



ICRC

The Astroparticle Physics Conference

34th International Cosmic Ray Conference

July 30 - August 6, 2015

The Hague, The Netherlands

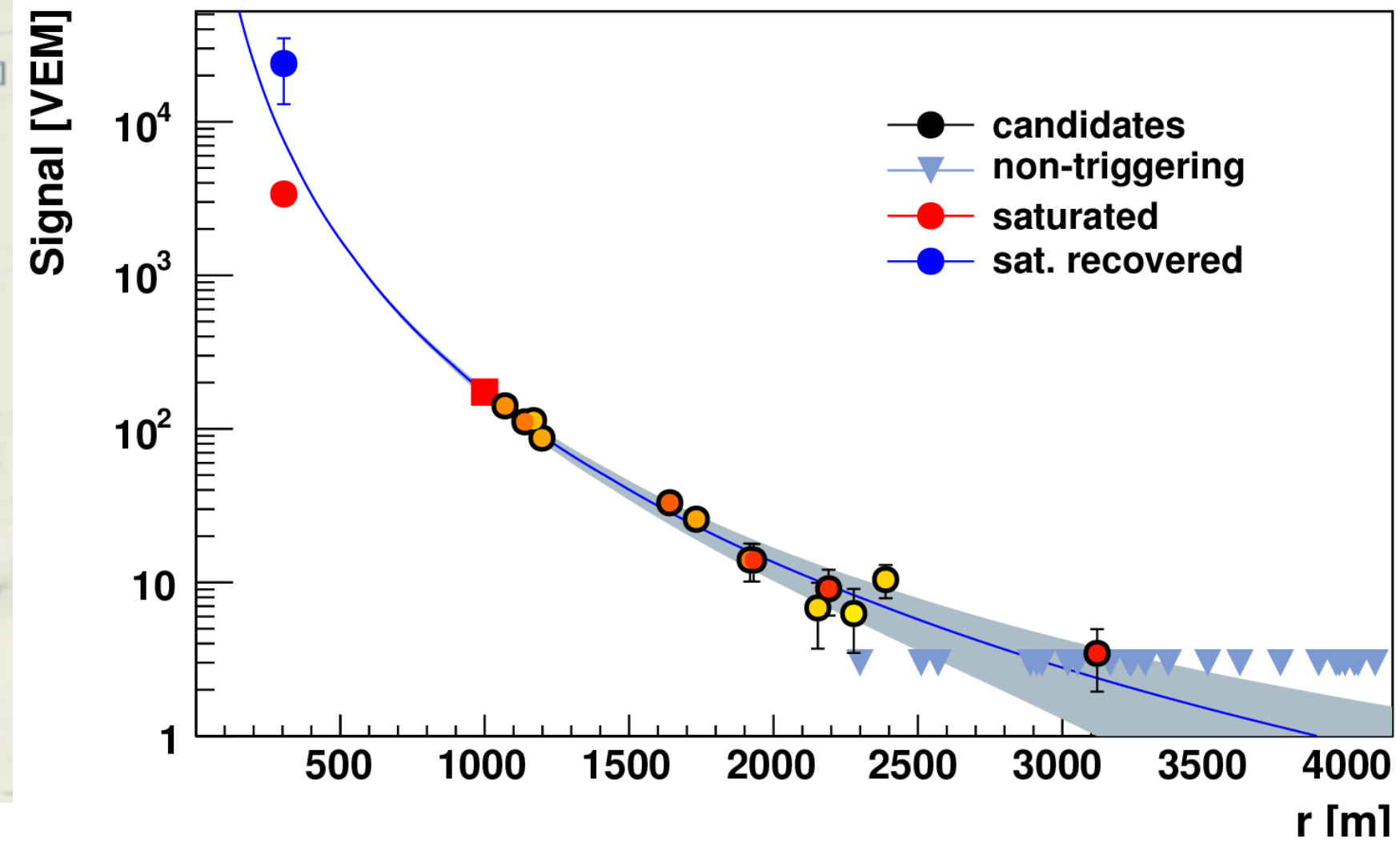
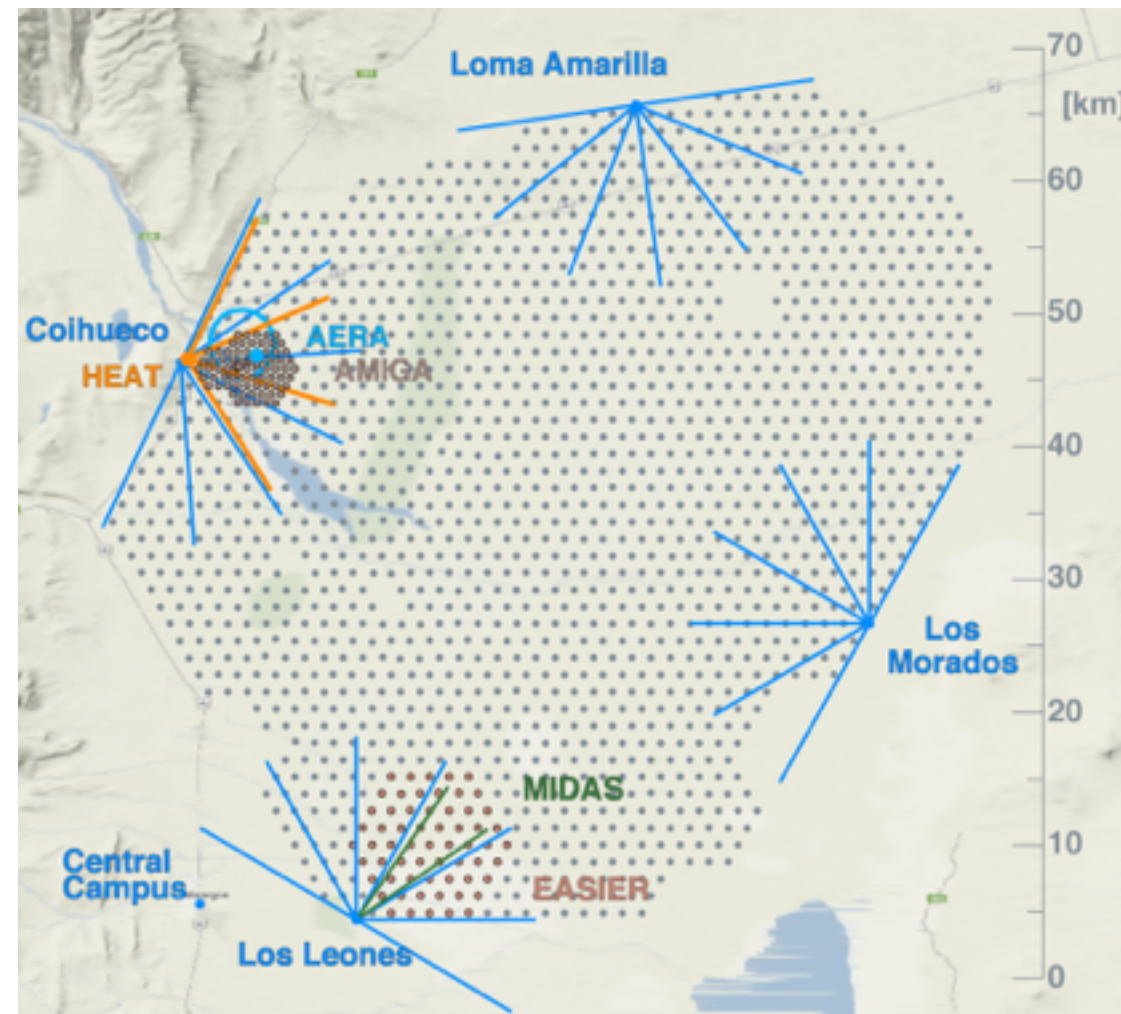


Karlsruher Institut für Technologie

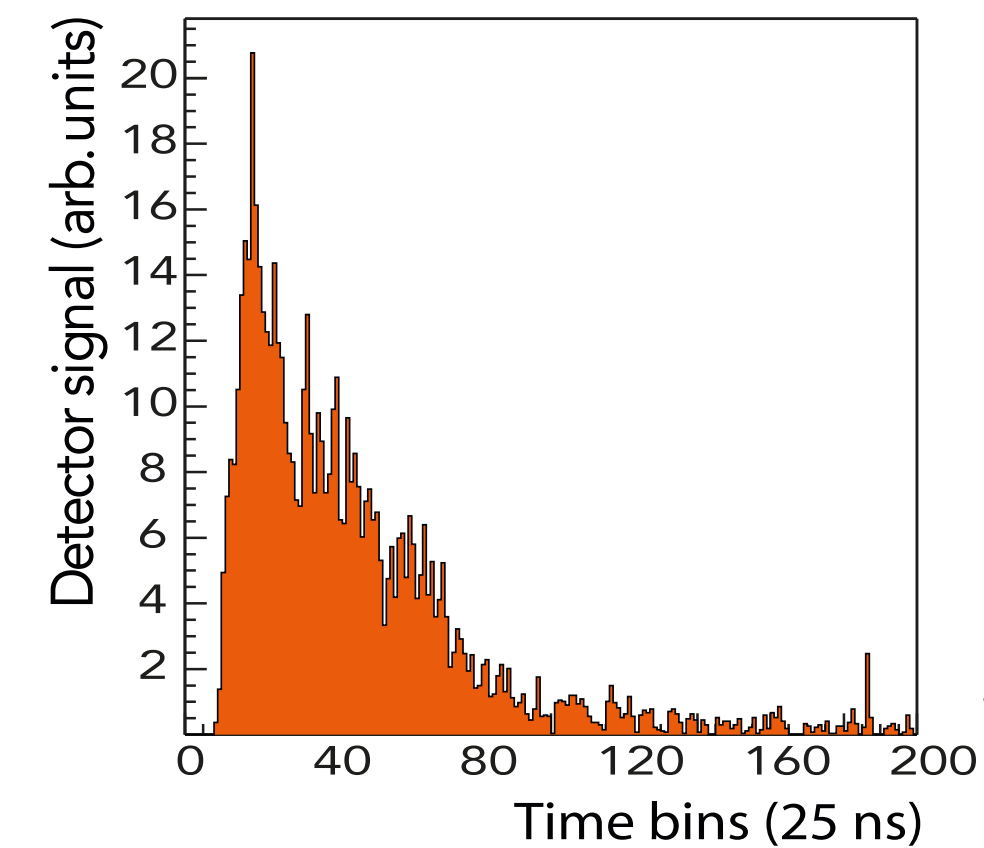
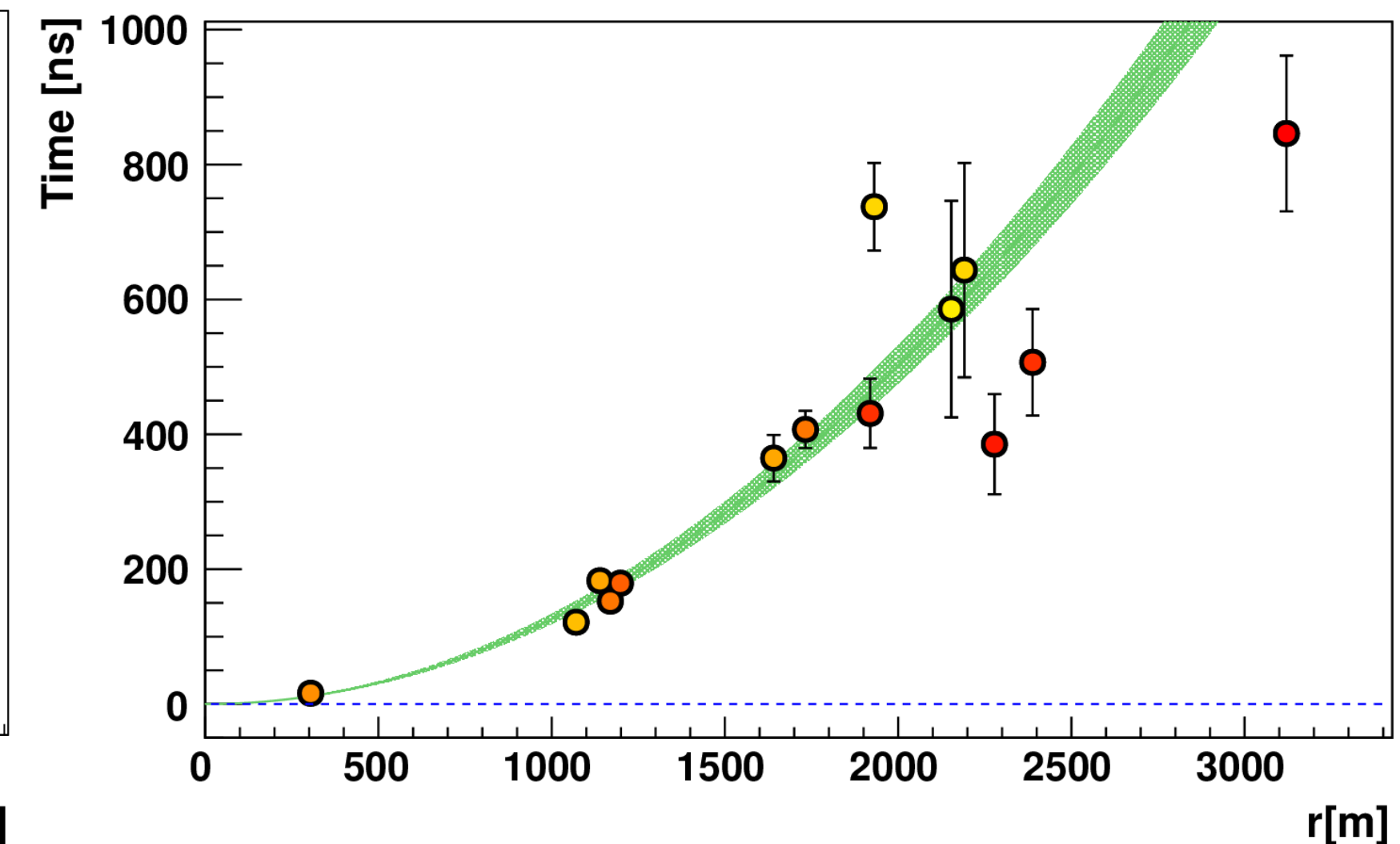
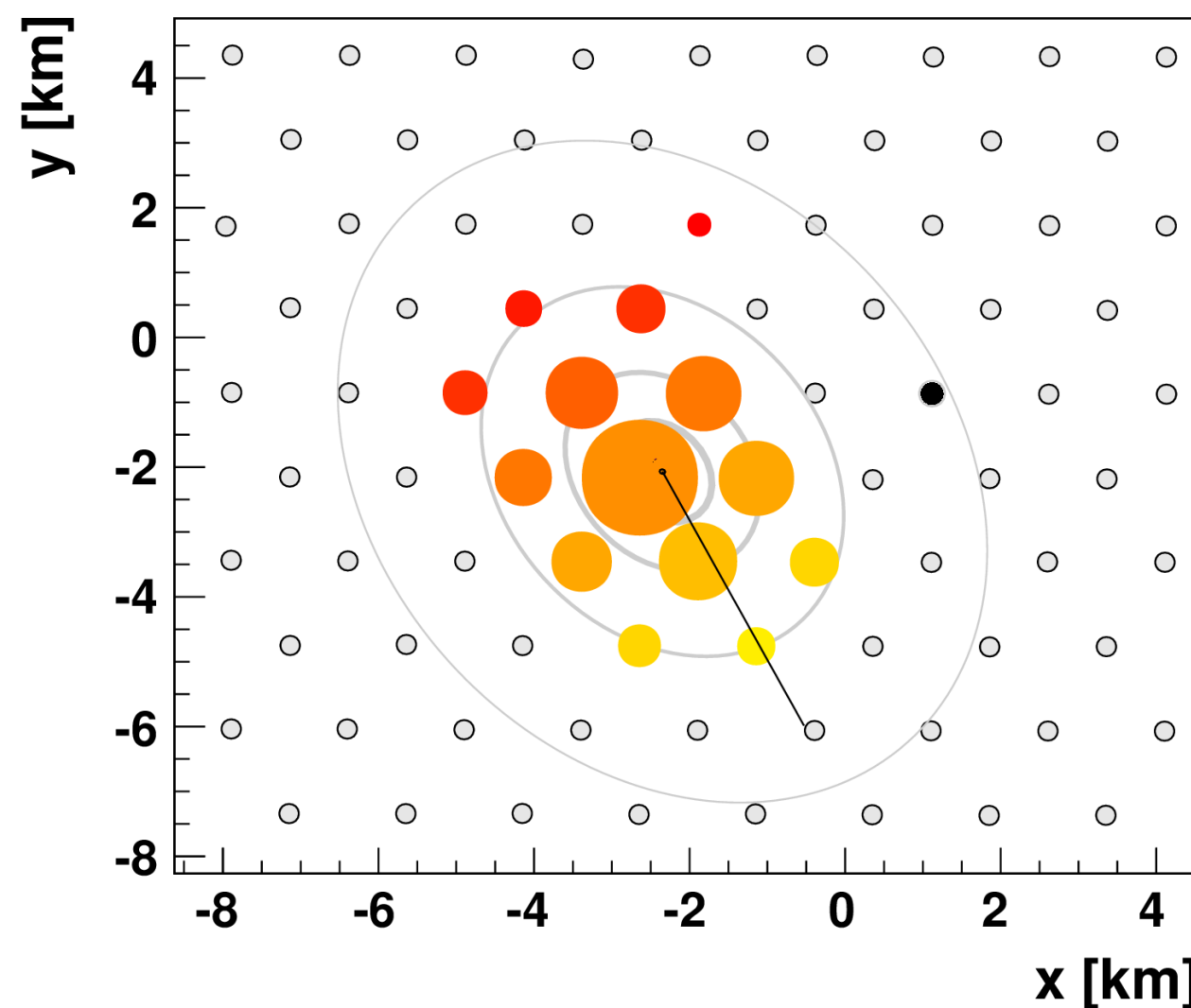
A universal description of temporal and spacial distributions of ground particles in extensive air showers

Maximo Ave Pernas, **Markus Roth**, Alexander Schulz
Karlsruher Institute of Technology

Reconstruction of surface detector data at e.g. Auger

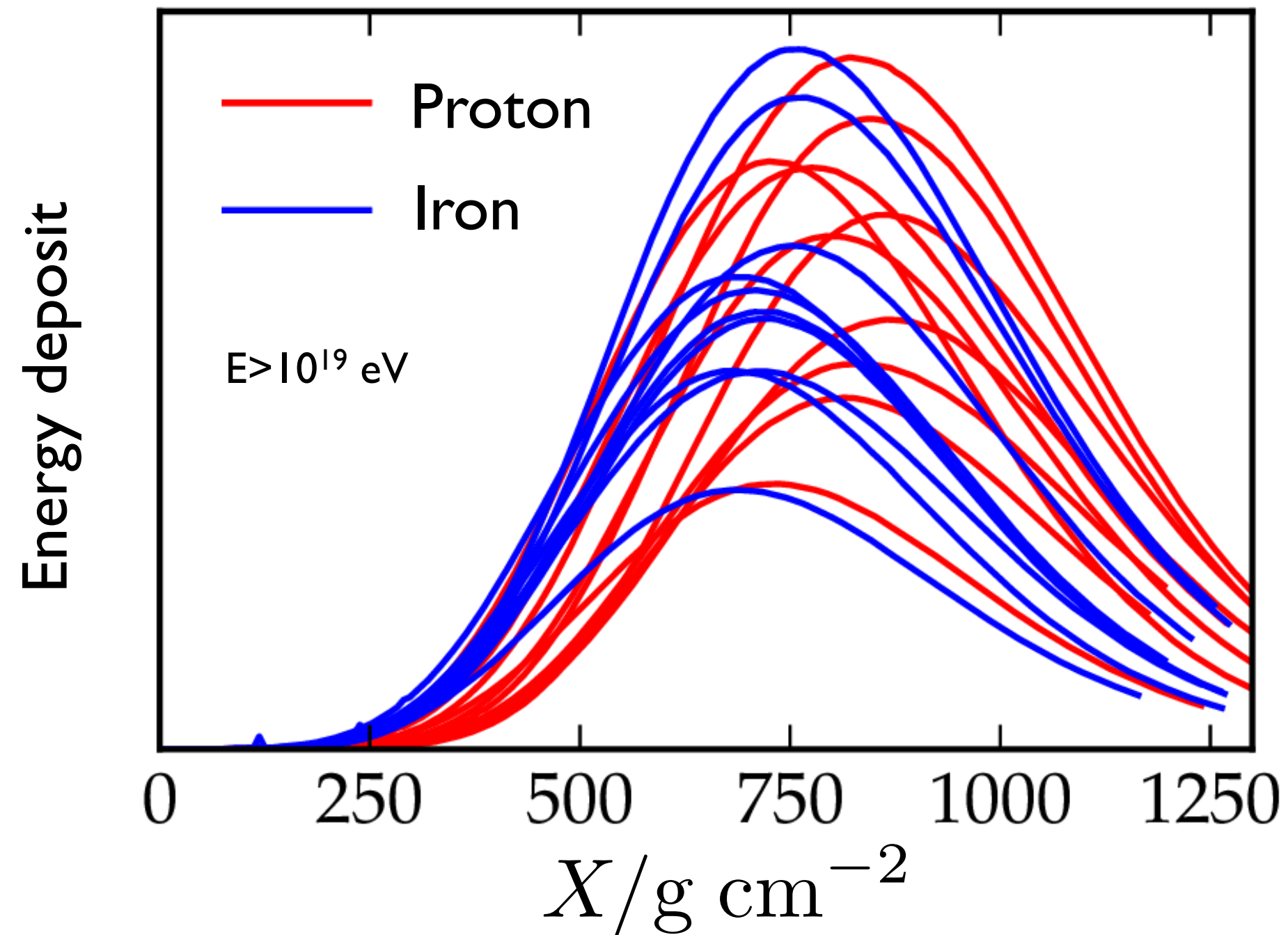


- **Particle density:** 1D lateral distribution
→ shower size, core position
- **Arrival times:** Spherical shower front
→ shower axis
- **Fit parameters** are connected only indirectly to mass and properties of UHE interactions, e.g. Rise-time, curvature, LDF slope



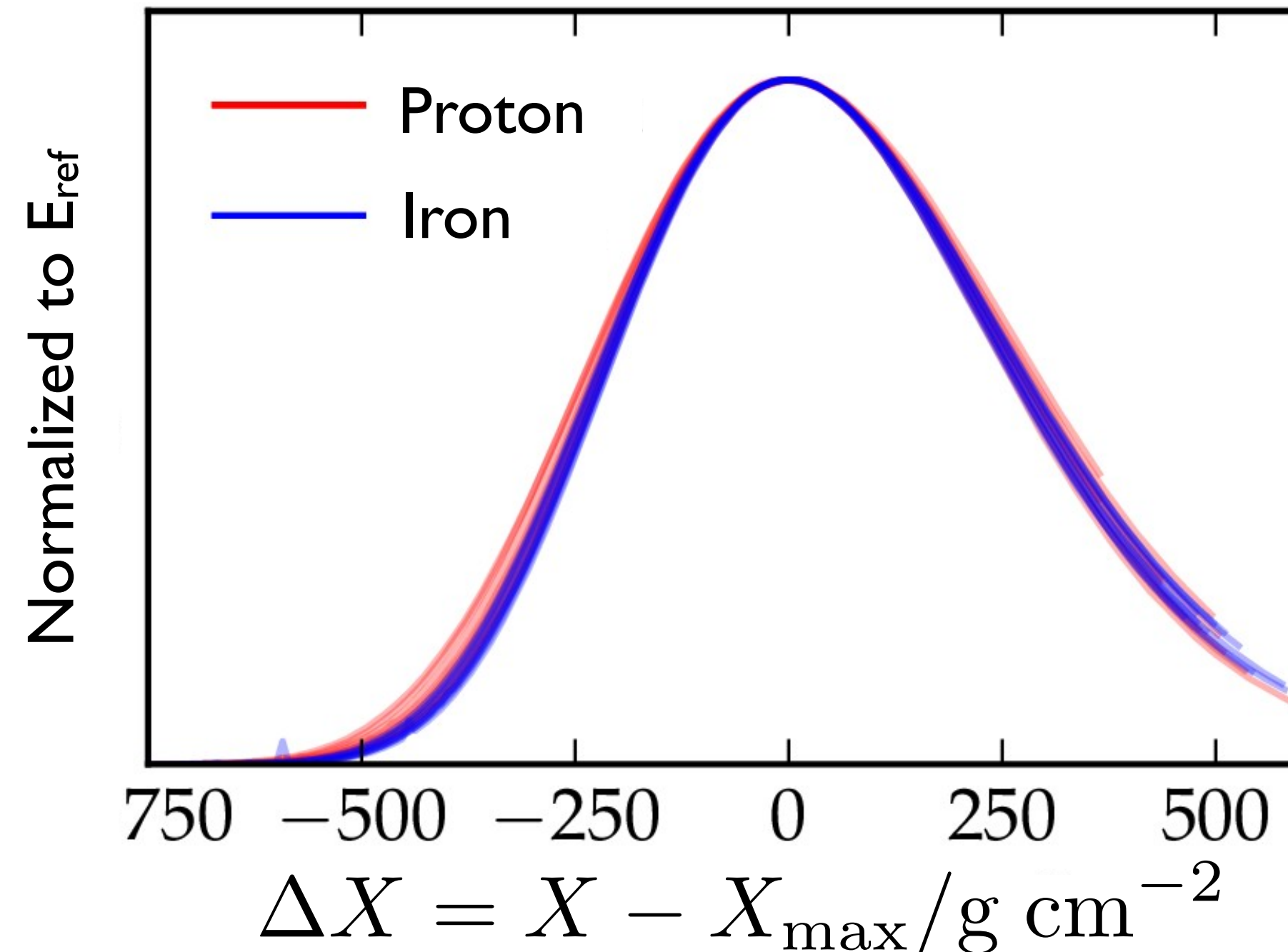
Rich timing information (40 MHz sampling) is not used in standard reconstruction

Universal longitudinal profiles of the electromagnetic component of hadron induced showers



Simulated **energy deposit** along shower axis

Large fluctuations in shower development apparent



Universal longitudinal profile

Longitudinal profiles depend on

- **Energy E**
- **Distance to X_{max} : ΔX**

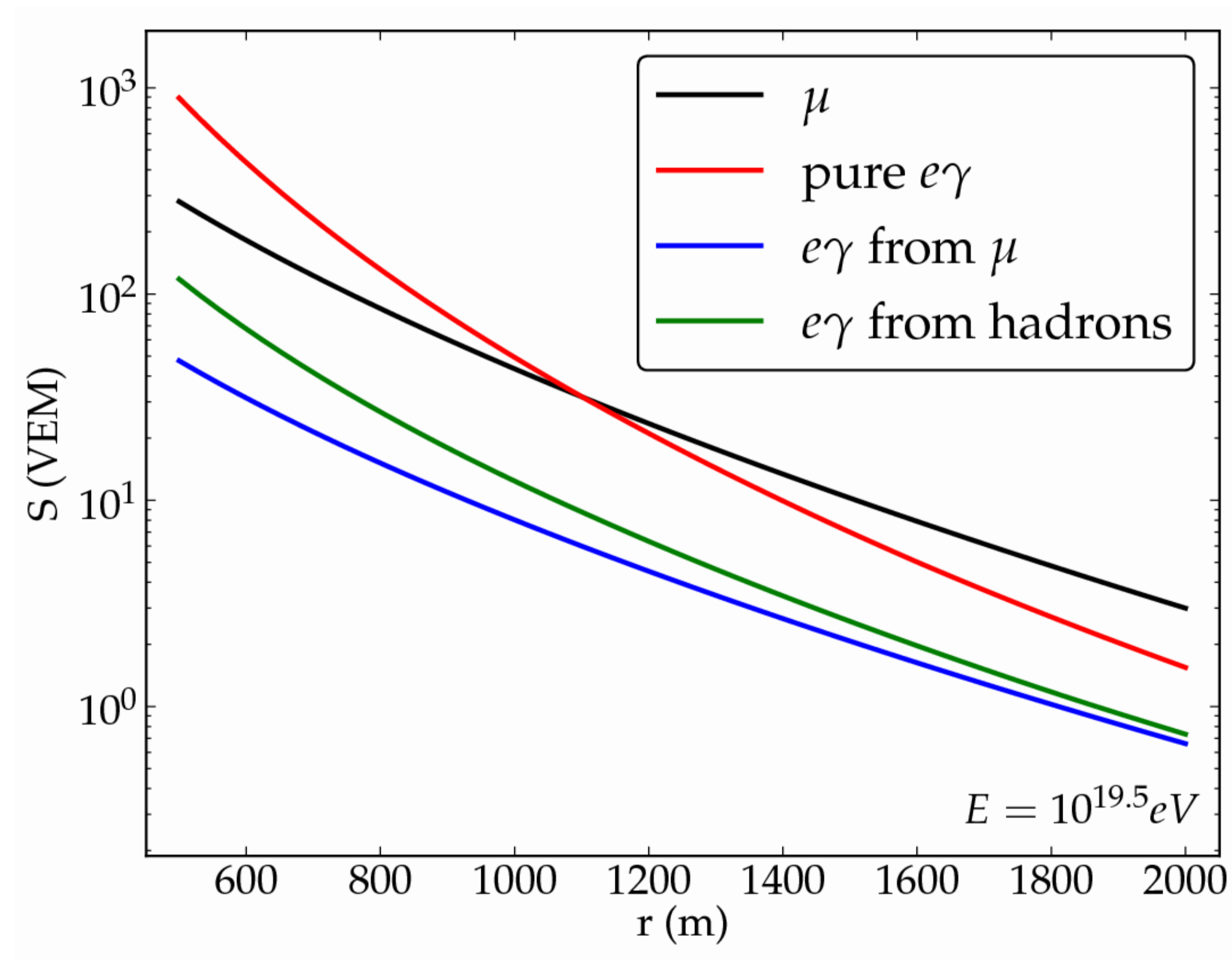
but **not on mass**

M.Ave, R.A.Vazquez, E. Zas, *Astropart. Phys.* 14 (2000) 109
M. Giller et al., *J. Phys. G* 31 (2005) 947
M. Giller, H. Stojek, G. Wieczorek, *Int. J. of Modern Physics A* 20 (2005) 6821
F. Nerling, J. Blümer, R. Engel, M. Risse, *Astropart. Phys.* 24 (2006) 421
D. Góra, R. Engel, D. Heck, et al., *Astropart. Phys.* 24 (2006) 484
P. Billoir, C. Roucelle, J.-C. Hamilton, *astro-ph/0701583*
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P. Lipari, *Phys. Rev. D* 79 (2008) 063001
S. Lafebvre, et al., *Astropart. Phys.* 31 (2009) 243
M.Ave, J. Gonzalez, D. Maurel, M.Roth ICRC 2011 Beijing, 1025

Extension to Lateral distributions of four signal components

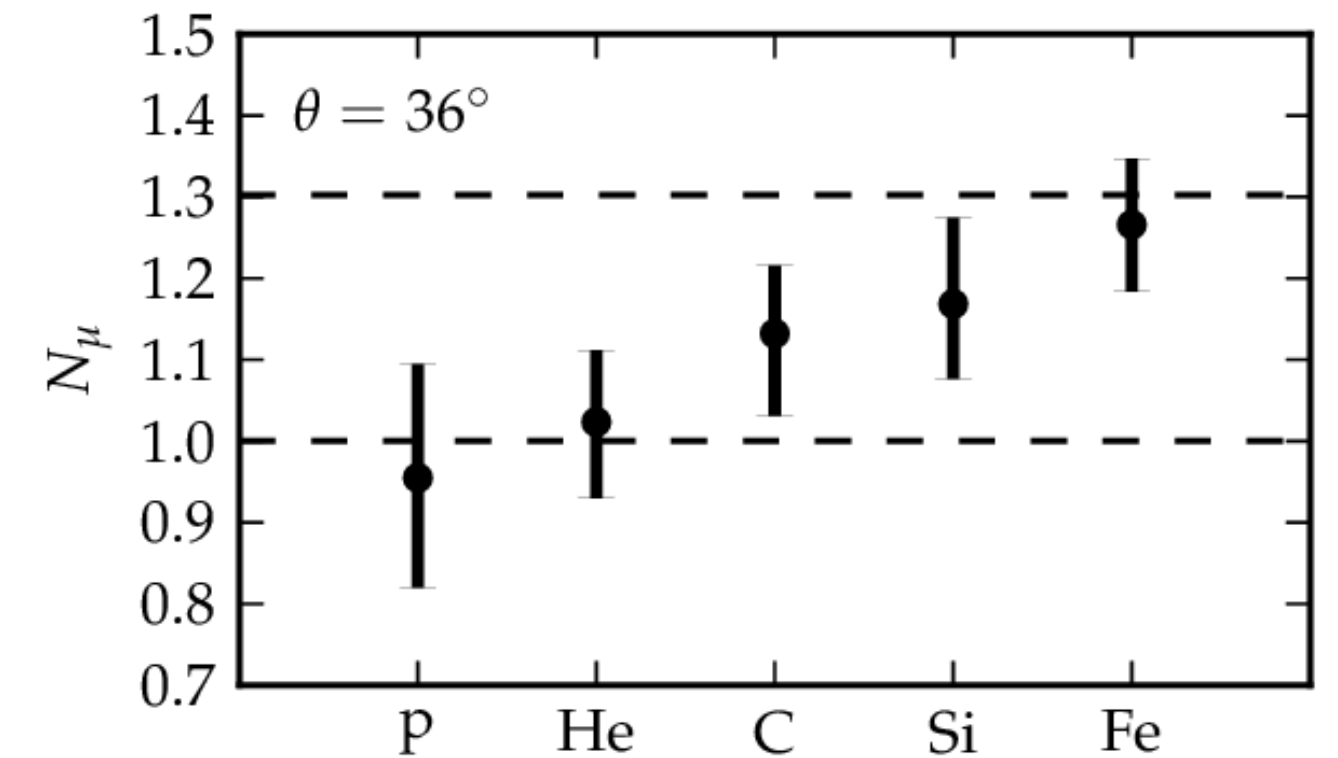
- Pure electromagnetic
- Muons
- Electromagnetic from muon decay
- Electromagnetic jets from hadrons

$$S_{\text{tot}} = S_{\text{em}}(DX, E) + N_{\mu}^{\text{rel}} S_{\mu}^{\text{QGS}}(DX, E) + N_{\mu}^{\text{rel}} S_{\text{em}}^{\mu}(DX, E) + (N_{\mu}^{\text{rel}})^{\alpha} S_{\text{em}}^{\text{jet}}(DX, E)$$

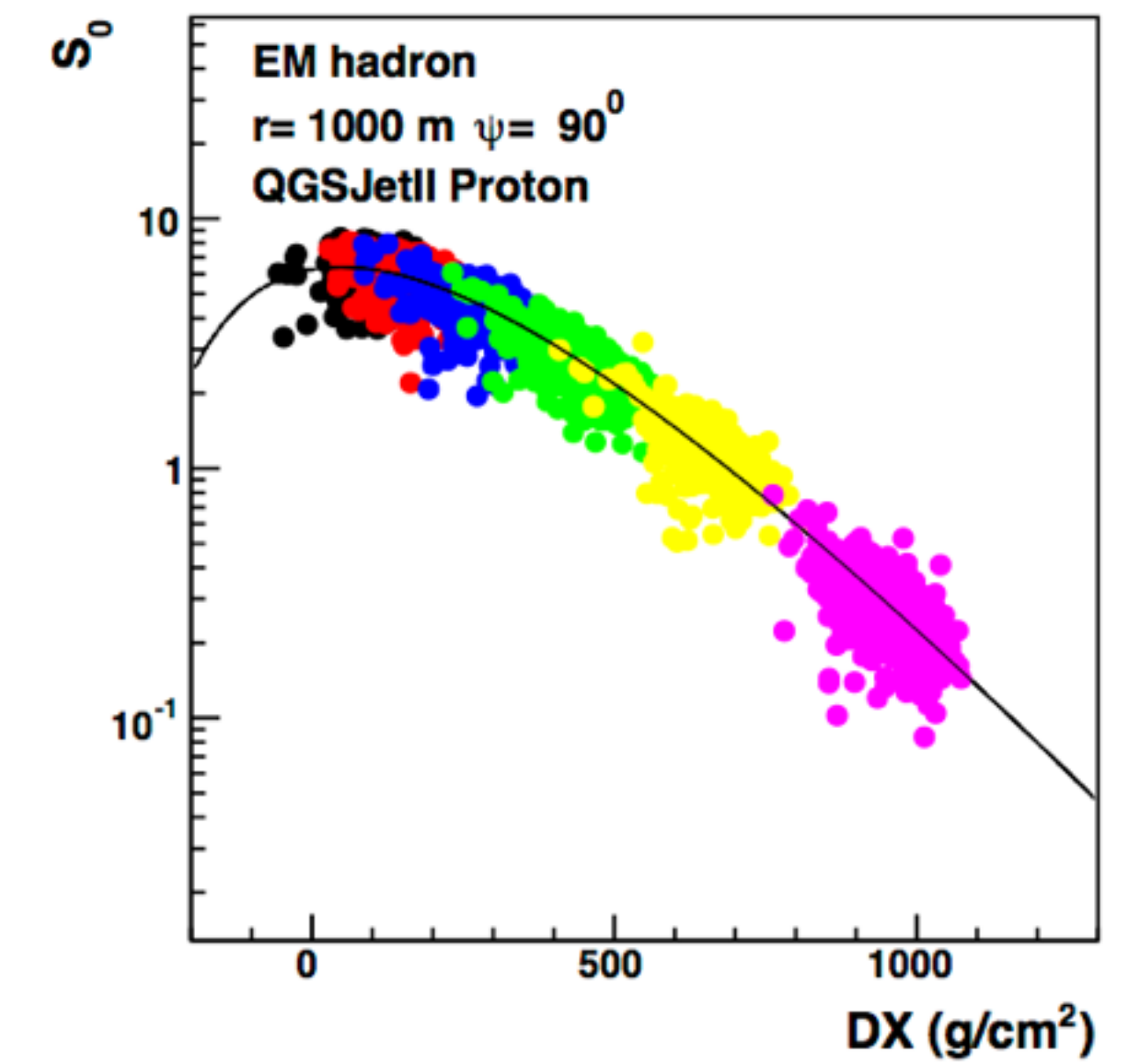
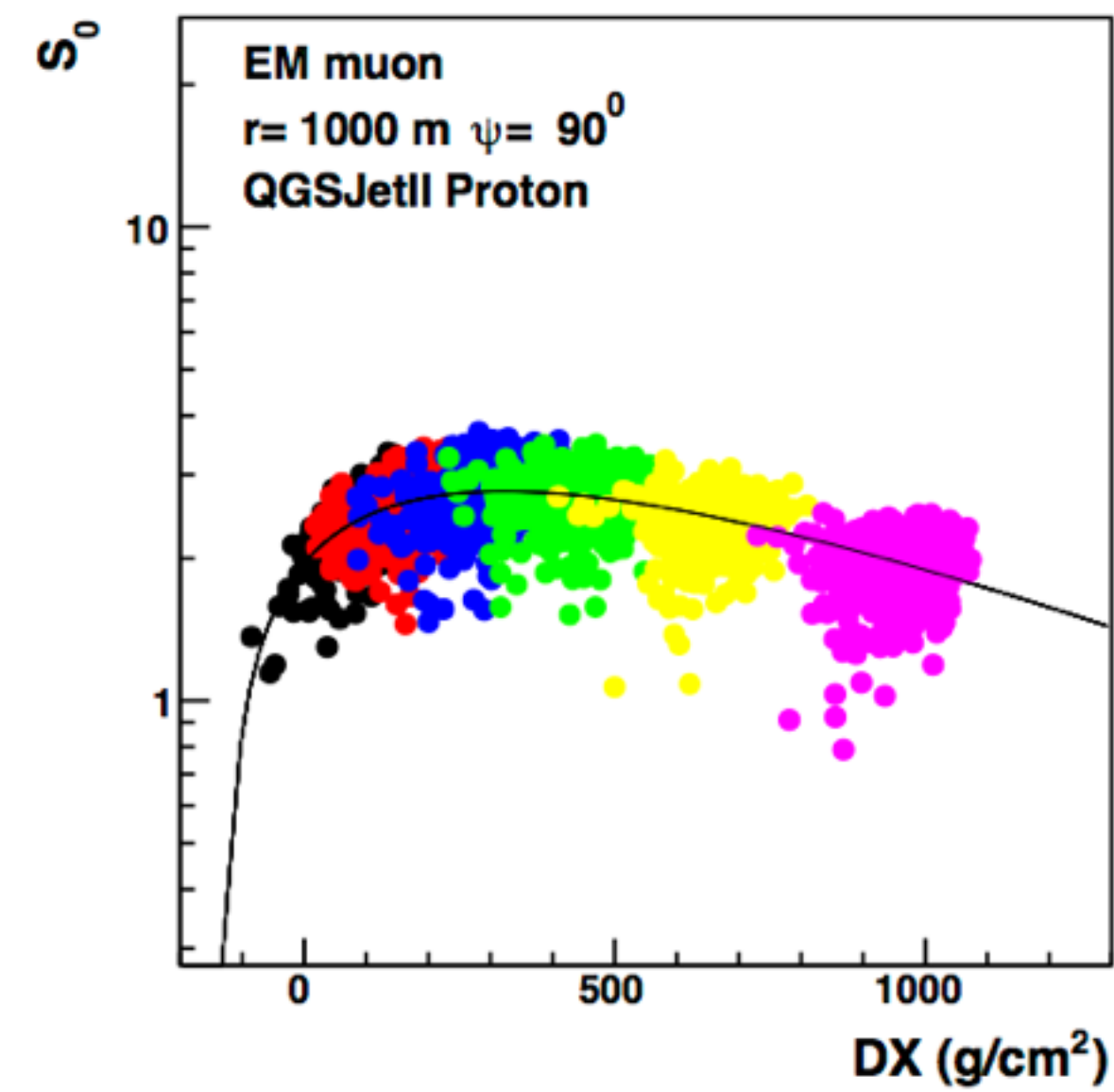
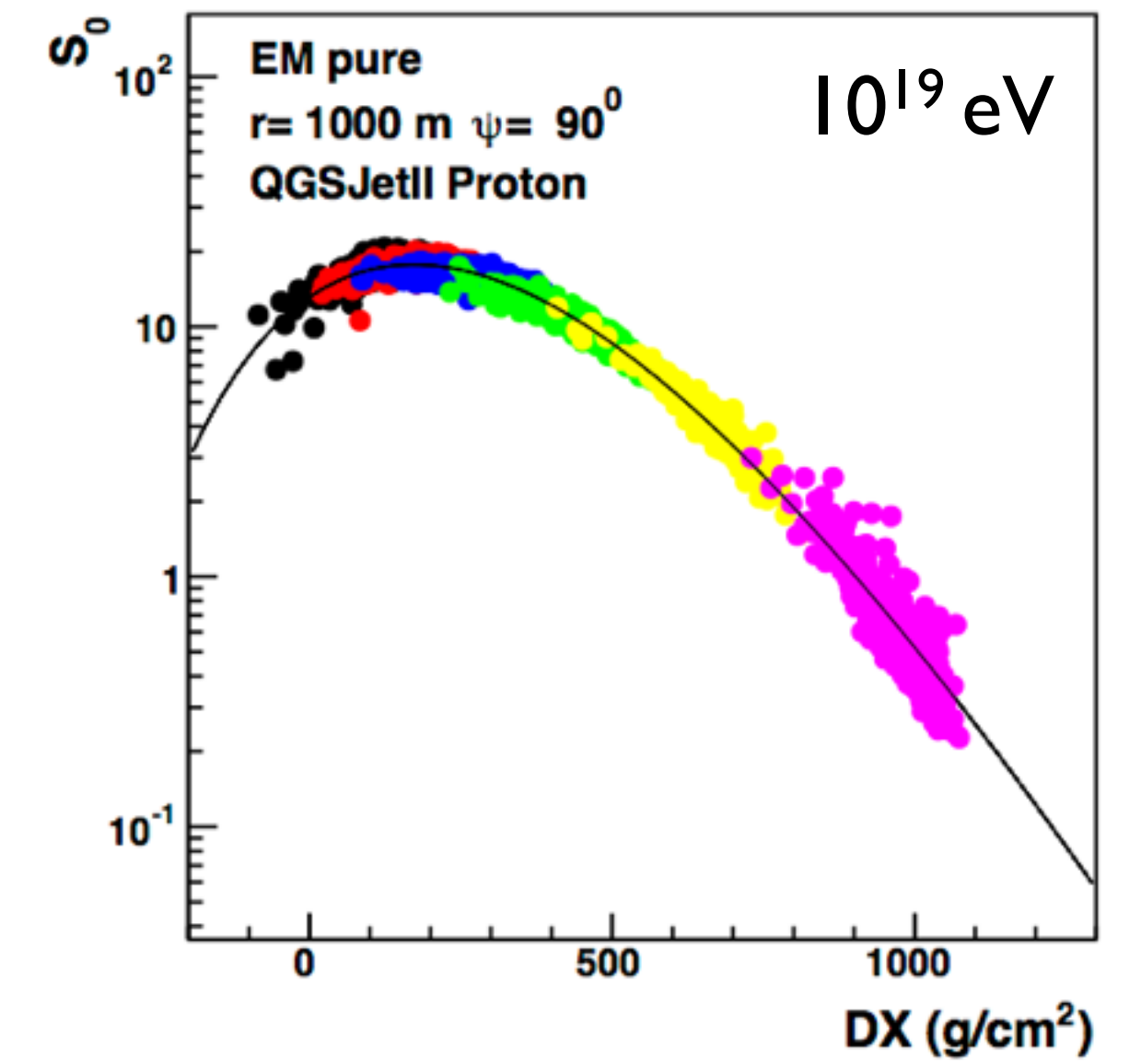
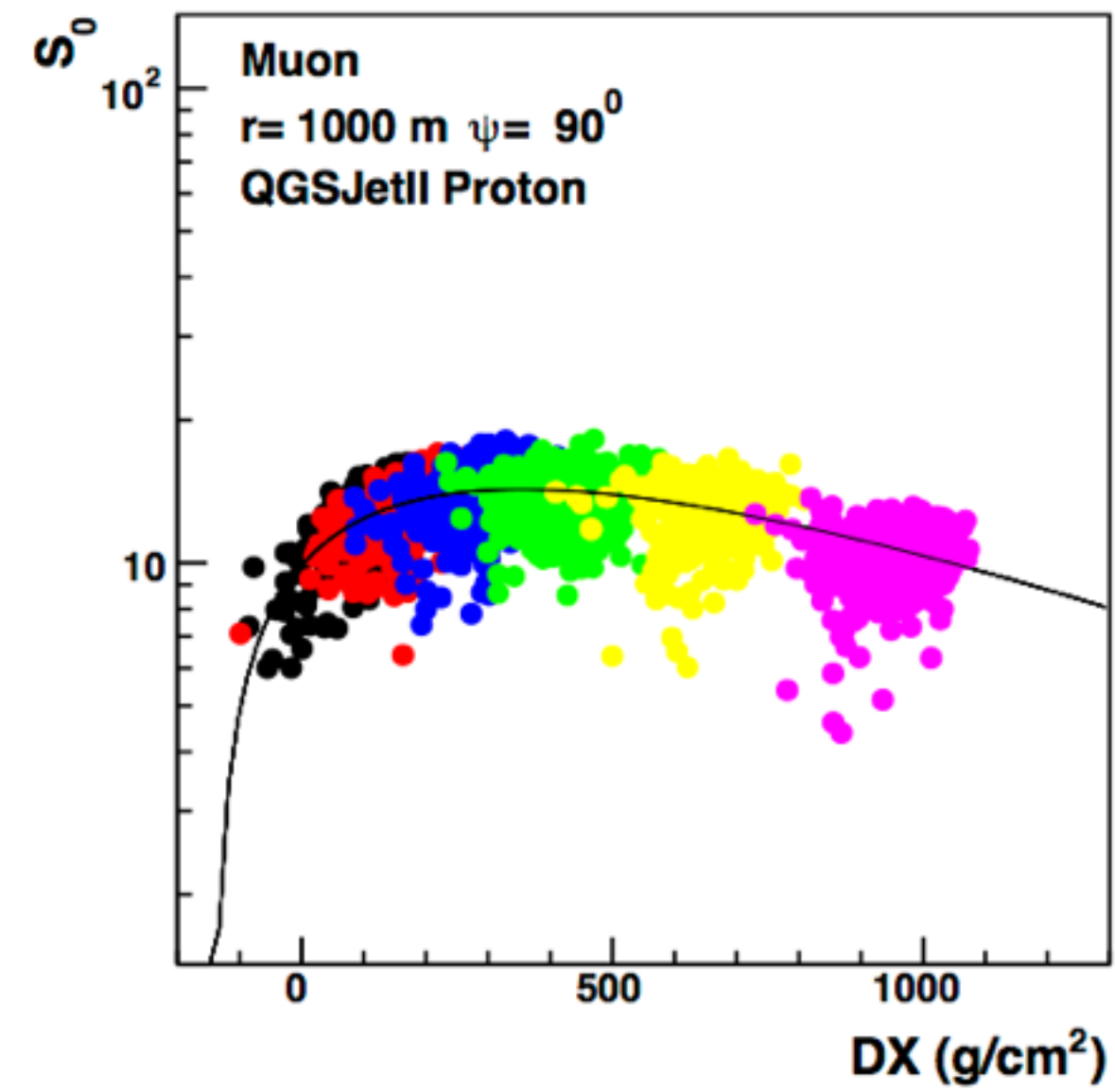
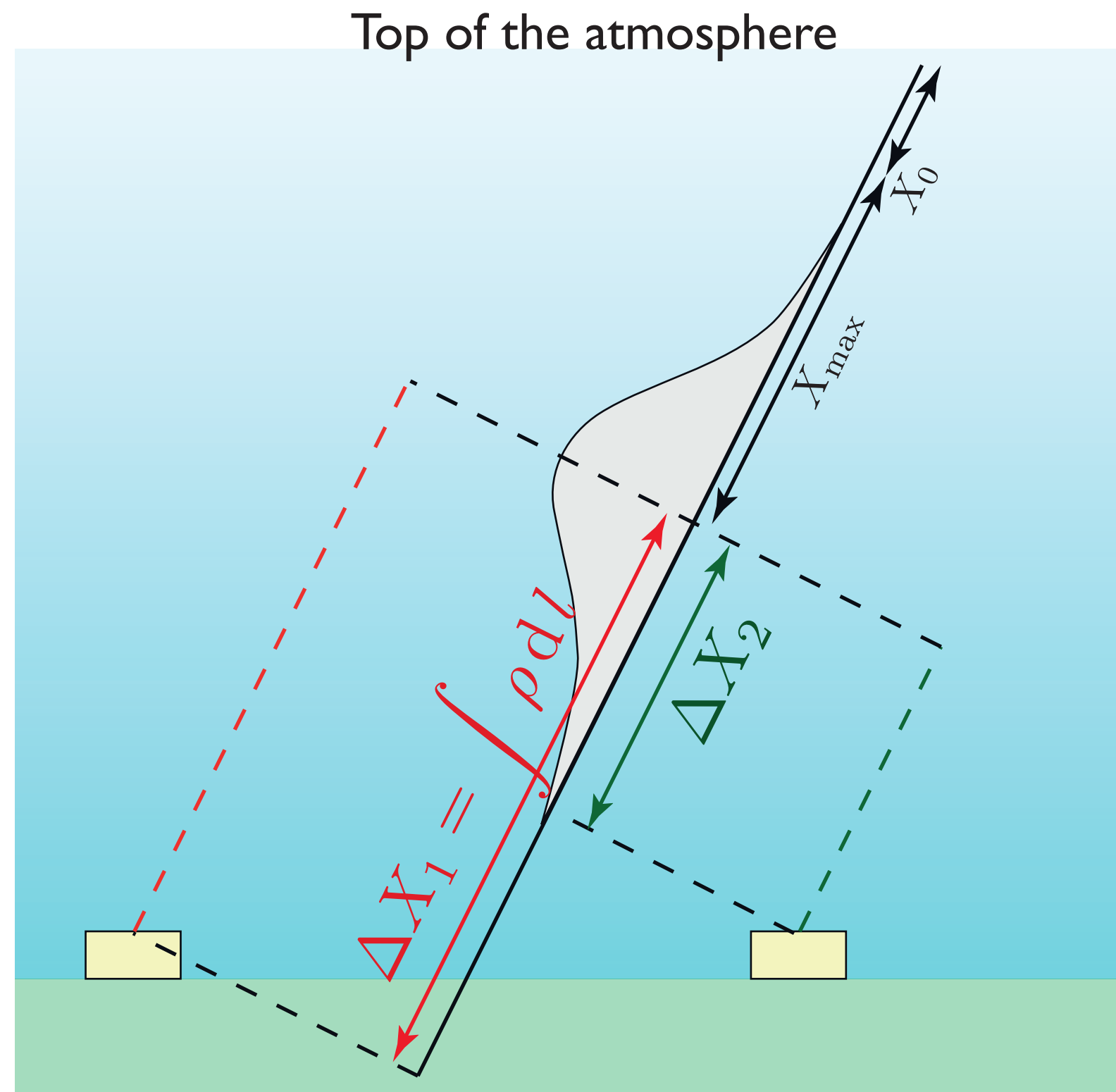


$$N_{\mu}^{\text{rel}} = \frac{S_{\mu}(DX, r, E)}{S_{\mu}^{\text{QGS}}(DX, r, E)}$$

determined by lateral distribution and thus correlated to S₁₀₀₀

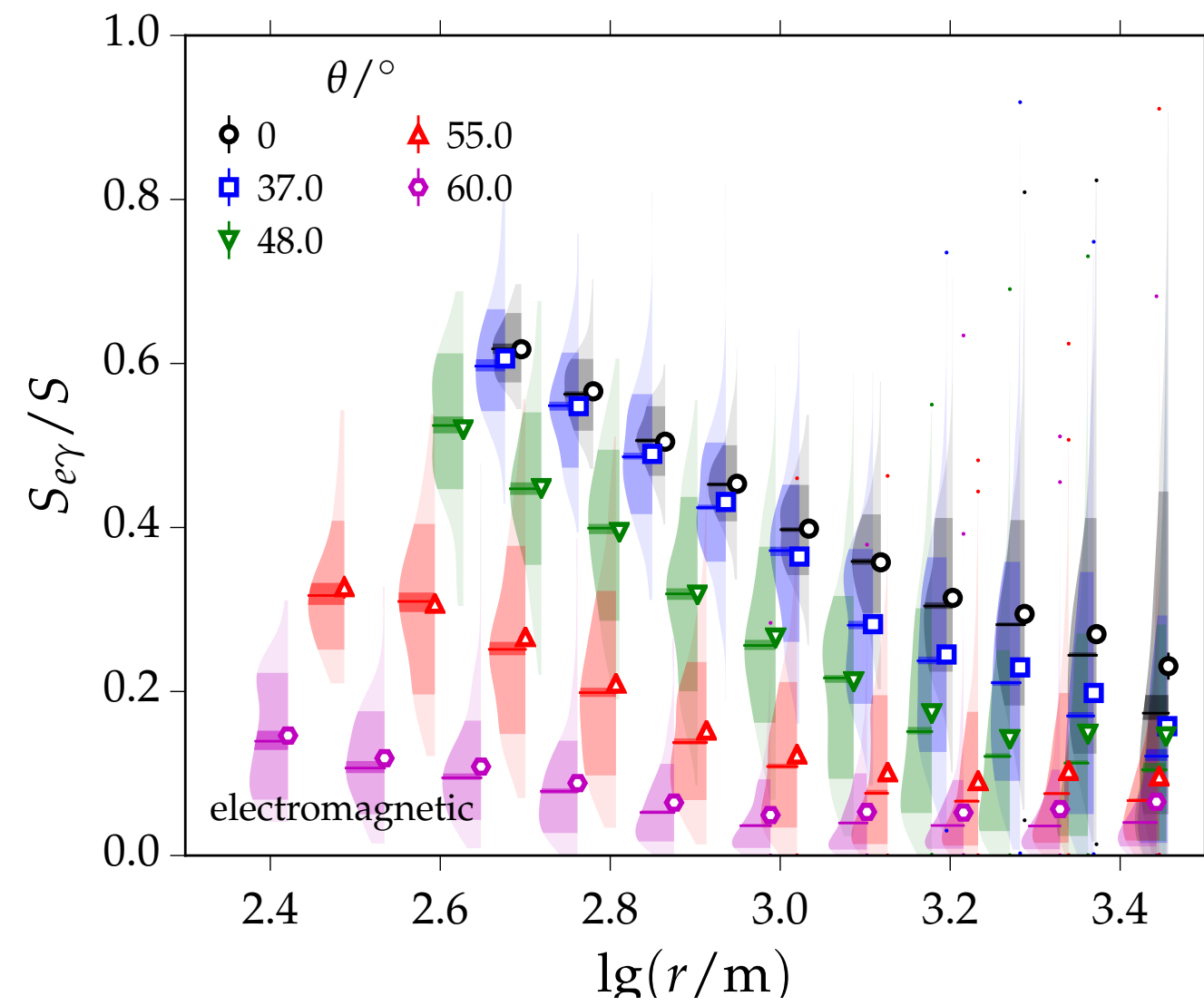


Signal contribution at 1000m

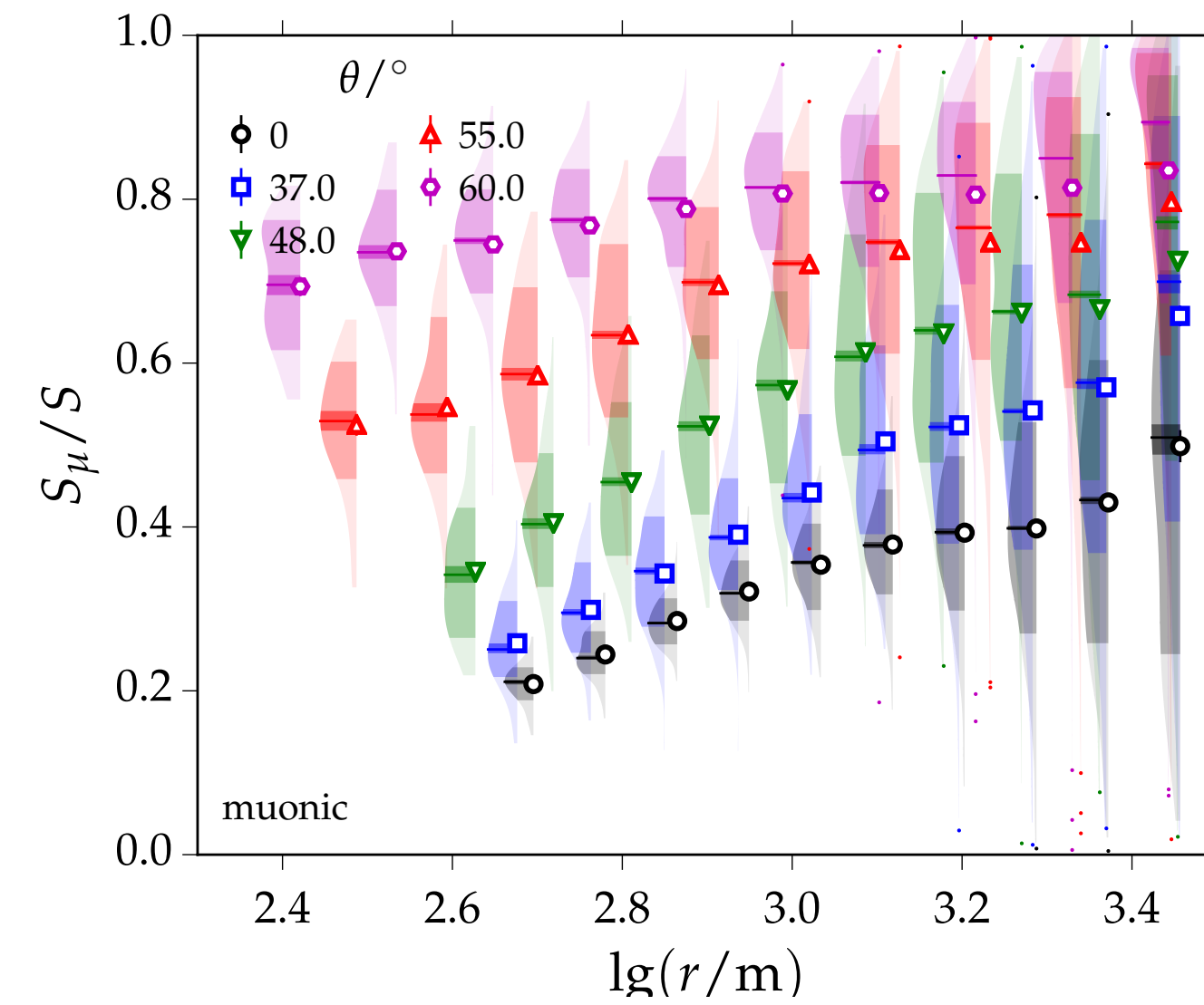


Relative signal contribution of individual components

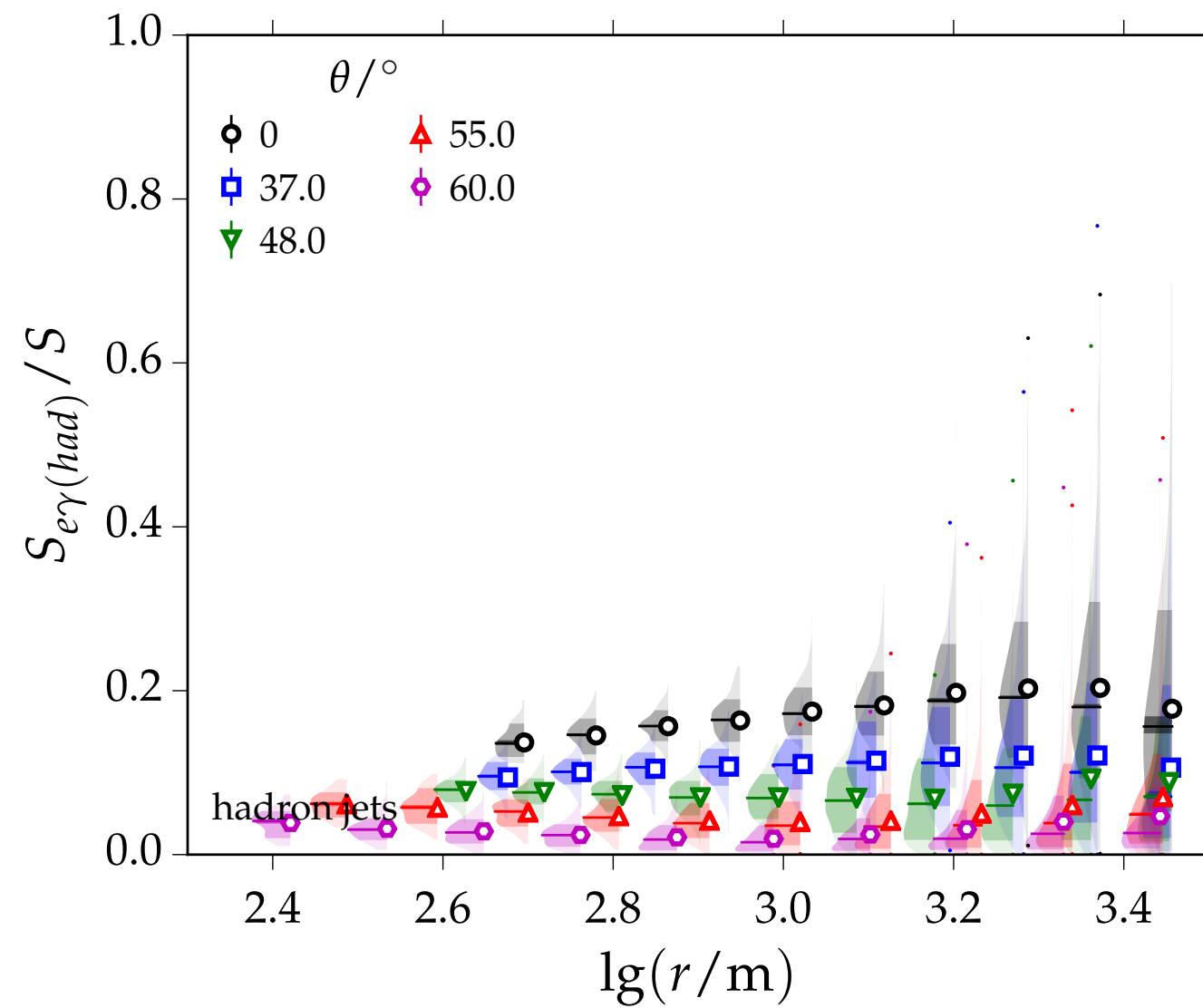
e/ γ



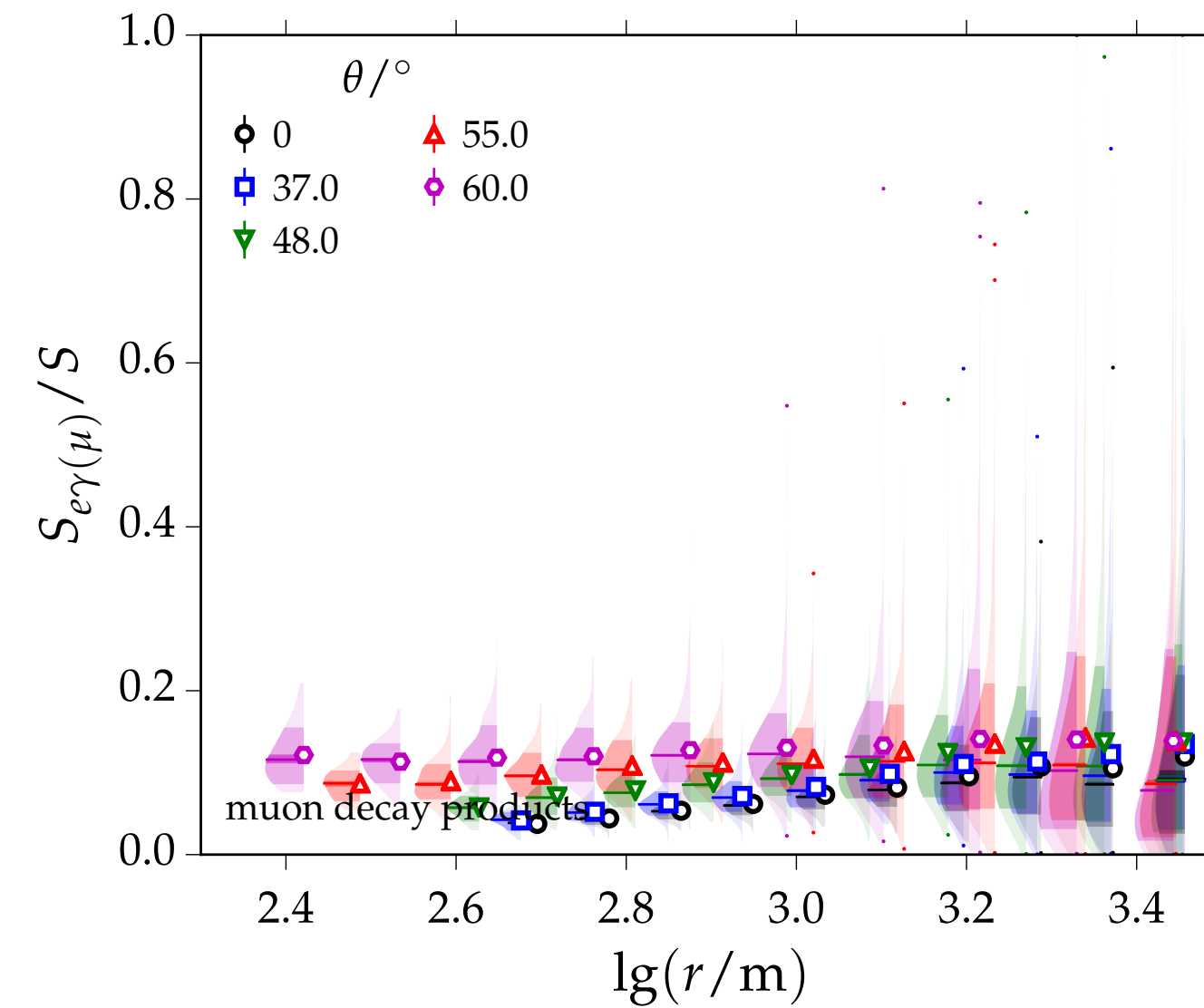
μ component



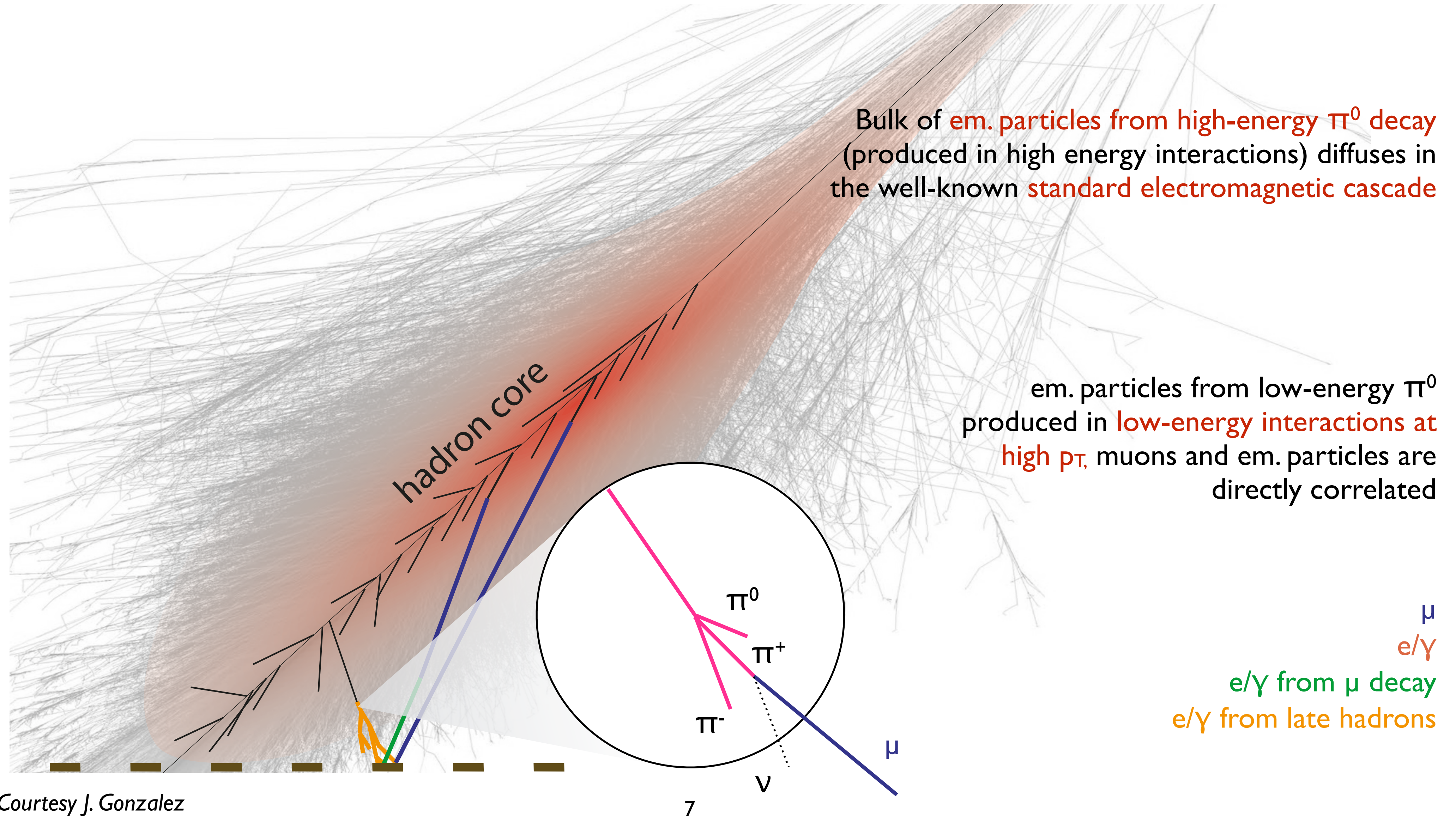
hadron jet



μ decay products



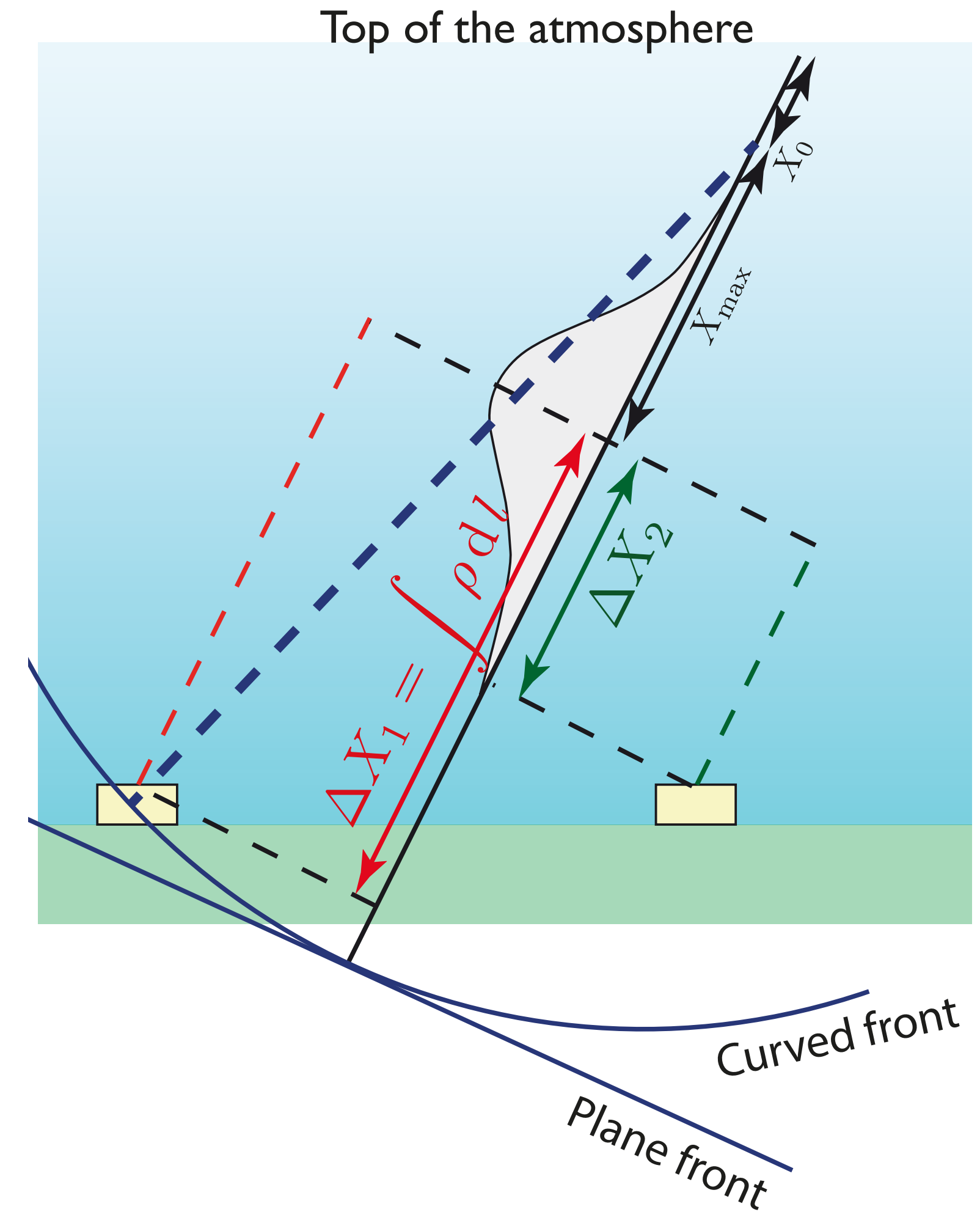
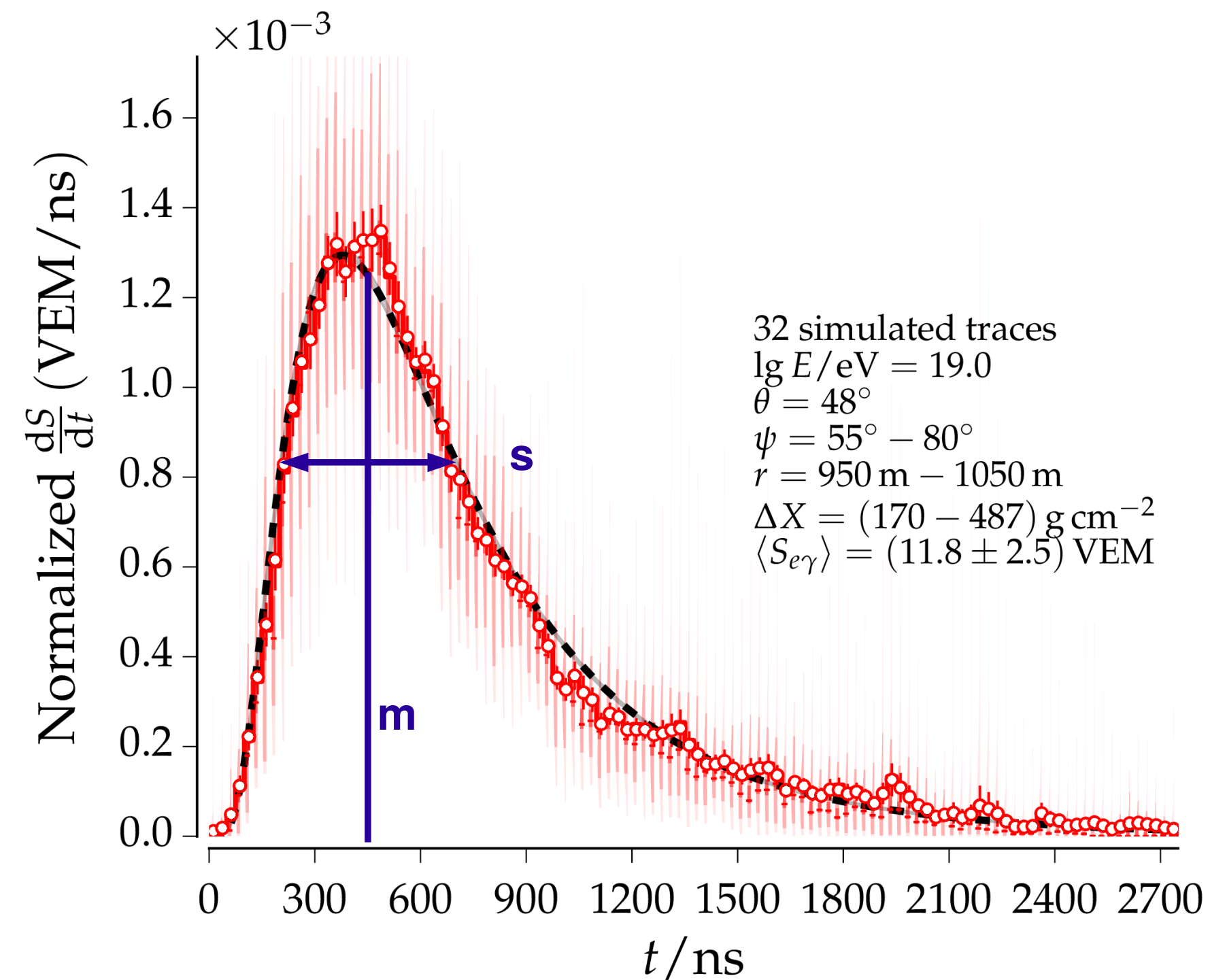
Four-component universality model



Parameterization of the time dependence

- Time-dependent detector response modeled by **log-normal distribution** (for each signal component)
- **Start time t_0** given by first interaction X_0 (spherical front)
- **Time shape (mean/spread)** depends on ΔX and position in shower plane:
 $S(t) = f(t, t_0, \Delta X, r, \psi)$
- Time structure is universal (no dependence on N_μ)

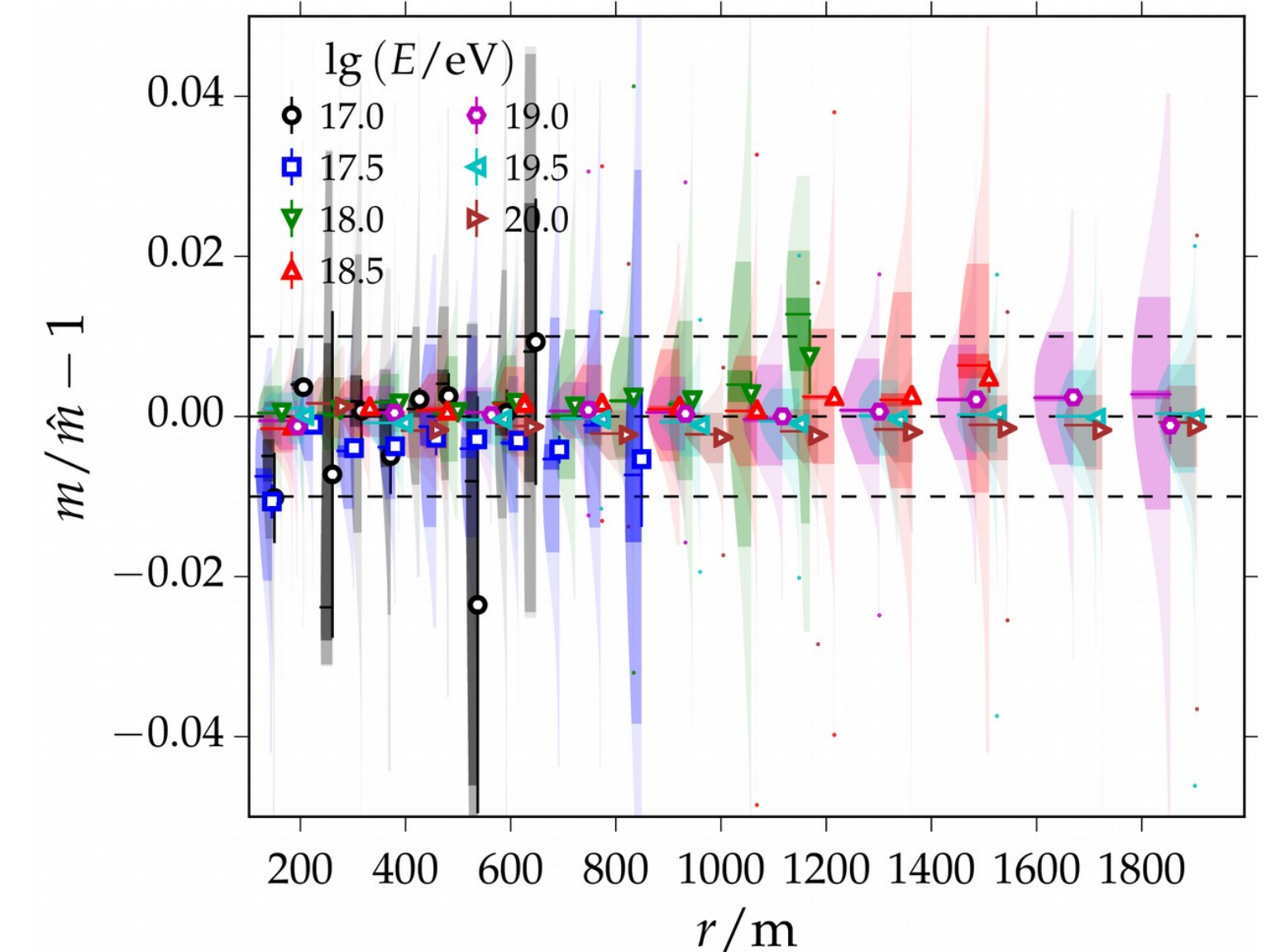
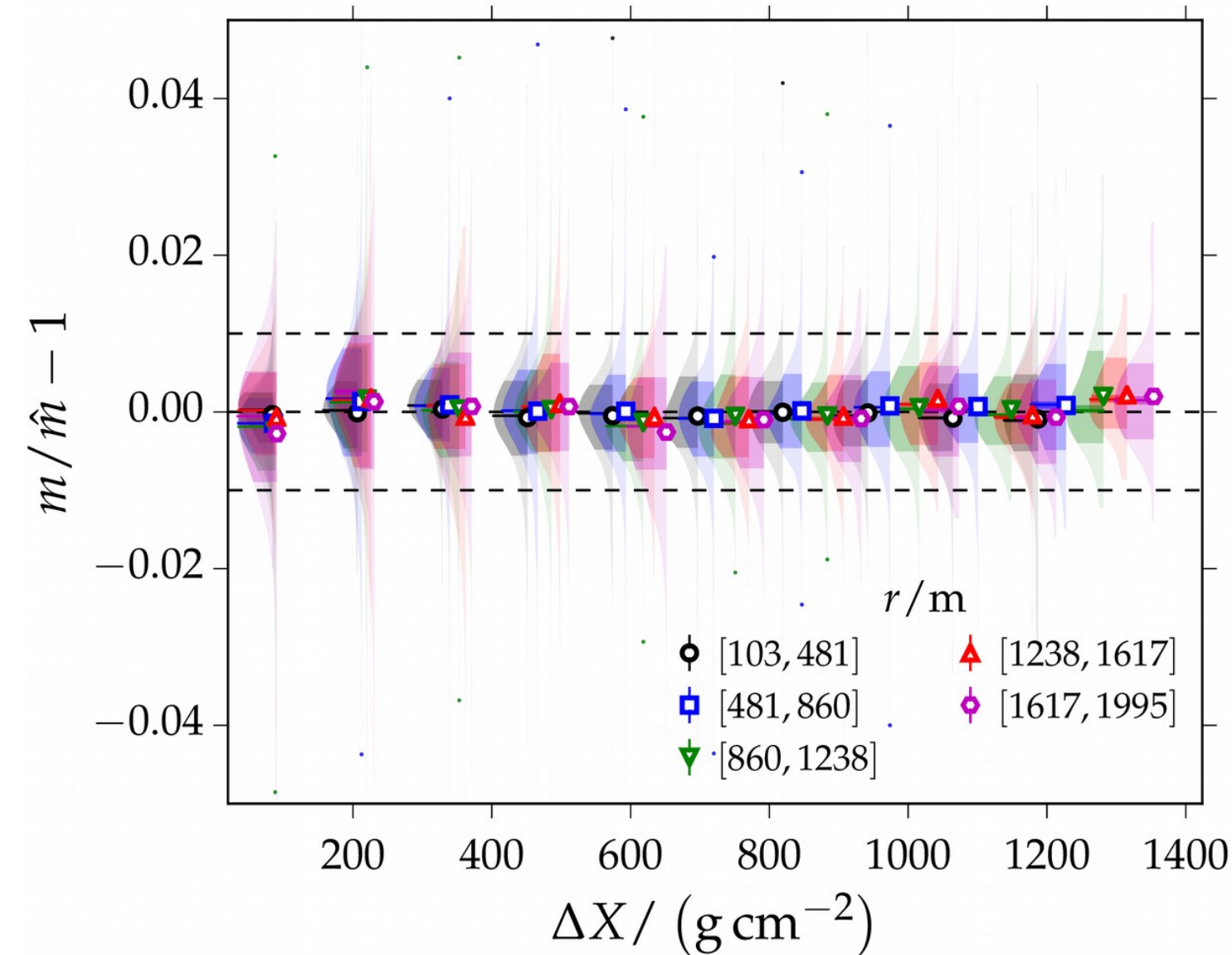
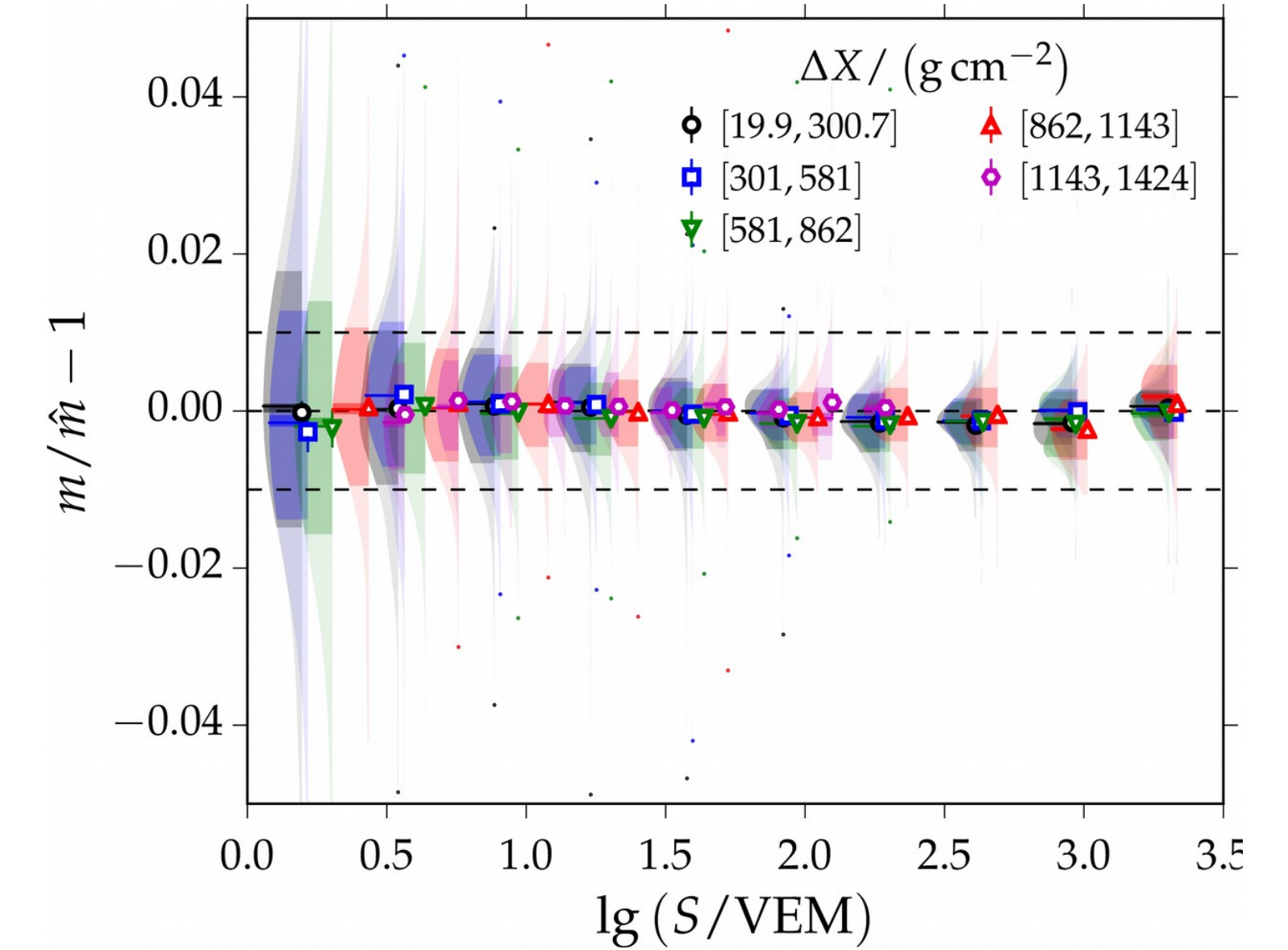
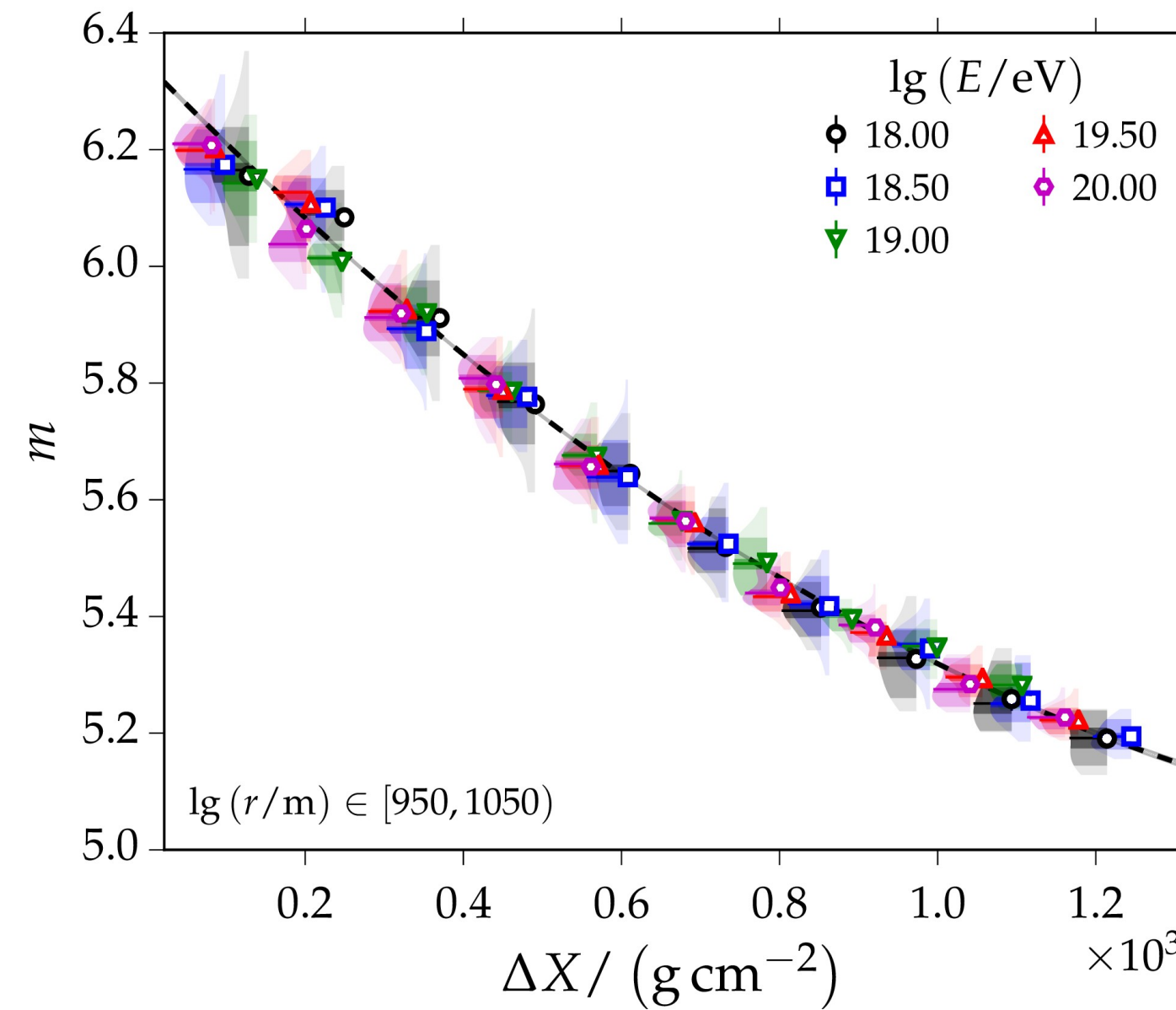
Example: e/ γ component



Parameter dependencies

Parameters m and s depend on

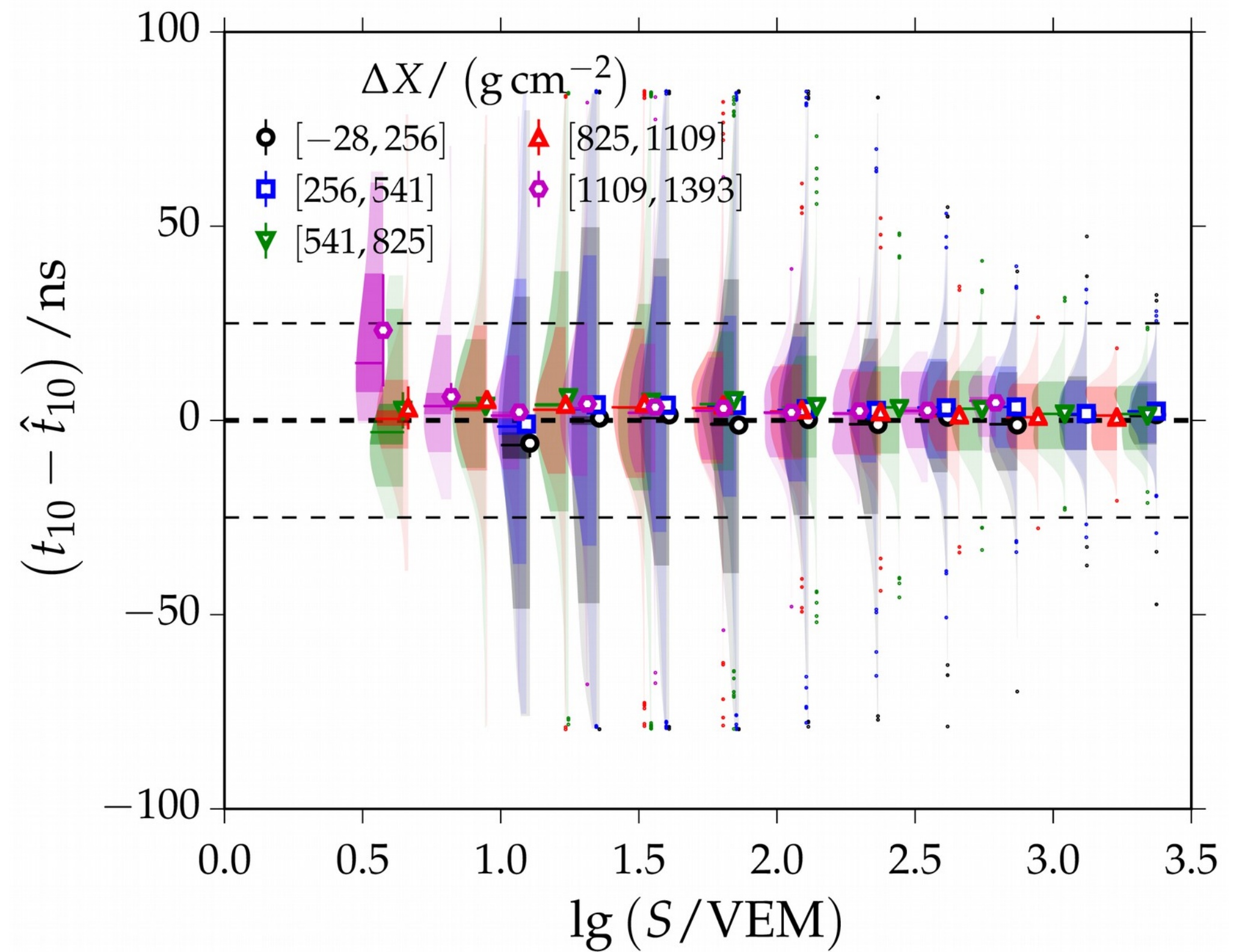
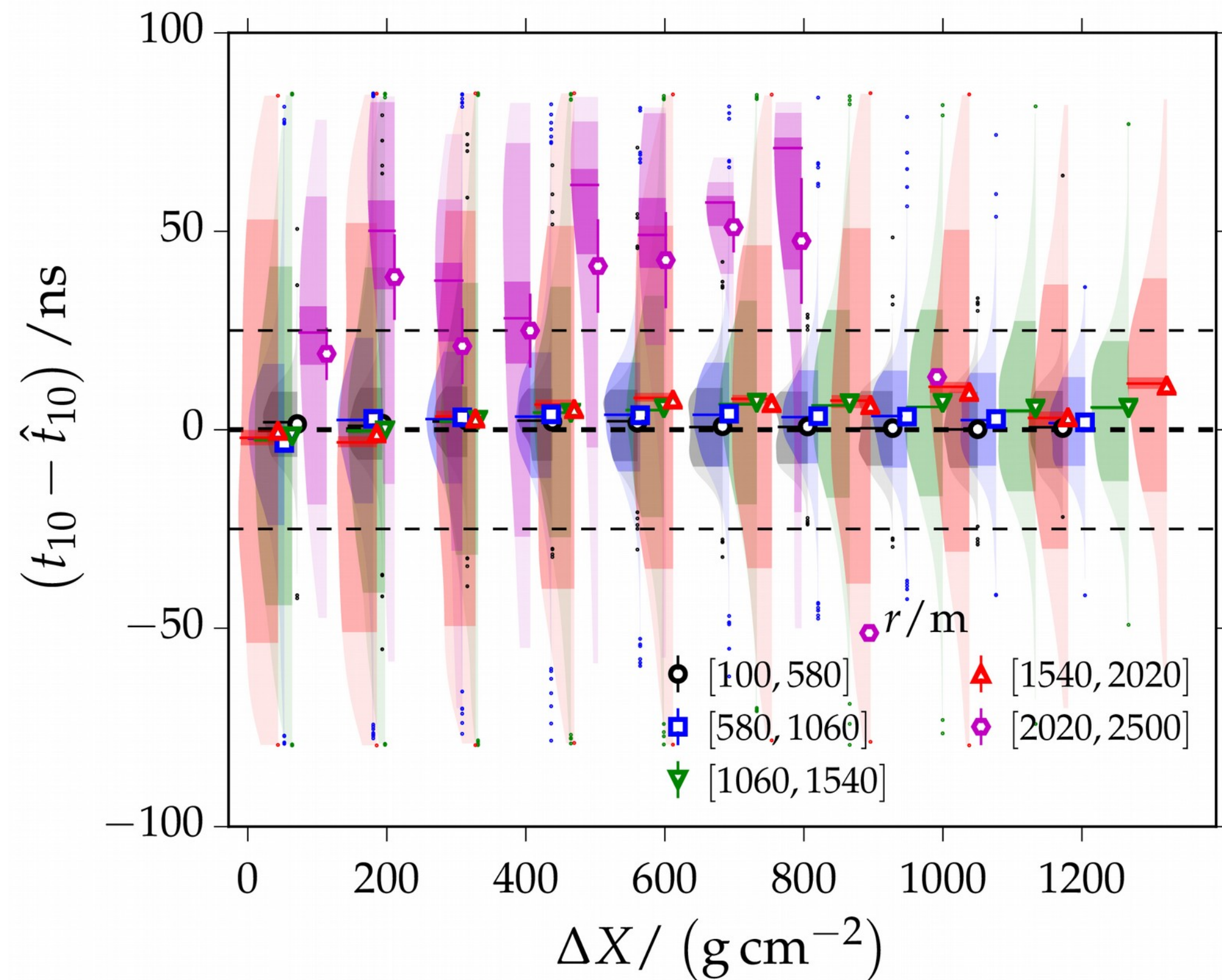
- Energy E
- Distance to X_{\max} : ΔX
- Core distance r
- Shower geometry θ, ϕ



No energy, ΔX nor core distance dependent bias apparent

Validation of the time model

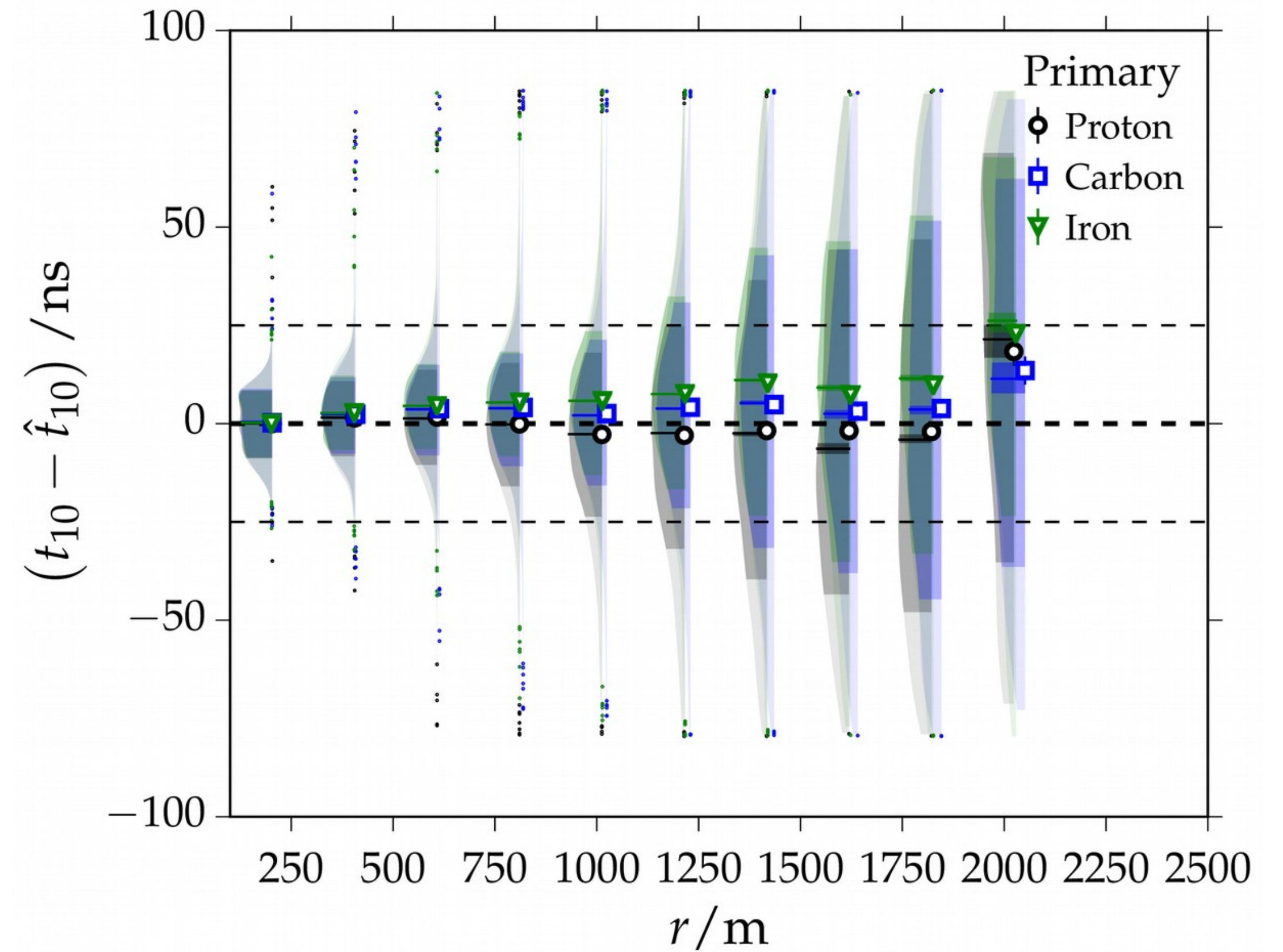
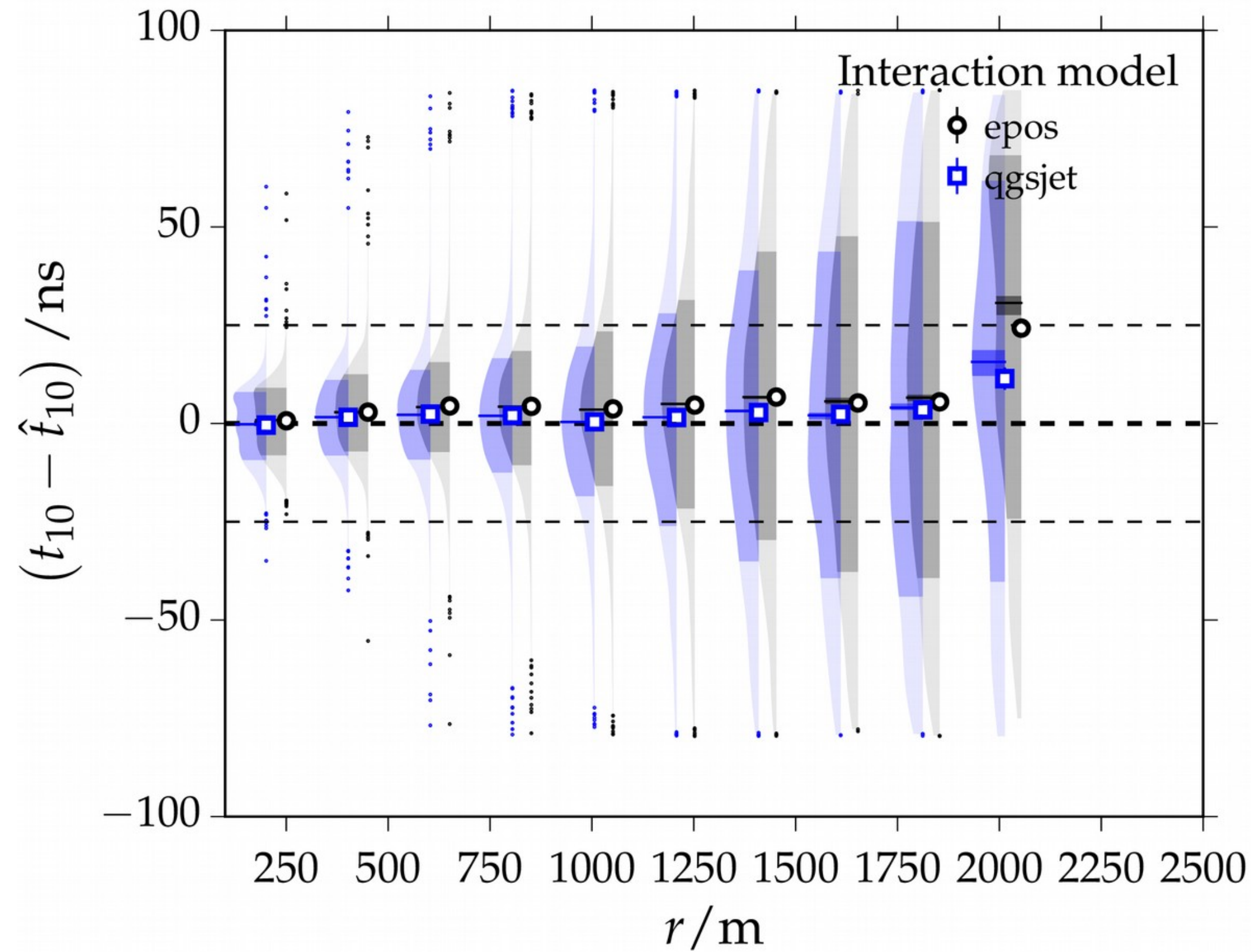
Time quantiles for muons



No bias for different ΔX nor signal size

Validation of the time model

Time quantiles for muons



No bias for different primaries nor interaction models used

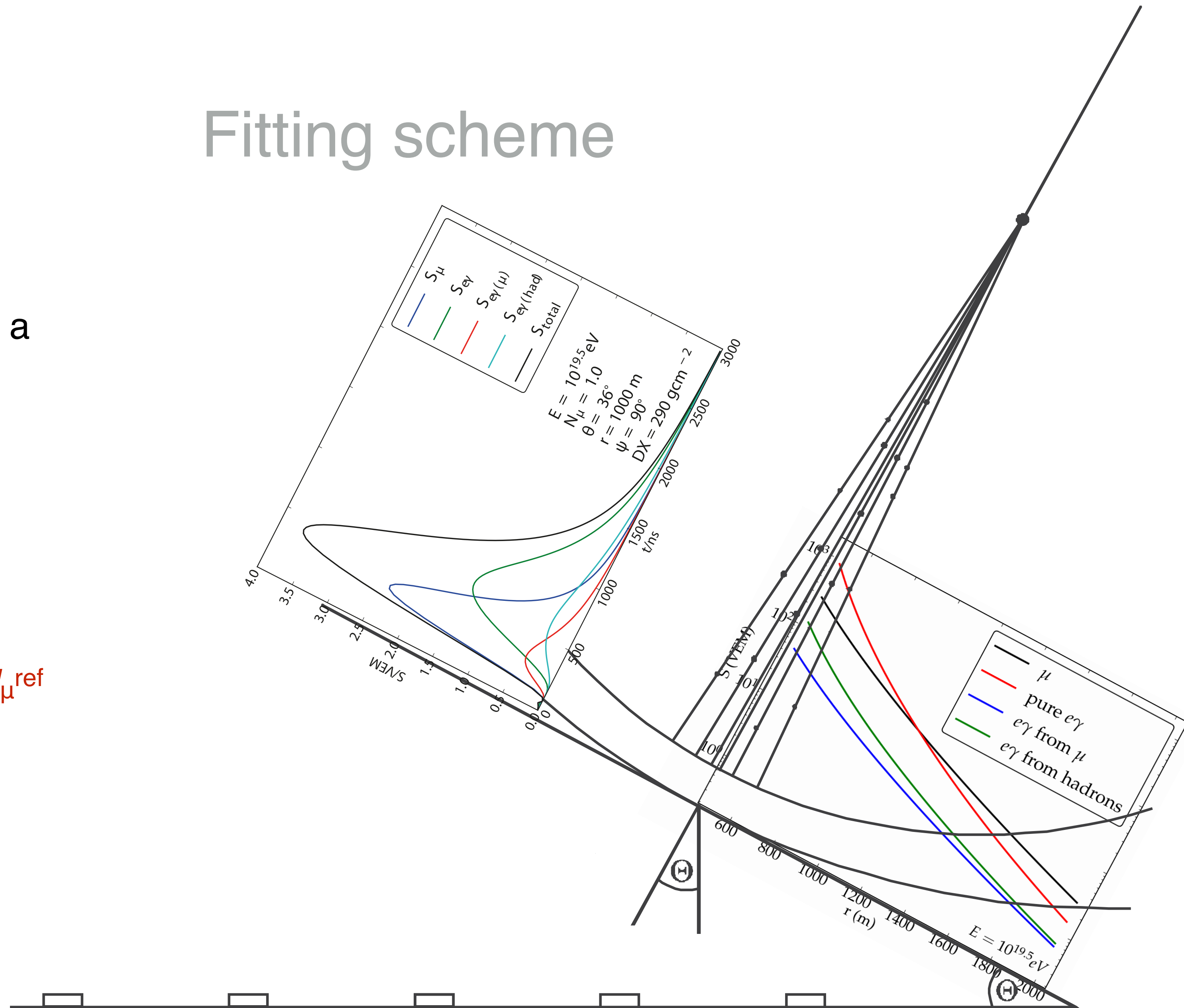
Fitting scheme

Parameterizations used in a likelihood procedure

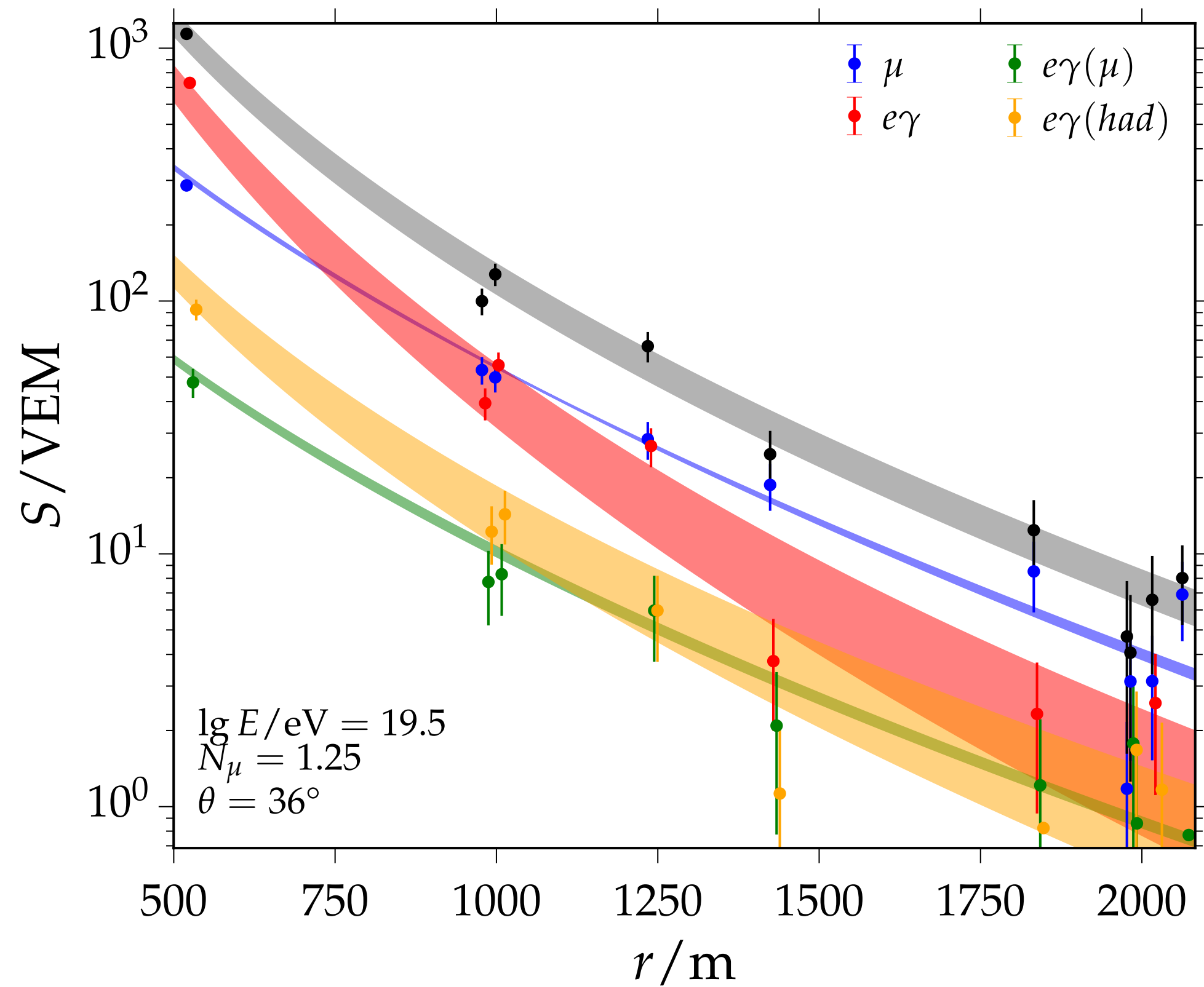
- Signal size S
- Start time t_0
- Shape m, s

as a function of

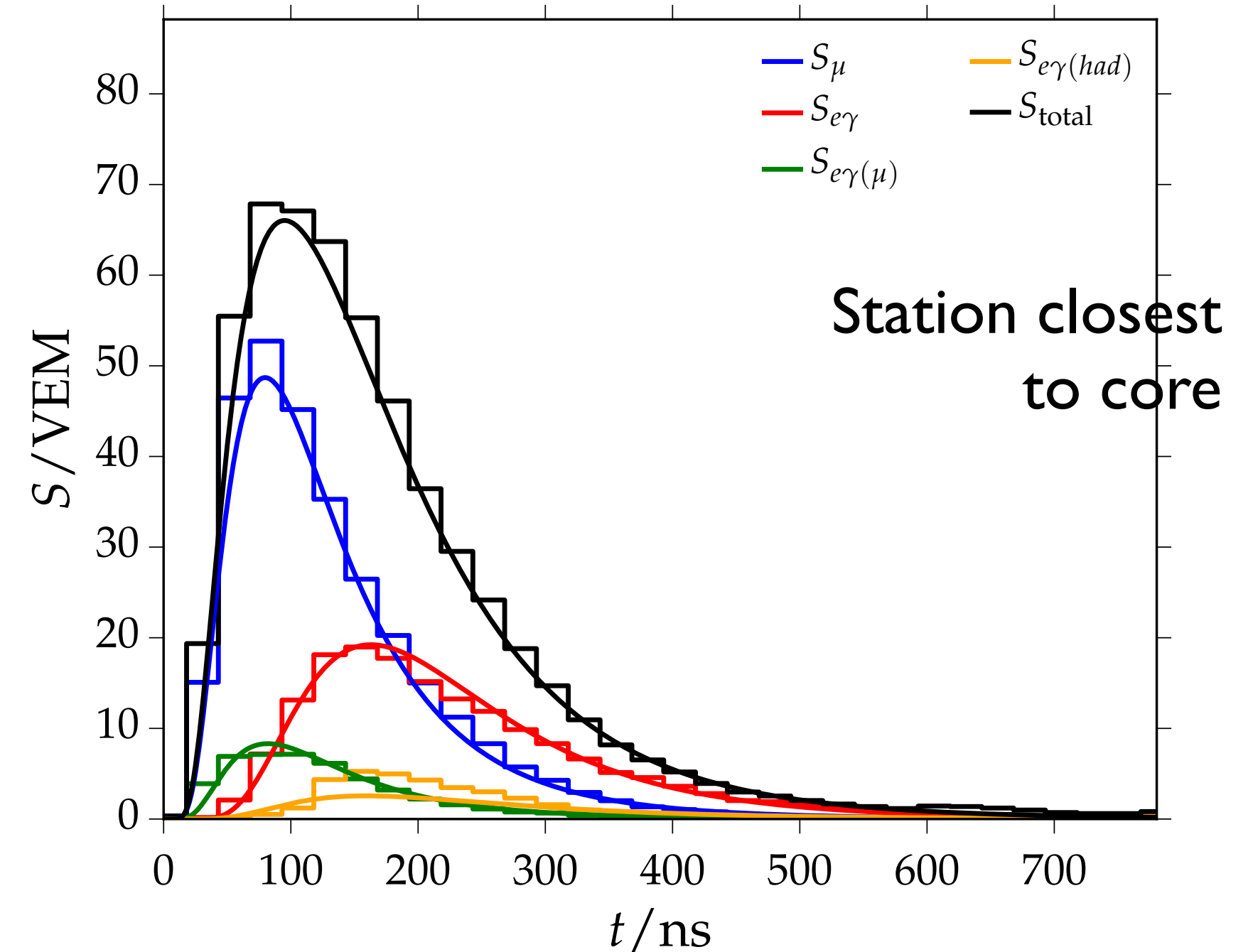
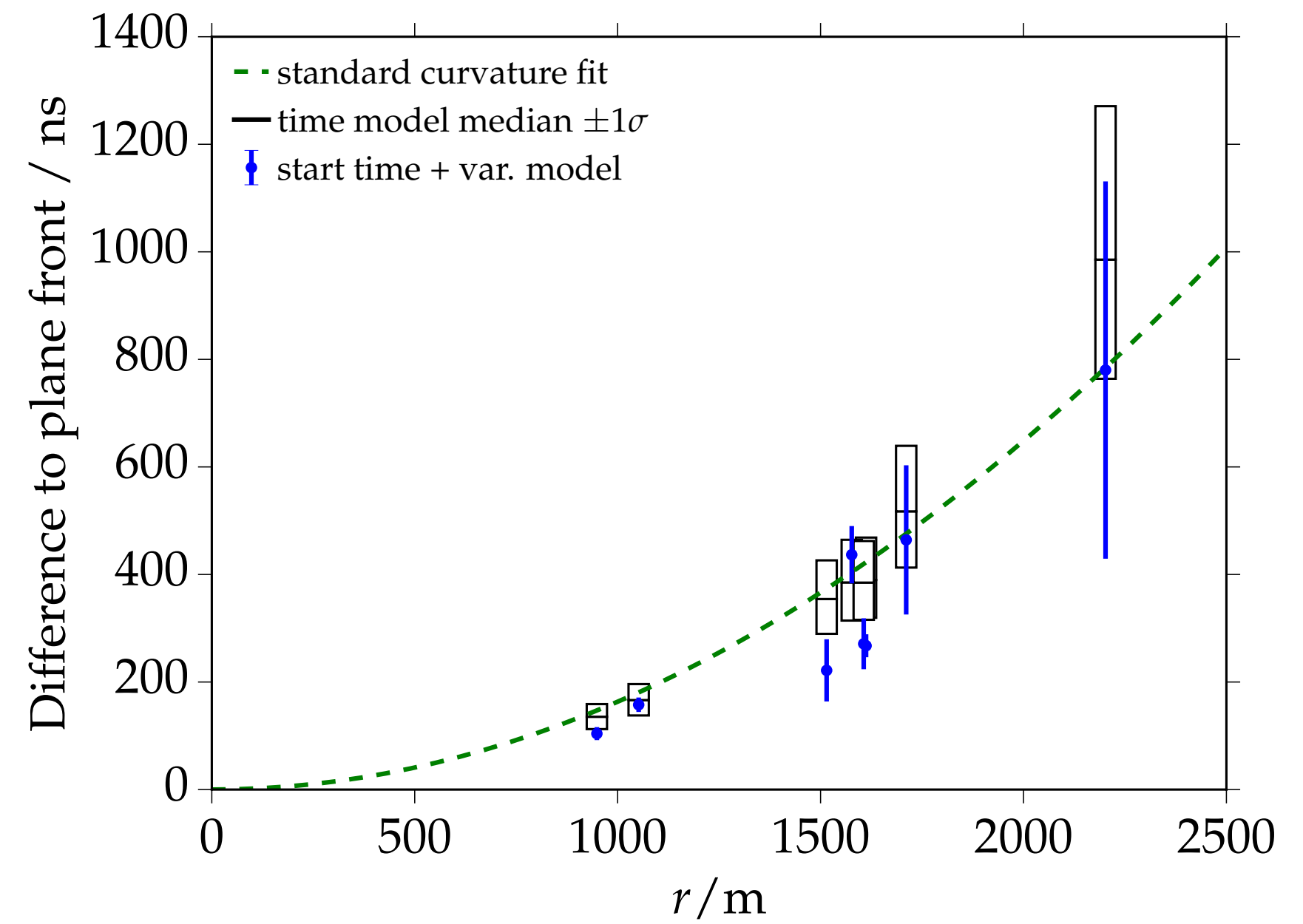
- Energy E
- Relative muon number N_{μ}^{ref}
- Distance to X_{max} : ΔX
- Shower geometry θ, ϕ



Fitting of a single event



Colored bands indicate corrections for **up- and downstream asymmetry**, e.g. different ΔX , ground screening, detector response



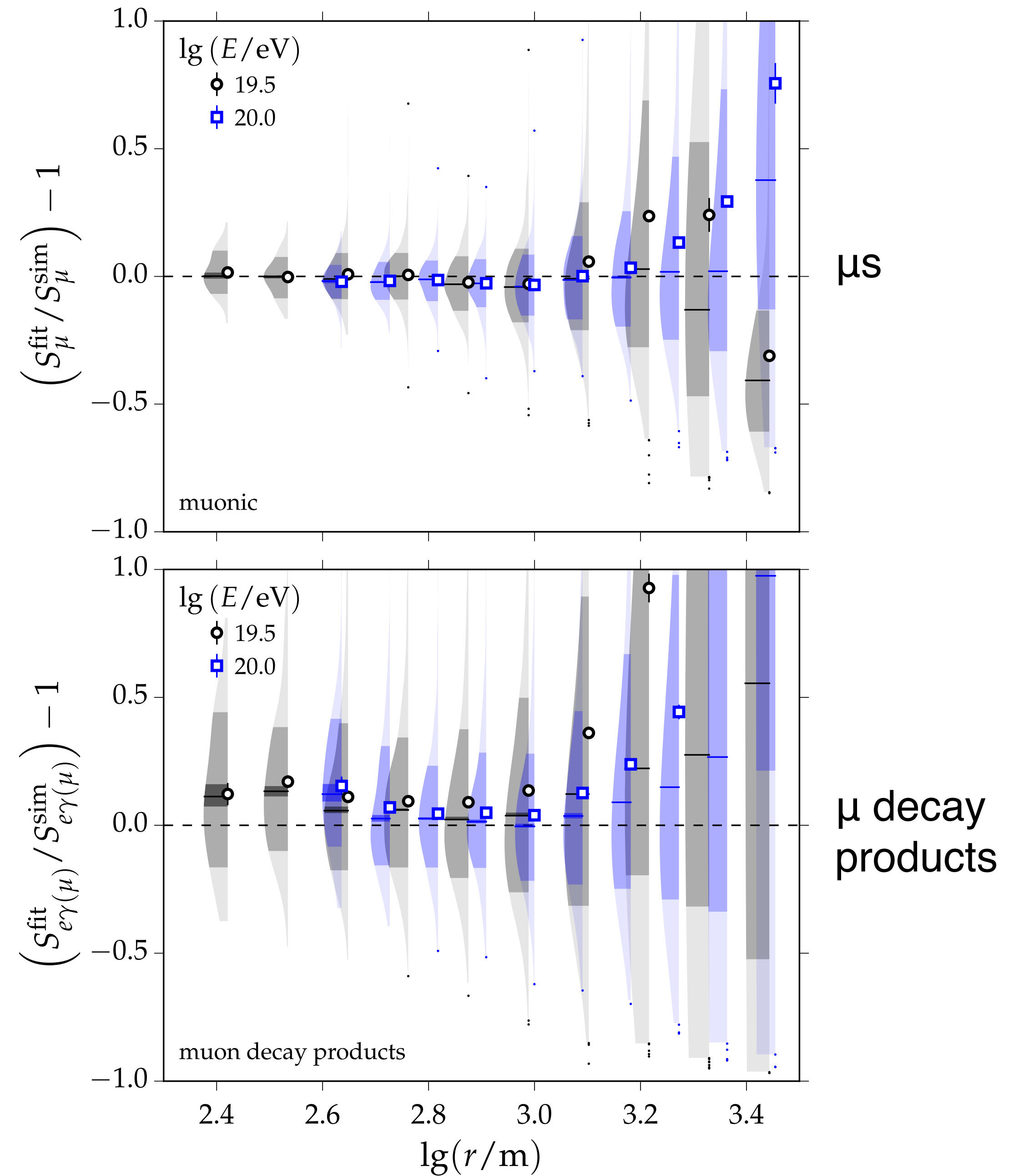
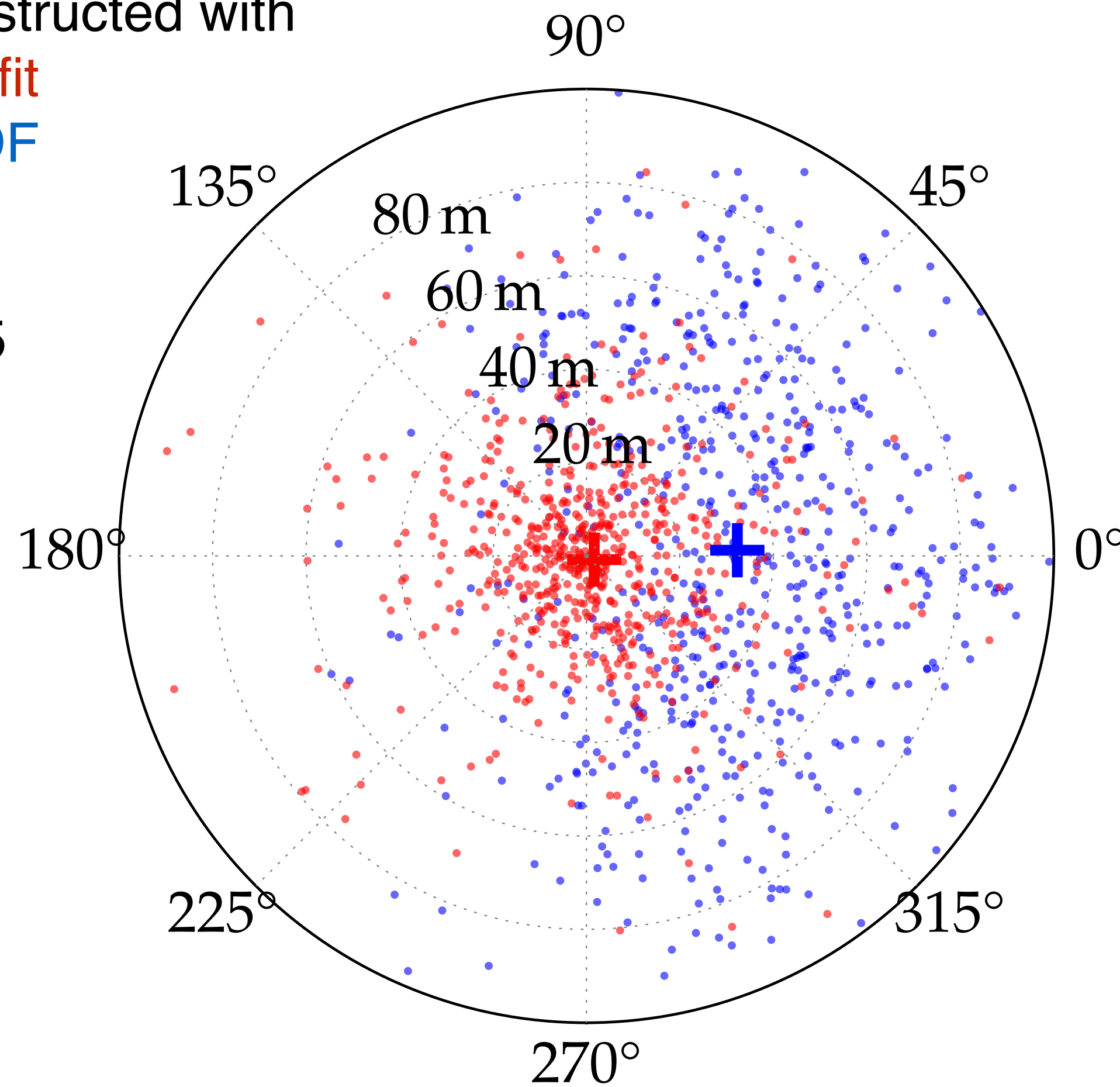
Reconstruction performance

Events reconstructed with

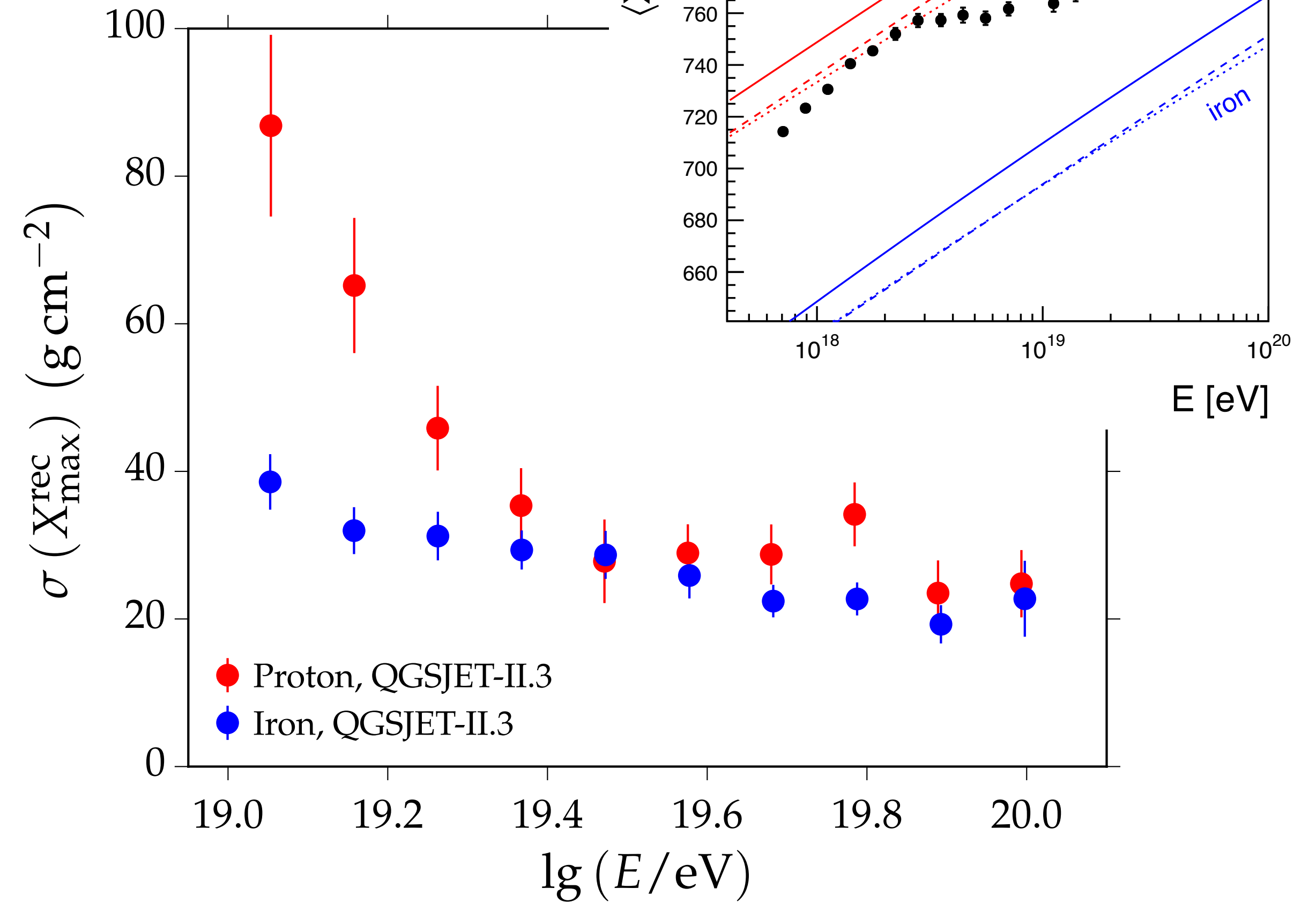
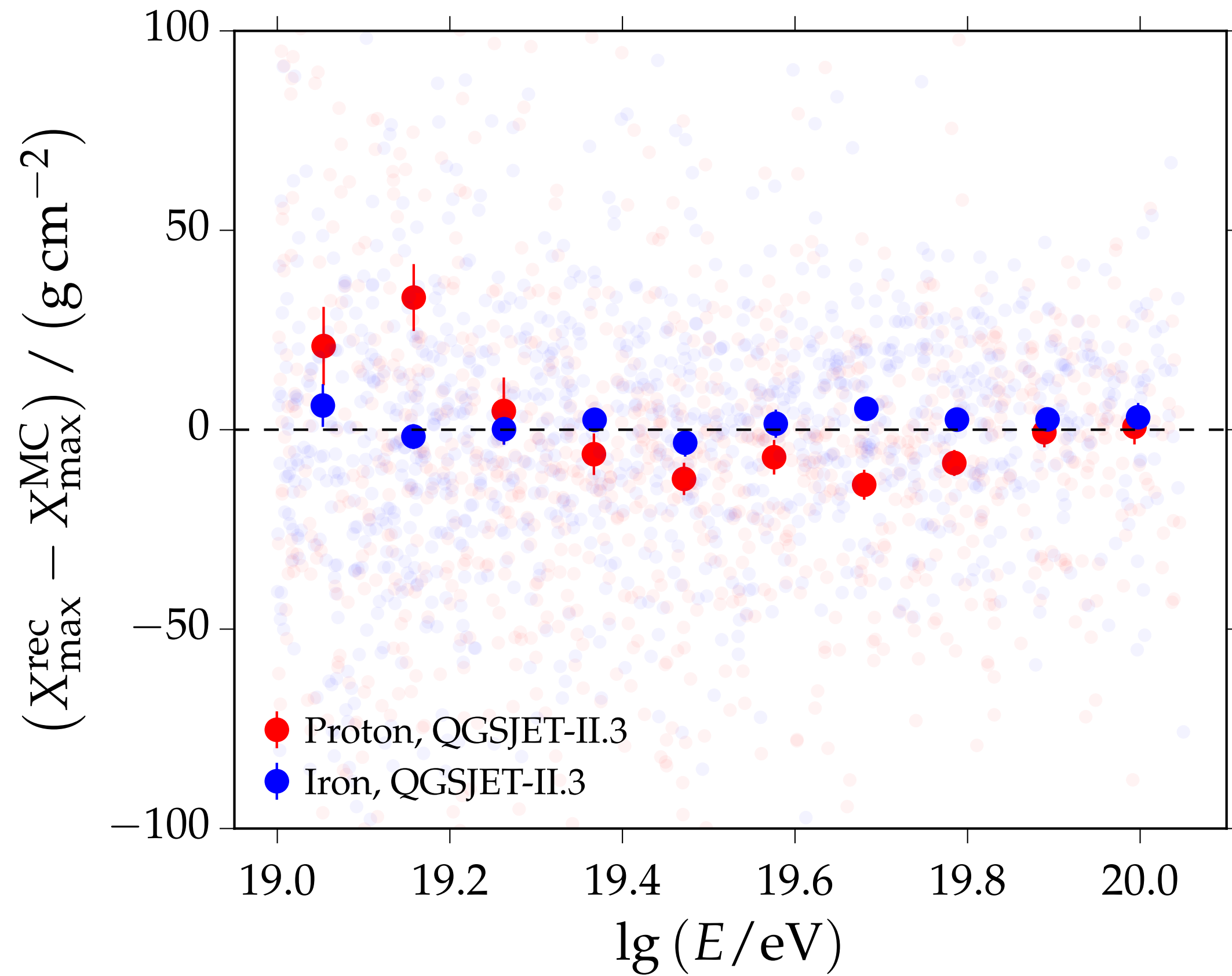
- **Universality fit**
- **Standard LDF**

$\theta=38^\circ$

$\lg(E/eV)=19.5$

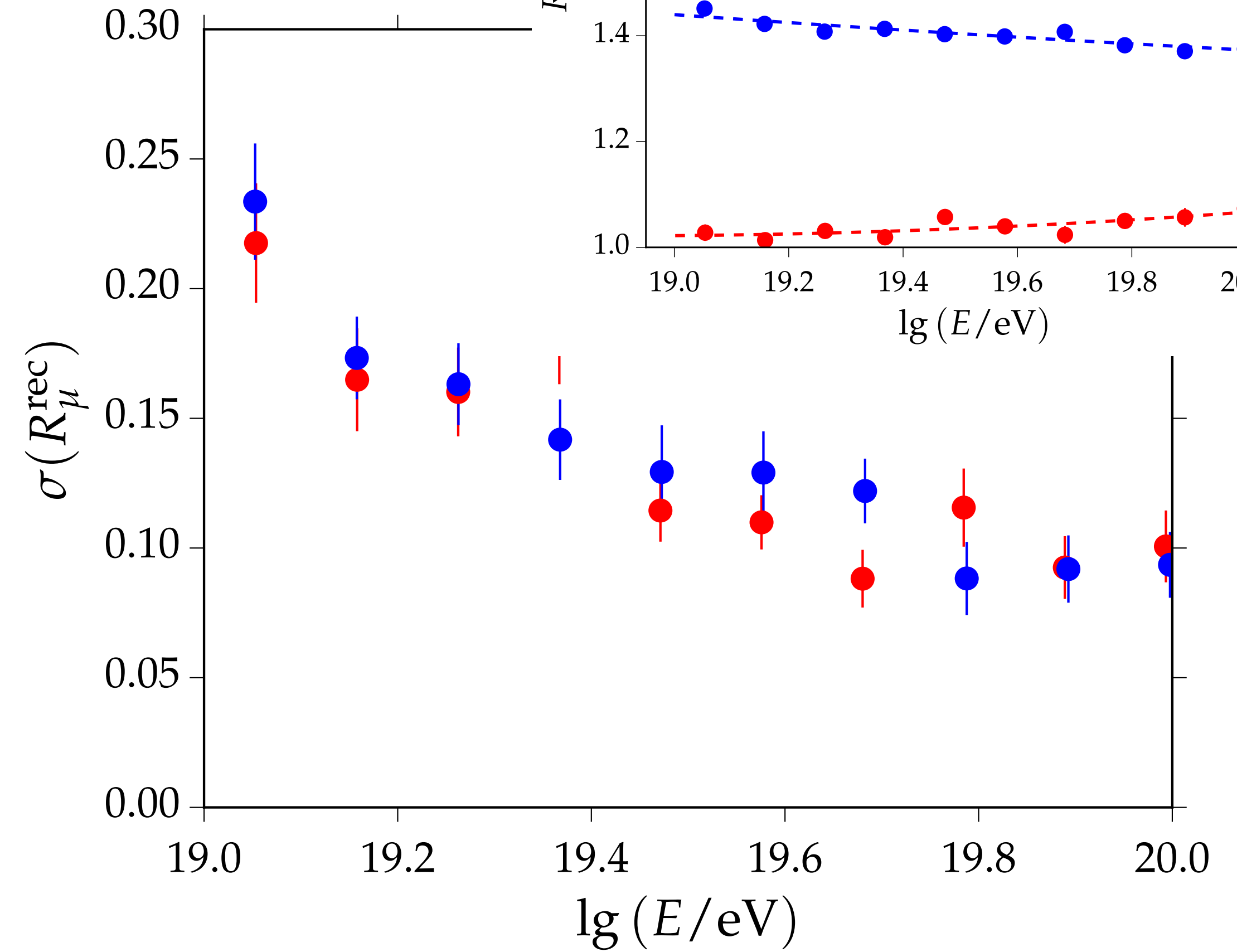
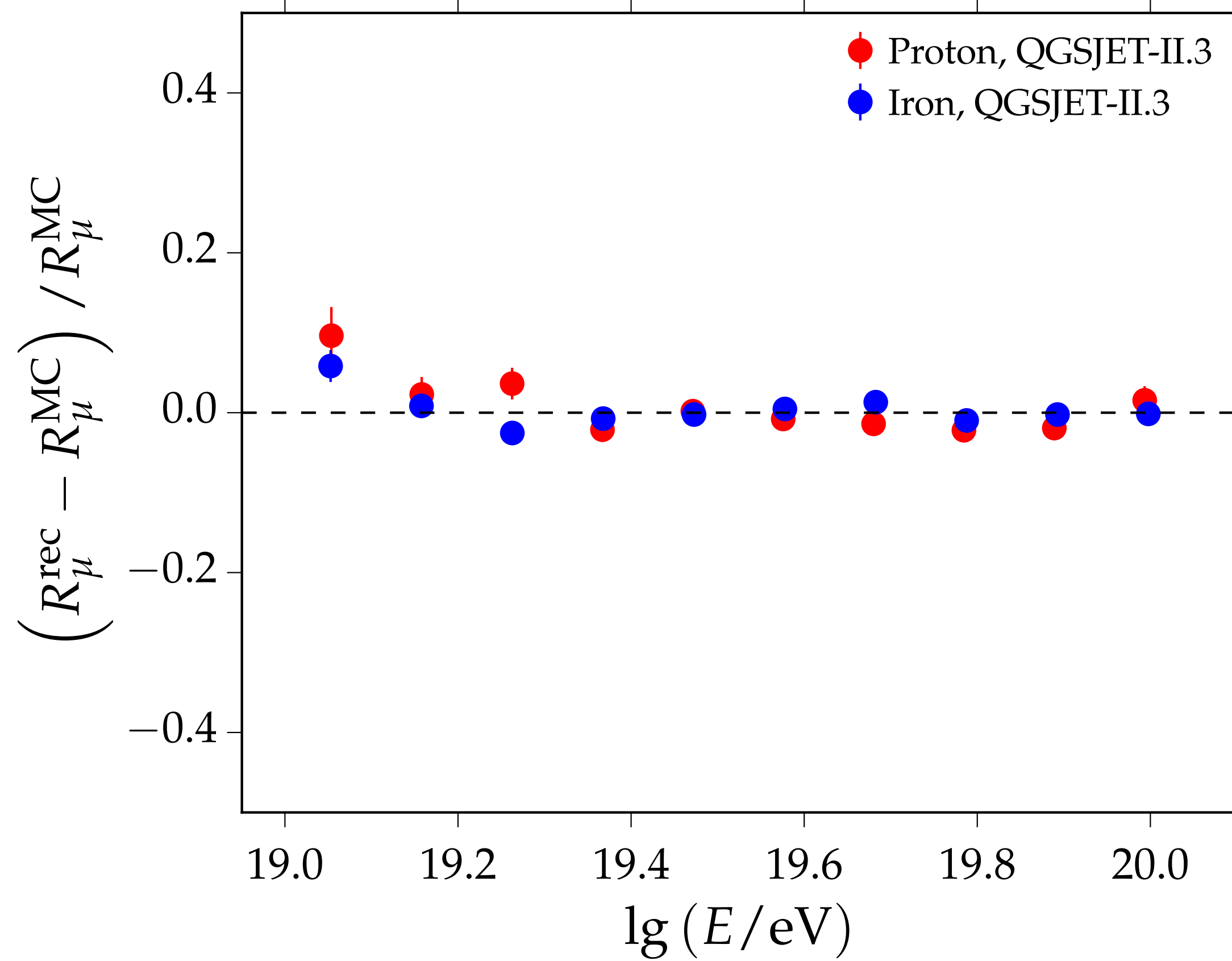


X_{\max} bias and resolution



Largest **systematic uncertainty**: Energy scale
 +10% shift in energy $\rightarrow \Delta X_{\max} = -10 \text{ g/cm}^2$

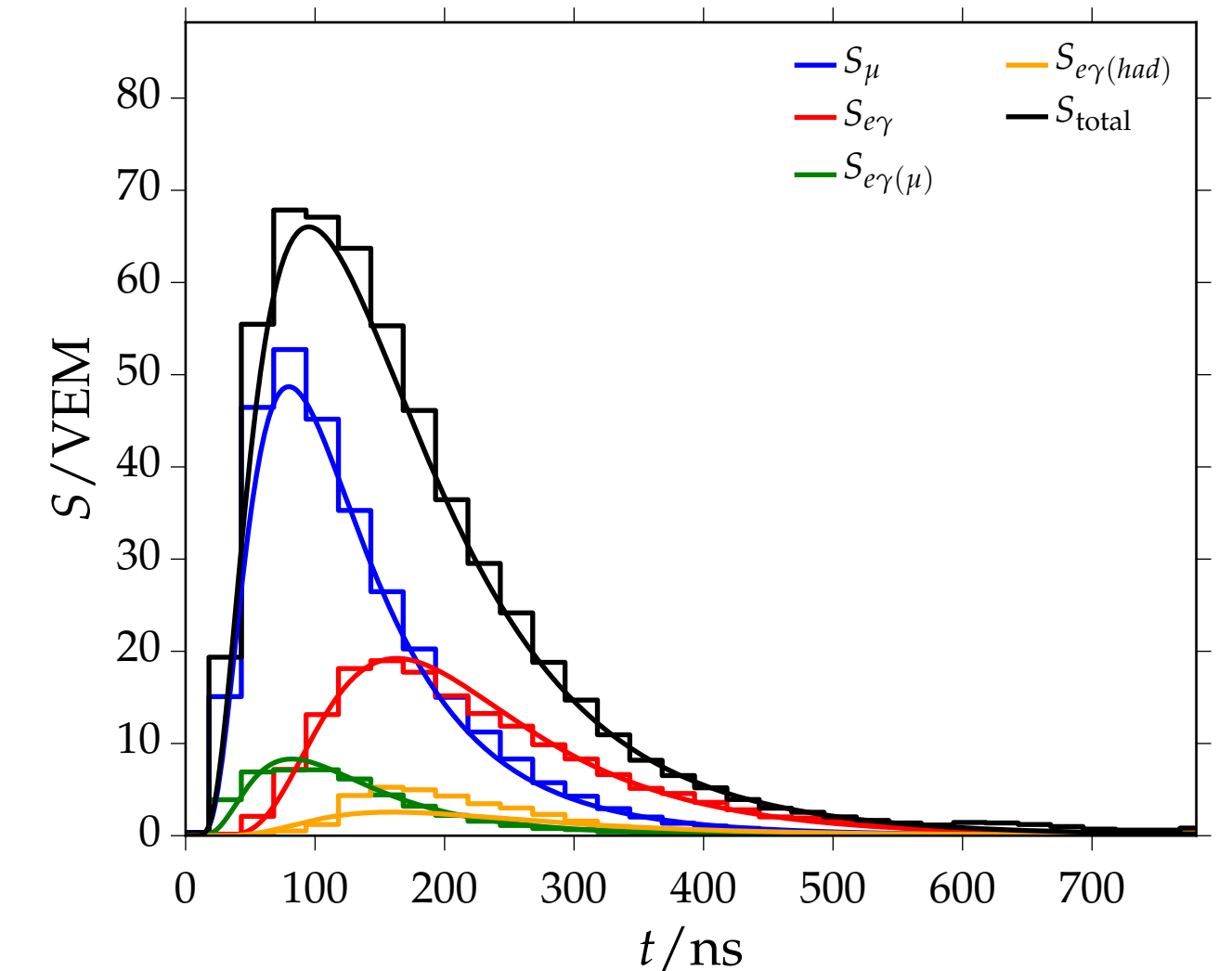
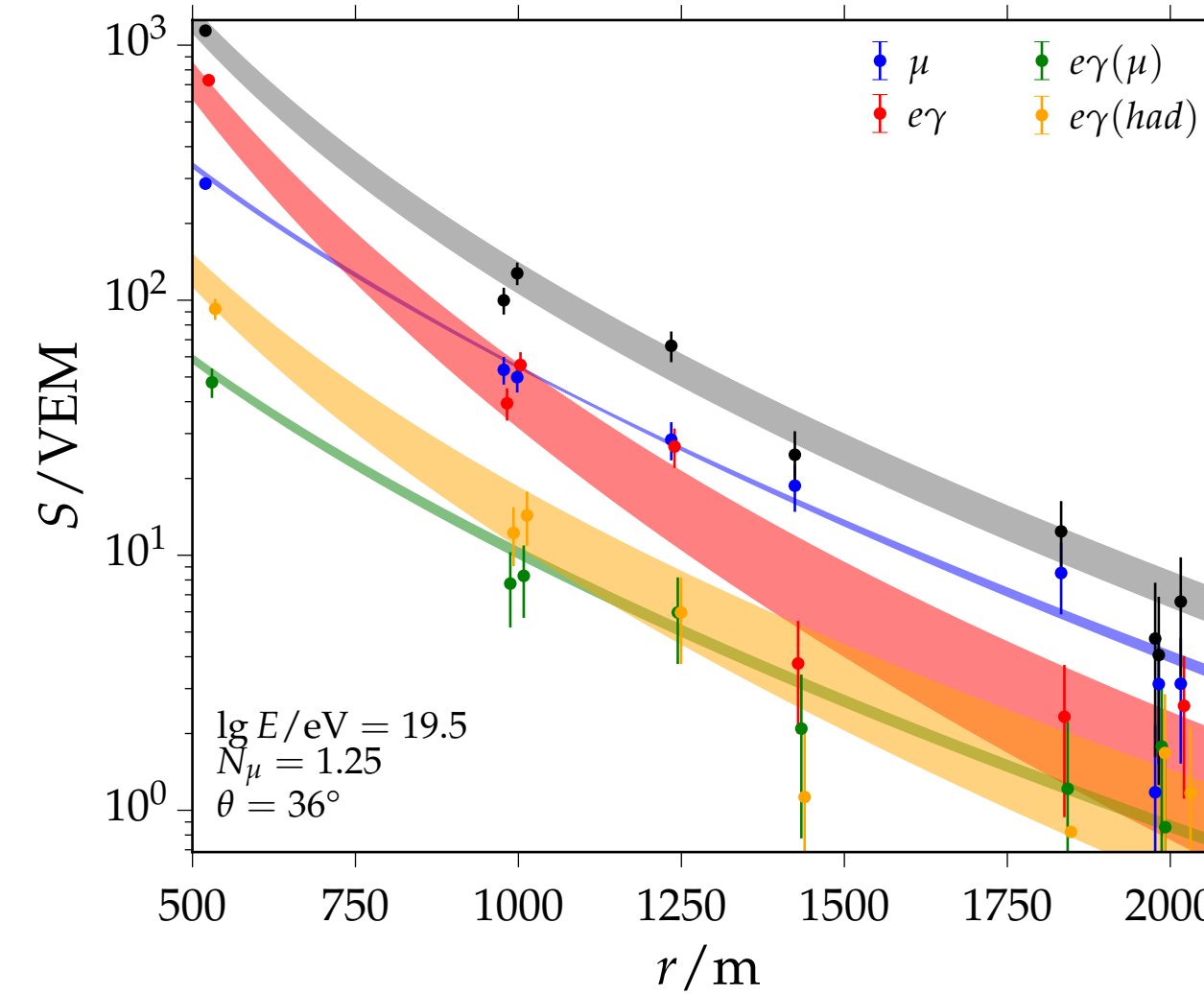
N_μ bias and resolution



Largest **systematic uncertainty**: Energy scale
 +10% shift in energy $\rightarrow \Delta N_\mu = -0.15$

Summary

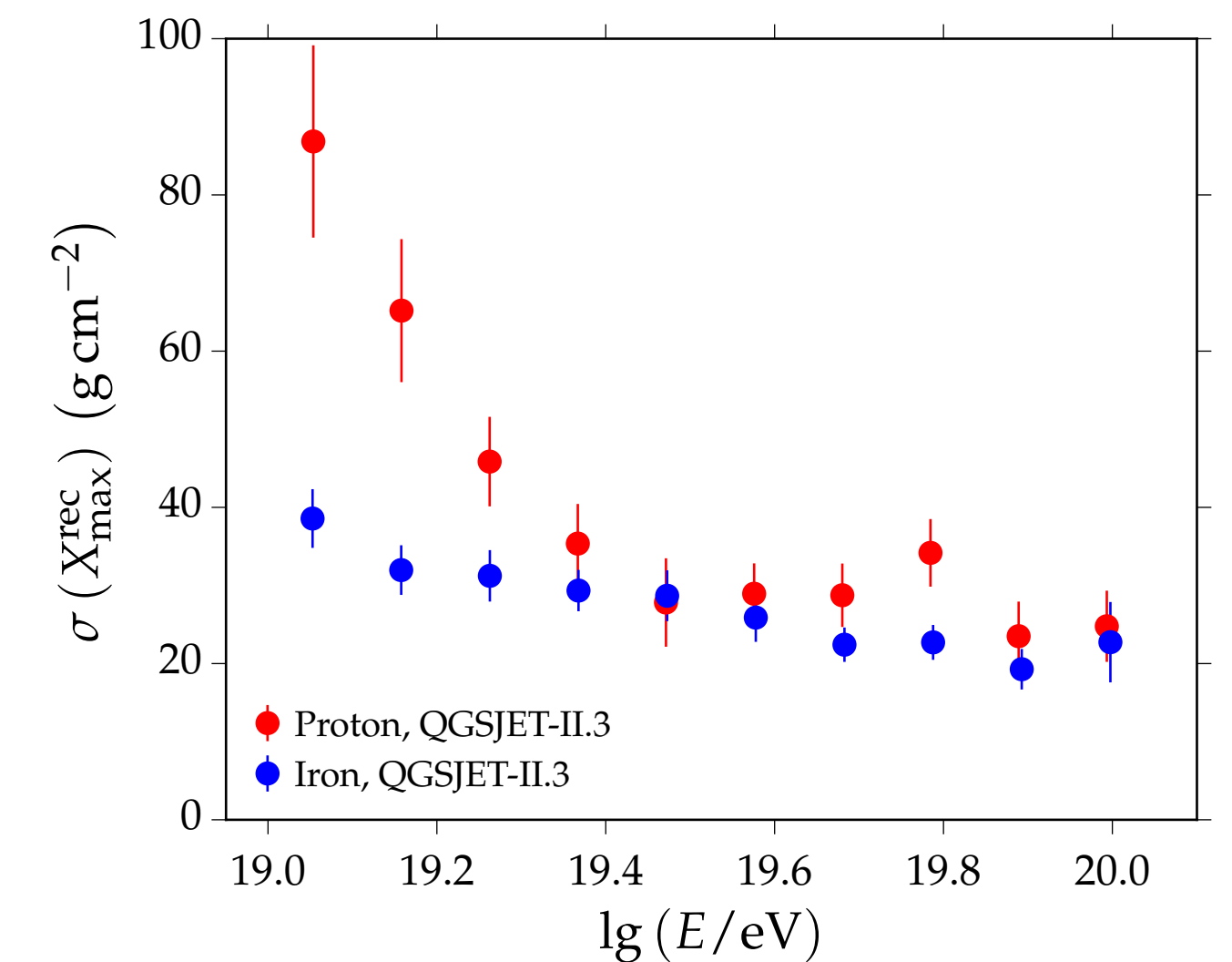
- **Signals and time structure** well described by shower universality
- Few parameters: N_μ , E , X_{\max} , X_0 (+geometry)
- Reconstruction using **only the time-dependent signal** of ground based detectors



Outlook

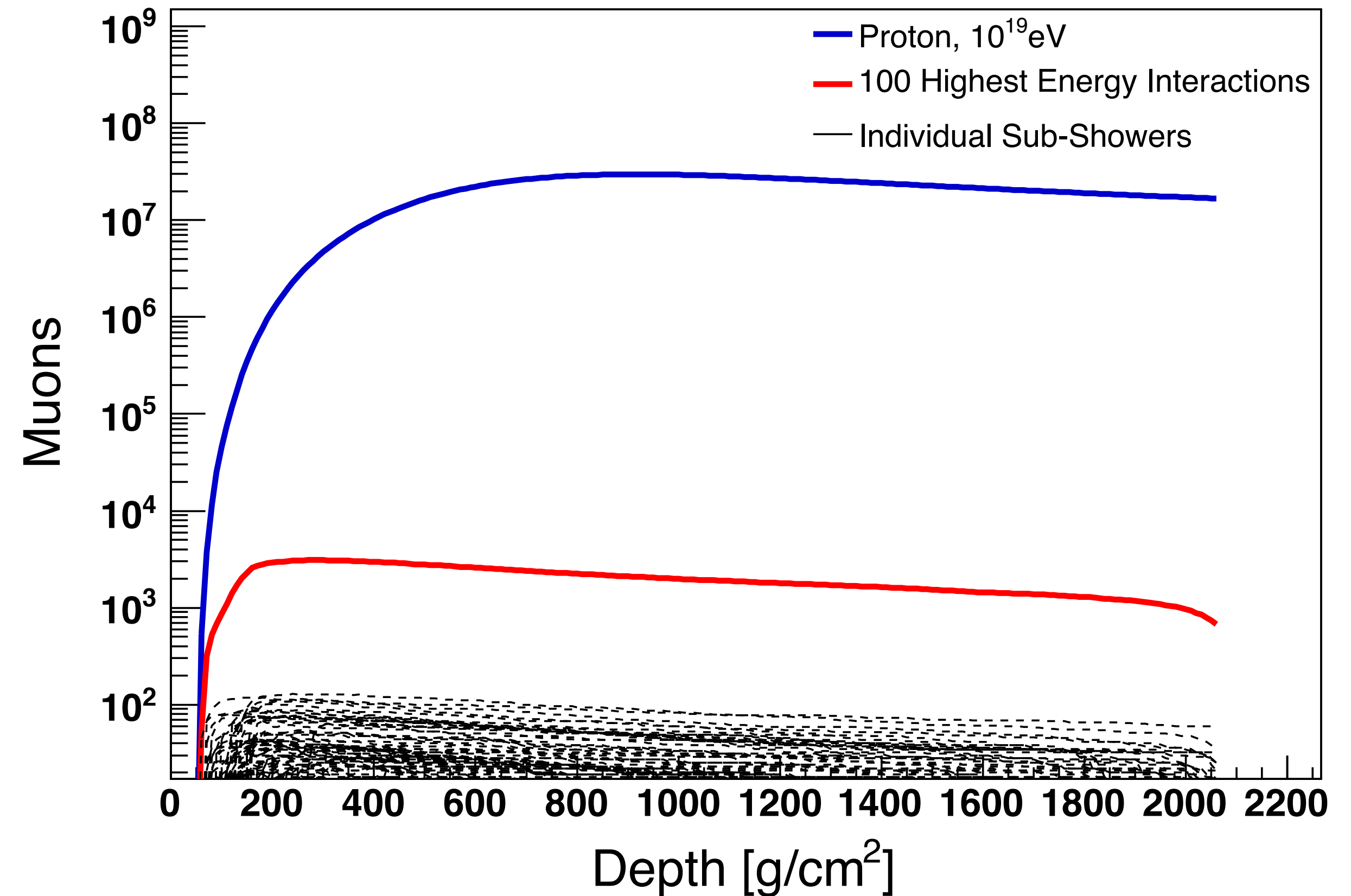
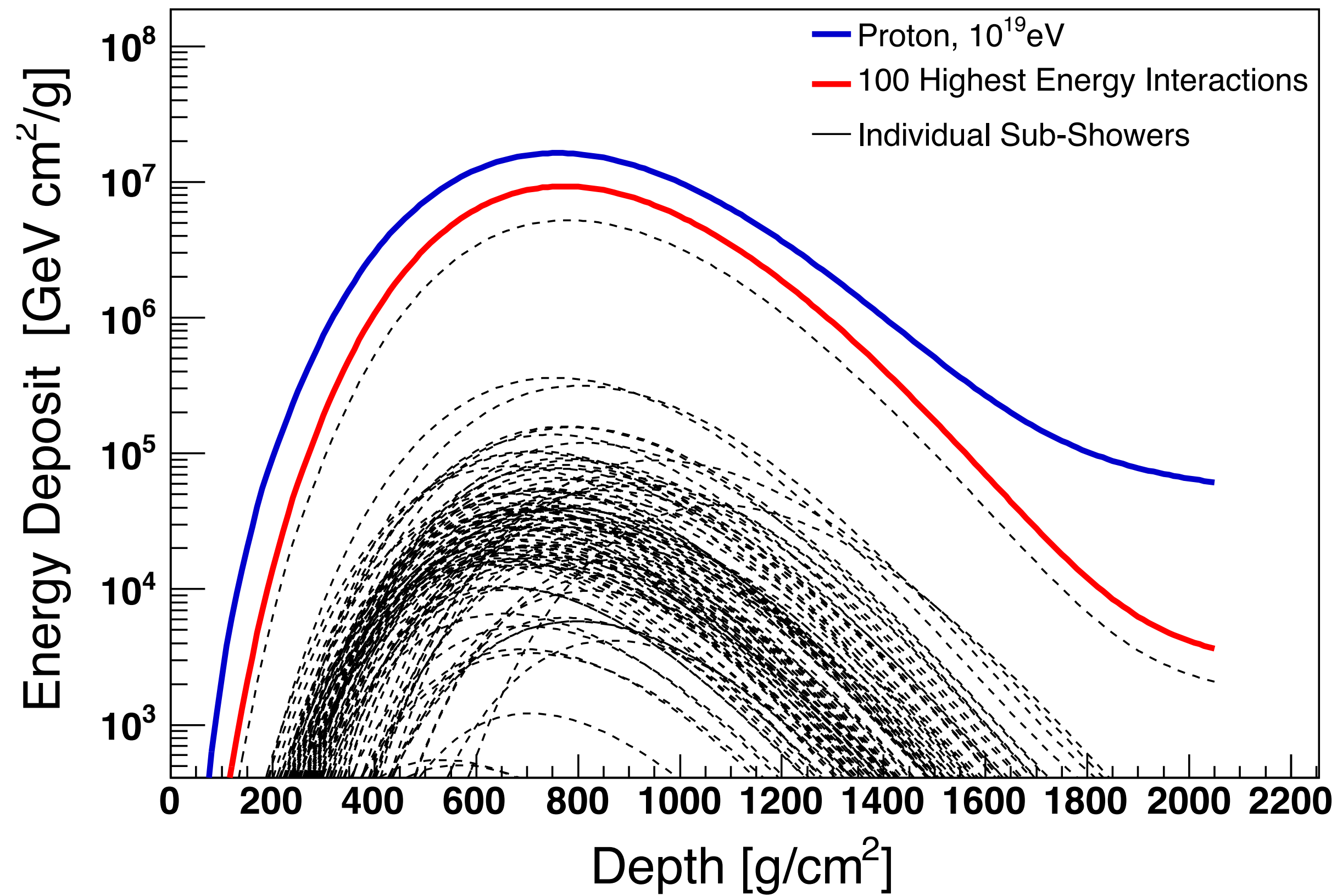
- Composition sensitive parameters *at 100% duty cycle*
- Mass composition at the **highest energies**
- **Mass-dependent search for correlations** to astrophysical sources
- Applicable to **other surface detector arrays**
- **Additional muon measurements** will reduce (N_μ , E) correlation

See R. Engel on the Auger upgrade PoS686 and M. Josebachuili on AMIGA POS409



Muonic component in air showers

R. Ulrich APS 2010



- X_{\max} is dominated by first interaction
Muons are produced late in the shower cascade
- Number of generations ~ 6 at 10¹⁹ eV
 - Amplified sensitivity to hadronic interactions

May allow to disentangle particle physics and composition using hybrid events (coincident measurement of SD and FD)