EFFECTIVE THEORIES FOR THE HIGGS: VIRTUES AND LIMITATIONS

Roberto Contino EPFL & CERN



Zurich Phenomenology Workshop, January 7-9, 2015

- At low energies, $E\!\ll\!m_*$, NP effects are well approximated by local operators

$$\mathcal{L} = \sum_{i} \bar{c}_i \, O_i(x)$$

 $\overbrace{\bigcirc} \rightarrow \overbrace{}$

Operators "generated" at the new physics scale m_* with coefficients

$$\bar{c}_i(m_*) \sim \left(\frac{1}{m_*}\right)^{d[O]-4}$$

- At low energies, $E\!\ll\!m_*$, NP effects are well approximated by local operators

$$\mathcal{L} = \sum_{i} \bar{c}_i \, O_i(x)$$



Operators "generated" at the new physics scale m_* with coefficients

$$\bar{c}_i(m_*) \sim \left(\frac{1}{m_*}\right)^{d[O]-4}$$

Assumptions:

- 1. There is a gap between the NP scale m_* and m_h
- 2. The new boson is part of an $SU(2)_L$ doublet

$$H = e^{i\pi/v} \begin{pmatrix} 0\\ v+h \end{pmatrix}$$

• Operators can be classified according to their dimension

$$\mathcal{L} = \mathcal{L}_{SM} + \sum_{i} \bar{c}_{i} O_{i} \equiv \mathcal{L}_{SM} + \Delta \mathcal{L}^{(6)} + \Delta \mathcal{L}^{(8)} + \dots$$

• Operators can be classified according to their dimension

$$\mathcal{L} = \mathcal{L}_{SM} + \sum_{i} \bar{c}_{i} O_{i} \equiv \mathcal{L}_{SM} + \Delta \mathcal{L}^{(6)} + \Delta \mathcal{L}^{(8)} + \dots$$

Leading effects from dim-6 operators

59 independent operators for 1 SM family

Buchmuller and Wyler NPB 268 (1986) 621 : Grzadkowski et al. JHEP 1010 (2010) 085

For a review see: RC, Ghezzi, Grojean, Muhlleitner, Spira JHEP 1307 (2013) 035

Giudice et al. JHEP 0706 (2007) 045

Giudice et al. JHEP 0706 (2007) 045

Assumption:

the UV physics is broadly characterized by 1 scale (m_{\ast}) and 1 coupling strength (g_{\ast})

Giudice et al. JHEP 0706 (2007) 045

Assumption:

the UV physics is broadly characterized by 1 scale (m_{st}) and 1 coupling strength (g_{st})

– each extra (covariant) derivative costs a factor $\frac{1}{m_*}$

- each extra power of H(x) costs a factor $\frac{g_*}{m_*} \equiv \frac{1}{f}$

Giudice et al. JHEP 0706 (2007) 045

Assumption:

the UV physics is broadly characterized by 1 scale (m_{st}) and 1 coupling strength (g_{st})

- each extra (covariant) derivative costs a factor $\frac{1}{m_{*}}$

- each extra power of H(x) costs a factor

s a factor
$$\ {g_*\over m_*}\equiv {1\over f}$$

For a strongly-interacting light Higgs (SILH): $g_* \gg 1 \quad \Longrightarrow \quad \frac{1}{f} \gg \frac{1}{m_*}$

Giudice et al. JHEP 0706 (2007) 045

Assumption:

the UV physics is broadly characterized by 1 scale (m_{st}) and 1 coupling strength (g_{st})

- each extra (covariant) derivative costs a factor $\frac{1}{m_{*}}$

each extra power of H(x) costs a factor

a factor
$$\ {g_*\over m_*}\equiv {1\over f}$$

For a strongly-interacting light Higgs (SILH): 1 1

$$g_* \gg 1 \quad \Longrightarrow \quad \frac{1}{f} \gg \frac{1}{m_*}$$

$$\bar{c}_W \sim \left(\frac{m_W^2}{m_*^2}\right)$$

 $O_W = \frac{ig}{2m_W^2} \left(H^{\dagger} \sigma^i D^{\mu} H \right) \left(D^{\nu} W_{\mu\nu} \right)^i$

 $(H^{\dagger}\sigma^{i}D^{\mu}H)$

Giudice et al. JHEP 0706 (2007) 045

Assumption:

the UV physics is broadly characterized by 1 scale (m_{st}) and 1 coupling strength (g_{st})



Secondary Assumptions:

1. The L	JV physics is minimally coupled
True holoç	for some of the popular models (e.g. weakly-coupled SUSY, graphic Higgs), but not necessarily so in more general contexts
see for	r example: Jenkins, Manohar, Trott JHEP 1309 (2013) 063
R S	operators generated at loop level suppressed by $(g_*^2/16\pi^2)$ ex: dipole operators

Secondary Assumptions:

1. The l	JV physics is minimally coupled
True holog	for some of the popular models (e.g. weakly-coupled SUSY, graphic Higgs), but not necessarily so in more general contexts
see fo	or example: Jenkins, Manohar, Trott JHEP 1309 (2013) 063
ß	operators generated at loop level suppressed by $(g_*^2/16\pi^2)$ ex: dipole operators

2. (light) SM fermions are weakly coupled to the UV dynamics

Equivalent to assuming "universality" of NP effects, easier to comply with LEP



Probing the strength of the EWSB dynamics

Although full reconstruction of their spectrum and couplings requires open producing the new states, EFT can give information on whether the UV dynamics is *strongly* or *weakly* coupled

Probing the strength of the EWSB dynamics

Although full reconstruction of their spectrum and couplings requires open producing the new states, EFT can give information on whether the UV dynamics is *strongly* or *weakly* coupled

• At LHC Run1 Higgs searches have focussed on single-Higgs on-shell production and decay

gives information on on-shell couplings at a fixed scale $Q = m_h$

$$\frac{\delta c}{c} \sim O\left(\frac{m_h^2}{m_*^2}\right) \text{ or } O\left(\frac{g_*^2}{g_{SM}^2} \times \frac{m_h^2}{m_*^2} = \frac{v^2}{f^2}\right) < 1$$

Probing the strength of the EWSB dynamics

Although full reconstruction of their spectrum and couplings requires open producing the new states, EFT can give information on whether the UV dynamics is *strongly* or *weakly* coupled

 At LHC Run1 Higgs searches have focussed on single-Higgs on-shell production and decay

→ gives information on on-shell couplings at a fixed scale $Q = m_h$

$$\frac{\delta c}{c} \sim O\left(\frac{m_h^2}{m_*^2}\right) \text{ or } O\left(\frac{g_*^2}{g_{SM}^2} \times \frac{m_h^2}{m_*^2} = \frac{v^2}{f^2}\right) < 1$$

• Next frontier for Run2: probe directly the strength of SSB dynamics at energies $E \gg m_h$ through 2 \rightarrow 2 scattering processes



Inspired by: R. Rattazzi, talk at "BSM physics opportunities at 100TeV", Cern 2014

In general:

$$A = g_{SM}^2 \times \left(1 + O\left(\frac{v^2}{f^2}\right)\right) + O\left(\frac{g_*^2 E^2}{m_*^2}\right) + \dots$$

Inspired by: R. Rattazzi, talk at "BSM physics opportunities at 100TeV", Cern 2014

In general:

$$A = g_{SM}^2 \times \left(1 + O\left(\frac{v^2}{f^2}\right)\right) + O\left(\frac{g_*^2 E^2}{m_*^2}\right) + \dots$$
$$= A_{SM}$$

Inspired by: R. Rattazzi, talk at "BSM physics opportunities at 100TeV", Cern 2014

In general:

Inspired by: R. Rattazzi, talk at "BSM physics opportunities at 100TeV", Cern 2014

In general:

dim-8 operators further suppressed by





Thus:
$$\frac{\delta A}{A_{SM}} \sim \frac{g(E)^2}{g_{SM}^2}$$
 can be > 1 if NP dynamics is strongly coupled ($g_* > g_{SM}$)





$$c_V = 1 - \frac{\bar{c}_H}{2}$$
$$c_{2V} = 1 - 2\bar{c}_H$$
$$c_3 = 1 - \frac{3}{2}\bar{c}_H + \bar{c}_6$$

$$A = c_V^2 \frac{m_h^2}{v^2} \left(1 + O(\delta_2, \delta_3) \right) - c_V^2 \frac{\hat{s}}{v^2} \delta_2 + \dots \qquad \qquad \delta_2 \equiv 1 - c_{2V} / c_V^2 \\ \delta_3 \equiv 1 - c_3 / c_V$$



$$c_V = 1 - \frac{\overline{c}_H}{2}$$
$$c_{2V} = 1 - 2\overline{c}_H$$
$$c_3 = 1 - \frac{3}{2}\overline{c}_H + \overline{c}_6$$

$$A = c_V^2 \underbrace{\frac{m_h^2}{v^2}}_{= A_{SM}} (1 + O(\delta_2, \delta_3)) - c_V^2 \frac{\hat{s}}{v^2} \delta_2 + \dots \qquad \delta_2 \equiv 1 - c_{2V}/c_V^2$$
$$\delta_3 \equiv 1 - c_3/c_V$$



$$c_V = 1 - \frac{\overline{c}_H}{2}$$
$$c_{2V} = 1 - 2\overline{c}_H$$
$$c_3 = 1 - \frac{3}{2}\overline{c}_H + \overline{c}_6$$

$$A = c_V^2 \underbrace{\frac{m_h^2}{v^2}}_{= A_{SM}} (1 + O(\delta_2, \delta_3)) - c_V^2 \frac{\hat{s}}{v^2} \delta_2 + \dots \qquad \delta_2 \equiv 1 - c_{2V}/c_V^2$$
$$\delta_3 \equiv 1 - c_3/c_V$$



$$c_V = 1 - \frac{\overline{c}_H}{2}$$
$$c_{2V} = 1 - 2\overline{c}_H$$
$$c_3 = 1 - \frac{3}{2}\overline{c}_H + \overline{c}_6$$

$$A = c_V^2 \underbrace{\frac{m_h^2}{v^2}}_{= A_{SM}} (1 + O(\delta_2, \delta_3)) - \underbrace{c_V^2 \frac{\hat{s}}{v^2} \delta_2}_{= g(E)^2} + \dots$$

= $g(E)^2 \sim \frac{E^2}{f^2}$

$$\delta_2 \equiv 1 - c_{2V}/c_V^2$$
$$\delta_3 \equiv 1 - c_3/c_V$$

 $A = c_{V}^{2} \begin{pmatrix} m_{h}^{2} \\ v^{2} \end{pmatrix} (1 + O(\delta_{2}, \delta_{3})) - \begin{pmatrix} c_{V}^{2} \hat{s} \\ v^{2} \sqrt{f^{2}} \end{pmatrix} = g(E)^{2} \sim \frac{E^{2}}{f^{2}}$ $c_{V} = 1 - \frac{\bar{c}_{H}}{2}$ $c_{2V} = 1 - 2\bar{c}_{H}$ $c_{3} = 1 - \frac{3}{2}\bar{c}_{H} + \bar{c}_{6}$ $O\left(\frac{E^{2}}{f^{2}} \times \frac{E^{2}}{m_{*}^{2}}\right)$ $\delta_{2} \equiv 1 - c_{2V}/c_{V}^{2}$ $\delta_{3} \equiv 1 - c_{3}/c_{V}$

 $A = c_{V}^{2} \begin{pmatrix} m_{h}^{2} \\ v^{2} \end{pmatrix} (1 + O(\delta_{2}, \delta_{3})) - \begin{pmatrix} c_{V}^{2} \hat{s} \\ v^{2} \sqrt{f^{2}} \end{pmatrix} = g(E)^{2} \sim \frac{E^{2}}{f^{2}}$ $c_{V} = 1 - \frac{\bar{c}_{H}}{2}$ $c_{2V} = 1 - 2\bar{c}_{H}$ $c_{3} = 1 - \frac{3}{2}\bar{c}_{H} + \bar{c}_{6}$ $O\left(\frac{E^{2}}{f^{2}} \times \frac{E^{2}}{m_{*}^{2}}\right)$ $\delta_{2} \equiv 1 - c_{2V}/c_{V}^{2}$ $\delta_{3} \equiv 1 - c_{3}/c_{V}$

If best sensitivity $(\delta_2)_{min}$ comes from events with invariant mass $m(hh) \sim \overline{E}$:

$$\sqrt{(\delta_2)_{min}} \,\frac{\bar{E}}{v} = g_{min} < g_* \lesssim 4\pi$$

10

pp colliders

p	p ightarrow l	hh_{J}	$jj \rightarrow 4$	bjj
	L	• /		•/ •/

	$(\delta_3)_{min}$	$(\delta_2)_{min}$	$ar{E}$	g_{min}
$14 \mathrm{TeV}, L = 300 \mathrm{fb}^{-1}$	8-9	0.35	$1.5\mathrm{TeV}$	3.6
$14 \mathrm{TeV}, L = 3 \mathrm{ab}^{-1}$	5 - 8	0.2	$1.5\mathrm{TeV}$	2.7
$100 \mathrm{TeV}, \ L = 3 \mathrm{ab}^{-1}$	4 - 6	0.04	$3.5\mathrm{TeV}$	2.8

work in progress with J. Rojo



$$\delta_2 \equiv 1 - c_{2V}/c_V^2$$
$$\delta_3 \equiv 1 - c_3/c_V$$

pp colliders $pp \rightarrow$

 $pp \rightarrow hh\,jj \rightarrow 4b\,jj$

	$(\delta_3)_{min}$	$(\delta_2)_{min}$	$ar{E}$	g_{min}
14 TeV, $L = 300 \text{fb}^{-1}$	8-9	0.35	$1.5{ m TeV}$	3.6
$14 \mathrm{TeV}, L = 3 \mathrm{ab}^{-1}$	5 - 8	0.2	$1.5\mathrm{TeV}$	2.7
$100 \mathrm{TeV}, \ L = 3 \mathrm{ab}^{-1}$	4 - 6	0.04	$3.5\mathrm{TeV}$	2.8

work in progress with J. Rojo

CLIC
$$e^+e^- \to hh\,\nu\bar\nu \to 4b\,\nu\bar\nu$$

	$(\delta_3)_{min}$	$(\delta_2)_{min}$	$ar{E}$	g_{min}
$3 \mathrm{TeV}, \ L = 1 \mathrm{ab}^{-1}$	0.3	0.05	$1.8\mathrm{TeV}$	1.6

RC, Grojean, Pappadopulo, Rattazzi, Thamm JHEP 1402 (2014) 006



$$\delta_2 \equiv 1 - c_{2V}/c_V^2$$
$$\delta_3 \equiv 1 - c_3/c_V$$

11

pp colliders $pp \rightarrow$

 $pp \rightarrow hh\,jj \rightarrow 4b\,jj$

	$(\delta_3)_{min}$	$(\delta_2)_{min}$	$ar{E}$	g_{min}
14 TeV, $L = 300 \text{fb}^{-1}$	8-9	0.35	$1.5{ m TeV}$	3.6
$14 \mathrm{TeV}, L = 3 \mathrm{ab}^{-1}$	5 - 8	0.2	$1.5{ m TeV}$	2.7
$100 \mathrm{TeV}, \ L = 3 \mathrm{ab}^{-1}$	4 - 6	0.04	$3.5\mathrm{TeV}$	2.8

work in progress with J. Rojo

CLIC
$$e^+e^- \to hh\,\nu\bar\nu \to 4b\,\nu\bar\nu$$

	$(\delta_3)_{min}$	$(\delta_2)_{min}$	$ar{E}$	g_{min}
$3 \mathrm{TeV}, \ L = 1 \mathrm{ab}^{-1}$	0.3	0.05	$1.8\mathrm{TeV}$	1.6

RC, Grojean, Pappadopulo, Rattazzi, Thamm JHEP 1402 (2014) 006

EFT better justified at high-precision machines (such as e⁺e⁻ colliders)



$$\delta_2 \equiv 1 - c_{2V}/c_V^2$$
$$\delta_3 \equiv 1 - c_3/c_V$$

$$A = g^{2} + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \,\bar{c}_{i=HB,(W-B)}\right) + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \,\bar{c}_{H\psi}\right)$$



$$O_{W} = \frac{ig}{2m_{W}^{2}} \left(H^{\dagger} \sigma^{i} \overleftarrow{D^{\mu}} H \right) (D^{\nu} W_{\mu\nu})^{i}$$
$$O_{B} = \frac{ig'}{2m_{W}^{2}} \left(H^{\dagger} \overleftarrow{D^{\mu}} H \right) \partial^{\nu} B_{\mu\nu}$$
$$O_{HB} = \frac{ig'}{m_{W}^{2}} (D^{\mu} H)^{\dagger} (D^{\nu} H) B_{\mu\nu}$$
$$O_{H\psi} = \frac{i}{v^{2}} \left(\bar{\psi} \gamma^{\mu} \psi \right) \left(H^{\dagger} \overleftarrow{D}_{\mu} H \right)$$

$$\begin{split} A &= g^{2} + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{i=HB,(W-B)}\right) + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{H\psi}\right) \\ &= A_{SM} \\ &= O\left(g^{2} \frac{E^{2}}{m_{*}^{2}}\right) \\ &= O\left(\lambda^{2} \frac{E^{2}}{m_{*}^{2}}\right) \\ \end{split}$$

$$O_{W} = \frac{ig}{2m_{W}^{2}} \left(H^{\dagger} \sigma^{i} \overleftrightarrow{D^{\mu}} H \right) (D^{\nu} W_{\mu\nu})^{i}$$
$$O_{B} = \frac{ig'}{2m_{W}^{2}} \left(H^{\dagger} \overleftrightarrow{D^{\mu}} H \right) \partial^{\nu} B_{\mu\nu}$$
$$O_{HB} = \frac{ig'}{m_{W}^{2}} \left(D^{\mu} H \right)^{\dagger} (D^{\nu} H) B_{\mu\nu}$$
$$O_{H\psi} = \frac{i}{v^{2}} \left(\bar{\psi} \gamma^{\mu} \psi \right) \left(H^{\dagger} \overleftrightarrow{D}_{\mu} H \right)$$

$$A = g^{2} + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{i=HB,(W-B)}\right) + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{H\psi}\right)$$

$$= A_{SM}$$

$$= O\left(g^{2} \frac{E^{2}}{m_{*}^{2}}\right) = O\left(\lambda^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$
Riva et al.
arXiv:1406.7320

$$Must have$$

$$\delta A/A_{SM} < 1$$
for EFT to be valid

$$O_W = \frac{ig}{2m_W^2} \left(H^{\dagger} \sigma^i \overleftrightarrow{D^{\mu}} H \right) (D^{\nu} W_{\mu\nu})^i$$
$$O_B = \frac{ig'}{2m_W^2} \left(H^{\dagger} \overleftrightarrow{D^{\mu}} H \right) \partial^{\nu} B_{\mu\nu}$$
$$O_{HB} = \frac{ig'}{m_W^2} (D^{\mu} H)^{\dagger} (D^{\nu} H) B_{\mu\nu}$$
$$O_{H\psi} = \frac{i}{v^2} \left(\bar{\psi} \gamma^{\mu} \psi \right) \left(H^{\dagger} \overleftrightarrow{D}_{\mu} H \right)$$

$$A = g^{2} + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{i=HB,(W-B)}\right) + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{H\psi}\right)$$

$$= A_{SM}$$

$$= O\left(g^{2} \frac{E^{2}}{m_{*}^{2}}\right) \qquad = O\left(\lambda^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$= O\left(\lambda^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$= O\left(\lambda^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$\frac{\uparrow}{\int}$$

$$\frac{M}{\int}$$
Riva et al.
arXiv:1406.7320
$$Must have$$

$$\delta A/A_{SM} < 1$$
if $\lambda \gg g$

$$O_W = \frac{ig}{2m_W^2} \left(H^{\dagger} \sigma^i \overleftrightarrow{D^{\mu}} H \right) (D^{\nu} W_{\mu\nu})^i$$
$$O_B = \frac{ig'}{2m_W^2} \left(H^{\dagger} \overleftrightarrow{D^{\mu}} H \right) \partial^{\nu} B_{\mu\nu}$$
$$O_{HB} = \frac{ig'}{m_W^2} (D^{\mu} H)^{\dagger} (D^{\nu} H) B_{\mu\nu}$$
$$O_{H\psi} = \frac{i}{v^2} \left(\bar{\psi} \gamma^{\mu} \psi \right) \left(H^{\dagger} \overleftrightarrow{D}_{\mu} H \right)$$

$$A = g^{2} + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{i=HB,(W-B)}\right) + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{H\psi}\right)$$

$$= A_{SM}$$

$$= O\left(g^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$= O\left(\lambda^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$= O\left(\lambda^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$\uparrow$$

$$Must have$$

$$\delta A/A_{SM} < 1$$
if $\lambda \gg g$
if $\lambda \gg g$

 $O_{W} = \frac{ig}{2m_{W}^{2}} \left(H^{\dagger} \sigma^{i} \overleftrightarrow{D^{\mu}} H \right) (D^{\nu} W_{\mu\nu})^{i}$ $O_{B} = \frac{ig'}{2m_{W}^{2}} \left(H^{\dagger} \overleftrightarrow{D^{\mu}} H \right) \partial^{\nu} B_{\mu\nu}$ $O_{HB} = \frac{ig'}{m_{W}^{2}} (D^{\mu} H)^{\dagger} (D^{\nu} H) B_{\mu\nu}$ $O_{H\psi} = \frac{i}{v^{2}} \left(\bar{\psi} \gamma^{\mu} \psi \right) \left(H^{\dagger} \overleftrightarrow{D}_{\mu} H \right)$

$$O_{HB}, (O_W - O_B)$$
 constrained by TGC

12

$$A = g^{2} + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{i=HB,(W-B)}\right) + O\left(g^{2} \frac{E^{2}}{m_{W}^{2}} \bar{c}_{H\psi}\right)$$

$$= A_{SM}$$

$$= O\left(g^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$= O\left(\chi^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$= O\left(\chi^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$= O\left(\chi^{2} \frac{E^{2}}{m_{*}^{2}}\right)$$

$$\int \left(\int \frac{\delta A}{A_{SM}} < 1\right)$$

$$\int \frac{\delta A}{A_{SM}} < 1$$

$$\int \frac{\delta A}{A_{SM}} > 1$$

$$O_{W} = \frac{ig}{2m_{W}^{2}} \left(H^{\dagger} \sigma^{i} \overleftarrow{D^{\mu}} H \right) (D^{\nu} W_{\mu\nu})^{i}$$

$$O_{B} = \frac{ig'}{2m_{W}^{2}} \left(H^{\dagger} \overleftarrow{D^{\mu}} H \right) \partial^{\nu} B_{\mu\nu}$$

$$O_{HB} = \frac{ig'}{m_{W}^{2}} (D^{\mu} H)^{\dagger} (D^{\nu} H) B_{\mu\nu}$$

$$O_{H\psi} = \frac{i}{v^{2}} \left(\bar{\psi} \gamma^{\mu} \psi \right) \left(H^{\dagger} \overleftarrow{D}_{\mu} H \right)$$

$$\longrightarrow \text{ constrained by Z-pole data at LEP1}$$

Experimental searches not yet sensitive to SM Higgs signal

ATLAS-CONF-2013-079 CMS PAS-HIG-13-012 D0, PRL 109 (2012) 121802

EFT <u>not</u> valid when setting limits on $\bar{c}_{HB}, (\bar{c}_W - \bar{c}_B)$

Riva et al. arXiv:1406.7320







work in progress with A. Azatov, G. Panico, M. Son



$$A(gg \to hh) \sim \left(\frac{\alpha_s}{4\pi}\right) \times \left[y_t^2 \left(1 + O\left(\frac{v^2}{f^2}\right)\right) + g_6^2(E) + g_8^2(E) + \dots\right]$$

work in progress with A. Azatov, G. Panico, M. Son



work in progress with A. Azatov, G. Panico, M. Son



suppressed by weak spurion λ for a pNGB Higgs

work in progress with A. Azatov, G. Panico, M. Son



14

work in progress with A. Azatov, G. Panico, M. Son

$$g_6^2(E) \sim \bar{c}_g \,\frac{4\pi}{\alpha_2} \,\frac{E^2}{v^2} \sim \frac{\lambda^2 E^2}{m_*^2} \qquad \qquad g_8^2(E) \sim \bar{c}_{gD0,2} \,\frac{4\pi}{\alpha_2} \,\frac{E^4}{v^2 m_W^2} \sim \frac{g_*^2 E^4}{m_*^4}$$

work in progress with A. Azatov, G. Panico, M. Son

$$g_6^2(E) \sim \bar{c}_g \,\frac{4\pi}{\alpha_2} \,\frac{E^2}{v^2} \sim \frac{\lambda^2 E^2}{m_*^2} \qquad \qquad g_8^2(E) \sim \bar{c}_{gD0,2} \,\frac{4\pi}{\alpha_2} \,\frac{E^4}{v^2 m_W^2} \sim \frac{g_*^2 E^4}{m_*^4}$$

• EFT valid for $E\!\ll\!m_*$:

$$g_6(E) < \lambda \lesssim g_* \qquad \qquad g_8(E) < g_*$$

work in progress with A. Azatov, G. Panico, M. Son

$$g_6^2(E) \sim \bar{c}_g \frac{4\pi}{\alpha_2} \frac{E^2}{v^2} \sim \frac{\lambda^2 E^2}{m_*^2}$$

$$g_8^2(E) \sim \bar{c}_{gD0,2} \, \frac{4\pi}{\alpha_2} \, \frac{E^4}{v^2 m_W^2} \sim \frac{g_*^2 E^4}{m_*^4}$$

- EFT valid for $E \ll m_*$: $g_6(E) < \lambda \lesssim g_*$ $g_8(E) < g_*$
- dim-8 operators are more important than dim-6 ones for $m_*(\lambda/g_*)\!<\!E\!<\!m_*$

Condition for neglecting dim-8 operators:

$$g_6(E) < \frac{\lambda^2}{g_*}$$

work in progress with A. Azatov, G. Panico, M. Son

$$g_6^2(E) \sim \bar{c}_g \, \frac{4\pi}{\alpha_2} \, \frac{E^2}{v^2} \sim \frac{\lambda^2 E^2}{m_*^2} \qquad g_8^2(E) \sim g_8^2(E)$$

$$\eta_8^2(E) \sim \bar{c}_{gD0,2} \, \frac{4\pi}{\alpha_2} \, \frac{E^4}{v^2 m_W^2} \sim \frac{g_*^2 E^4}{m_*^4}$$

- EFT valid for $E \ll m_*$: $g_6(E) < \lambda \lesssim g_*$ $g_8(E) < g_*$
- dim-8 operators are more important than dim-6 ones for $m_*(\lambda/g_*)\!<\!E\!<\!m_*$

Condition for neglecting dim-8 operators:

$$g_6(E) < \frac{\lambda^2}{g_*}$$

• Let δ be the precision obtained on $\bar{c}_g(4\pi/\alpha_2)$ using events with invariant mass $m(hh)\sim \bar{E}$

$$g_6(E)_{min} \equiv g_{min} \sim \sqrt{\delta} \, \frac{\bar{E}}{v}$$

$$\lambda > g_{min} \qquad \frac{\lambda^2}{g_*} > g_{min}$$

work in progress with A. Azatov, G. Panico, M. Son



- Let δ be the precision obtained on $\bar{c}_g(4\pi/\alpha_2)$ using events with invariant mass up to \bar{E}

$$g_6(E)_{min} \equiv g_{min} \sim \sqrt{\delta} \, \frac{\bar{E}}{v}$$

$$\lambda > g_{min} \qquad \frac{\lambda^2}{g_*} > g_{min}$$

work in progress with A. Azatov, G. Panico, M. Son



• Let δ be the precision obtained on $\bar{c}_g(4\pi/\alpha_2)$ using events with invariant mass up to \bar{E}

$$g_6(E)_{min} \equiv g_{min} \sim \sqrt{\delta} \, \frac{\bar{E}}{v}$$

$$\lambda > g_{min} \qquad \frac{\lambda^2}{g_*} > g_{min}$$

work in progress with A. Azatov, G. Panico, M. Son



• Let δ be the precision obtained on $\bar{c}_g(4\pi/\alpha_2)$ using events with invariant mass up to \bar{E}

$$g_6(E)_{min} \equiv g_{min} \sim \sqrt{\delta} \, \frac{\bar{E}}{v}$$

$$\lambda > g_{min} \qquad \qquad \frac{\lambda^2}{g_*} > g_{min}$$

work in progress with A. Azatov, G. Panico, M. Son



Notice:

Fully composite t_R ($\lambda = y_t$) and partially composite t_R and t_L ($\lambda = \sqrt{g_* y_t}$) can be probed only if $g_{min} < y_t$ (requires sensitivity to the SM cross section)

$$\lambda > g_{min} \qquad \qquad \frac{\lambda^2}{g_*} > g_{min}$$

work in progress with A. Azatov, G. Panico, M. Son



Notice:

Fully composite t_R ($\lambda = y_t$) and partially composite t_R and t_L ($\lambda = \sqrt{g_* y_t}$) can be probed only if $g_{min} < y_t$ (requires sensitivity to the SM cross section) Results for channel $pp \rightarrow hh \rightarrow b \overline{b} \gamma \gamma$

$14 \mathrm{TeV},300 \mathrm{fb}^{-1}$	$14\mathrm{TeV},3\mathrm{ab}^{-1}$	$100{ m TeV},3{ m ab}^{-1}$
$\delta \simeq 0.15$ $\bar{E} \simeq 1.0 \mathrm{TeV}$	$\delta \simeq 0.1$ $\bar{E} \simeq 1.0 \mathrm{TeV}$	$\delta \simeq 0.05$ $\bar{E} \simeq 1.0 \mathrm{TeV}$
$g_{min} = 1.6$	$g_{min} = 1.3$	$g_{min}=0.9$

Higgs Effective Lagrangian is the tool for future precision physics

Higgs Effective Lagrangian is the tool for future precision physics

 Power counting needed to estimate coefficients and extract information on the underlying UV dynamics. SILH power counting useful to test weak vs strong EWSB dynamics

Higgs Effective Lagrangian is the tool for future precision physics

Power counting needed to estimate coefficients and extract information on the underlying UV dynamics. SILH power counting useful to test weak vs strong EWSB dynamics

Next frontier for LHC Run2: $2 \rightarrow 2$ scattering processes to probe strength of EWSB dynamics at $E \gg m_h$

Higgs Effective Lagrangian is the tool for future precision physics

Power counting needed to estimate coefficients and extract information on the underlying UV dynamics. SILH power counting useful to test weak vs strong EWSB dynamics

 Next frontier for LHC Run2: 2→2 scattering processes to probe strength of EWSB dynamics at E≫m_h

Assessing the validity of EFT requires power counting.
 Dim-8 operators can be relevant if dim-6 ones are suppressed



On-shell single-Higgs cannot distinguish the top loop from a point-like interaction:

$$A(gg \to h) = A_{SM} \left(1 - \bar{c}_u + 12 \left(\frac{4\pi}{\alpha_2} \right) \bar{c}_g \right) + \dots$$
$$O(v^2/f^2) \qquad O\left(\frac{\lambda^2}{y_t^2} \frac{m_t^2}{m_*^2} \right)$$



Banfi et al. arXiv:1308.4771

Azatov, Paul arXiv:1309.5273 Grojean et al. arXiv:1312.3317 Schlaffer et al. arXiv:1405.4295

On-shell single-Higgs cannot distinguish the top loop from a point-like interaction:

$$A(gg \to h) = A_{SM} \left(1 - \bar{c}_u + 12 \left(\frac{4\pi}{\alpha_2} \right) \bar{c}_g \right) + \dots$$
$$O(v^2/f^2) \qquad O\left(\frac{\lambda^2}{y_t^2} \frac{m_t^2}{m_*^2} \right)$$

An extra hard jet can probe the top loop and break the degeneracy:



$$A(gg \to gh) = A_{SM} \left(1 - \bar{c}_u + 12 \left(\frac{4\pi}{\alpha_2} \right) \bar{c}_g \times f\left(\frac{p_T}{m_t} \right) \right) + \dots$$
$$O\left(\frac{\lambda^2}{y_t^2} \frac{E^2}{m_*^2} \right) \quad \text{for } p_T \gg m_t$$





For the effective theory to be valid one needs:

 $\overline{0}$

$$12\left(\frac{4\pi}{\alpha_2}\right)\bar{c}_g \times \frac{p_T^2}{m_t^2} \approx \frac{\lambda^2}{y_t^2} \frac{E^2}{m_*^2} < \frac{\lambda^2}{y_t^2}$$

An extra hard jet can probe the top loop and break the degeneracy:

-

Banfi et al. arXiv:1308.4771 Azatov, Paul arXiv:1309.5273 Grojean et al. arXiv:1312.3317 Schlaffer et al. arXiv:1405.4295

$$A(gg \to gh) = A_{SM} \left(1 - \bar{c}_u + 12 \left(\frac{4\pi}{\alpha_2} \right) \bar{c}_g \times f\left(\frac{p_T}{m_t} \right) \right) + \dots$$

$$O\left(\frac{\lambda^2}{y_t^2} \frac{E^2}{m_*^2} \right) \quad \text{for } p_T \gg m_t$$

 \sim



For the effective theory to be valid one needs:

$$12\left(\frac{4\pi}{\alpha_2}\right)\bar{c}_g \times \frac{p_T^2}{m_t^2} \approx \frac{\lambda^2}{y_t^2}\frac{E^2}{m_*^2} < \frac{\lambda^2}{y_t^2}$$

For a cut $p_T > 650 \,\mathrm{GeV}$ (as done in arXiv:1312.3317)

0000

$$3 < \frac{\lambda^2}{y_t^2}$$

 $O\left(\frac{\pi}{u_t^2}\frac{z}{m_t^2}\right)$

$$1.7 \, y_t \lesssim \lambda < g_* < 4\pi$$

An extra hard jet can probe the top loop and break the degeneracy:

0000

Azatov, Paul arXiv:1309.5273 Grojean et al. arXiv:1312.3317 Schlaffer et al. arXiv:1405.4295

Banfi et al. arXiv:1308.4771

for $p_T \gg m_t$

$$A(gg \to gh) = A_{SM} \left(1 - \bar{c}_u + 12 \left(\frac{4\pi}{\alpha_2} \right) \bar{c}_g \times f\left(\frac{p_T}{m_t} \right) \right) + \dots$$

