Search for an enhanced emission of neutrinos from the Southern Sky with the ANTARES telescope

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On behalf of the ANTARES Collaboration



Overview

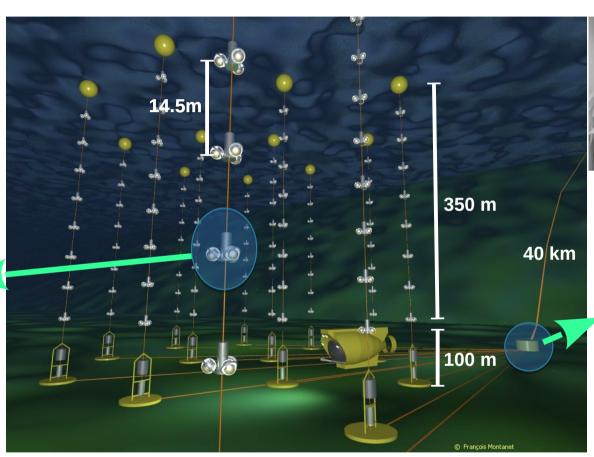
- The ANTARES neutrino telescope
- Search for diffuse fluxes of neutrinos from an enhanced emission scenario
 - Trying to explain the IceCube signal
 - Southern Sky emission region hypothesis
- Sensitivity estimation and unblinding results

The ANTARES detector

885 10" PMTs on 12 lines 25 storeys/line 3 PMTs / storey



storey



shore station



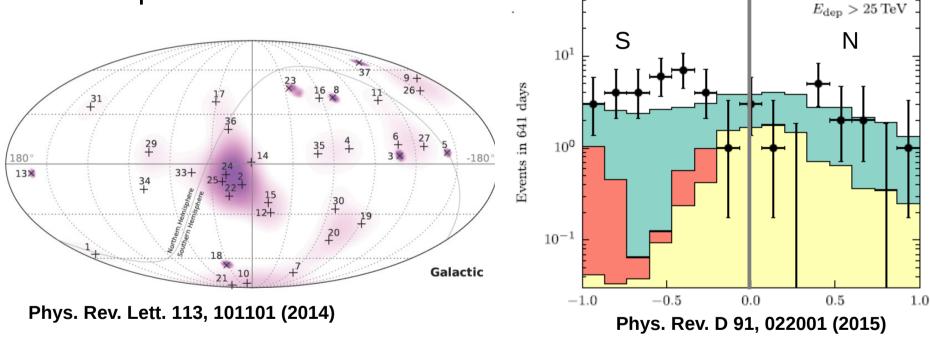
junction box

In the Mediterranean Sea (Toulon): Southern sky visibility Depth of ~2500 m: atmospheric muon background reduced

The IceCube Signal

First evidence of a diffuse cosmic flux of neutrinos → HESE





Some (possible) excess in the Southern Sky → Galactic Plane?

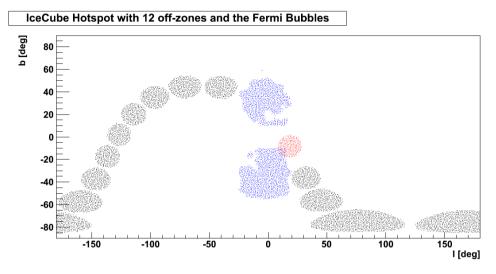
An extended emission region

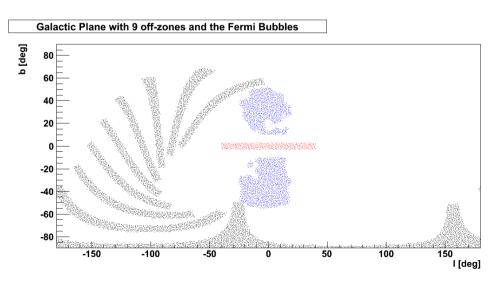
- IceCube signal mainly in the shower channel
 - Angular resolution ~20°
 - → Signal is compatible with isotropy
 - Smaller scale regions can produce part of the signal
- ANTARES looks at the Southern Sky
 - Upgoing muon tracks: much better angular resolution
 - 2007-2013 data, ~1500 days livetime
 - Effective areas in the Southern Sky are compatible in low energy regime
 - An isotropic all sky signal from IceCube can appear as coming from an extended region.

Search method

- Assuming a signal spectrum proportional to E^{-2.4÷2.5}
 - Optimised event selection with the Model Rejection Factor procedure:
 - Track quality parameters to reject wrongly reconstructed atmospheric muons
 - Energy estimation to suppress atmospheric background
- Signal region to be defined:
 - 10° circle around the IceCube "hotspot"
 - Galactic plane area
- Background estimated from data:
 - Off-zones for which the telescope has the same exposure
 - Compare to the event rates from the signal region to the average from off-zones

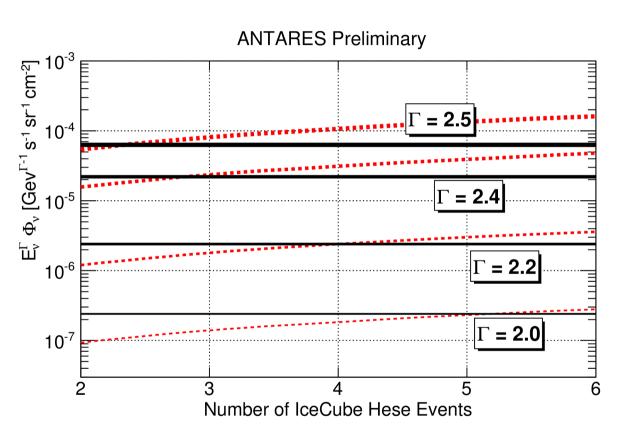
Search method





- Sensitivity depends on the N_{bkg}/Ω
 - Similar exposure for different regions gives same background expectations
- One optimised selection for both regions and spectral indices
 - Differences ~5% with respect to full optimisation
 - The result still holds for analogous regions of the sky

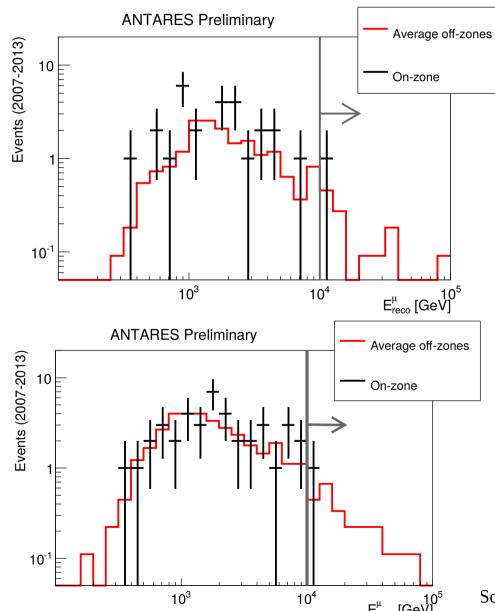
Sensitivity to the IceCube signal



Sensitivity compared to the expected flux producing a certain number of events in the IceCube HESE coming from a 0.1 sr region

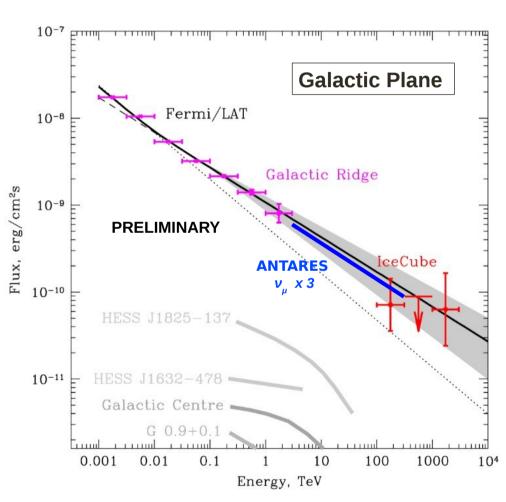
ANTARES can exclude many hypothesis for the origin of IceCube signal

Unblinded results



- 1.0 and 2.5 events expected from background only for the HS and the GP respectively
- 1 event is observed in both the two signal regions
- Fully compatible with background expectations
- Upper limits corresponding to the sensitivity

Upper limits compared to Galactic Plane models



ANTARES, $III < 40^{\circ} lbI < 3^{\circ}$ **Upper limit** $\nu_{\mu} \times 3$ ♦ spectral index: -2.4 10^{-6} $E_{\nu}^{2} d\Phi/dE_{\nu} [\text{GeV cm}^{-2}\text{s}^{-1}]$ \blacksquare spectral index: -2.5 10^{-7} **PRELIMINARY** 10^{-8} KRA_{γ} (cut 5×10^7 GeV) KRA_{γ} (cut 5×10^6 GeV) KRA'(cut 5x107 GeV) KRA (cut 5x106 GeV 10⁰ 10^{2} 10^{1} E_{ν} [TeV]

Lines from Phys. Rev. D 89, 103002 (2014)

Lines from arXiv:1504.00227 (2015)

Conclusions and outlook

- Antares is the largest underwater neutrino telescope, in its 8th year of operation.
- Good performance in Southern Sky searches thanks to its resolution and effective area.
- Stringent upper limits on the possible origin of the IceCube signal from an extended region in the Southern Sky.