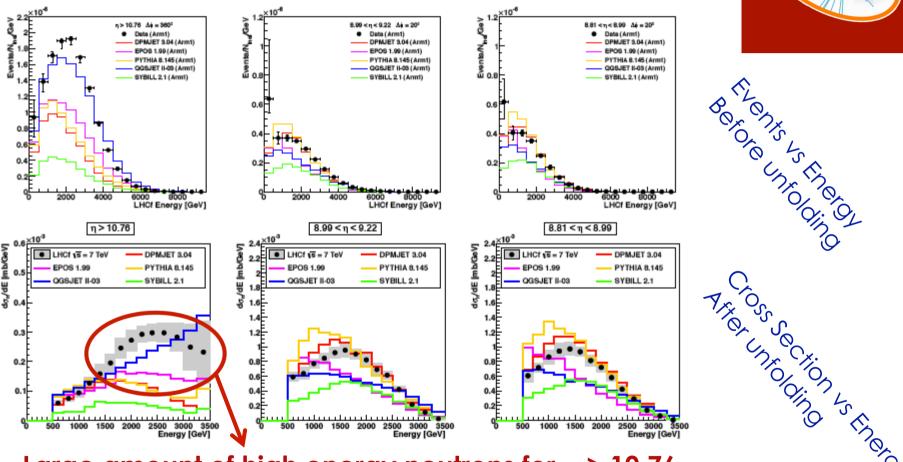
p_ [GeV]

0.1 0.2 0.3 0.4 0.5 0.6

NMF grows with PT as expected (p-Pb spectra are harder than p-p spectra) 11

Neutrons analysis in p-p collisions at 7 TeV

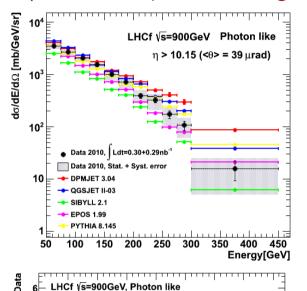
(2.6x10¹⁶ eV proton hitting the atmosphere)



Large amount of high energy neutrons for $\eta > 10.76$ (only predicted by QGSJET) -> small inelasticity in the very forward region

LHCf: p-p run 0.9 and 7 TeV Photons spectra in p-p at **900 GeV**

(4.3x10¹⁴ eV proton hitting the atmosphere)



 $\eta > 10.15 (<\theta> = 39 \mu rad)$

◆ DPMJET 3.04

QGSJET II-03
SIBYLL 2.1

EPOS 1.99

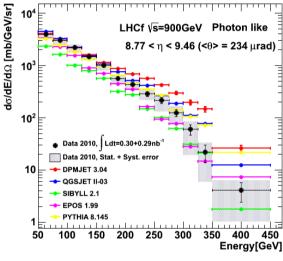
PYTHIA 8.145

150

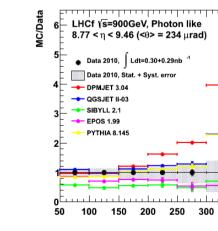
200

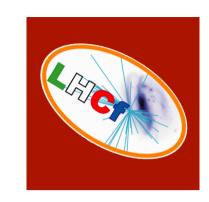
250

300



350





- DPMJET SYBILL show reasonable agreement of shape
- None of the models reproduces the data within the error bars

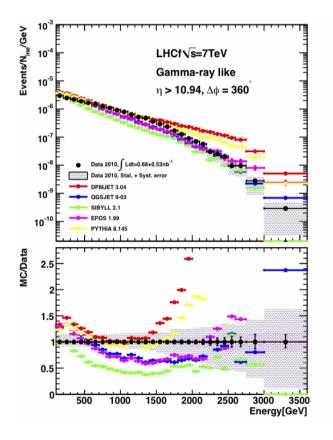
PLB 715 298 (2012)

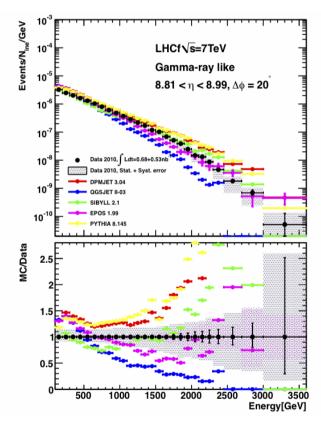
No strong evidence of η -dependence

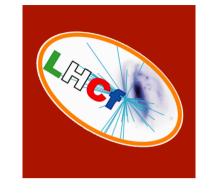
LHCf: p-p run 0.9 and 7 TeV

Photons spectra in p-p at **7 TeV**

(2.6x10¹⁶ eV proton hitting the atmosphere)







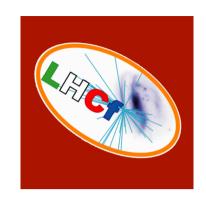
- **DPMJET PYTHIA** are in good agreement at high- η only for $E_{\gamma} < 1.5 \text{TeV}$.
- QGSJET SIBYLL
 are in reasonable
 agreement for
 high- η
- **EPOS** has less η dependency.

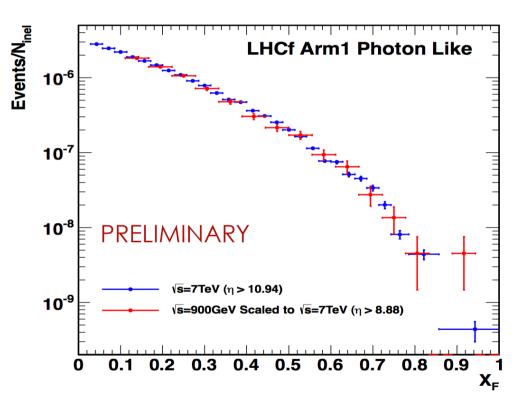
No models reproduce the LHCf data perfectly

PLB 703 128 (2011)

LHCf: p-p run 0.9 and 7 TeV Feynman X_F scaling

The Feynman scaling is important to extrapolate the cross sections information at cosmic ray high energy.





- p-p runs: at \sqrt{s} = 7 TeV and at \sqrt{s} = 900 GeV spectrum.
- the two spectra were normalized by the inelastic events for X_F > 0.1

hep-ex: 1507.08764

- Forward photon spectra measured by LHCf at 7 TeV and 900 GeV support X_F scaling
- The same analysis will be done for RICHf at 510 GeV

LPM effect (PRELIMINARY)

At high energies multiple scattering reduces the cross sections for bremsstrahlung and pair production processes (the Landau Pomeranchunk Migdal effect)

-> Delay and significant elongation of electromagnetic shower.

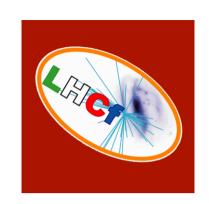
LPM effect is evident for k>E_{LPM}

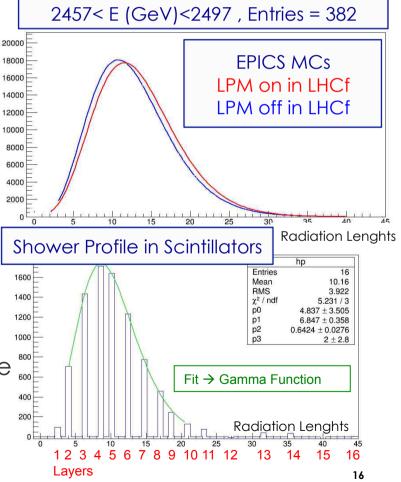
- $E_{LPM} = m^2 c^3 \alpha X_o / 4\pi \hbar$
- k= photon energy

And small emission angle θ ->

LHCf has the optimal set up to detect It directly (by measuring the shape of longitudinal profile of the shower):

we use EPICS software for simulations with and without LPM effect and parametrize the shower shape using the scintillators signal in the layers of ARM2





LPM effect (PRELIMINARY)

SIn= S2+S3 -> more in the poster session

We define the beginning of shower as the signals recorded in Layer 2 and 3

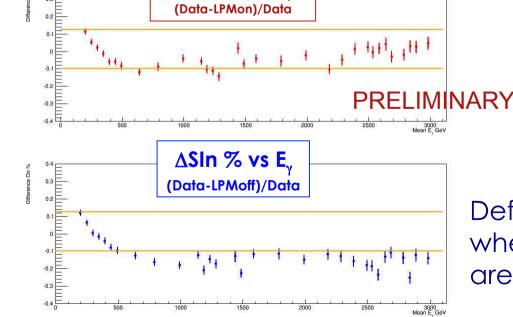


It will be smaller for LPM on than LPM off because of shower delay

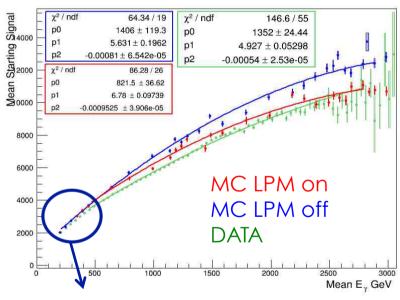
MCs do not reproduce the Data at low energy (<600 GeV)

MCs are tested for LPM effect up to energies of ~300 GeV _____

ΔSIn % vs E,



Sin vs E_{γ}



Define a compatibility range where MC LPM on and off are equal

LHCf: p-p run at 13 TeV

(9x10¹⁶ eV proton hitting the atmosphere)

Special physics run from 8 to 12 June: low luminosity (10²⁹ cm⁻²s⁻¹, 5 orders of magnitudes below the nominal one of LHC) to avoid pileup and improve angular precision



■ First collisions for physics (STABLE BEAMS) were at the midnight of 10 June and until morning of 13 June: 6 STABLE BEAM PERIODS

LHCf takes successfully more than 32 hours of physics Data.

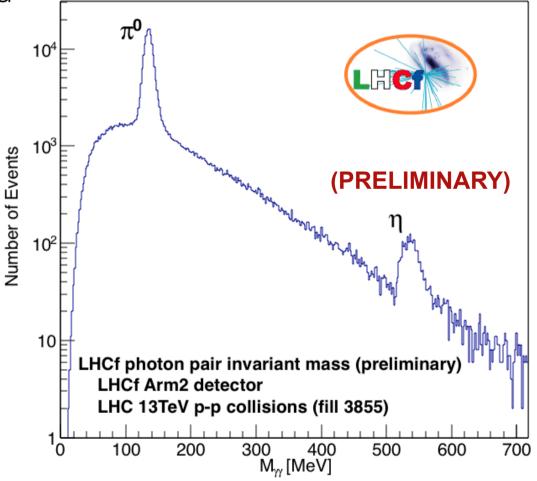
- This operation was in collaboration with the ATLAS experiment, the trigger signals of LHCf were sent to ATLAS that recorded their data accordingly. The analyses of common events will permit to classify the events based on the nature of the processes (diffractive dissociation and non-diffractive).
- The LHCf detectors were uninstalled from the LHC tunnel on 15 June during the first technical stop of LHC to avoid radiation damage expected with the higher luminosity lasting in the LHC RUN2.

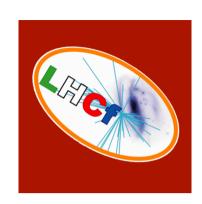
LHCf: p-p run at 13 TeV

(9x10¹⁶ eV proton hitting the atmosphere)

With a preliminary analysis during the operation not only π_0 mesons but also η mesons were clearly

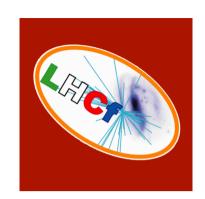
defected





Conclusion and Future plan

LHCf is a small experiment with the aim to understand the hadronic interaction taking place in the collisions between high-energy cosmic rays and earth's atmosphere:



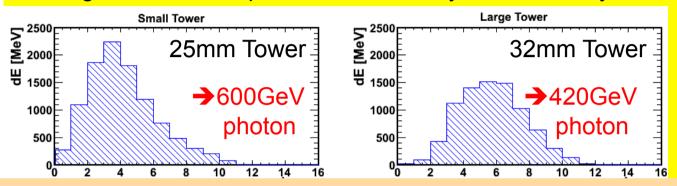
- We measured photons, neutral pions and neutrons for p-p interactions at \sqrt{s} = 900 GeV, 2.76 TeV, 7 TeV, 13 TeV and p-Pb interactions at \sqrt{s} =5.02 TeV
- The analysis of hadronic component for p-p 7TeV is completed
- The π_0 analysis (NMF) for p-Pb run is already published
- Atlas/LHCf common analysis on data taking in 2013 is ongoing. Its follow up in the 2015 is under discussion
- The analysis of data taken in p-p run at 13 TeV is ongoing
- Data taking p-Pb at LHC (2016-2017) is been considered
- LHCf->RICHf: operation at RICH for p-p at 510 GeV (2017)



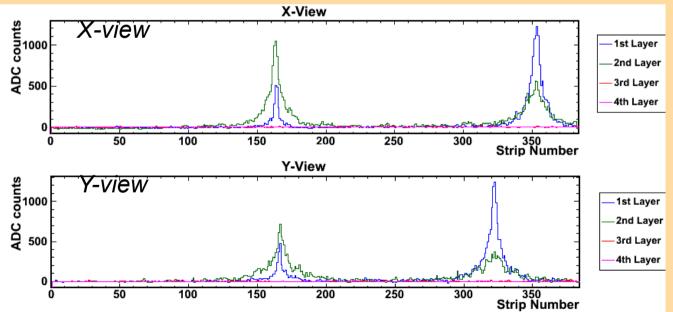
Bk Slides

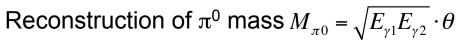
π^0 analysis reconstruction

Longitudinal development measured by scintillator layers



Transverse profile measured by silicon μ–strip layers







Determination of **energy** from total energy release

PID from shape

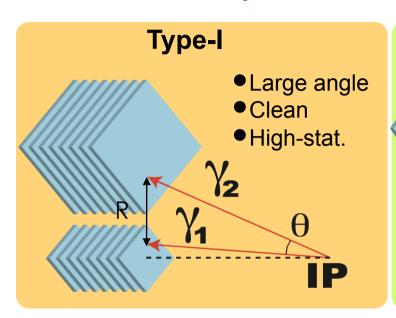
Determination of the **impact point**

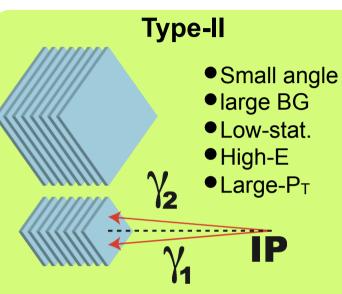
Measurement of the **opening angle** of gamma pairs

Identification of **multiple hit**

22

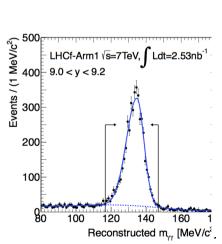
π^0 analysis reconstruction

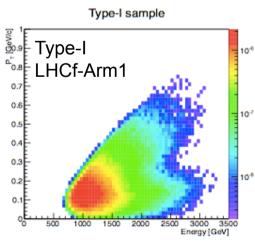


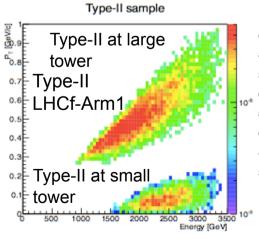


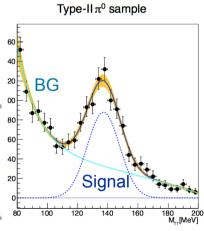


$$\theta = \frac{R}{140 \text{ m}}$$

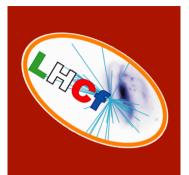


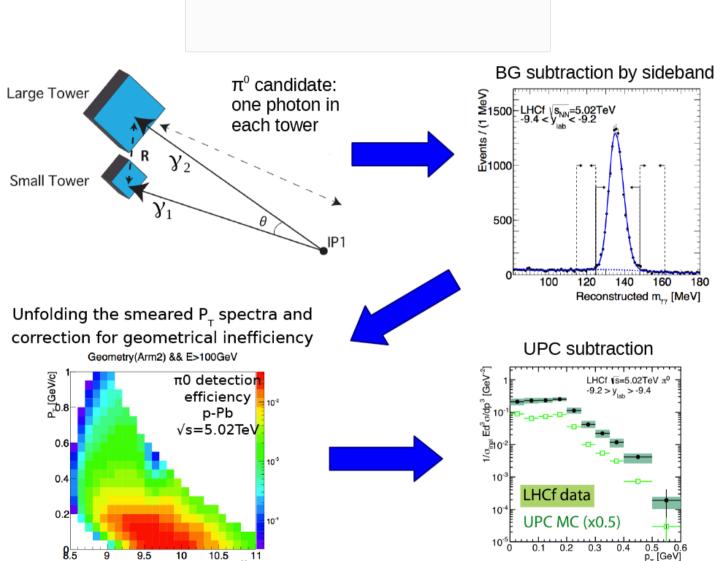






π⁰ analysis reconstruction



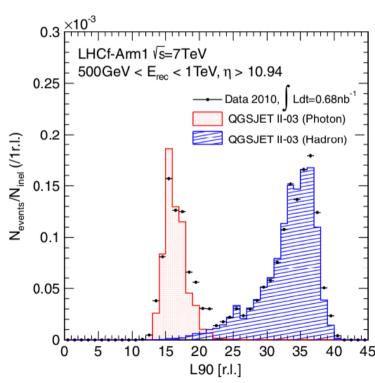


Photons and hadron selection



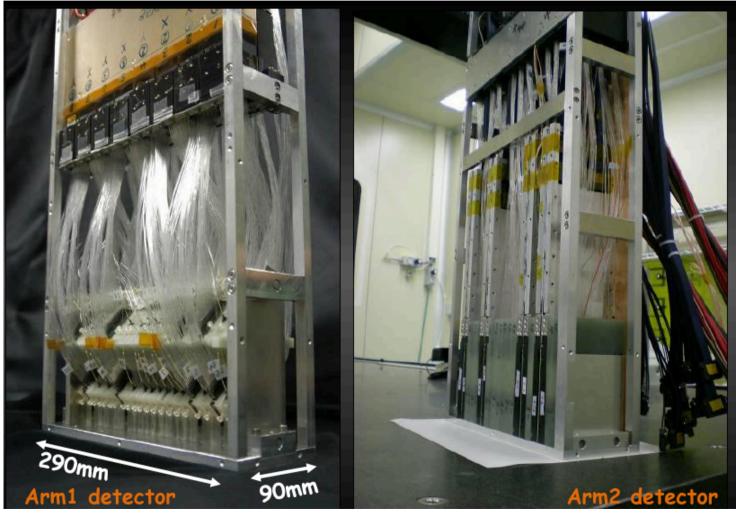
EM and hadronic showers are separated with a method based on a difference of the longitudinal shower development

- L90%: depth in X₀ where 90% of the deposited energy is contained
- Energy-dependent threshold to keep photon detection efficiency at 90%
- Events with L90% less than threshold are recognized as photons

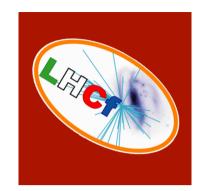


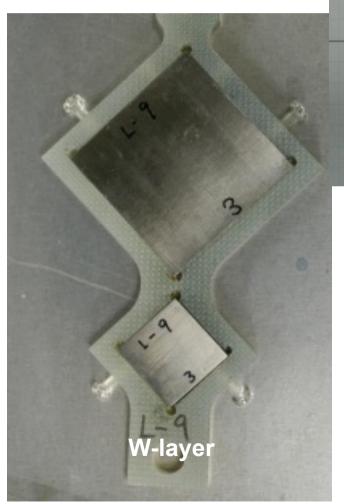
PHOtOS

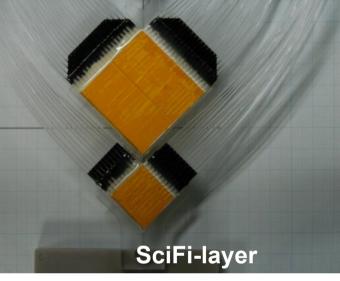


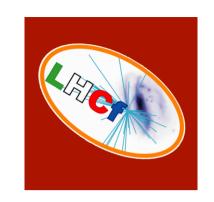


PHOtOS





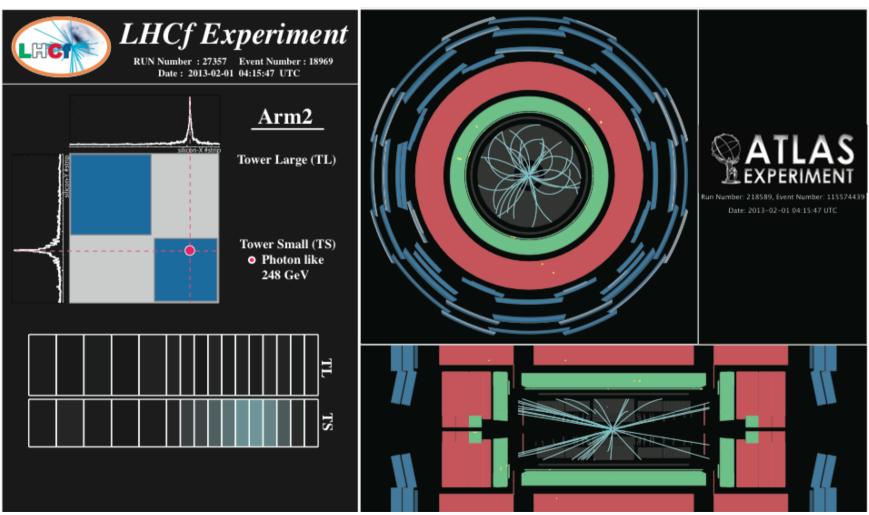




year		Equivalent proton energy in LAB (eV)	γ	n	π^0
	SPS test beams		NIM A 671 129 (2012)	JINST 9 P03016 (2014)	
2009	p-p 900 GeV	4.3x10 ¹⁴	PLB 715 298 (2012)		
2009/2010	p-p 7 TeV	2.6x10 ¹⁶	PLB 703 128 (2011)	In preparation	PRD 86 092001 (2012)
2013	p-p 2.76 TeV	4.1x10 ¹⁵			Phys. Rev. C
2013	p-Pb 5.02 TeV	1.3x10 ¹⁶			89, 065209 (2014)
2015	p-p 13 TeV	9.0x10 ¹⁶		Data taken	







Neutrons analysis in p-p collisions at 7 TeV

(2.6x10¹⁶ eV proton hitting the atmosphere)



- $\Delta x \sim (0.2 \div 1) \text{mm}$
- Δ E/E ~(35 ÷ 40)% -> high leakage
- True spectrum is smeared by detector response
 - Unfolding is needed to extract physics results



