

Status of cosmic ray antideuteron searches



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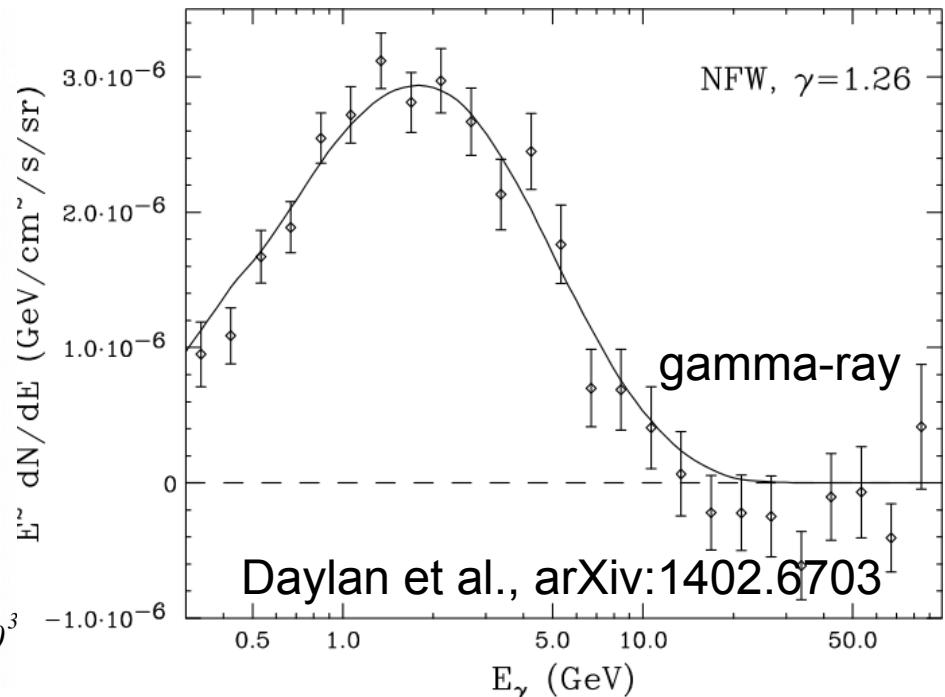
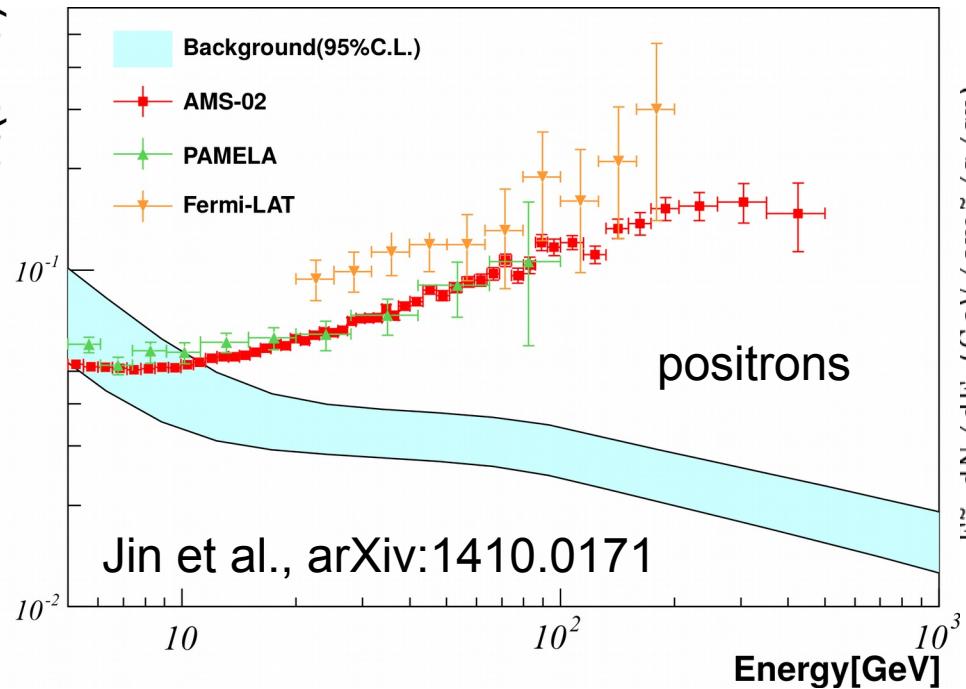


Review of the theoretical and experimental status of dark matter identification with cosmic-ray antideuterons

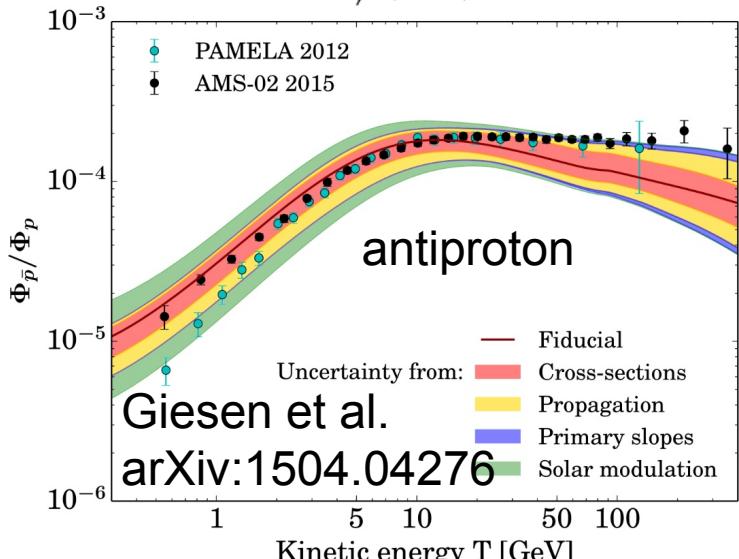
under review at Physics Reports: arXiv:1505.07785

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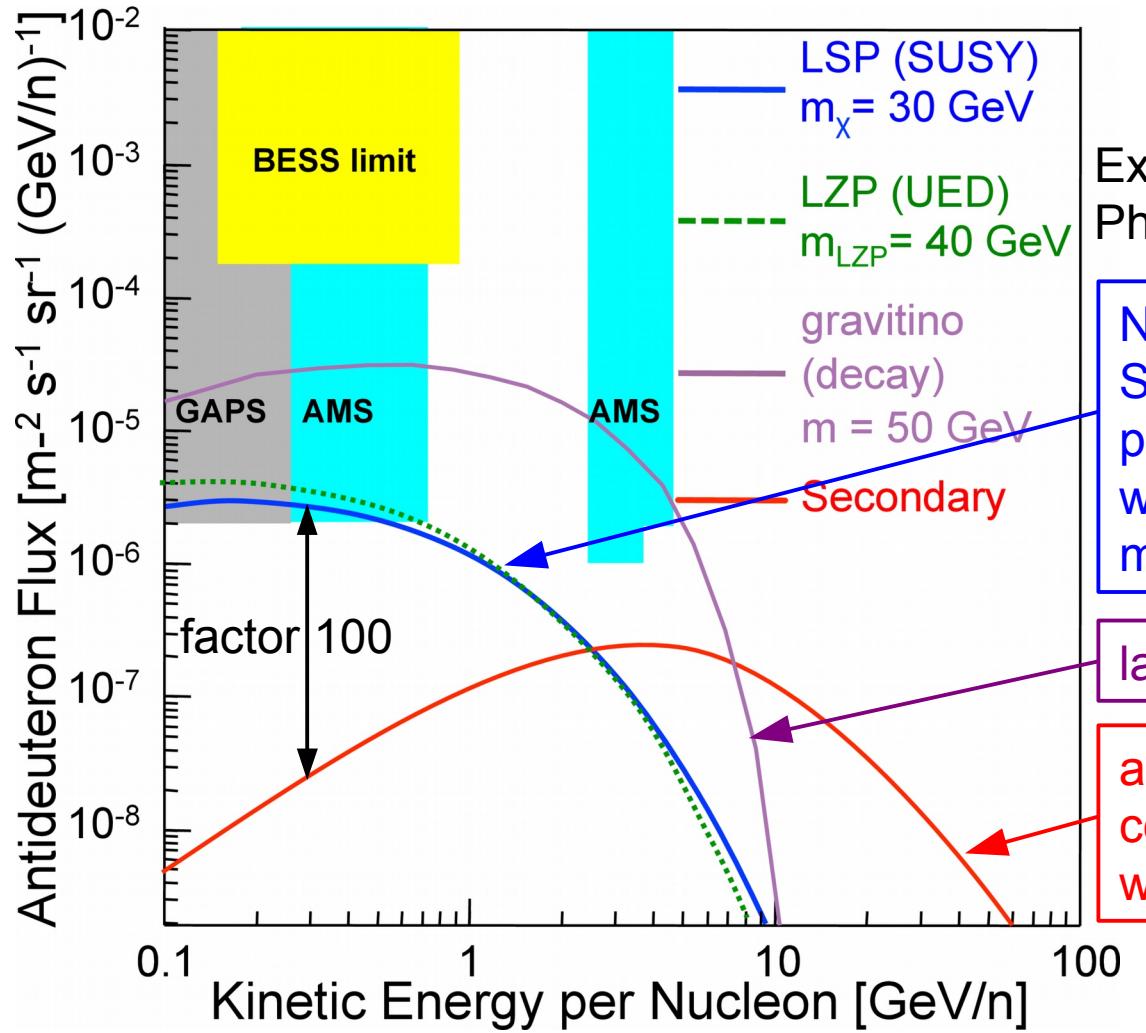
Dark matter signal in cosmic rays?



- unexplained features in positrons
- proposed theories:
 - astrophysical origin → pulsars
 - SNR acceleration
 - **dark matter self-annihilation**
- gamma-ray excess at the galactic center
→ 30GeV dark matter particle?
- **No (?) excess for antiprotons → inconclusive**



Status of cosmic ray antideuterons



Examples for beyond-standard-model Physics:

Neutralino:
SUSY lightest supersymmetric particle, decay into bb , compatible with signal from Galactic Center measured by Fermi

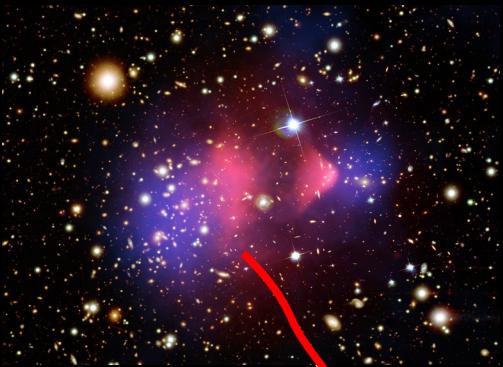
late decays of unstable gravitinos

astrophysical background:
collisions of protons and antiprotons with interstellar medium

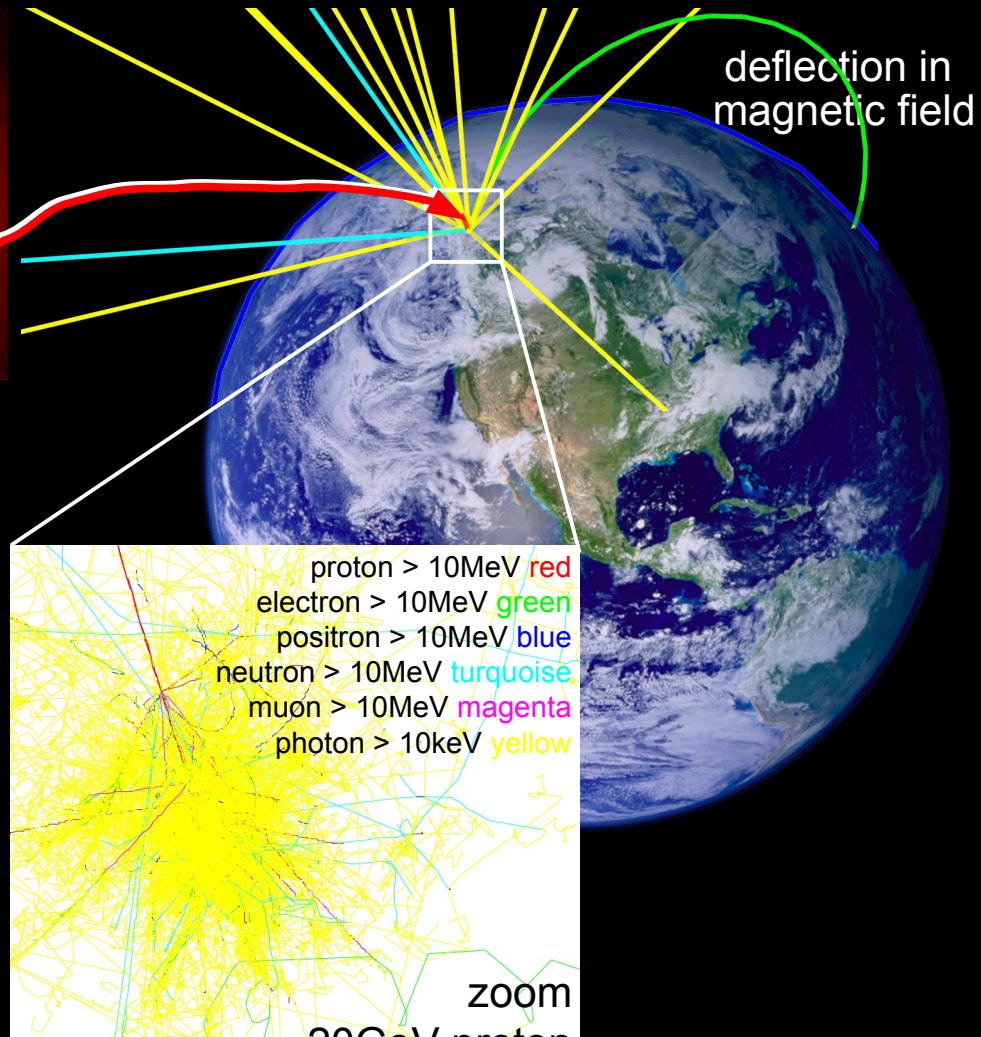
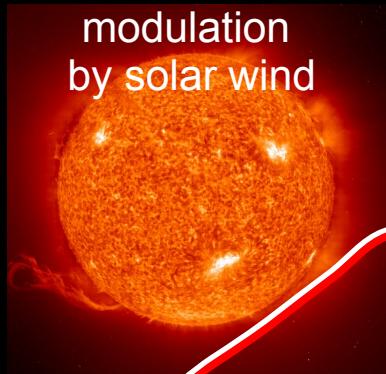
+ models with heavy dark matter

Antideuterons are the most important unexplored indirect detection technique!

Uncertainties

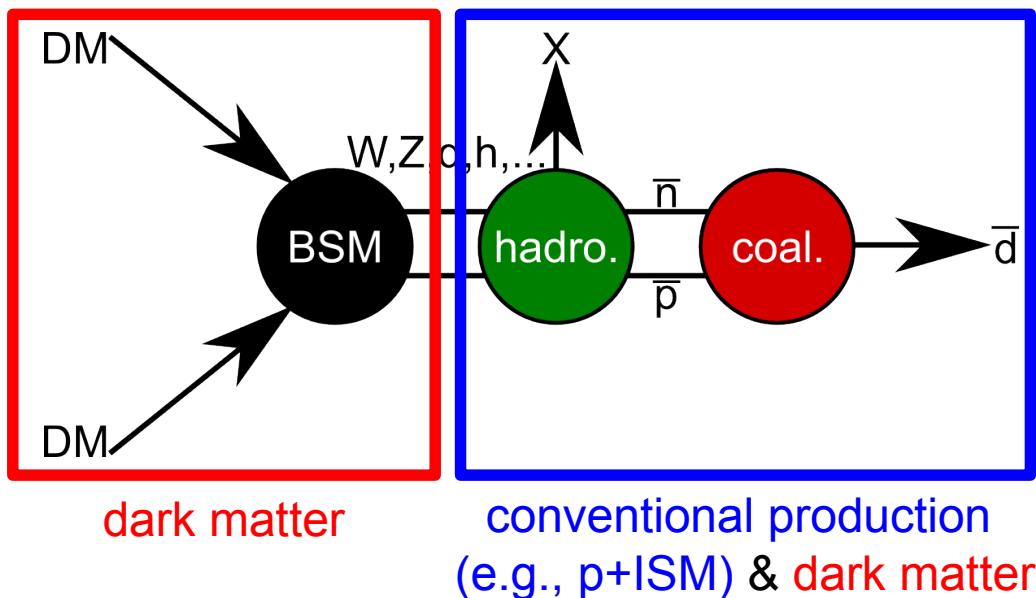


scattering in magnetic fields,
interaction with
interstellar medium

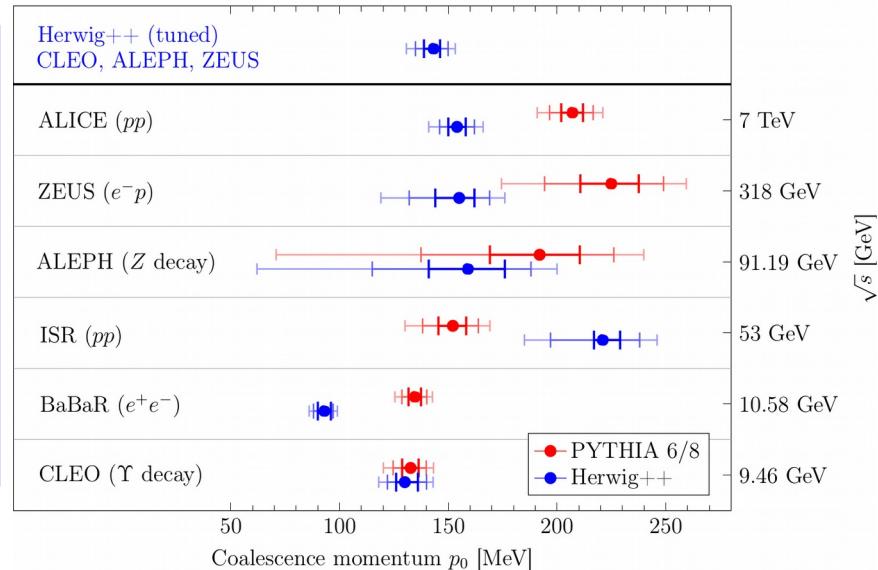


- Dark matter annihilation or decay
- Dark matter clumping
- **Antideuteron production**
- **Galactic propagation**
- Solar modulation
- Geomagnetic deflection
- Atmospheric interactions
- Interactions in detector

Antideuteron formation



Fitting p_0 to data on \bar{d} production

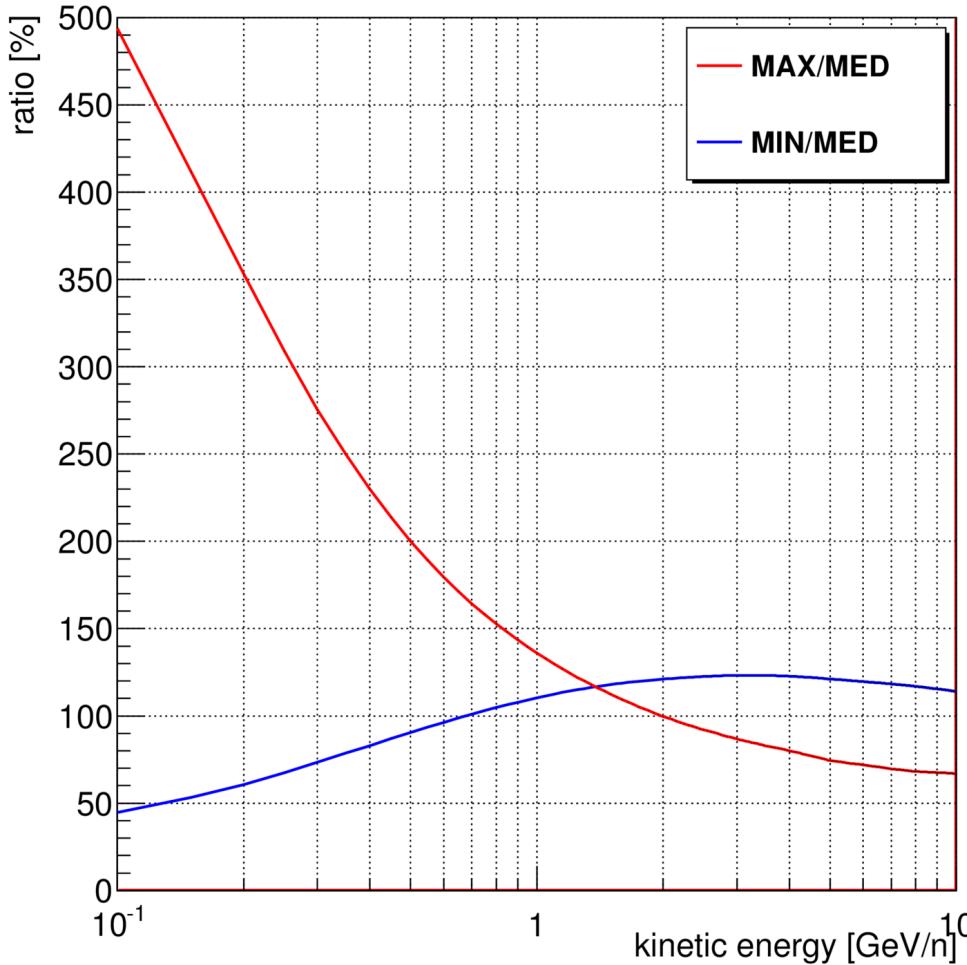


- antideuterons can be formed by an antiproton-antineutron pair if relative momentum is small (coalescence momentum p_0)

$$\frac{dN_{\bar{d}}}{dT_{\bar{d}}} = \frac{p_0^3}{6} \frac{m_{\bar{d}}}{m_{\bar{n}} m_{\bar{p}}} \frac{1}{\sqrt{T_{\bar{d}}^2 + 2m_{\bar{d}}T_{\bar{d}}}} \frac{dN_{\bar{n}}}{dT_{\bar{n}}} \frac{dN_{\bar{p}}}{dT_{\bar{p}}}$$

- important differences for different experiments and MC generators exist → more data would help

Propagation uncertainty



- Propagation is a large uncertainty source for low-energy antideuterons:
halo size for diffusion calculation is poorly constrained
- More data on different cosmic nuclei are needed (and hope that they do not need more complicated modeling for interpretation!)

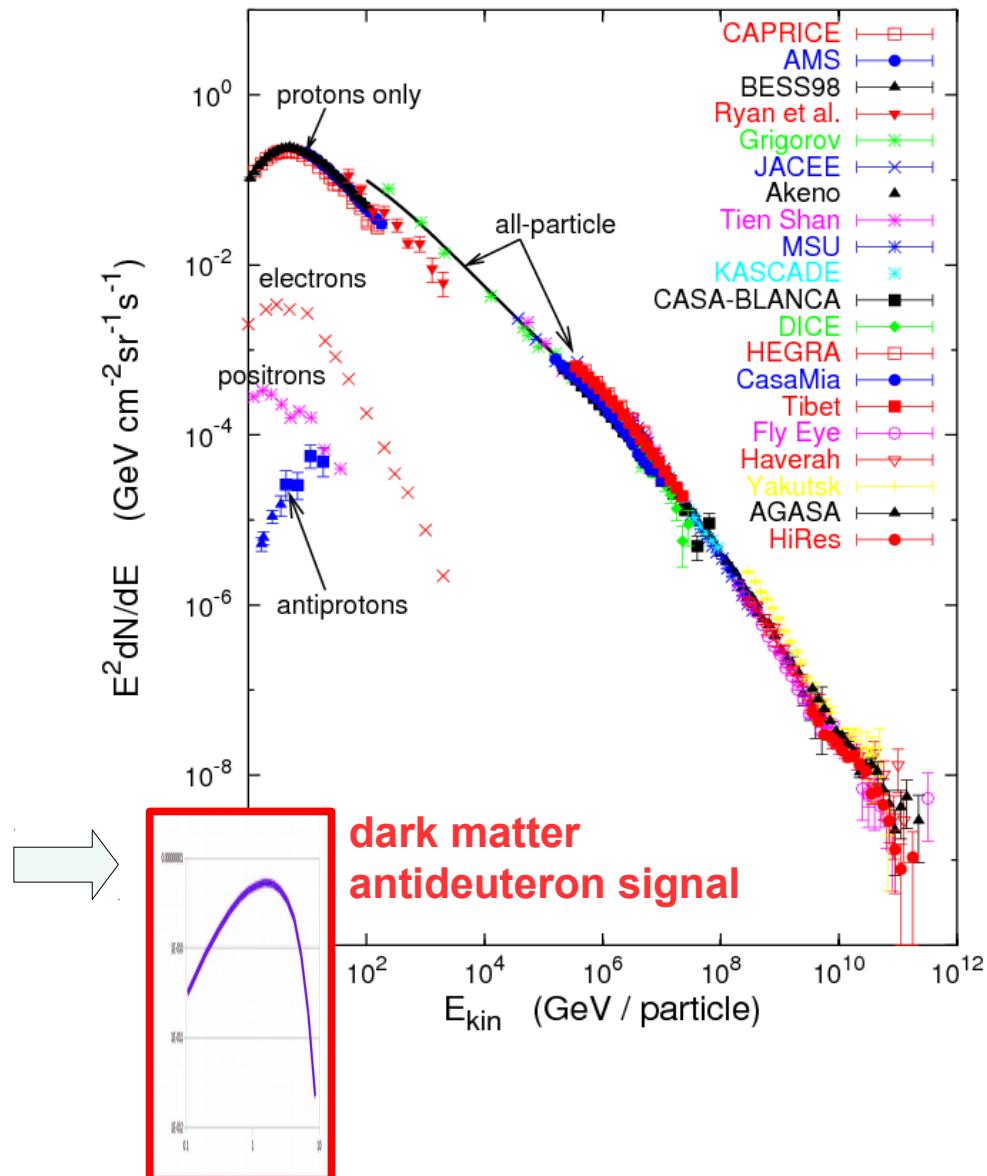
Identification challenge

Required rejections for antideuteron detection:

- **protons**: $> 10^8 - 10^{10}$
- **He-4**: $> 10^7 - 10^9$
- **electrons**: $> 10^6 - 10^8$
- **positrons**: $> 10^5 - 10^7$
- **antiprotons**: $> 10^4 - 10^6$

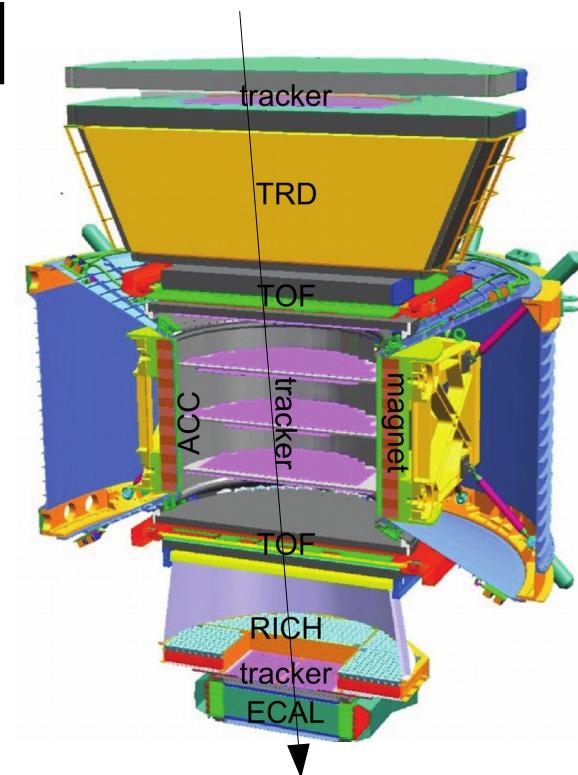
Antideuteron measurement with balloon and space experiments require:

- **strong background suppression**
- **long flight time and large acceptance**



AMS-02 antideuteron analysis

| | e^- | p | He,Li,Be,..Fe | γ | e^+ | \bar{p}, \bar{d} | $\bar{\text{He}}, \bar{\text{C}}$ |
|----------------------------------|-------|-----|---------------|----------|-------|--------------------|-----------------------------------|
| TRD $\gamma = E/m$ | | | | | | | |
| TOF $dE/dx, velocity$ | | | | | | | |
| Tracker $dE/dx, momentum$ | | | | | | | |
| RICH precise velocity | | | | | | | |
| ECAL shower shape, energy det | | | | | | | |



- Operating on the ISS since 2011

- **antideuteron identification:**

- lower velocities: **Time Of Flight** scintillator system
- higher velocities: **Ring Image Cherenkov** detector

- **self-calibrated analysis:**

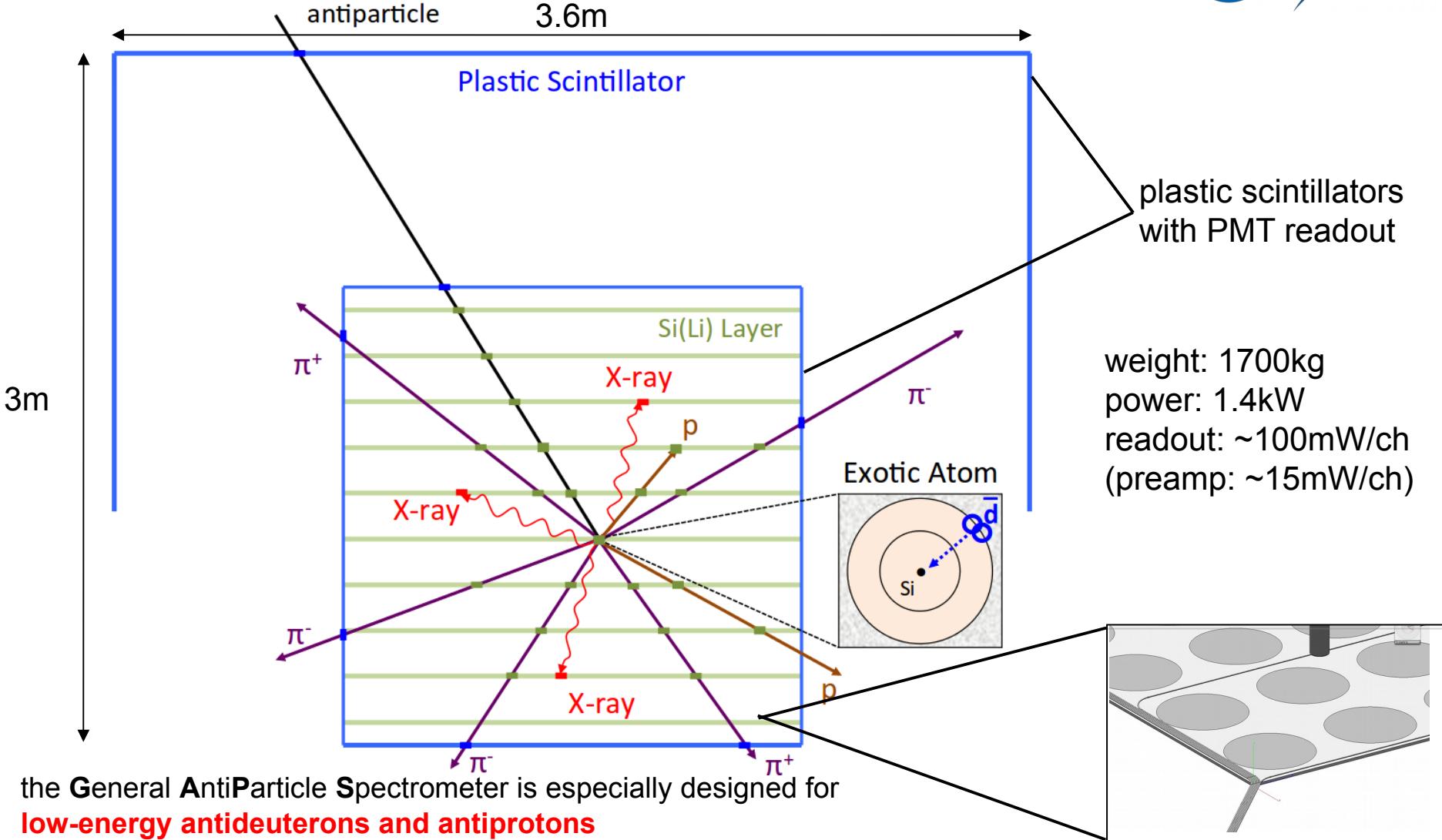
- calibrate antideuteron analysis with deuterons and antiprotons (simulations and data)
- geomagnetic cut-off and solar effects: study much more abundant low-energy protons, antiprotons, and deuterons for calibration

$$m = R \cdot Z \sqrt{\frac{1}{\beta^2} - 1}$$

Analysis ongoing!

The GAPS experiment

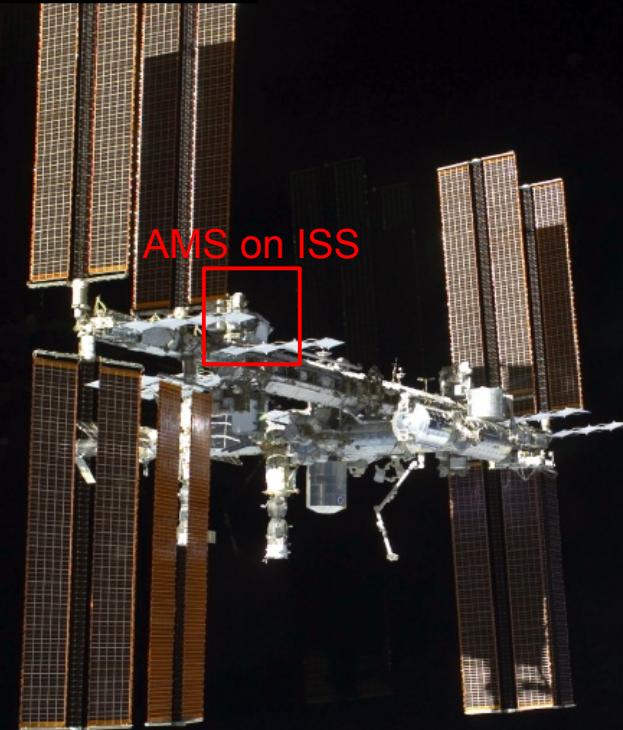
Columbia U, UC Berkeley
UCLA, U Hawaii,
Haverford



- the General AntiParticle Spectrometer is especially designed for **low-energy antideuterons and antiprotons**
- identification by stopping and creation of an exotic atom
[KEK testbeam measurements → Astropart. Phys. 49, 52 (2013)]
- LDB flights from Antarctica proposed to NASA

Conclusion

- measurement of antideuterons is a promising way for indirect dark matter search
- AMS on the ISS is currently the best instrument for the study of antideuterons
- future GAPS is specifically designed for low-energetic antideuterons
- more exchange between theory and experiments:
We started a bigger community effort last year.



GAPS from
Antarctica

