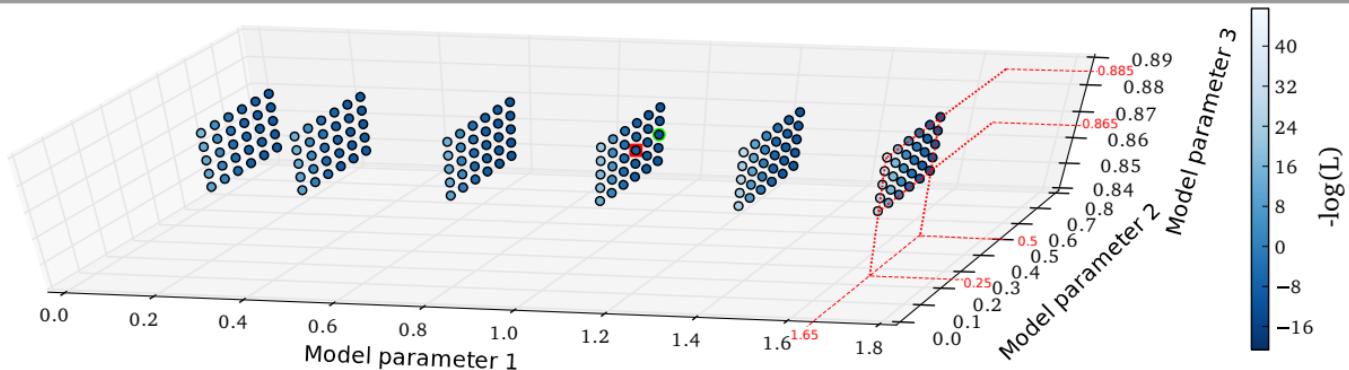




Combined analysis of accelerator and ultra-high energy cosmic ray data

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Motivation

- Hadronic interaction models introduce large systematic uncertainties for air shower measurements
 - Phenomenological
 - Over 100 parameters
- Tuning of parameters
 - To low-energy fixed target measurements and
 - New data forward phase-space by the LHC
 - Air shower measurements contain information about hadronic interactions
 - Tedious parameter tuning process → optimisation
- How to use air shower data?
 - Interpretation always depends on the composition
 - Either rely on weak constraints or
 - Reduce dependencies by dedicated selections

Method and benefit

Multiple measurements
(LHC and cosmic rays)



$$\mathcal{L} = \mathcal{L}_{\text{astro}} \mathcal{L}_{\text{LHC}}$$

Agreement Data-MC



Modify model
parameters



Sensitivity of parameters

- Likelihood-based method
 - More flexible than χ^2 -approach
 - Include asymmetric uncertainties and complex functions for CR data
 - treat many measurements in a consistent way
- Allows to decide which parameters are important
- Connection between changes in LHC prediction and CR prediction
- Later: optimisation of parameters

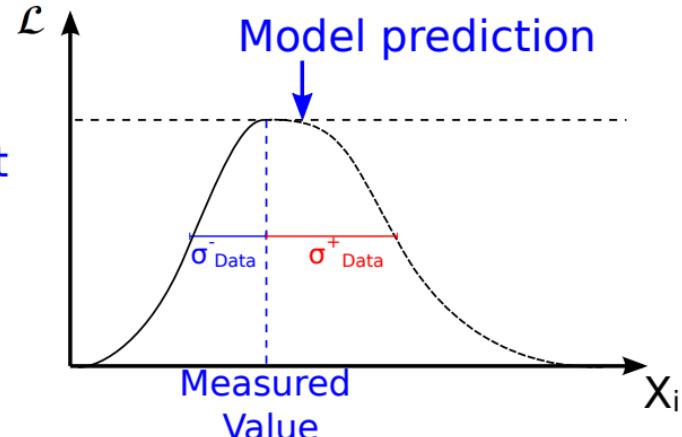
EPOS model: parameters for diffraction and cross section

- EPOS-LHC model
 - widely used for cosmic ray physics
 - tuned to LHC data (up to 7 TeV)
- Current parameters under investigation
 - Name: “wdiff(2)”: $w_{\text{diff}} = \{0.865, 0.870, \textcolor{red}{0.875}, 0.880, 0.885\}$
 - Probability of **diffraction** for protons
 - Name “epscrx”: $\epsilon_{\text{crx}} = \{0.25, 0.3, 0.35, \textcolor{red}{0.4}, 0.45, 0.5\}$
 - Scale parameter for saturation scale in scattering amplitude. Influences high-energy **cross section**
 - Name “alpdi”: $\alpha_{\text{diff}} = \{0.2, 0.4, 0.725, \textcolor{red}{1.05}, 1.35, 1.65\}$
 - Modifies the **mass distribution** of **diffractive** events $\sim 1/M^{(2 \alpha_{\text{diff}})}$
 - Larger values of α_{diff} → more low-mass states (influences forward multiplicity)

➡ Grid of 180 parameter sets

Analysing collider measurements

- Asymmetric gaussian distribution for probability function for each data point
- Included LHC measurements
 - Cross sections:
 - p-p at 7 and 8 TeV (TOTEM)
 - p-Pb at 5 TeV/nn (CMS)
 - Multiplicity: forward and central charged particles at 8 TeV (CMS+TOTEM)
 - Soft diffractive cross section at 7 TeV (CMS)
 - Very forward neutrons at 7 TeV (LHCf)
- Started to include pion charge ratio (NA49) for low-energy data



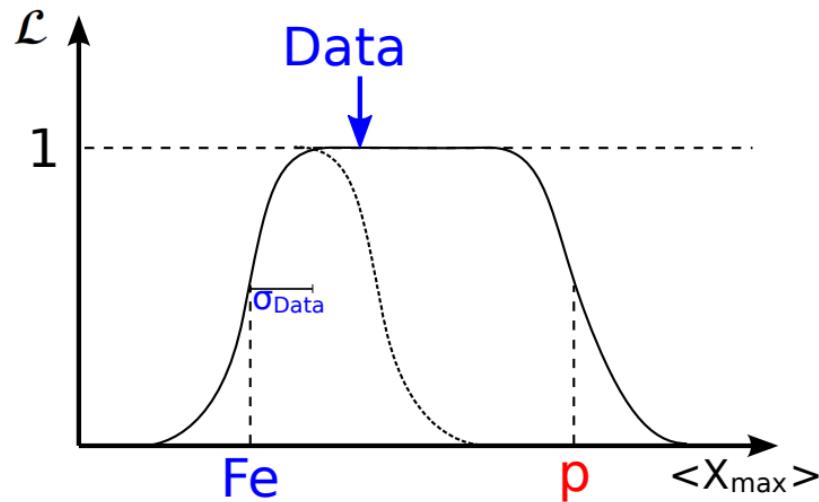
Integrating ultra-high energy cosmic ray measurements

- Mass composition unknown at highest energies
- Use measurements of longitudinal shower profile
- Constrain data without any assumptions
 - Physically allowed regions
→ Predictions of proton and iron must bracket data
 - Fit tail of distribution (only light particles, even He less important)
- Constrain with weak assumption on mass
 - $\ln(A)$ and $V[\ln(A)]$ cannot be below 0
 - Fit distributions with templates for multiple components
- Later: Muon measurements

Cross section
Elasticity/Diffraction
Multiplicity

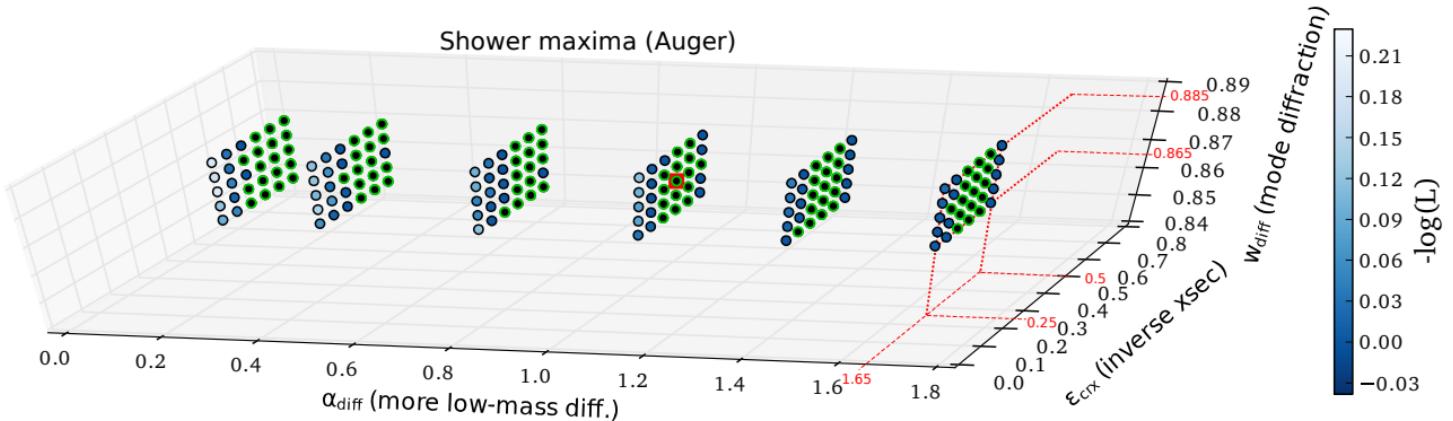
Diff. mass distribution
Meson/Baryon production

Average shower maximum

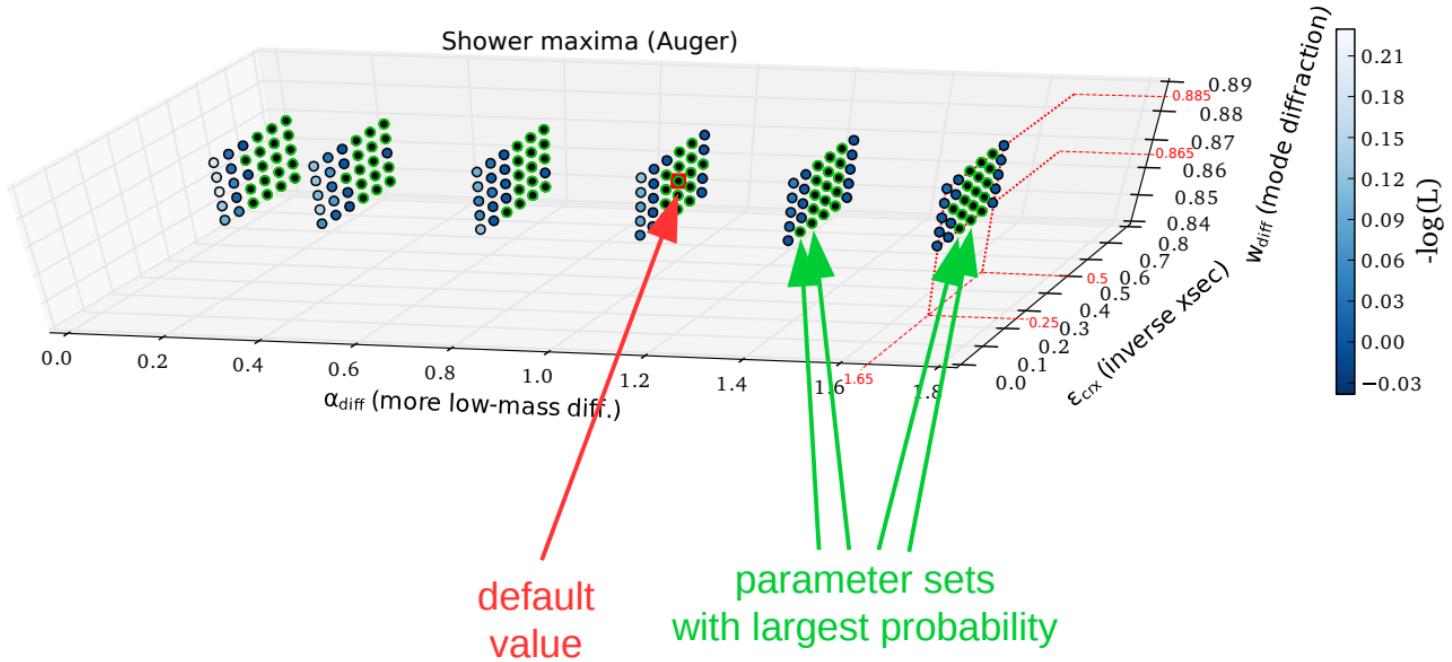


$$P(X_{\max}^{\text{Fe/p}} | X_{\max}, \sigma) = \begin{cases} c_1 \text{Gaus}(X_{\max} + \sigma - X_{\max}^{\text{p}}, \sigma) & \text{for } X_{\max} \geq X_{\max}^{\text{p}} - \sigma \\ c_2 & \text{for } X_{\max}^{\text{Fe}} + \sigma < X_{\max} < X_{\max}^{\text{p}} - \sigma \\ c_1 \text{Gaus}(X_{\max} - \sigma - X_{\max}^{\text{Fe}}, \sigma) & \text{for } X_{\max} \leq X_{\max}^{\text{Fe}} + \sigma. \end{cases}$$

Average shower maximum



Average shower maximum



Average shower maximum

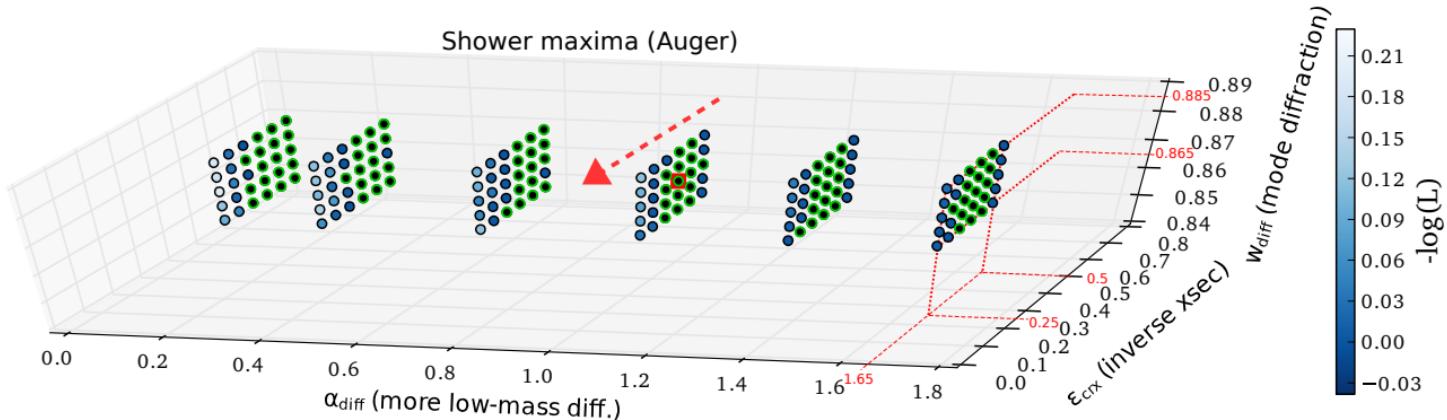
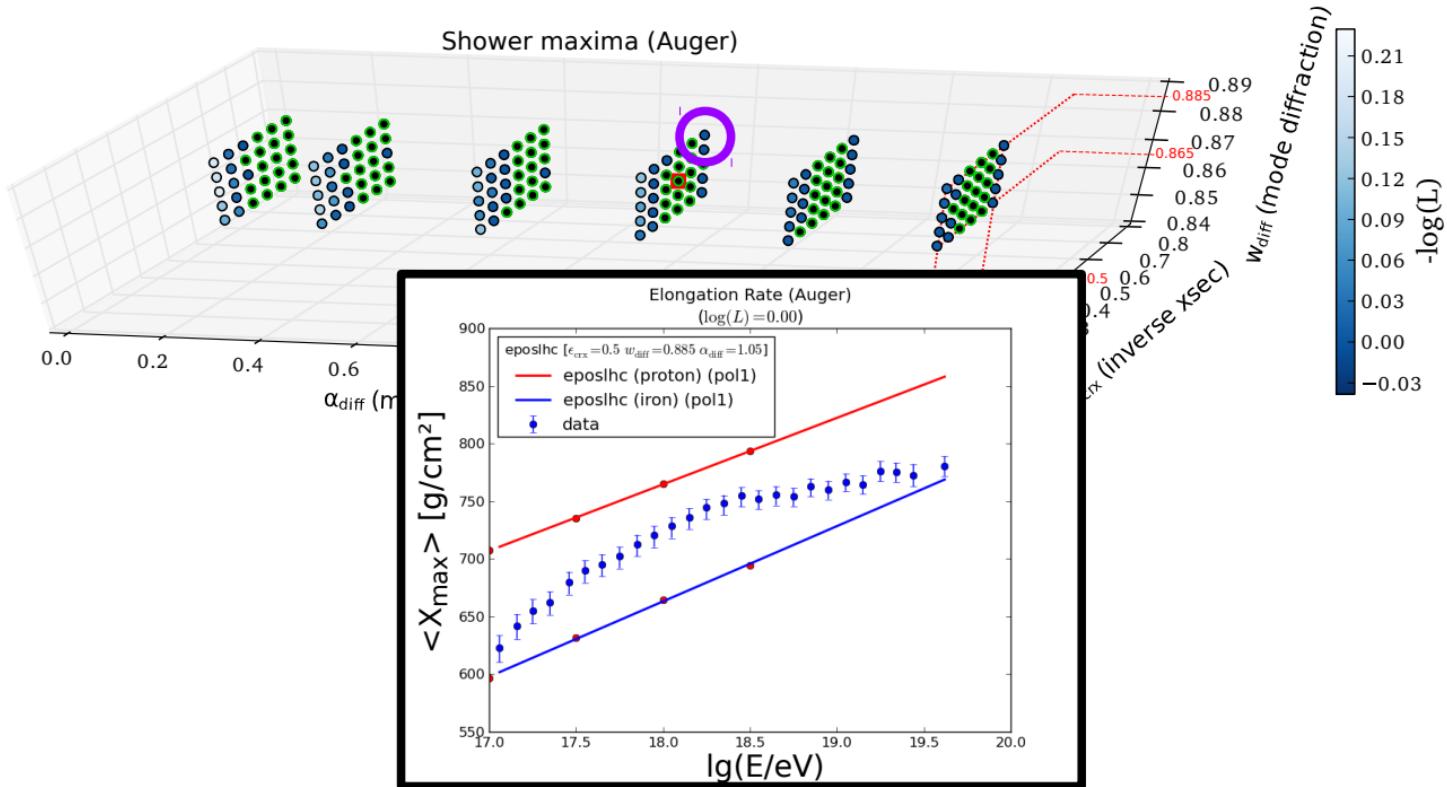
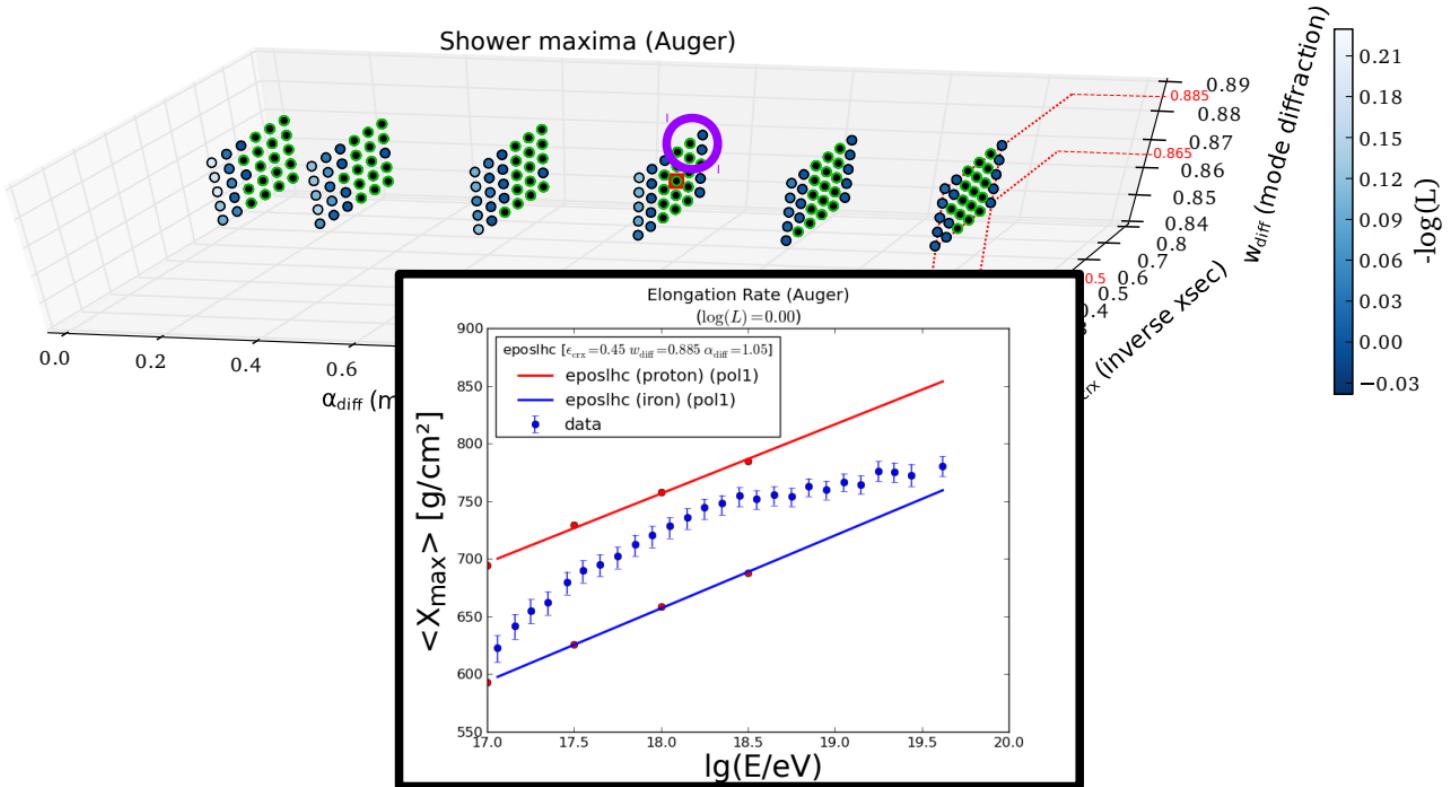


Illustration for increasing
extrapolated cross section
 \Rightarrow lower X_{max} values

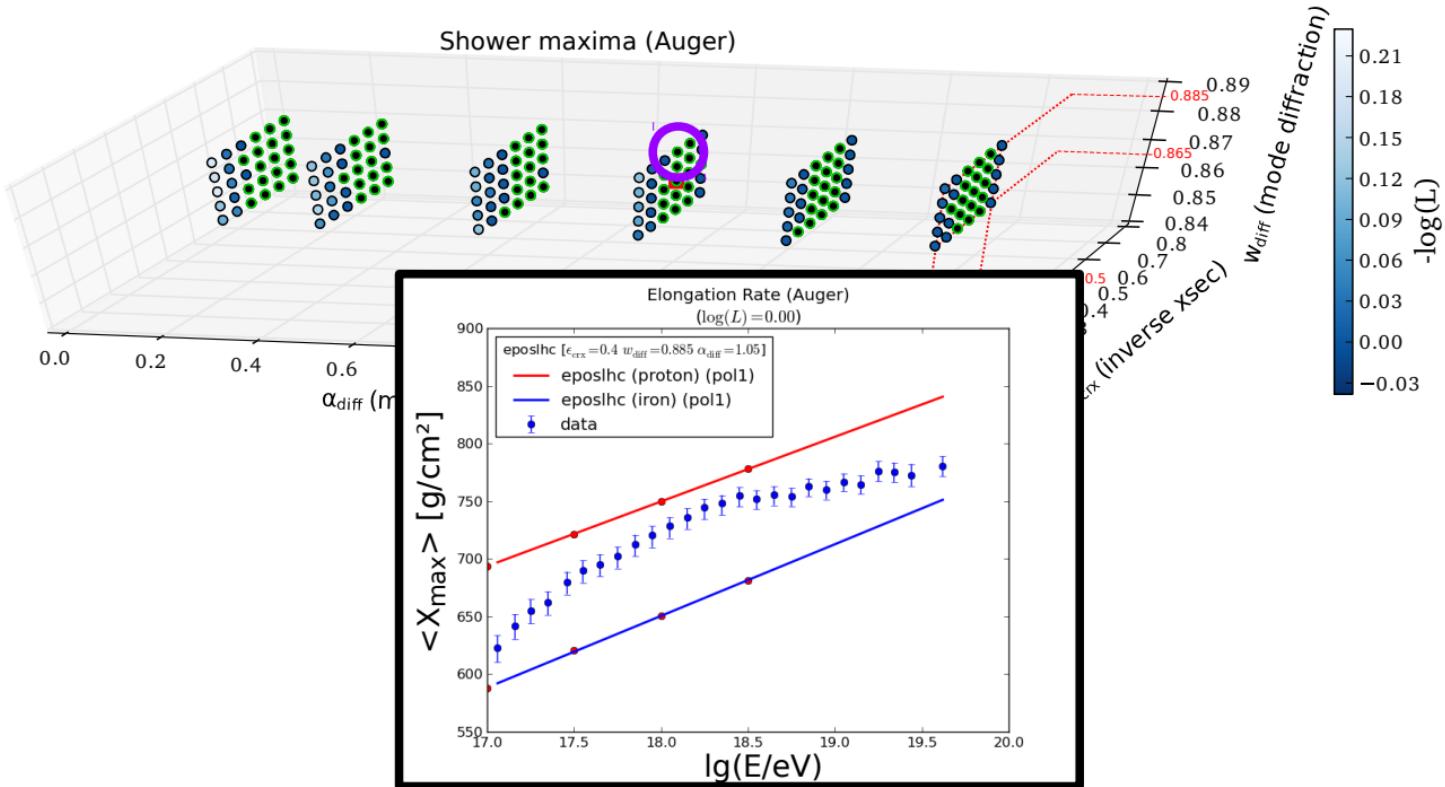
Average shower maximum



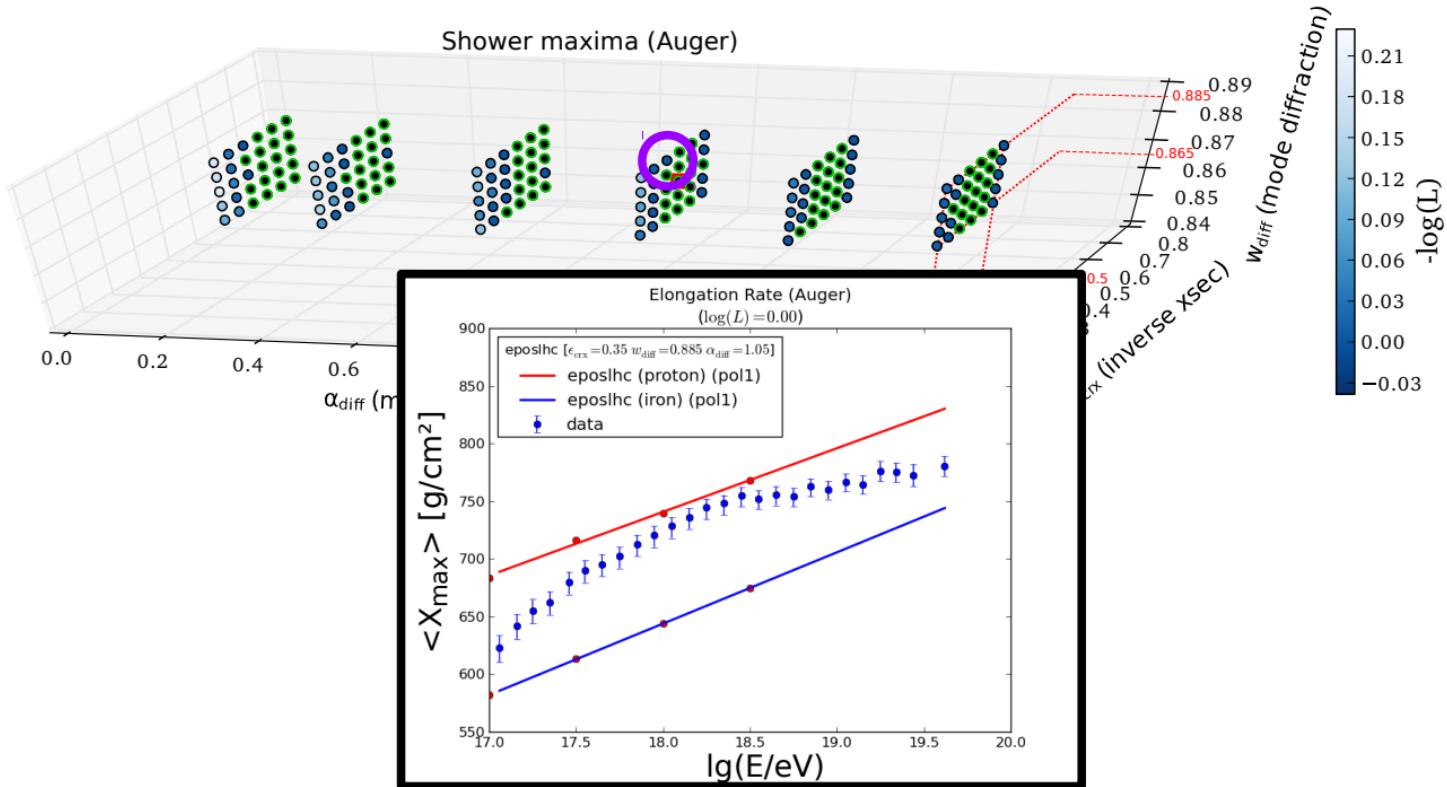
Average shower maximum



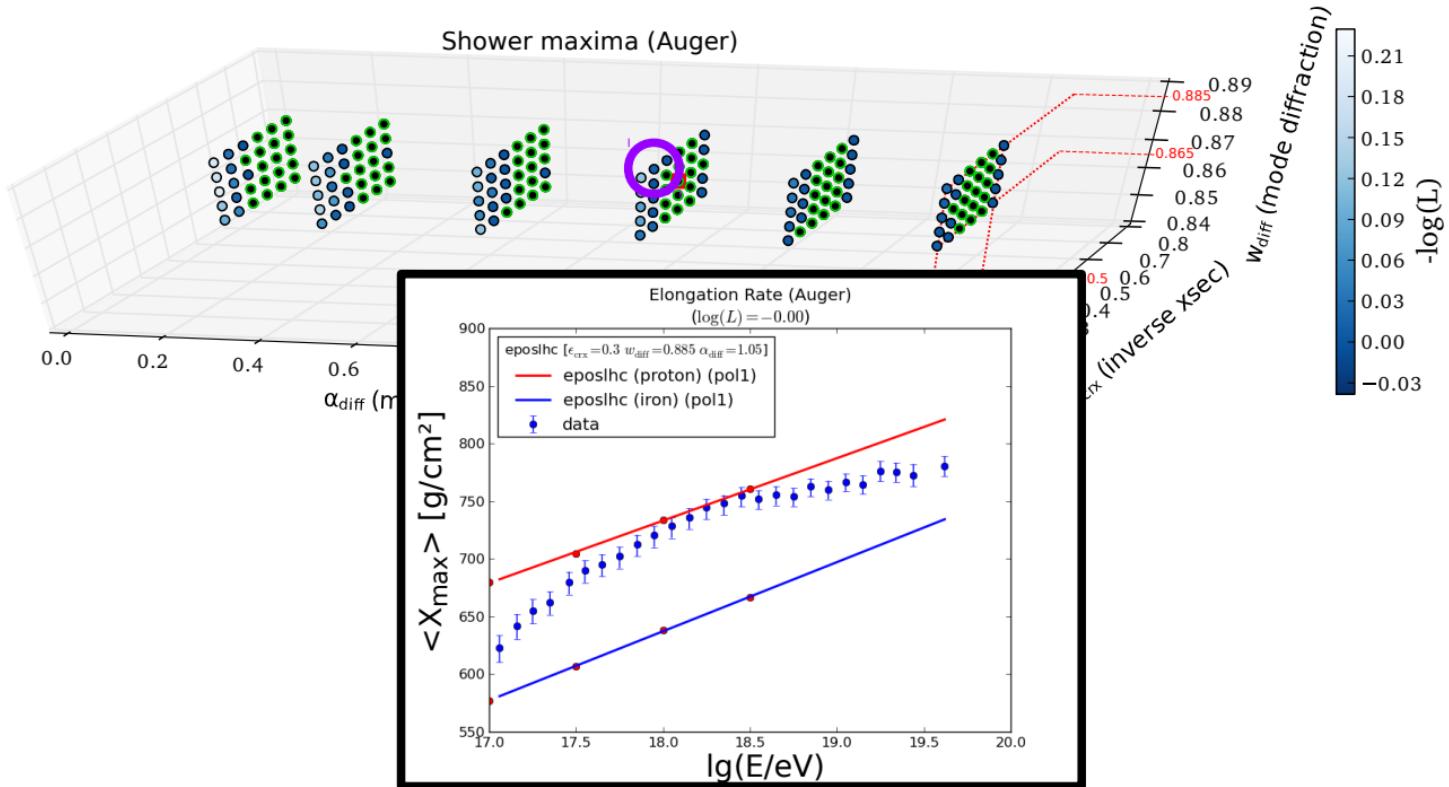
Average shower maximum



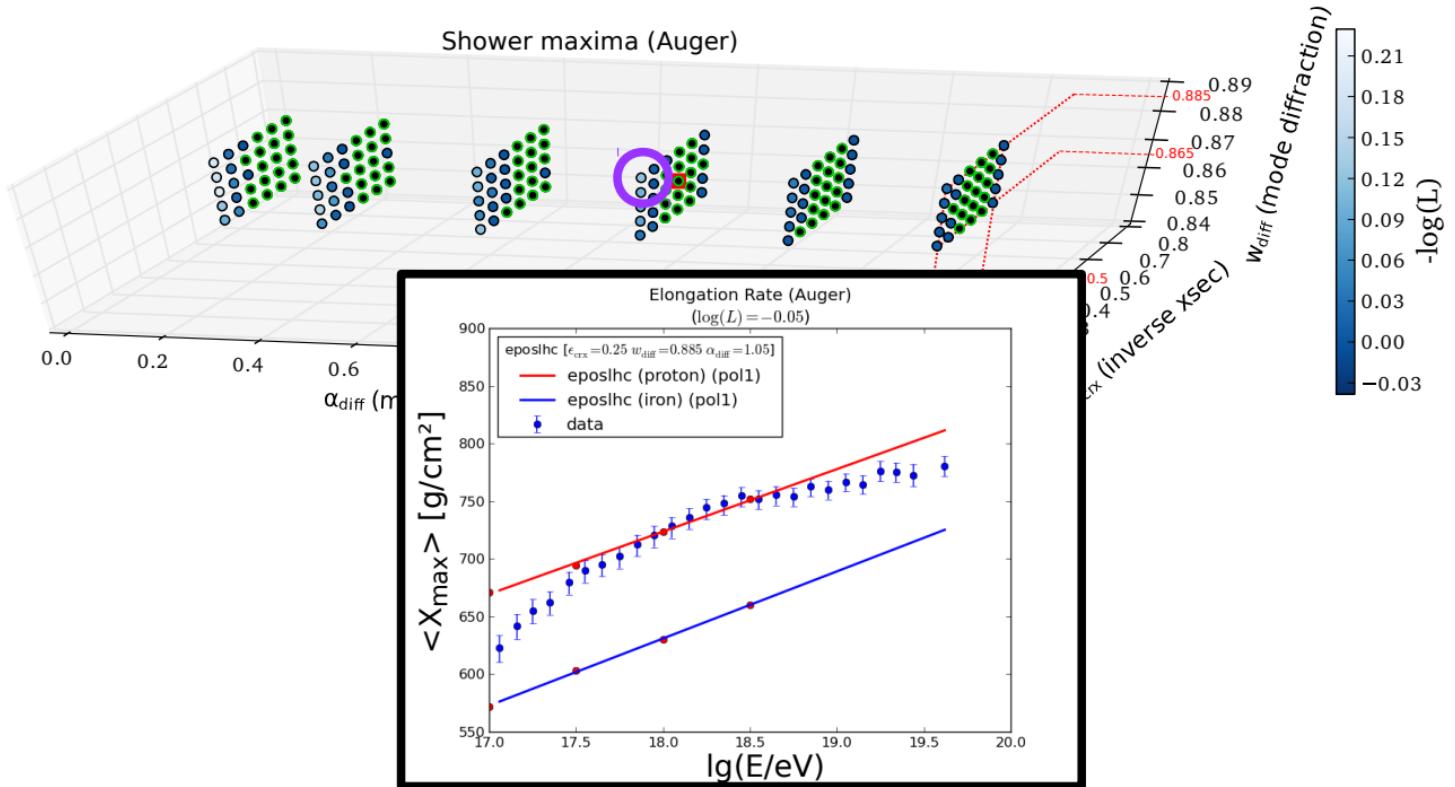
Average shower maximum



Average shower maximum

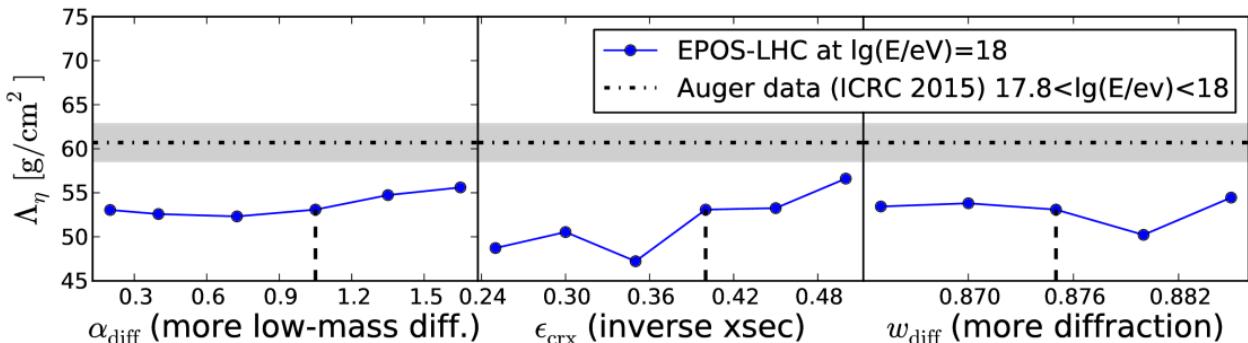
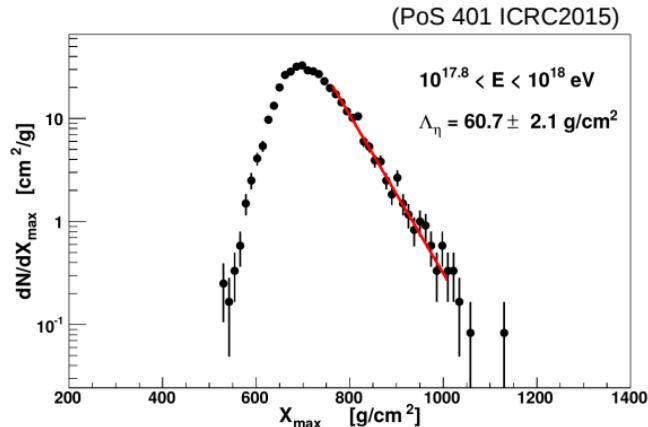


Average shower maximum



Tails of shower maxima

- Use additional information from distribution of shower maxima
- Tail of distribution at low Auger energies dominated by protons
- Unbinned \mathcal{L} -fit of 20% deepest penetrating showers (weights for p and He)
- Helium fraction always small (<10%) and no visible influence on Λ_η



Technical Challenges

- Air shower
 - 1D-approach with CONEX
 - 10^{17} to 10^{19} eV (p,He,N,Fe)
 - each 1,000-10,000 showers
- LHC-type
 - CRMC (Cosmic Ray Monte Carlo) program
 - developed to simulate single collisions in fixed or center-of-mass frames
 - keeping particle ancestry information in common output formats
 - 4 different energies for p-p and p-Pb collisions

CPU time



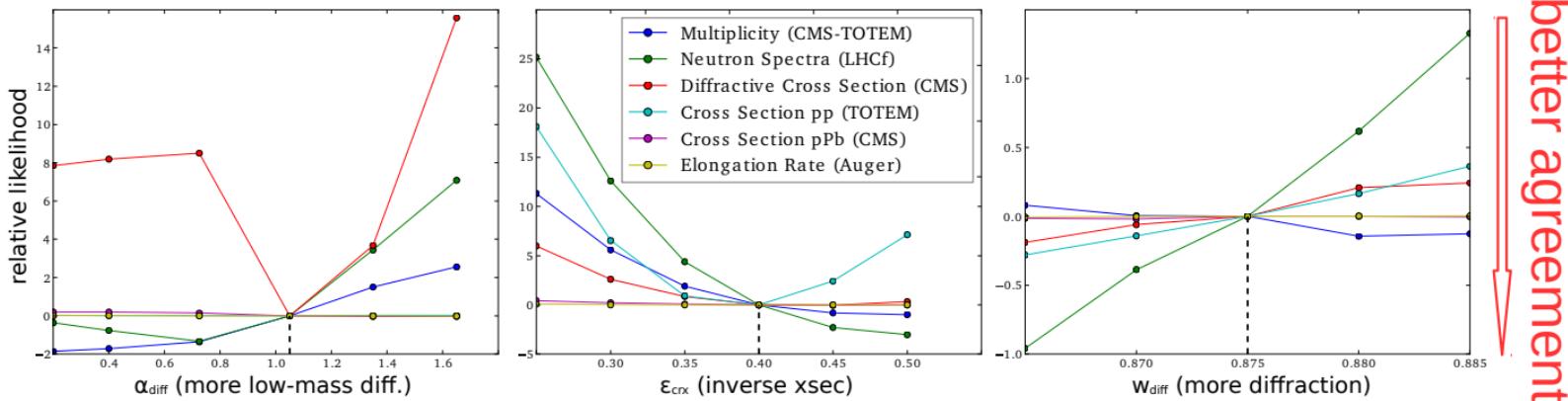
Storage



CERN cluster

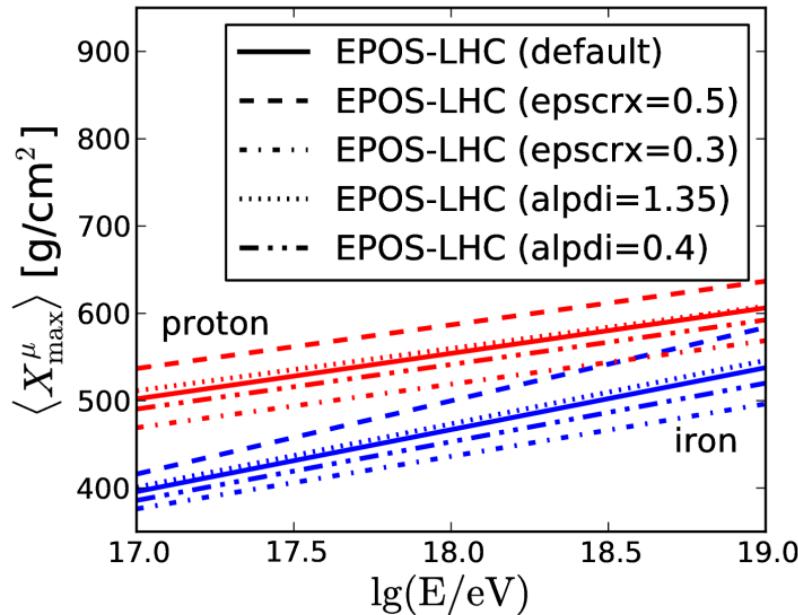
Comparison to data

- Construct likelihood relative to default and “marginalised” (average over dimensions)
- Parameters well constrained except for diffractive probability in parameter range
- Default parameters reasonable
- Some tension between the data



Studying parameter influence on possible muon deficit

- One interesting aspect is the possible muon deficit with data from Auger, IceTop, Kascade
- Muon Production Depth (MPD):
 - For data comparison: 3D shower simulations necessary
 - Measurements for muons in certain lateral regions
 - Strong influence on hadronic physics and low prospective to use for mass composition



Conclusion

- We could show that hadronic interactions can be constrained by astrophysical data
- We propose and tested multiple methods with weak assumptions on the mass composition
- We developed a framework to combine these (air shower simulations) with collider measurements (p - p , p - Pb simulations)
- With this framework we derived constraints on three parameters on the recent tune of the EPOS model
- EPOS-LHC parameters in general reasonable

BACKUP

Helium fraction

