

Rules in S'Cool LAB

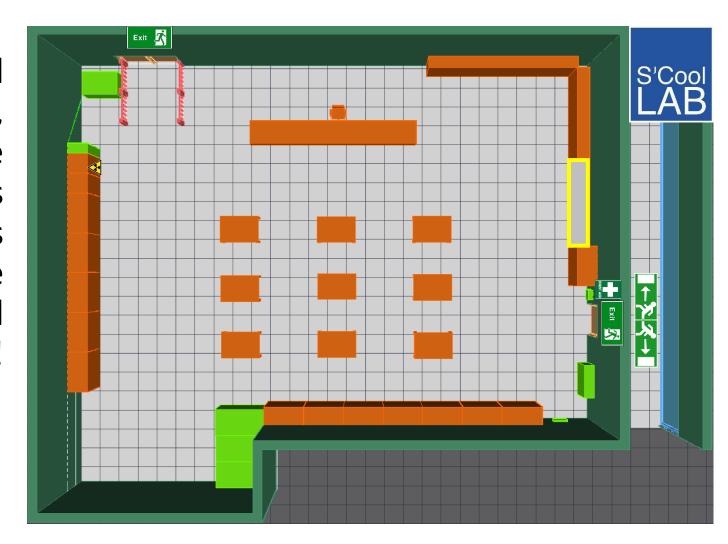




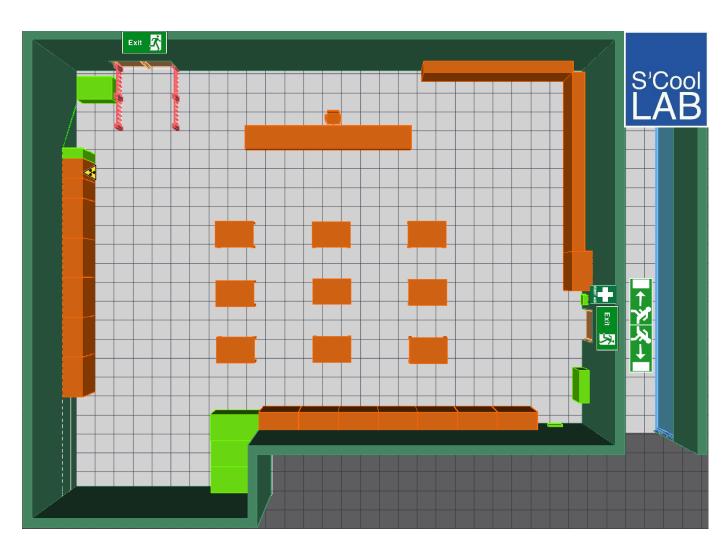


Bags

To avoid collisions, please place your bags and jackets in the designated shelf!



Emergency exits



Assembly point



Rest rooms



Cloud Chamber Workshop

Outline

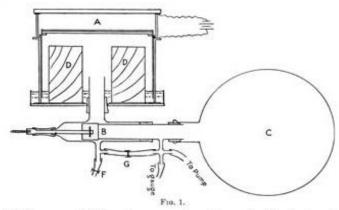
- History
- Step by step tutorial
- Build your own particle detector
- Tidying up
- Discussion and explanations

History

History

Charles T. R. Wilson (1869 - 1959)

This Scottish physicist perfected the first (expansion) cloud chamber in 1911 and received the Nobel Prize in 1927.

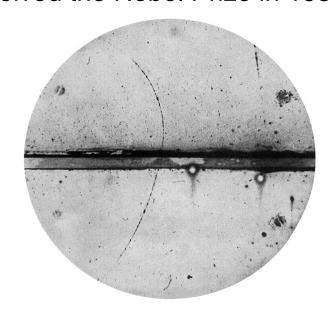


A diagram of Wilson's apparatus. The cylindrical cloud chamber ('A') is 16.5cm across by 3.4cm deep.

C. T. R. WILSON: On an Expansion Apparatus for Making Visible the Tracks of Ionising Particles in Gases and Some Results Obtained by Its Use. Proc. R. Soc. Lond. A. 1912 87 277-292 DOI:10.1098/rspa.1912.0081

Carl Anderson (1905 - 1991)

This physicist discovered the positron in 1932 and the muon in 1936 using a cloud chamber. He received the Nobel Prize in 1936.



Carl D. Anderson (1905–1991) - Anderson, Carl D. (1933). "The Positive Electron". Physical Review 43 (6): 491–494. <u>DOI:10.1103/PhysRev.43.491</u>.

Step by step tutorial







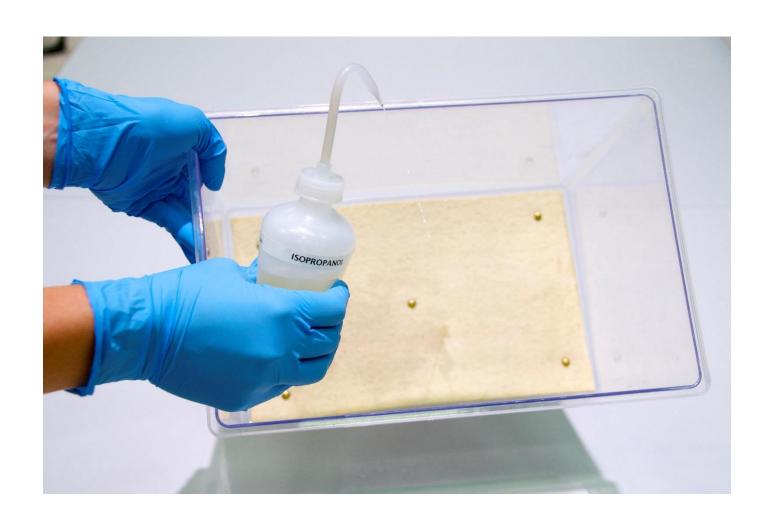






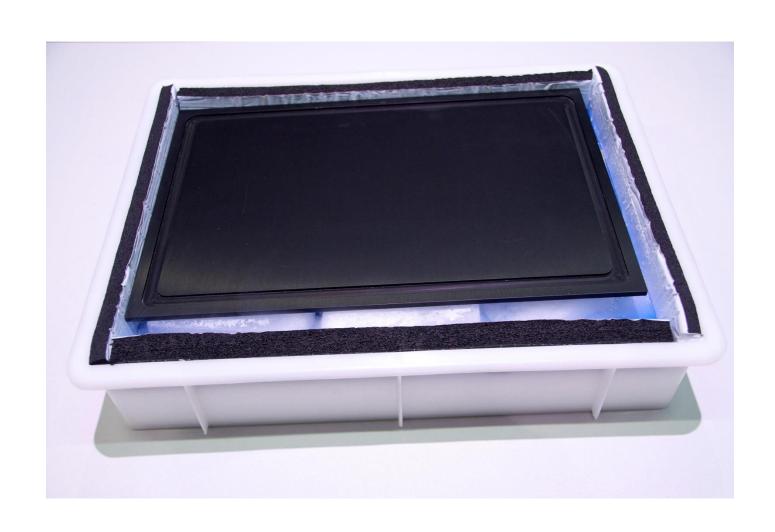










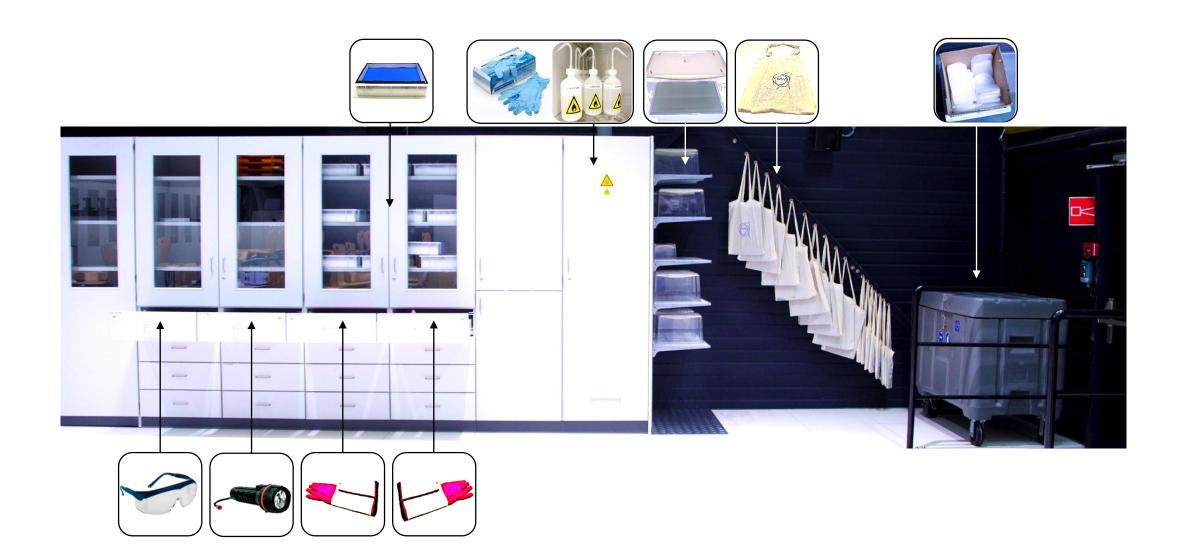








Build your own particle detector!

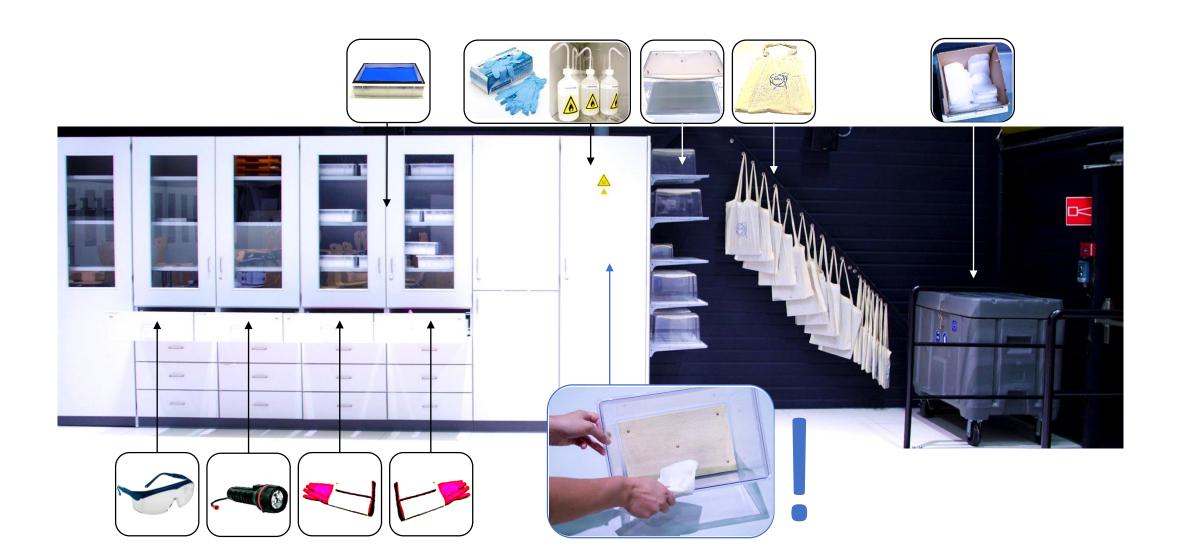


Build your own particle detector!

Tasks

- Observe your Cloud Chamber
- Find the optimal torch position and the optimal observation position
- Describe visible tracks (shape, length, width, ...)
- Discuss the reason for these tracks
- Count the number of tracks you can see for 1 minute, repeat this measurement 2 times

Tidying up



Discussion and explanations

Whiteboard

Additional Material

Air Shower Simulation

Cosmic Ray Air Shower Pictures

by H.-J. Drescher <u>drescher@th.physik.uni-frankfurt.de</u>.

Air showers are cascades of secondary particles induced in the atmosphere by high energy cosmic rays. What you see here is a **visualisation of realistic simulations of these showers**. Of course, not all of the particles in a shower are displayed, there are far too many! The **fraction displayed here is about 1e-6**, sampled with a **thinning algorithm**.

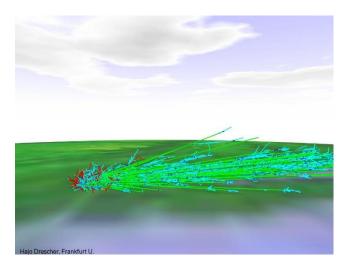
blue:electrons/positrons

cyan:photons

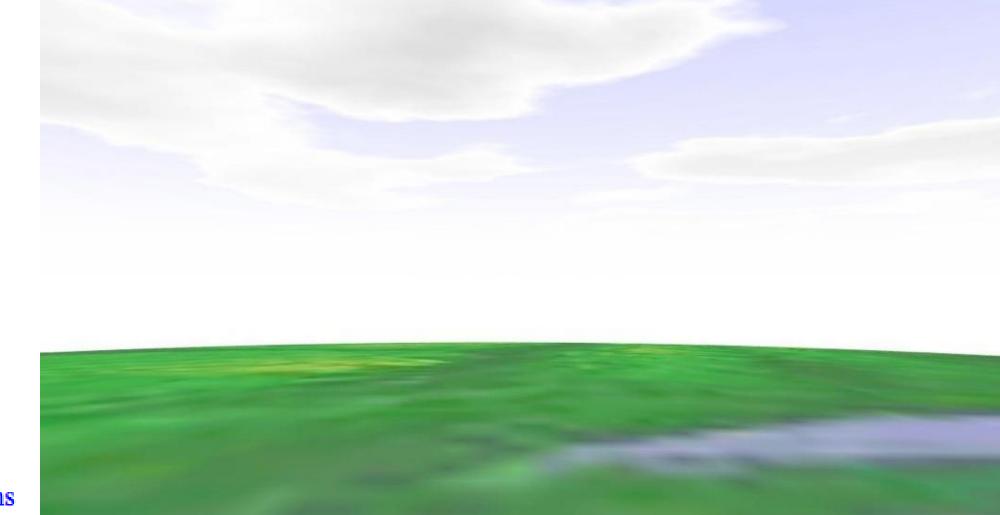
red:neutrons

orange: protons

gray: mesons



http://th.physik.uni-frankfurt.de/~drescher/CASSIM/



cyan:photons

red:neutrons

orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U. time = -1000 μs



cyan:photons red:neutrons

orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U. time = -900 μs



cyan:photons red:neutrons

orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U. time = -800 μs



cyan:photons red:neutrons

orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U. time = -700 μs



cyan:photons red:neutrons

orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U. time = -600 μs



Hajo Drescher, Frankfurt U.

cyan:photons

red:neutrons

orange: protons

gray: mesons

green:muons

time = -500 μs



cyan:photons red:neutrons

orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U. time = -400 μs



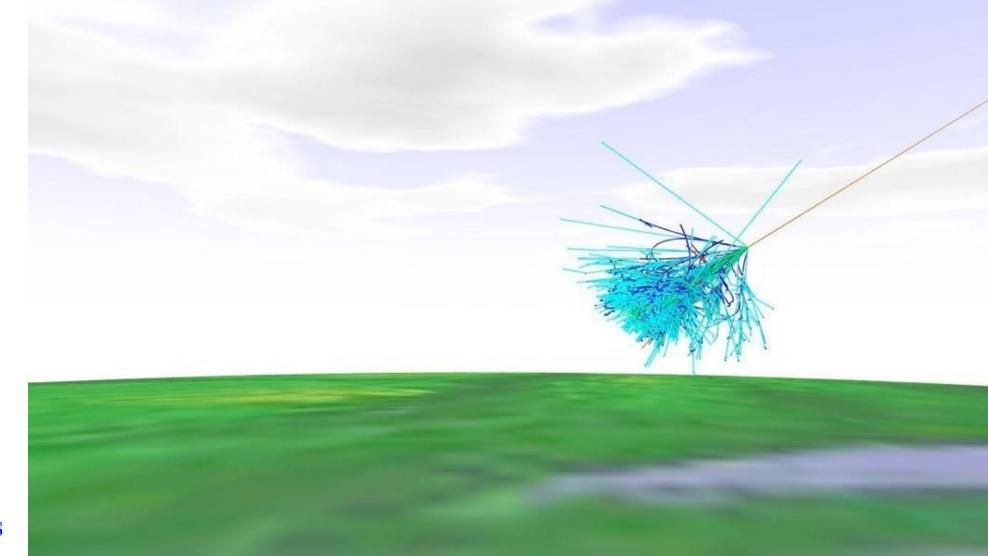
cyan:photons

red:neutrons

orange: protons

gray: mesons





cyan:photons

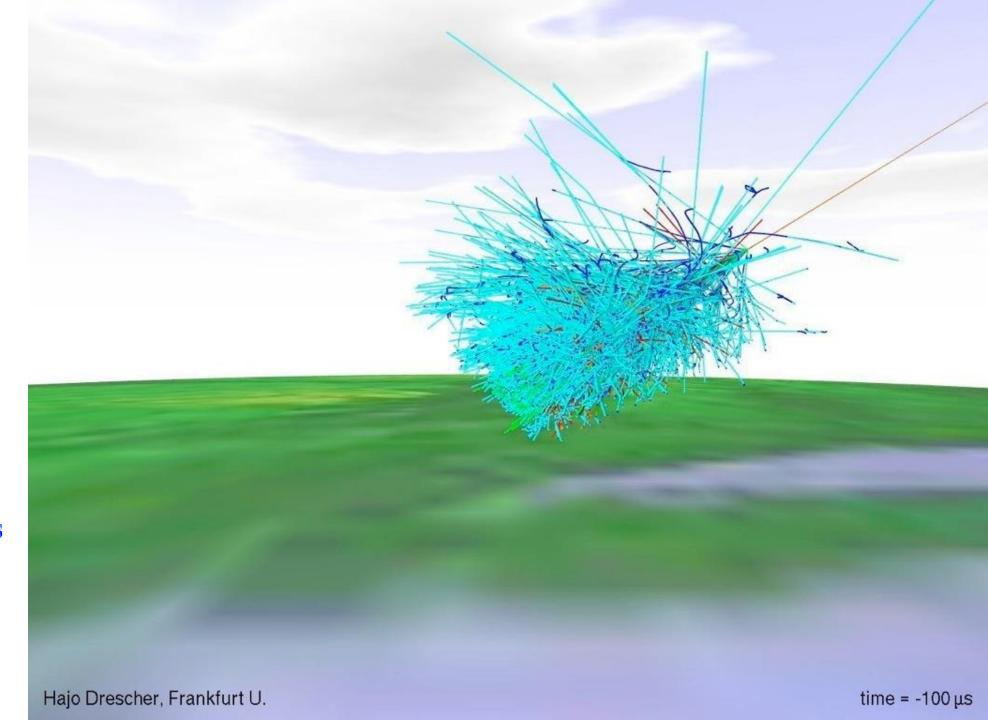
red:neutrons

orange: protons

gray: mesons

green:muons

Hajo Drescher, Frankfurt U. time = -200 μs

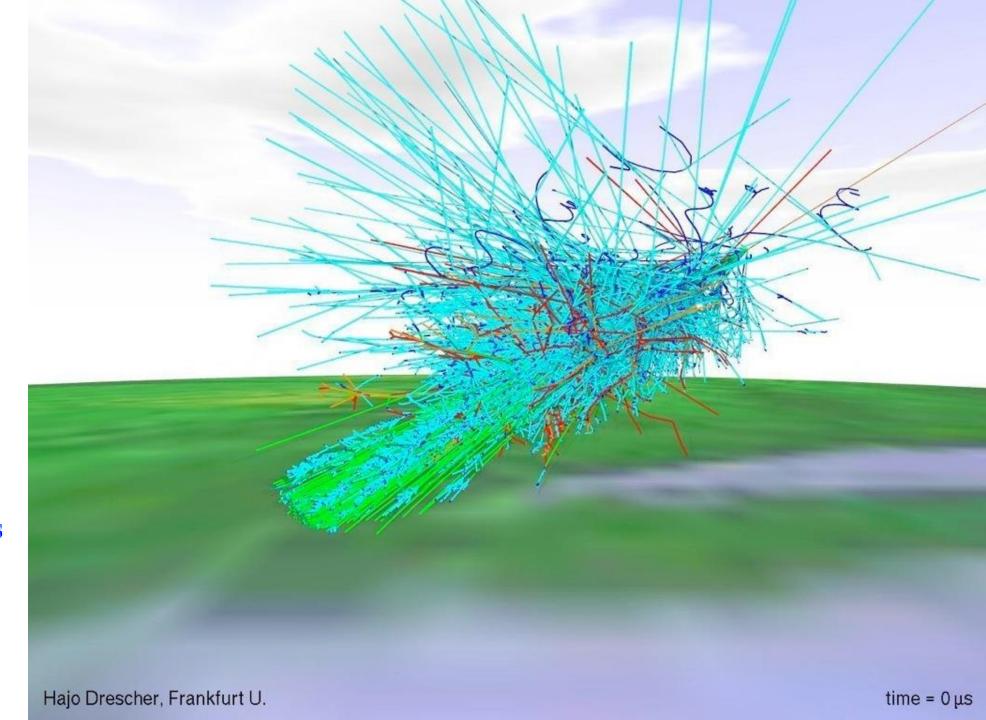


cyan:photons

red:neutrons

orange: protons

gray: mesons

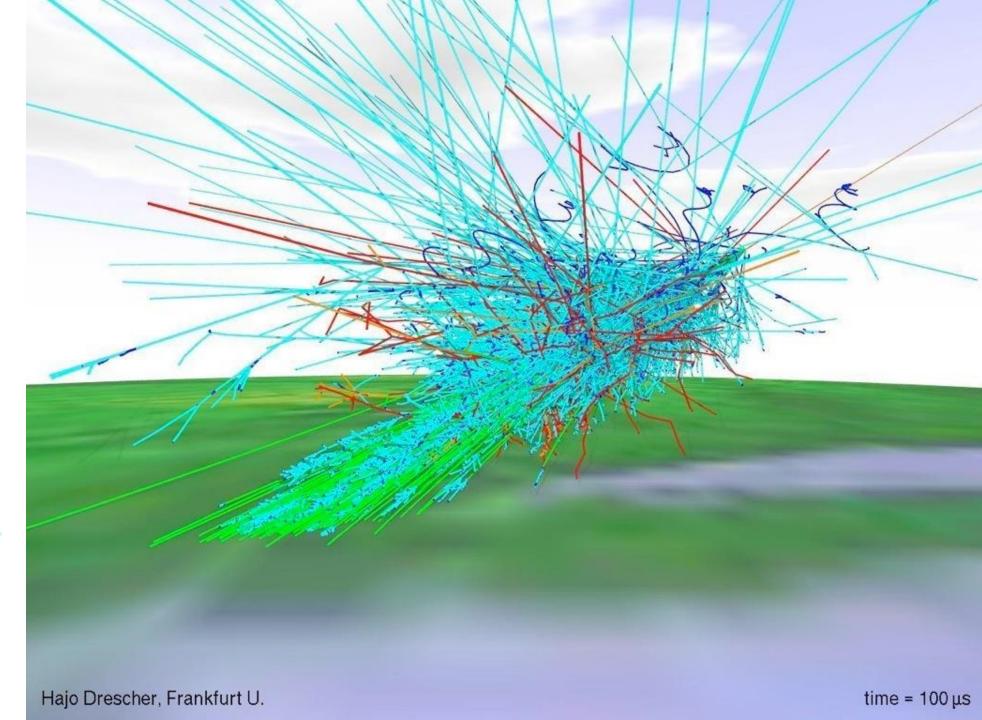


cyan:photons

red:neutrons

orange: protons

gray: mesons

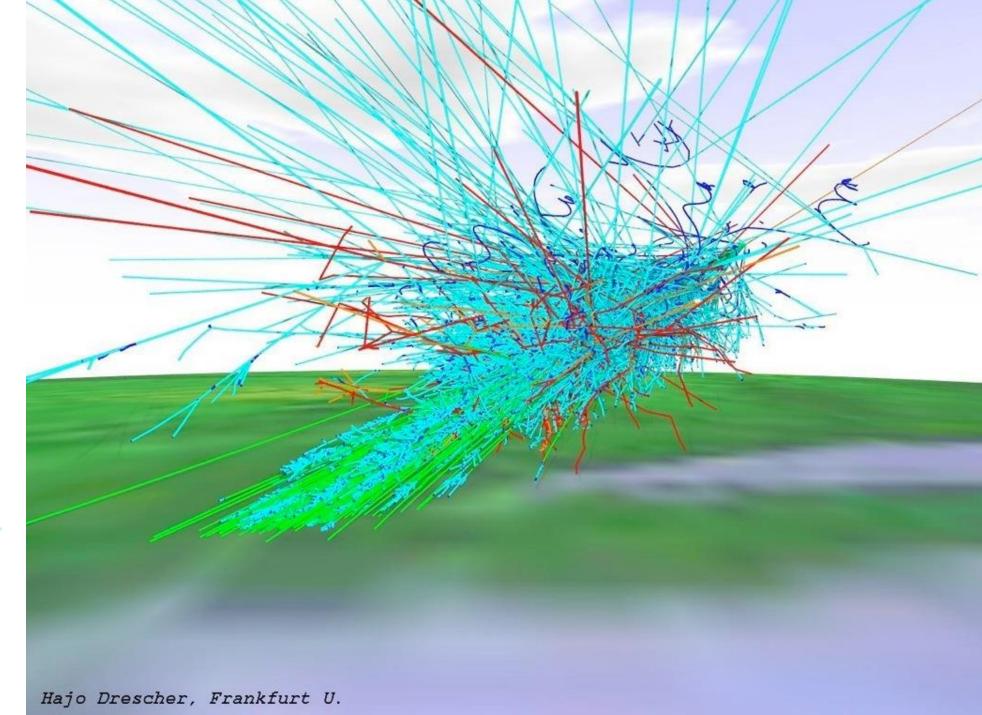


cyan:photons

red:neutrons

orange: protons

gray: mesons

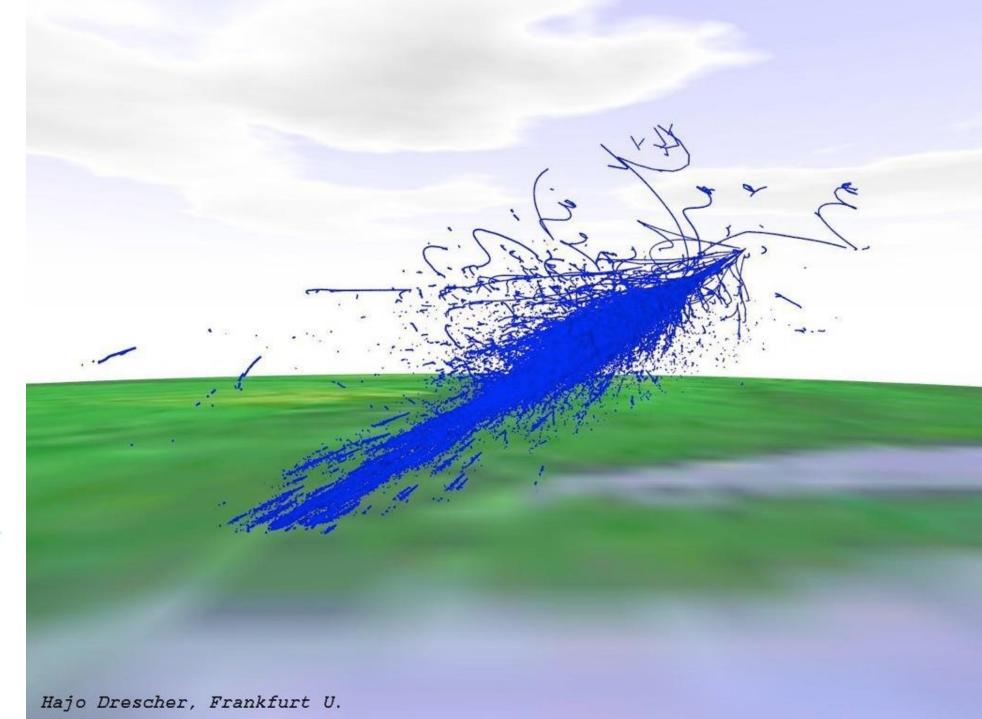


cyan:photons

red:neutrons

orange: protons

gray: mesons

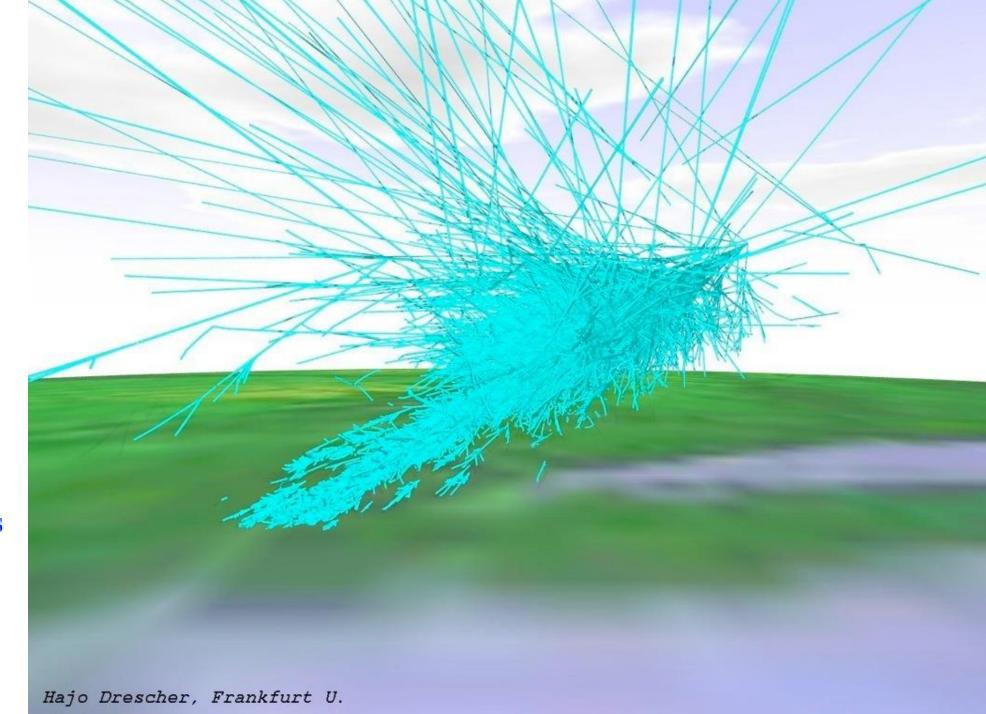


cyan:photons

red:neutrons

orange: protons

gray: mesons

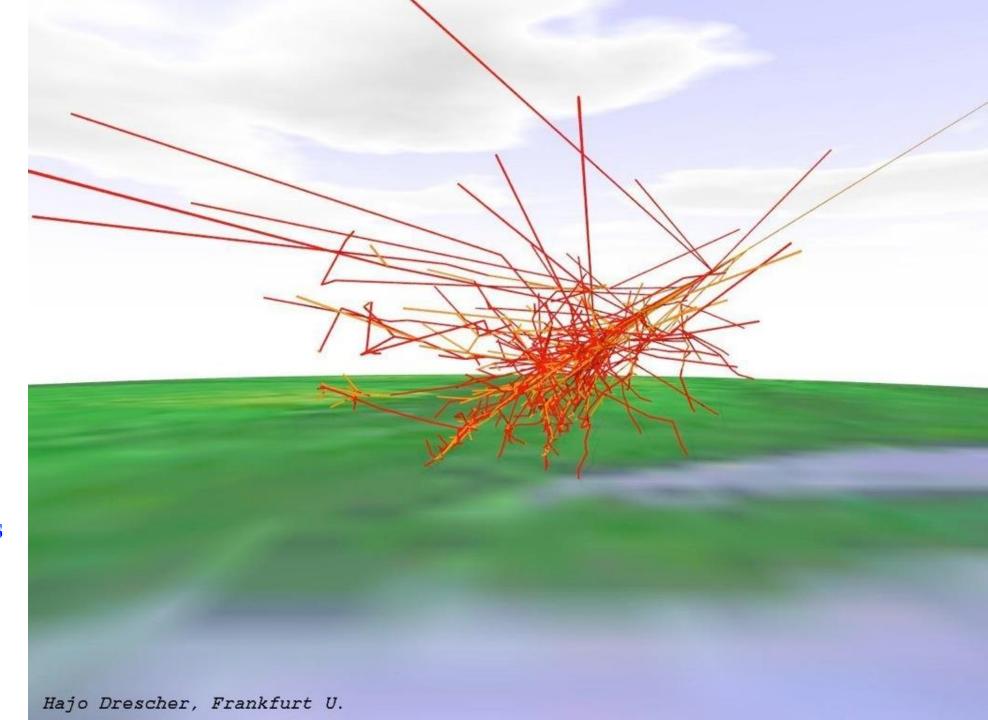


cyan:photons

red:neutrons

orange: protons

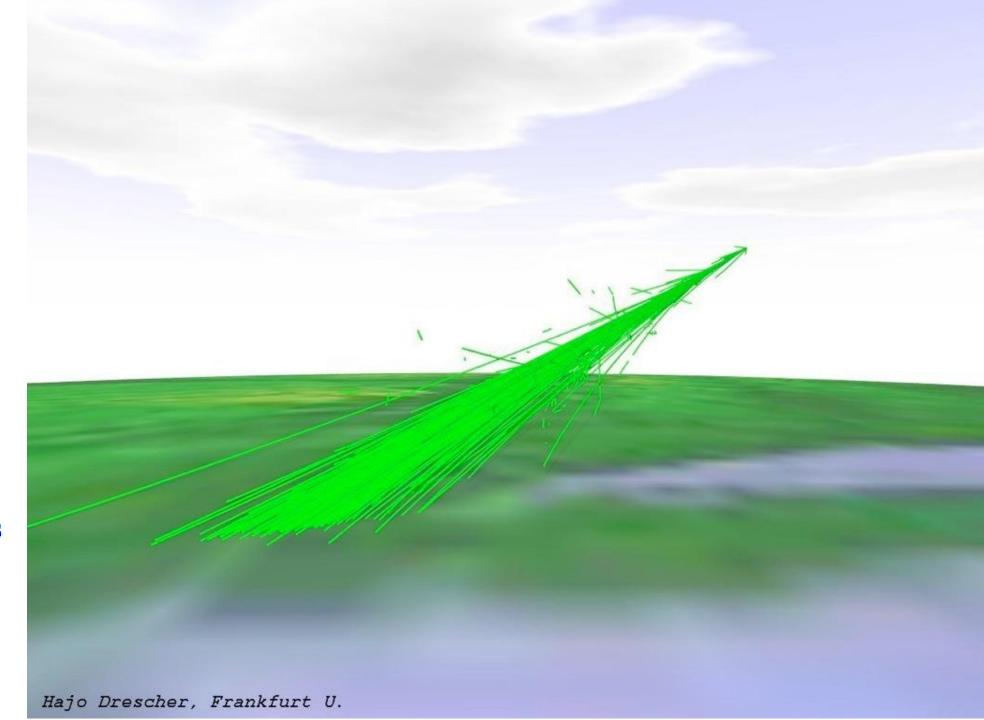
gray: mesons



cyan:photons red:neutrons

orange: protons

gray: mesons



cyan:photons

red:neutrons

orange: protons

gray: mesons