# Antineutrino Spectra and Nuclear Databases

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#### **ENDF/B-VII.1 Decay Data Sublibrary**

- Based on the latest ENSDF file.
- EEM and ELP data for most of Greenwood nuclides and all of the Valencia results.
- CGM calculations by Kawano and Moller, with the latest Q-values, for nuclides with incomplete data in ENSDF (not available for antineutrinos due to a format issue).

#### **ENDF/B-VII.1.1 (in progress)**

Contains I $\beta$  from Greenwood and Valencia TAGS. Contains I $\beta$  from a fit to Rudstam/Tengblad data for a few nuclides.

I $\beta$  (gs) for 92Rb set to 95.1% (from Lhesornneau et al).



# Recent work on Antineutrino Spectra Calculations

Antineutrino spectra are calculated as (Vogel 1981)

Spectrum =  $\Sigma$  CFY<sub>i</sub> Spectrum<sub>i</sub>

Cumulative Fission Yields from JEFF-3.1

No adjustable parameters

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# Recent work on Antineutrino Spectra Calculations

Nuclides with CGM calculations contribute about 4% for 235U, 14% for 238U, 7% for 239Pu and 12% for 241Pu, to the total <σl>.





# Recent work on Antineutrino Spectra Calculations

Comparison with Rudstam data for the two most important nuclides for energies higher than 5 MeV, assuming allowed shape.

New measurements are needed to test deviation from allowed shapes for first forbidden transitions.



# Conflicting Results on <sup>92</sup>Rb decay





# **Antineutrino multiplicities**



#### Antineutrino Energy (MeV)

The integral of antineutrino multiplicity has a similar behavior as the delayed neutron multiplicity.







# Systematics of <σl> with (3Z-A)

LFF: Light fission fragments (Z<47)

HFF: Heavy Fission Fragments (Z >= 47)

EE: Even-Z, Even-N EO: Even-Z, Odd-N OE: Odd-Z, Even-N OO: Odd-Z, Odd-N



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## **Q**β- using liquid drop model parameters

Qβ-(Z,A)=Mn-Mp-a<sub>c</sub>(2Z+1)/A<sup>2/3</sup>+a<sub>sym</sub>(4N-2Z-4)/A+ $\Delta$ (Z+1,A)- $\Delta$ (Z,A) If Z,A= EE -> Z+1,A=OO ->  $\Delta$ (Z+1,A)- $\Delta$ (Z,A)=-2δ/A<sup>1/2</sup> If Z,A= EO -> Z+1,A=OE ->  $\Delta$ (Z+1,A)- $\Delta$ (Z,A)=0 If Z,A= OE -> Z+1,A=EO ->  $\Delta$ (Z+1,A)- $\Delta$ (Z,A)=0 If Z,A= OO -> Z+1,A=EE ->  $\Delta$ (Z+1,A)- $\Delta$ (Z,A)=+2δ/A<sup>1/2</sup> δ≈12 MeV, 2δ/A<sup>1/2</sup> ≈ 2.4 MeV for A=100

		98Mo STABLE 24.39%	99Mo 65.976 H	100Mo 7.3E+18 Y 9.9295	101Mo 14.61 M	102Mo 11.3 M	103 <b>M</b> o 67.5 S	104 <b>M</b> o 60 S	105 <b>M</b> o 35.6 S	106 <b>M</b> o 8.73 S
	z	24.3770	β-: 100.00%	2β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%
		-1684	1357.8	-169.6	2825	1000	3635	2151	4.95E+3	3635
		97Nb 72.1 M	98Nb 2.86 S	99 <b>Nb</b> 15.0 S	100 <b>N</b> Б 1.5 S	101 <b>N</b> Б 7.1 S	102 <b>№</b> 4.3 S	103 <b>N</b> Б 1.5 S	104Nb 4.9 S	105Nb 2.95 S
	41	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%
		1935.5	4584	3637	6386	4628	7260	5942	р-л: 0.06%5 8531	р-л: 1.70жо 7431
		96Zr 2.35E+19 Y	972r 16.749 H	98Zr 30.7 S	99Zr 2.1 S	100Zr 7.1 S	101Zr 2.3 S	102Zr 2.9 S	103Zr 1.32 S	1042r 0.87 S
		2.80% 2β-	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%
		162	2659.7	2238	4707	3421	5717	4717	р-дз 1.00мв 7204	6095
	39	95Y 10.3 M	96Y 5.34 S	97Y 3.75 S	98Y 0.548 S	99Y 1.484 S	100Y 735 MS	101Y 0.45 S	102Y 0.36 S	103Y 0.23 S
		β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%
		4450	7103	б821	р-л: 0.33ж 8992	6969	р-л: 0.92жо 9049	р-л: 1.94% 8104	р-л: 4.90% 10420	9364
	38	94Sr 75.3 S	95Sr 23.90 S	96Sr 1.07 S	975r 429 MS	98Sr 0.653 S	99Sr 0.269 S	100Sr 202 MS	1015r 118 MS	102Sr 69 MS
		β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%	β-: 100.00%
		3507	6089	5412	7545	5875	8144	7503	9.51E+3	8.82E+3
		56	57	58	59	60	61	62	63	N
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# (3Z-A) Average antineutrino energy dependence



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Delayed nu-bar= $\Sigma$  CFY<sub>i</sub> Pn<sub>i</sub> using JEFF-3.1 yields and ENDF/B-VII.1 decay data



## **Relation to Decay Heat**

Power released as a function of time following a single fission event.







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# Conclusions

Due to incomplete decay schemes and not very precise fission yields, the summation method can't be used to calculate antineutrino spectra with high-precision.

However, one can learn a lot by exploiting from it, and relate it with the decay heat and delayed nu-bar phenomena.

Odd-Z, Odd-N nuclides are the main contributors to the antineutrino spectra, due to a) large  $Q\beta$  and existence of low-spin ground state/isomer.

Similarly, the light fission fragment contributes more antineutrinos than the heavy group.

Fission yields are needed in summation and decay heat calculations. However, there are discrepancies between the different libraries.

In particular, for 92Rb a measurement from Lowell puts the CFY at around 7%.



# **Back up slides**

# 239Pu Thermal

20 most important nuclides at T=500 seconds





Time x EEM Decay Heat (MeV)

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20 most important nuclides at T=500 seconds





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