

The Little Review on Leptogenesis

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DISCRETE08, 13 December 2008, Valencia, Spain

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- **Matter–AntiMatter Asymmetry and Leptogenesis**
- **The Flavourdynamics of Leptogenesis**
- **Resonant Leptogenesis (RL)**
- **Phenomenology of RL Models**
- **Conclusions**

• Matter–AntiMatter Asymmetry and Leptogenesis

$$\eta_B^{\text{CMB}} = \frac{n_B}{n_\gamma} = 6.1_{-0.2}^{+0.3} \times 10^{-10} \quad (\eta_B^{\text{BBN}} = 3.4\text{--}6.9 \times 10^{-10}, \text{ at 95\% CL})$$

Sakharov's conditions for generating the BAU:

[A.D. Sakharov, JETP Lett. 5 (1967) 24.]

- B-violating interactions
- C and CP violation
- Out-of-equilibrium dynamics

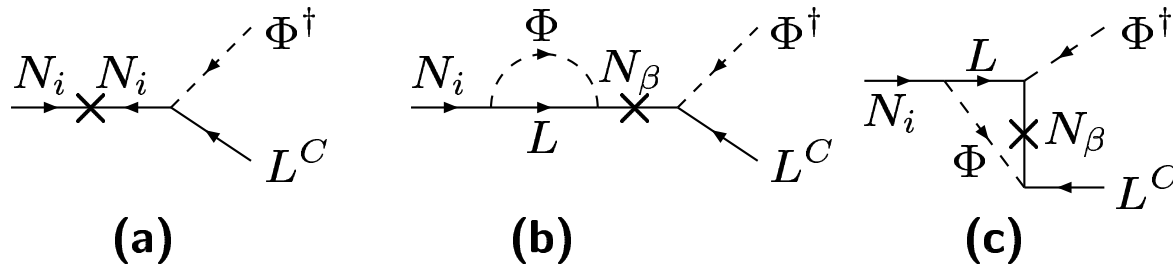
Typical Scenarios for Baryogenesis:

- **Baryogenesis through the decay of a heavy particle**
Out-of-equilibrium, *B*-violating decay of a heavy GUT particle, e.g. in $SO(10)$.
[M. Yoshimura, PRL**41** (1978) 281; S. Dimopoulos and L. Susskind, PRD**18** (1978) 4500.]
- **Baryogenesis at the electroweak phase transition**
BAU generated by $(B + L)$ -violating sphaleron interactions at $T \sim T_c \approx 140$ GeV, through a 1st order phase transition.
[V.A. Kuzmin, V.A. Rubakov, M.E. Shaposhnikov, PLB**155** (1985) 36;
MSSM: M. Carena, M. Quirós, C. Wagner '96; K. Rummukainen, M. Laine '98 . . .]

- **Baryogenesis through Leptogenesis**

Out-of-equilibrium L -violating decays of heavy Majorana neutrinos produce a net lepton asymmetry, converted into the BAU through $(B + L)$ -violating sphaleron interactions.

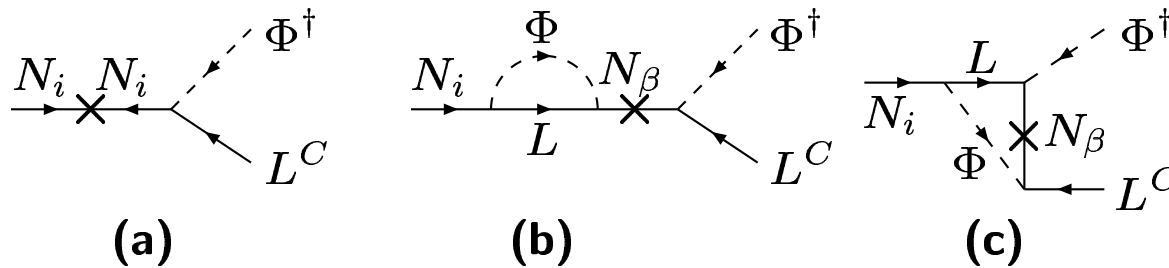
[M. Fukugita and T. Yanagida, PLB174 (1986) 45.]



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Models of Leptogenesis:

[For a review, see, R.N. Mohapatra et al, RPP70 (2007) 1757]

- **Resonant Leptogenesis**

- **Dirac Leptogenesis**

[K. Dick, M. Lindner, M. Ratz, D. Wright, PRL84 (2000) 4039.]

- **Other scenarios:** Non-thermal leptogenesis, Affleck–Dine, spontaneous leptogenesis, . . .

[For a review, see, M. Dine and A. Kusenko, Rev. Mod. Phys. 76 (2004) 1.]

• The Flavourdynamics of Leptogenesis

BAU can be generated from and protected in a single lepton flavour:

$$\frac{1}{3}B - L_{e,\mu,\tau}.$$

[e.g. J.A. Harvey, M.S. Turner, PRD42 (1990) 3344;
H. Dreiner, G.G. Ross, NPB410 (1993) 188;
J.M. Cline, K. Kainulainen, K.A. Olive, PRD49 (1994) 6394.]

Two sources of flavour effects:

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Two sources of flavour effects:

- Charged-lepton Yukawa couplings $h_{e,\mu,\tau}$

[R. Barbieri, P. Creminelli, A. Strumia, N. Tetradis, NPB575 (2000) 61;
A. Pilaftsis, T.E.J. Underwood, PRD72 (2005) 113001;
E. Nardi, Y. Nir, J. Racker, E. Roulet, JHEP0601 (2006) 068;
A. Abada, S. Davidson, F. X. Josse-Michaux, M. Losada, A. Riotto, JCAP0604 (2006) 004.]

Modify BAU predictions by up to 1-order of magnitude.

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Modify BAU predictions by up to 1-order of magnitude.

- Heavy-neutrino Yukawa couplings $h_{l\alpha}^\nu$

[A. Pilaftsis, PRL95 (2005) 081602 [hep-ph/0408103];
T. Endoh, T. Morozumi and Z. h. Xiong, PTP111 (2004) 123;
A. Pilaftsis, T.E.J. Underwood, PRD72 (2005) 113001;
P. Di Bari, NPB727 (2005) 318; O. Vives, PRD73 (2006) 073006.]

Modify BAU predictions by many orders of magnitude, e.g. $> 10^6!$

Boltzmann Equations

[A.P., T.E. Underwood, NPB692 (2004) 303; PRD72 (2005) 113001.]

$$\frac{d\eta_{N_\alpha}}{dz} = \frac{z}{H(z=1)} \left[\left(1 - \frac{\eta_{N_\alpha}}{\eta_{N_\alpha}^{\text{eq}}}\right) \sum_{k=e,\mu,\tau} (\Gamma^D(\alpha k) + \Gamma_{\text{Yukawa}}^S(\alpha k) + \Gamma_{\text{Gauge}}^S(\alpha k)) \right. \\ \left. - \frac{2}{3} \sum_{k=e,\mu,\tau} \eta_{\Delta L_k} \delta_{N_\alpha}^k (\hat{\Gamma}^D(\alpha k) + \hat{\Gamma}_{\text{Yukawa}}^S(\alpha k) + \hat{\Gamma}_{\text{Gauge}}^S(\alpha k)) \right],$$

$$\frac{d\eta_B}{dz} = -\frac{z}{H(z=1)} \left[\eta_B + \frac{28}{51} \sum_{j=e,\mu,\tau} \eta_{L_j} \right. \\ \left. + \frac{225}{561} \frac{v^2(T)}{T^2} \left(\eta_B + \frac{108}{225} \sum_{j=e,\mu,\tau} \eta_{L_j} \right) \right] \Gamma_{\Delta(B+L)},$$

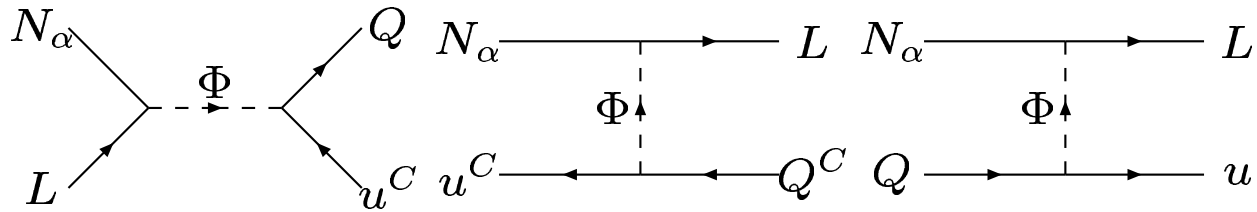
[S.Y. Khlebnikov, M.E. Shaposhnikov, NPB308 (1988) 885;
L. Carson, X. Li, L.D. McLerran, R.T. Wang, PRD42 (1990) 2127;
M. Laine, M.E. Shaposhnikov, PRD61 (2000) 117302.]

$$\frac{d\eta_{L_i}}{dz} = \frac{3}{2} \frac{d\eta_{\Delta L_i}}{dz} - \frac{2}{21} \frac{d\eta_{\Delta L}}{dz} + \frac{1}{3} \frac{d\eta_B}{dz},$$

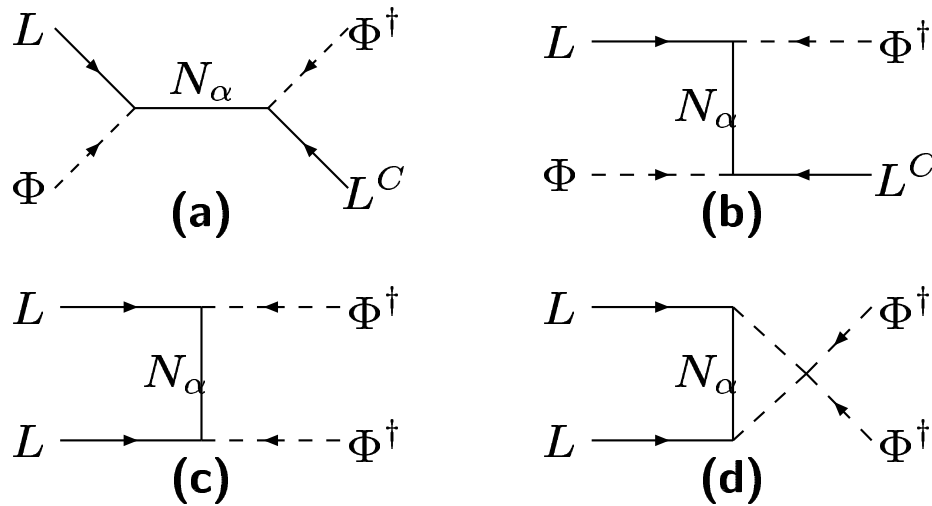
$$\begin{aligned}
\frac{d\eta_{\Delta L_j}}{dz} &= \frac{z}{H(z=1)} \\
&\times \left\{ \sum_{\alpha=1}^3 \delta_{N_\alpha}^j \left(\frac{\eta_{N_\alpha}}{\eta_{N_\alpha}^{\text{eq}}} - 1 \right) \sum_{k=e,\mu,\tau} \left(\Gamma^D(\alpha k) + \Gamma_{\text{Yukawa}}^S(\alpha k) + \Gamma_{\text{Gauge}}^S(\alpha k) \right) \right. \\
&- \frac{2}{3} \eta_{\Delta L_j} \left[\sum_{\alpha=1}^3 B_{N_\alpha}^j \left(\tilde{\Gamma}^D(\alpha j) + \tilde{\Gamma}_{\text{Yukawa}}^S(\alpha j) + \tilde{\Gamma}_{\text{Gauge}}^S(\alpha j) + \Gamma_{\text{Yukawa}}^W(\alpha j) + \Gamma_{\text{Gauge}}^W(\alpha j) \right) \right. \\
&\quad \left. \left. + \sum_{k=e,\mu,\tau} \left(\Gamma_{\text{Yukawa}}^{\Delta L=2}(jk) + \Gamma_{\text{Yukawa}}^{\Delta L=0}(jk) \right) \right] \right. \\
&- \frac{2}{3} \sum_{k=e,\mu,\tau} \eta_{\Delta L_k} \left[\sum_{\alpha=1}^3 \delta_{N_\alpha}^j \delta_{N_\alpha}^k \left(\Gamma_{\text{Yukawa}}^W(\alpha k) + \Gamma_{\text{Gauge}}^W(\alpha k) \right) \right. \\
&\quad \left. \left. + \Gamma_{\text{Yukawa}}^{\Delta L=2}(kj) - \Gamma_{\text{Yukawa}}^{\Delta L=0}(kj) \right] \right\}
\end{aligned}$$

Computational package: **LeptoGen**

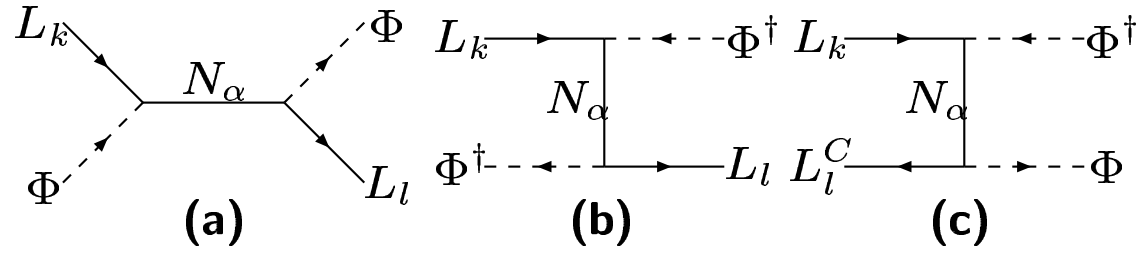
$\Delta L = 1$ scatterings involving L , N_i and quarks



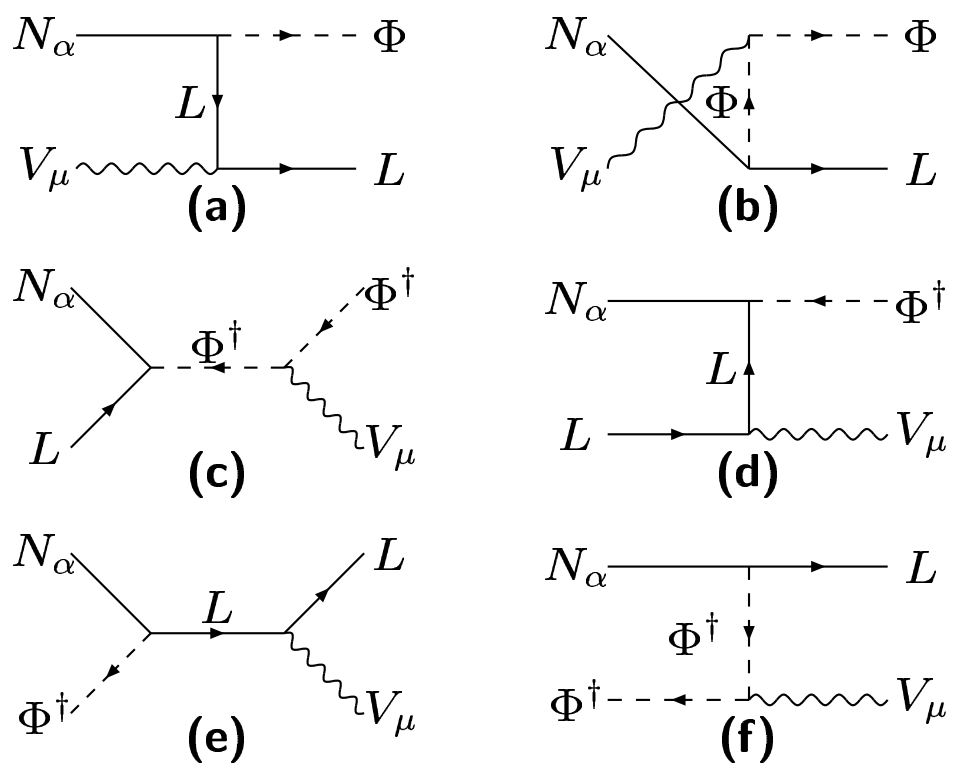
$\Delta L = 2$ scatterings involving L , Φ and N_α



$\Delta L = 0$ scatterings involving L , Φ and N_α

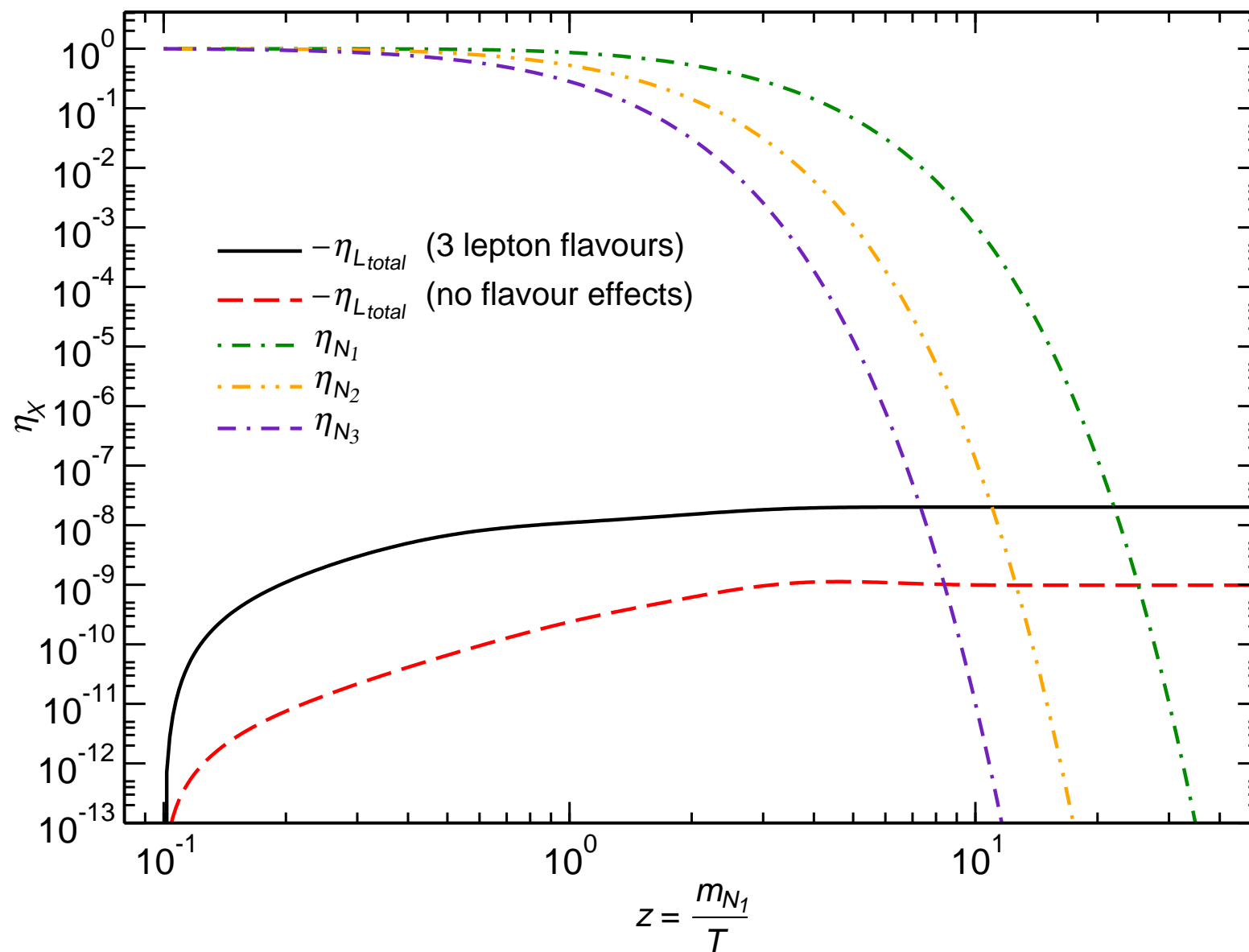


Gauge-mediated $\Delta L = 1$ scatterings



$m_{N_1} = 10^{10}$ GeV, $m_{N_2} = 2m_{N_1}$, $m_{N_3} = 3m_{N_1}$:

LeptoGen



• Resonant Leptogenesis

[A.P., PRD56 (1997) 5431; A.P. and T. Underwood, NPB692 (2004) 303;
inspired by the $K^0 \bar{K}^0$ system: T.D. Lee, R. Oehme, C.N. Yang, PR106 (1957) 340.]

RL is based on the observation that selfenergies dominate the lepton asymmetries if $|m_{N_1} - m_{N_2}| \ll m_{N_{1,2}}$.

[J. Liu, G. Segré, PRD48 (1993) 4609;
M. Flanz, E. Paschos, U. Sarkar, PLB345 (1995) 248;
L. Covi, E. Roulet, F. Vissani, PLB384 (1996) 169.]

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Several variants:

- Soft RL

[L. Boubekour, hep-ph/0208003 (unpublished);
Y. Grossman, T. Kashti, Y. Nir, E. Roulet, PRL91 (2003) 251801;
G. D'Ambrosio, G. F. Giudice, M. Raidal, PLB575 (2003) 75.]

- Radiative RL

[R. Gonzalez Felipe, F. R. Joaquim and B. M. Nobre, PRD70 (2004) 085009;
G. C. Branco, A. J. Buras, S. Jager, S. Uhlig, A. Weiler, JHEP0709 (2007) 004.]

- Coherent RL (via sterile neutrino oscillations)

[E. K. Akhmedov, V. A. Rubakov, A. Y. Smirnov, PRL81 (1998) 1359;
T. Asaka, M. Shaposhnikov, PLB620 (2005) 17;
M. E. Shaposhnikov, arXiv:0804.4542]

The Field-Theory of Resonant Leptogenesis:

LSZ-type formalism for mixing and decay of heavy Majorana neutrinos

[A.P., PRD56 (1997) 5431; NPB504 (1997) 61.]

$$\times [S_{ii}(\not{p})]^{-1} u_{N_i}(p) = \mathcal{T}_{N_i}$$

2- N Model:

$$S_{ij}(\not{p}) = \begin{pmatrix} \not{p} - m_{N_1} + \Sigma_{11}(\not{p}) & \Sigma_{12}(\not{p}) \\ \Sigma_{21}(\not{p}) & \not{p} - m_{N_2} + \Sigma_{22}(\not{p}) \end{pmatrix}^{-1}$$

[For 3- N mixing, see, A.P., T. Underwood, NPB692 (2004) 303.]

ϵ' -type CP violation :

$$\epsilon'_{N_i} = \frac{\text{Im} (h^{\nu\dagger} h^\nu)_{ij}^2}{(h^{\nu\dagger} h^\nu)_{ii} (h^{\nu\dagger} h^\nu)_{jj}} \left(\frac{\Gamma_{N_j}}{m_{N_j}} \right) f \left(\frac{m_{N_j}^2}{m_{N_i}^2} \right),$$

where

$$\Gamma_{N_j} = \frac{(h^{\nu\dagger} h^\nu)_{jj}}{8\pi} m_{N_j}$$

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$$\epsilon_{N_i} = \frac{\text{Im} (h^{\nu\dagger} h^\nu)_{ij}^2}{(h^{\nu\dagger} h^\nu)_{ii} (h^{\nu\dagger} h^\nu)_{jj}} \frac{(m_{N_i}^2 - m_{N_j}^2) m_{N_i} \Gamma_{N_j}}{(m_{N_i}^2 - m_{N_j}^2)^2 + m_{N_i}^2 \Gamma_{N_j}^2}$$

Note that $\epsilon_{N_{1,2}}$ are of the same sign!

Resonant conditions for $O(1)$ leptonic asymmetries:

[A.P., PRD56 (1997) 5431.]

$$\Rightarrow m_{N_2} - m_{N_1} \sim \frac{1}{2} \Gamma_{N_{1,2}}$$

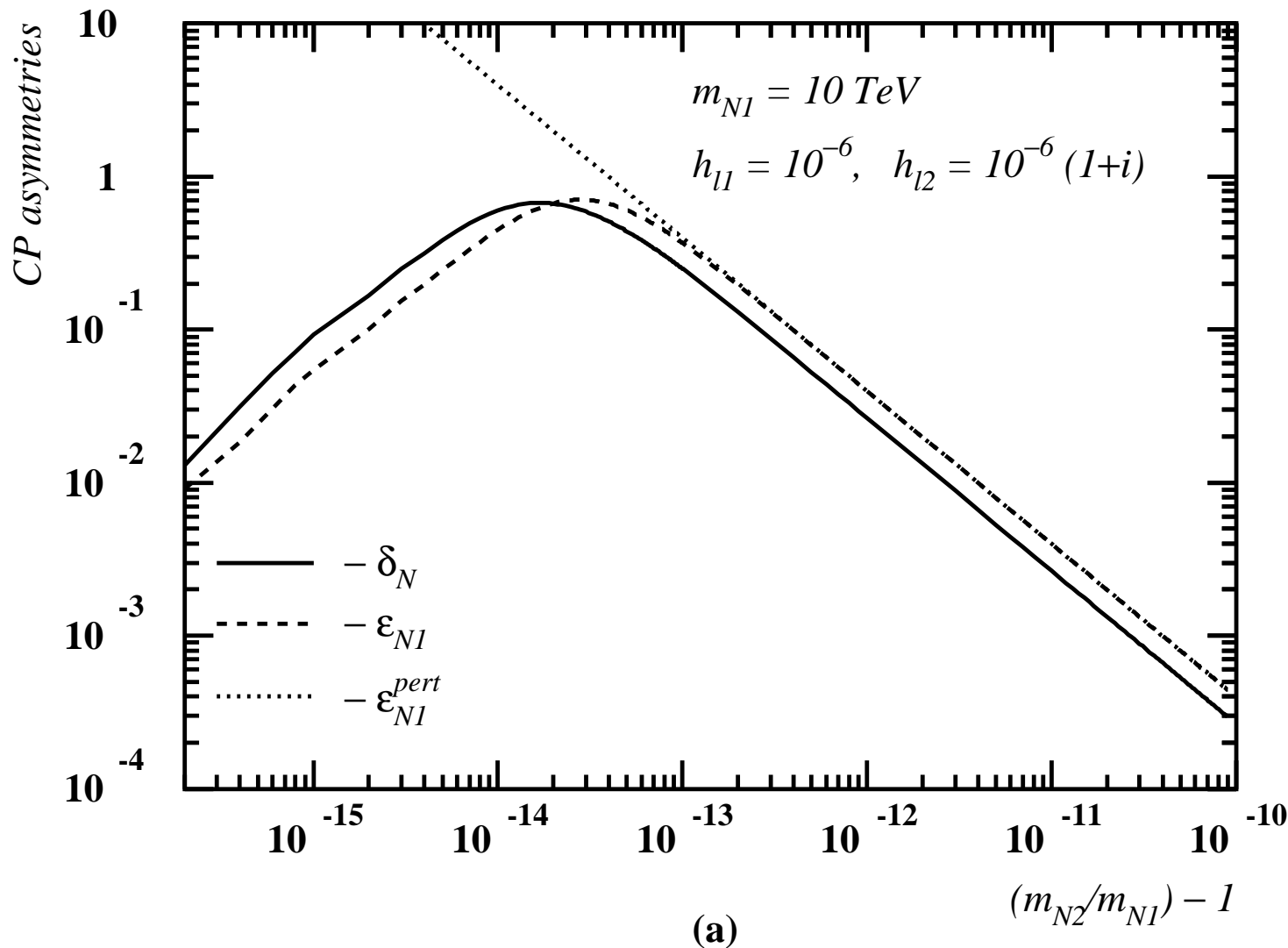
$$\Rightarrow \frac{\text{Im} (h^{\nu\dagger} h^\nu)_{ij}^2}{(h^{\nu\dagger} h^\nu)_{ii} (h^{\nu\dagger} h^\nu)_{jj}} \sim 1$$

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Flavour Effects in Resonant Leptogenesis

– The Non-Seesaw Paradigm

[A.P., PRL95 (2005) 081602 [hep-ph/0408103];

based on A.P., ZPC55 (1992) 275;

R.N. Mohapatra, J.W.F. Valle, PRD34 (1986) 1642.]

Break $SO(3)$ and $U(1)_l$ flavour symmetries:

$$SO(3) \xrightarrow{\sim h_\tau} SO(2) \simeq U(1)_l \xrightarrow{\sim h_e} \mathbf{1}$$

$U_l(1)$ -broken Yukawa sector:

$$m_D = \frac{v}{\sqrt{2}} \begin{pmatrix} \varepsilon_e & a e^{-i\pi/4} & a e^{i\pi/4} \\ \varepsilon_\mu & b e^{-i\pi/4} & b e^{i\pi/4} \\ \varepsilon_\tau & c e^{-i\pi/4} & c e^{i\pi/4} \end{pmatrix},$$

with $a \sim b \sim 10^{-2} \sim h_\tau$, $c \lesssim 10^{-4}$ & $|\varepsilon_l| \sim 10^{-7} \sim h_e$.

$$\implies m_\nu^{\text{light}} \sim \frac{m_e^2}{m_N} \sim 0.1 \text{ eV} \implies m_N \sim 100 - 500 \text{ GeV}$$

\implies 3 nearly degenerate heavy Majorana neutrinos.

Light neutrino-mass spectrum:

[A.P., T. Underwood, PRD72 (2005) 113001.]

$$m_{\nu}^{\text{light}} = \frac{v^2}{2m_N} \begin{pmatrix} \frac{\Delta m_N}{m_N} a^2 - \varepsilon_e^2 & \frac{\Delta m_N}{m_N} ab - \varepsilon_e \varepsilon_{\mu} & \frac{\Delta m_N}{m_N} ac - \varepsilon_e \varepsilon_{\tau} \\ \frac{\Delta m_N}{m_N} ab - \varepsilon_e \varepsilon_{\mu} & \frac{\Delta m_N}{m_N} b^2 - \varepsilon_{\mu}^2 & \frac{\Delta m_N}{m_N} bc - \varepsilon_{\mu} \varepsilon_{\tau} \\ \frac{\Delta m_N}{m_N} ac - \varepsilon_e \varepsilon_{\tau} & \frac{\Delta m_N}{m_N} bc - \varepsilon_{\mu} \varepsilon_{\tau} & \frac{\Delta m_N}{m_N} c^2 - \varepsilon_{\tau}^2 \end{pmatrix},$$

where

$$\Delta m_N = 2(\Delta m_M)_{23} + i[(\Delta m_M)_{33} - (\Delta m_M)_{22}], \quad \frac{b}{a} = \frac{19}{50},$$

and (in $\sim 10^{-7}$ units)

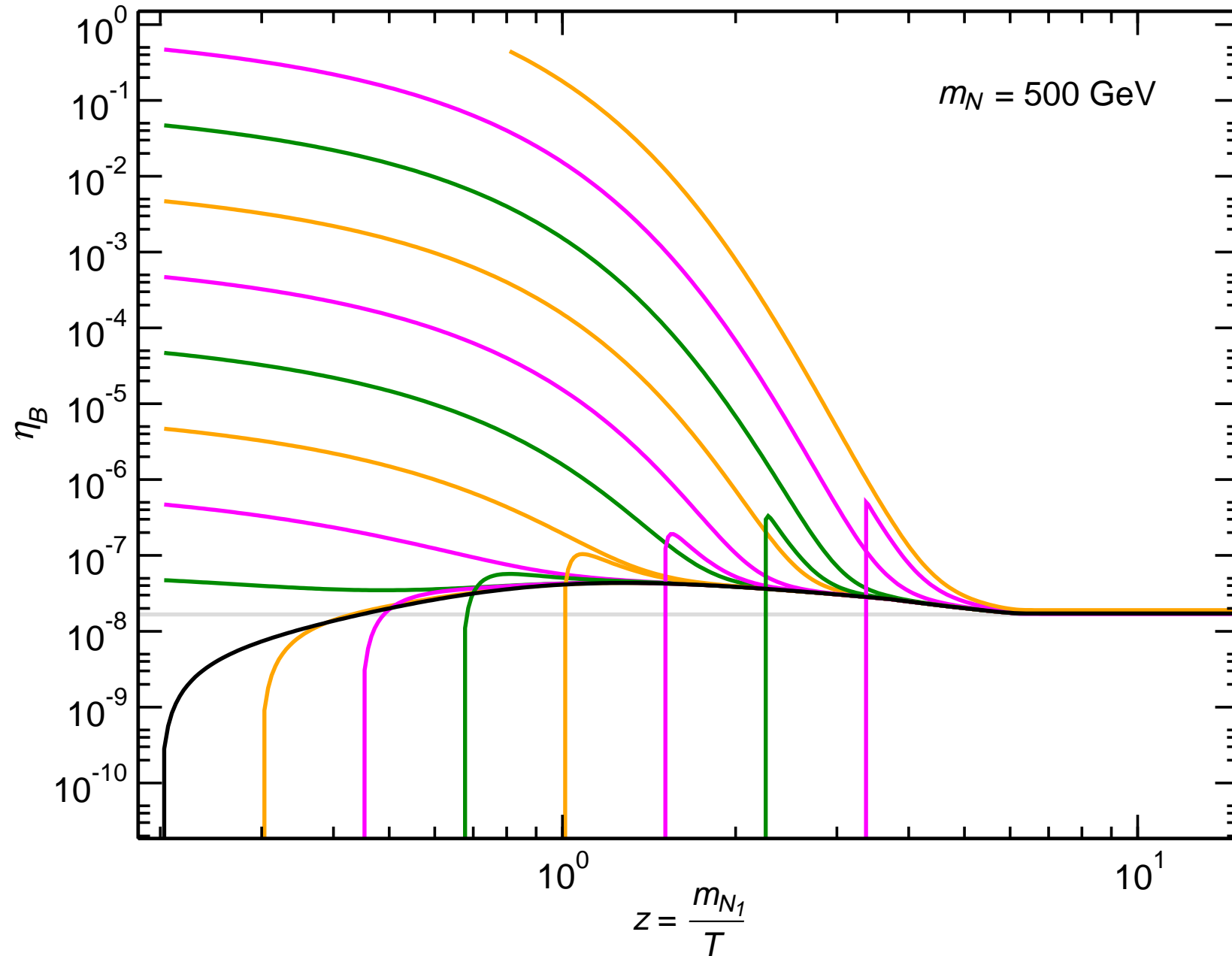
$$\sqrt{\frac{\Delta m_N}{m_N}} a = 2, \quad \varepsilon_e = 2 + \frac{21}{250}, \quad \varepsilon_{\mu} = \frac{13}{50}, \quad \varepsilon_{\tau} = -\frac{49}{128}.$$

Prediction: inverted mass hierarchy, $m_{\nu_3} < m_{\nu_1} < m_{\nu_2}$, with

$$m_{\nu_2}^2 - m_{\nu_1}^2 = 7.54 \times 10^{-5} \text{ eV}^2, \quad m_{\nu_1}^2 - m_{\nu_3}^2 = 2.45 \times 10^{-3} \text{ eV}^2, \\ \sin^2 \theta_{12} = 0.362, \quad \sin^2 \theta_{23} = 0.341, \quad \sin^2 \theta_{13} = 0.047.$$

Resonant Leptogenesis

[A.P. and T. Underwood, PRD72 (2005) 113001.]



Qualification of Models for Baryogenesis

Q-Factor (a measure of the dependence of the BAU on initial conditions):

$$Q = \ln \left| \frac{\delta\eta_B^{\text{in}}}{\delta\eta_B^{\text{fin}}} \right| .$$

T_{in} is a typical initial temperature of the baryogenesis mechanism,
e.g. $T_{\text{in}} \sim m_N$ or $T_{\text{in}} \sim T_c^{\text{EW}}$.

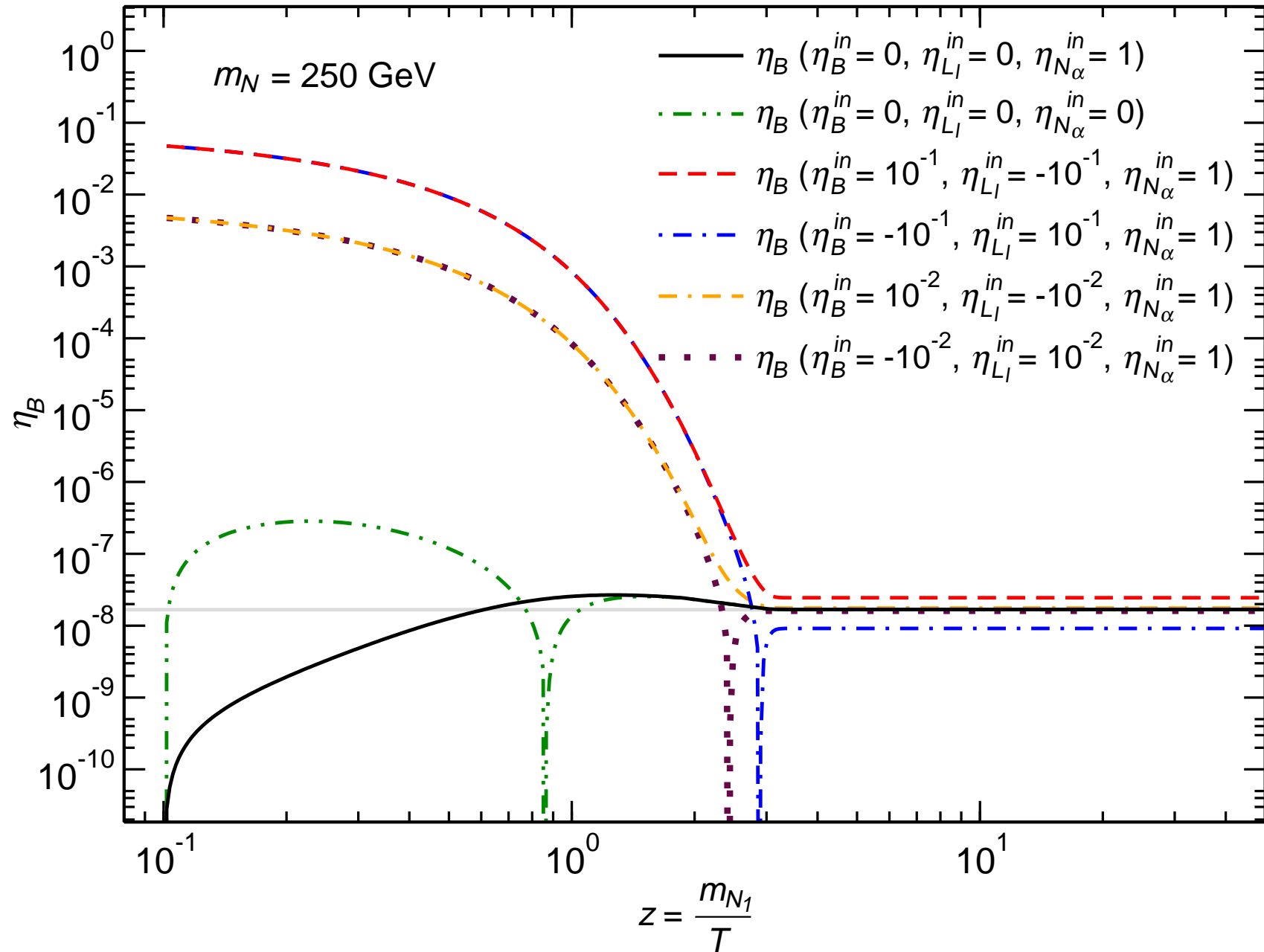
T_{fin} is the freeze-out temperature of the baryogenesis mechanism.

If $Q \gg 1 \implies$ no-strong dependence of the BAU on the initial conditions.

If $Q < 0 \implies$ BAU does strongly depend on the initial conditions.

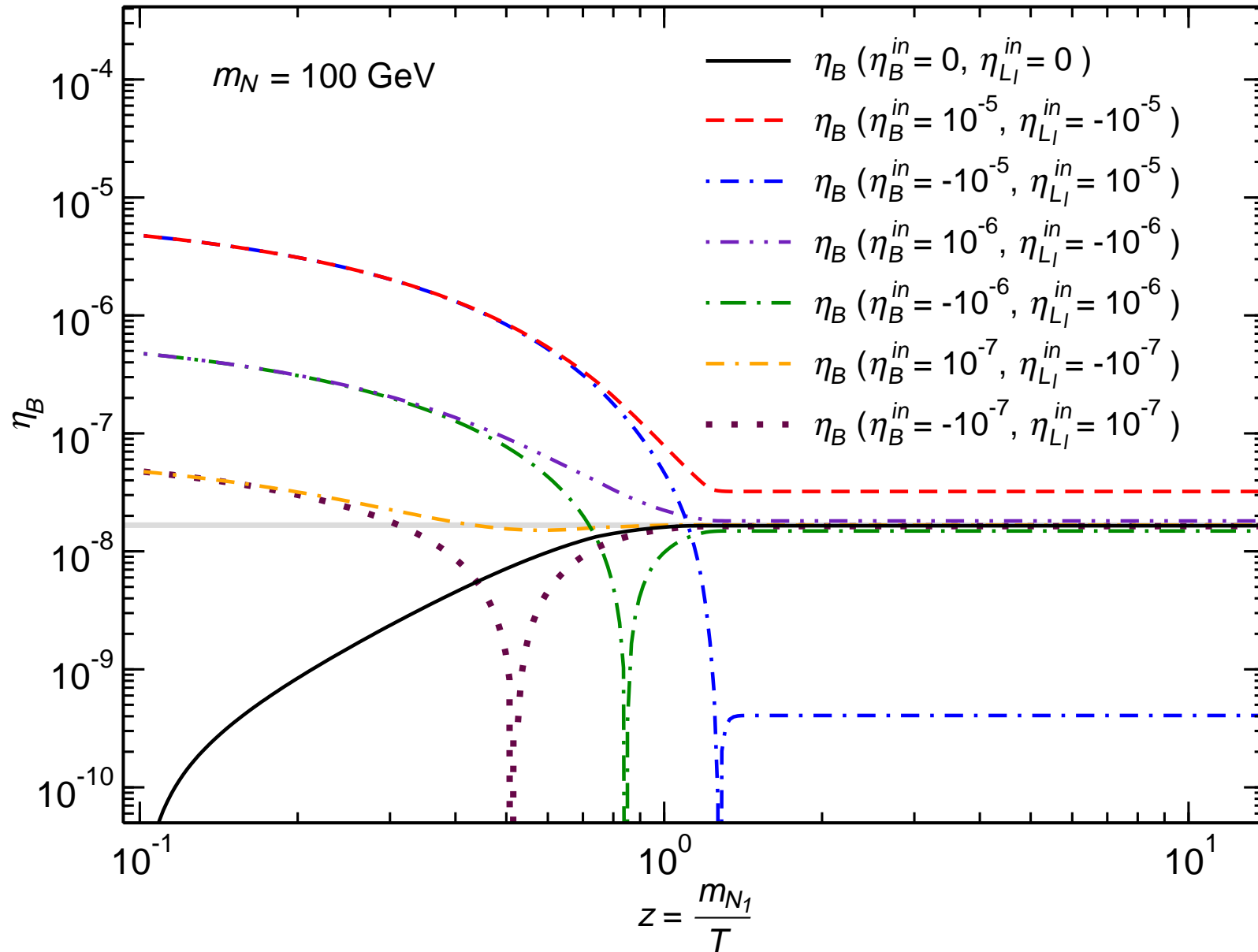
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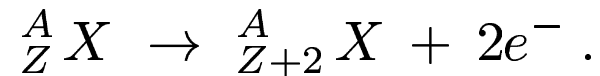


Calculation must be done in the broken phase

[A.P., PRD78 (2008) 013008]

• Phenomenology of **RL** Models

• $0\nu\beta\beta$ Decay



Half-life for $0\nu\beta\beta$ decay:

$$[T_{1/2}^{0\nu\beta\beta}]^{-1} = \frac{|\langle m \rangle|^2}{m_e^2} |\mathcal{M}_{0\nu\beta\beta}|^2 G_{01} .$$

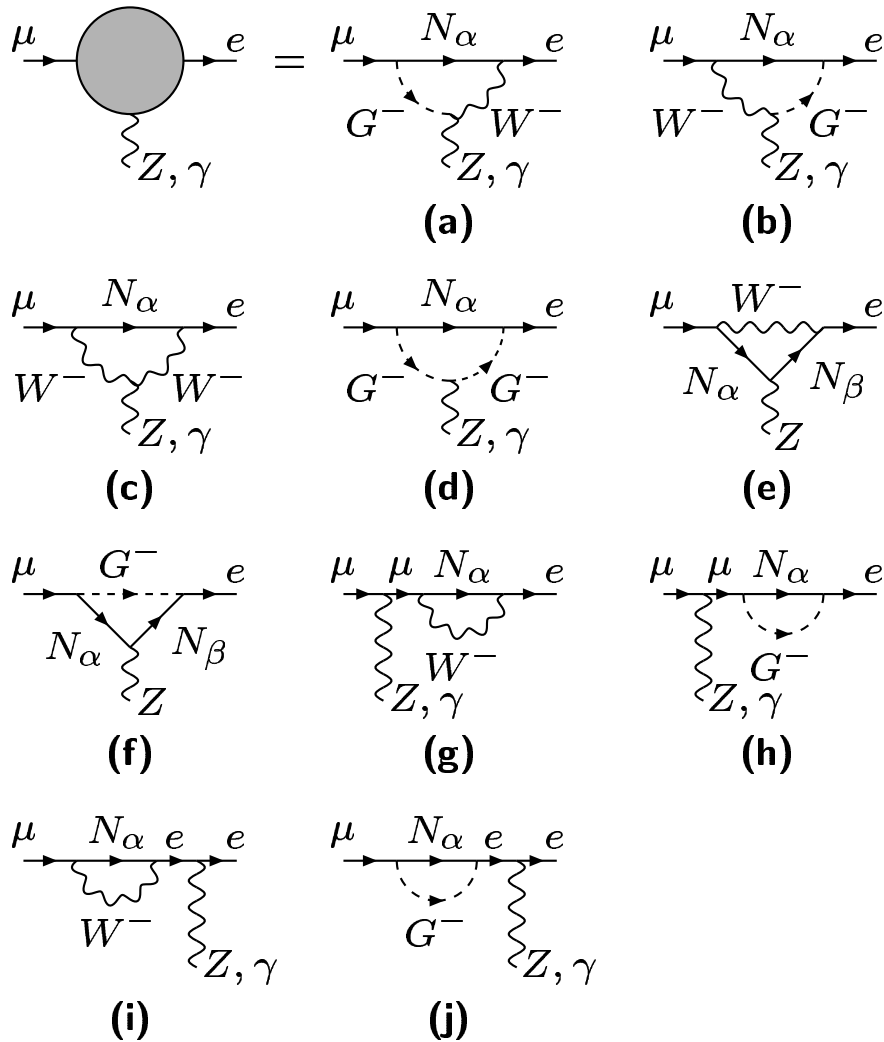
RTL realizes **inverted** light-neutrino hierarchy, with:

$$|\langle m_{0\nu\beta\beta} \rangle| = |(\mathbf{m}^\nu)_{ee}| = \frac{v^2}{2m_N} \left| \frac{\Delta m_N}{m_N} a^2 - \epsilon_e^2 \right| \approx 0.013 \text{ eV} .$$

Future $0\nu\beta\beta$ experiments will be sensitive to $|\langle m \rangle| \sim 0.01\text{--}0.05 \text{ eV}$.

• $\mu \rightarrow e\gamma$

[T.P. Cheng, L.F. Li, PRL45 (1980) 1908.]



For $h_{eN_{2,3}}^\nu = h_{\mu N_{2,3}}^\nu = 8 \times 10^{-3}$
and $m_N = 250$ GeV:

$$B(\mu \rightarrow e\gamma)$$

$$\sim 7 \cdot 10^{-4} \times \frac{(h_{eN}^\nu h_{\mu N}^\nu)^2 v^4}{m_N^4}$$

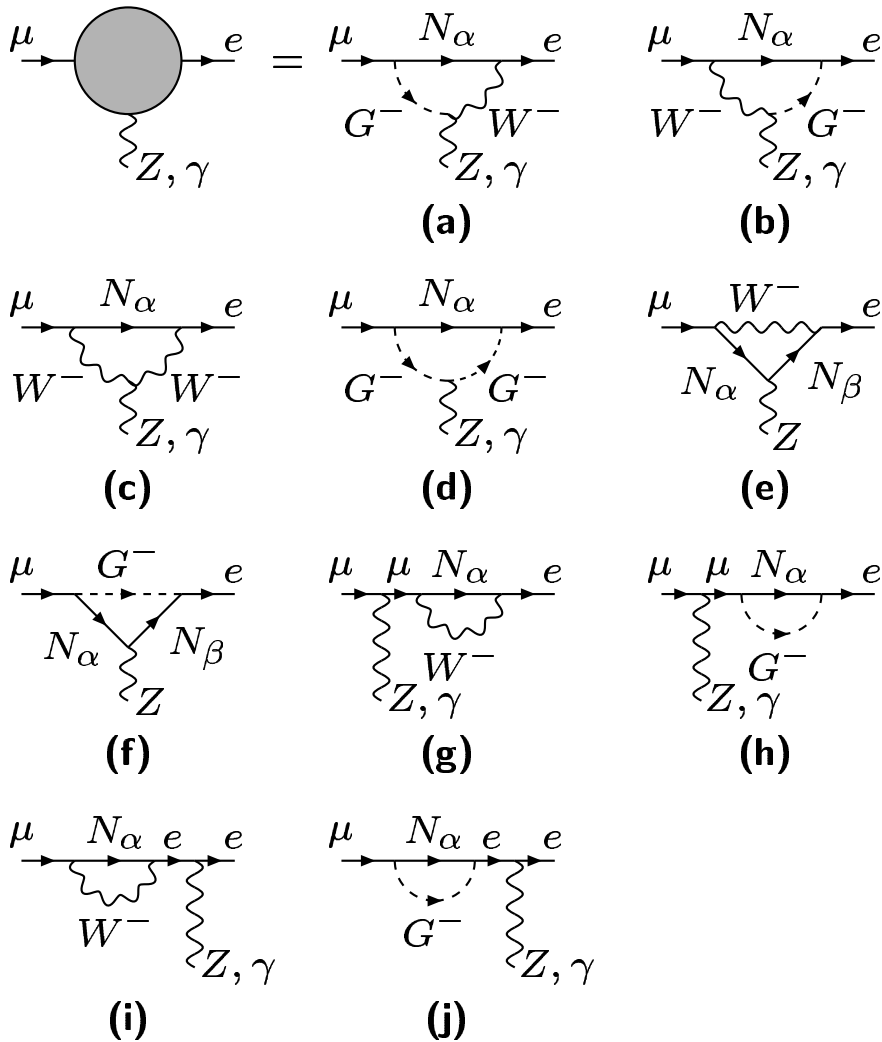
$$\sim 10^{-12} .$$

MEG sensitivity:

$$B(\mu \rightarrow e\gamma) \sim 10^{-13} - 10^{-14} .$$

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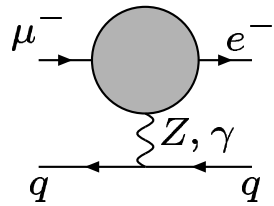
$$B(\mu \rightarrow e\gamma) \sim 10^{-13} - 10^{-14} .$$

Other effects at the Z -boson pole, such as **breaking** of **lepton universality** are **small**.

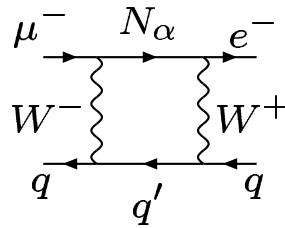
[J. Bernabéu, J.G. Körner, A.P., K. Schilcher, PRL71 (1993) 2695;
J. Bernabéu, A. Santamaria, J. Vidal, A. Mendez, J.W.F. Valle, PLB187 (1987) 303;
J.G. Körner, A.P., K. Schilcher, PLB300 (1993) 381.]

• **Coherent $\mu \rightarrow e$ Conversion in Nuclei (${}^{48}_{22}\text{Ti}$)**

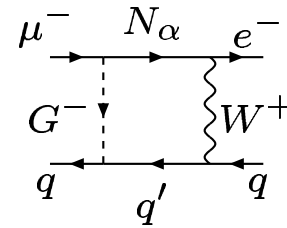
[A.P., T. Underwood, PRD72 (2005) 113001;
hadronic part from R. Pla, J. Bernabéu, An. Fís. 67 (1971) 455.]



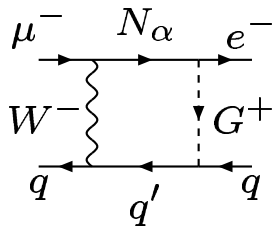
(a)



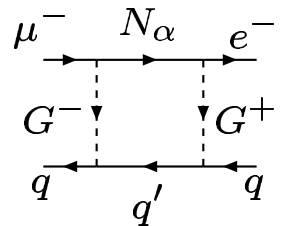
(b)



(c)



(d)



(e)

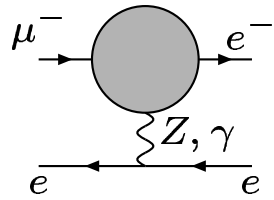
$m_N = 250$ GeV:

$$B(\mu \rightarrow e) \approx 0.5 \times B(\mu \rightarrow e\gamma) \sim 5 \times 10^{-13}.$$

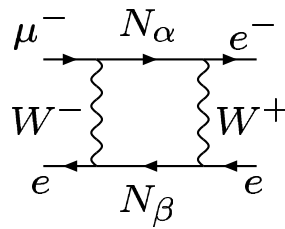
PRISM/PRIME ... will be sensitive to $B(\mu \rightarrow e) \sim 10^{-13}$ – 10^{-18} .

• $\mu \rightarrow eee$

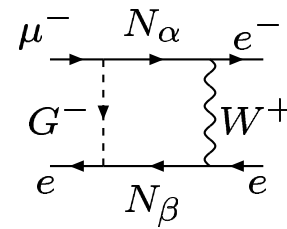
[A. Ilakovac, A.P., NPB437 (1995) 491.]



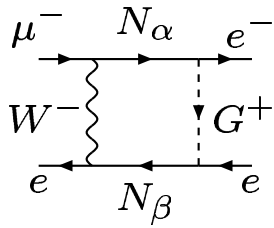
(a)



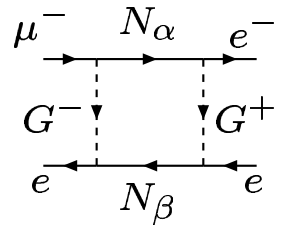
(b)



(c)



(d)



(e)

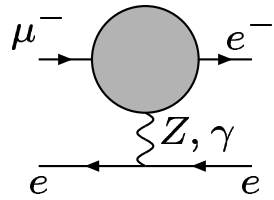
+ ($e \leftrightarrow e^-$)

$m_N = 250 \text{ GeV}$:

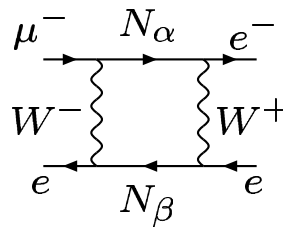
$$B(\mu \rightarrow eee) \approx 1.4 \cdot 10^{-2} \times B(\mu \rightarrow e\gamma) \sim 1.4 \times 10^{-14} .$$

• $\mu \rightarrow eee$

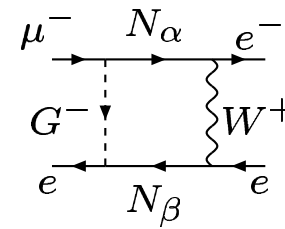
[A. Ilakovac, A.P., NPB437 (1995) 491.]



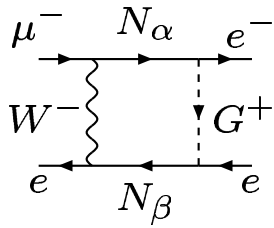
(a)



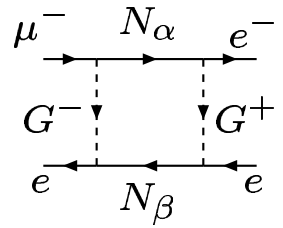
(b)



(c)



(d)



(e)

+ ($e \leftrightarrow e^-$)

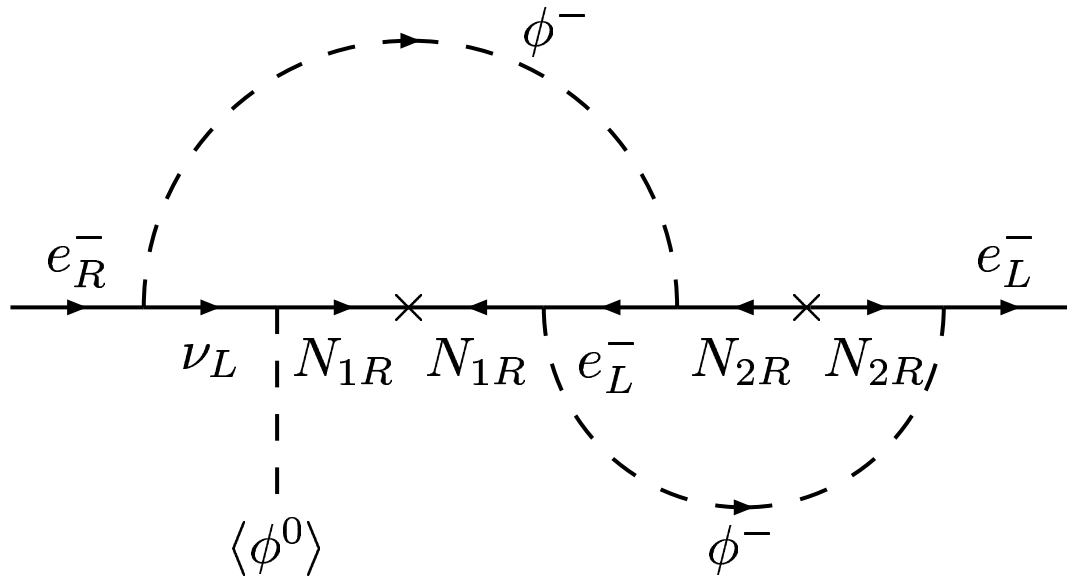
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No new experiment proposed yet!

• Electric Dipole Moment (EDM) Predictions

[A.P., IJMPA14 (1999) 1811; J.P. Archaumbault, A. Czarnecki, M. Pospelov, PRD70 (2004) 073006.]



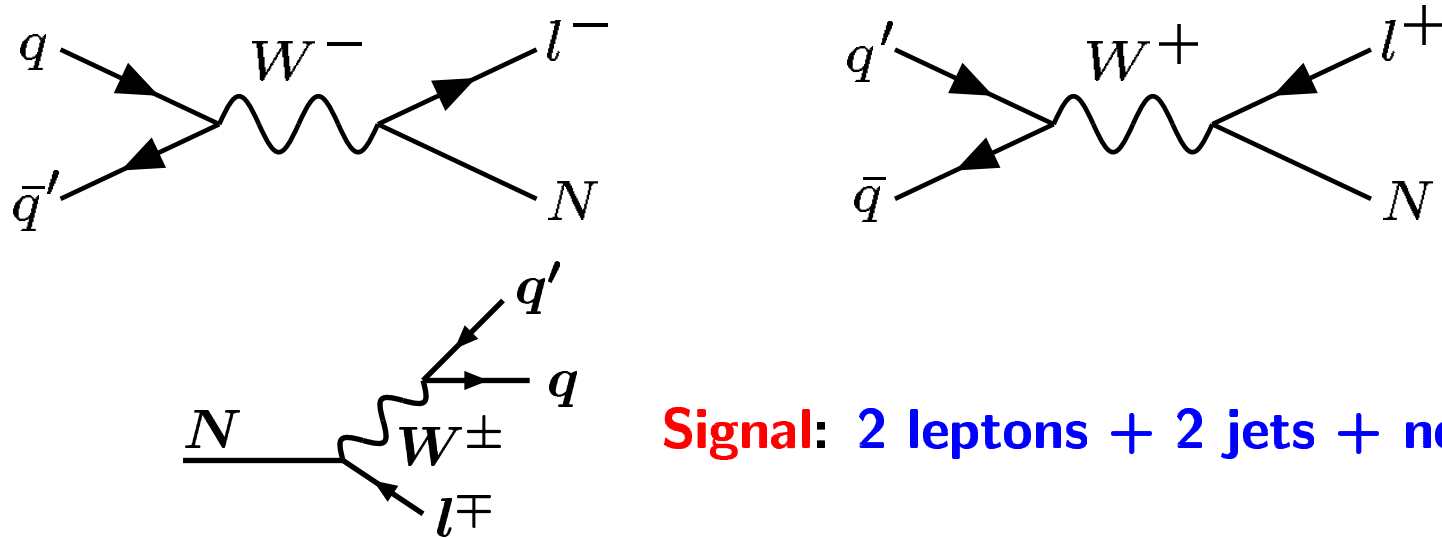
$$\frac{d_e}{e} \sim \frac{1}{(16\pi^2)^2} \frac{m_e}{m_{N_1}^2} \frac{\text{Im}(h^{\nu\dagger} h^\nu)_{12}^2 m_{N_1} m_{N_2} (m_{N_2}^2 - m_{N_1}^2)}{(m_{N_1}^2 + m_{N_2}^2)^2} \ln \frac{m_{N_1}}{M_W}$$

$$\sim (10^{-25} \text{ cm}) \times \frac{M_W^2}{m_{N_1} m_{N_2}} \text{Im}(h^{\nu\dagger} h^\nu)_{12}^2 \lesssim 10^{-32} \text{ cm}$$

\therefore Observation of an electron EDM $d_e \gtrsim 10^{-32} e \cdot \text{cm}$ will rule out leptogenesis in the SM with singlet neutrinos.

• Heavy Majorana Neutrino Production at the LHC

[A.P., ZPC55 (1992) 275; A. Datta, M. Guchait, A.P., PRD50 (1994) 3195;
 T. Han, B. Zhang, PRL97 (2006) 171804;
 F. del Aguila, J. A. Aguilar-Saavedra, R. Pittau, JHEP0710 (2007) 047.]

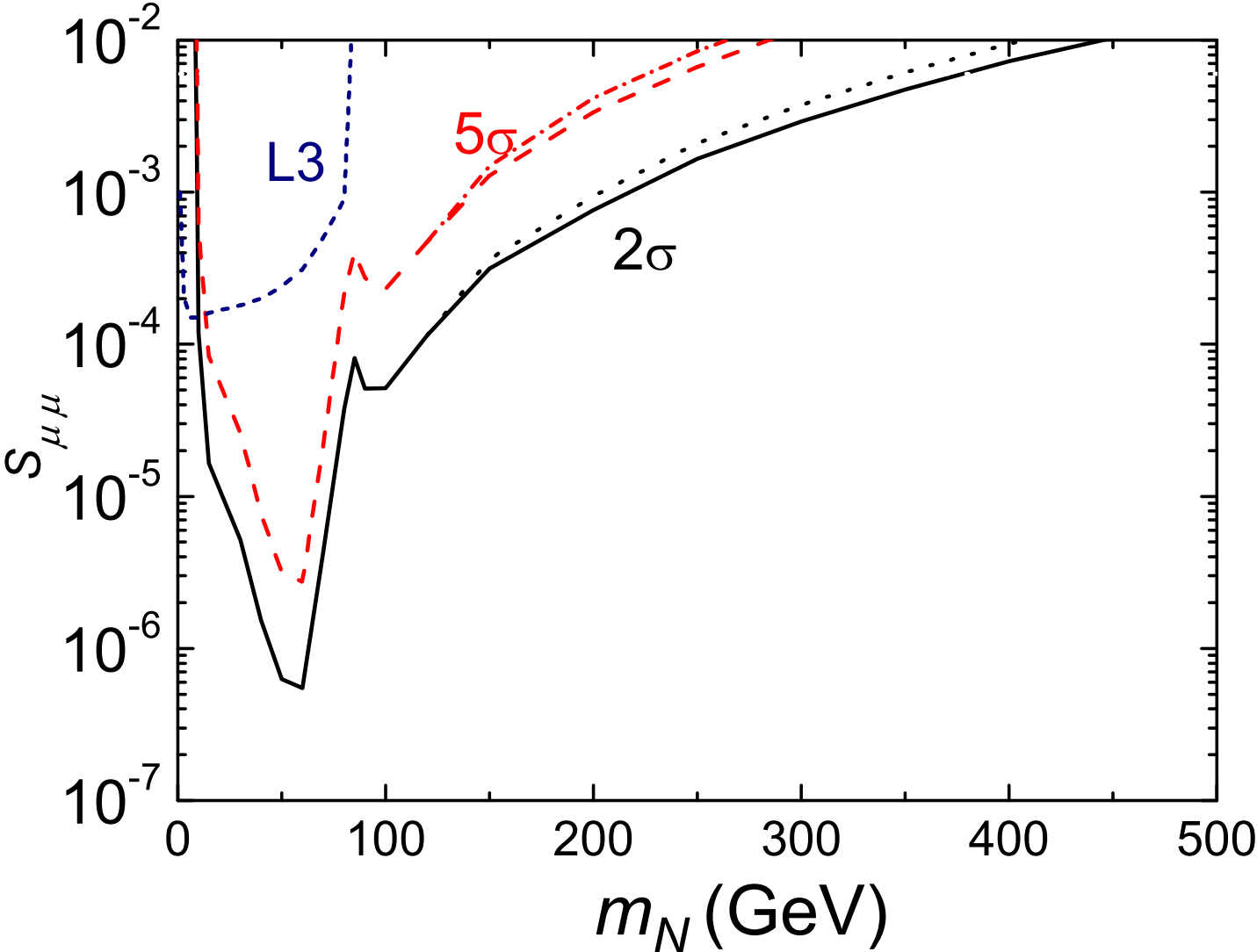


Signal: 2 leptons + 2 jets + no \cancel{p}_T

- **LVN signatures:** $pp \rightarrow e^+e^+, e^+\mu^+, e^-e^-, e^-\mu^-, e^-\tau^- \dots$
- **LFV signatures:** $pp \rightarrow e^+\mu^-, e^-\mu^+, e^-\tau^+ \dots$
- **CP Asymmetries**

[S. Bray, J.S. Lee, A.P., NPB786 (2007) 95.]

LHC sensitivity with 100 fb^{-1}



[T. Han, B. Zhang, PRL97 (2006) 171804.]

- Lepton Number Violation:**

$$A_{\text{CP}}(\text{LNV1}) = \frac{\sigma(pp \rightarrow e^+e^+W^-X) - K\sigma(pp \rightarrow e^-e^-W^+X)}{\sigma(pp \rightarrow e^+e^+W^-X) + K\sigma(pp \rightarrow e^-e^-W^+X)},$$

$$A_{\text{CP}}(\text{LNV2}) = \frac{\sigma(pp \rightarrow e^+\mu^+W^-X) - K\sigma(pp \rightarrow e^-\mu^-W^+X)}{\sigma(pp \rightarrow e^+\mu^+W^-X) + K\sigma(pp \rightarrow e^-\mu^-W^+X)},$$

$$R_{\text{CP}}(\text{LNV}) = \frac{\frac{\sigma(pp \rightarrow e^+e^+W^-X)}{\sigma(pp \rightarrow e^+\mu^+W^-X)} - \frac{\sigma(pp \rightarrow e^-e^-W^+X)}{\sigma(pp \rightarrow e^-\mu^-W^+X)}}{\frac{\sigma(pp \rightarrow e^+e^+W^-X)}{\sigma(pp \rightarrow e^+\mu^+W^-X)} + \frac{\sigma(pp \rightarrow e^-e^-W^+X)}{\sigma(pp \rightarrow e^-\mu^-W^+X)}}.$$

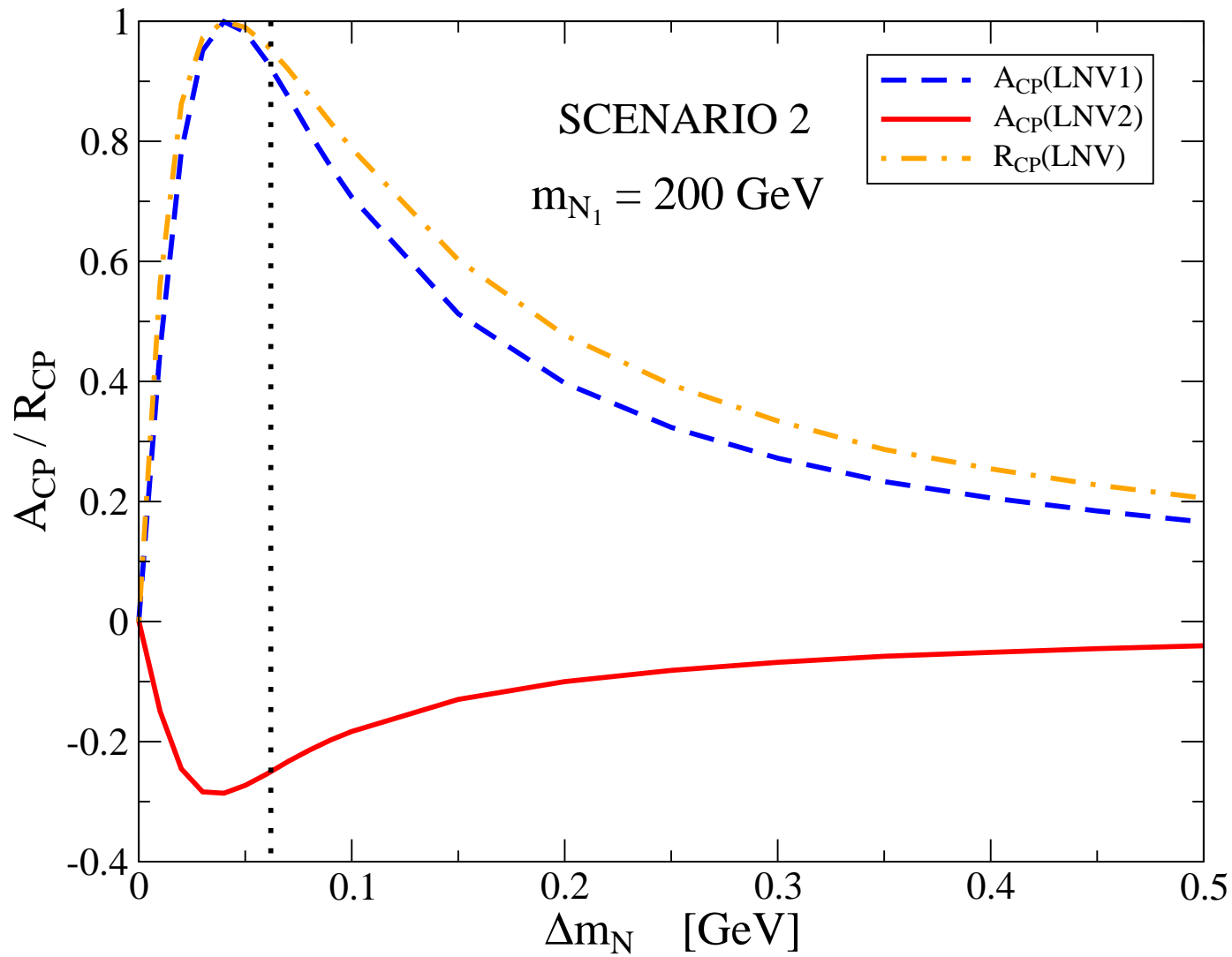
- Lepton Flavour Violation:**

$$A_{\text{CP}}(\text{LNC}) = \frac{\sigma(pp \rightarrow e^+\mu^-W^\pm X) - \sigma(pp \rightarrow e^-\mu^+W^\pm X)}{\sigma(pp \rightarrow e^+\mu^-W^\pm X) + \sigma(pp \rightarrow e^-\mu^+W^\pm X)},$$

$$R_{\text{CP}}(\text{LNC}) = \frac{\frac{\sigma(pp \rightarrow e^+\mu^-W^\pm X)}{\sigma(pp \rightarrow e^-\mu^+W^\pm X)} - \frac{\sigma(pp \rightarrow e^-\mu^+W^\pm X)}{\sigma(pp \rightarrow e^+\mu^-W^\pm X)}}{\frac{\sigma(pp \rightarrow e^+\mu^-W^\pm X)}{\sigma(pp \rightarrow e^-\mu^+W^\pm X)} + \frac{\sigma(pp \rightarrow e^-\mu^+W^\pm X)}{\sigma(pp \rightarrow e^+\mu^-W^\pm X)}}.$$

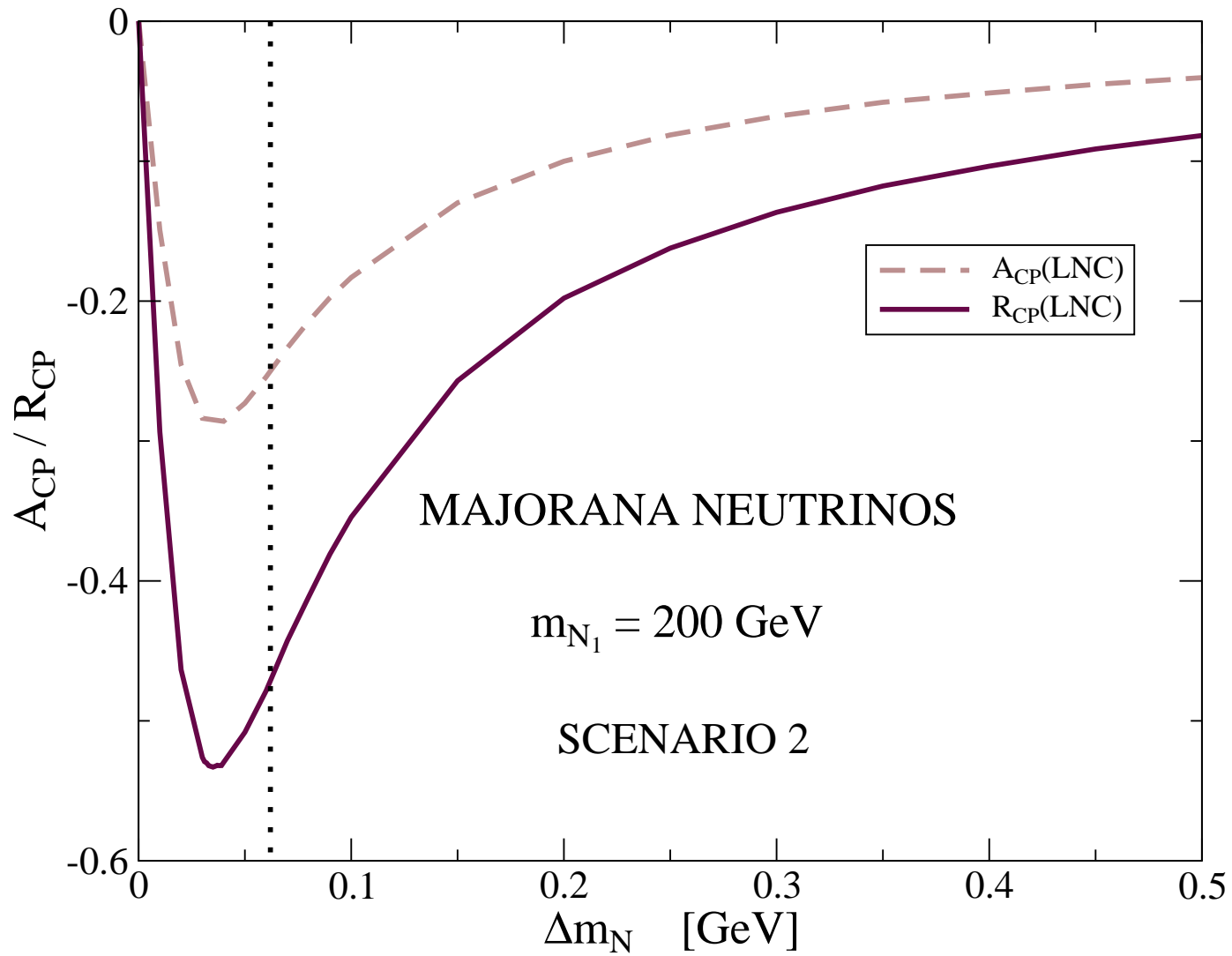
Resonant CP Violation through Mixing of Heavy Majorana Neutrinos

[S. Bray, J.S. Lee, A.P., NPB786 (2007) 95.]



Resonant CP Violation through Mixing of Heavy Majorana Neutrinos

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$$B(\mu \rightarrow eee) \approx 1.4 \cdot 10^{-2} \times B(\mu \rightarrow e\gamma),$$

$$|\langle m_{0\nu\beta\beta} \rangle| \approx 0.01 \text{ eV}.$$